



# Testing General Relativity with Gravitational Wave Signals

**Noah Wolfe**, Jacob Golomb, Colm Talbot



# GR is wrong!

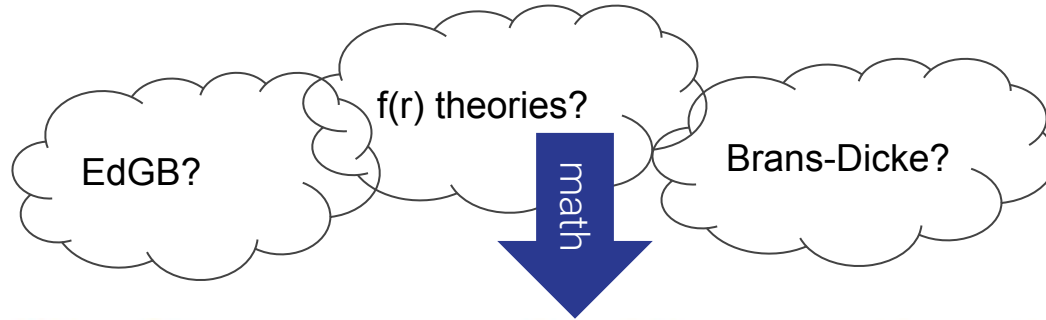
...in the signals I injected assuming GR is wrong

# Alternatives and Post-Newtonian Formalism

- General Relativity (GR) is successful in predicting GW properties
- Many alternative theories
- Use **generic post-Newtonian (PN) expansion** of GR metric



# Alternatives and Post-Newtonian Formalism



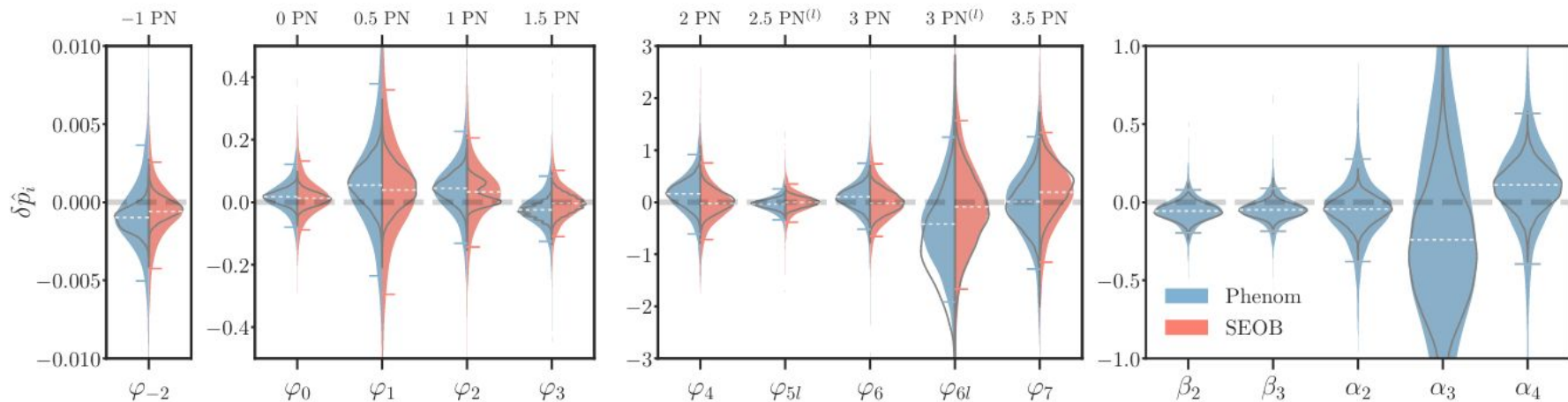
$$g_{00} = -1 + 2U - 2\beta U^2 - 2\xi\Phi_W + (2\gamma + 2 + \alpha_3 + \zeta_1 - 2\xi)\Phi_1 + 2(3\gamma - 2\beta + 1 + \zeta_2 + \xi)\Phi_2 + 2(1 + \zeta_3)\Phi_3 + 2(3\gamma + 3\zeta_4 - 2\xi)\Phi_4 - (\zeta_1 - 2\xi)\mathcal{A} - (\alpha_1 - \alpha_2 - \alpha_3)w^2U - \alpha_2 w^i w^j U_{ij} + (2\alpha_3 - \alpha_1)w^i V_i + \mathcal{O}((v^2/c^2)^3)$$



$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)} \quad \psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 \varphi_i v^i$$

Approximants like TaylorF2

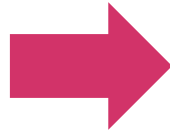
# Strong-Field Tests of General Relativity



Abbott et al. (2021)

# Hybrid Sampling

Simple

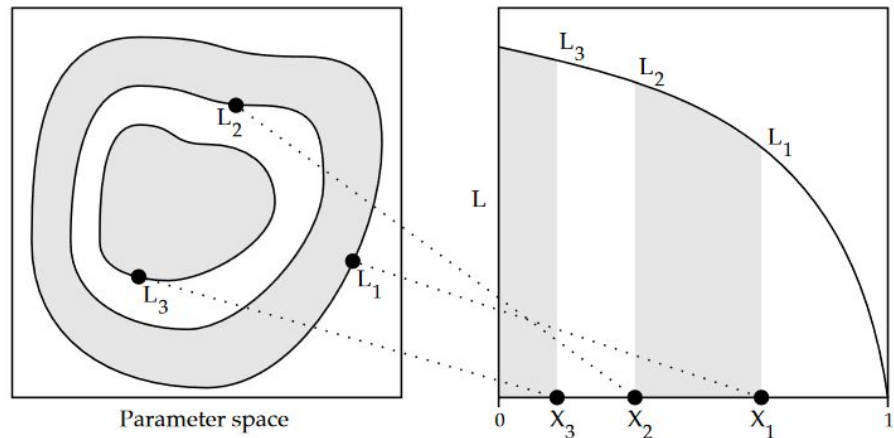


Complex



# Nested Sampling

- Fast to initialize
- We know it works!
- Fixed number of samples
- Can't start from other points



Skilling (2006)

# Monte Carlo Markov Chains

- Random walker explores parameter space to reconstruct posterior
- Possible to continue  $\sim$ indefinitely
- Can initialize at any point in the space\*
  - Fast if **well-initialized**





# Ensemble Monte Carlo Markov Chains

- ~~Random walker explores parameter space to reconstruct posterior~~
- Random **walkers** explore parameter space to reconstruct posterior
- Possible to continue ~infinitely
- Can initialize at any point in the space\*
  - Fast if **well-initialized**



# Parallel-Tempered Ensemble Monte Carlo Markov Chains

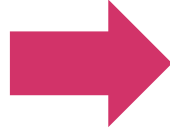
- ~~Random walker explores parameter space to reconstruct posterior~~
- ~~Random **walkers** explore parameter space to reconstruct posterior~~
- Random **walkers**, at different “temperatures”, explore parameter space to reconstruct posterior
- Possible to continue ~infinitely
- Can initialize at any point in the space\*
  - Fast if **well-initialized**



# Back to GR!

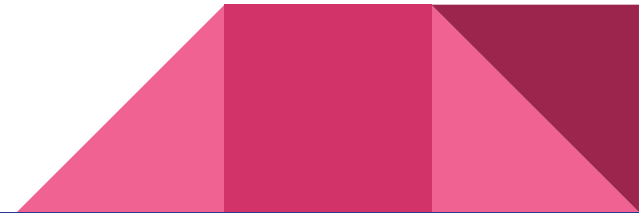
Simple  
GR

(almost?) correct



Complex  
beyond-GR

possibly more  
correct



# Back to GR!

Simple  
GR

(almost?) correct

$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)}$$
$$\psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 \varphi_i v^i$$

$$v = \left( \pi f \frac{GM}{c^3} \right)^{1/3}$$



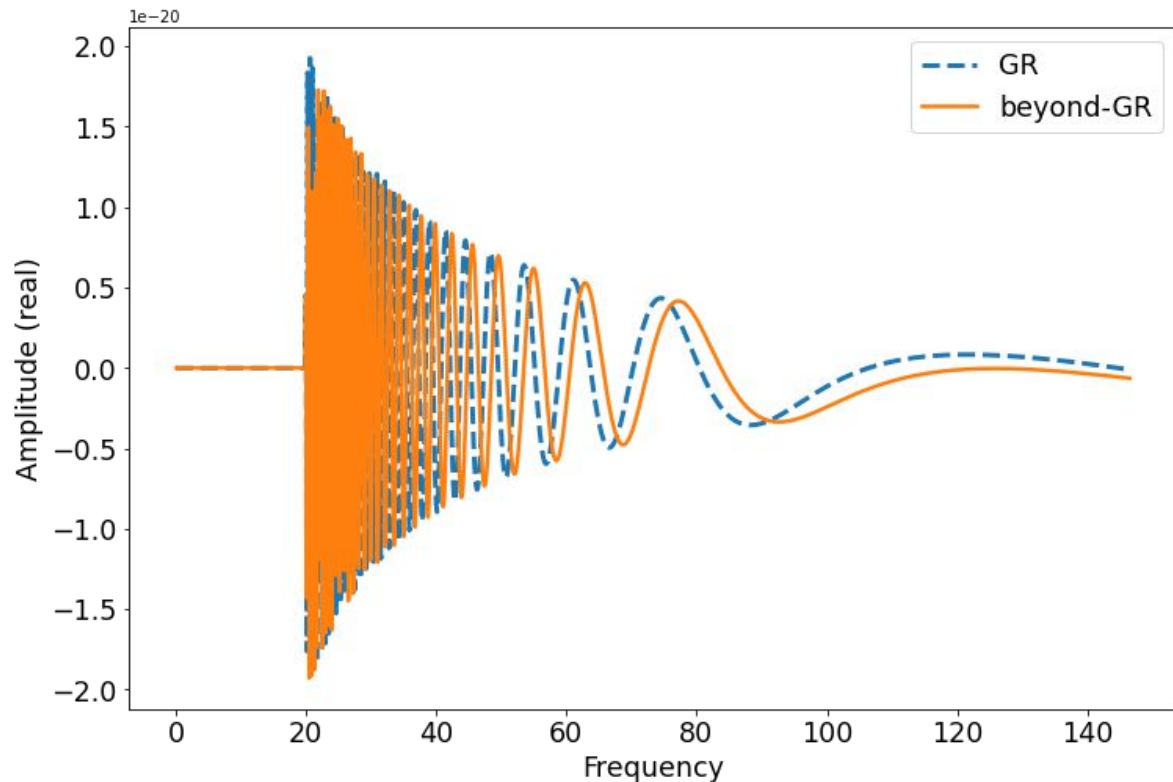
Complex  
beyond-GR

possibly more  
correct

$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)}$$
$$\psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 (\varphi_i + \delta\varphi_i) v^i$$

# Waveform Generation

- $d_L = 200$  Mpc
- $q = 1$
- $\delta\varphi_1 = 10^{-5}$
- No spin
- Chirp masses in  $\{8, 15, 30\}$



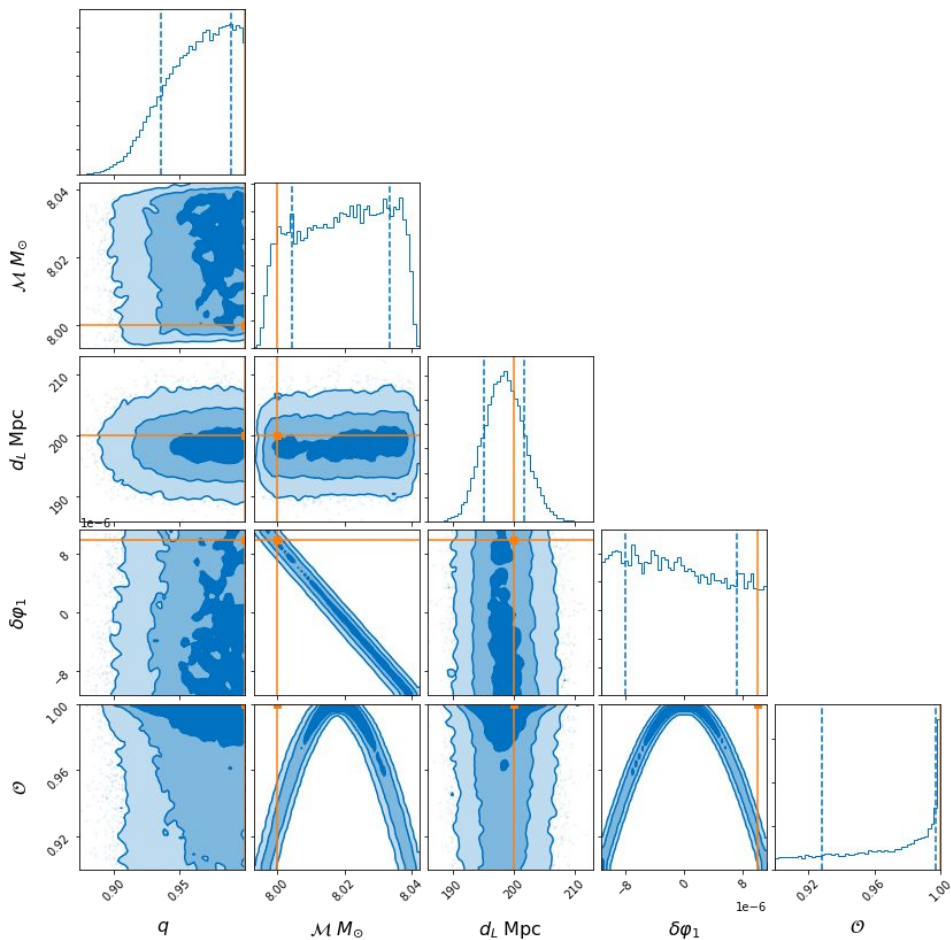
# Initial Results

- Limited set of parameters
- Check that hybrid sampling returns the same as dynesty
- Investigate chirp mass, signal-to-noise ratio dependence
- Prior limit of  $\mathcal{O} \geq 0.9$



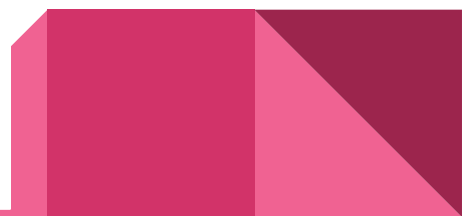
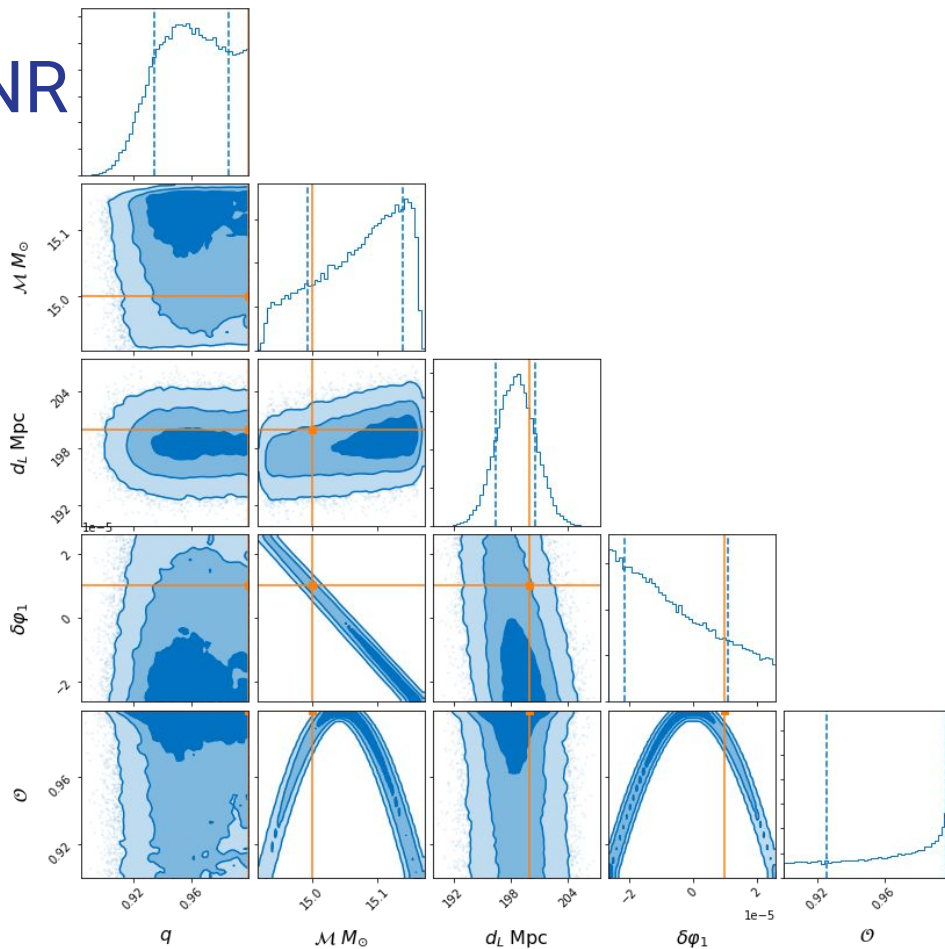
“Low” SNR

Chirp Mass = 8  
SNR  $\sim$  50



# “Medium” SNR

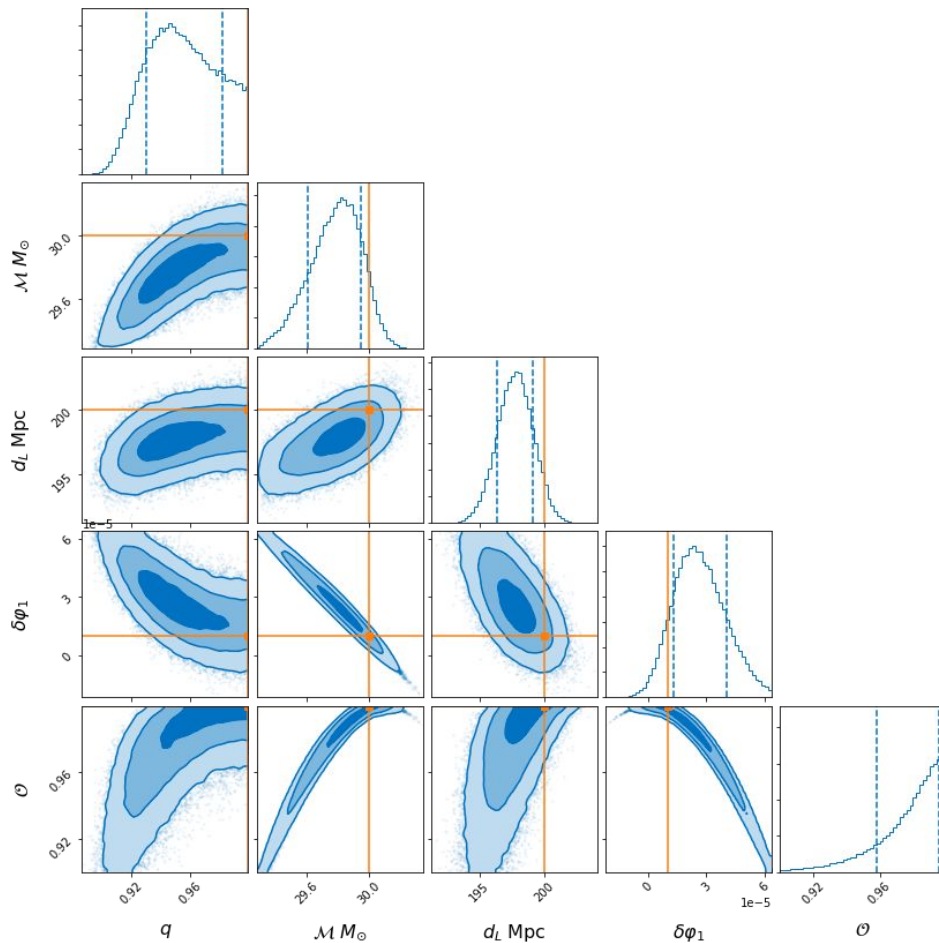
Chirp Mass = 15  
SNR  $\sim$  80





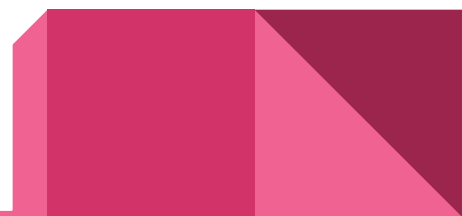
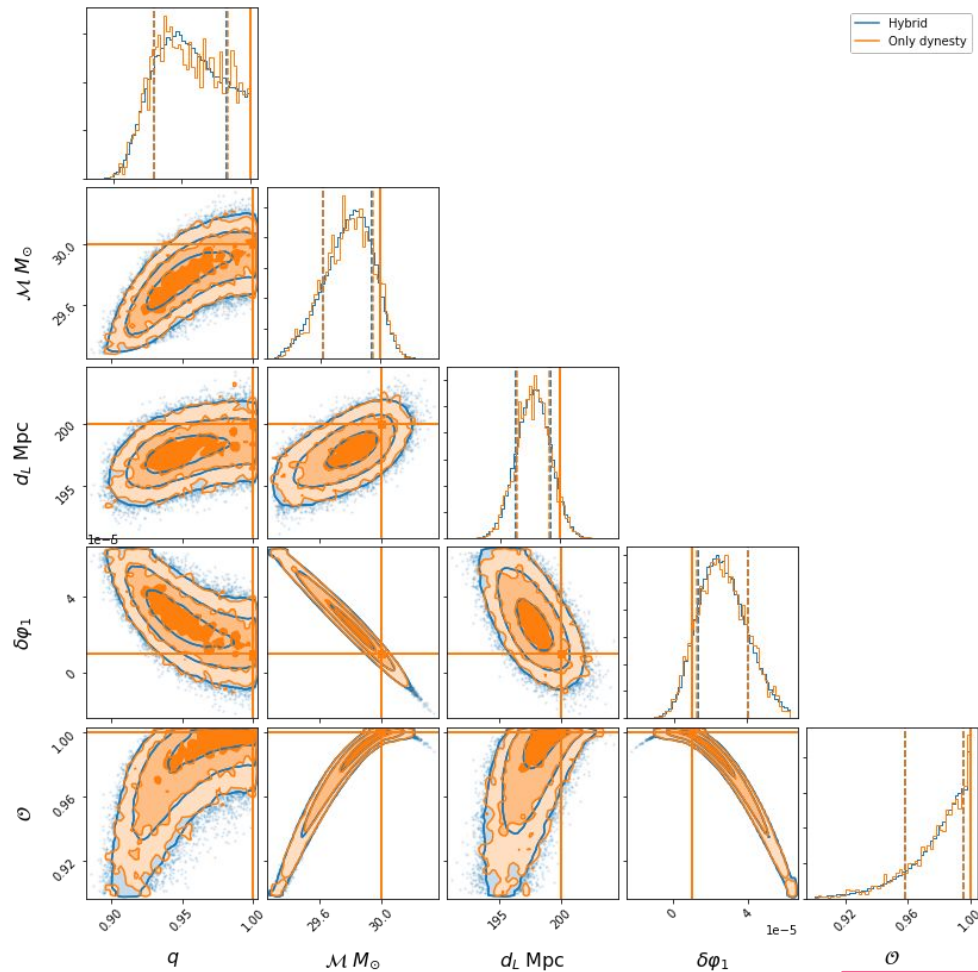
# High SNR

Chirp Mass = 30  
SNR ~ 140



# High SNR

Chirp Mass = 30  
SNR ~ 140



# Summary

- ✓ Generate beyond-GR phase variations
- ✓ Test hybrid parameter estimation scheme
- ✓ Implement modified source model with beyond-GR waveform
- ✓ Compare hybrid sampling & dynasty on beyond-GR waveform
- Hybrid sampling with all GR parameters
- Hybrid sampling with real data
- Vary multiple deviation coefficients simultaneously
- 🤔 Prove GR wrong

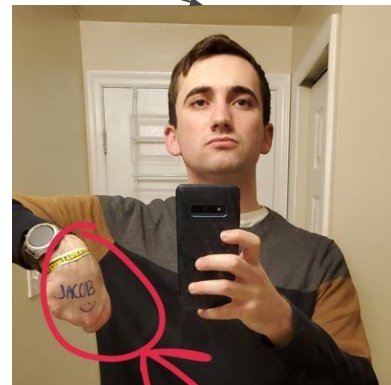
# Acknowledgements

Paid me

# Caltech



Paid me  
(in knowledge)



Me (about to be taught  
nested sampling)

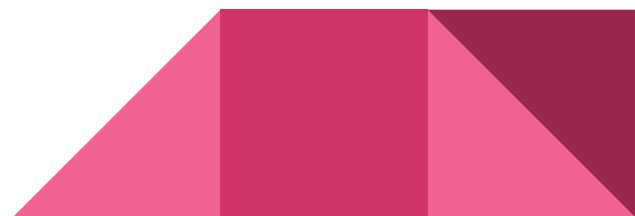
LIGO-T2100339

# Backup Slides

- Overlap function

$$\mathcal{O} = \frac{\langle \tilde{h}_1(f), \tilde{h}_2(f) \rangle}{\sqrt{\langle \tilde{h}_1(f), \tilde{h}_1(f) \rangle \langle \tilde{h}_2(f), \tilde{h}_2(f) \rangle}}$$

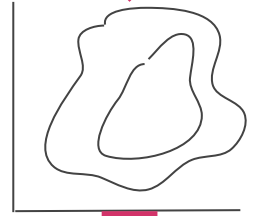
$$\langle \tilde{h}_1(f), \tilde{h}_2(f) \rangle = 4\Delta f \sum_i^N \frac{\tilde{h}_{1,i} \tilde{h}_{2,i}^*}{s_i}$$



# Hybrid Sampling Scheme

- Nested sampling
  - Directly calculate the Bayesian evidence, then posteriors
  - Traveling up the likelihood surface, accumulating prior mass
- Ensemble Monte Carlo Markov Chain
  - Multiple walkers, in-concert, exploring the parameter space
- Possibility to ‘temper’ both methods
  - Increase the ‘temperature’ of the likelihood to more easily explore all modes

Nested sampling  
for mass, spin,  
etc.



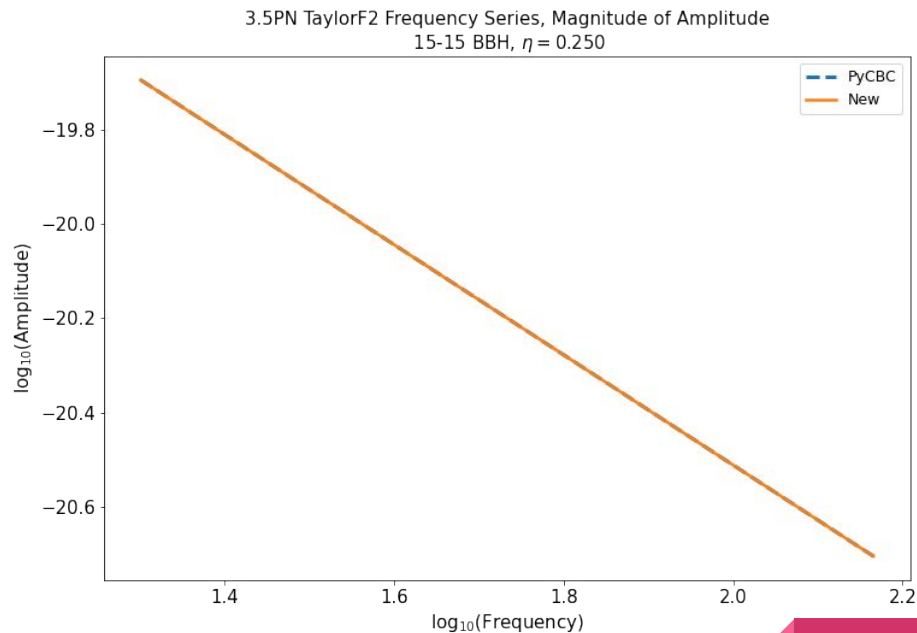
Ensemble  
MCMC for  
beyond-GR  
terms

# TaylorF2 Waveform Generation

$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)}$$

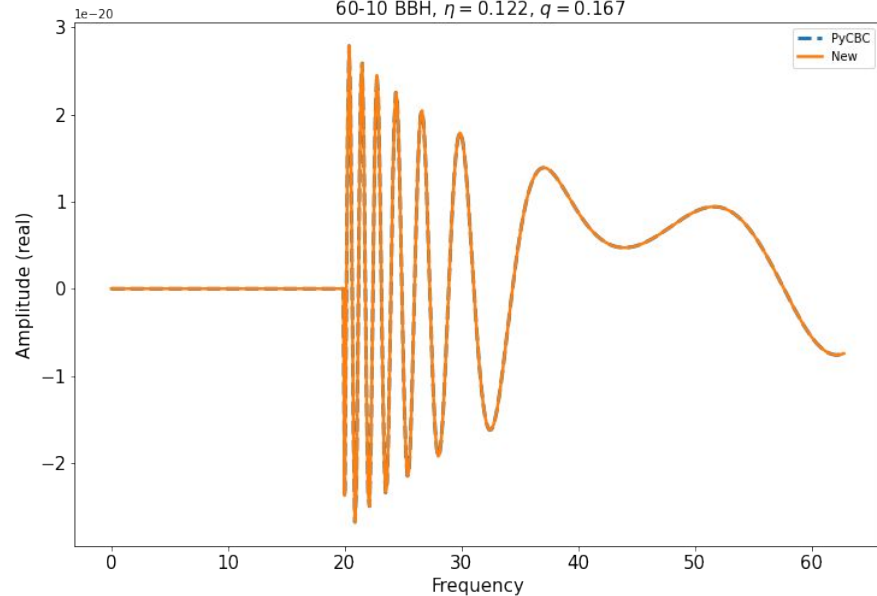
$$\psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 \varphi_i v^i$$

$$v = (\pi M f G c^{-3})^{1/3}$$

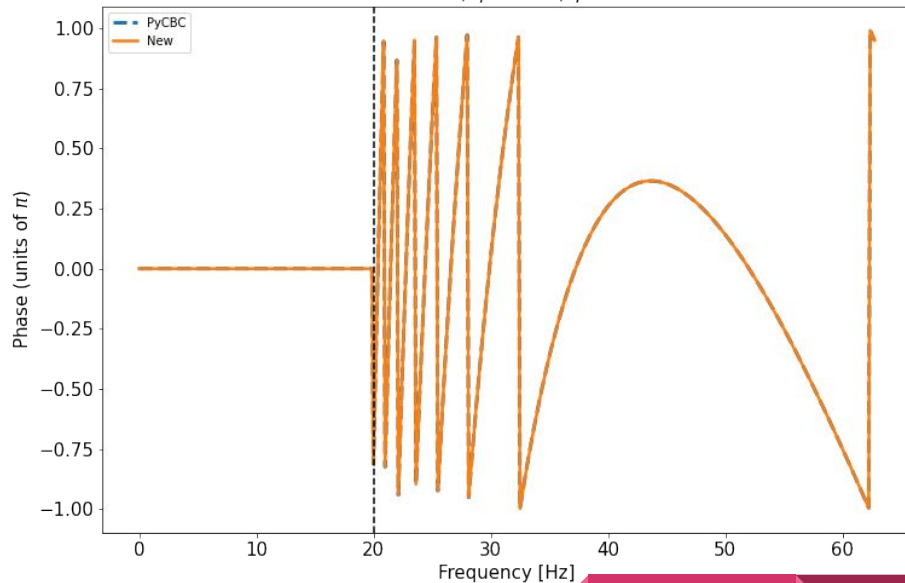


# Phase Evolution Agreement

3.5PN TaylorF2 Frequency Series, Real Part of Amplitude  
60-10 BBH,  $\eta = 0.122$ ,  $q = 0.167$

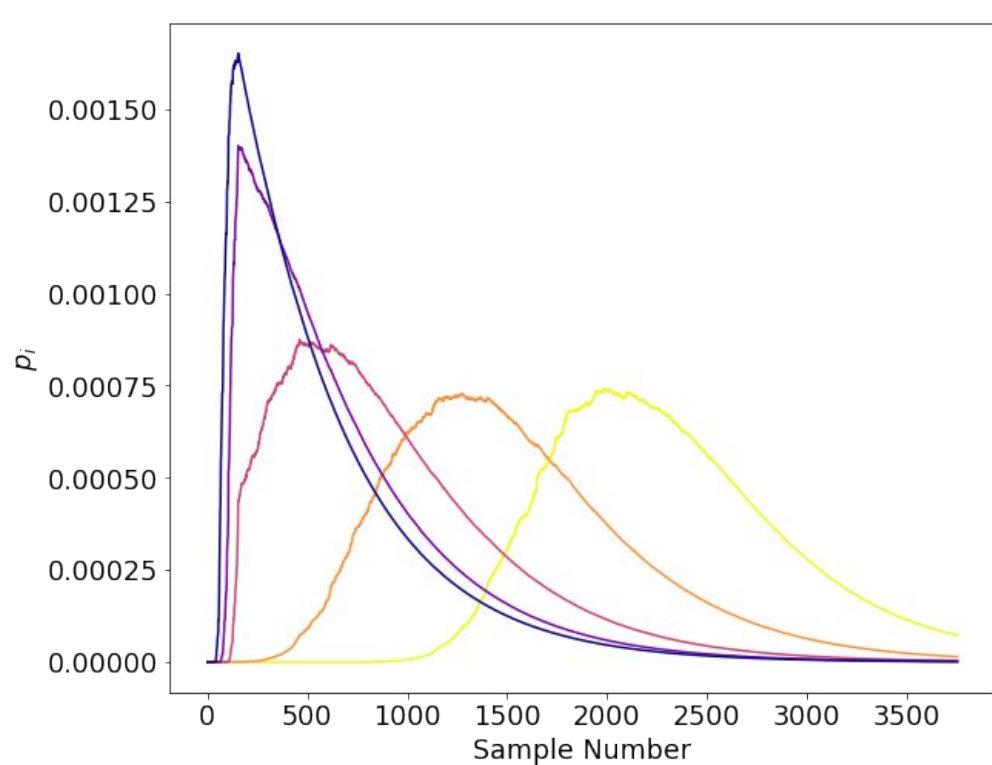


3.5PN TaylorF2 Frequency Series, Phase  
60-10 BBH,  $\eta = 0.122$ ,  $q = 0.167$

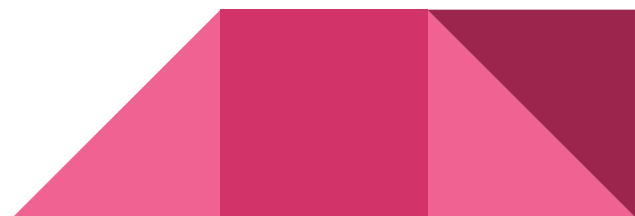




# Tempered Posterior Weights



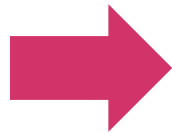
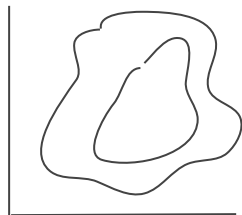
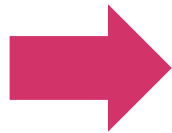
$$\beta_T p_{i, \beta_T} = \left( \frac{L_i w_i}{Z} \right)^{\beta_T}$$



# Hybrid Sampling

## Finish this?

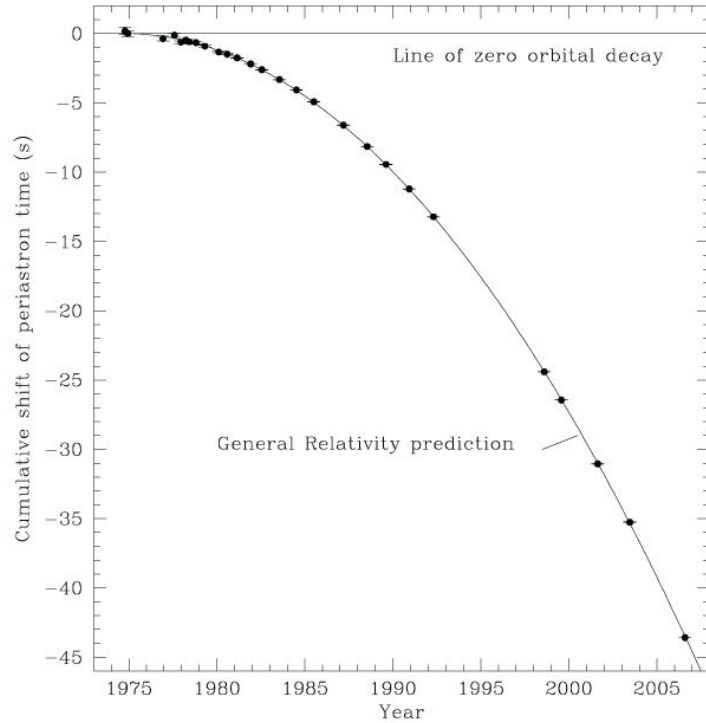
Simple  
Dynesty



Complex

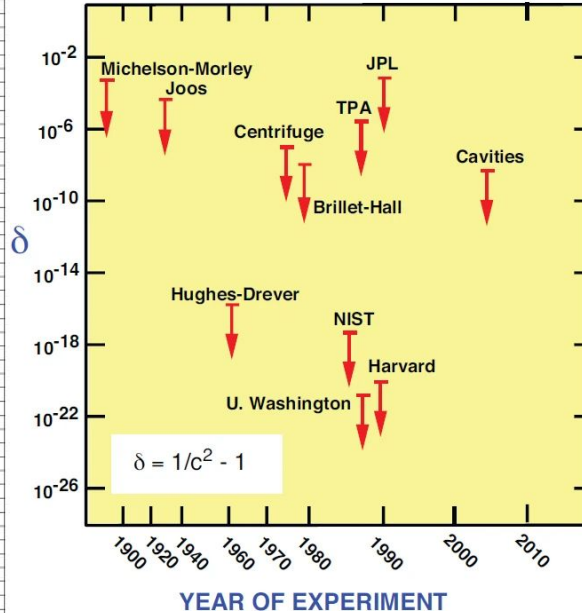


# Weak-Field Tests of General Relativity

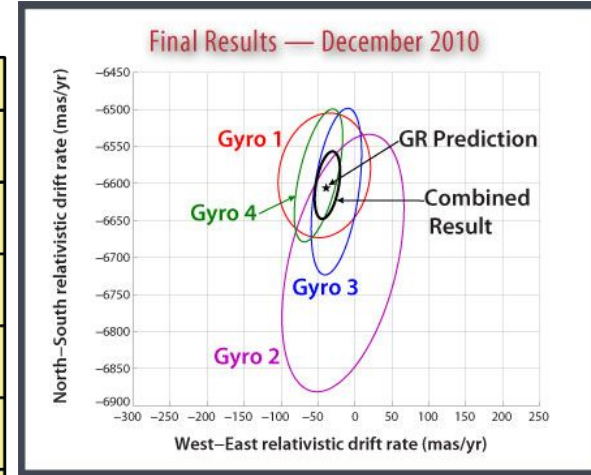


Weisberg et al. (2010)

## TESTS OF LOCAL LORENTZ INVARIANCE



Will (2014)



Everitt et al. (2011)

# Our Goal

- Previous test with GWTC-2
  - Only **varied one beyond-GR parameter** at a time
  - Required full parameter estimation on **15 GR parameters + 10 beyond-GR parameters**
- Hybrid nested sampling and ensemble monte carlo markov chain scheme
  - More efficient parameter estimation

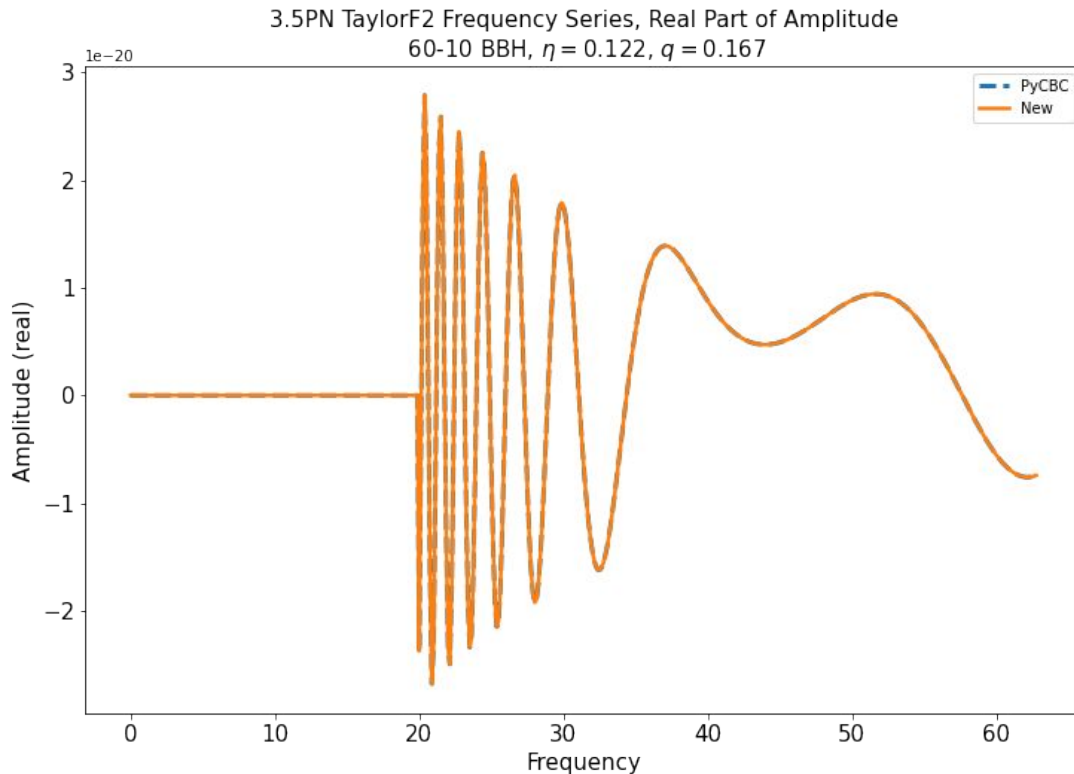


# TaylorF2 Waveform Generation

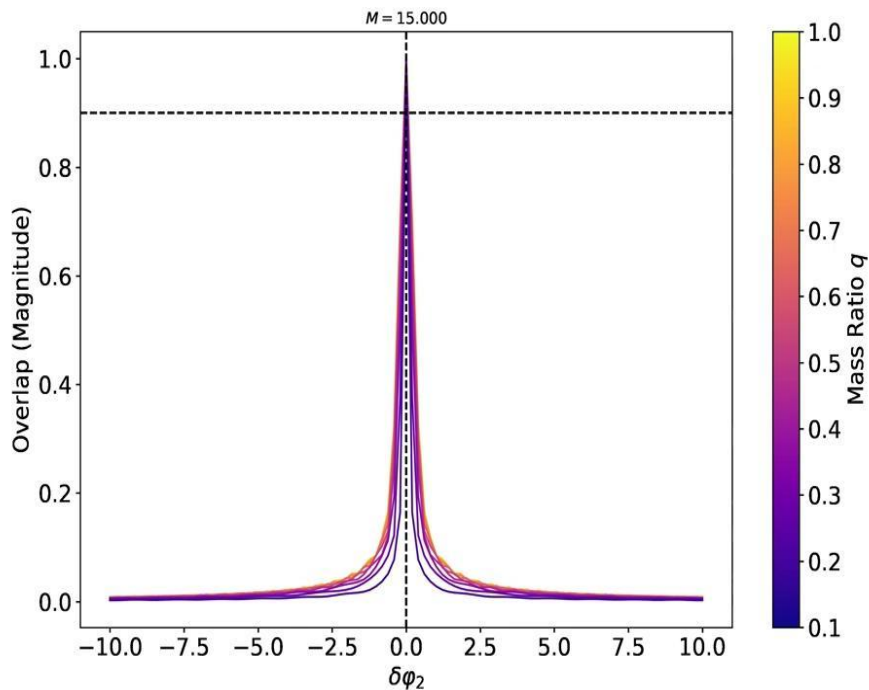
$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)}$$

$$\psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 \varphi_i v^i$$

$$v = (\pi M f G c^{-3})^{1/3}$$



# Overlap Phase Variations



$$\tilde{h}(f) = \mathcal{A} f^{-7/6} e^{i\psi(f)}$$

$$\psi(f) = -\frac{\pi}{4} + \frac{3}{128\eta v^5} \sum_i^7 (\varphi_i + \delta\varphi_i) v^i$$

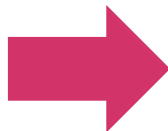
Left: variation of the  $i = 2$  term

## Toy Model

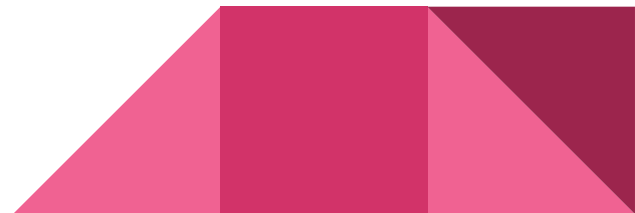
$$P(x) = \frac{1}{\sigma\sqrt{2\pi}} e^{-((x-\mu)/(\sigma\sqrt{2}))^2} \longrightarrow P(x) = \frac{\beta}{2\alpha\Gamma(1/\beta)} e^{-(|x-\mu|/\alpha)^\beta}$$

$\alpha = \sigma\sqrt{2}$

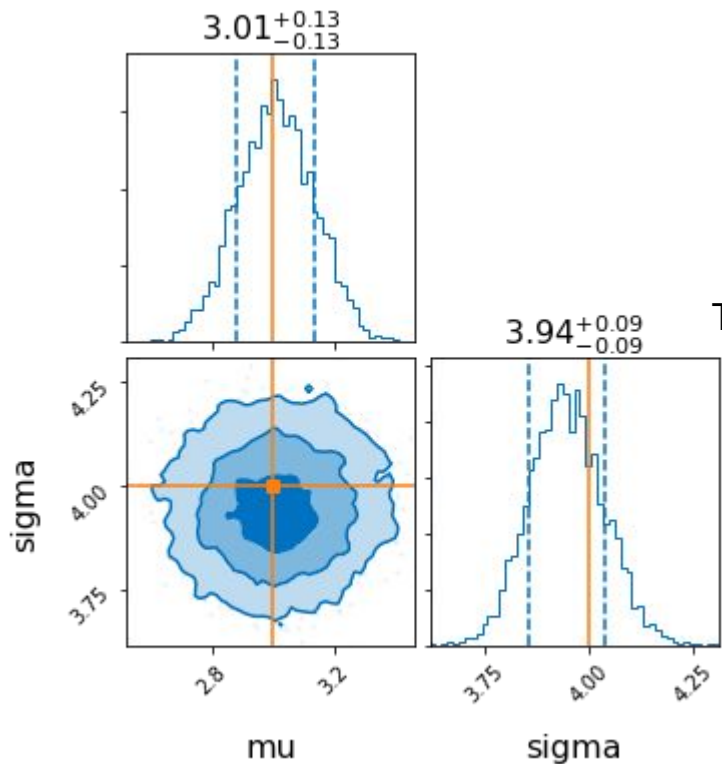
Sample in  $\mu, \sigma$   
(dynesty)



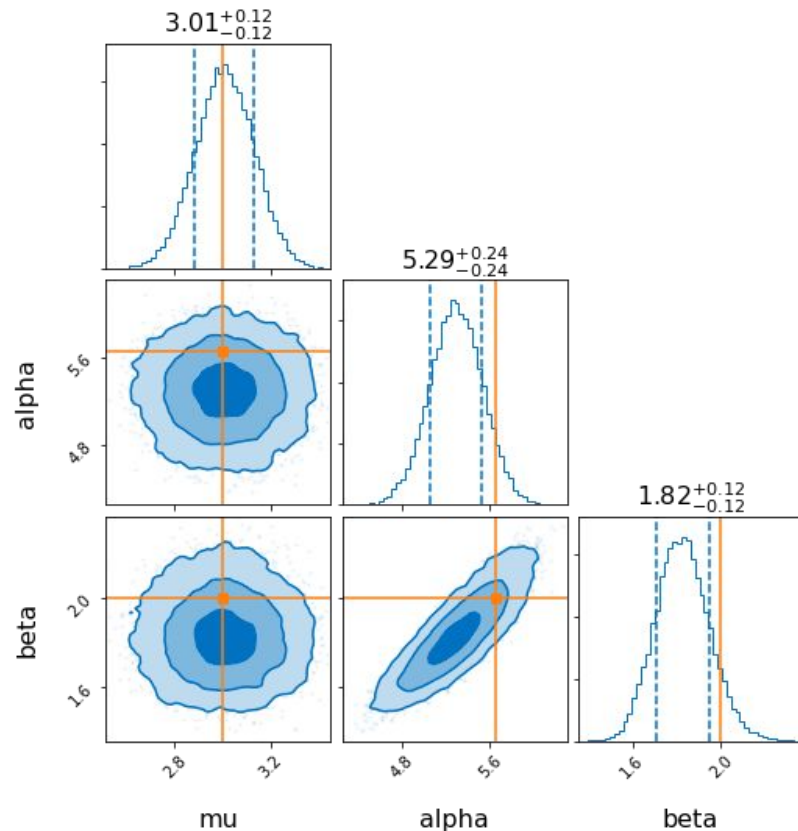
Sample in  $\mu, \alpha, \beta$   
(ptemcee)



# Toy Model Results

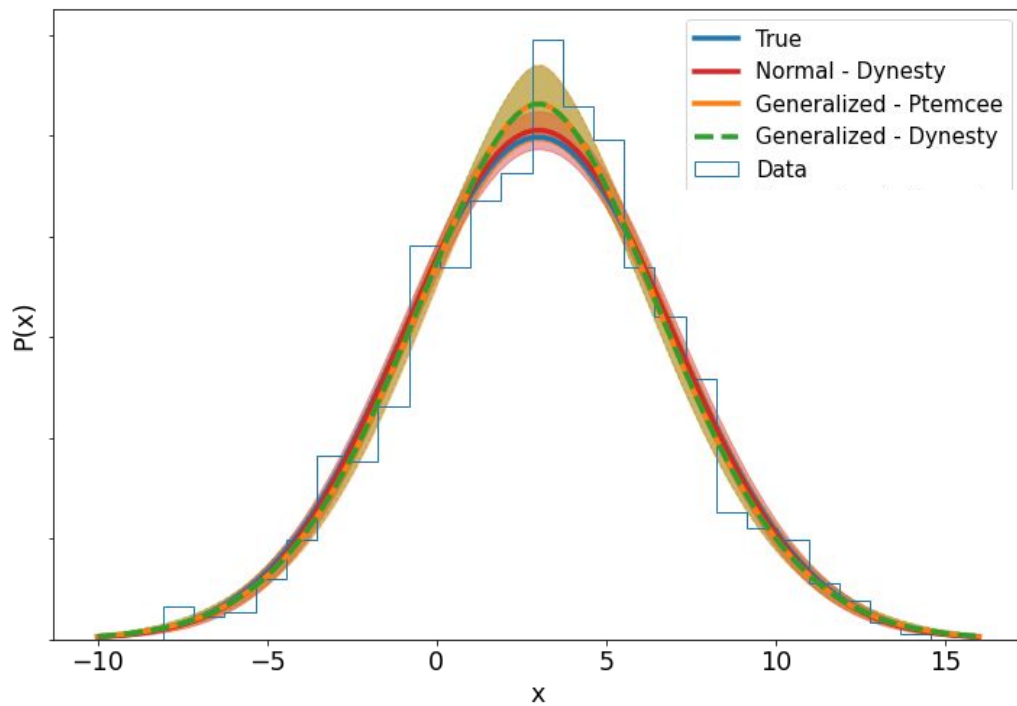
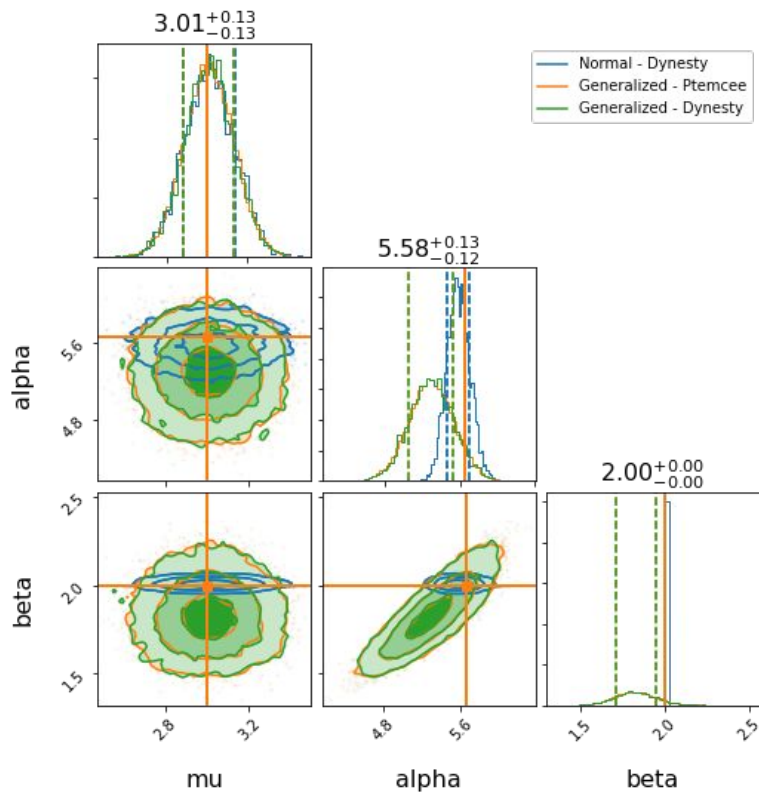


Tempered Posterior

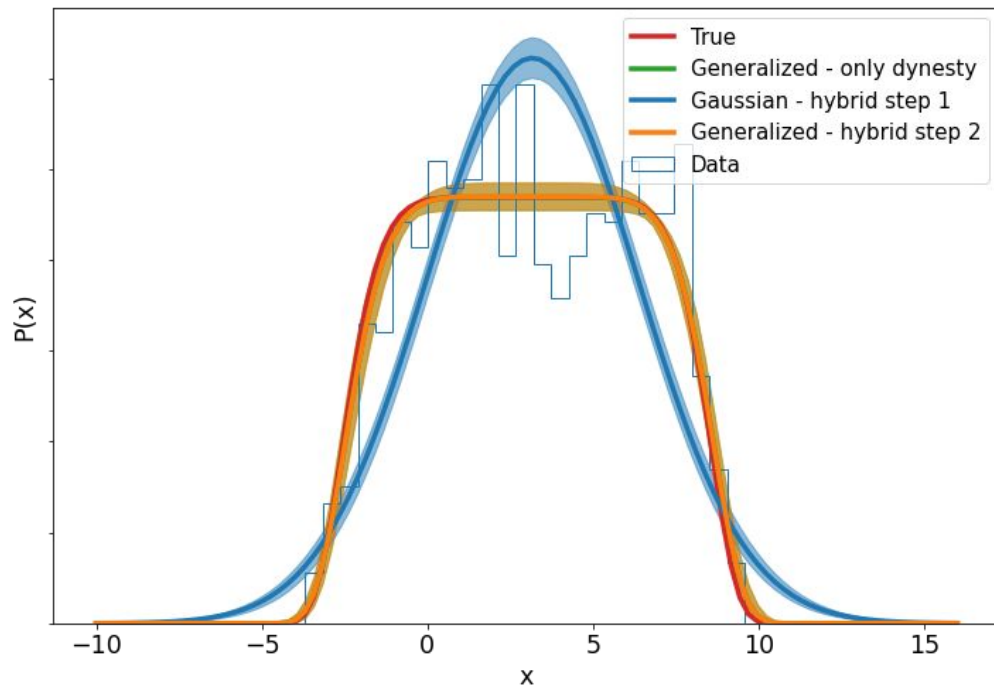
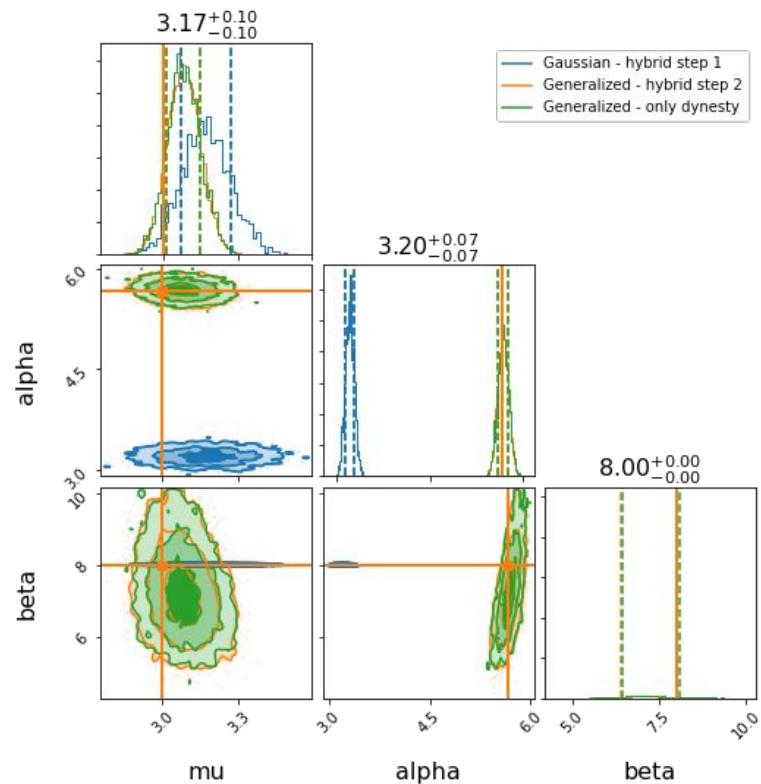





# Toy Model Results



# Model Misspecification



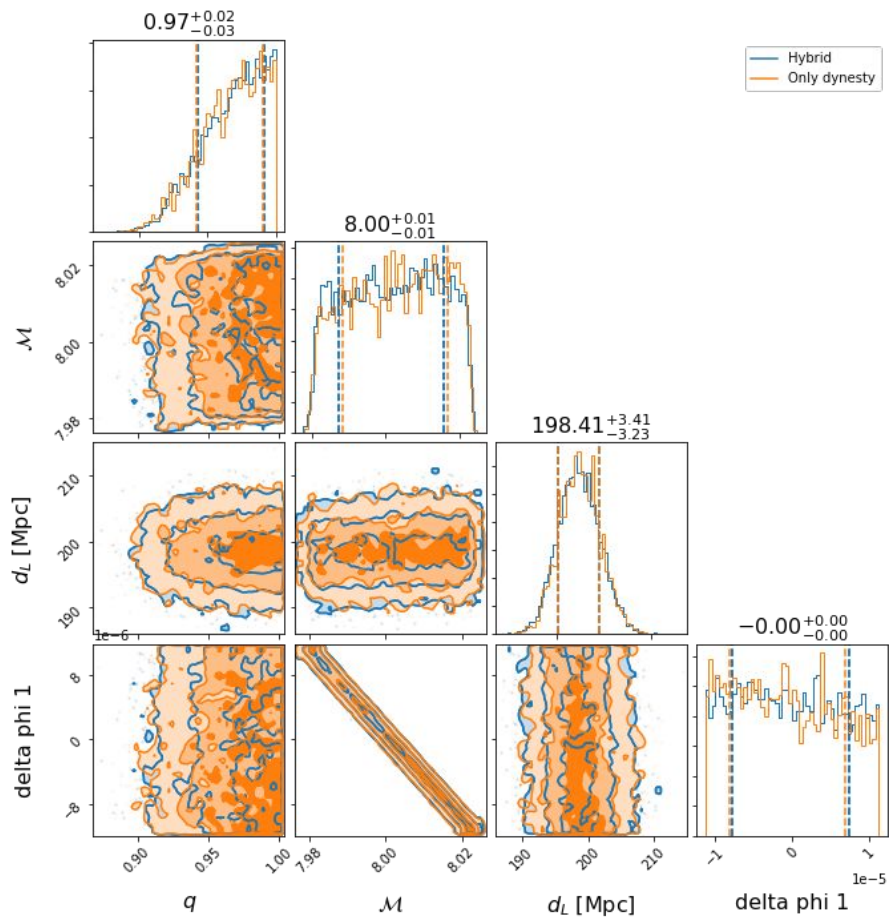
## Overlap Calculation

$$\mathcal{O} = \frac{\langle \tilde{h}_1(f), \tilde{h}_2(f) \rangle}{\sqrt{\langle \tilde{h}_1(f), \tilde{h}_1(f) \rangle \langle \tilde{h}_2(f), \tilde{h}_2(f) \rangle}}$$
$$\langle \tilde{h}_1(f), \tilde{h}_2(f) \rangle = 4\Delta f \sum_i^N \frac{\tilde{h}_{1,i} \tilde{h}_{2,i}^*}{s_i}$$


# CUT

“Low” SNR

SNR  $\sim$  50



# CUT

## “Medium” SNR

Chirp Mass = 15  
SNR  $\sim$  80

