

Optical Cavity Inference Techniques for Low Noise Interferometry

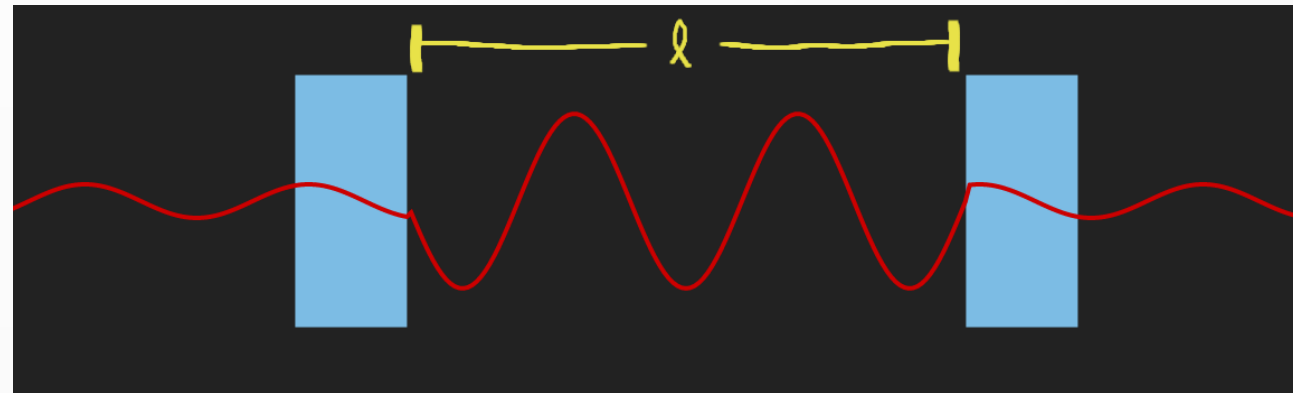
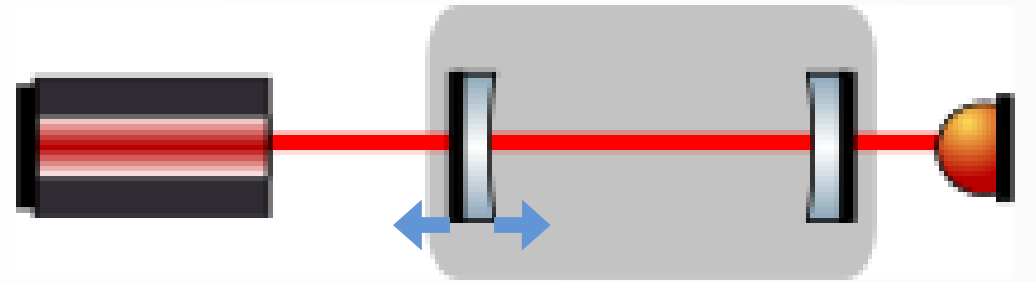
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What is a Fabry-Pérot Cavity?

A Fabry-Pérot cavity is a type of cavity that consists of two planar mirrors separated by an adjustable distance ℓ .

If the mirrors are separated by an integer number of laser wavelengths, the laser will constructively interfere inside the optical cavity and no light will be reflected.



Fabry-Perot Cavity Simulator

Fabry-Pérot Parameters

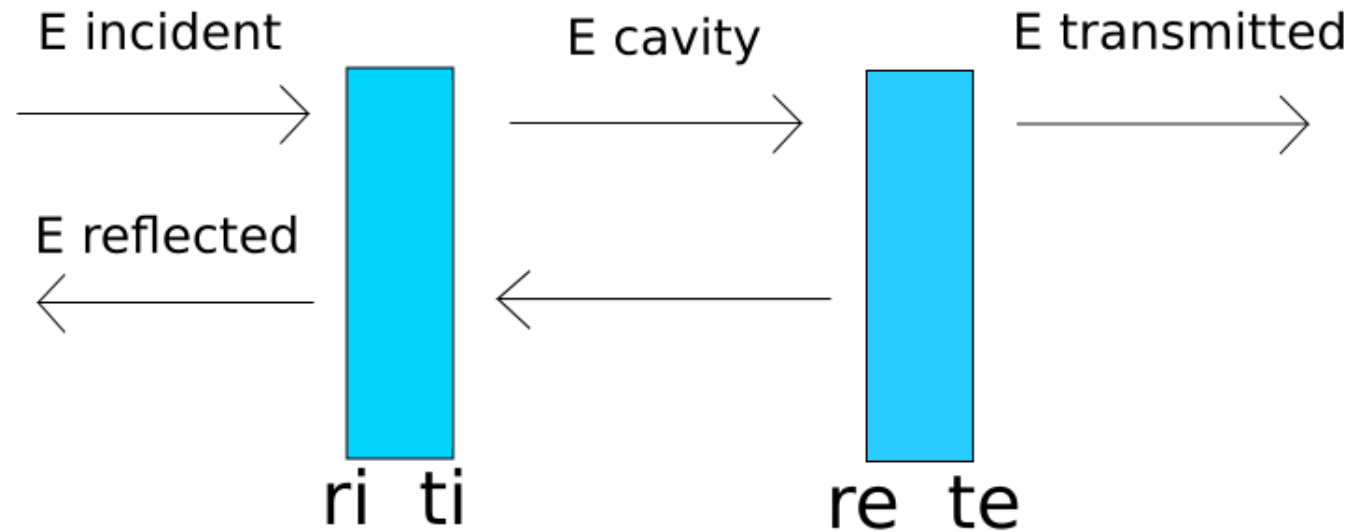
<u>Parameter</u>	<u>Sym</u>	<u>Description</u>	<u>Values</u>
Cavity Length	ℓ	Space between mirrors	~3.68 cm
Cavity Poles	f_{pole}	Frequencies at which cavity is resonant [Hz]	$f_{\text{pole}} = \frac{f_{\text{FSR}}}{2\pi} \ln\left(\frac{1}{r_i r_e}\right)$
Free Spectral Range	f_{FSR}	Distance between cavity poles [Hz]	$f_{\text{FSR}} = \frac{c}{2L}$
Finesse	\mathcal{F}	Gives an estimate of cavity sensitivity	$\mathcal{F} = \frac{\pi\sqrt{R}}{(1-R)^2}$
Reflectivity & Transmissivity	r t	Intrinsic properties of the mirror coatings and mirror material	Advanced LIGO $T_{\text{input}} = 0.000004$ $T_{\text{end}} = 0.14$

Fabry-Pérot Cavity Equations

$$E_{tran} = E_{inc} \frac{t_i t_e e^{2i\phi}}{1 - r_i r_e e^{2i\phi}}$$

$$E_{cav} = E_{inc} \frac{t_i}{1 - r_i r_e e^{2i\phi}}$$

$$E_{refl} = E_{inc} \left(r_i - \frac{t_i^2 r_e e^{2i\phi}}{1 - r_i r_e e^{2i\phi}} \right)$$



Why use a Fabry-Pérot Cavity at LIGO?

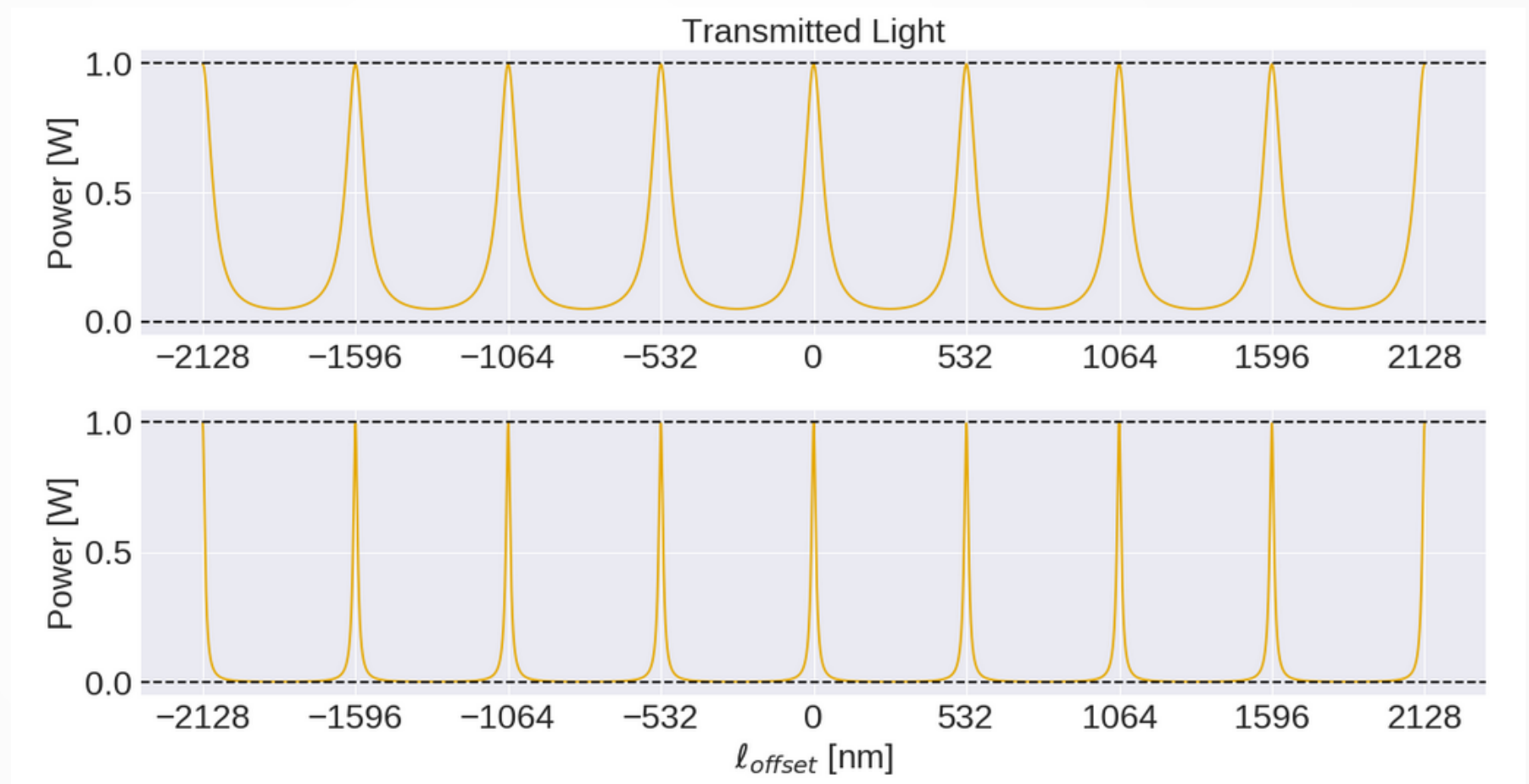
A Fabry-Pérot is extremely sensitive to microscopic shifts in length- sound familiar?

$$R = 0.8$$

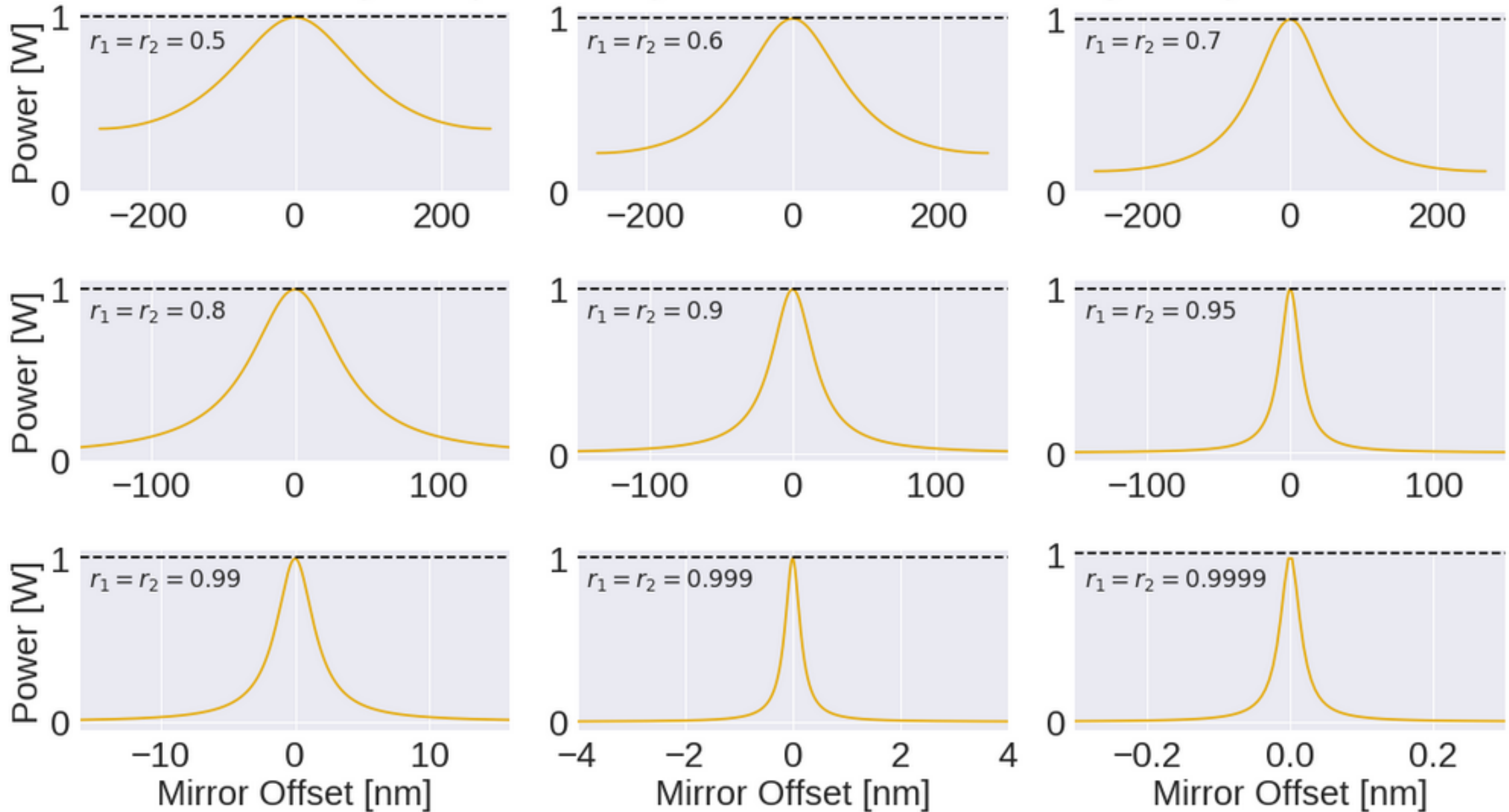
$$\mathcal{F} = 14$$

$$R = 0.95$$

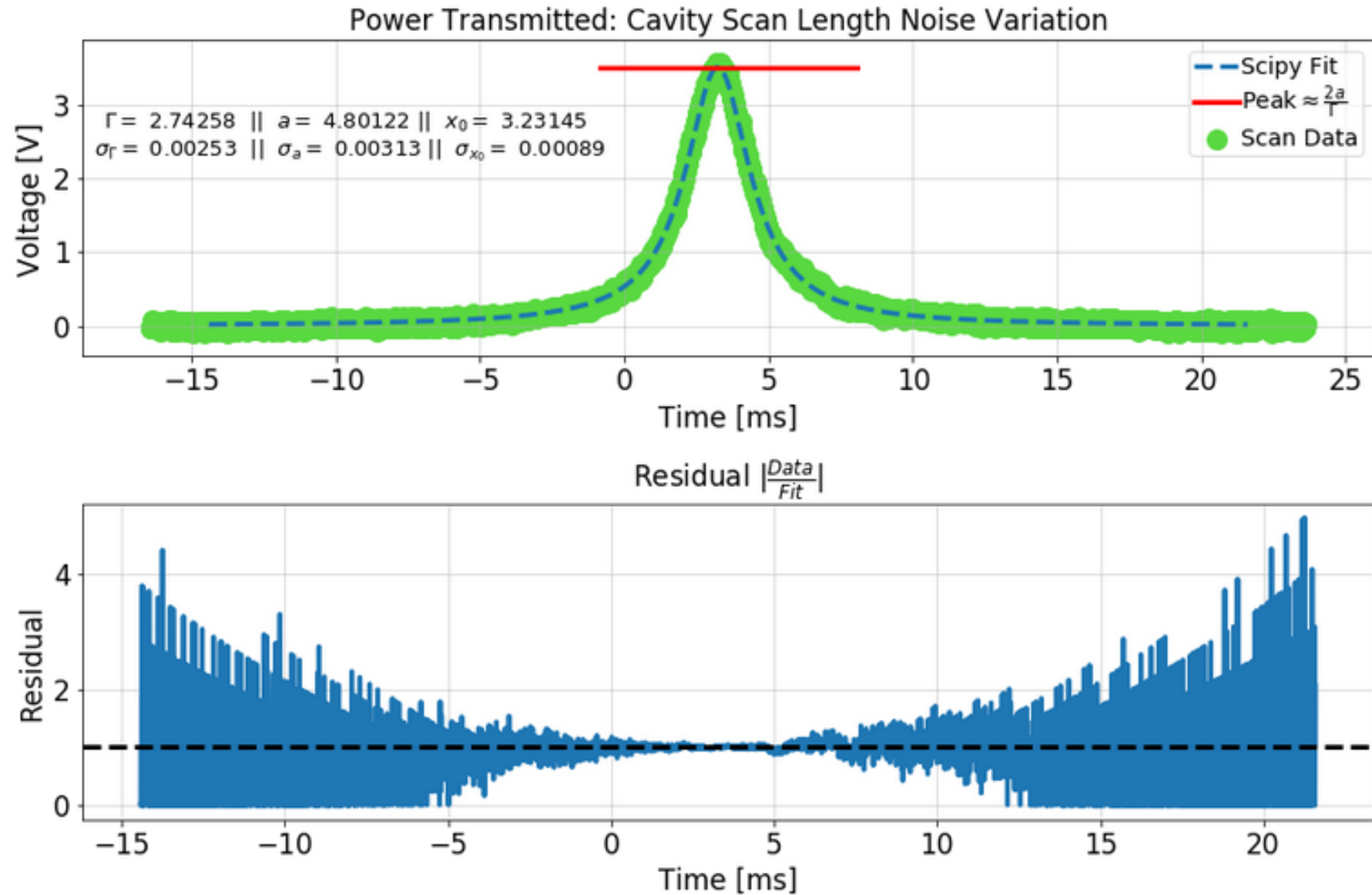
$$\mathcal{F} = 312$$



Critically Coupled Fabry-Perot Mirror Reflectivity Analysis



Real Cavity Scan Fit



What are some noise sources?

Utilizing the Fabry-Pérot in a LIGO detector for its extreme sensitivity to length changes means we must isolate the power signal from the cavity from outside influences.

Seismic Noise - vibrations in the ground supporting the instrument

Thermal Noise - thermal “wiggling” of the matter composing the mirrors

Laser Noise - “dirty” laser beams whose purported frequency can wander

Shot Noise - poissonic nature of counting photons leading to discrepancies

Length Noise - “catch-all” for unwanted Gaussian noise in movement of mirrors during a sweep

Laser Noise: Pound-Drever-Hall Technique

PDH uses the laser to regulate its own frequency by forming a negative feedback control loop

Phase modulate the main laser beam, use the newly-created sidebands as reference points for how close you are to resonance

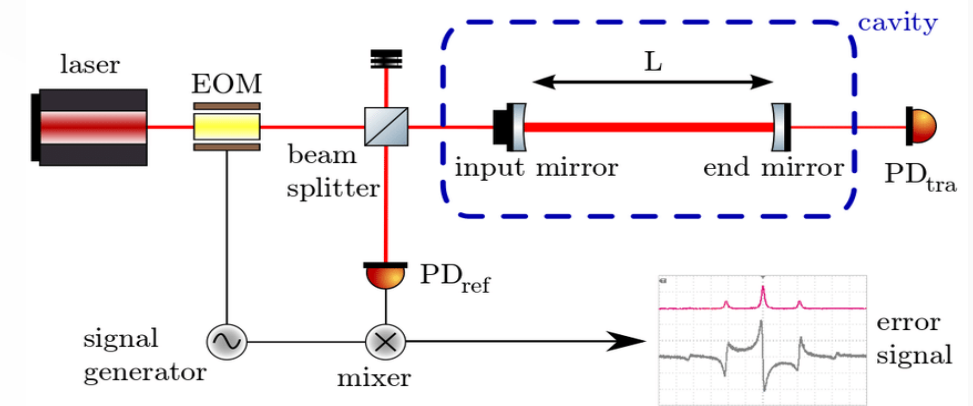
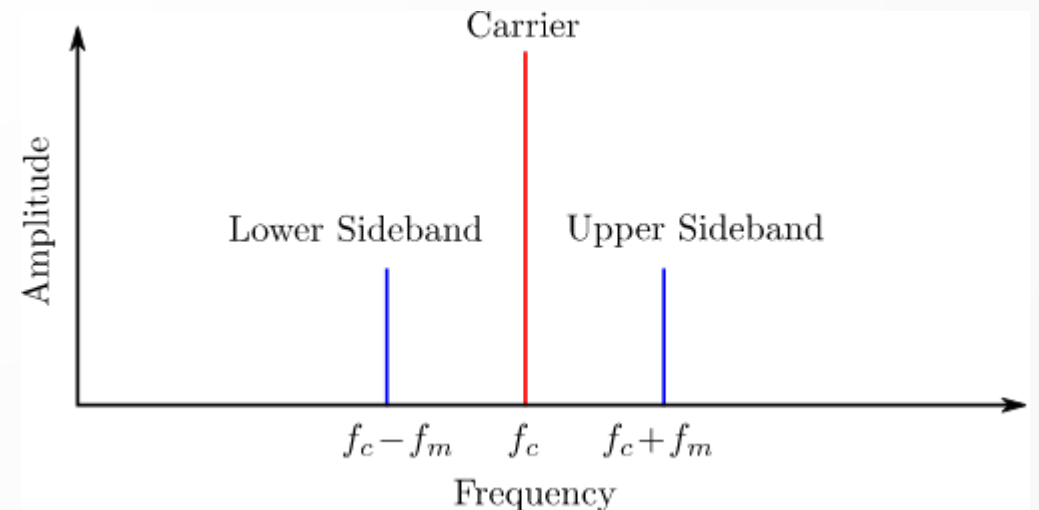
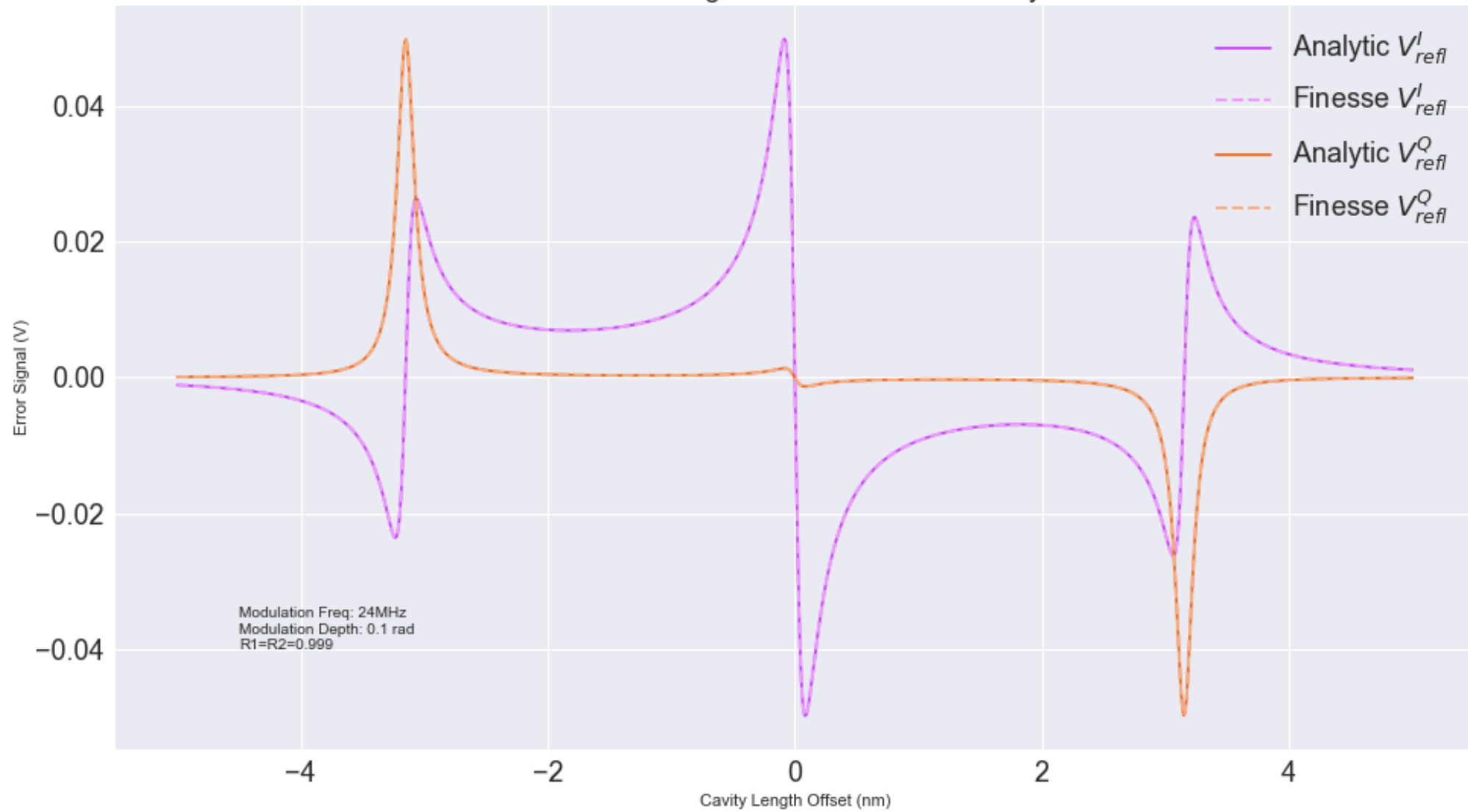


Image Source: Lock Acquisition & Commissioning of the Advanced Virgo detector



PDH Error Signal for a Two Mirror Cavity

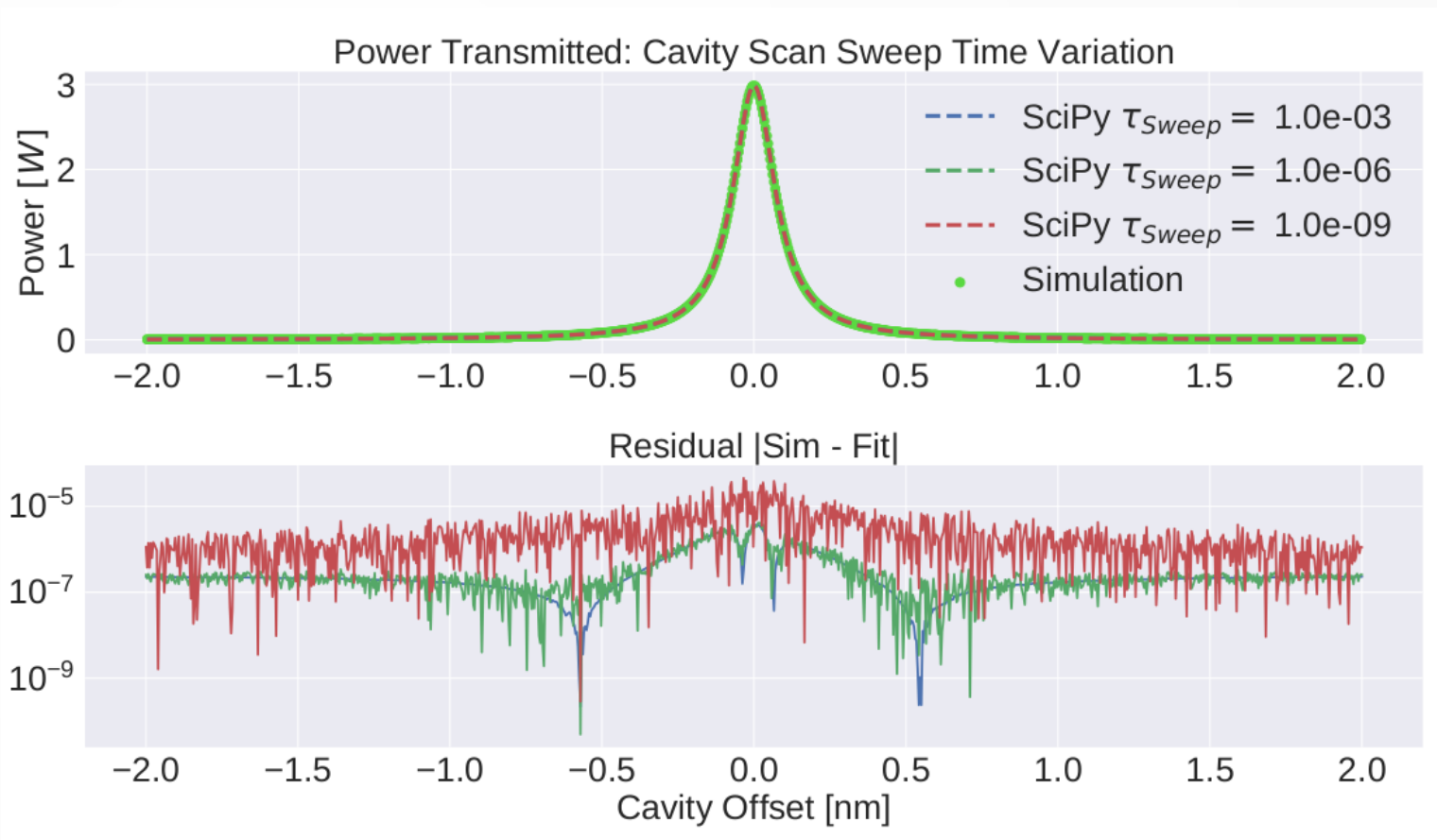


Shot Noise

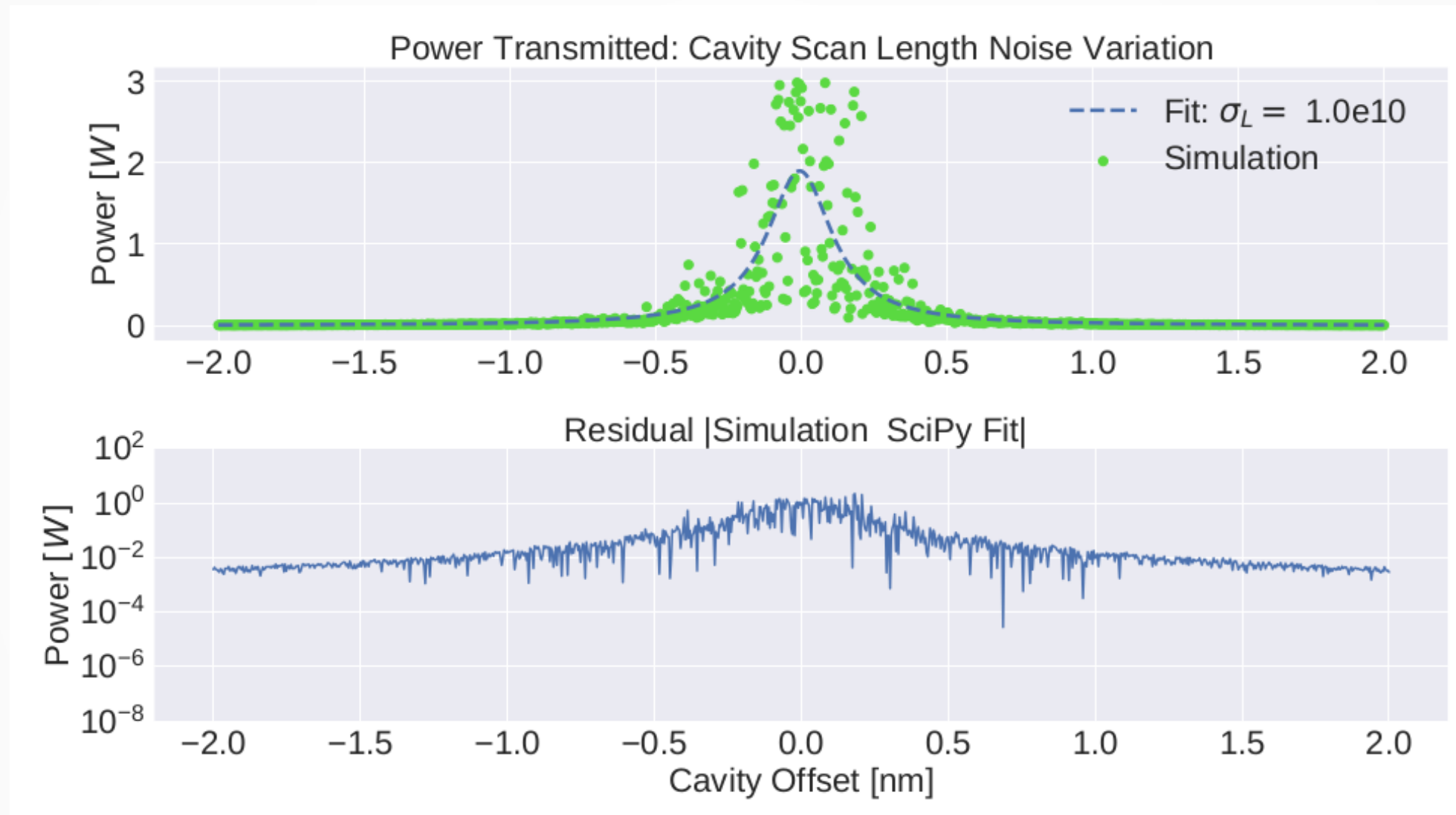
Simulating a cavity scan with a 100MHz sampling rate, three sweeps at 1ms, 1 μ s, and 1ns intervals are shown.

Slower sweep leads to more counts, but introduce their own technical problems

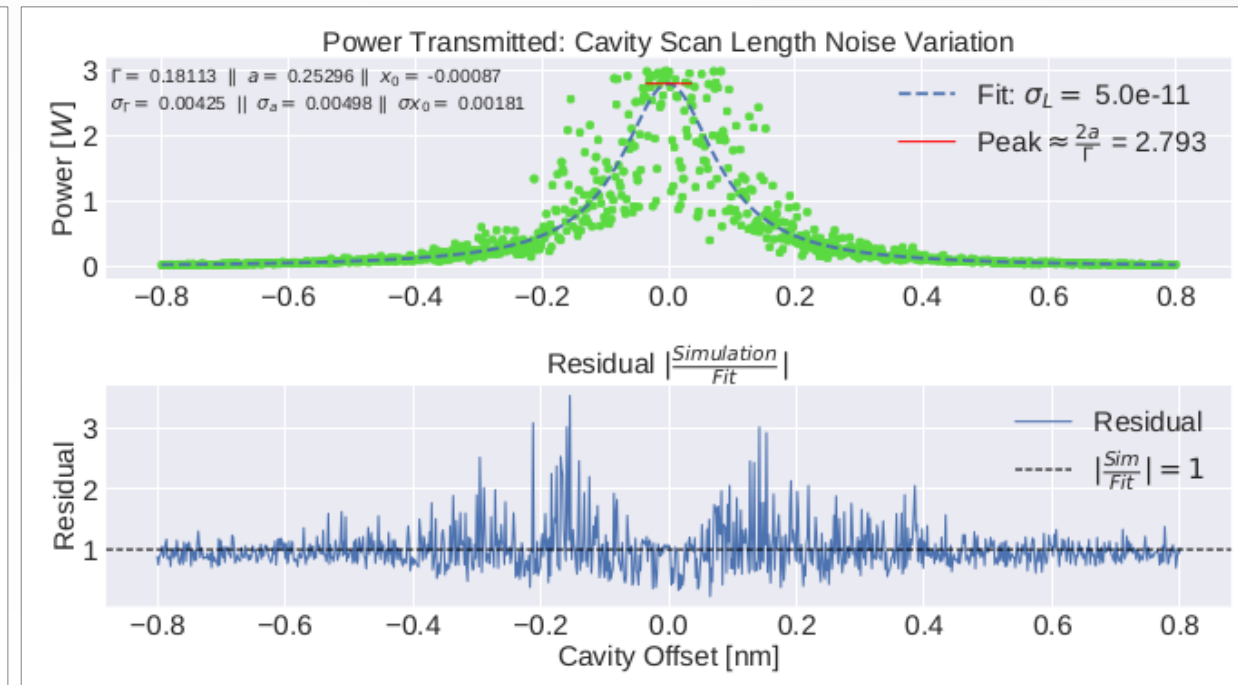
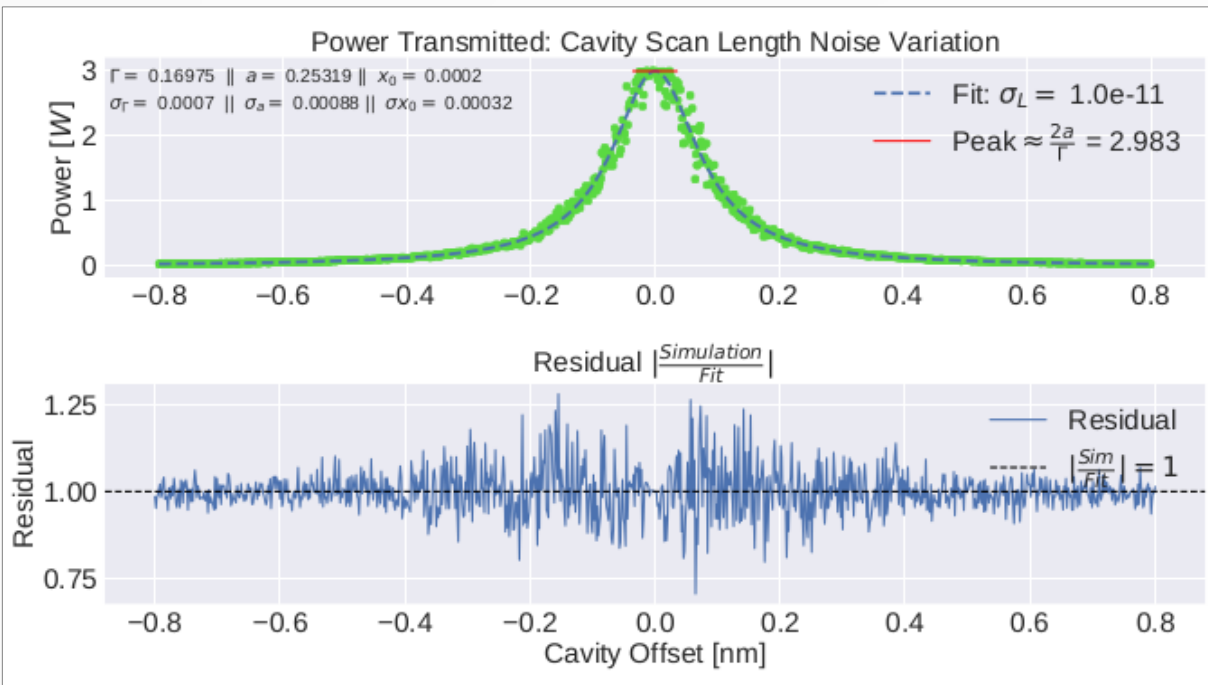
$$\sigma_{shot} = \sqrt{\hbar\omega P\tau_s}$$



Length Noise



Servo controlling ℓ isn't perfect, this creates an uncertainty in the x-value of each data point. Above is fit with no weights



Propagation of Errors for Length Noise

$$E_{tran} = \frac{[TR - T\cos(2\phi)] - [T\sin(2\phi)]i}{1 - 2R\cos(2\phi) + R^2} \text{ where } T = t_i t_e \text{ and } R = r_i r_e$$

$$P_{tran} = E_{tran}(E_{tran}^*) = \frac{T^2}{(1-R)^2 + 4R\phi^2} \text{ using } \cos(2\phi) = 1 - 2\phi^2$$

$$\sigma_P = \sigma_L \frac{\partial \phi}{\partial L} \frac{\partial P}{\partial \phi}$$

$$\sigma_P = \sigma_L \frac{2\omega}{c} \frac{-8RT^2\phi}{((1-R)^2 + 4R\phi^2)^2}$$

Lesson learned:
 when fitting,
 assigning and
 weighing errors is
 critical!

```
-----  
Fit Parameters (gamma, a, x_0) for noise 8e-11 :
```

```
[ 0.192  0.26  -0.005]
```

```
St. Dev:
```

```
[0.008, 0.009, 0.003]
```

```
gamma: 2.921 deviations || a: 0.798 deviations || x0: -1.43 deviations
```

```
-----  
Fit Parameters (gamma, a, x_0) for noise 1e-10 :
```

```
[ 0.202  0.266 -0.006]
```

```
St. Dev:
```

```
[0.011, 0.012, 0.004]
```

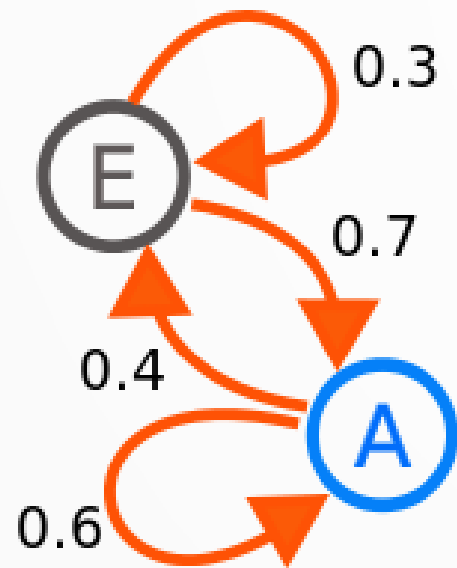
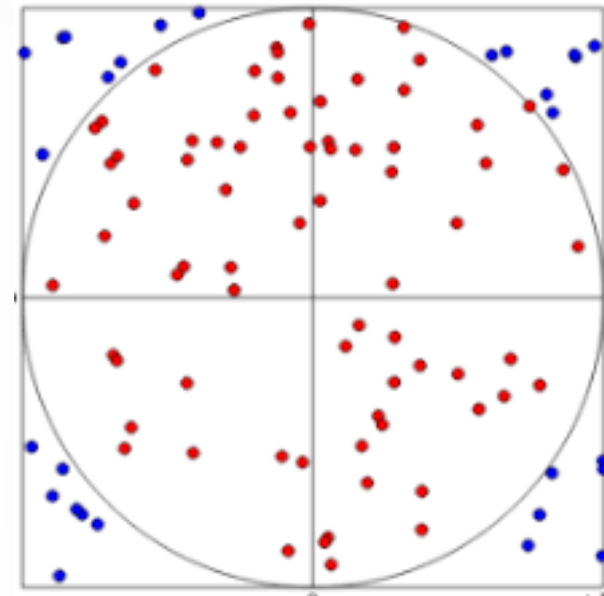
```
gamma: 2.976 deviations || a: 1.095 deviations || x0: -1.452 deviations
```

$$\mathcal{L} = a * \frac{\frac{1}{2} \Gamma}{(x - x_0)^2 + \left(\frac{1}{2} \Gamma\right)^2}$$

Still, even after carefully propagating error and including it in the least squares fitting algorithm, estimates for cavity parameters weren't good enough!

Markov Chain Monte Carlo (MCMC)

Monte Carlo Method - utilizing random sampling to generate sets of non-random data by relying on the *Law of Large Numbers* (for a large amount of trials, results will tend towards the mean)



Markov Chain - a chain of events where the probability of future events is uncorrelated with past events

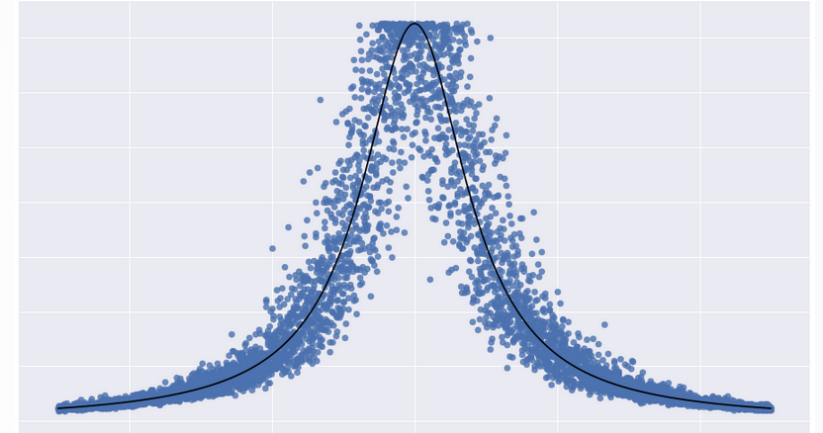
(1) Define a model and generate noisy data

$$E_{tran} = \frac{-t_i t_e e^{2i\phi}}{1 - r_i r_e e^{2i\phi}}$$

$$P_{tran} = E_{tran}(E_{tran}^*)$$

$$T = t_i t_e$$

$$R = r_i r_e$$



(2) Define a parameter space with boundaries which will form the basis for the Markov chain

$$\theta_0 = \begin{cases} p_1 \\ p_2 \\ \dots \\ p_n \end{cases} = \begin{cases} T \\ R \\ P \\ \lambda \\ \gamma \end{cases} \in \begin{cases} 0 < T < 1 \\ 0 < R < 1 \\ 2.5 < P < 3.5 \\ 1060 < \lambda < 1070 \\ 0.18 < \gamma < 0.22 \end{cases}$$

$$T = t_i t_e$$
$$R = r_i r_e$$

(3) Define a likelihood that nudges each “walker” in the markov-chain in the correct direction

$$\text{Likelihood} = -\frac{1}{2} \sum (\text{data} - \text{model}(\theta_0))^2$$

(4) Set MCMC parameters, set initial guesses, and define the size of each step:

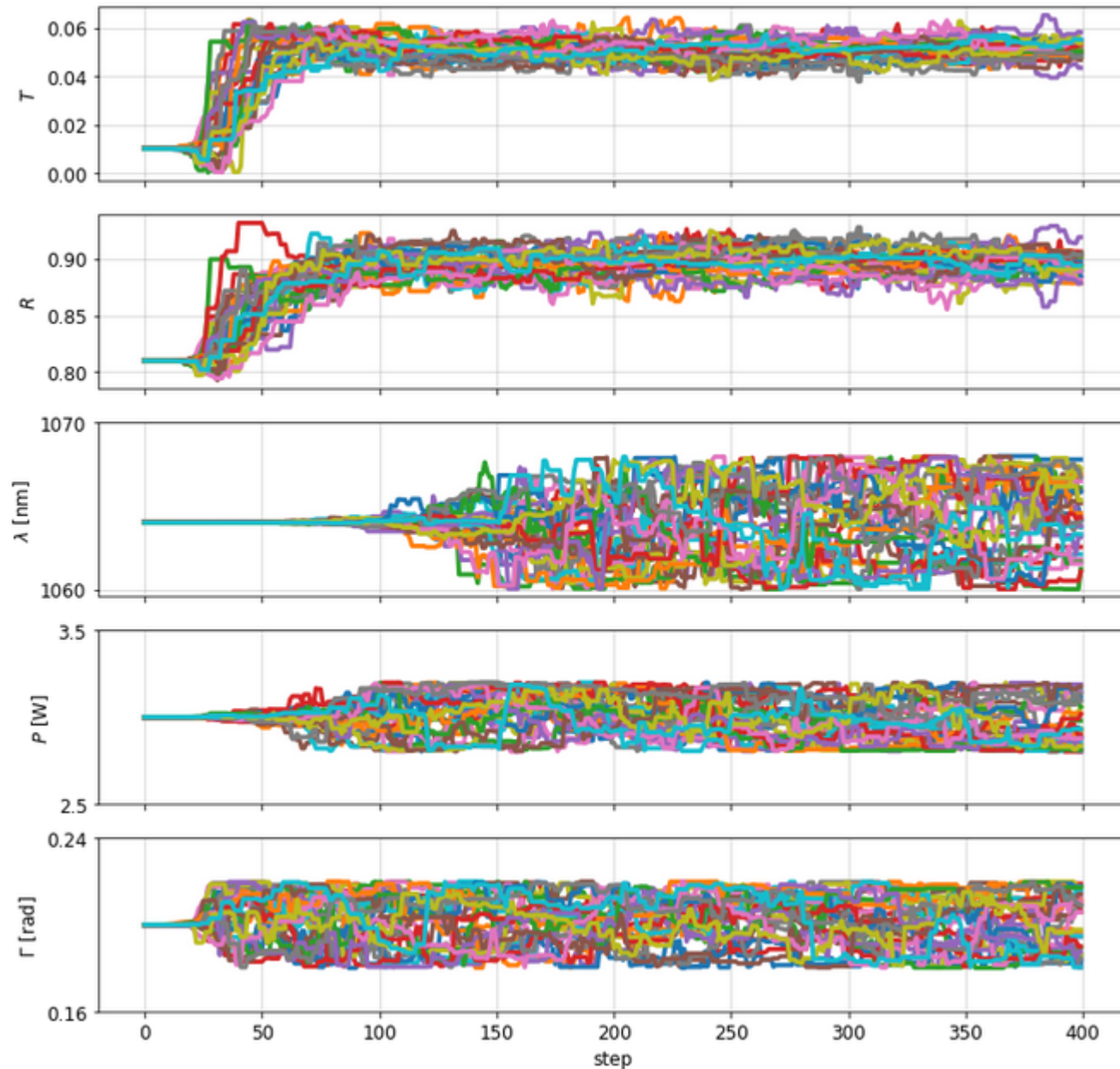
```
#--- MCMC settings
ndim      = 5 # n
nwalkers  = 1000
nsteps    = 600 #
nburn     = 100 #
nthreads  = 10 #

#--- tell each walker where to start
# theta =[T, R, laser wavelength, laser_power, mod_depth]
Tguess, Rguess, LWguess, LPguess, MDguess = (0.1**2), (0.9**2), 1064, 3, 0.2
initial = np.array([Tguess, Rguess, LWguess, LPguess, MDguess])

p0 = [np.array(initial) + 1e-6 * np.random.randn(ndim) for i in range(nwalkers)]
```

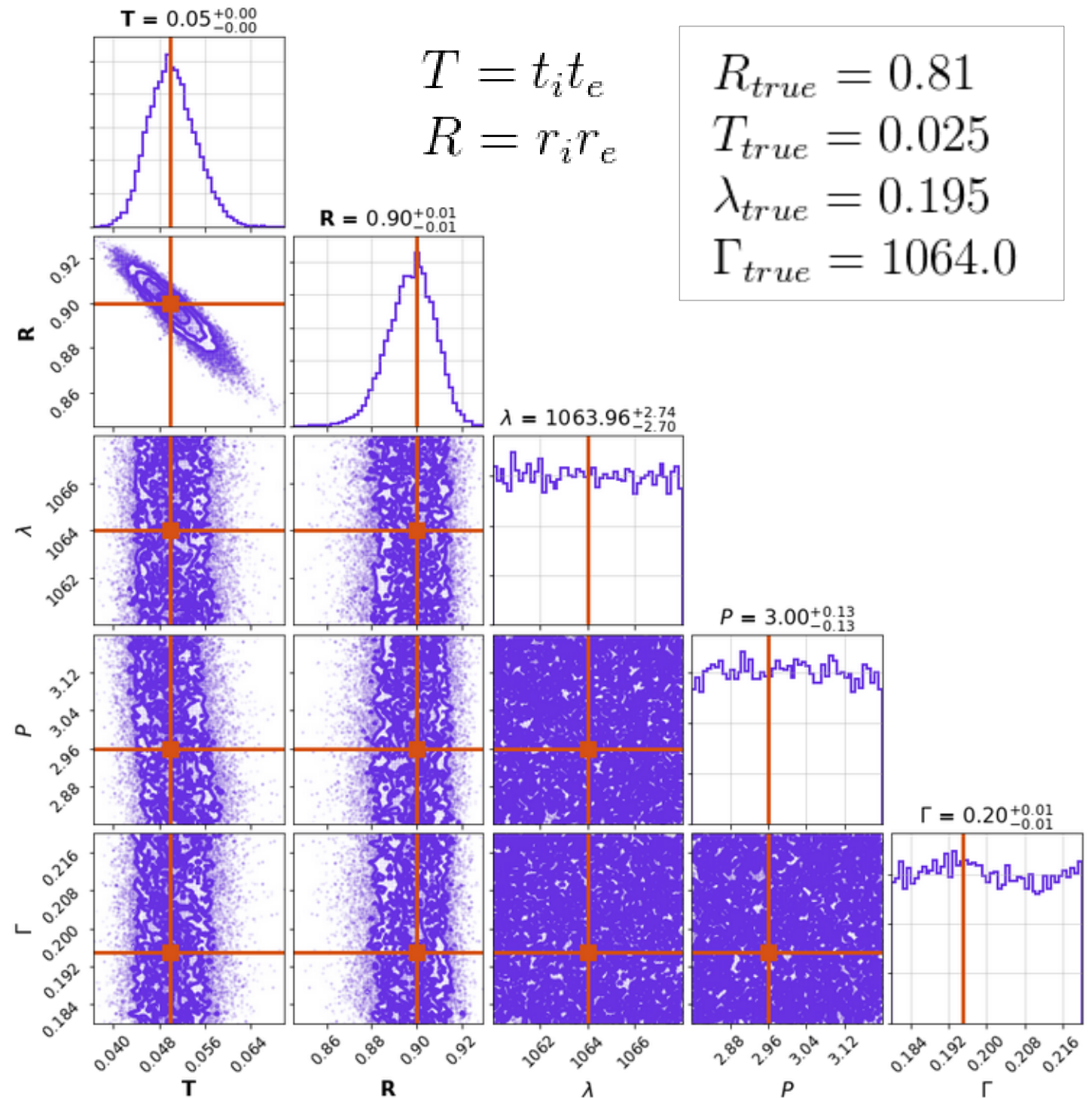
(5) Let them loose! Walkers will start at their initial guess, change it by a little bit every step, and then move towards the parameters with highest likelihood (and turning around if they ever step out of bounds)

Results



- T and R converge relatively quickly and nicely
- Laser Wavelength has a lot of variations, but most cluster around 1064
- Laser Power and Modulation Depth do not converge

- T vs R scatter plot shows a negative covariance
- Not obvious from previous plot, but wavelength converged around 1063.9
- Laser Power and Modulation Depth do not converge



Future Work

- Improving the MCMC algorithm by introducing a prior distribution that assigns probabilities to parameters
- Support for adjusting which parameters of the cavity we're trying to estimate
- Adding estimated uncertainties to likelihood function to improve the efficiency each step

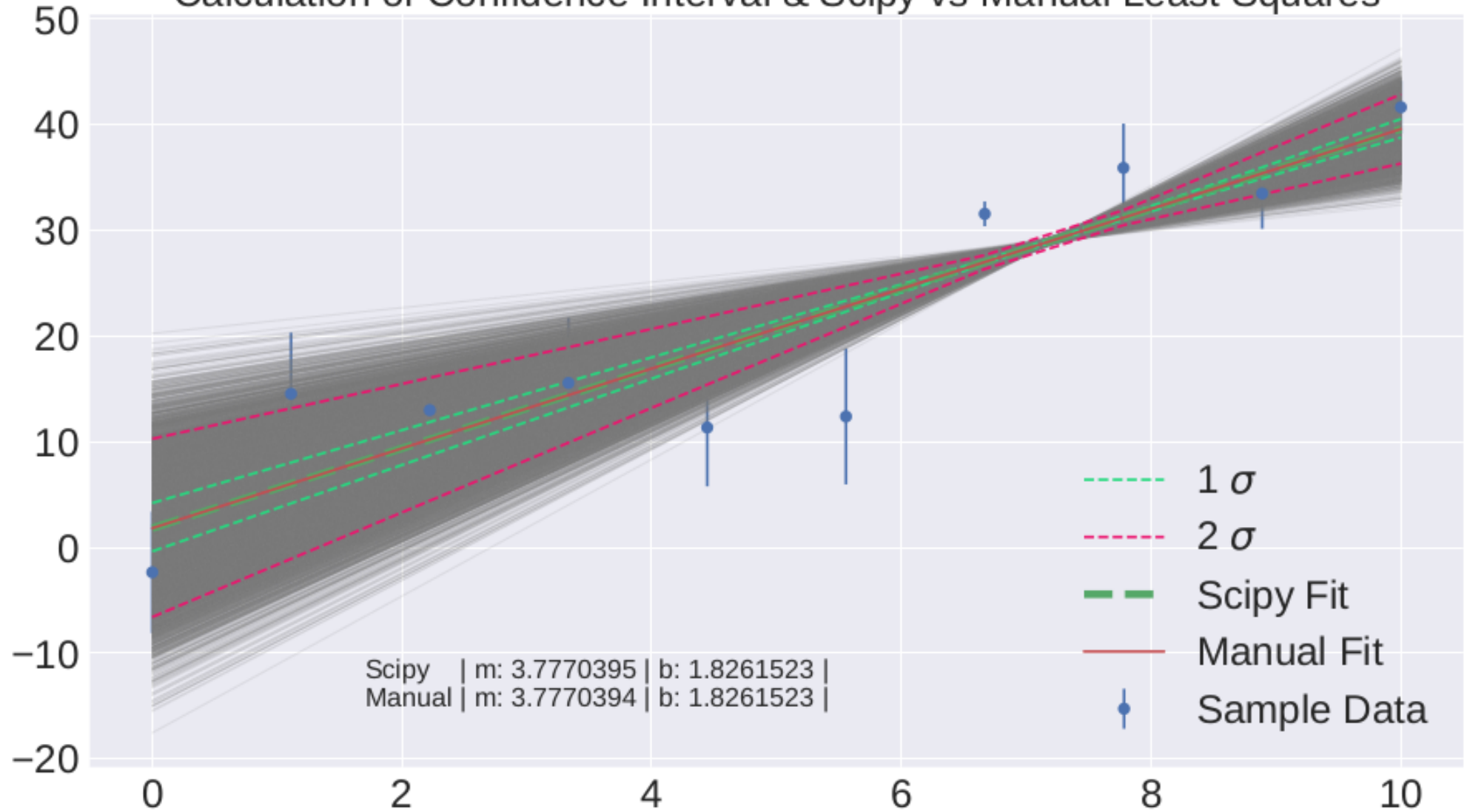
- I'd like to thank Rana for his mentorship and guidance, as well as helping me realize that science isn't only about getting the right answer, but asking the right questions
- Craig for all his advice, lessons, and for supporting me whenever I felt like I was in way over my head with this project.

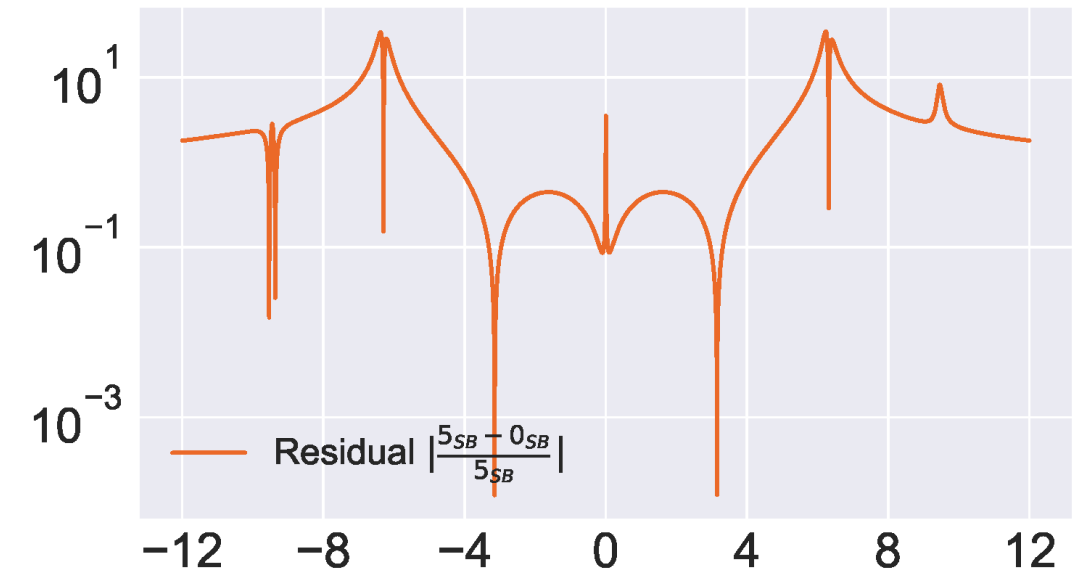
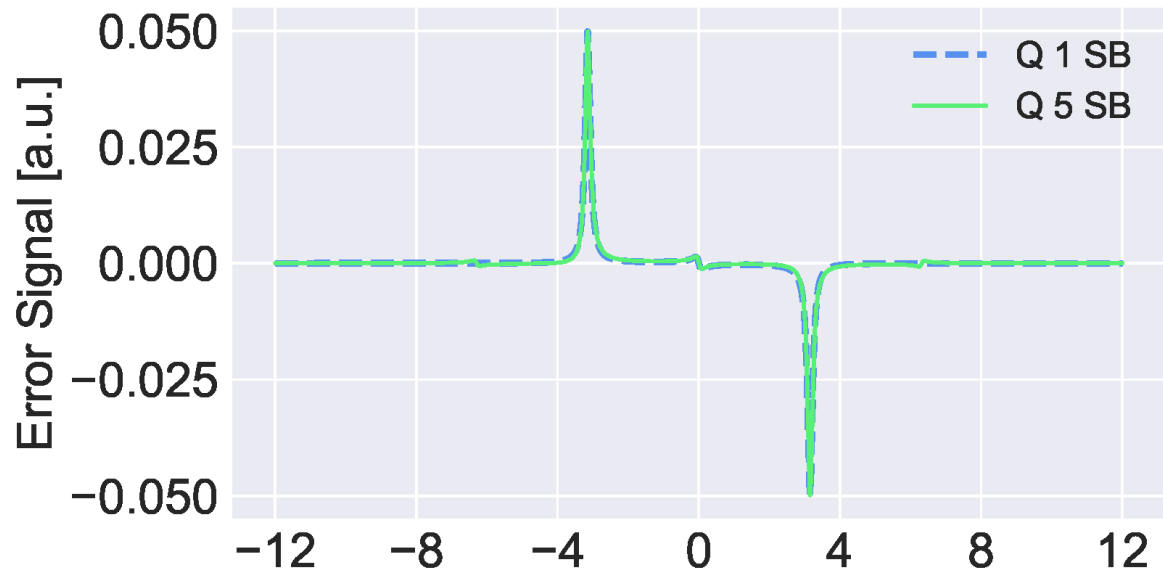
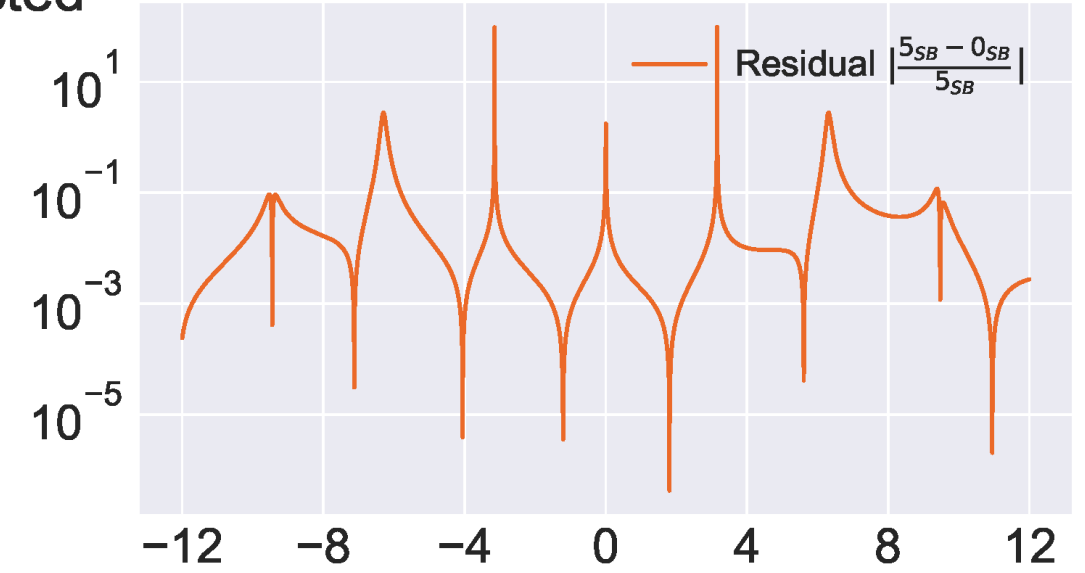
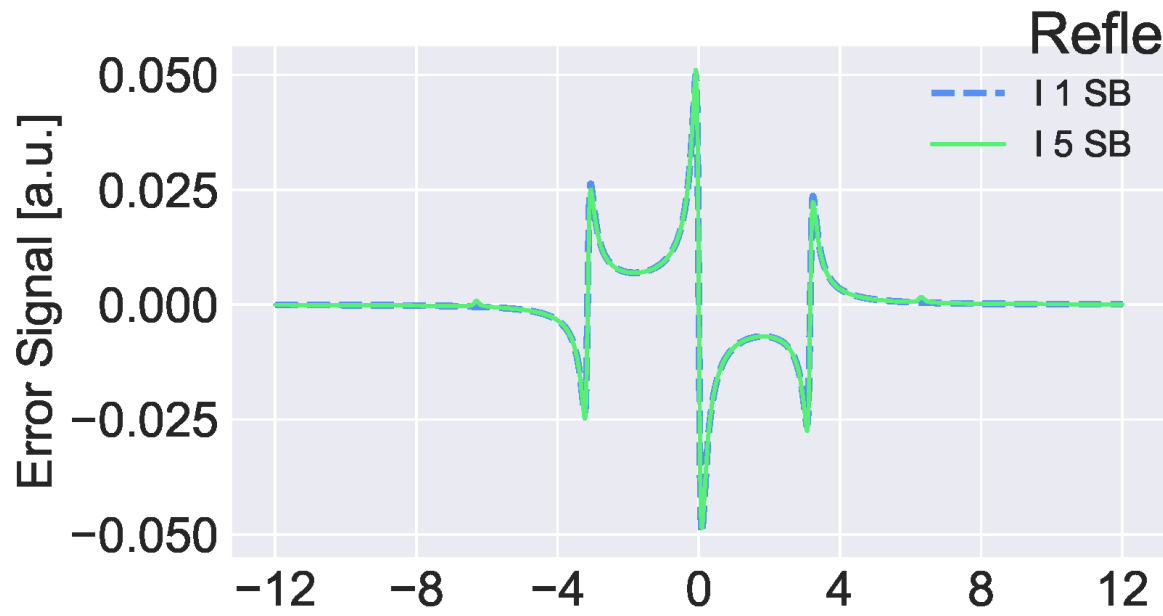
Thank you!



Extra Slides

Calculation of Confidence Interval & Scipy vs Manual Least Squares





residual min: 0.9975015620659504
residual max: 0.9986615326479165

