

Data Quality and GWpy

Open Data Workshop #3

May 27, 2020

Marissa Walker



Acknowledgments

Many slides adapted from Jess McIver and Duncan Macleod [https://
dcc.ligo.org/LIGO-G1800649/public](https://dcc.ligo.org/LIGO-G1800649/public)
(2018 Open Data Workshop)

Outline

What does LIGO strain data look like?

- Time domain
- Frequency domain
- Time-frequency representations

Data quality: noise artifacts in LIGO data

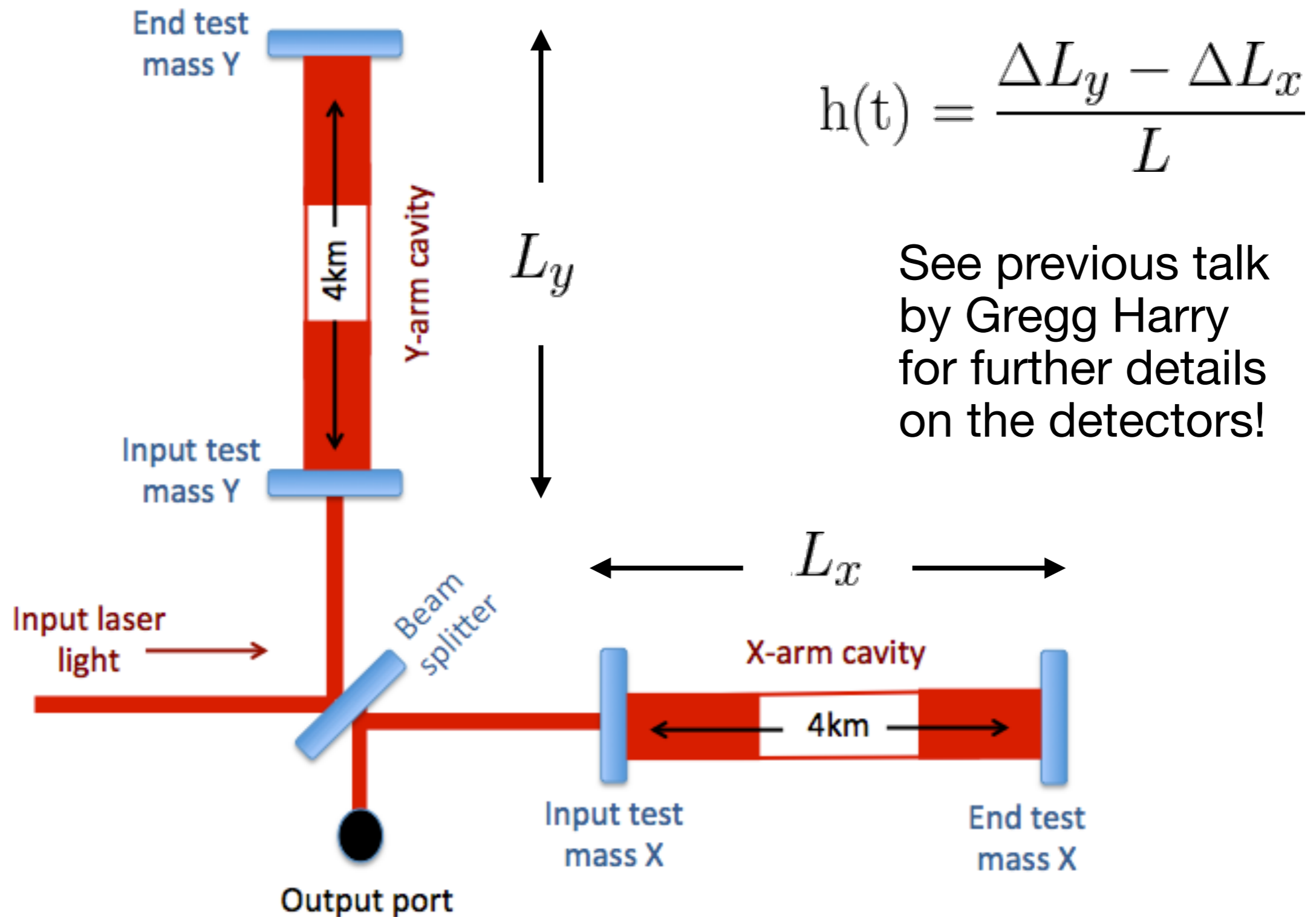
- Glitches
- Lines

Data quality information

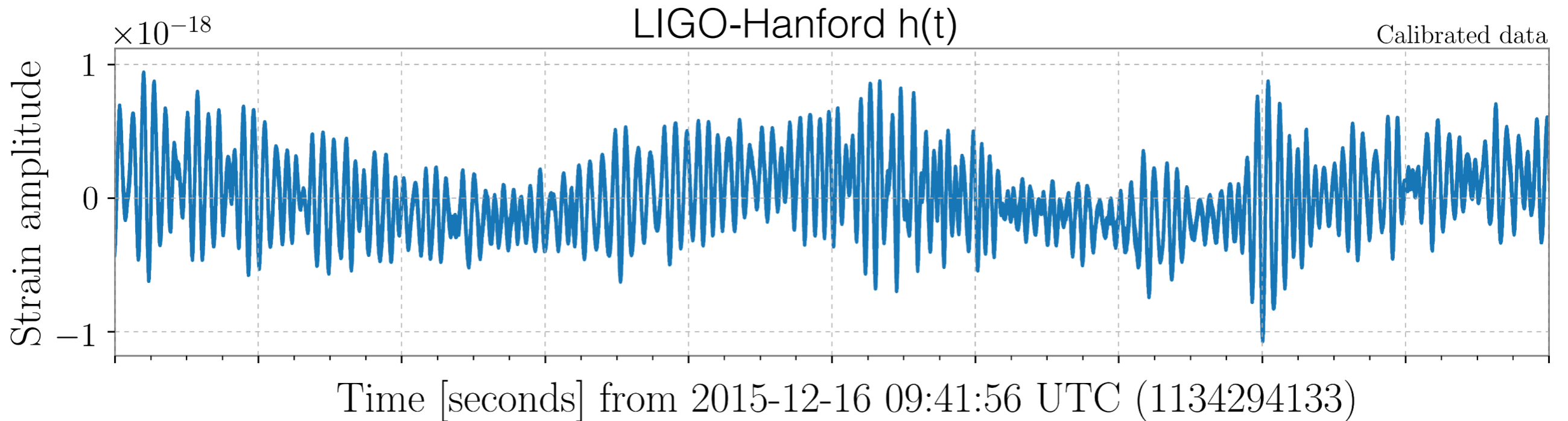
- Detector status summary pages
- Data quality segments
- Event validation

Summary of resources and references

What is strain data, $h(t)$?



What does LIGO data look like?



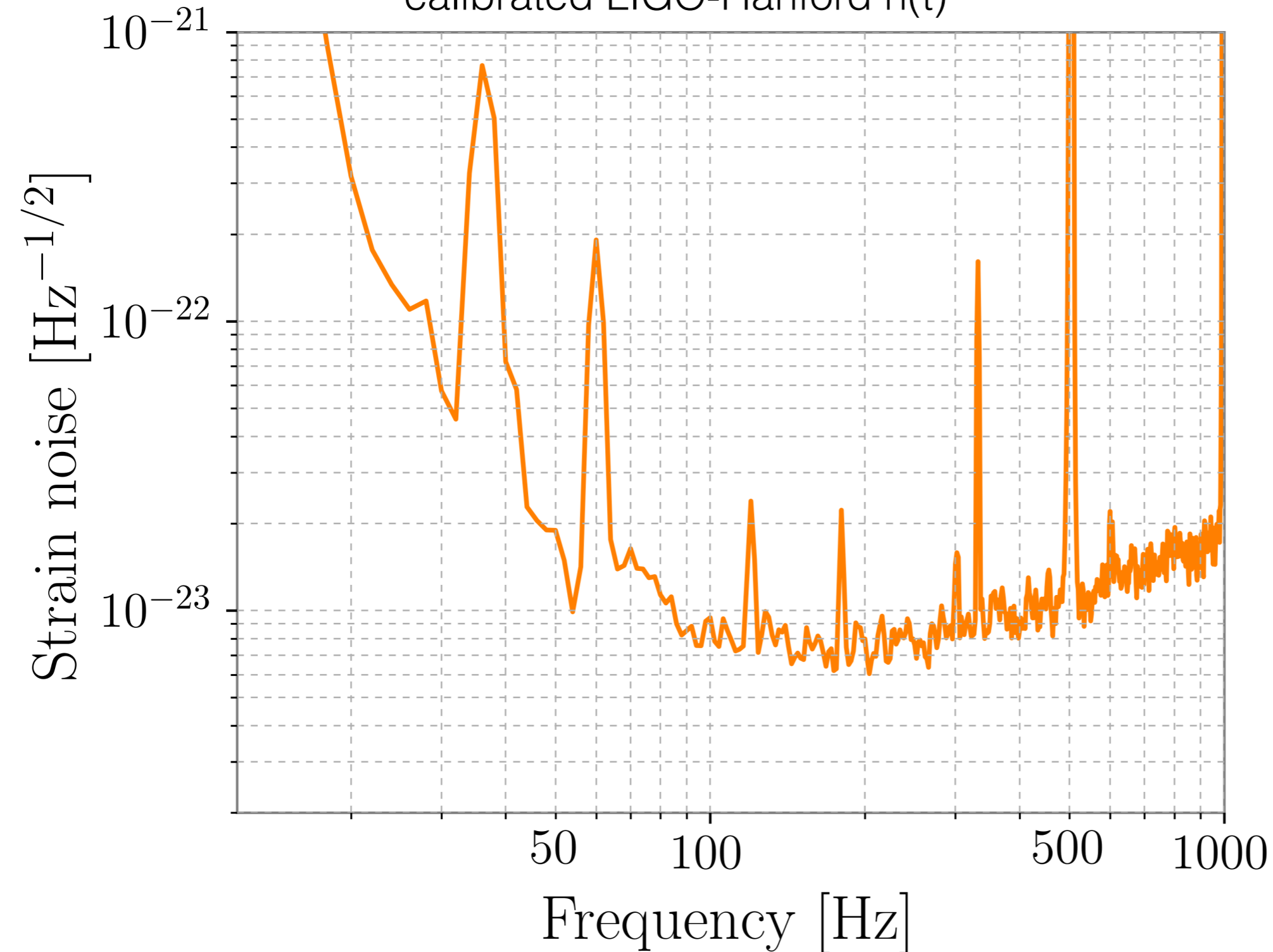
$h(t)$ **sampling rate** for LIGO detectors: 16384 Hz

Open data: 16384 or 4096 Hz

In Tutorials 1.1 and 1.2 you will learn how to get
and plot time series data

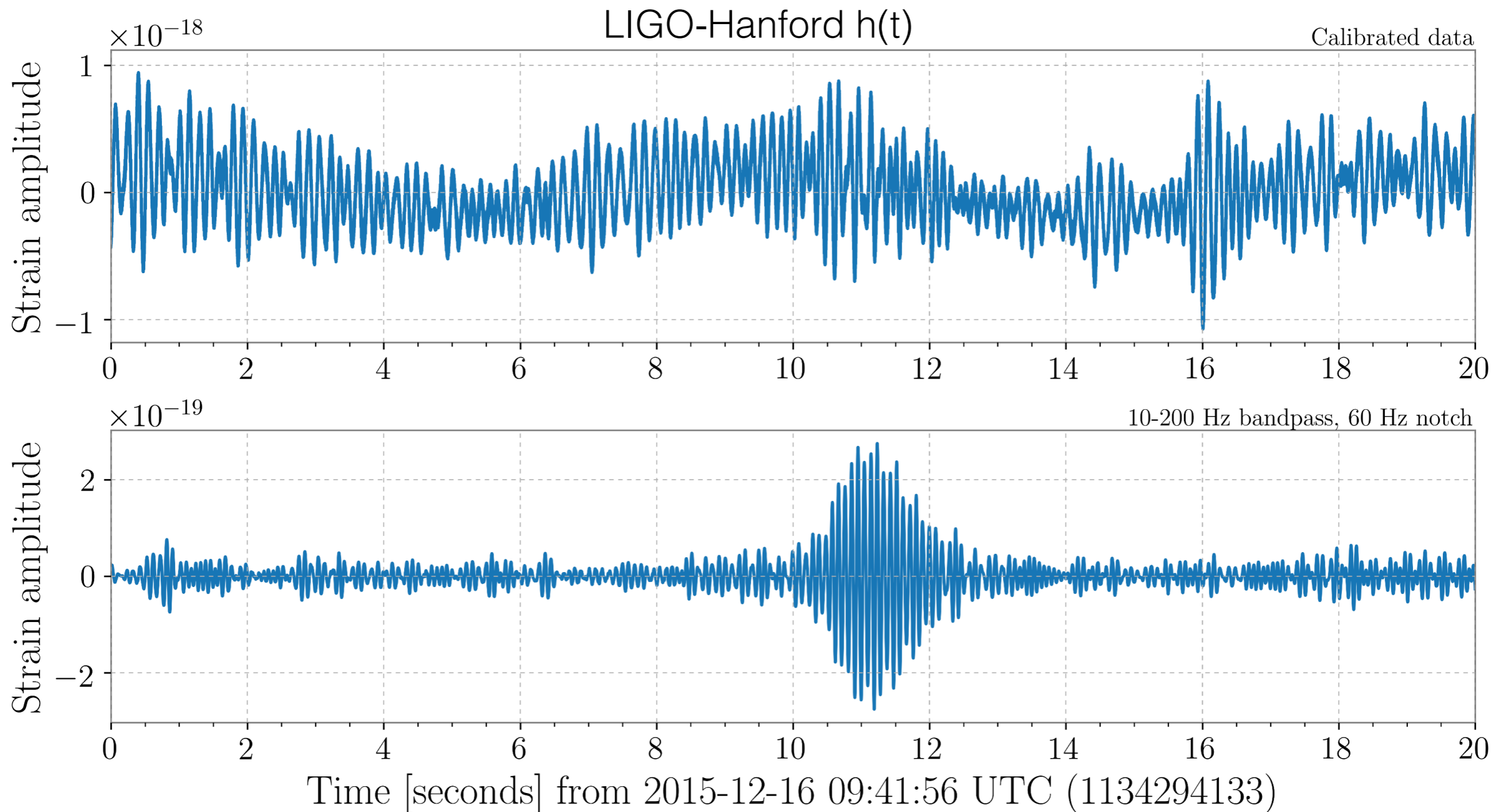
LIGO data in the frequency domain

Amplitude Spectral Density (ASD) of
calibrated LIGO-Hanford $h(t)$

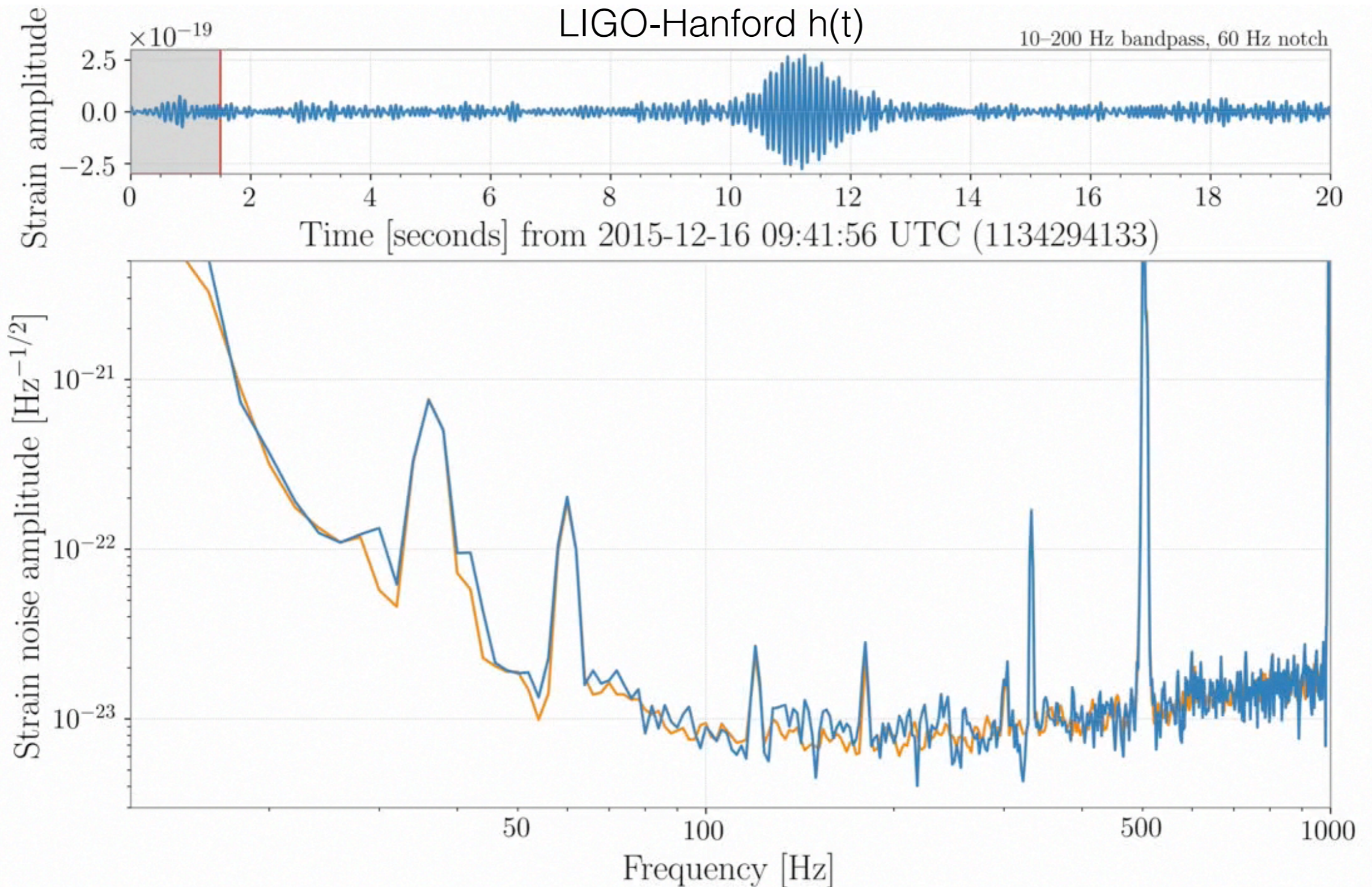


In Tutorial 1.2 you
will learn how to
calculate the
amplitude spectral
density using
GWpy

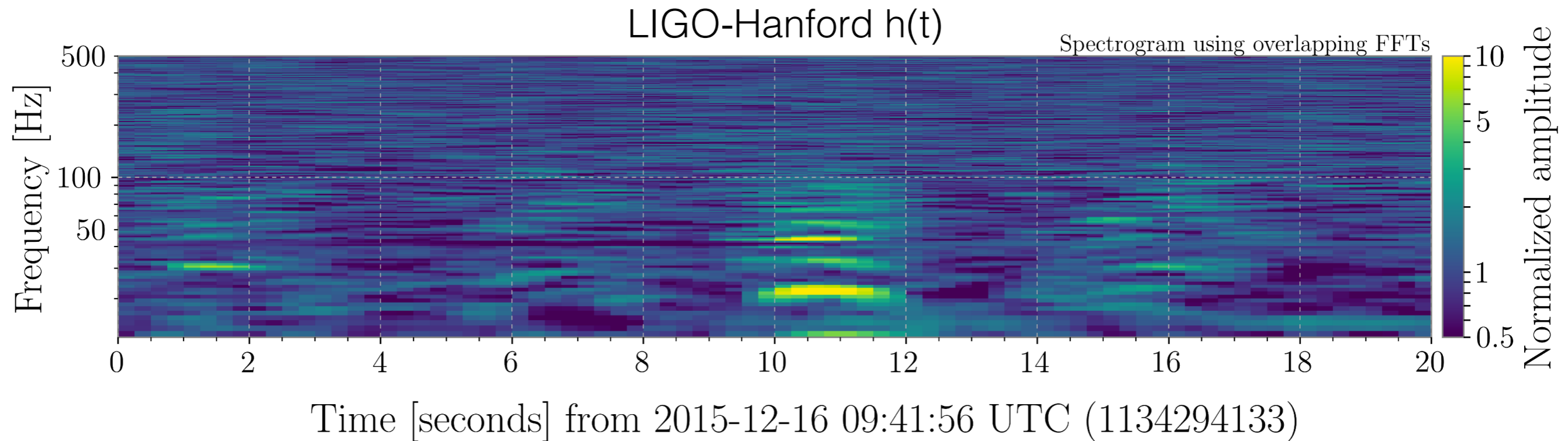
What does LIGO data look like?



LIGO data in the frequency domain

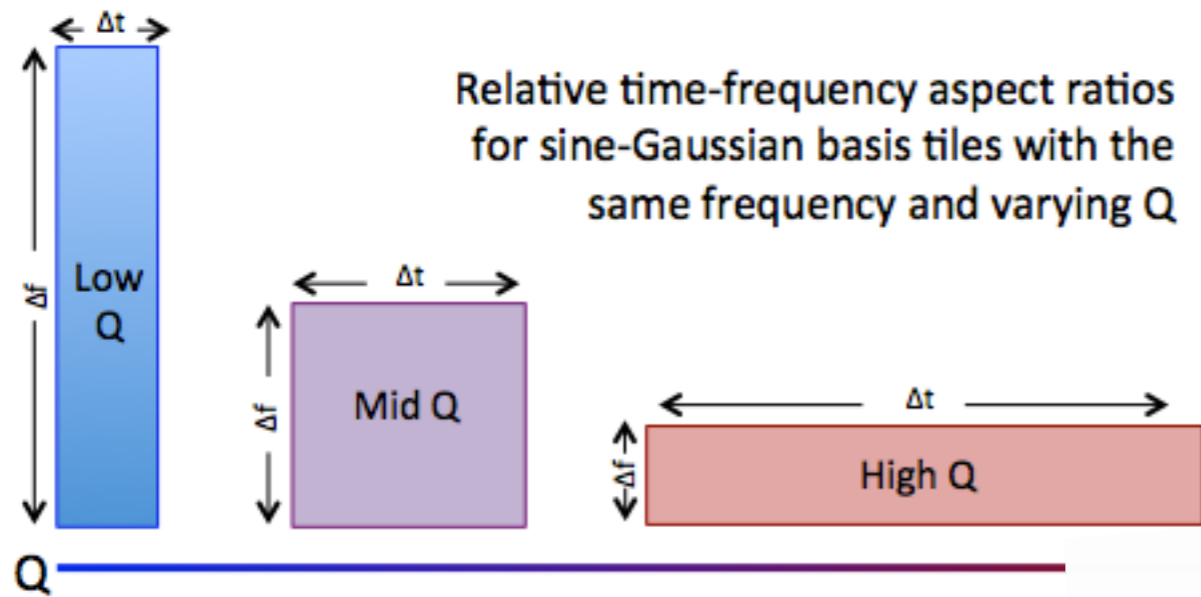


Time-frequency spectrogram

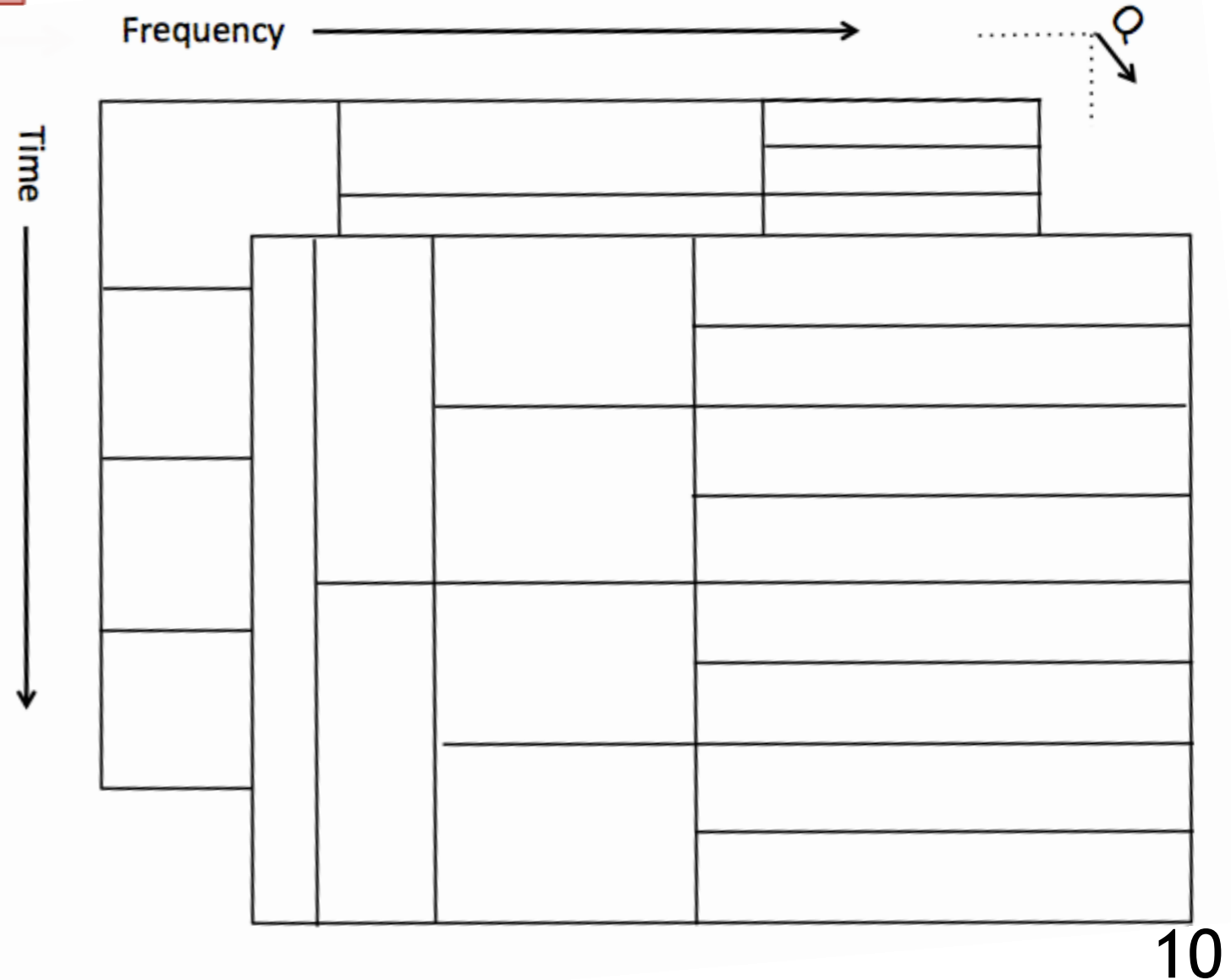
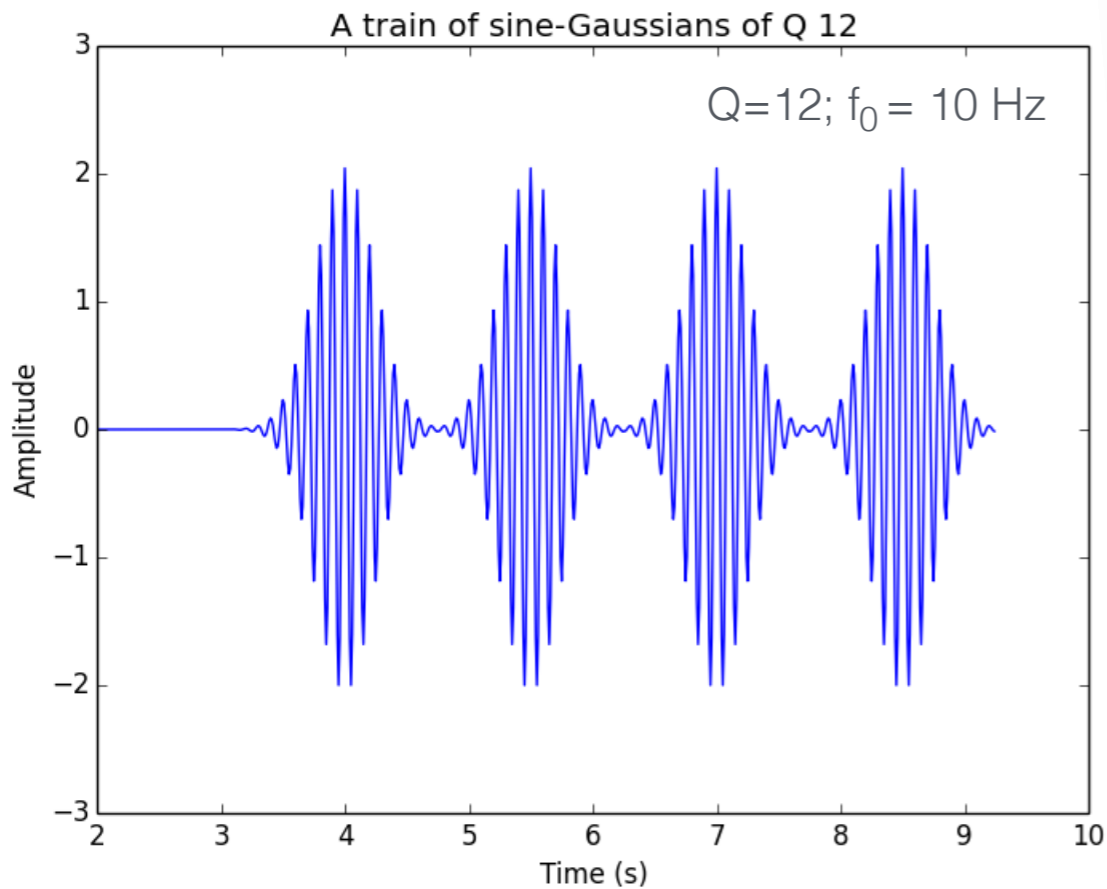


The Q transform

S. Chatterji et al. CQG (2010)
Images: McIver

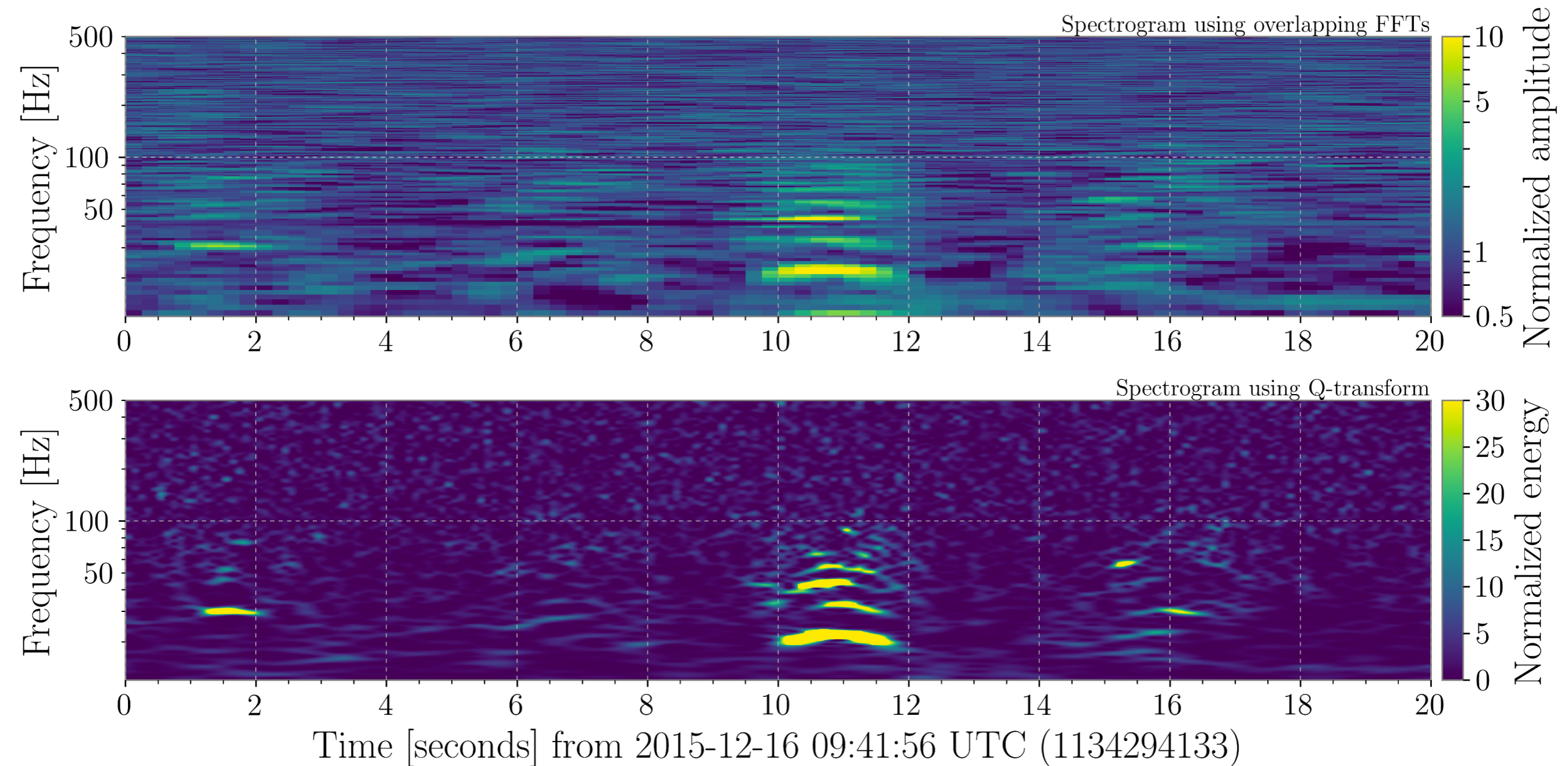


$$Q = \frac{f_0}{\Delta f}$$



Time-frequency spectrograms

LIGO-Hanford $h(t)$



In Tutorial 1.3 you will learn how to plot spectrograms and use the q-transform

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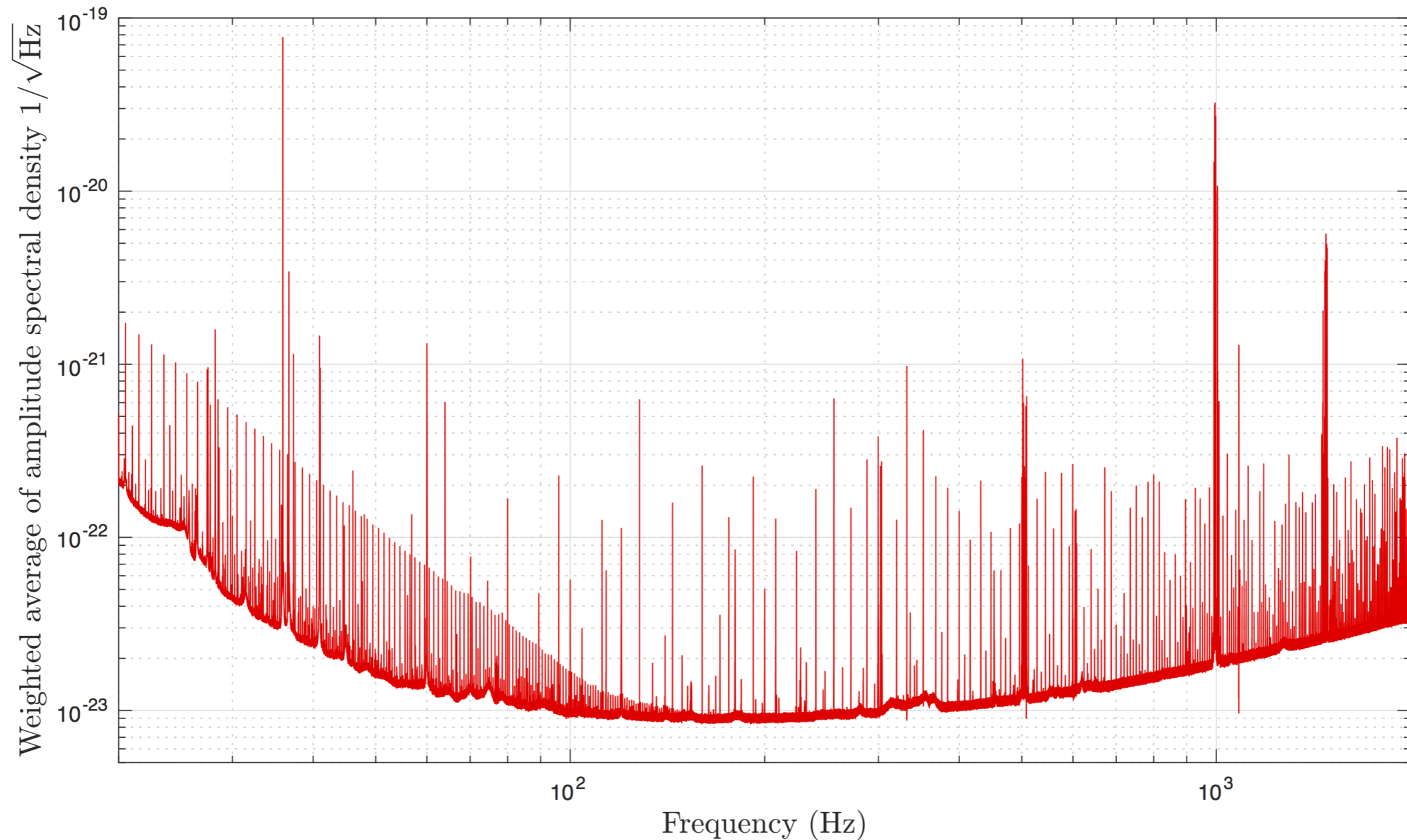
Data quality information

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Summary of resources and references

Combs of lines in LIGO data

H1 O1 combs (7200 second averaging)



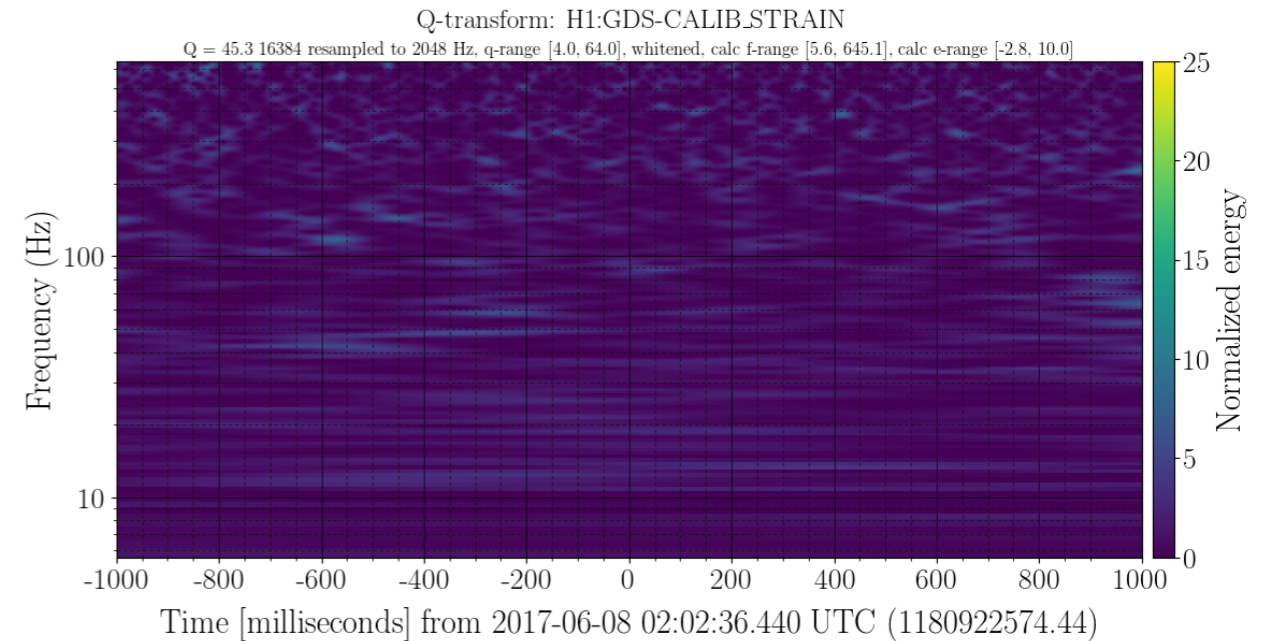
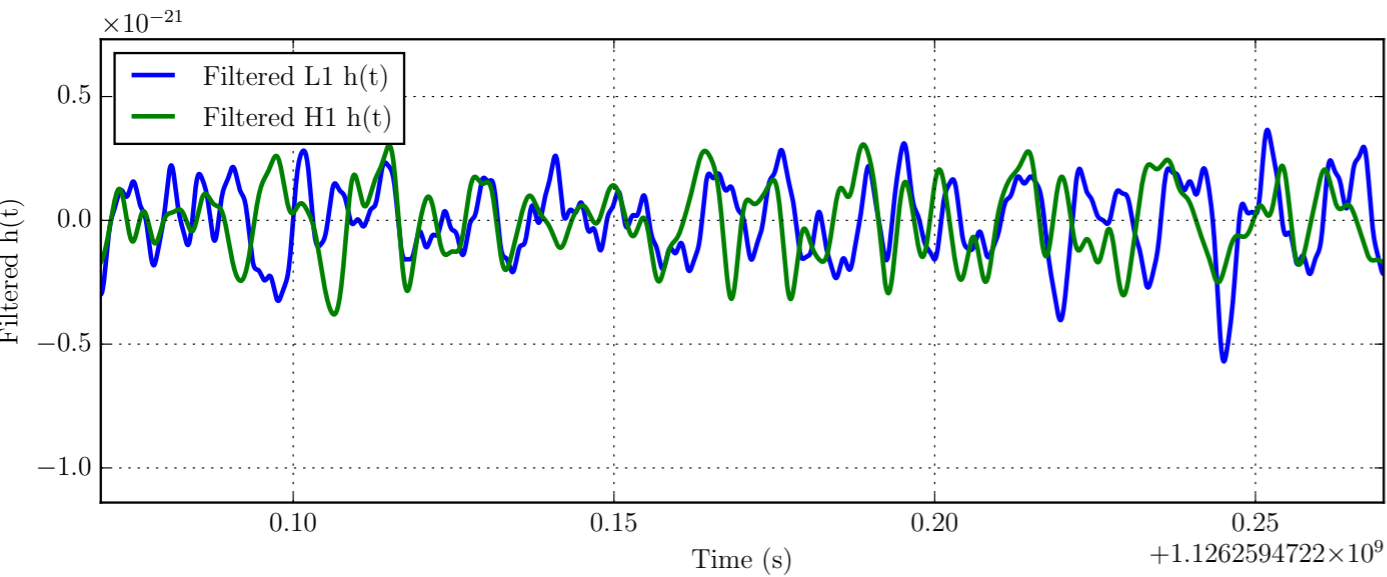
O1 and O2 noise lines paper: Covas et al. (2017) arXiv 1801.07204

O2 Instrumental lines:

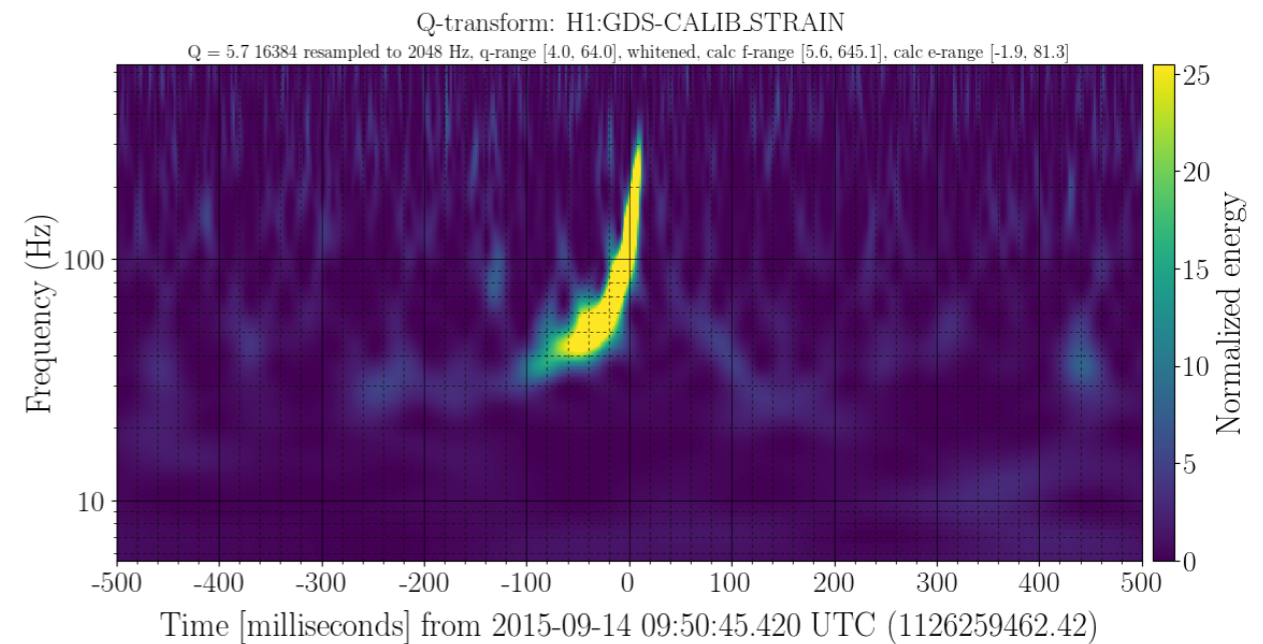
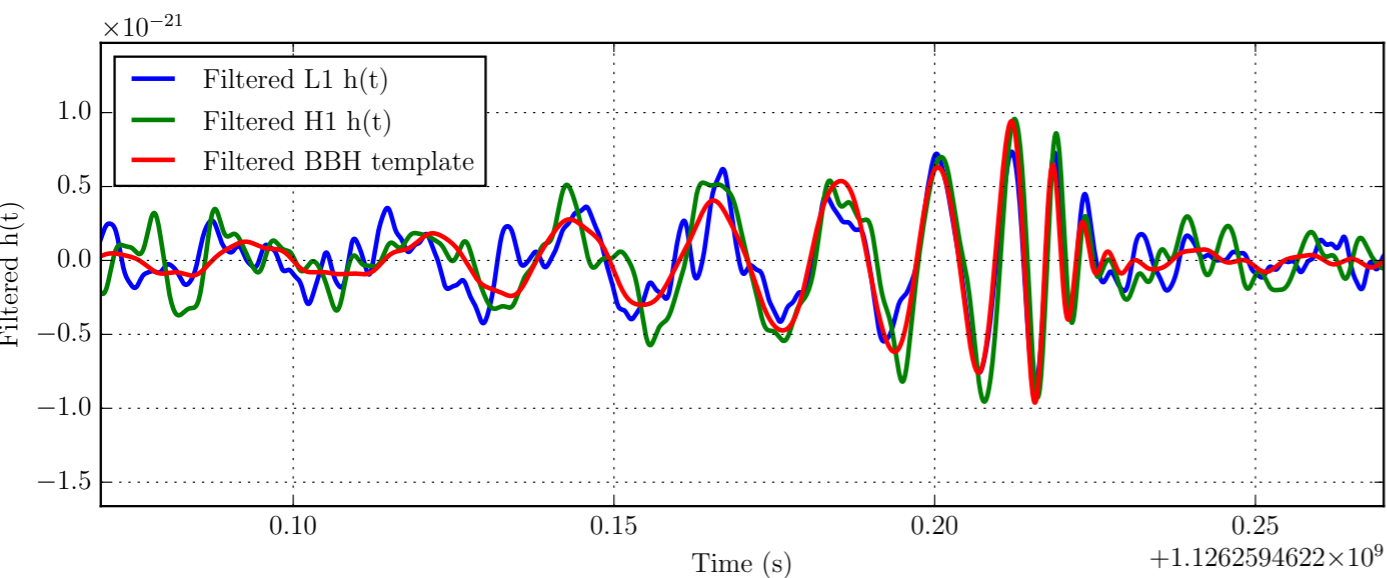
<https://www.gw-openscience.org/o2speclines/>

GW Detector data in a perfect world

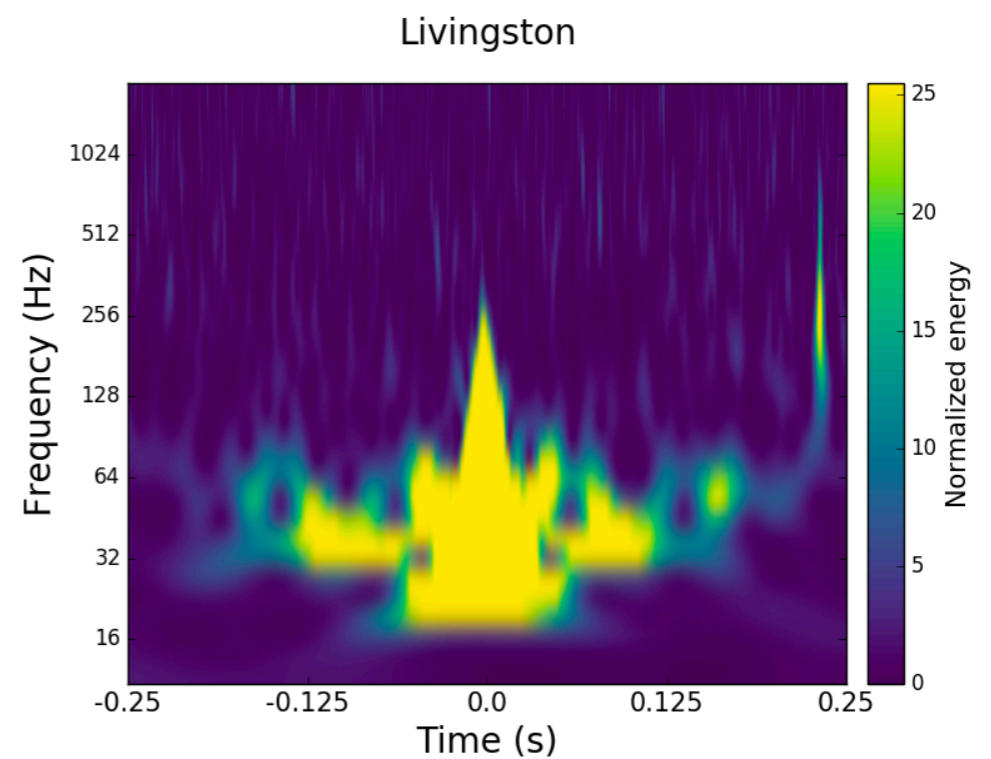
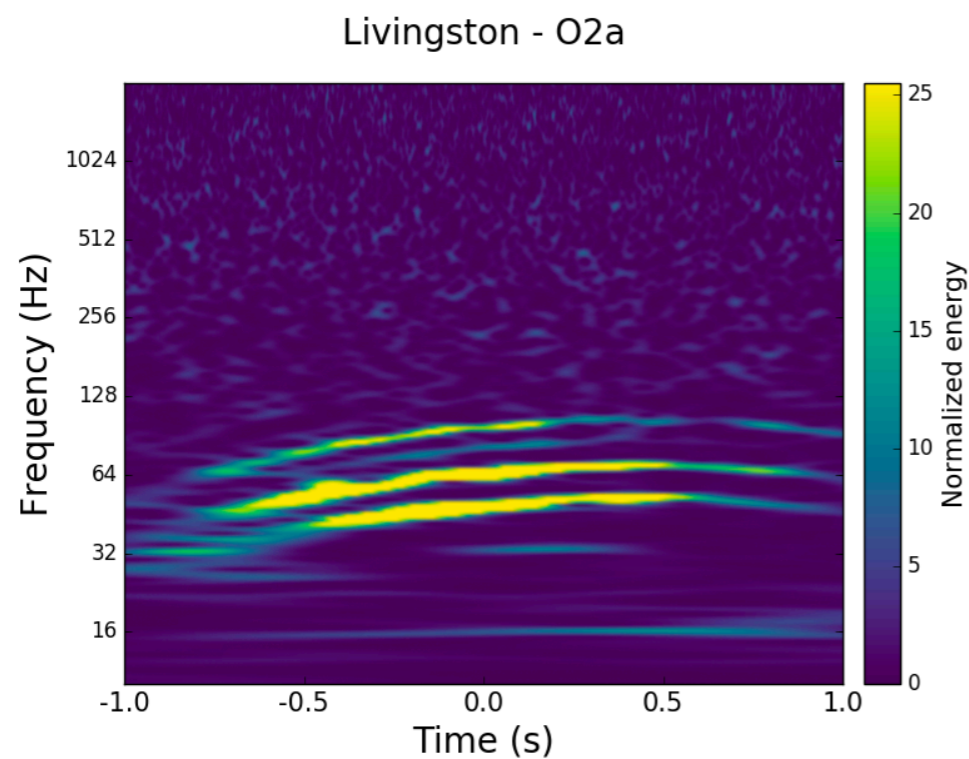
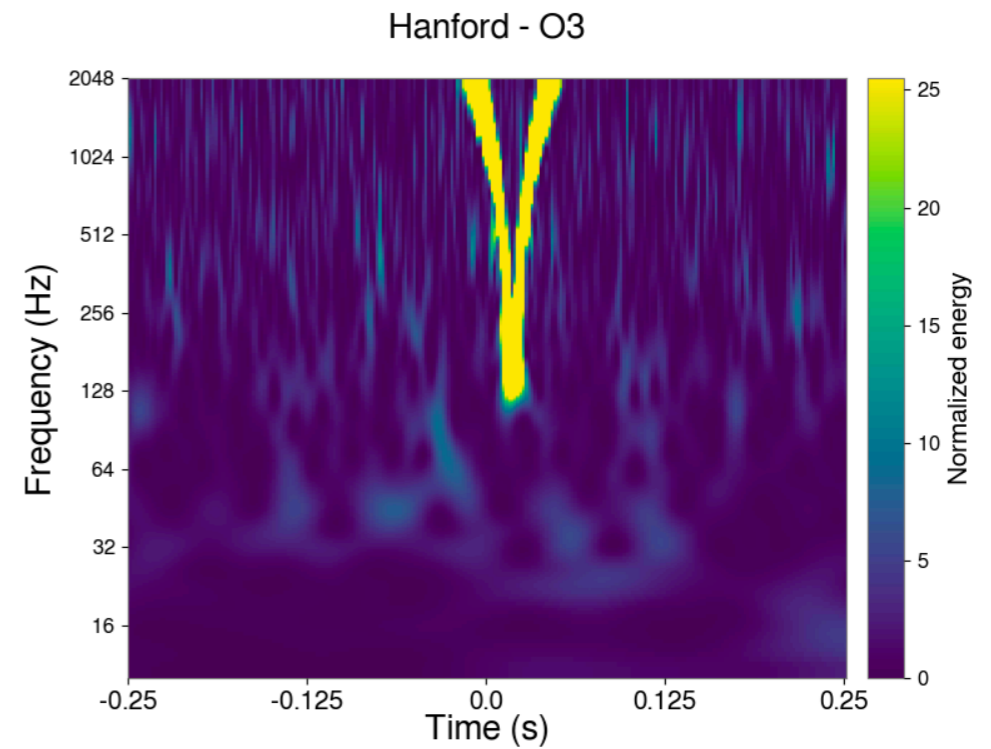
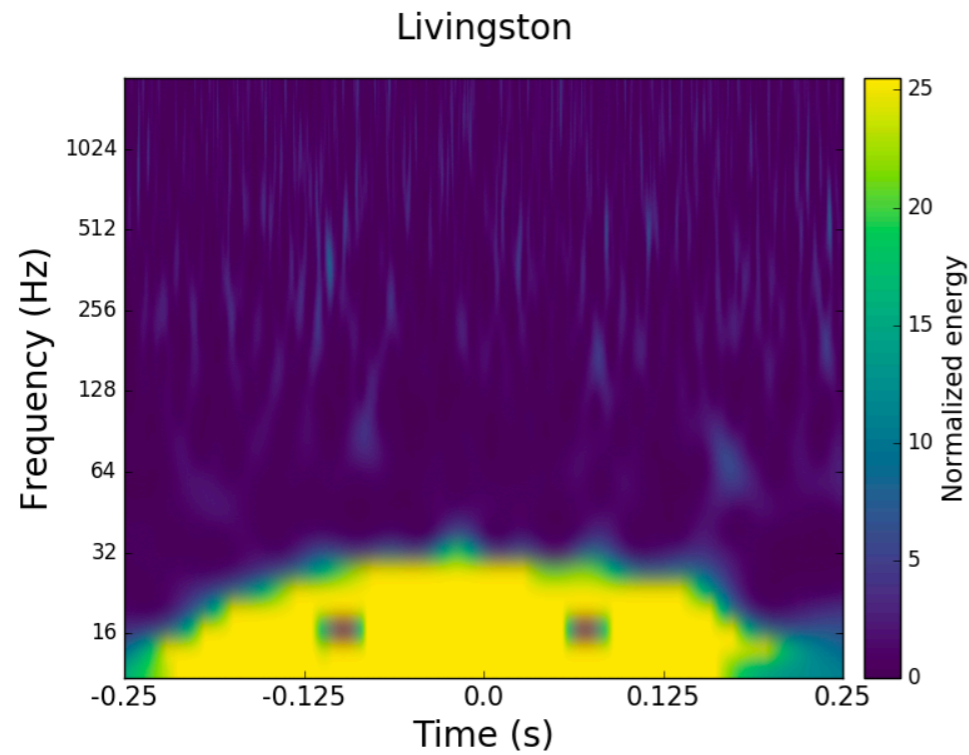
No signal



Signal



But in the real world we also have...



Common glitch types

gravityspy.org Zevin et al, 2017, CQG

GRAVITY SPY

ABOUT

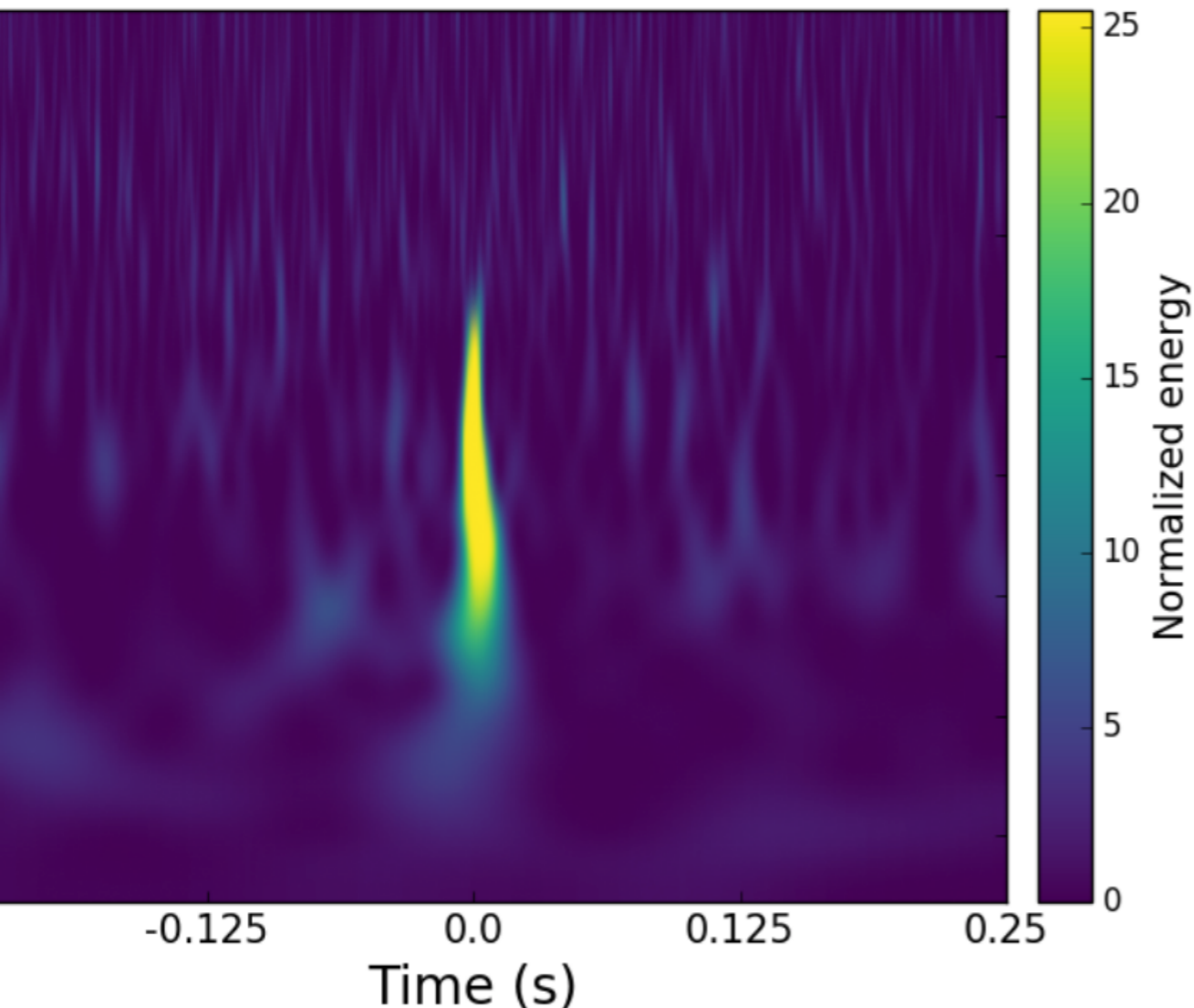
CLASSIFY

TALK

COLLECT

BLOG

Livingston



Duration

Frequency

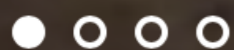
Event

| | |
|---|---|
| <input type="checkbox"/> Air Compressor (50 Hz) | <input type="checkbox"/> No Glitch |
| <input type="checkbox"/> Blip | <input type="checkbox"/> Paired Doves |
| <input type="checkbox"/> Chirp | <input type="checkbox"/> Power Line (60 Hz) |
| <input type="checkbox"/> Extremely Loud | <input type="checkbox"/> Repeating Blips |
| <input type="checkbox"/> Helix | <input type="checkbox"/> Scattered Light |
| <input type="checkbox"/> Koi Fish | <input type="checkbox"/> Scratchy |
| <input type="checkbox"/> Light Modulation | <input type="checkbox"/> Tomte |
| <input type="checkbox"/> Low Frequency Burst | <input type="checkbox"/> Violin Mode Harmonic |
| <input type="checkbox"/> Low Frequency Line | <input type="checkbox"/> Wandering Line |
| <input type="checkbox"/> None of the Above | <input type="checkbox"/> Whistle |

Showing 20 of 20.

Done & Talk

Done



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Data quality: noise artifacts in LIGO data

- Glitches
- Lines

Data quality information

- Detector status summary pages
- Data quality segments
- Event validation

Summary of resources and references

Daily detector status (available for O2 & O3)

https://www.gw-openscience.org/detector_status/day/20170817/

« August 17 2017 ▾ »

Summary

Home

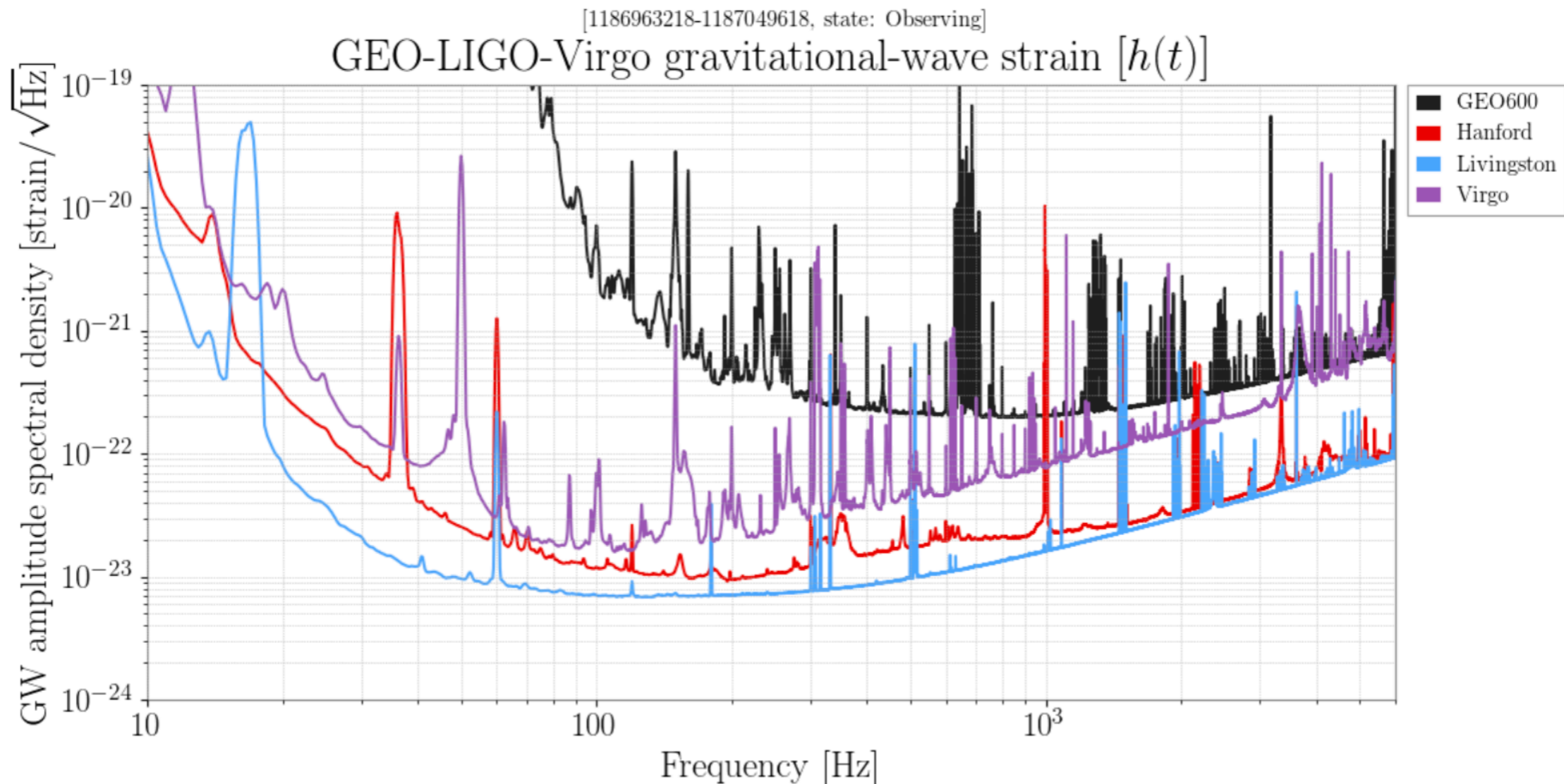
Environment ▾

Instrument performance ▾

Summary

Date selection

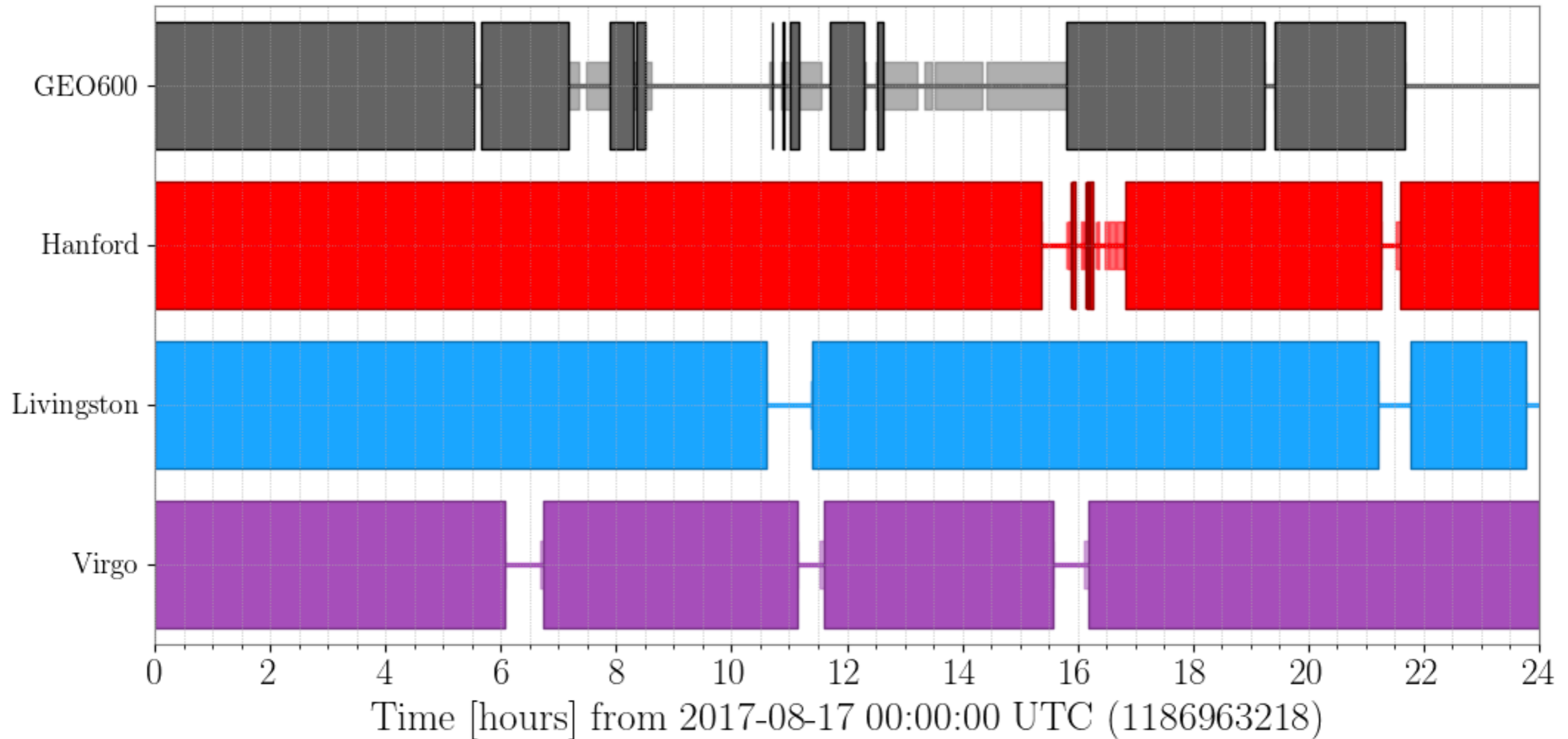
The plots shown below characterize the sensitivity and status of each of the LIGO interferometers as well as the [Virgo](#) detector in Cascina, Italy and the [GEO600](#) detector in Hanover, Germany. For more information about the plots listed below, click on an image to read the caption. Use the tabs in the navigation bar at the top of the screen for more detailed information about the LIGO, Virgo, and GEO interferometers.



Daily detector status (available for O2 & O3)

https://www.gw-openscience.org/detector_status/day/20170817/

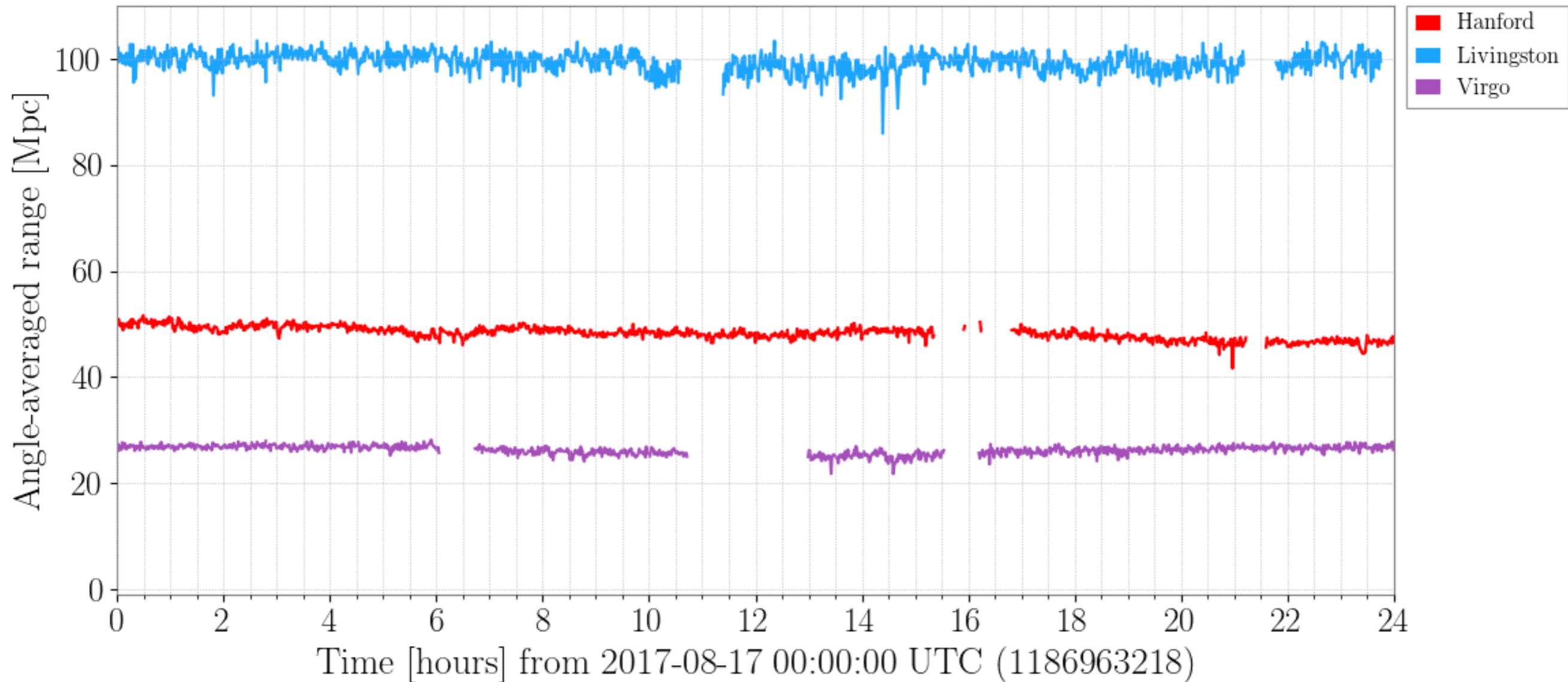
GEO-LIGO-Virgo operating segments



Daily detector status (available for O2 & O3)

https://www.gw-openscience.org/detector_status/day/20170817/

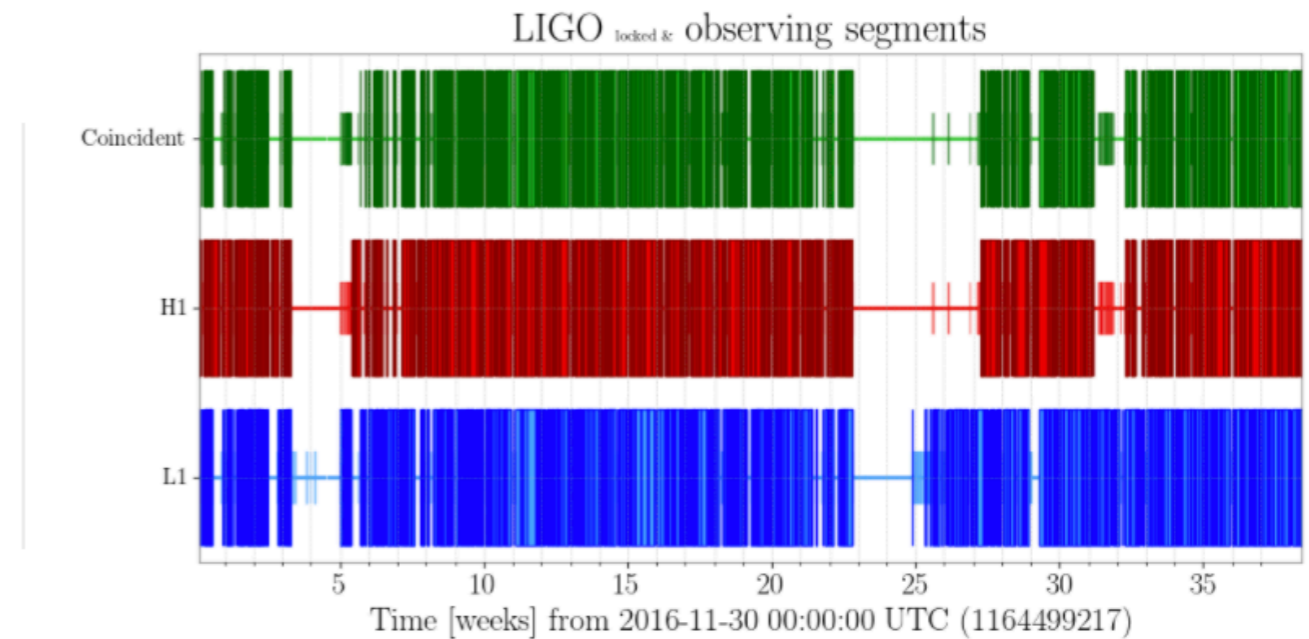
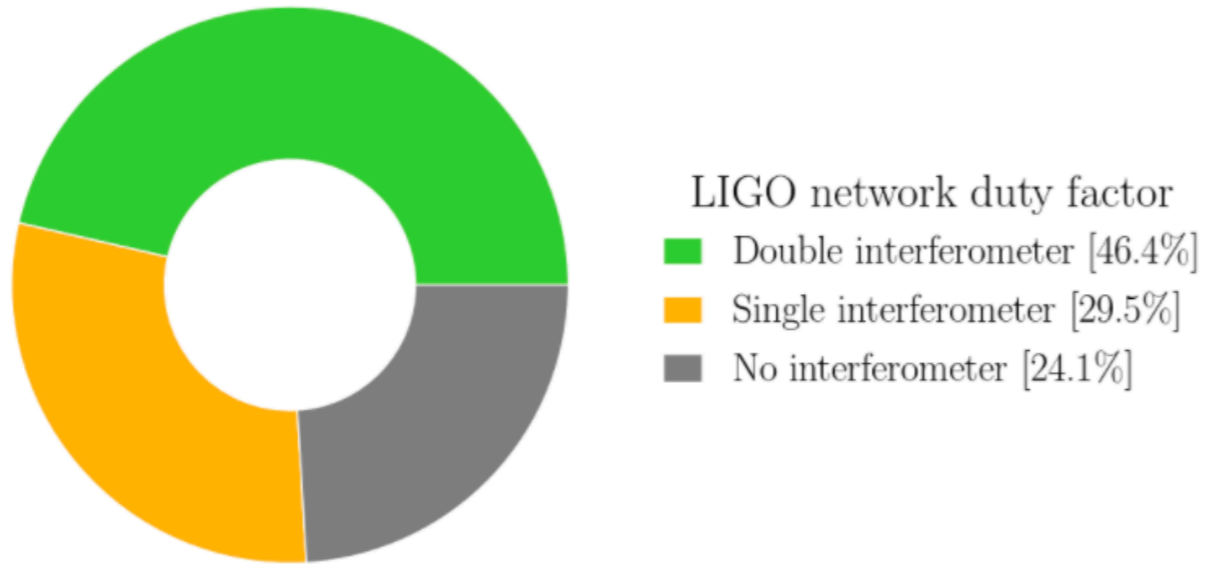
LIGO-Virgo binary neutron star inspiral range



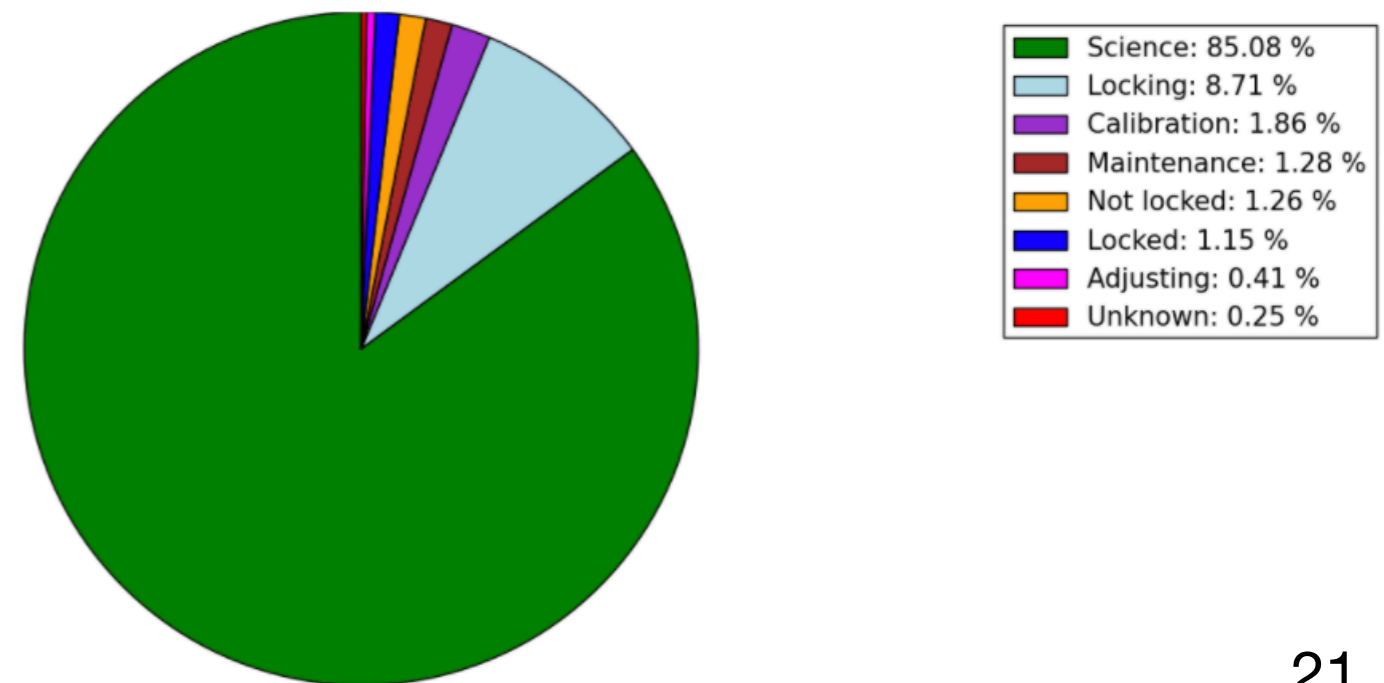
Observing run summaries (O1 or O2)

https://www.gw-openscience.org/detector_status/O2/

Includes summary plots of LIGO segments and sensitivity over the run



Also linked is a page showing Virgo's status in O2 (<http://www.virgo-gw.eu/O2.html>):



Data quality information

DATA (Data Available): Failing this level indicates that LIGO data are not publicly available because the instruments or data calibration were not operating in an acceptable condition.

CAT1 (Category 1): Failing a data quality check at this category indicates **a critical issue with a key detector component not operating in its nominal configuration.**

- These times are identical for each data analysis group.
- *Times that fail CAT1 flags are not available as LIGO open data.*

CAT2 (Category 2): Failing a data quality check at this category indicates times when there is a **known, understood physical coupling to the gravitational wave channel.** For example, high seismic activity.

CAT3 (Category 3): Failing a data quality check at this category indicates times when there is **statistical coupling to the gravitational wave channel** which is not fully understood.

Data quality levels are defined in a cumulative way: a time which fails a given category automatically fails all higher categories.

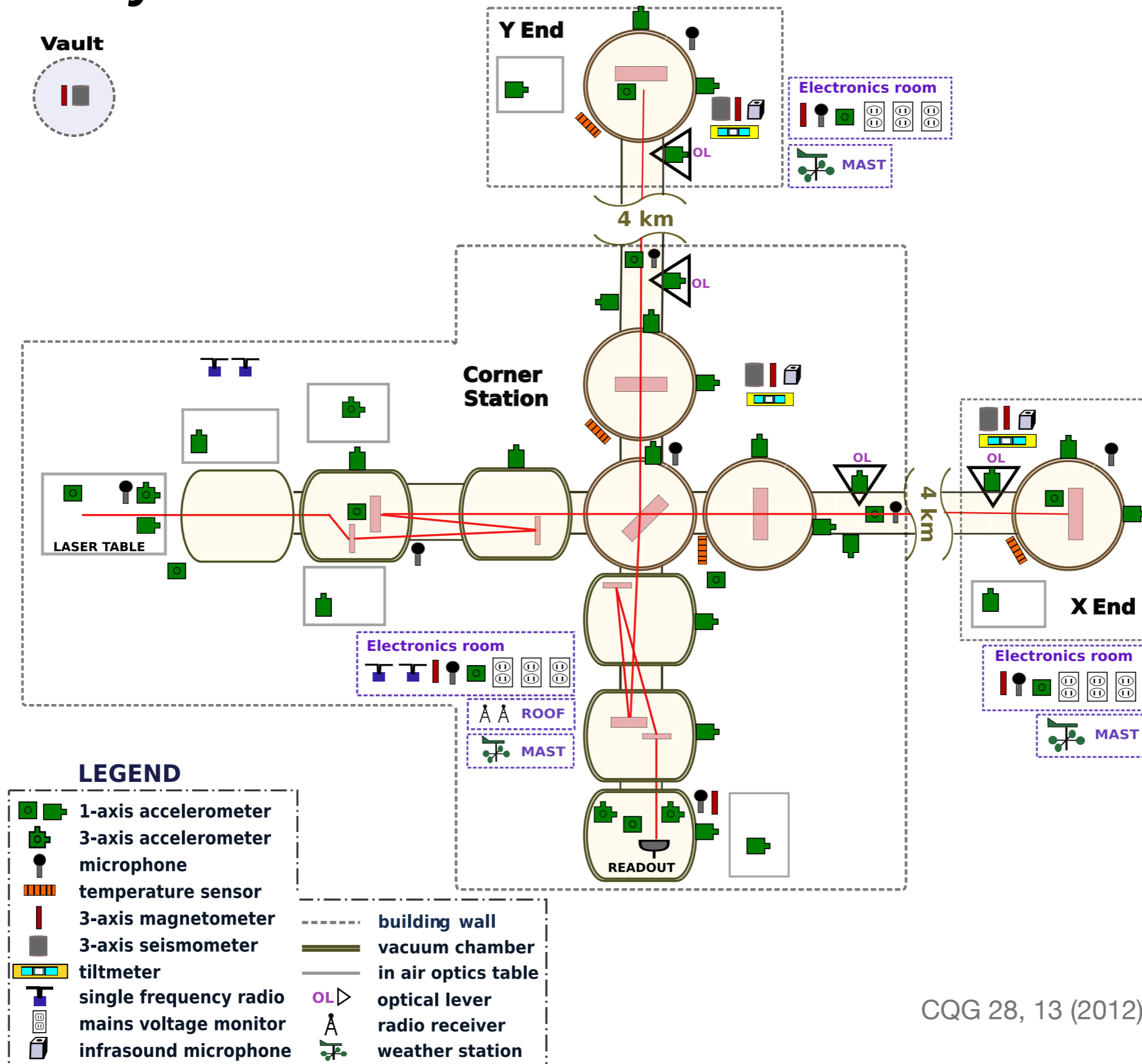
Data quality categories are defined independently for different analysis groups: if something fails at CAT2_BURST, it could pass CAT2_CBC.

Data quality information

Available via the GWOSC

| Bit Short Name | Description |
|--------------------------|-------------------------------|
| Data Quality Bits | |
| 0 DATA | data present |
| 1 CBC_CAT1 | passes the cbc CAT1 test |
| 2 CBC_CAT2 | passes cbc CAT2 test |
| 3 CBC_CAT3 | passes cbc CAT3 test |
| 4 BURST_CAT1 | passes burst CAT1 test |
| 5 BURST_CAT2 | passes burst CAT2 test |
| 6 BURST_CAT3 | passes burst CAT3 test |
| Injection Bits | |
| 0 NO_CBC_HW_INJ | no cbc injection |
| 1 NO_BURST_HW_INJ | no burst injections |
| 2 NO_DETCHAR_HW_INJ | no detchar injections |
| 3 NO_CW_HW_INJ | no continuous wave injections |
| 4 NO_STOCH_HW_INJ | no stoch injections |

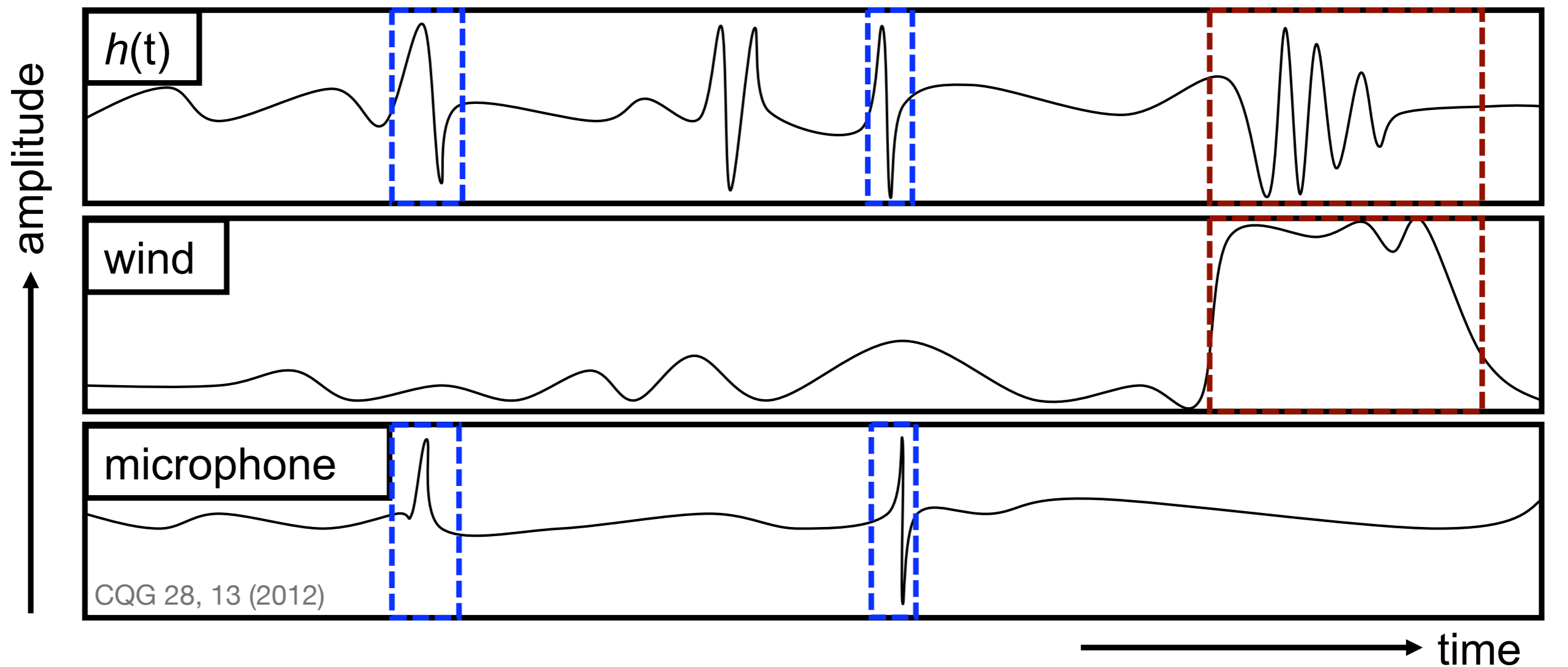
Physical environment channels



Auxiliary channels

We record **over 200,000 channels per detector** that monitor the environment and detector behavior.

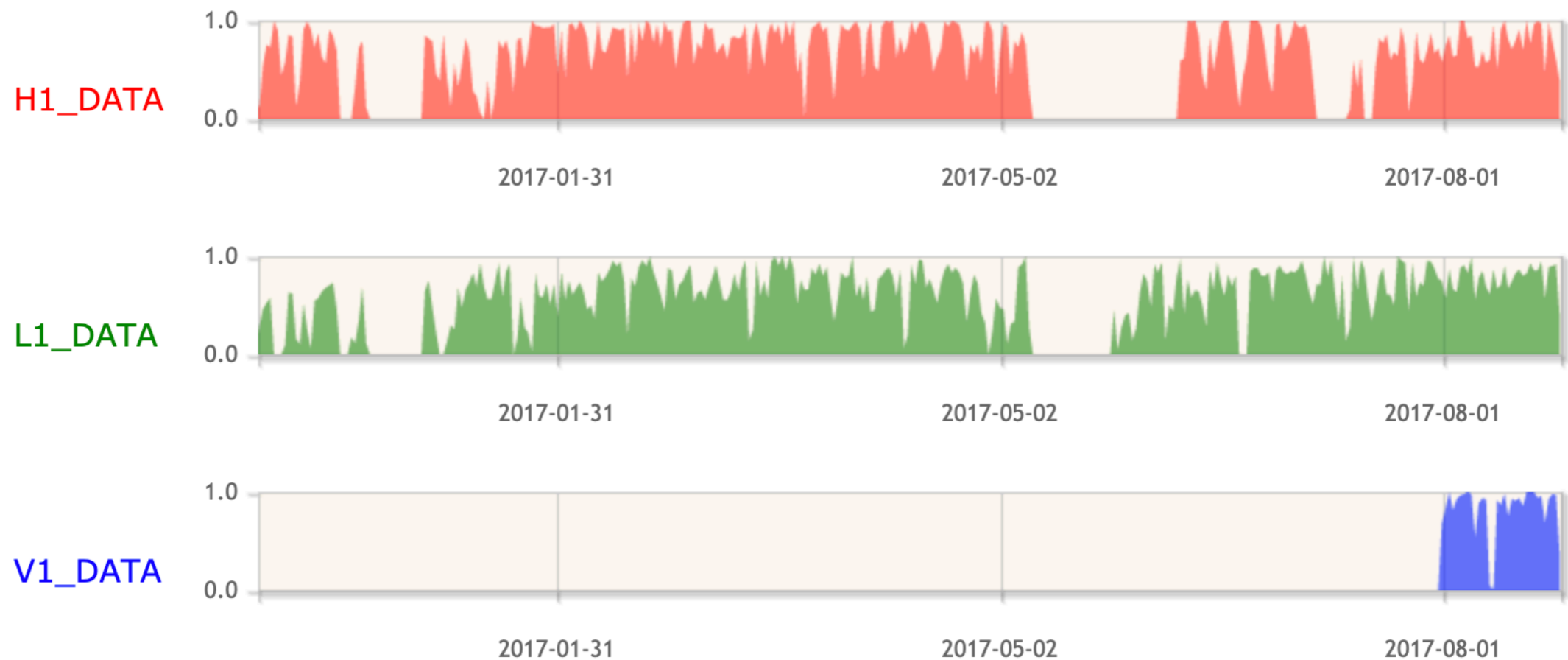
We can use these to **help trace the instrumental causes of glitches** that pollute the search backgrounds.



How to get Data Quality Segments Using GWOSC Timeline Query

<https://www.gw-openscience.org/timeline/query/Run/>

Here I selected the O2 data flags for H1, L1, and V1



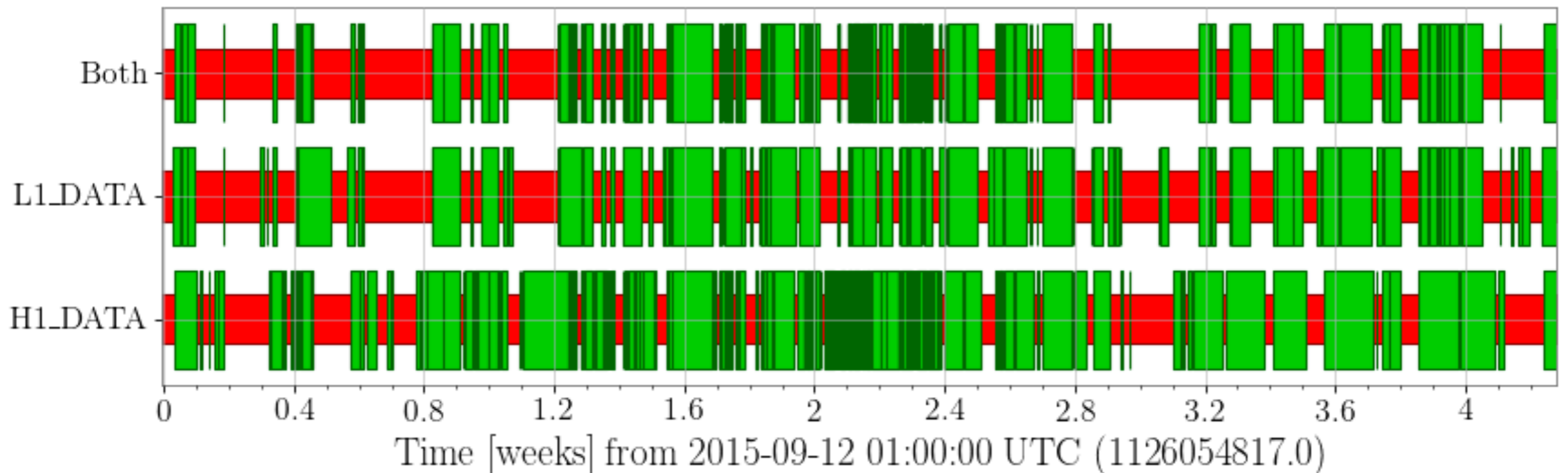
Segments can be plotted (with interactive zooming) or downloaded₂₆

How to get Data Quality Segments

Using GWpy

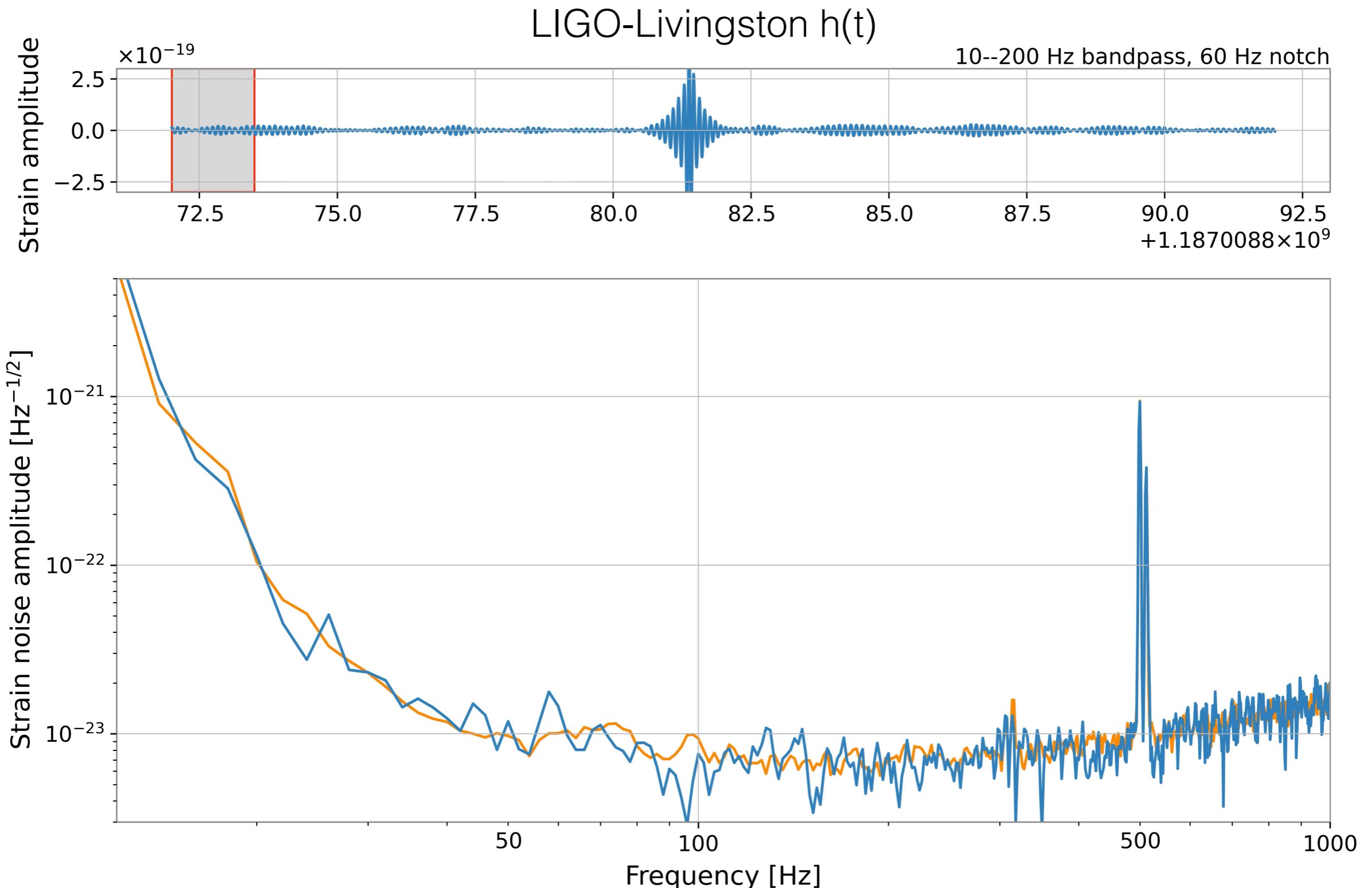
Example showing how to find and plot data quality segments from O1:

<https://gwpy.github.io/docs/stable/examples/segments/open-data.html>



Search-specific data quality segments

Example: GW170817

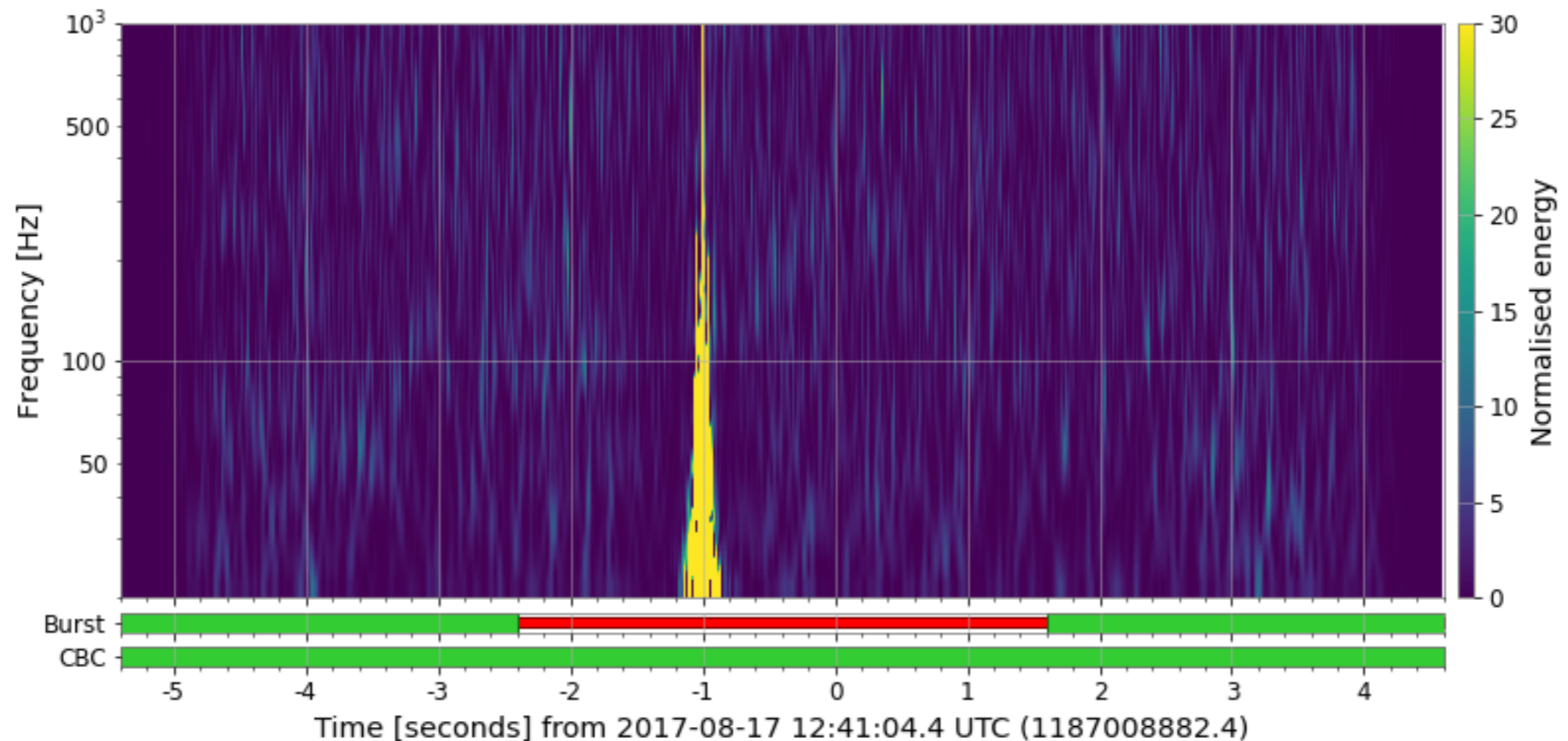


Modified from Duncan's original script. Notebook: https://github.com/marissawalker/data-quality-odw-2020/blob/master/LIGO_data_animation.ipynb

Search-specific data quality segments

Example: GW170817

LIGO-Livingston $h(t)$



Notebook: https://github.com/marissawalker/data-quality-odw-2020/blob/master/GW170817_glitch_qscans.ipynb

In Tutorial 1.3 you will learn to make a version of this plot that will show the BNS signal in this plot, not just the glitch

Data quality for event validation: examples from GWTC-1

GWTC-1: A Gravitational-Wave Transient Catalog of Compact Binary Mergers Observed by LIGO and Virgo during the First and Second Observing Runs

B.P. Abbott *et al.* (LIGO Scientific Collaboration and Virgo Collaboration)

Phys. Rev. X **9**, 031040 – Published 4 September 2019

DOI: <https://doi.org/10.1103/PhysRevX.9.031040>

See talks by Derek Davis and Alan Weinstein tomorrow for details on the searches for compact binary coalescences (CBC)

Example: MC170720

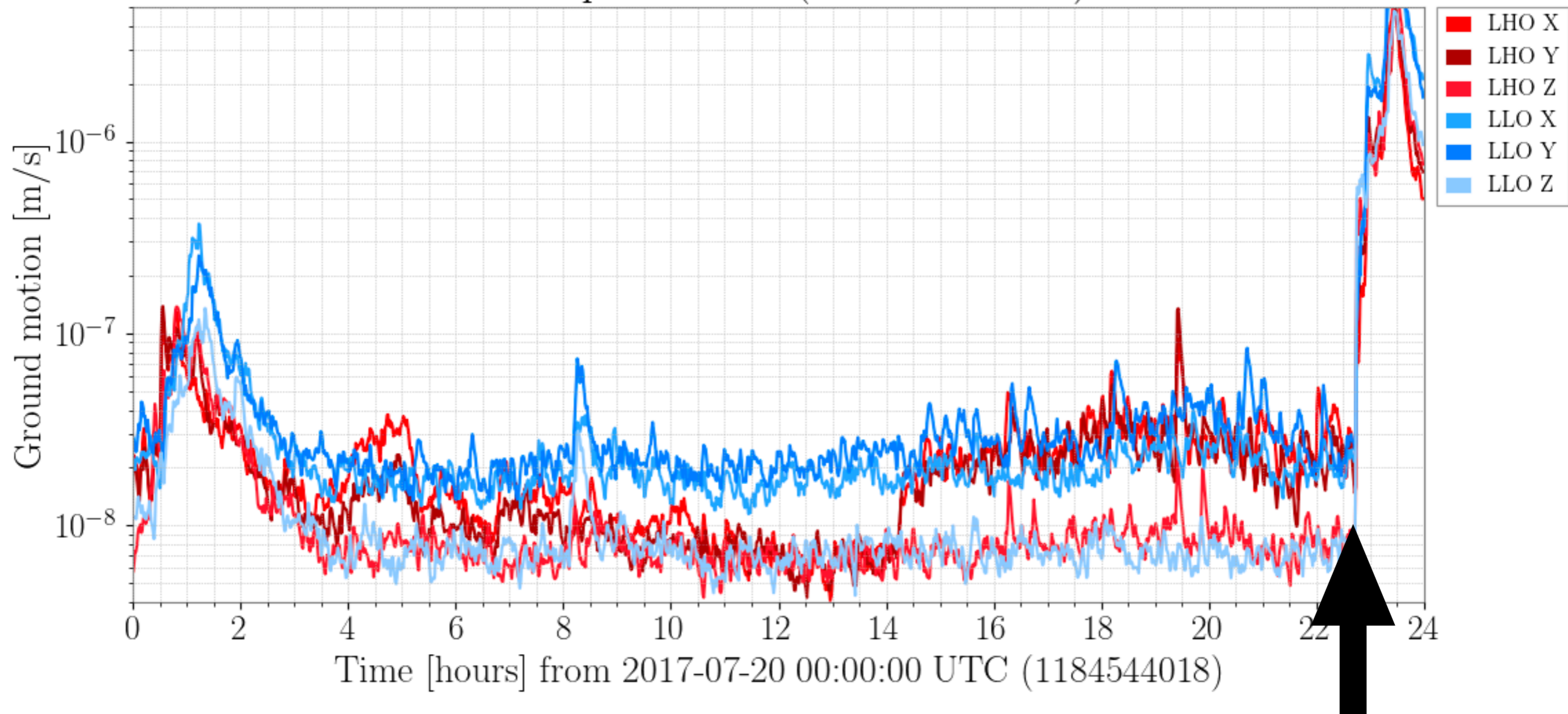
Marginal trigger most likely of terrestrial origin

Within 20 s of trigger 170720, excess ground motion from earthquakes forced the Livingston detector to drop out of its nominal mode of operation. Before the detector dropped out of the observing state, the data are heavily polluted with scattering artifacts that could account for the SNR of the triggers. As the PyCBC search does not consider times near the edges of observing periods, this time period is also not analyzed by that search. Artifacts related to scattered light are also observed at Hanford at this time. GWTC-1 p. 11

Example notebook: <https://github.com/marissawalker/data-quality-odw-2020/blob/master/MC170720.ipynb>

Example: MC170720

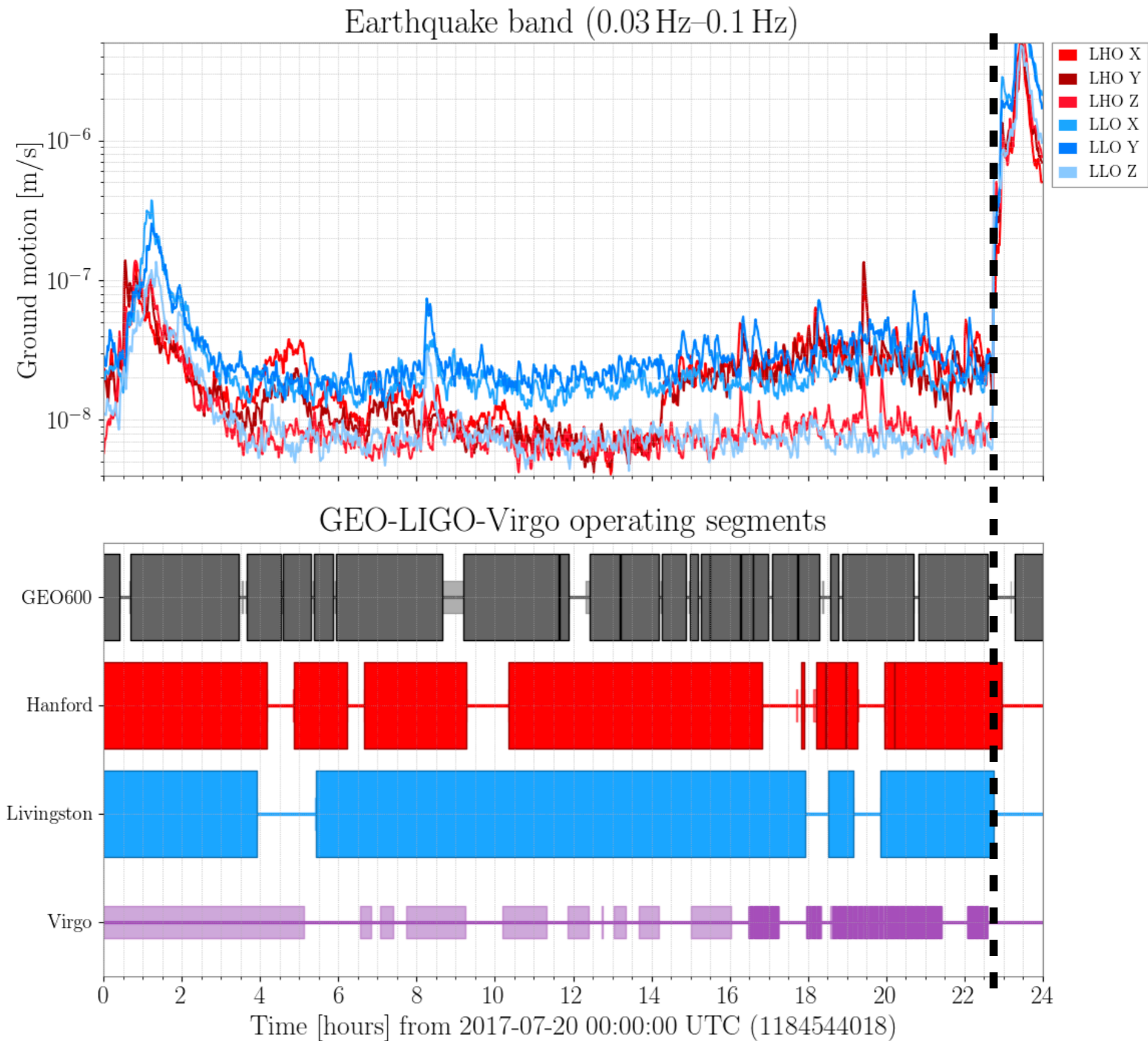
Earthquake band (0.03 Hz–0.1 Hz)



https://www.gw-openscience.org/detector_status/day/20170720/environment/ground_motion/

Within 20 s of trigger 170720, excess ground motion from earthquakes forced the Livingston detector to drop out of its

Example: MC170720



Example: MC170720

Marginal trigger most likely of terrestrial origin

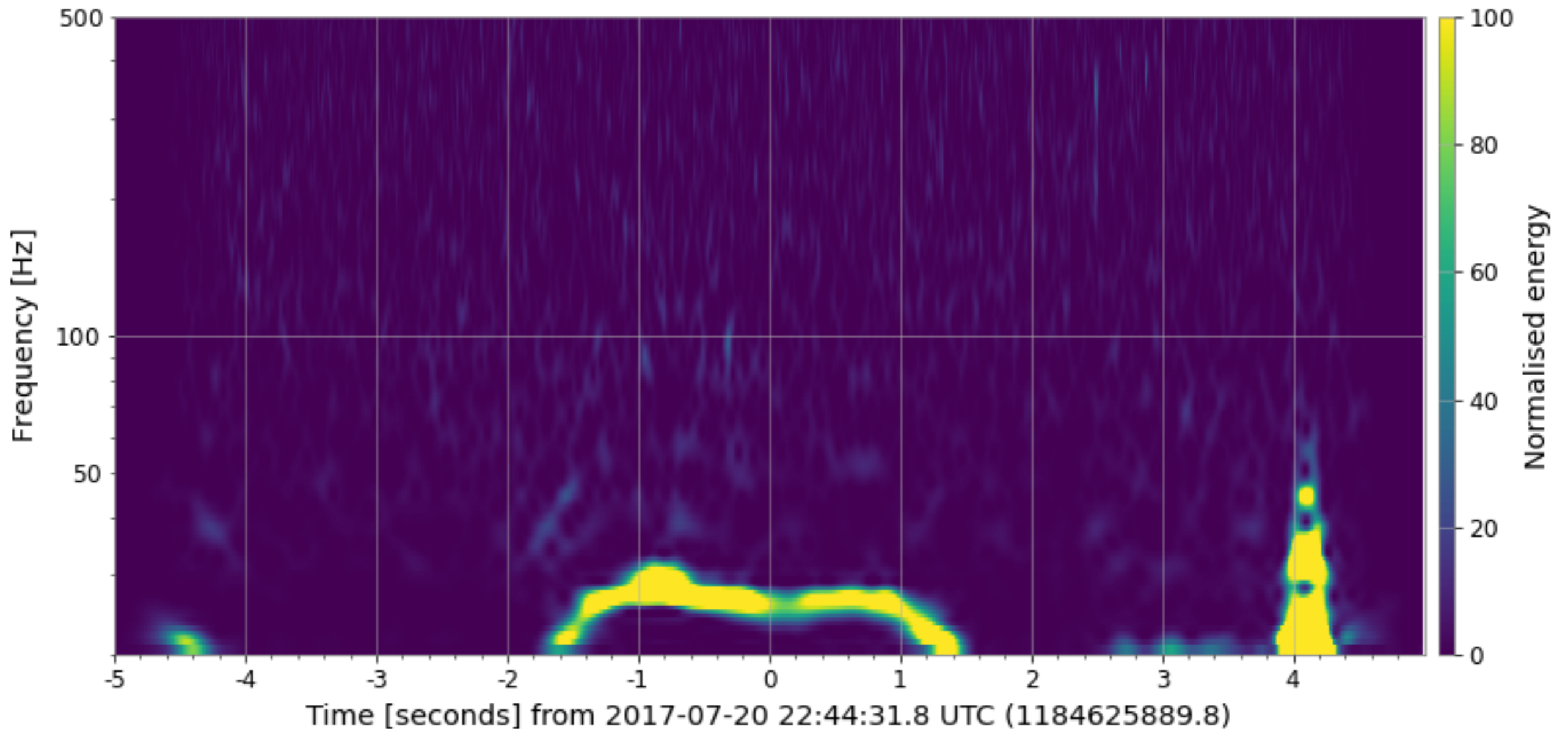
marginal trigger. Because of the short observing duration, this time period is not analyzed by the PyCBC search.

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Example: MC170720

Marginal trigger most likely of terrestrial origin

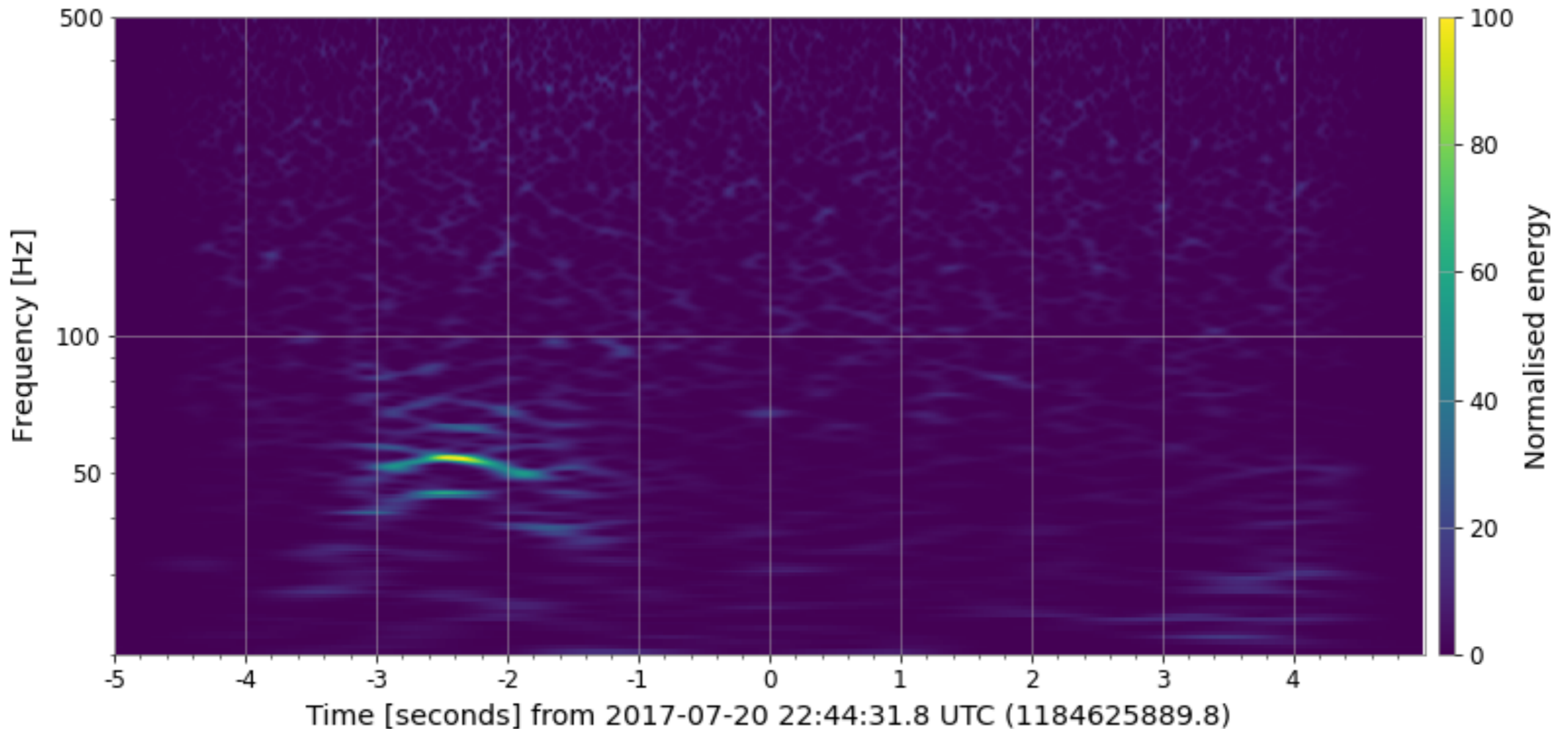
Livingston



Example: MC170720

Marginal trigger most likely of terrestrial origin

Hanford



Useful data quality references

For glitches:

GW150914 Detector characterization paper: [arXiv 1602.03844](https://arxiv.org/abs/1602.03844)

O1 CBC DQ paper; CQG (2018): <http://iopscience.iop.org/article/10.1088/1361-6382/aaaafa/meta>

Gravity Spy: gravityspy.org

For lines:

O1/O2 lines paper: [arXiv 1801.07204](https://arxiv.org/abs/1801.07204)

O2 lines catalog on the GWOSC: <https://www.gw-openscience.org/o2speclines/>

Data quality segments:

Data quality timelines: <https://www.gw-openscience.org/timeline/>

Public interferometer status monitoring: https://www.gw-openscience.org/detector_status/

O3 public alerts: <https://gracedb.ligo.org/superevents/public/O3/>

Example notebooks used in this talk:

<https://github.com/marissawalker/data-quality-odw-2020>

GWpy documentation: <https://gwpy.github.io/>

Thank you!

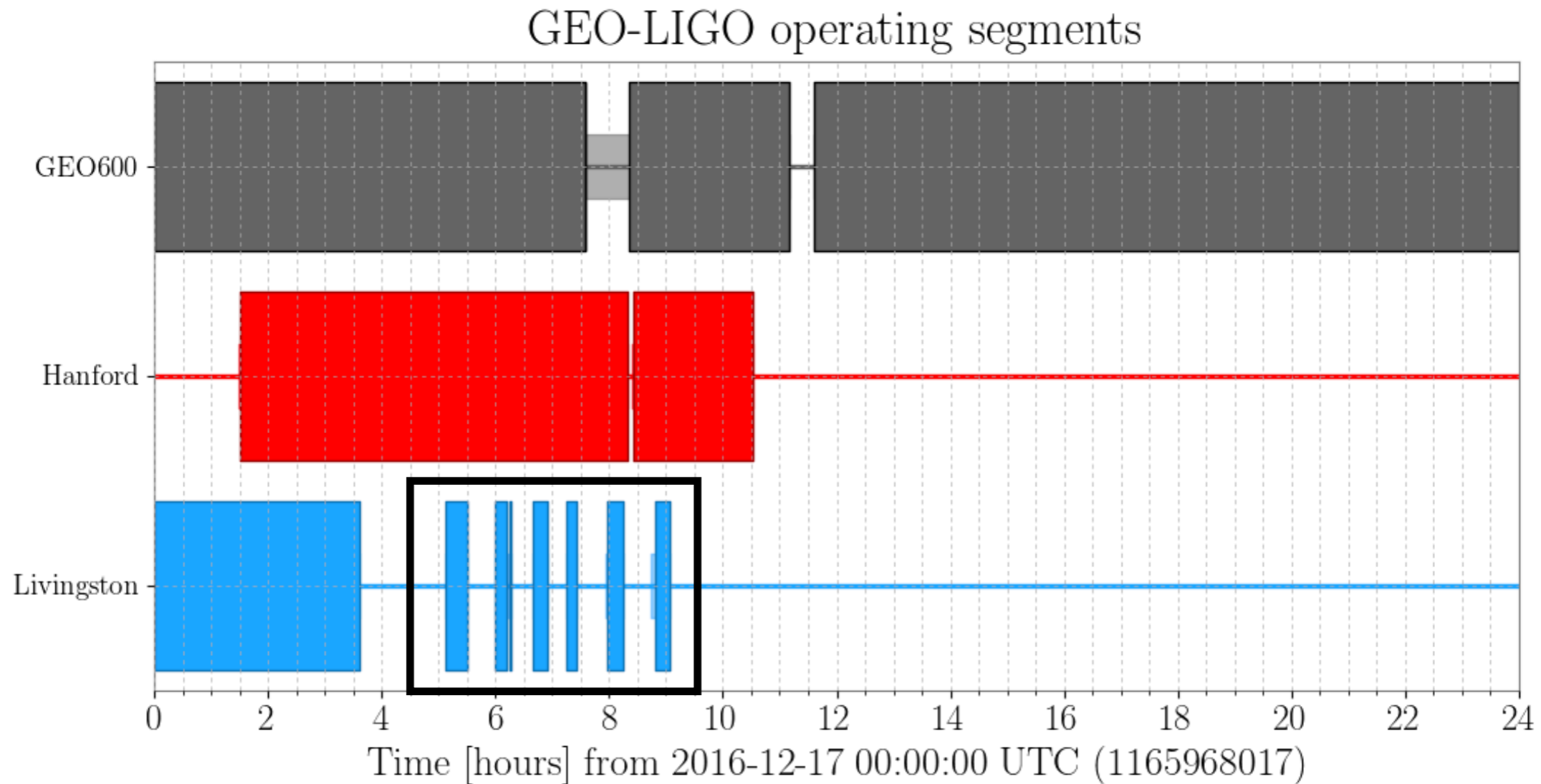
Example: MC161217

Marginal trigger most likely of terrestrial origin

the observatories at the times in question identified high-amplitude ground motion correlated with the scattering.

The marginal trigger 161217 occurs during a period of high-amplitude ground motion at Livingston caused by storm activity. During this storm activity, the Livingston detector is not able to maintain a stable interferometer for periods longer than 10 min. The presence of intense scattering artifacts contribute to the unstable state of the interferometer and could account for the SNR of the marginal trigger. Because of the short observing duration, this time period is not analyzed by the PyCBC search.

Example: MC161217



https://www.gw-openscience.org/detector_status/day/20161217/instrument_performance/analysis_time/

storm activity. During this storm activity, the Livingston detector is not able to maintain a stable interferometer for periods longer than 10 min. The presence of intense

Example: MC161217

Marginal trigger most likely of terrestrial origin

the observatories at the times in question identified high-amplitude ground motion correlated with the scattering.

The marginal trigger 161217 occurs during a period of high-amplitude ground motion at Livingston caused by storm activity. During this storm activity, the Livingston detector is not able to maintain a stable interferometer for periods longer than 10 min. The presence of intense scattering artifacts contribute to the unstable state of the interferometer and could account for the SNR of the marginal trigger. Because of the short observing duration, this time period is not analyzed by the PyCBC search.

Within 20 s of trigger 170720 excess ground motion from

Try it yourself! MC161217 data quality

Using the example notebook about MC170720, explore the data quality of MC161217, and compare with the statements made in GWTC-1:

- Check the length of Livingston's data segment
- Using the q transform, make a spectrogram to look for evidence of scattering