

# GWANW 2023 Group Update WWU

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# Western Washington University (WWU)



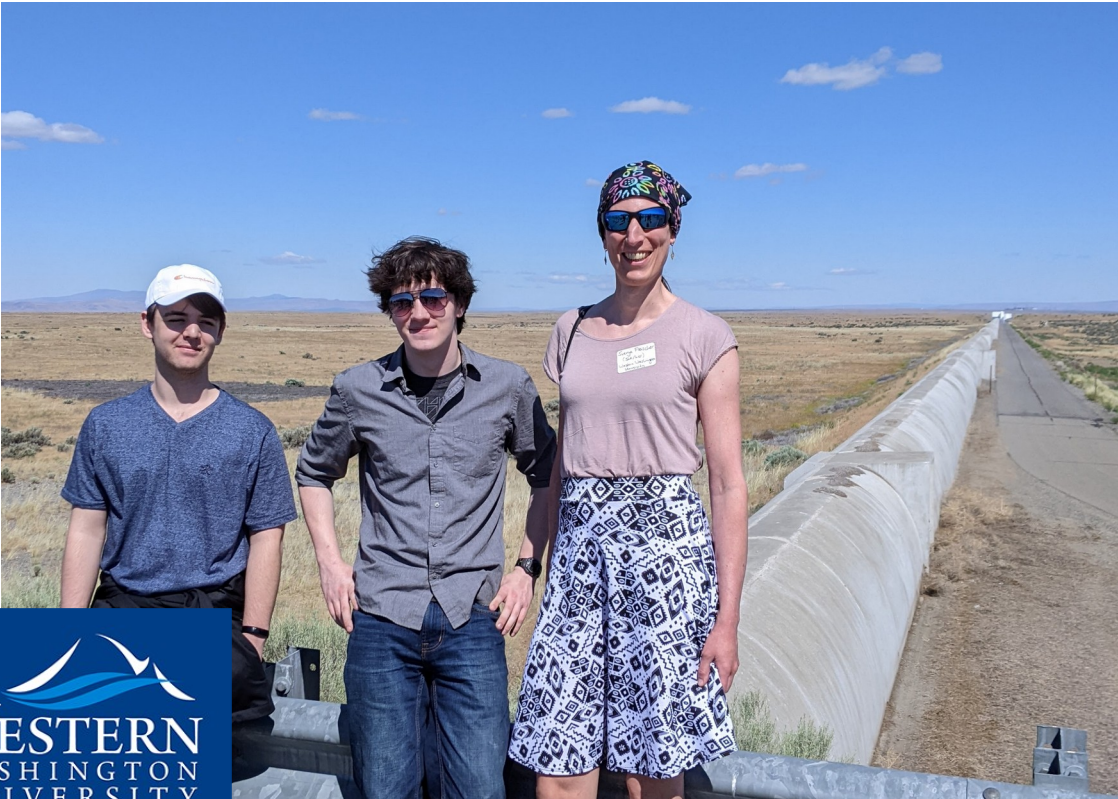
The Salish Sea



MAKE WAVES.

- Public university in Bellingham, WA, primarily undergraduate institution, ~16,000 students
- northernmost university in the Lower 48

# LIGO @ Western Washington University



- WWU has a LIGO group since 2021: SF and currently 4 undergraduate students
- also some torsion balance work with UW's Eotwash group

not pictured:

- Leah Vizmeg
- Veronica Russ

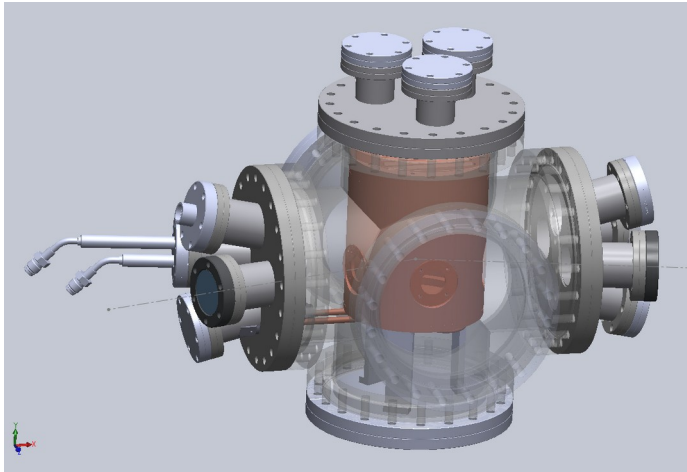


MAKE WAVES.



# Icy mirrors in cryogenic GW detectors

- Currently building an optical cryostat to study ice films on cold optical surfaces (and possible mitigation strategies)



- Significant progress in lab setup (SF started at WWU in fall 2020)

# Background: Cryogenic GW detectors

Operating future gravitational wave detectors at cryogenic temperatures has some very attractive benefits:

- improved material properties (given the right materials)
- e.g.: high thermal conductivity of mirror substrate allows for much greater circulating laser power
- decreased thermal noise

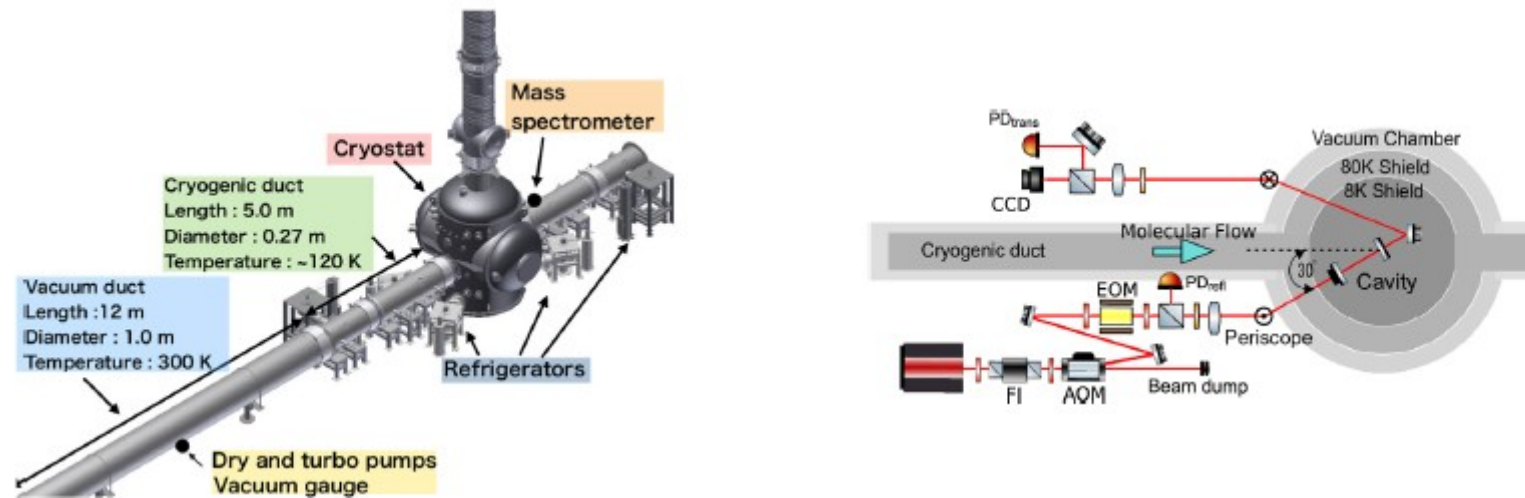
On the other hand: many new challenges!

**Current cryogenic GW detector concepts:**

KAGRA, Einstein Telescope (ET), LIGO Voyager

# Background: Ice layers on cold optics

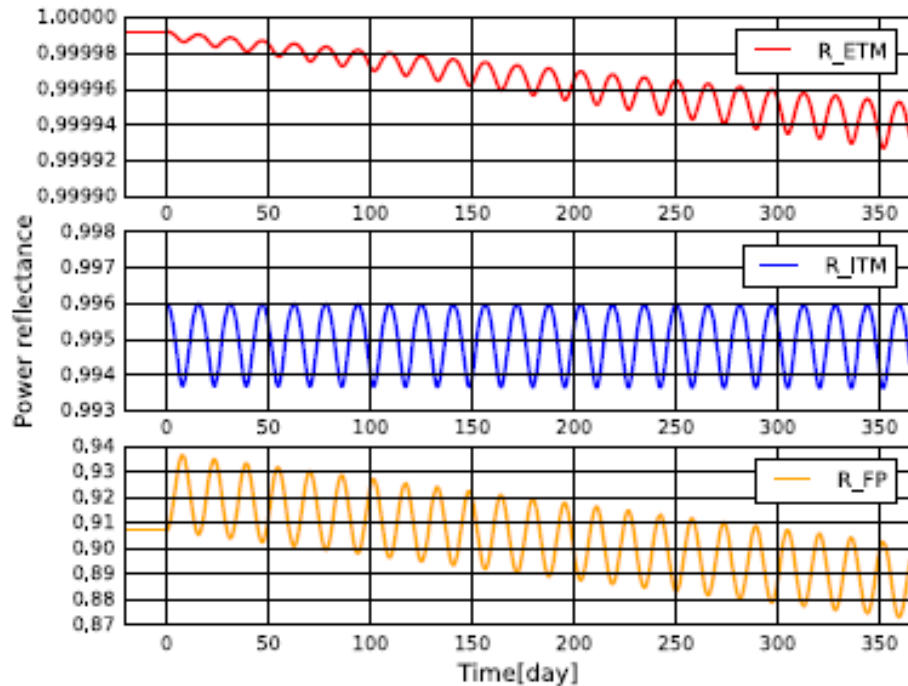
- Even in ultra-high vacuum, residual gas will freeze out on the cold surfaces – initially observed at KAGRA:  $27 \pm 1.9$  nm/day
- Mostly amorphous water ice
- Very limited experimental data available on amorphous water ice



from Hasegawa et al., *Phys. Rev. D* **99**, 022003 (2019)

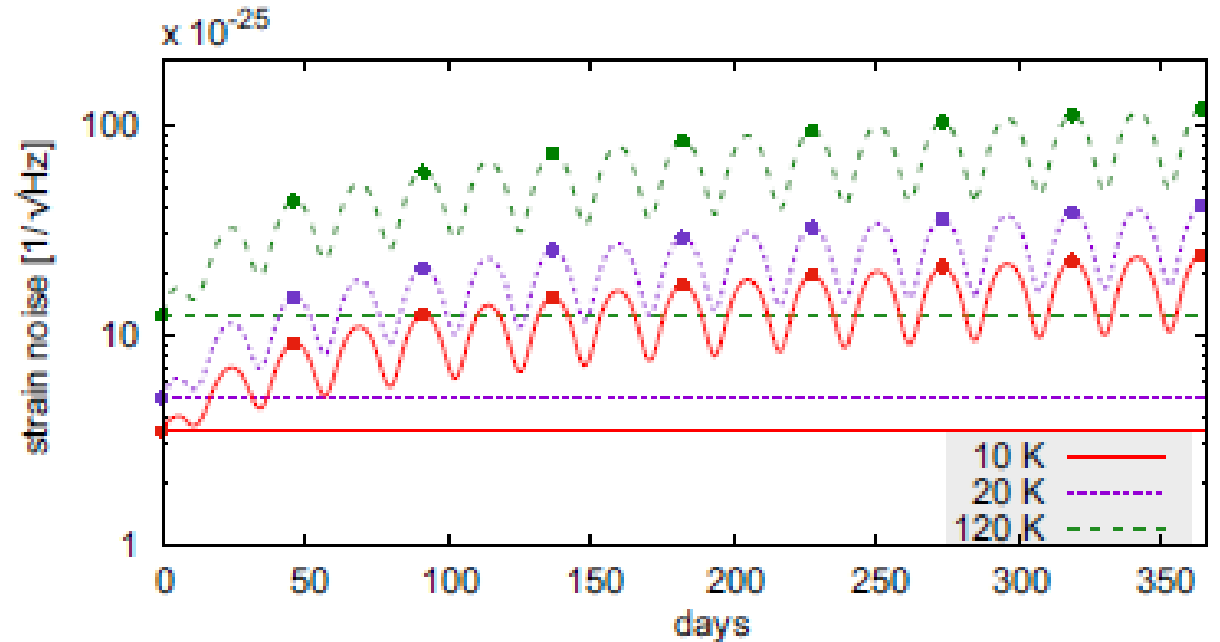
# Effects of Ice Layers I

- changing reflectance



from Hasegawa *et al.*, *Phys. Rev. D* **99**, 022003 (2019)

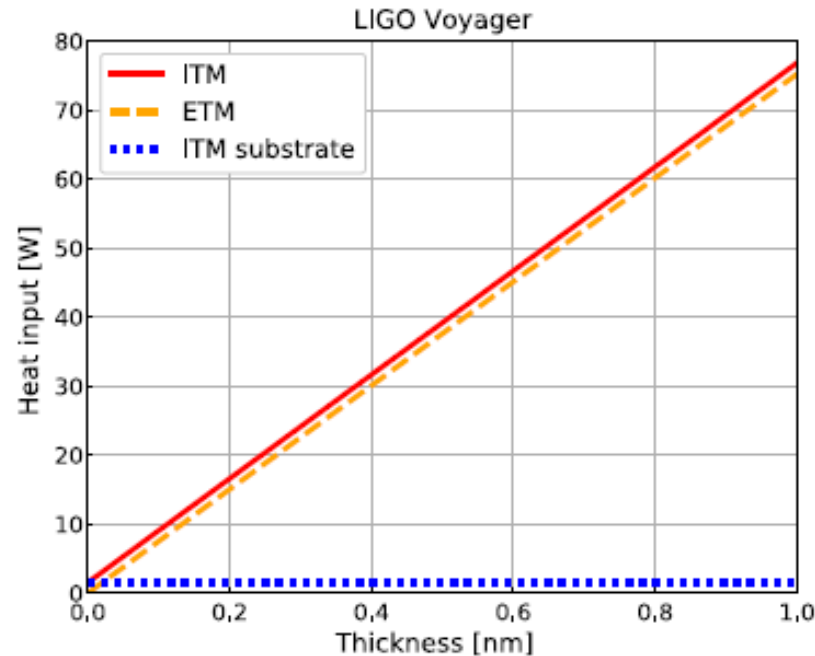
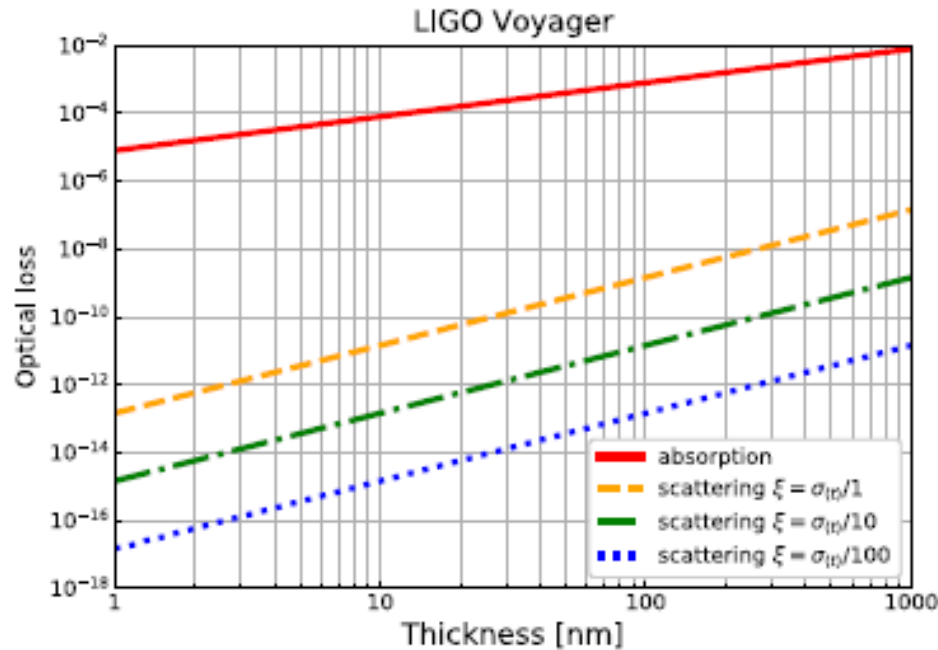
- increased coating thermal noise



from J. Steinlechner, I. Martin,  
*Phys. Rev. Res.* **1**, 013008 (2019)

# Effects of Ice Layers II

- probably most concerning: optical absorption @ 2 $\mu\text{m}$



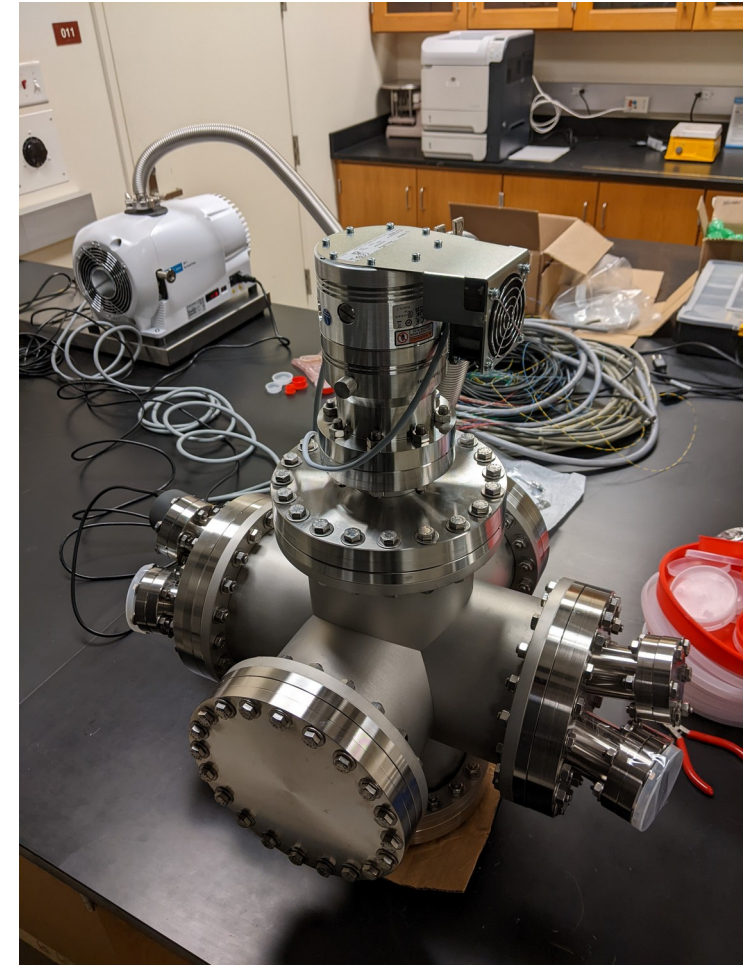
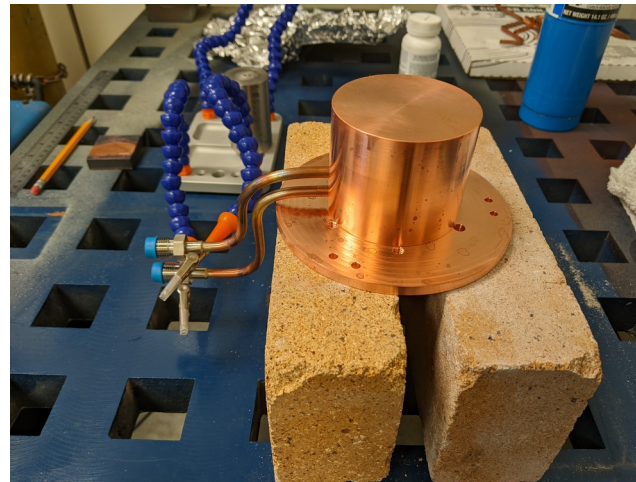
from Tanioka *et al.*,  
*Phys. Rev. D.* **102**,  
 022009 (2020)

For LIGO Voyager, this means that a thin layer of ice could increase the heat input by a factor of 1000. This appears to overestimate the absorption in thin layers – but still: maximum tolerable ice layer estimated to be around 20-30nm!



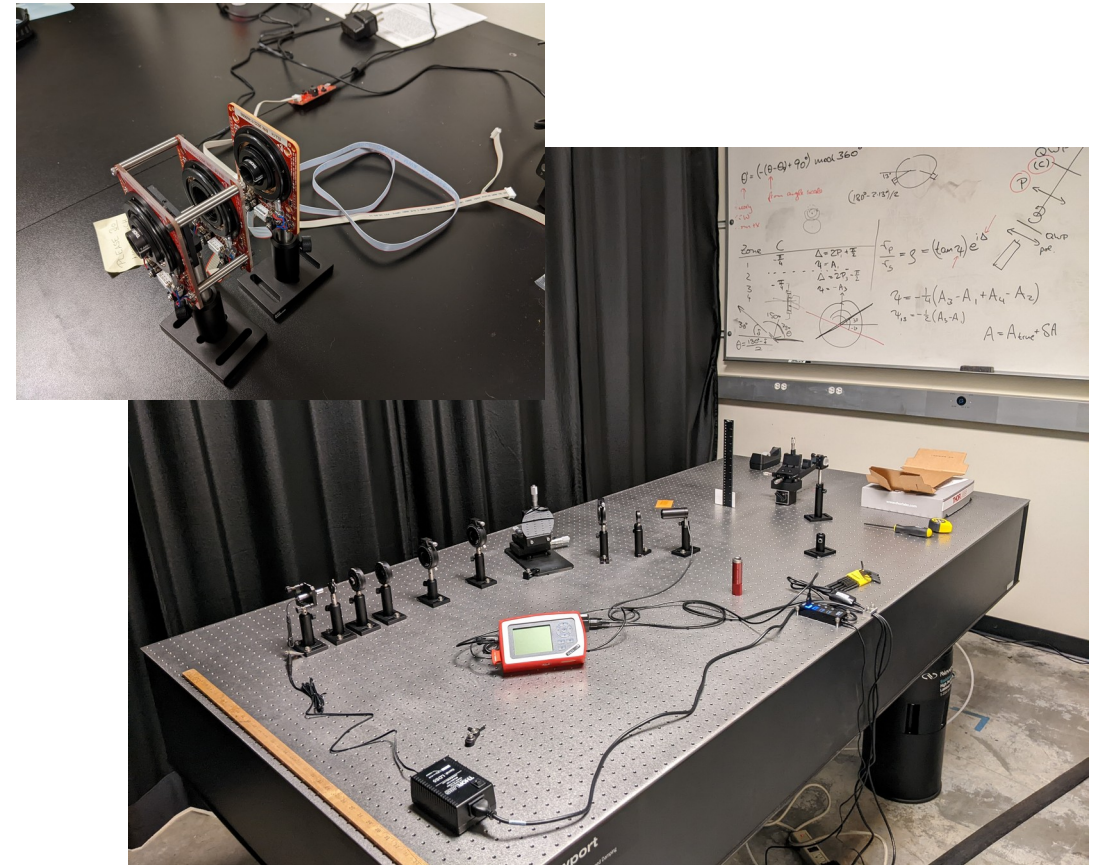
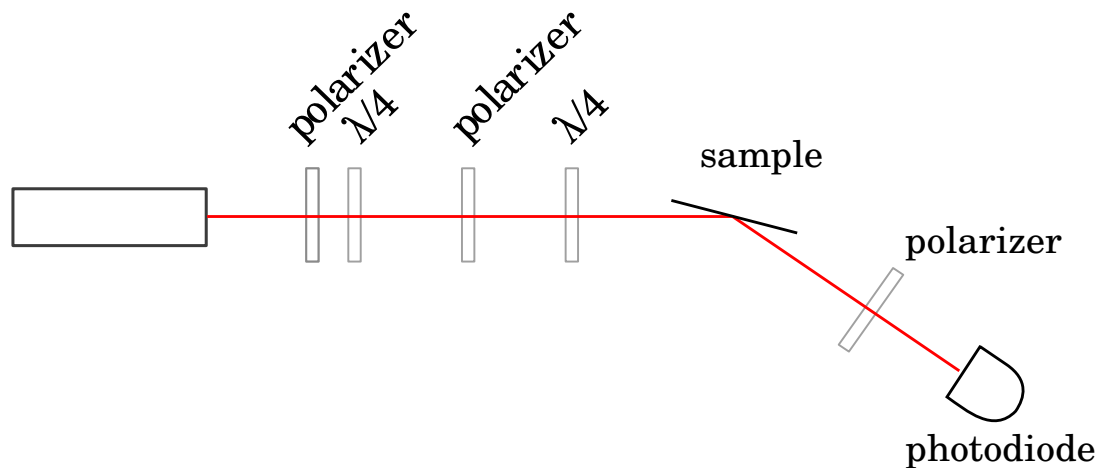
# Optical Cryostat: Current Status

- Ultra-high vacuum system to allow for ice film growth under controlled conditions:
  - vacuum chamber assembled
  - $7 \times 10^{-9}$  mbar without baking
- Thermal shield / LN2 system in progress



# Diagnostics: Monitoring Ice Layer Thickness

- Built a simple single-wavelength ellipsometer
- Ellipsometry can optically measure (sub-)nanometer layers:
- Currently testing with calibration target (13.3nm SiO<sub>2</sub> on Si wafer)

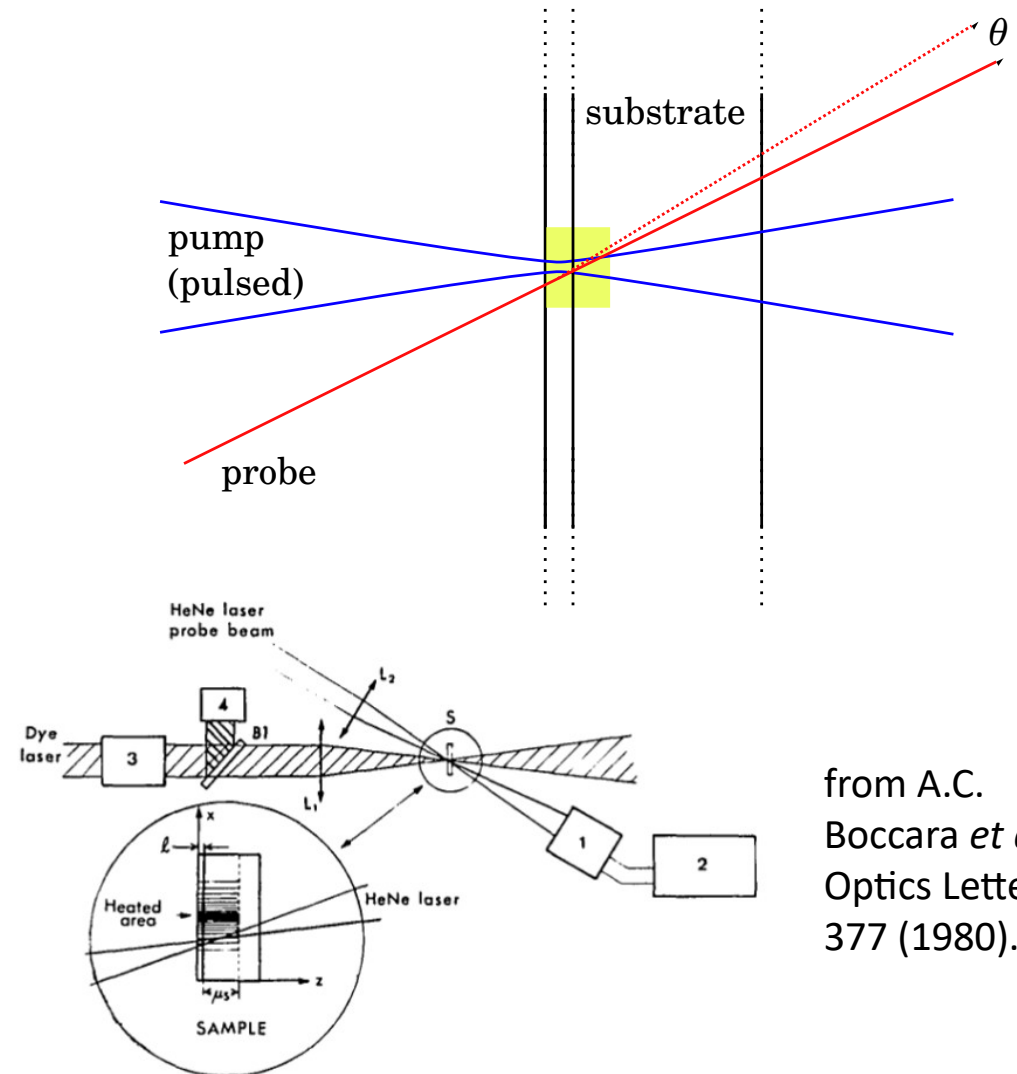


# Optical Cryostat: Next Steps

- Add ion pump to vacuum system and bake out
- Assemble LN2 system and temperature sensing/control
- Continue ellipsometer tests, integrate with cryostat; check effects of window birefringence
- Look at ice buildup from residual gas, consider controlled dosing of clean water vapor

# Next Steps: Studying Ice Films on Cold Surfaces

- Little experimental data, various parameters of interest
- What to try first...?
- Very interesting would be a direct absorption measurement using photothermal deflection spectroscopy (PDS)...



from A.C.  
Boccara *et al.*,  
Optics Letters 5,  
377 (1980).



# Thank you for your attention!

Questions?

## Students on this project:

- Jackson Larsen
- Douglas Slater
- Leah Vizmeg

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