



# Minimising the Effect of Mirror Perturbations on Quantum Decoherence

LIGO SURF 2020

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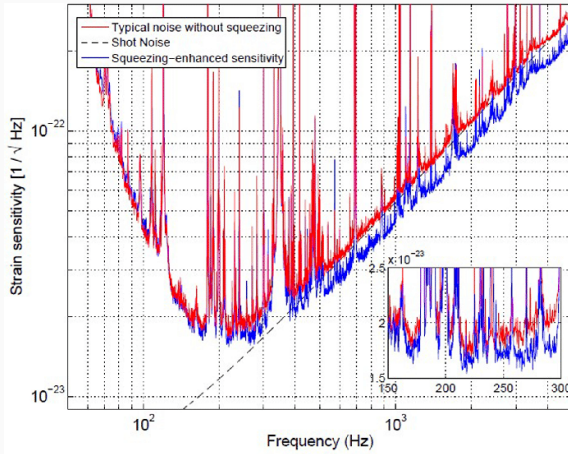
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# Introduction

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# Motivation



**Figure 1:** The strain sensitivity of the LIGO detectors is currently limited above approximately 200 Hz by quantum noise.<sup>1</sup>

<sup>1</sup>Publicly available image at <https://www.ligo.caltech.edu/>

## Squeezed states of light: our saviour

$$\Delta A \Delta \phi \geq \frac{\hbar}{2} \rightarrow \text{Unevenly distributed}$$



3 dB  $\uparrow \implies 3 \times$  Event Rate!

What is limiting the effective level of squeezing?

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Optical losses



What do optical losses result in?

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Squeezed  $\rightarrow$  squeezed + unsqueezed



What is causing these losses?



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What is causing these losses?

Mode-mismatches! (More than 10% in aLIGO<sup>[1]</sup>)



What causes these mode-mismatches?

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Optical losses



What do optical losses result in?

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What is causing these losses?

Mode-mismatches! (More than 10% in aLIGO<sup>[1]</sup>)



What causes these mode-mismatches?

Perturbation in apparatus ( $R = 5.932418$  m?!)

Can we find an optimal set of design parameters such that the interferometer becomes minimally sensitive to design perturbations?

Perturbations → curvatures and positions of the optical elements

# Background

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# Spatial Modes<sup>[2, 3]</sup>

Laser beam  $\rightarrow$  Spatial intensity distribution

$$E(t, x, y, z) = \sum_j \sum_{n,m} a_{jnm} u_{nm}(x, y, z) \exp(i(\omega_j t - k_j z)) \quad (1)$$

$u_{nm}(x, y, z) \rightarrow$  set of Hermite-Gauss (HG) or Laguerre-Gauss (LG) polynomials

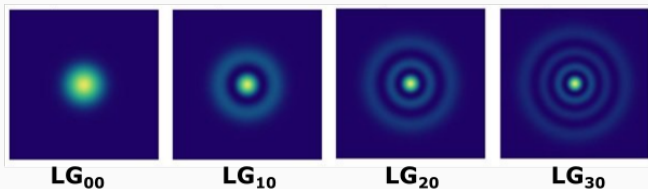


Figure 2:  $u_{nm}$  spatial distributions in the LG basis

# The Fundamental Gaussian Mode

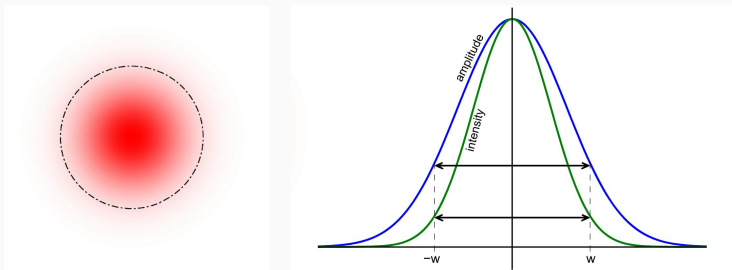
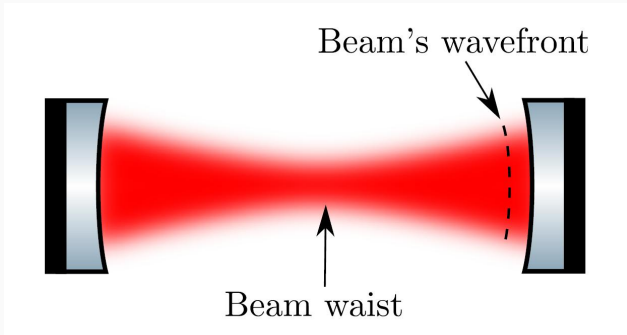


Figure 3: Intensity pattern of a Gaussian beam (left) and the intensity and amplitude distributions of a normalised Gaussian beam (right).<sup>[3]</sup>

# Cavity Eigenmode and Mode-Mismatch



**Figure 4:** Cavity eigenmode: The beam curvature must be equal to the curvature of the mirrors at the mirror positions.<sup>[3]</sup>

# Goals

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- Algorithm for numerically optimising any optical setup
- Implement algorithm to optimise the aLIGO design
- Analytic formalism to calculate loss

# Analytic Formalism

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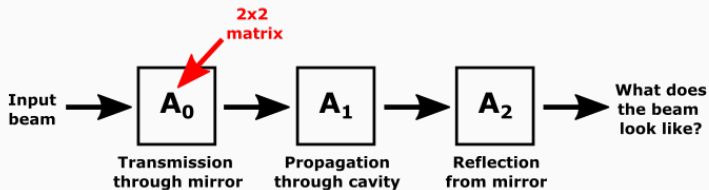
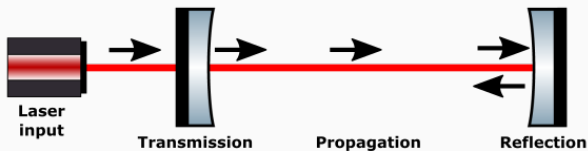
To develop an analytic formalism to calculate the total mode-matching loss in a complex optical system as a function of small perturbations of the optic positions and curvatures.

Small curvature and position perturbations:  $LG_{00} \leftrightarrow LG_{10}^{[4,1]}$

Electric field at any point

$$|\Psi\rangle = \begin{pmatrix} \alpha \\ \beta \end{pmatrix} = \alpha |LG_{00}\rangle + \beta |LG_{10}\rangle \quad (2)$$

# Mode-Mixing Matrices



$$|\Psi_{\text{out}}\rangle = A_2 \times A_1 \times A_0 \times |\Psi_{\text{in}}\rangle \quad (3)$$

# Formalism and Simulations: Do They Agree?

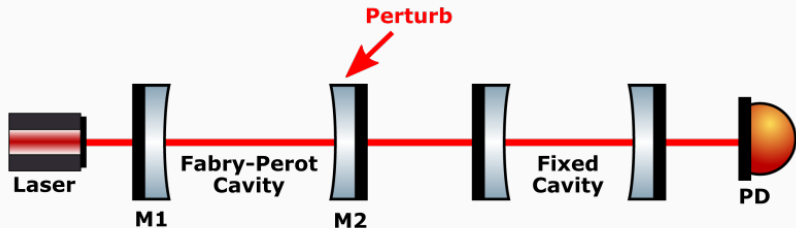
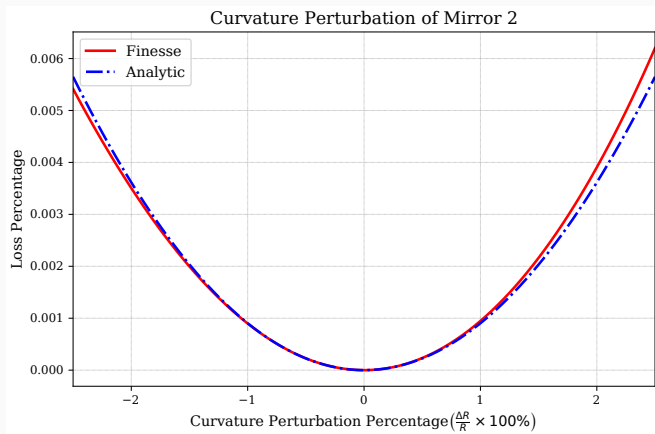


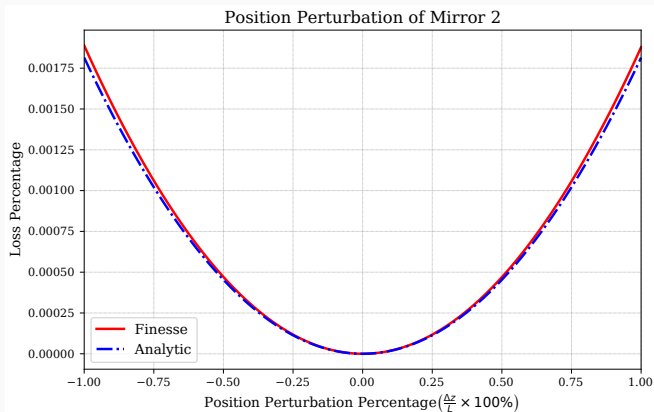
Figure 5: Fabry-Perot cavity before the OMC.

# Formalism and Simulations: Do They Agree?



**Figure 6:** Analytic and simulation results for power loss percentage as a function of curvature perturbation percentage at M2.

# Formalism and Simulations: Do They Agree?



**Figure 7:** Analytic and simulation results for power loss percentage as a function of position perturbation percentage at M2.



# Numerical Optimisation of the aLIGO System

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# Optimising the aLIGO System

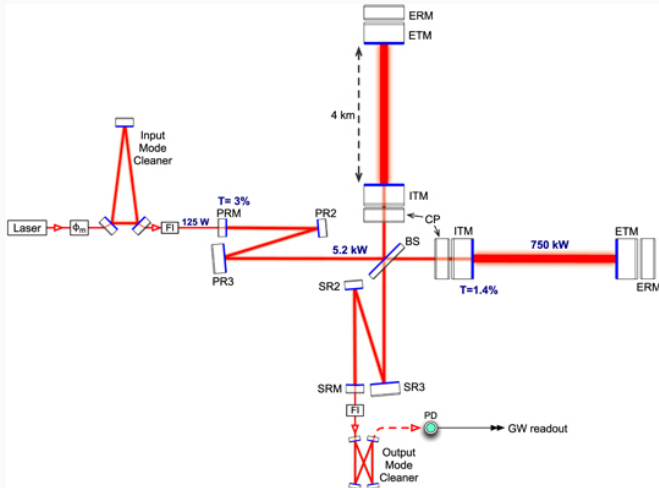


Figure 8: aLIGO setup

# Optimising the Signal Recycling Cavity

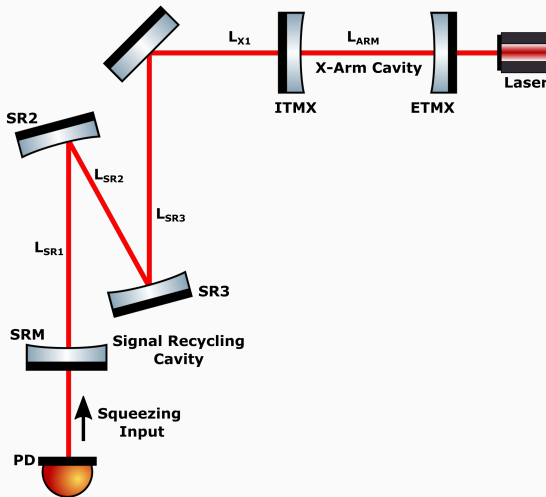


Figure 9: aLIGO X-arm cavity and Signal Recycling Cavity (SRC)

# Optimisation Routine - Particle Swarm Optimisation

Ensure the setup is mode-matched



Ensure beam-size remains small ( $< 1\text{cm}$ ) and total length of the SRC remains fixed



Perturb curvature and position of SRC mirrors and observe degradation in squeezing level<sup>2</sup>



Minimise this degradation

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<sup>2</sup>Laser and squeezer inputs remain mode-matched to the unperturbed cavity

# Monte Carlo Analysis for the Expected Squeezing Level

- $\Delta R \sim N(\mu = 0, \sigma = 0.01R)$  and  $\Delta z \sim N(\mu = 0, \sigma = 3mm)$ , where  $\sim N$  indicates a normal distribution.
- Repeat 1,000 times
- Plot probability distribution of the squeezing level
- Compute the 85<sup>th</sup> percentile

# Expected Squeezing Level for aLIGO SRC

Effective squeezing level of unperturbed cavity = 9.8dB

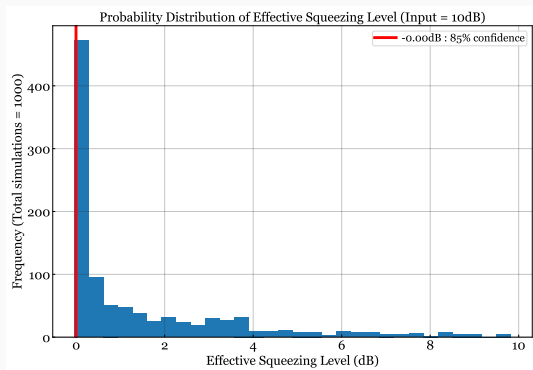


Figure 10: Monte Carlo analysis for perturbed cavity

	$R_{SR3}$	$R_{SR2}$	$R_{SRM}$	$L_{SR3}$	$L_{SR2}$	$L_{SR1}$
aLIGO (m)	35.97	-6.41	-5.69	19.37	15.44	15.76

# Perturbation Analysis for $R_{SRC}$

Effective squeezing level of unperturbed cavity = 9.8dB

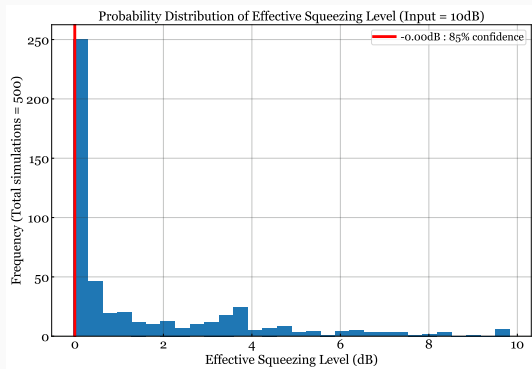


Figure 11: Perturbing only the  $R_{SRC}$

	$R_{SR3}$	$R_{SR2}$	$R_{SRM}$	$L_{SR3}$	$L_{SR2}$	$L_{SR1}$
aLIGO (m)	35.97	-6.41	-5.69	19.37	15.44	15.76

# Expected Squeezing Level for Optimised SRC

Effective squeezing level of unperturbed cavity = 9.7dB

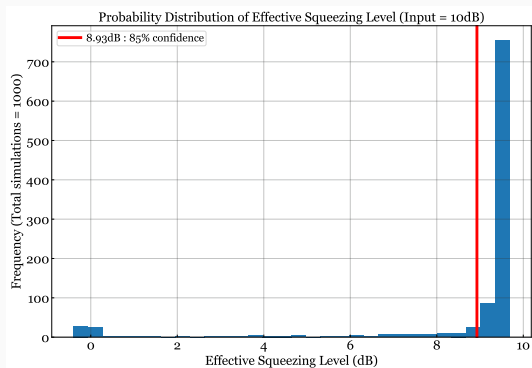


Figure 12: Monte Carlo analysis for perturbed cavity

	$R_{SR3}$	$R_{SR2}$	$R_{SRM}$	$L_{SR3}$	$L_{SR2}$	$L_{SR1}$
Optimised (m)	59.66	-233.60	-3.53	27.69	18.98	9.87



## Conclusion

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- Developed algorithm to minimise sensitivity of any optical setup.
- Optimised the aLIGO Signal Recycling Cavity
- Quantified the improvement.
- Developed analytic formalism.
- Checked agreement between formalism and simulations.

## Next Steps

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The next steps for this project are:

- Apply to the entire aLIGO setup.
- Take into account the LIGO thermal compensation system.

# Thank You

My sincere thanks and gratitude to my mentors and to LIGO Laboratory for giving me this opportunity.

Questions?

Helmholtz equation:

$$\nabla^2 \mathbf{E} - \frac{\ddot{\mathbf{E}}}{c^2} = 0 \quad (4)$$

The fundamental Gaussian mode at a position  $z$ :

$$u_{00}(r, z) = \sqrt{\frac{2}{\pi}} \frac{1}{\omega(z)} \exp(i\psi(z)) \exp\left[-r^2 \left(\frac{1}{w^2(z)} + i \frac{\pi}{\lambda R(z)}\right)\right] \quad (5)$$

# Reflection Matrix

Reflection from a mirror having amplitude reflectivity coefficient  $r$

$$\mathbf{r} = \begin{pmatrix} r & 0 \\ 0 & r \end{pmatrix} \quad (6)$$

Modified reflection matrix for a small curvature perturbation  $\delta R$  in a mirror of curvature  $R$

$$\mathbf{r}' = \mathbf{a} \cdot \mathbf{r} = \begin{pmatrix} r\sqrt{1-a^2} & -\imath ra \\ -\imath ra & r\sqrt{1-a^2} \end{pmatrix}, \quad a = \frac{\pi\omega^2(z_m)}{2\lambda R^2} \delta R \quad (7)$$

where  $\omega(z_m)$  is the beam size at the mirror position  $z_m$



# Transmission Matrix

The scattering matrix for transmission of a beam through a mirror having amplitude transmissivity coefficient  $t$  is given by

$$\mathbf{t} = \begin{pmatrix} t & 0 \\ 0 & t \end{pmatrix} \quad (8)$$

## Propagation Matrix





If one the one-way propagation of a beam across a distance or cavity accumulates a phase  $\phi_0$  in the  $LG_{00}$  mode and  $\phi_1$  in the  $LG_{10}$  mode, this phase accumulation during propagation can be represented by the scattering matrix

$$\phi = \begin{pmatrix} e^{i\phi_0} & 0 \\ 0 & e^{i\phi_1} \end{pmatrix} \quad (9)$$

The modified propagation matrix for a small position perturbation  $\delta z$  in a mirror of radius of curvature  $R$

$$\phi' = \mathbf{b} \cdot \phi = \begin{pmatrix} \sqrt{1-b^2} e^{i\phi_0} & -b e^{i\phi_1} \\ -b e^{i\phi_0} & \sqrt{1-b^2} e^{i\phi_1} \end{pmatrix}, \quad b = \frac{\delta z}{R} \quad (10)$$

# References

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