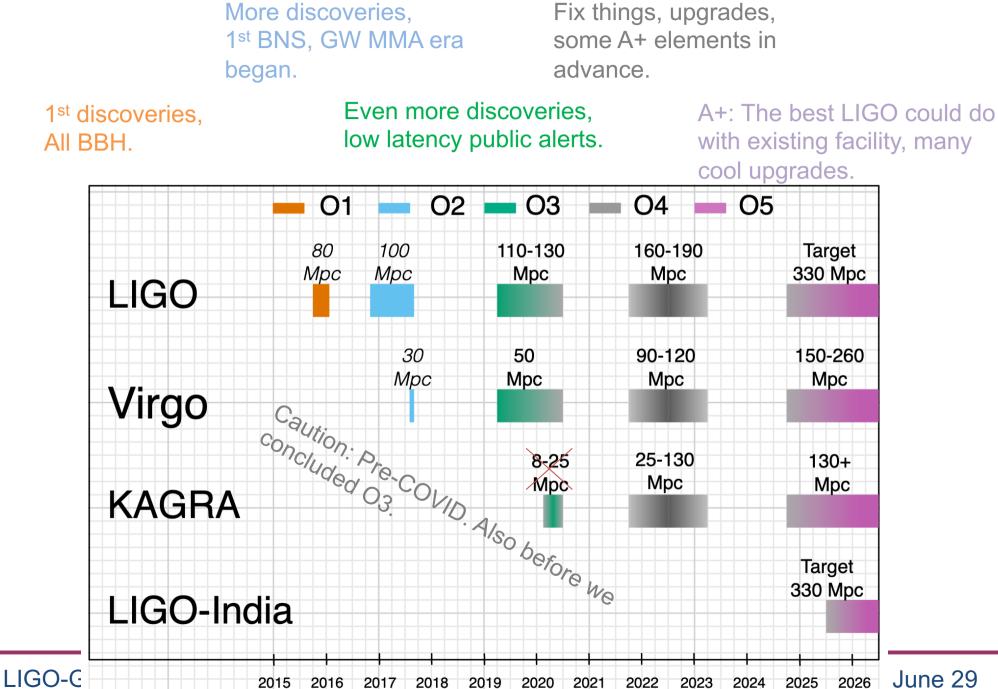
# LVK Status, June 2020

Keita Kawabe



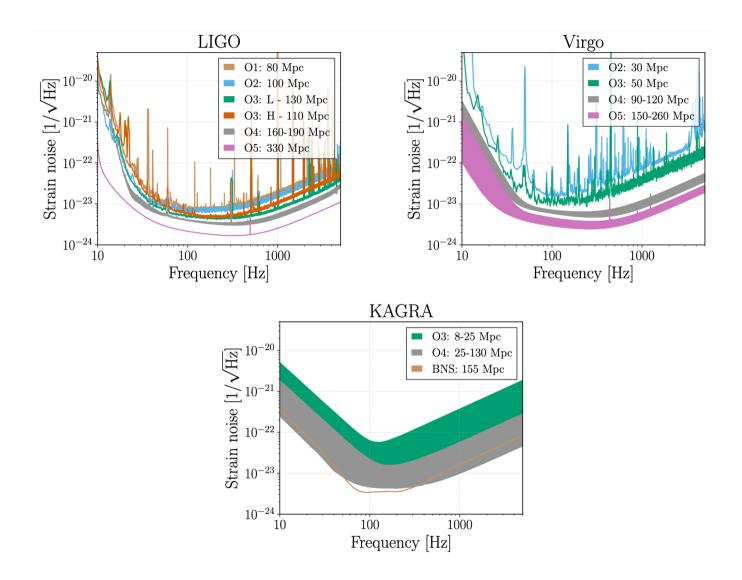
LIGO-G2000997

# **Big-ish picture**



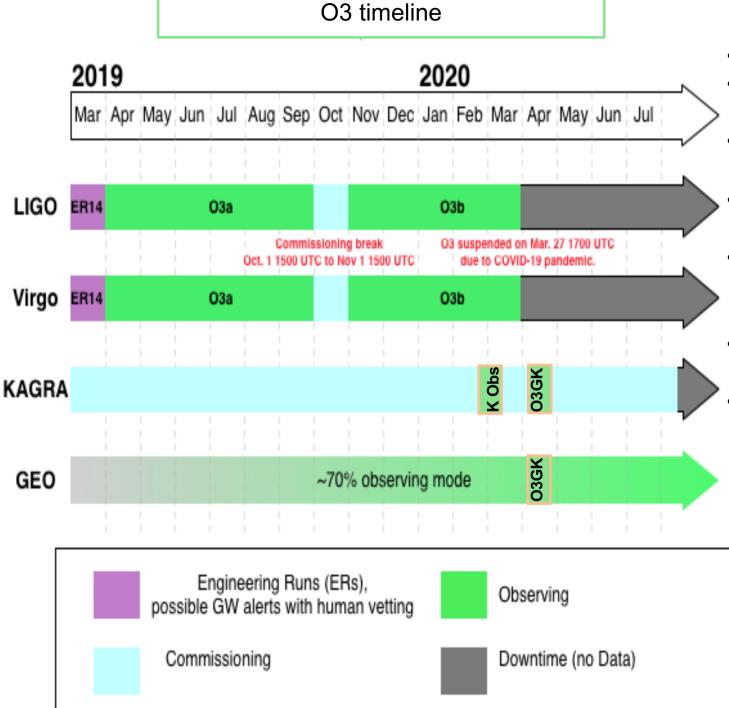
June 29

# (I'll skip this: Strain noise of the detectors for observation runs past and future)



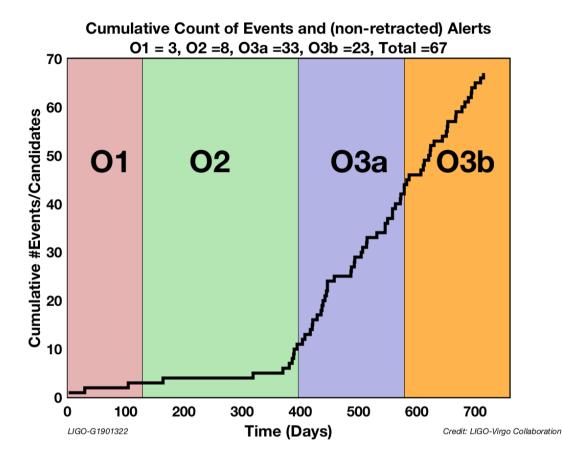
# I'll talk about

- LIGO-centric report of O3.
- Future, COVID, how LIGO, Virgo and KAGRA are working (or not).



- O3 began April 1, 2019
- 1 month commissioning break in October 2019.
- Was scheduled to end on April 30, 2020.
- Suspended due to COVID-19 on March 27, 2020.
- KAGRA observation Feb.
  25 Mar. 10, 2020 (not a part of O3.)
- GEO-KAGRA short run
  O3GK, Apr. 7-21, 2020.
- June 18, 2020: Decision not to pursue resuming O3.

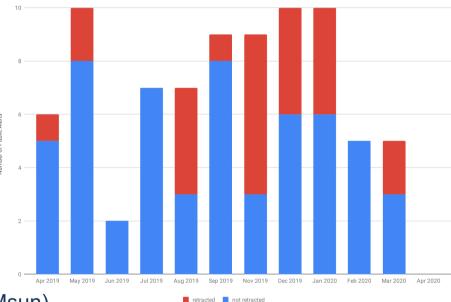
# O3: A great success! Many public alerts, many discoveries.



- <u>GW190412</u>: Large mass ratio BBH (~30 and ~8 Msun).
- <u>GW190425</u>: Likely BNS @ ~ 160+-70 Mpc.
- <u>GW190814</u>: "Mass gap" event (~2.6+-0.1 and ~23+-1 Msun).
- GW190521 near future.
- O3a catalog near future.

Total Public Alerts in O3	80
Not Retracted	56
Retracted	24

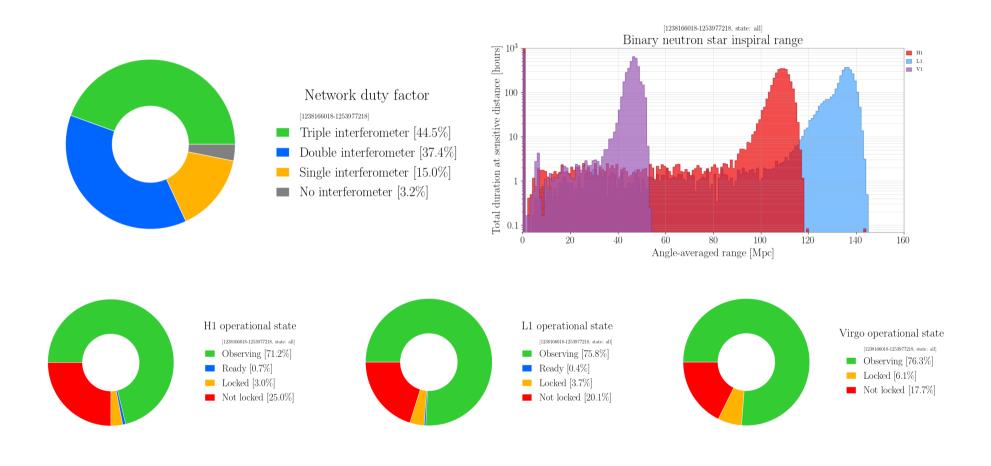
O3 Public Alerts (to date) by Month



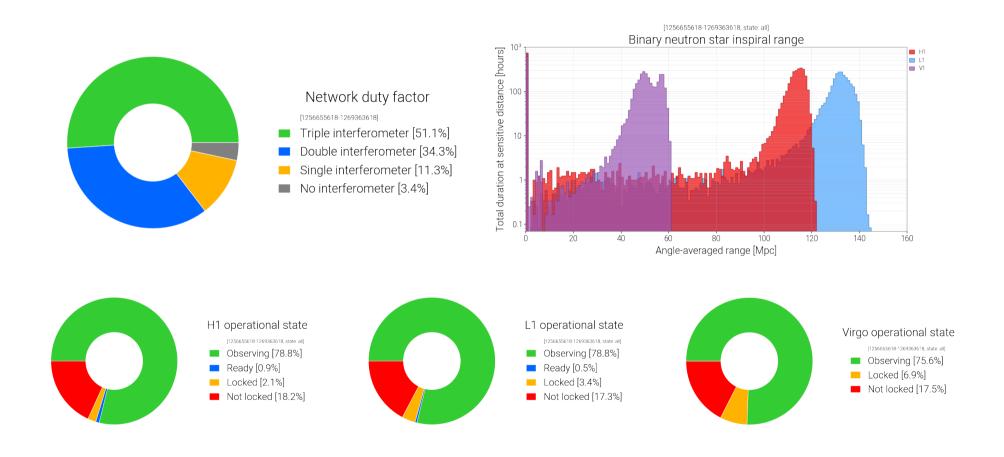
GWANW 2020 June 29

### LIGO-G2000997

# O3a Performance



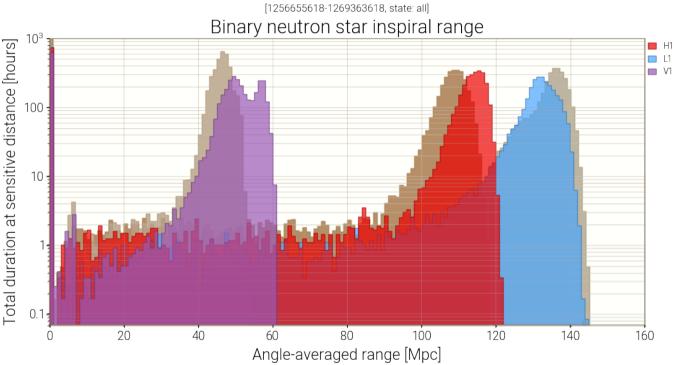
# O3b Performance



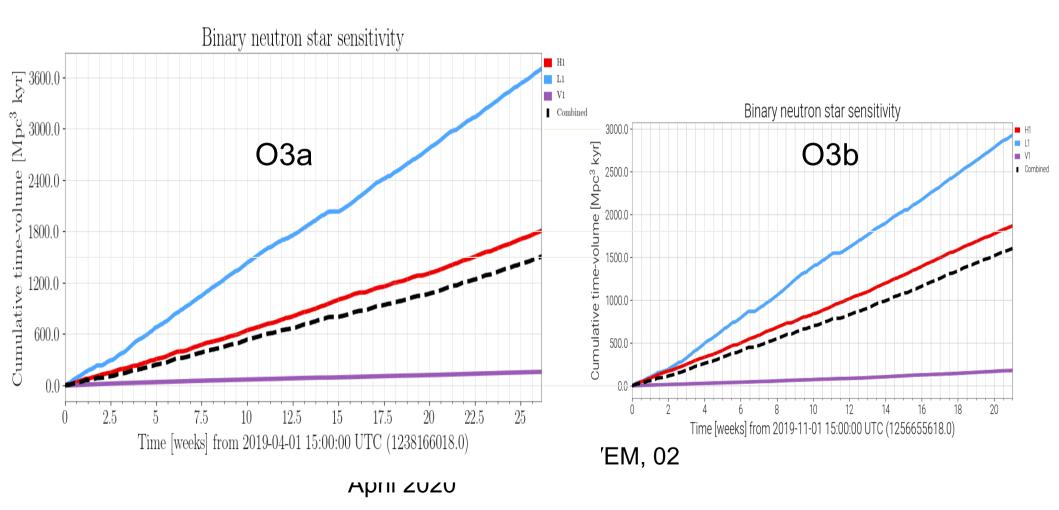
(KAGRA BNS range ~600kpc in Feb-Mar, 700k~1Mpc in O3GK)

### **October Commissioning Break Impact**

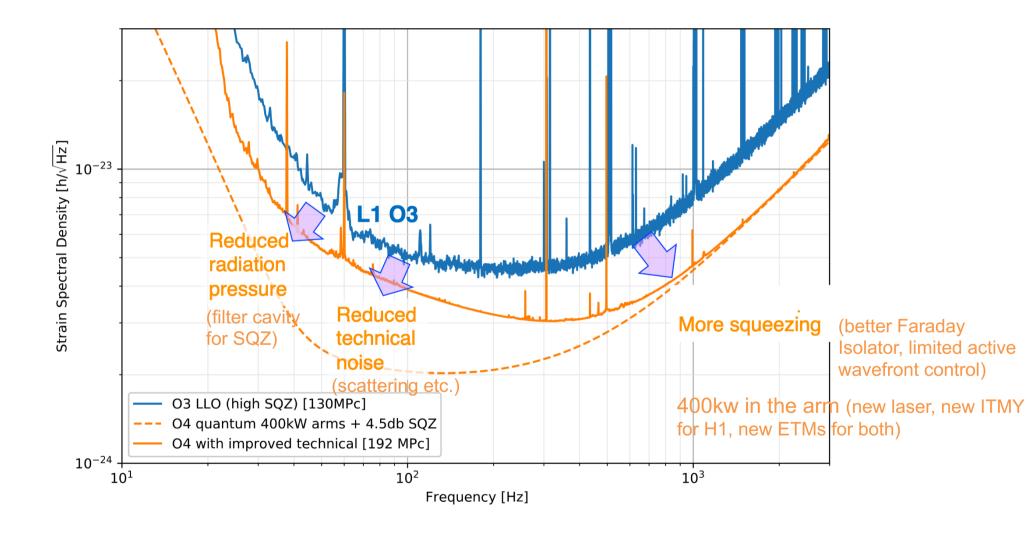
- V1 range improved.
- H1 range improved:
  - Various SQZ improvements.
  - Other tunings (e.g. spot position)
- L1 range reduced:
  - New point absorbers appeared which limited power buildup in the arms.
  - Some of the lost sensitivity was recovered by improved tuning of the squeezer.
- Some other improvements (e.g. glitch rate), see bonus slides, Jenne's talk, Detchar talks.



# Still a win: More BNS volume-time coverage in less than 5 months (148.08 days) than in 6 months (183 days).



## O4: This is what LIGO wants to implement.



#### FC Simple cartoon : Filter PSL Cavity TPD OPO spatial reference SHG 🧵 neoVAN-4S-HP #2 neoVAN-4S-HP #1 pre-mode-IFO cleaner (PMC) sqz NPRO EOM<sub>1</sub> FI FI Strain Spectral Density [h/ $\sqrt{H_z}$ 140W 125W 2W 10<sup>-23</sup> 60W L1 O3 Mirro Reduced Flexure radiation PZT pushes on flexure, which bends mirror PZT pressure Reduced (filter cavity technical for SQZ) Accommodates Fine thread noise PZT misalignments for preloading Mass = 1.77 kg (scattering etc.) **Output Faraday Isolator** Shroud around ETM Transmission Monitor 1 installed, 3 more to go End station chamber Mechanical layout Just one example of additional baffles for O4 $10^{-24}$ $10^{1}$ ETA

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#### Deiore

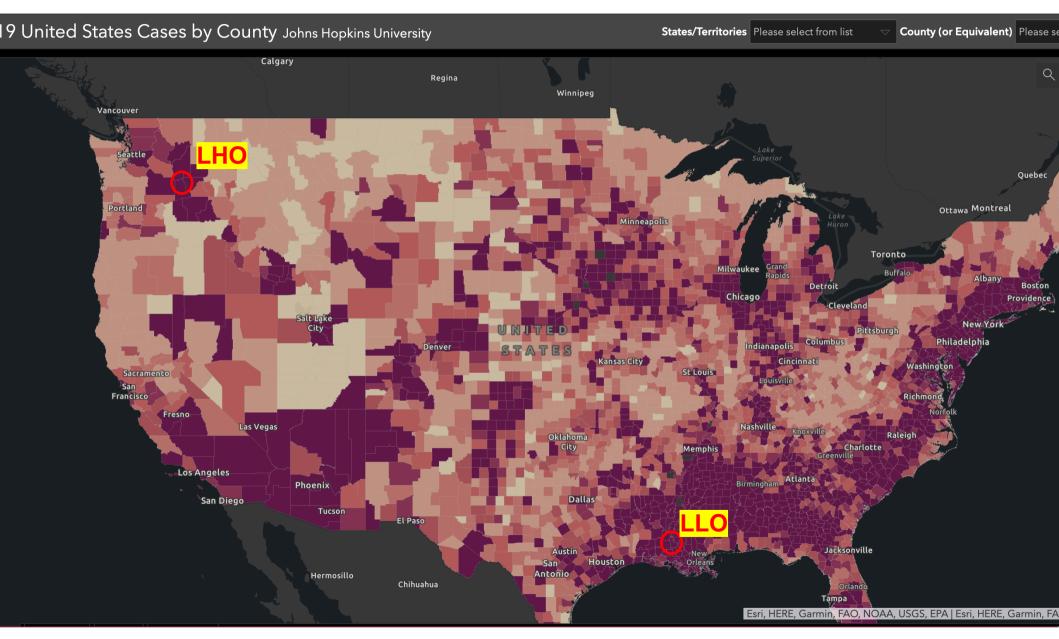
LIGO-G2000997

### New mirrors, a bunch of technical upgrades.

# But, COVID.

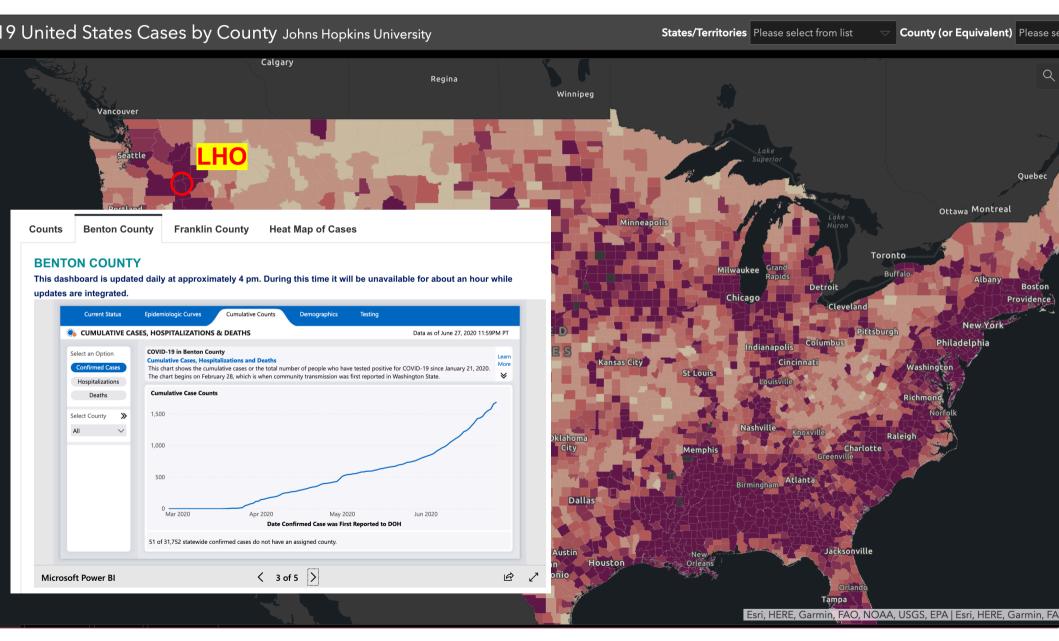
- O3 was cut short by 1 month.
- LHO and LLO: LIGO-Phase 3 (<u>L2000119</u>) as of now, i.e. still shut down. No site activities except to maintain the integrity of the detector and the observatory facilities.
  - WA: Benton and Franklin Co (LHO): 2 of 3 Counties in WA that is still in WA Phase 1 (most restrictive).
  - Louisiana (LLO): Sate-wide Phase 2 since June 5 (one step looser than Phase 1). On June 22, transition to Phase 3 was postponed for at least 22 days.

# LIGO locations on COVID map.



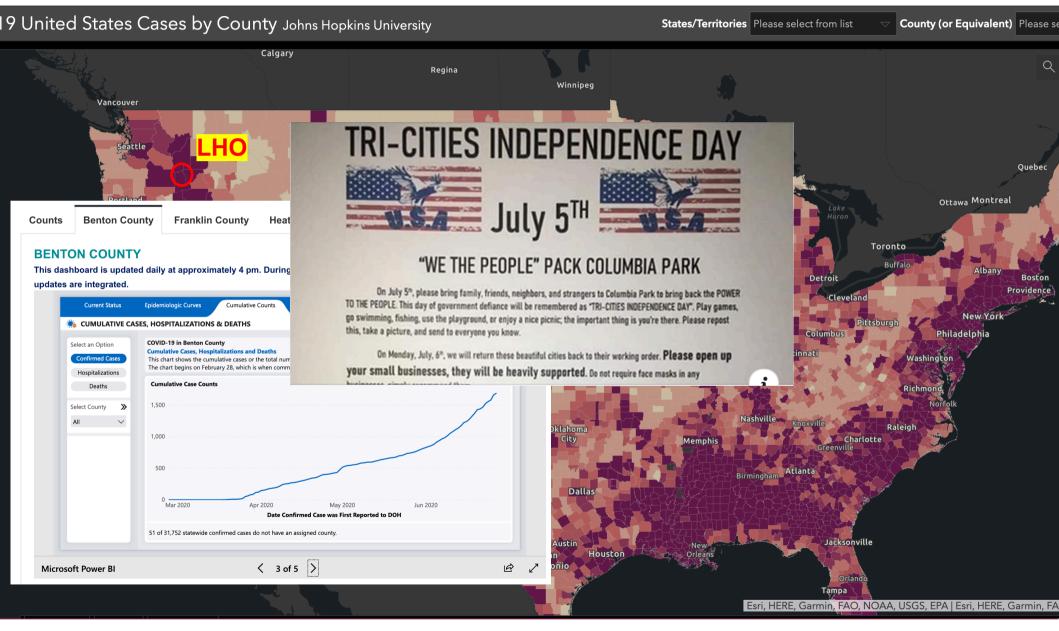
### LIGO-G2000997

# 1<sup>st</sup> wave isn't gone at LHO



#### LIGO-G2000997

# It will take some time until we flatten the curve here. But we'll do it eventually.



### LIGO-G2000997

# Impact on O4 (BNS 160~190Mpc) and A+ (BNS ~330Mpc target).

- Day-by-day schedule slip. May/June are already gone, at least a few weeks of July too.
- No solid date for LIGO-Phase 2 transition yet.
  - LIGO MIGHT be able to transition without waiting for State- or County-level transition, though that's not preferred.
- Even in LIGO-Phase 2, work won't be as efficient due to physical distancing and limited number of people in chamber etc.
- People are making plans e.g. safety guideline, modified work protocol etc. to transition to Phase 2, modify installation schedule etc.
- Both LHO and LLO will first resurrect H1 and L1 to O3-level. Jenne's talk about LHO status.

# Virgo in Italy. Saner.

- Situation is improving every week. People are working at the site.
- Adding Signal Recycling (or RSE)
  - Mirror arrived, gluing spacers and magnets.
  - SR payload will be assembled and installed inside the tower at mid-July.
- Adding guided lock acquisition like LIGO/KAGRA
  - Components of the auxiliary laser system are in production; the first complete SHG system should be ready and tested by end July to be installed in one terminal building;
- Infrastructure works for Frequency Dependent Squeezing (filter cavity etc.) are in progress.
  - Increasing the height of the tunnel for filter cavity (FC).
  - Central 150 m of the FC vacuum pipe have been installed.
  - All the other components (opticsl, mechanics, electronics) either are in production or have been ordered;

# KAGRA in Japan. Saner.

- U-Tokyo-Level 1 (30% of full capacity).
- Almost no new cases in Toyama, Japan, so 50% capacity is allowed at the site. Might loosen up further on July 6.
- O4 upgrade work starts in July.
  - Fixing/modifying all suspensions.
  - Dual Recycling with ASC (as opposed to no RSE, no full ASC).
  - Adding heater to cryostat.
  - Fixing leaks.
  - Adding baffles.
- Commissioning for sensitivity improvement scheduled from July 2021.

# Pre-COVID Plan again. Lots of uncertainties. But we'll come back with greater performance than O3.

	01	<b>—</b> 02	<b>O</b> 3	04	O5
LIGO	80 Мрс	100 Мрс	105-130 Mpc	160-190 Mpc	Target 330 Mpc
Virgo		30 Мрс	50 Mpc	90-120 Mpc	150-260 Mpc
KAGRA			8-25 Mpc	25-130 Mpc	130+ Mpc
LIGO-India	2				Target 330 Mpc
201	1 5 2016	2017 2018 2	1 I I 019 2020 202	1 2022 2023	2024 2025 2026

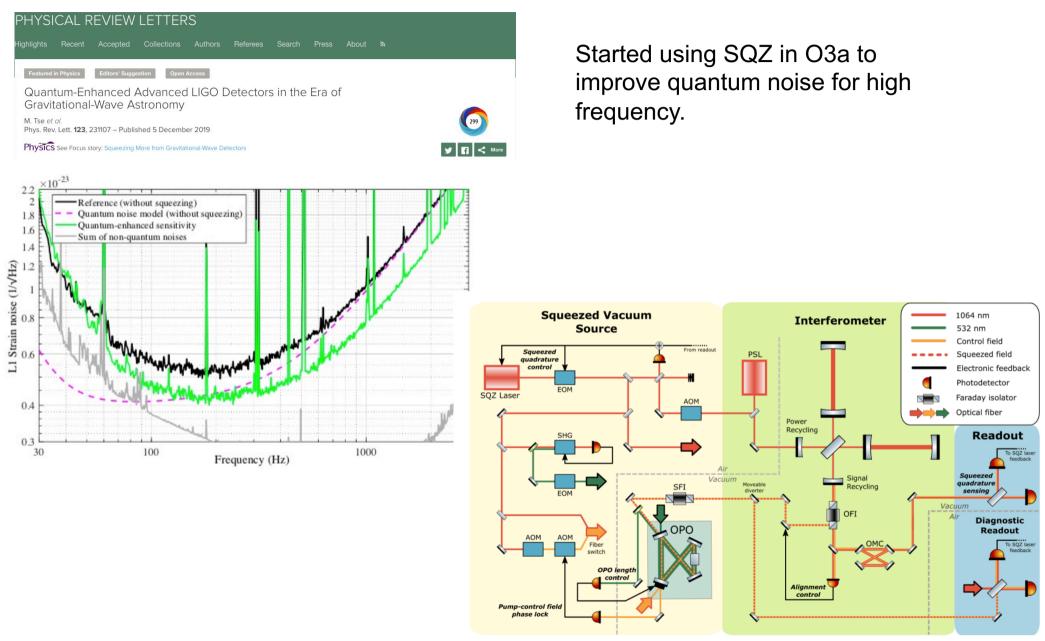
# Thank you, let's listen to Jenne's LHO Status talk.



# Bonus: KAGRA

- A detector next to Super-K, Kamioka mine, Japan.
  - Cryogenic sapphire mirrors to reduce thermal noise.
  - Underground for smaller seismic displacement and seismic Newtonian noise.
- Everything installed, but no successful Signal Recycling (or rather Resonant Sideband Extraction) yet.
- Cryogenic operation!
  - Increased absorption at 30k and below: "Ice" layer of the residual gas molecules. Engineering solution under study.
  - This is an important problem to solve for ALL 3G detectors.
- Inhomogeneous birefringence of sapphire turns out to be a serious problem. Mitigation strategy under study.
- BNS range ~600kpc in observation, Feb 25-Mar 10, 2020. Didn't join O3 at this point.
- **O3GK** with GEO600, Apr 7-Apr 21 2020, 0.7~1Mpc(?).

# (Bonus: Squeezing)

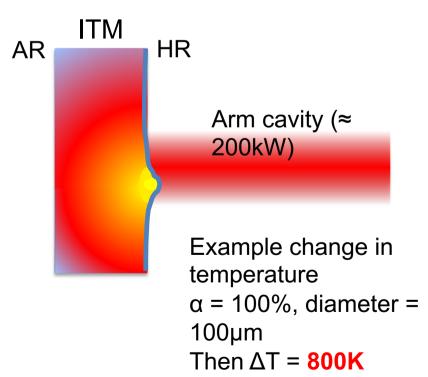


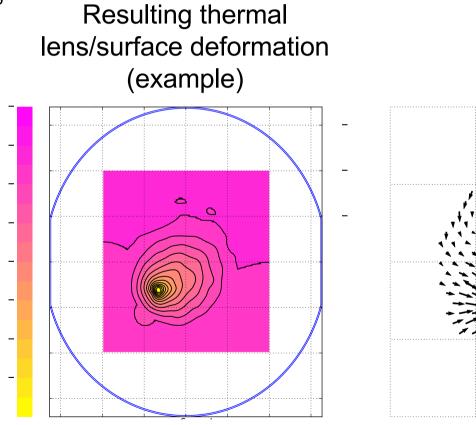
### GWANW 2020 June 29

### LIGO-G2000997

Bonus: What is a point absorber?

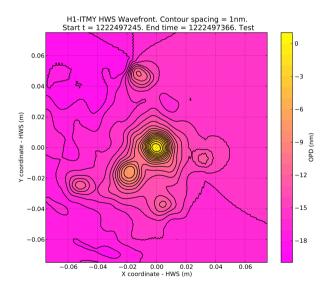
- Localized small ( $\leq 100 \mu m$ ),
- highly absorbing (> 1E4 ppm)
- on test mass HR surface

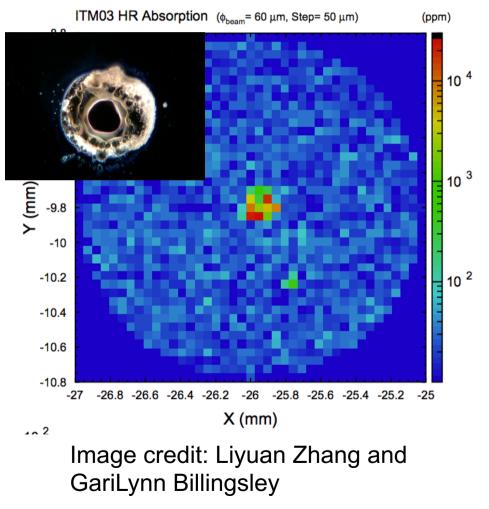




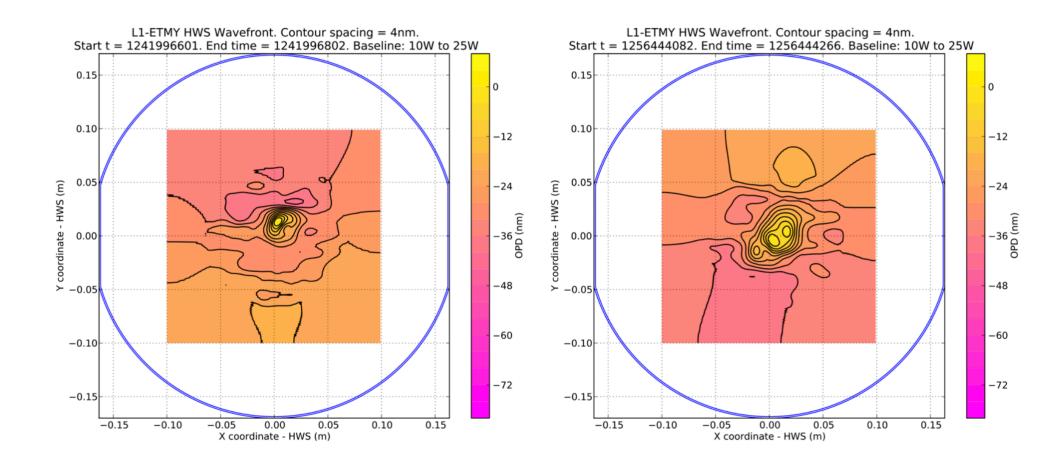
## **Measured Point Absorbers**

- Majority of optics measured show one or more point absorbers
  - Hartmann sensor (in-situ)
  - Photothermal common-path interferometry
  - Thermal imaging camera (under investigation)





# (Bonus: New Point Absorbers at LLO in O3b)

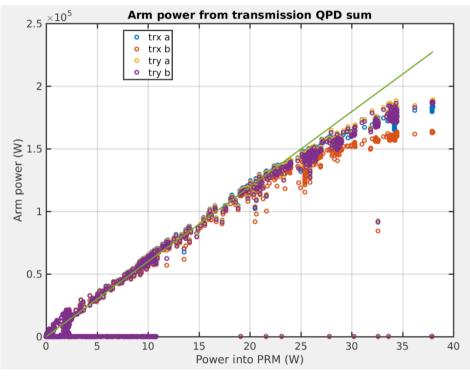


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# (Bonus: Point Absorbers)

- Hanford's large point absorber on ITMY limits circulating power and shot noise limited sensitivity
- Livingston circulating power reduced by 25% for O3b
  - Tuning squeezing recovered BNS
    range
- MIT and CIT groups are investigating possible point absorber mitigation strategies



LLO alog 42639 Valera

### **Scattered Light Upgrades During October Vent**

LHO · Angled septum

#### LLO

- **Pcal baffles** ٠ (EX & EY)
- TMSX • shroud Nozzle •
- baffles



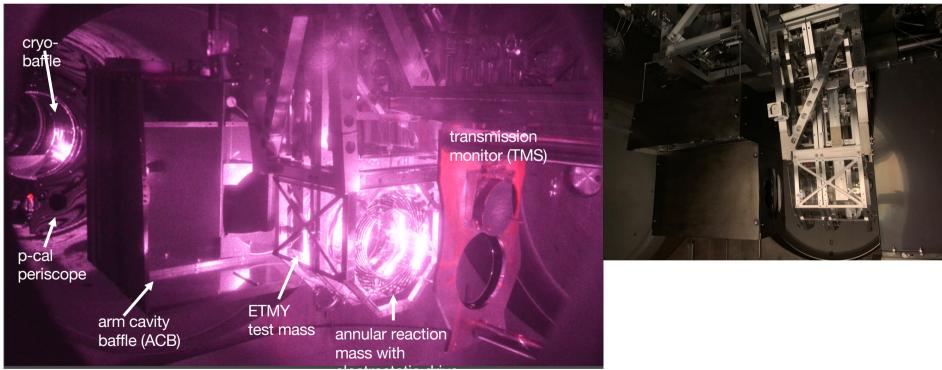
Images from: https://www.ligo.caltech.edu/WA/news/ligo20191104

LIGO-G2000997

## O3 Lessons Learned: Scattered Light

ETMY

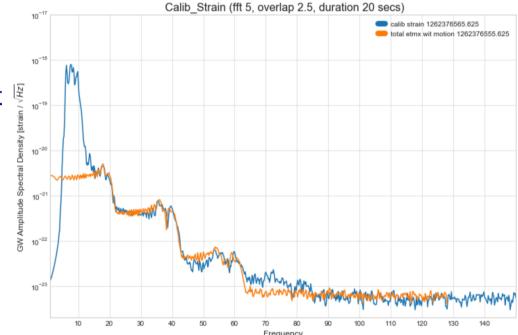


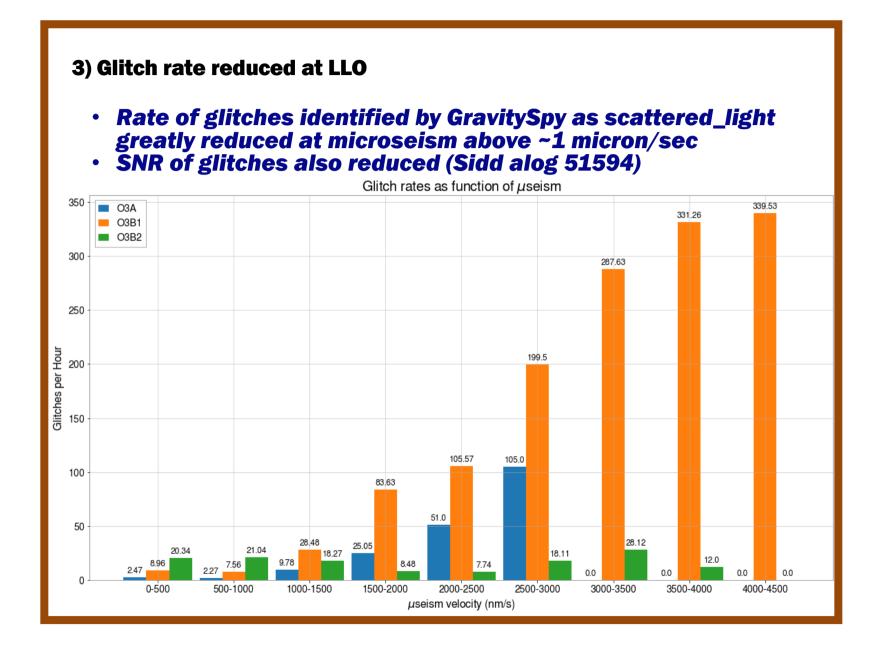


### **6/20005**79-V3

# O3 Lessons Learned: R0 Tracking

 Feed L2 Length OSEM signal back to R0 to reduce relative motion between test and reaction mass

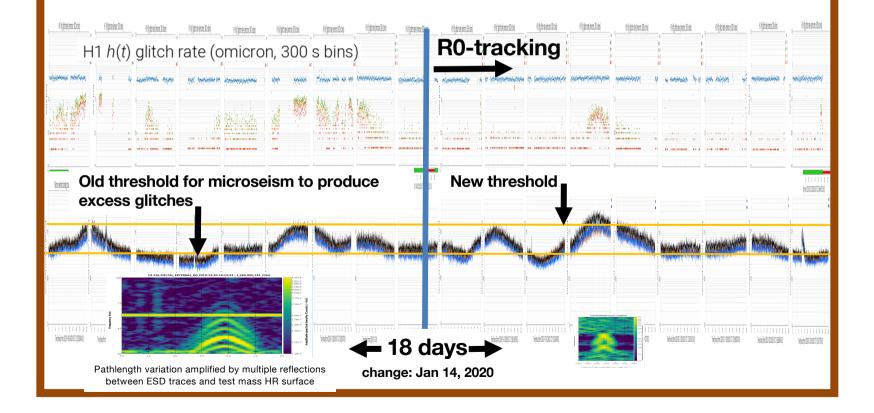




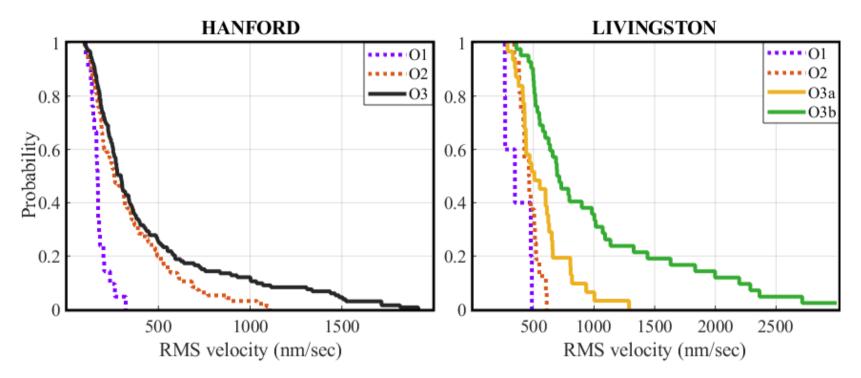
### LIGO-G2000997

#### 4) Glitch rate reduced at LHO

 Next level of scattering requires ~4 times higher microseism to produce excess glitching; investigating new dominant source – ACB, TMS, etc.

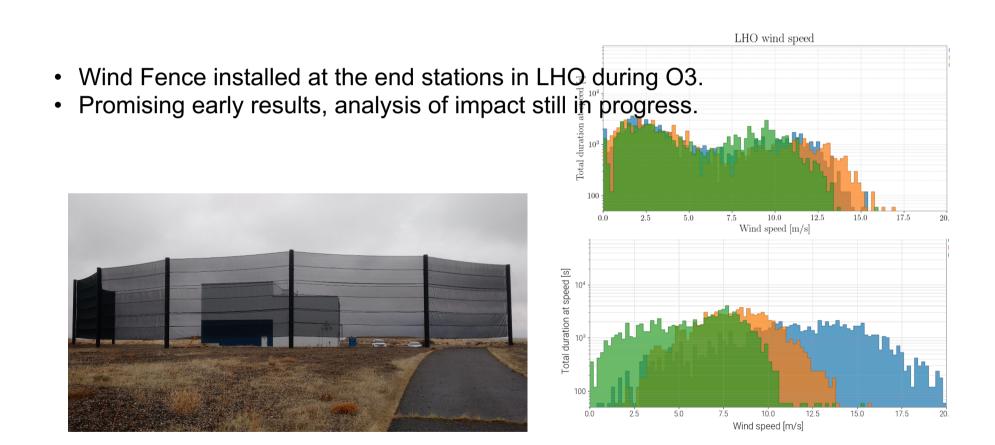


# O3 Lessons Learned: Earthquake Robustness



Source: E. Schwartz via email (from paper in progress)

## O3 Lessons Learned: Wind Resistance

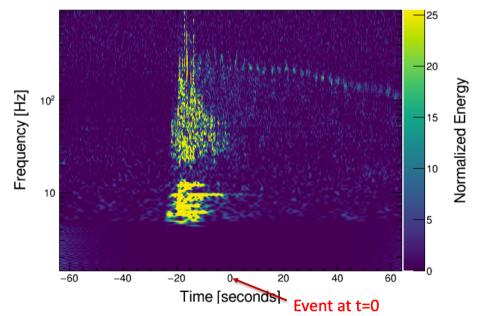


# Automation and Remote Operation

- H1 and L1 have high duty factors. Stay locked for ~80% of shifts.
- H1 relocks without intervention about half of the time.
- Livingston interferometer usually requires some minimal intervention. This was improving before run was suspended.
- CDS teams prepared remote work infrastructure to allow relocking from home.
- Automated alerts of various problems via text and email.
- Unattended operations trial at both LIGO sites, midnight to 8 am local at Hanford and 4 am to 8 am local at Livingston.
- Expect much more development for O4.

## Bonus: Retraction example: S190518bb

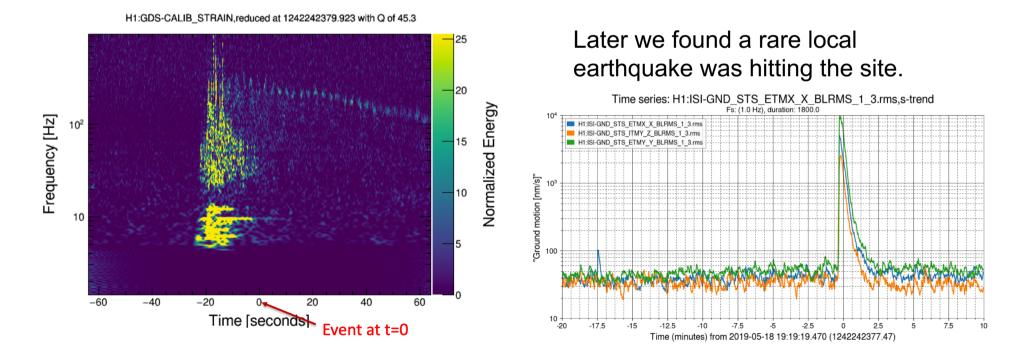
- Obvious clues:
  - [H1, L1, V1, netrowk] SNR = [11.7, 3.2, 5.0, 13.1].
  - [H1, L1, V1] chisquared=[2.7, "none", 1.05].
  - Huge scattering noise in H1 starting -20s.
  - Tight localization because the source needs to be in dark spot of L1.



H1:GDS-CALIB\_STRAIN, reduced at 1242242379.923 with Q of 45.3

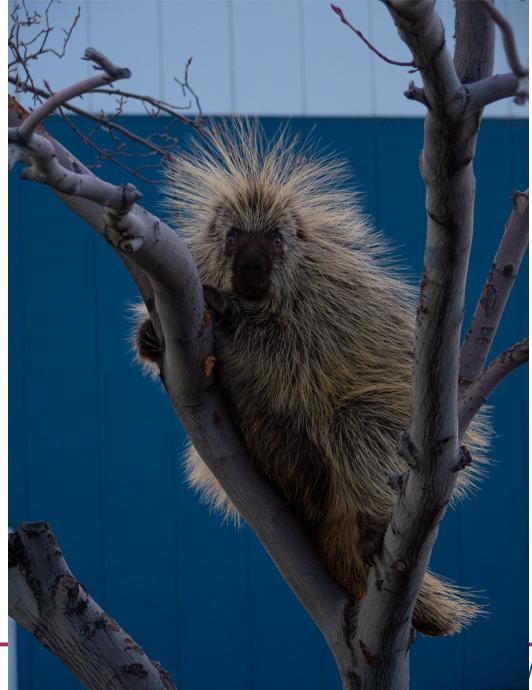
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  - Huge scattering noise in H1 starting -20s.
  - Tight localization because the source needs to be in dark spot of L1.



### LIGO-G2000997

# End of Bonus Slides



LIGO-G2000997

### VANW 2020 June 29