

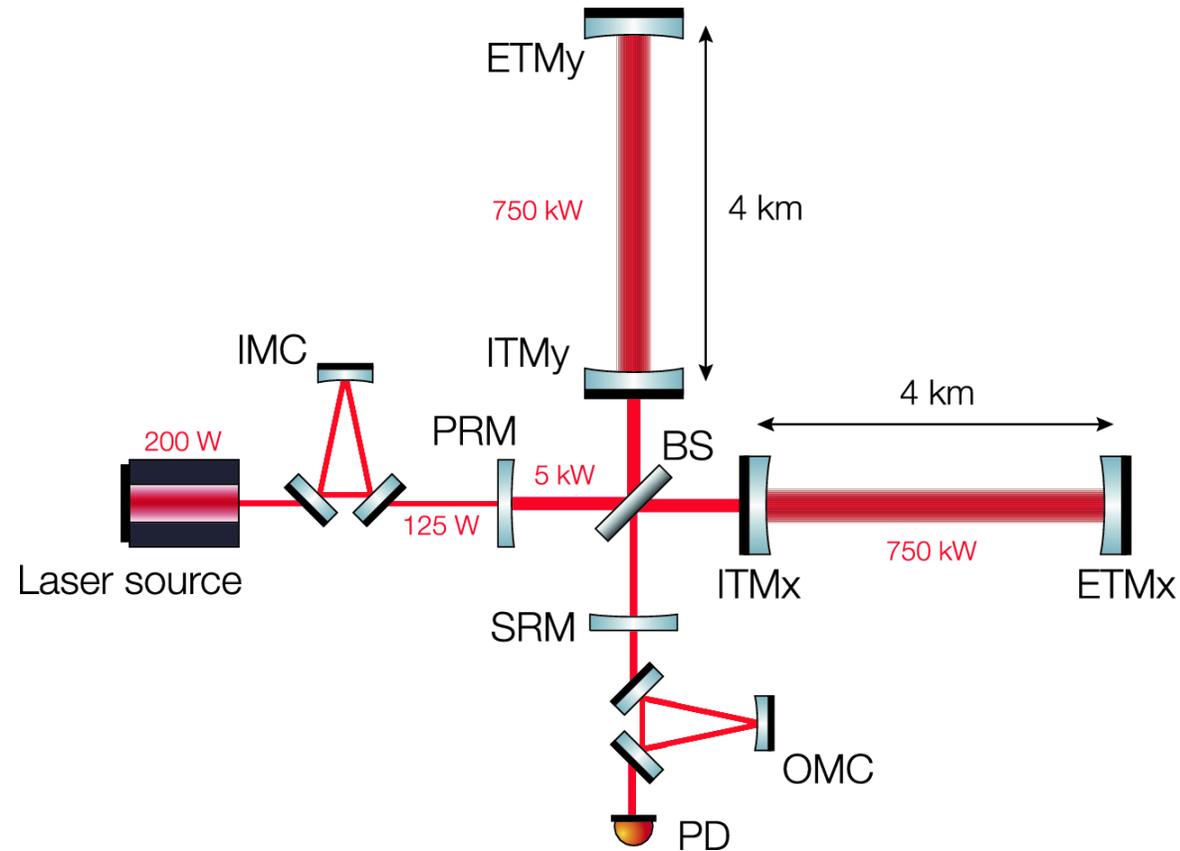


Mitigating LIGO Point Absorbers with High-Order Thermal Compensation

Edita Bytyqi

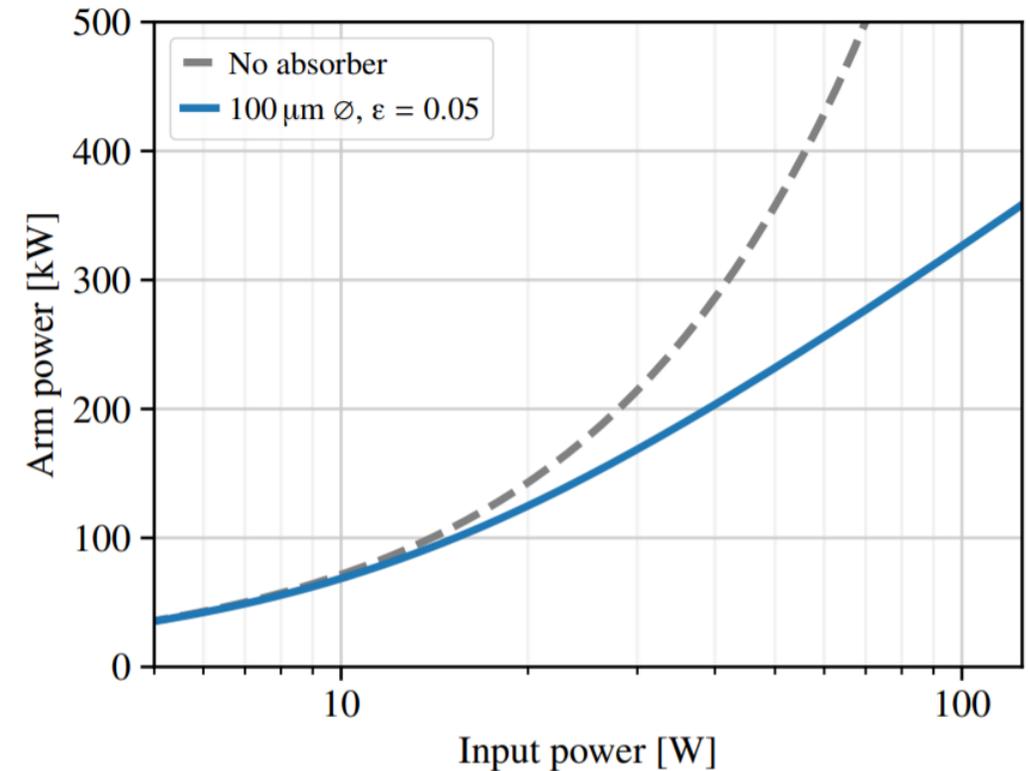
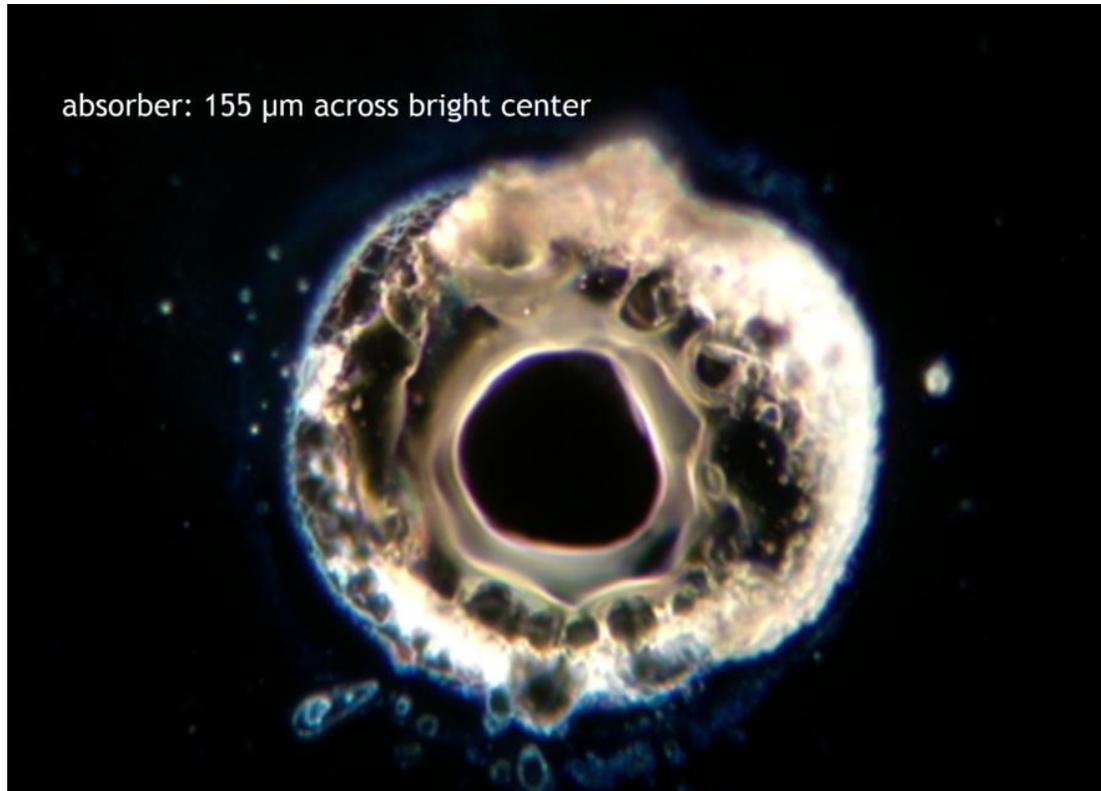
Mentors: Jon Richardson, Aidan Brooks, Rana X Adhikari

The role of optical cavities in GW detection



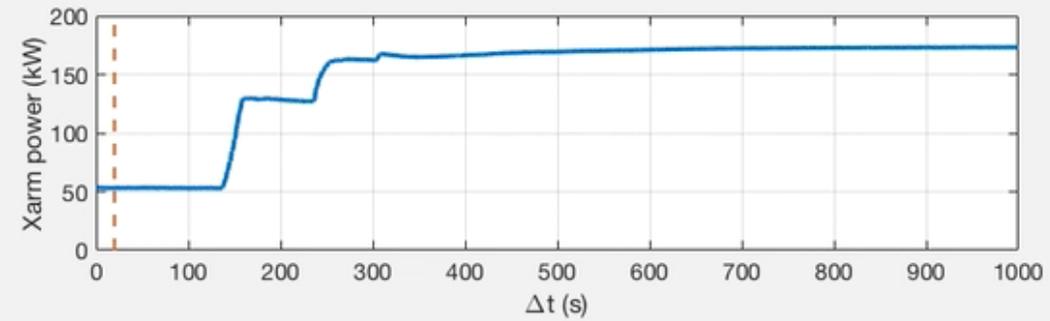
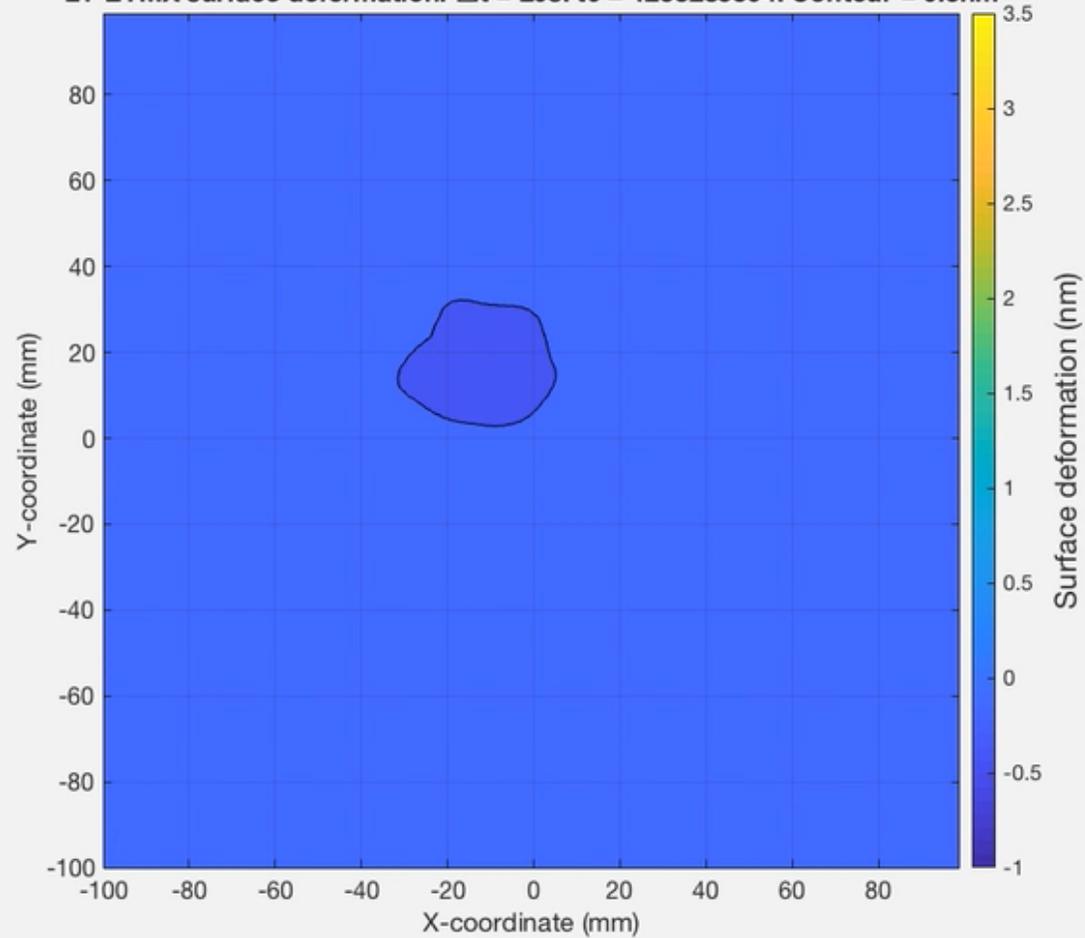
Álvarez M.D. (2019) Near-Unstable Cavities for Future Gravitational Wave Detectors. In: Optical Cavities for Optical Atomic Clocks, Atom Interferometry and Gravitational-Wave Detection. Springer Theses (Recognizing Outstanding Ph.D. Research). Springer, Cham

Point absorbers on the aLIGO test masses

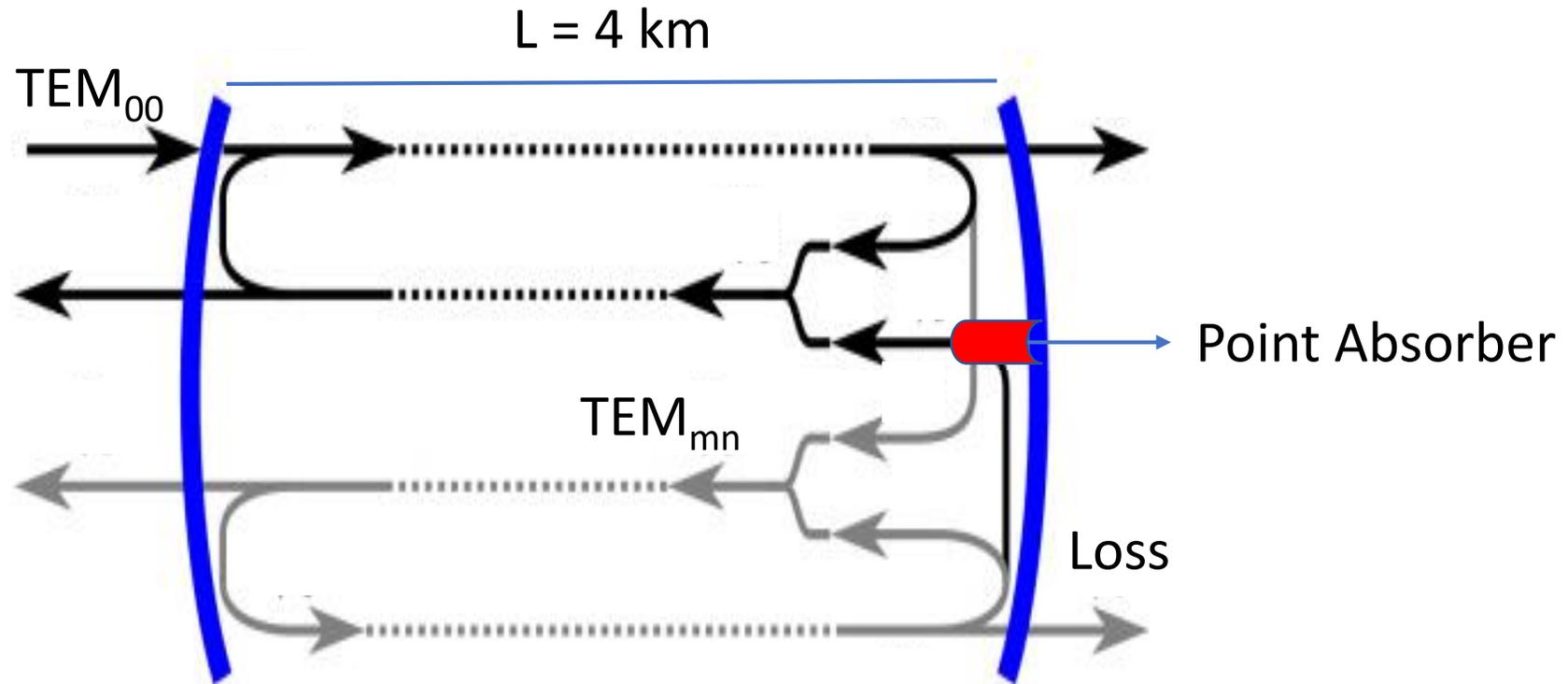


1. A. Brooks, G. Vajente et al., "The point absorbers," (2019), G1900203-v3.
2. E. D. Hall, "The effect of a point absorber in an arm cavity," (2019).

L1-ETMX surface deformation. $\Delta t = 20\text{s}$. $t_0 = 1233285394$. Contour = 0.3nm

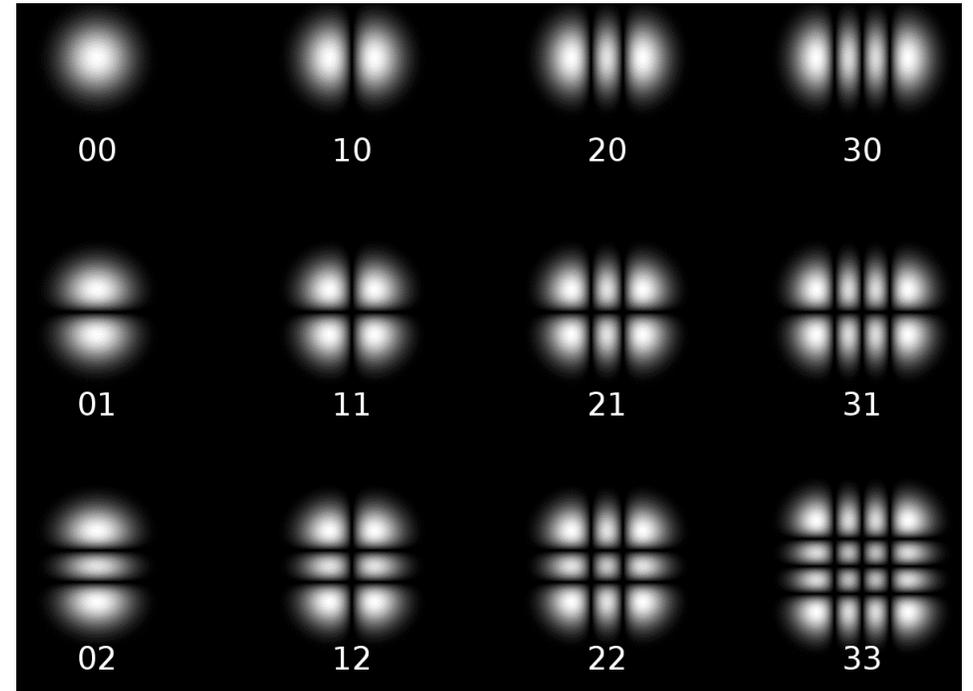
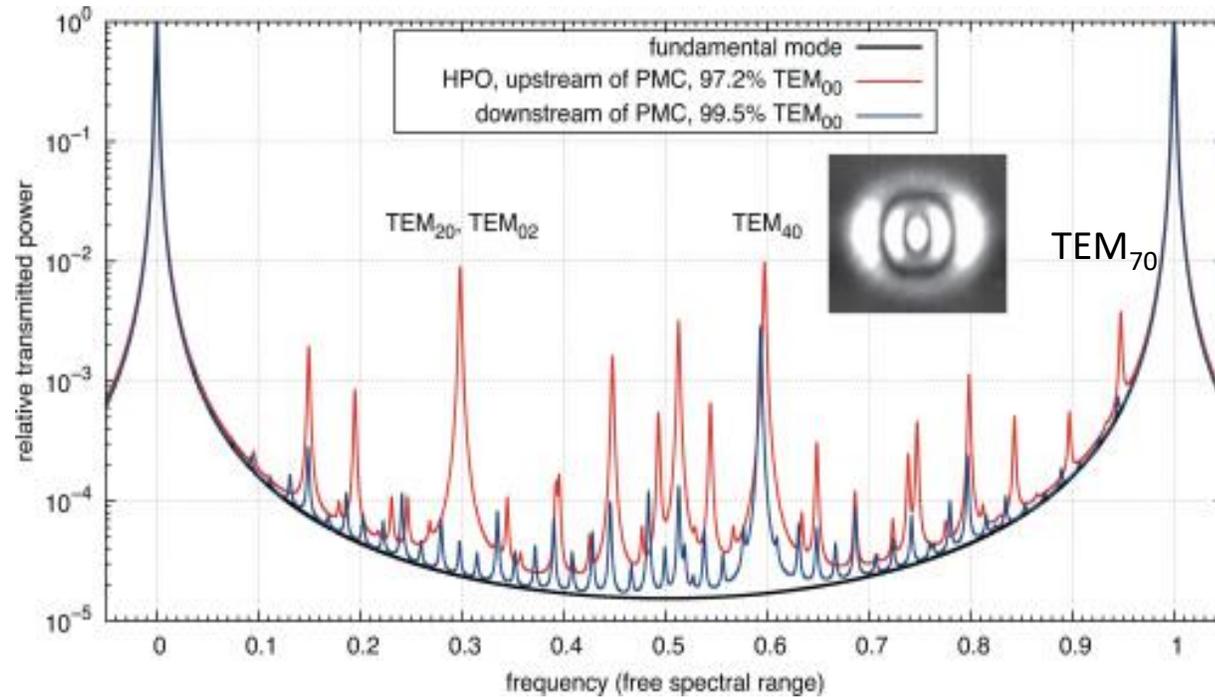


Higher order mode generation



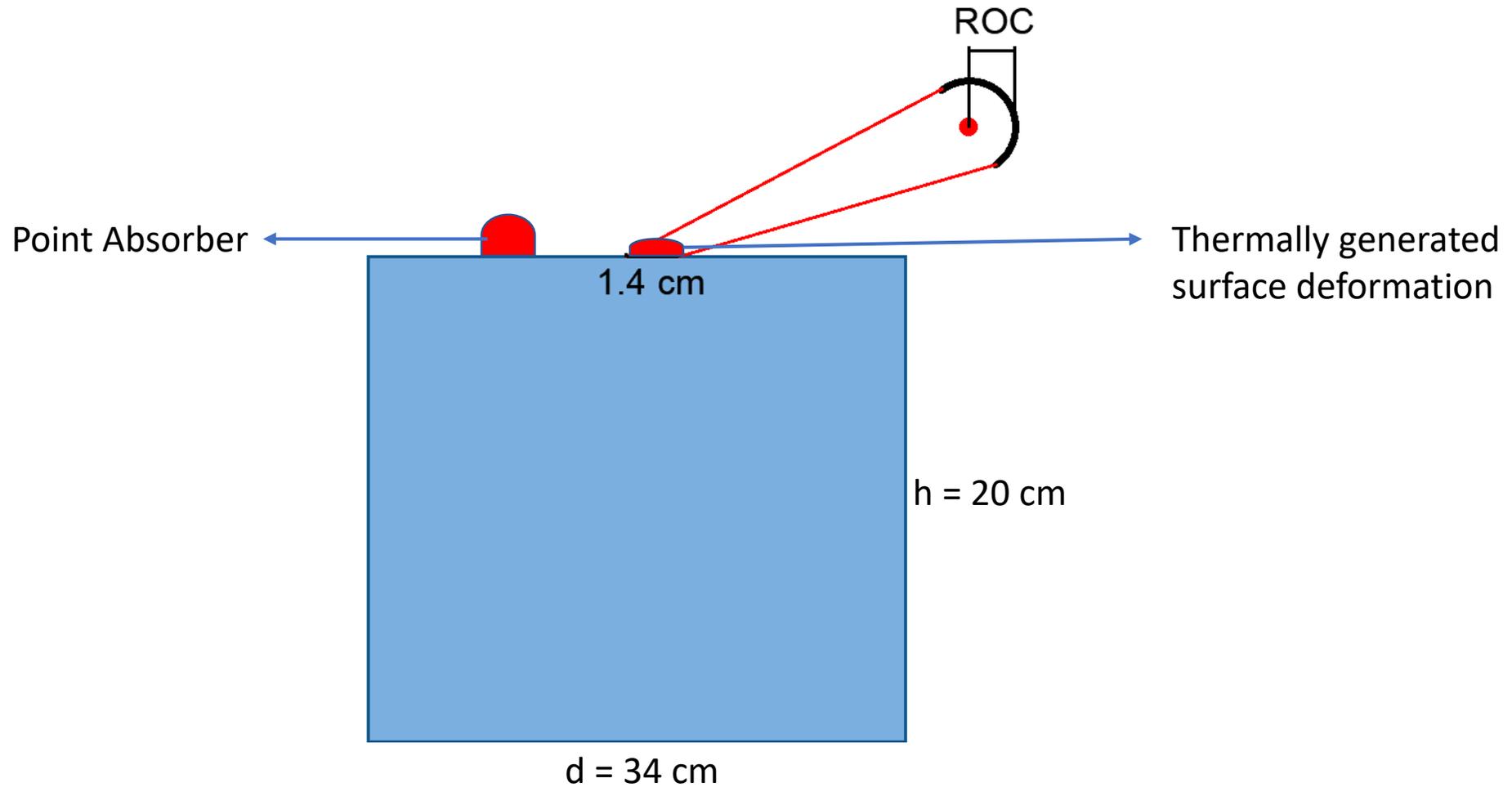
G. Vajente, "In situ correction of mirror surface to reduce round-trip losses in Fabry-Perot cavities," Appl. Opt. 53, 1459-1465 (2014) [modified]

TEM Hermite Gauss propagation modes



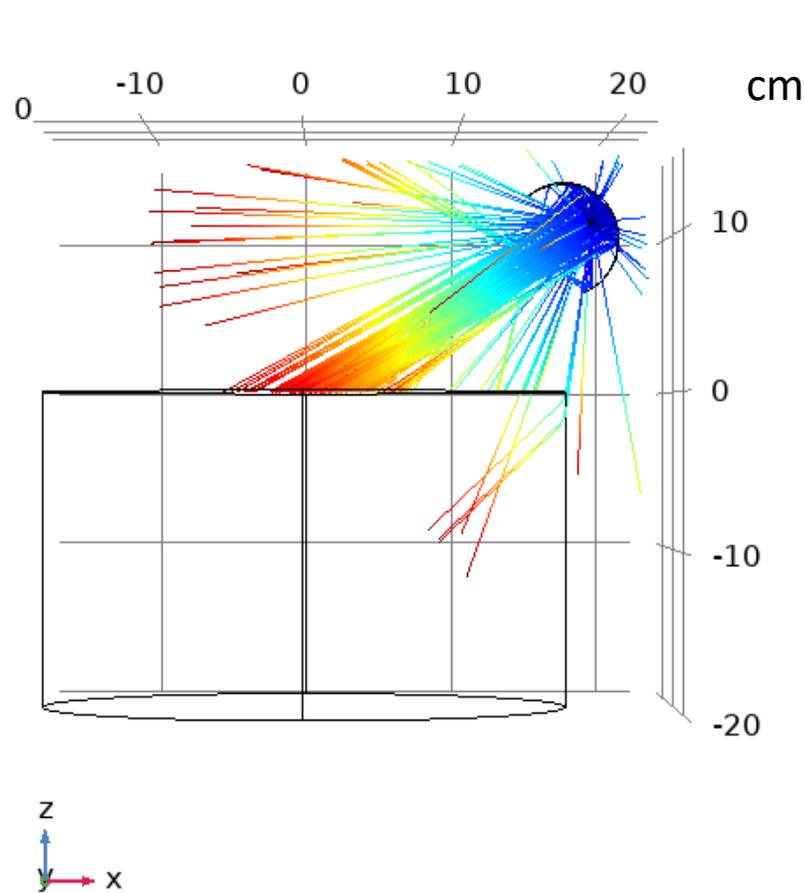
P. Kwee, C. Bogan, K. Danzmann, M. Frede, H. Kim, P. King, J. Pöld, O. Puncken, R. L. Savage, F. Seifert, P. Wessels, L. Winkelmann, and B. Willke, "Stabilized high-power laser system for the gravitational wave detector advanced LIGO," *Opt. Express* 20, 10617-10634 (2012)

Thermal compensation system design

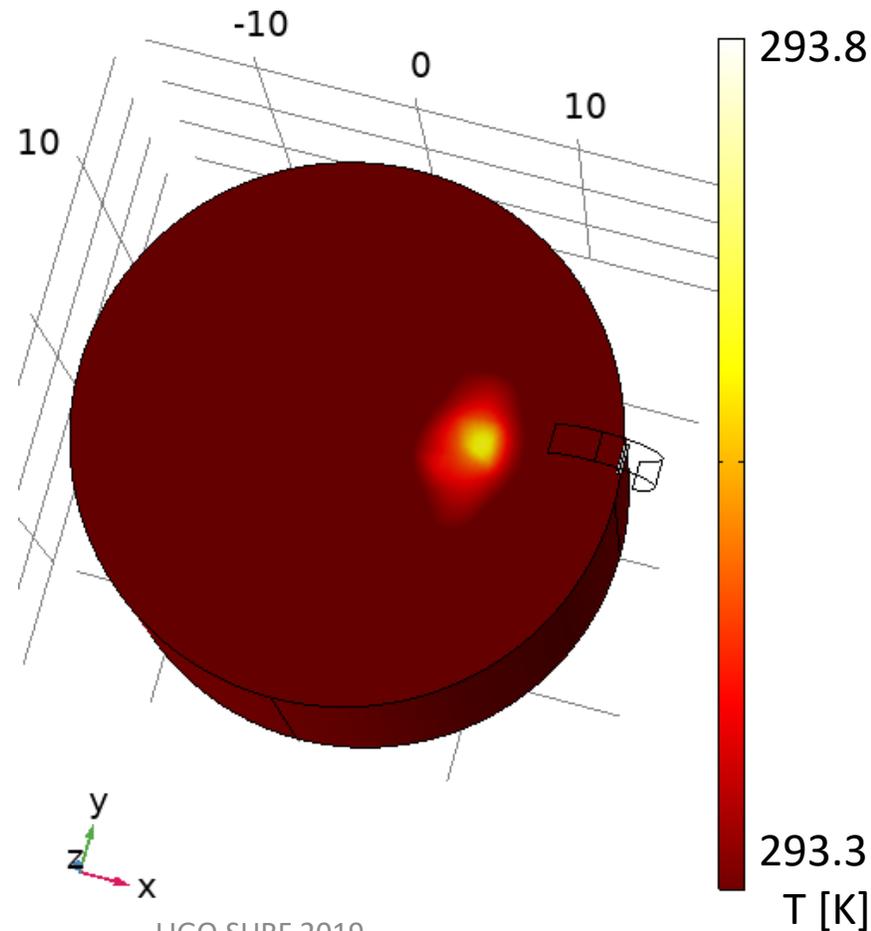


Finite-element analysis of design

Ray Trajectories



3D Temperature Profile



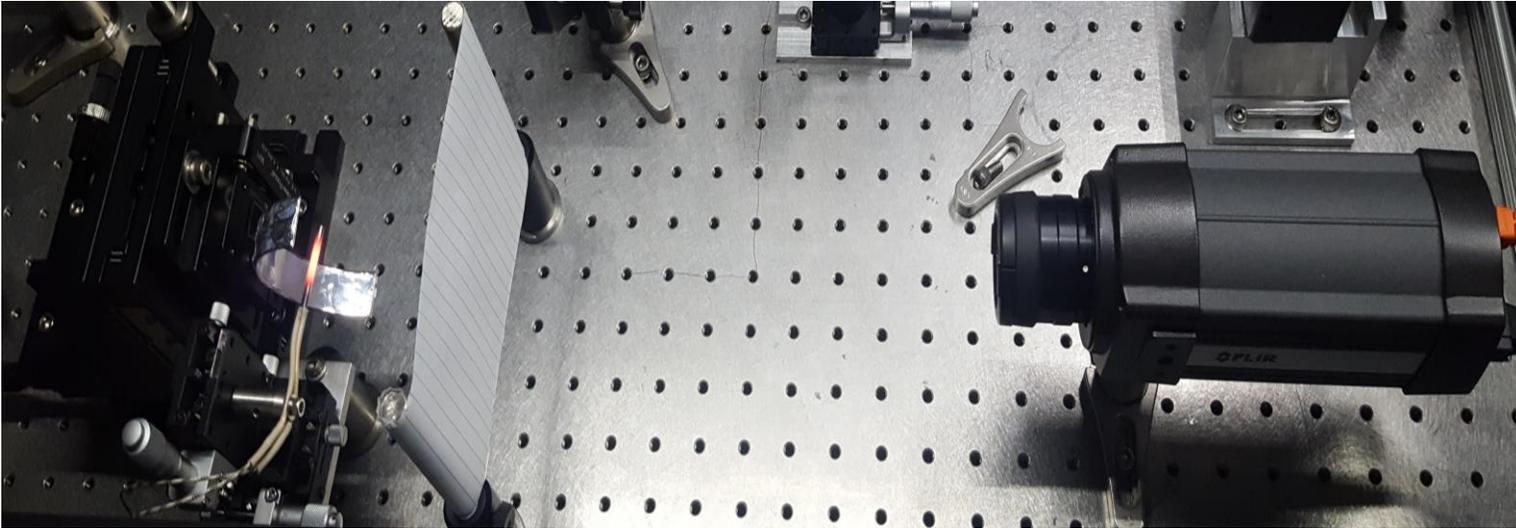
Optimized Parameters:

1. ROC
2. Reflector Geometry
3. Height (z)

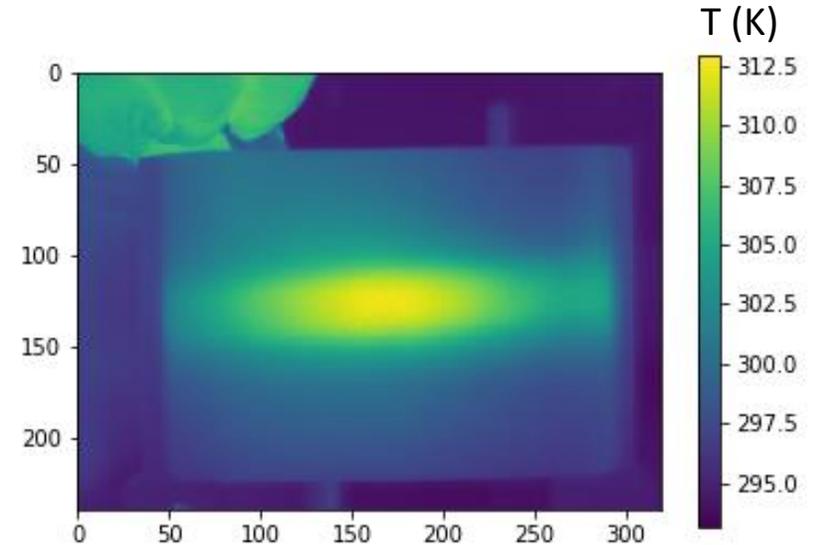
Constraints:

1. Delivery distance
2. Heater position
3. $x\text{-position} > R_{\text{test mass}}$

Measuring the heat distribution



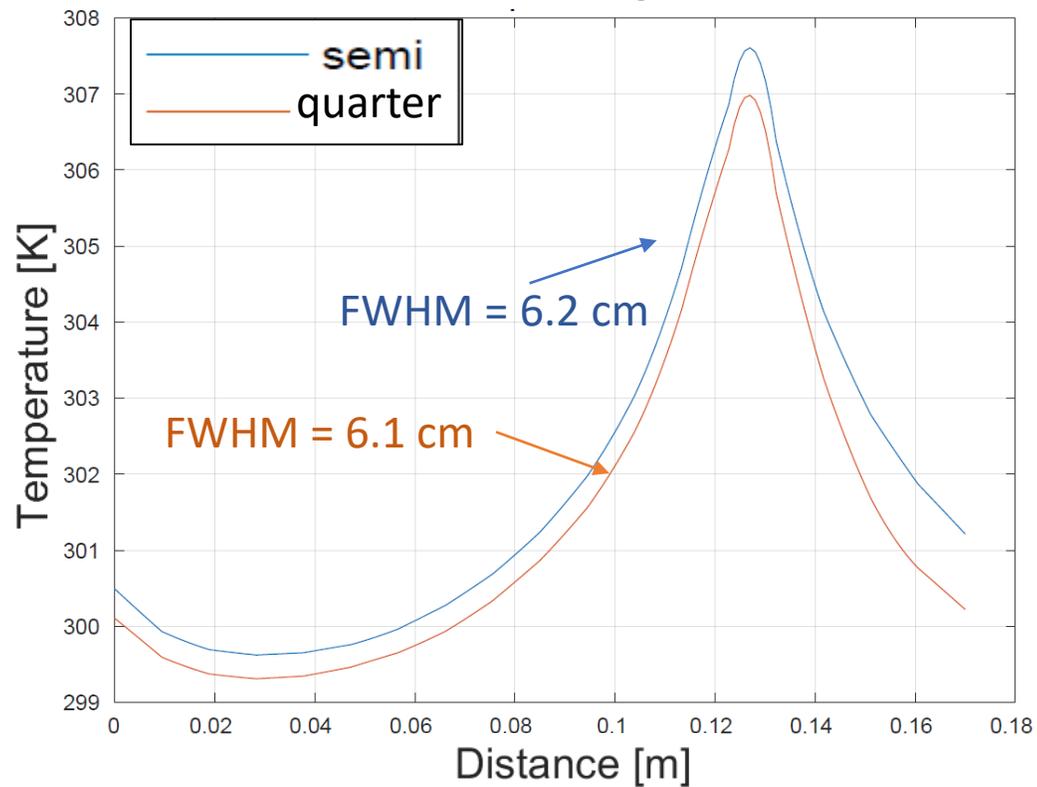
Reflector ROC (cm)	3.6	1.75
Delivery distance (cm)	10	5
Thermal Profile FWHM (cm)	2.85	1.9



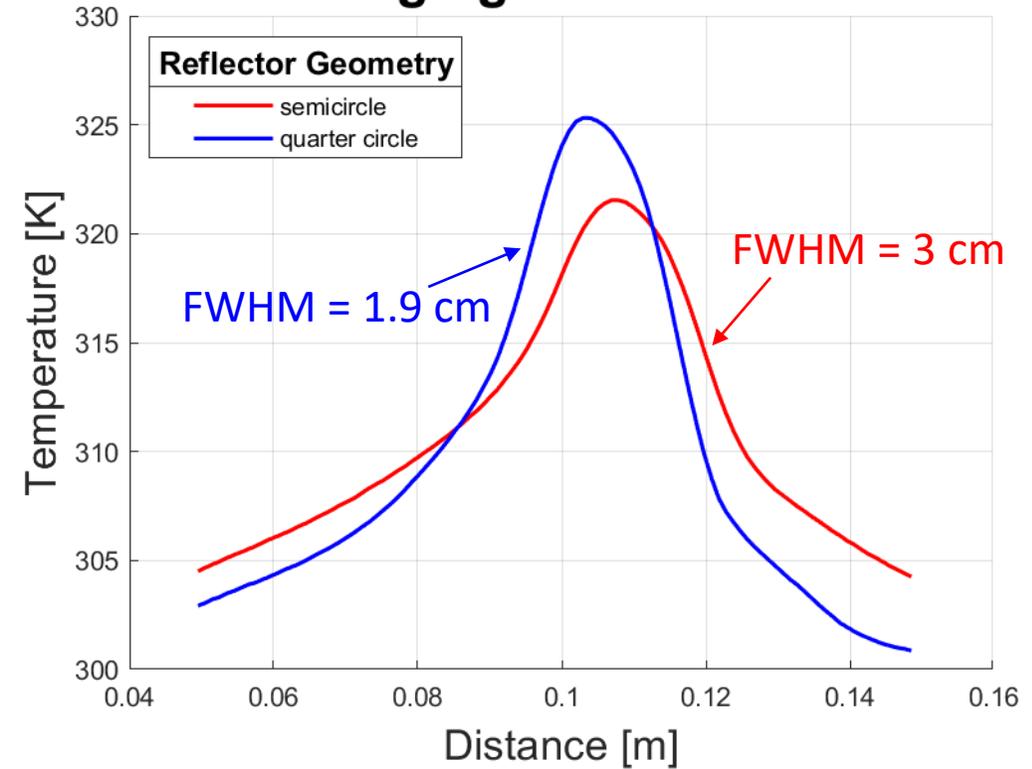
IR Image: Heat distribution pattern produced by TCS

Simulation vs Experiment Temperature Profile

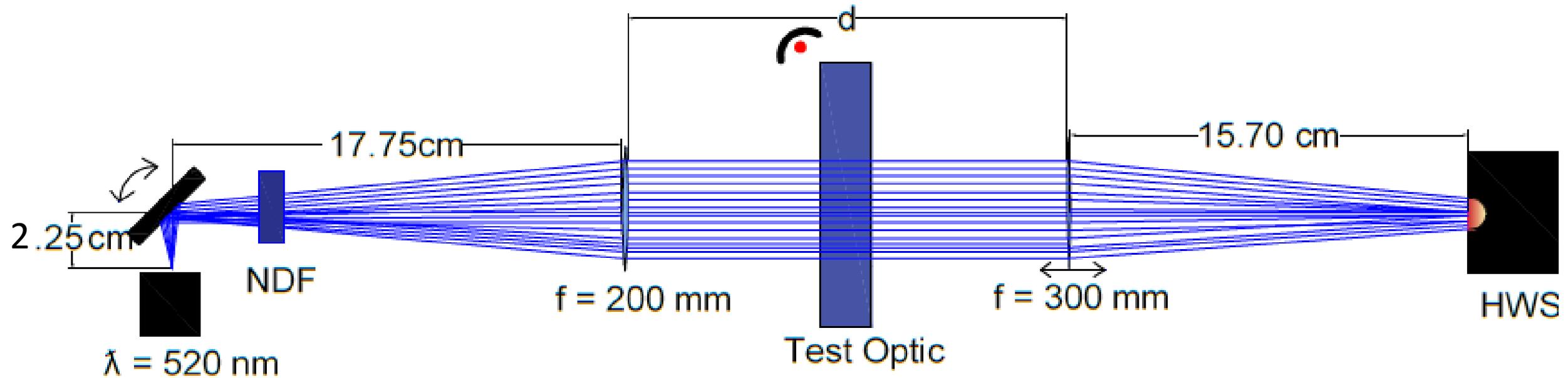
COMSOL Radial Temperature Profile



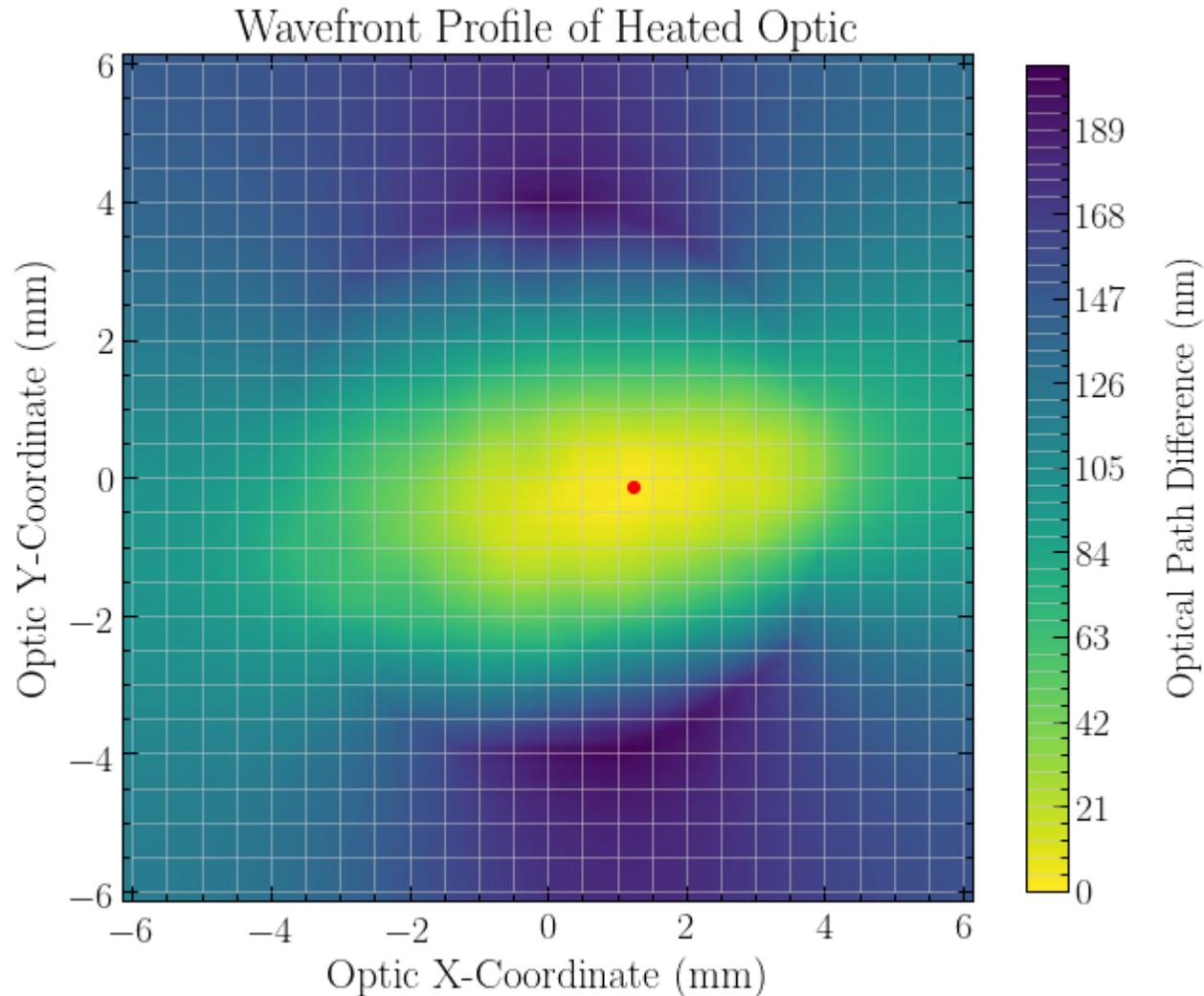
FLIR Imaging Heat Distribution



Experimental Setup



Hartmann Wavefront Sensor Measurement



Experimental Setup:

- ROC = 1.75 cm
- Radiated Power = 25 W
- Delivery distance = 6 cm

Conclusions and future work

- Obtained thermal profile can be used to actuate 7th order modes
- Further parameter optimization is necessary for a better system
- LIGO noise requirements
- Vacuum compatibility considerations

Acknowledgements

I would like to thank my mentors: Jonathan Richardson, Aidan Brooks, and Rana Adhikari for their help and guidance through all parts of the project.

Thank you for your attention!
Questions?

