

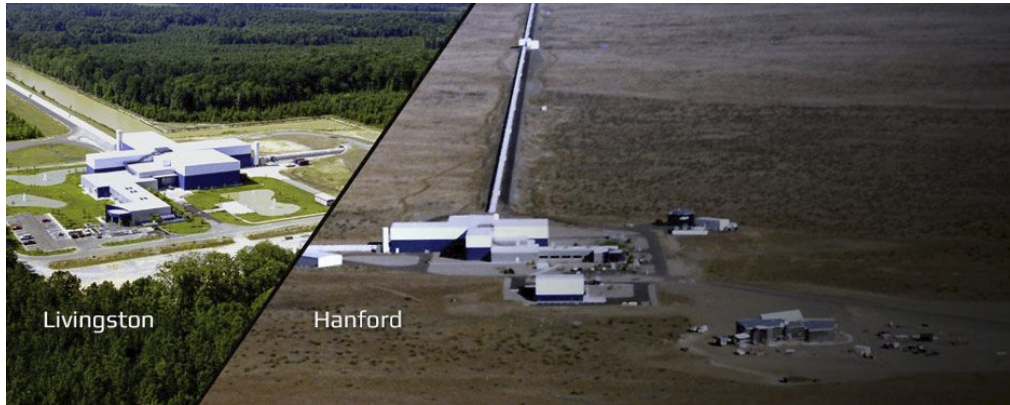


# Mitigating Laser Damage in LIGO Optics

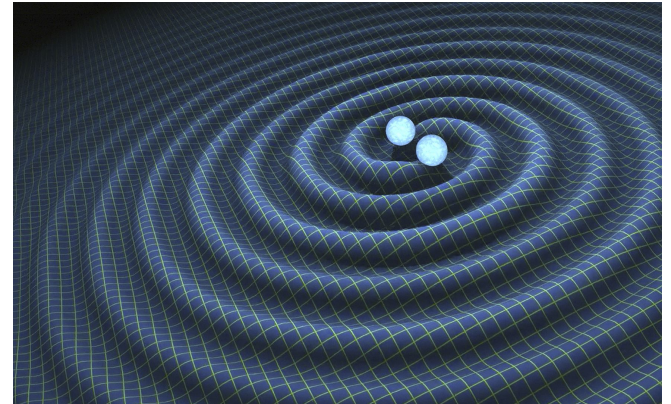
Jasmine Terrones  
Mechanical Engineering '22



# Laser Interferometer Gravitational-Wave Observatory



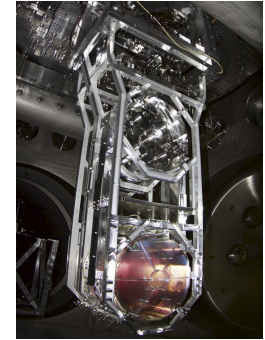
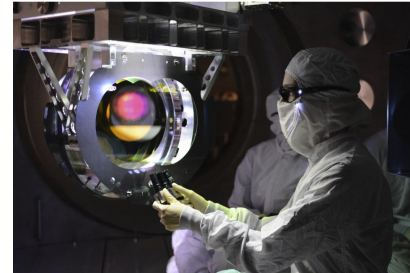
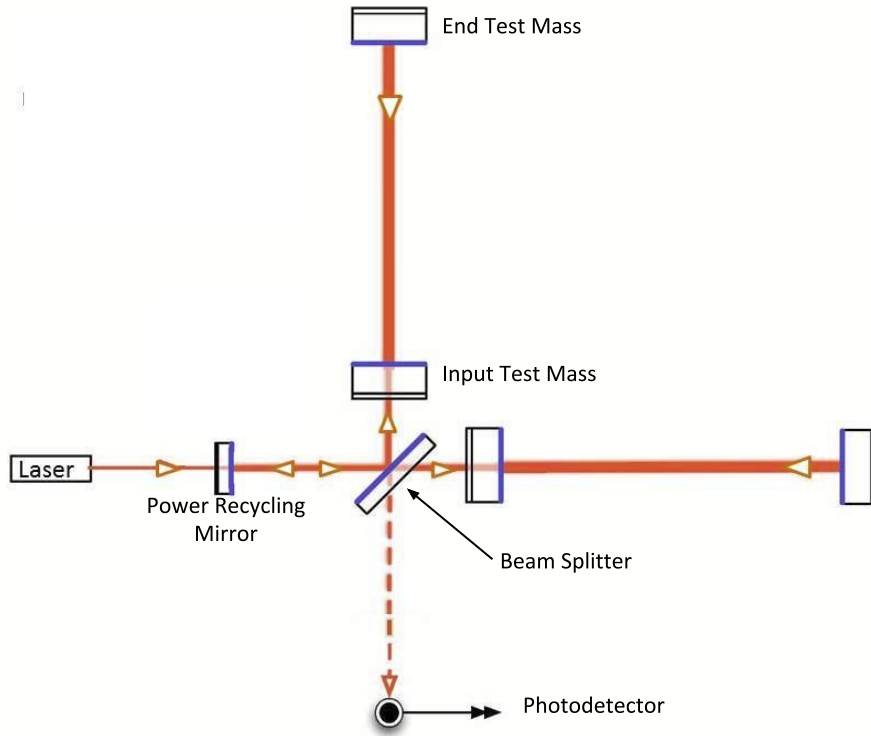
<https://www.ligo.caltech.edu/>



<https://www.nasa.gov/feature/goddard/2016/nsf-s-ligo-has-detected-gravitational-waves/>

# LIGO Optics and Laser Intensity Levels

<https://www.ligo.caltech.edu/news/ligo20191001>



Optic	Estimated Intensity [W/mm <sup>2</sup> ]
Beam Splitter	0.3445
ETM	33.98
ITM	46.84
PRM	199.9

<https://www.ligo.caltech.edu/page/ligos-ifo>

# Stray Light Control - Absorbing Materials

- Beam dumps and baffles absorb scattered infrared light
- Black glass used in critical areas
  - Very low reflectivity compared to other materials
  - Extremely fragile and difficult to install
  - Overheats at high power
- Different materials used at higher power
- Search for better materials and coatings ongoing

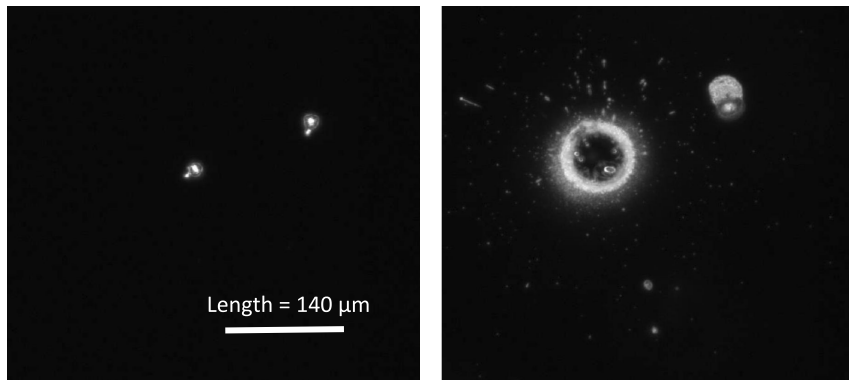
<https://www.ligo.caltech.edu/news/ligo20191104>



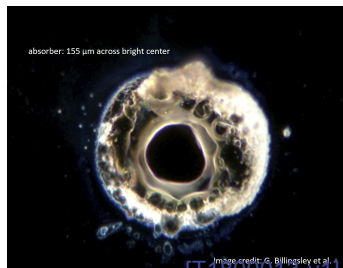
DLC coated SSTL baffle  
installation

# Contamination and Laser Damage

Particle pre and post irradiation at 400 W/mm<sup>2</sup>



Gushwa et al. [\[P1400205-v8\]](#)



Billingsley et al. [\[P1400015-v1\]](#)

- Dust particles melt into optic coatings at 92 W/mm<sup>2</sup> [Billingsley et al., G1400209-v2]
  - Test masses, power recycling mirror, and mode cleaner regularly reach higher intensities
  - Multiple defects found on these optics
- Larger particles (10 μm) and higher power more likely to yield bigger pits
- Low level contaminants can still cause damage in optic coatings
- Increasing cleanliness procedures can only reduce contamination by so much

# Laser Damage Facility (Testing Set-Up)

- In vacuum and in air testing capabilities
- Testing intensities ranging from 50 mW/mm<sup>2</sup> to 1 kW/W/mm<sup>2</sup>
- Equipped with XY translation stages for raster scans
- Damage captured with Nikon Scanning Darkfield microscope
- Higher power laser needed for contamination tests (and more resistant filters, beam dumps, etc.)

Vacuum system



Ananyeva et al.  
[\[G1800420-v1\]](#)

Sample for testing

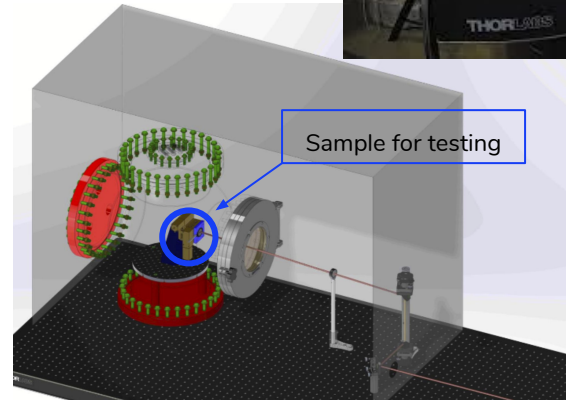


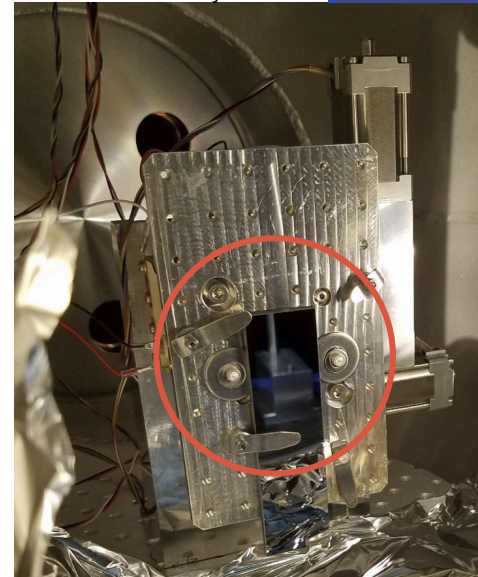
Image of the setup at Caltech and  
modelled in Solidworks

# Evaluating Different Absorbing Materials

- Studying five different materials for upgrading stray light absorbing optics [Ananyeva et al., G1800420-v1]
- Irradiated with CW 1064 nm
- Each sample irradiated at 10 points (5 min per point) at increasing power levels
  - Power increased until damage occurred
- Samples were cleaned prior to irradiation

Material	Cost	Reflectivity	Backscatter
Black glass	\$\$	4%	$2 \times 10^{-4}$
Diamond-like Carbon Coated SSTL	\$\$	6%	$2 \times 10^{-4}$
Silicon	\$\$\$\$	10%	$8 \times 10^{-5}$
Sintered Silicon Carbide	\$\$\$	Diffuse	High

Ananyeva et al. [\[E1700388-v3\]](#)



Silicon sample installed in vacuum chamber

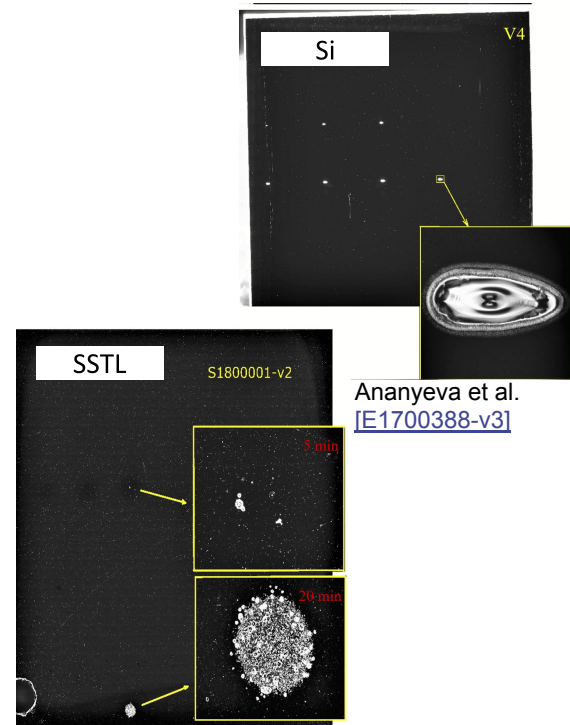
Ananyeva et al. [\[T1800001-v2\]](#)

# Damage Thresholds and Analysis

Material	Damage Threshold [W/mm <sup>2</sup> ]	Total Power [W]
Black glass	0.3	2.86
Diamond-like Carbon Coated SSTL	5.04	31.8
Silicon	371.13	16
Sintered Silicon Carbide	736.97	37.05

Ananyeva et al. [\[T1800001-v2\]](#)

- SiC can sustain high temperatures but is more reflective than black glass
  - A good option for areas with high power
  - Alternatively, sintered SiC is inexpensive but cannot be polished
- DLC SSTL should be used at low power areas
  - Has low reflectivity at small incident angles

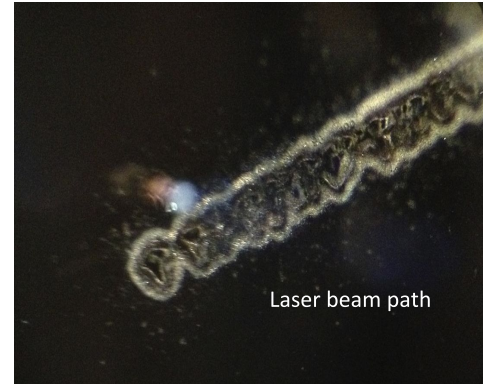


Ananyeva et al.  
[\[E1700388-v3\]](#)



# Research on Contamination and Laser Damage Ongoing

- Currently preparing for contamination experiments
  - How clean should LIGO optics be?
  - Utilizing similar set up as the absorbing material experiment
- Updating laser damage facility and additional components for high power testing (100 W)
- Modelling new laser set up on Zemax
  - Running simulations to model damage with Comsol



Billingsley et al. [\[G1400209-v2\]](#)

# Conclusion



Thank you for listening!