

Update on the Frequency Noise Budgeting and Controls

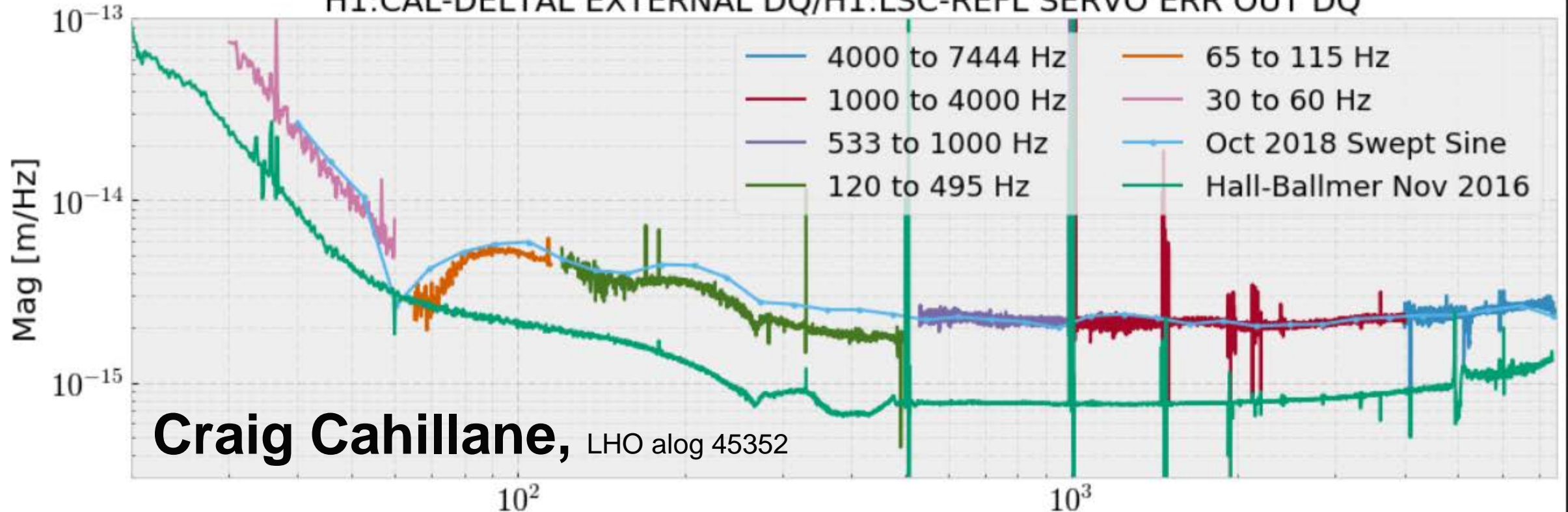
Daniel Sigg

LIGO Hanford Observatory

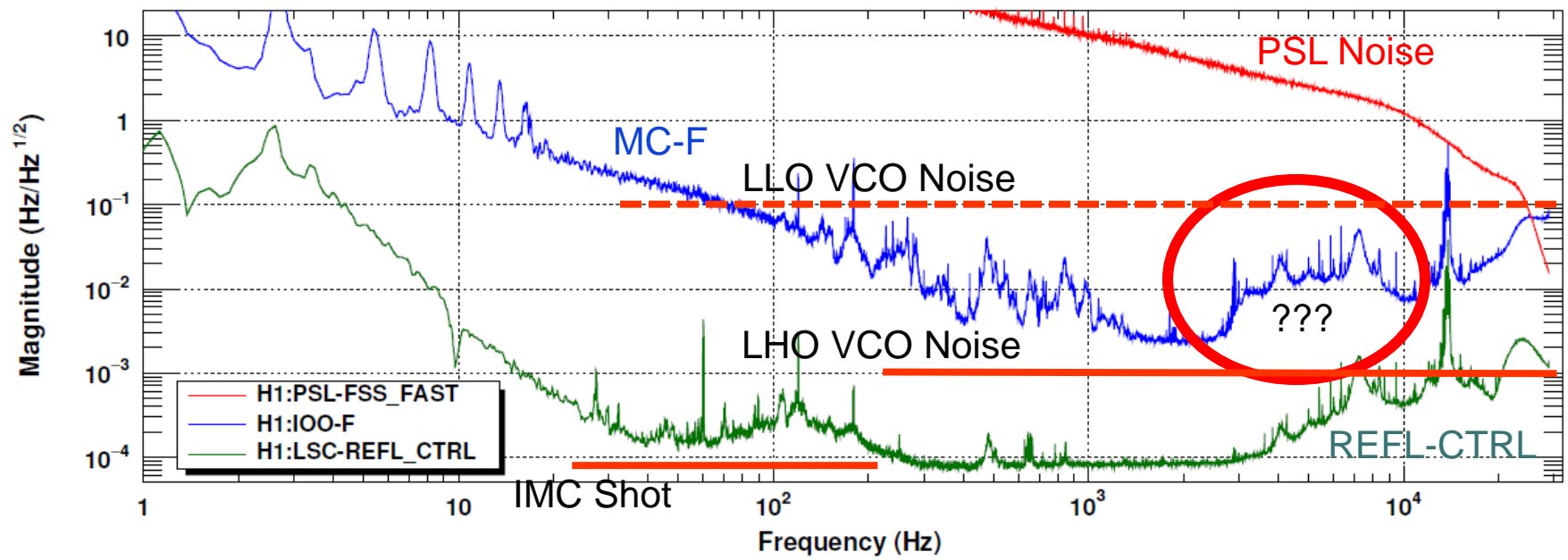
**Commissioning Meeting
July 12, 2019**

- Frequency noise coupling to DARM tends to be non-negligible for $f > \sim 1\text{kHz}$
- Frequency noise coupling to DARM flattens out due to higher order mode couplings at $>100\text{Hz}$, measured $1\text{-}2 \times 10^{-15} \text{ m/Hz}$ (also \sim flat in W/rad)
- Frequency noise coupling depends on differential heating
Can only get worse the higher the power

Band Limited Frequency Noise Injections TFs - Nov 16, 2018
H1:CAL-DELTA EXTERNAL DQ/H1:LSC-REFL SERVO ERR OUT DQ



Power spectrum

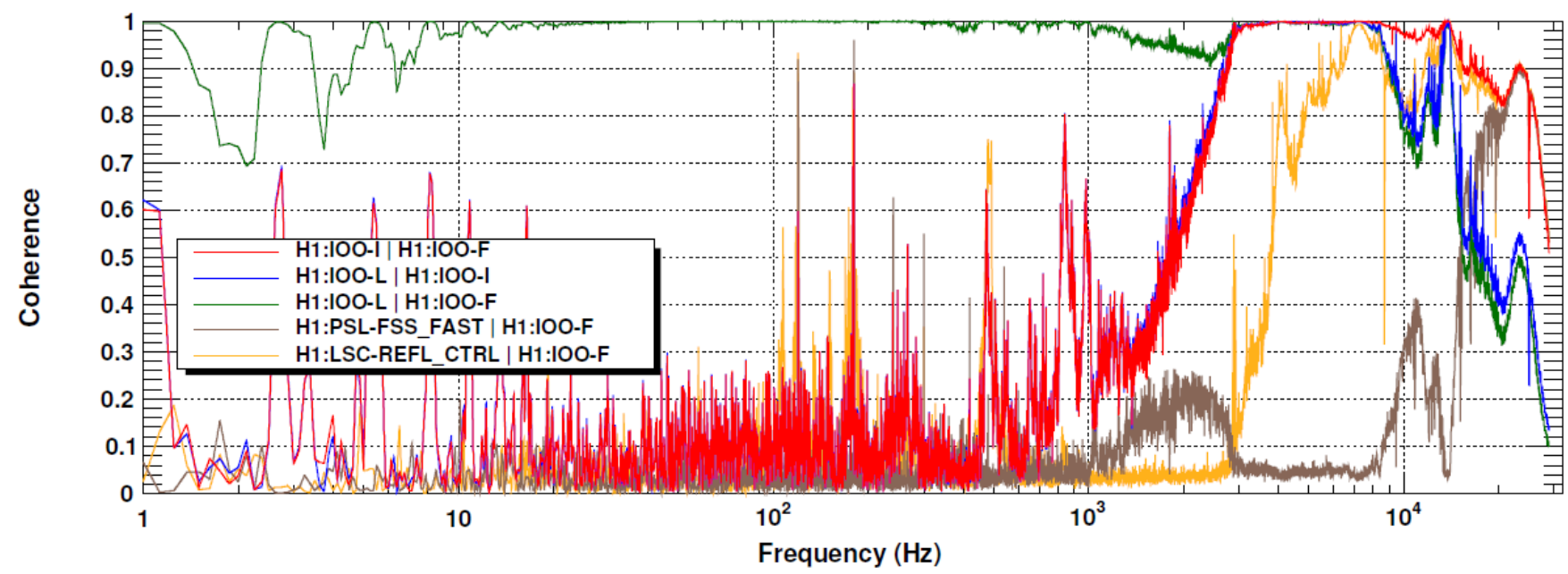


37W at input of IMC

Feature at 3-10kHz:

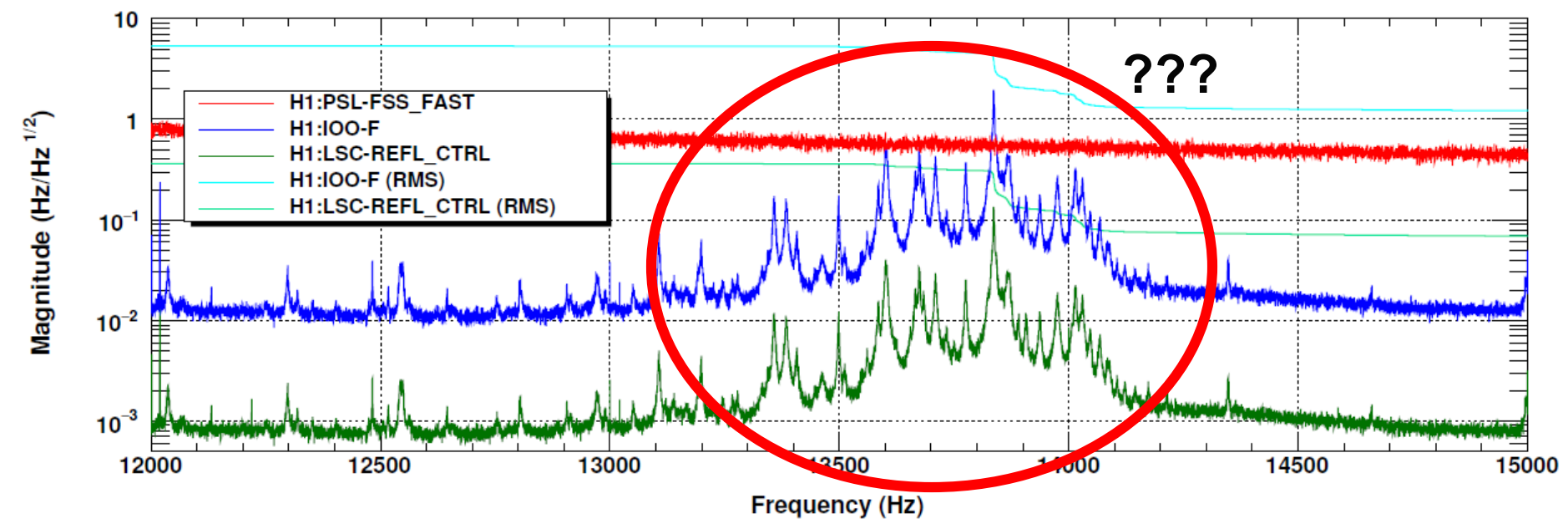
- Not NPRO noise
 - Ref. cavity noise?
 - From PSL?
 - From IMC?
 - Not IMC sensing noise
- IFO-REFL BW ~10-15kHz

Coherence



IMC sensing noise limited by poor modulation index:
 $20 \times \Gamma \rightarrow 10 \times$ lower sens. noise
 PD: 750mW unlocked
 9mW locked (99% carrier)

Power spectrum



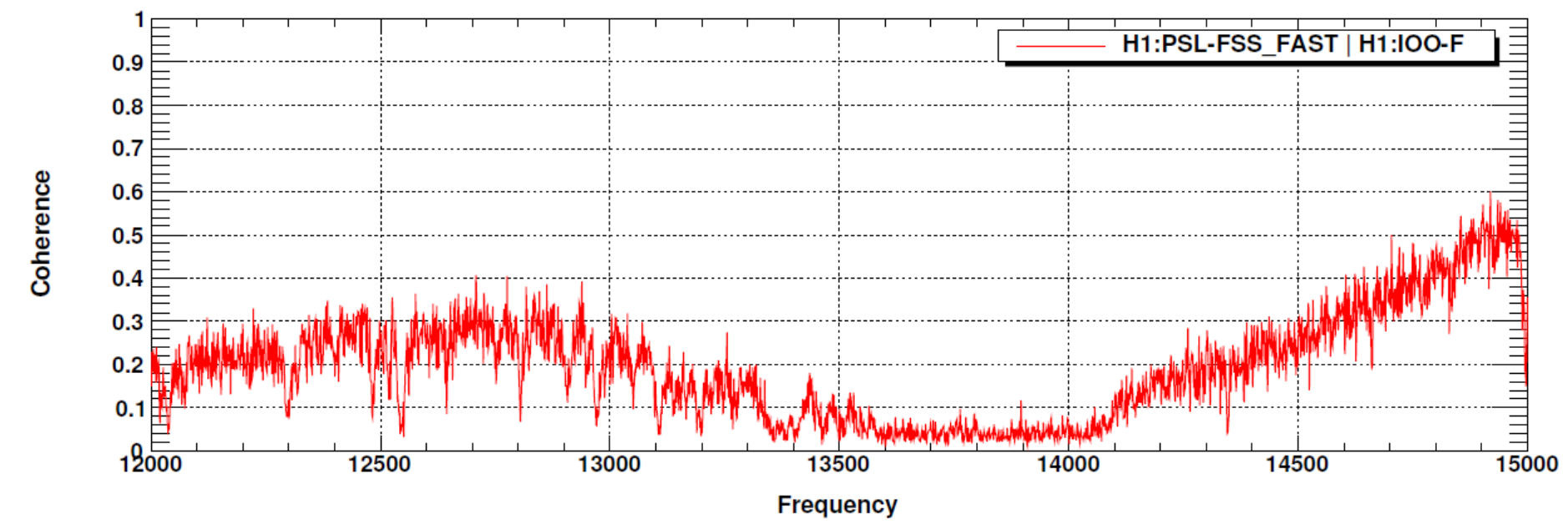
Peaks not in FSS FAST
Ref. cavity resonance?
Why so many?

T0=08/07/2019 17:25:44

Avg=30/Bin=2

BW=0.187465

Coherence



T0=08/07/2019 17:25:44

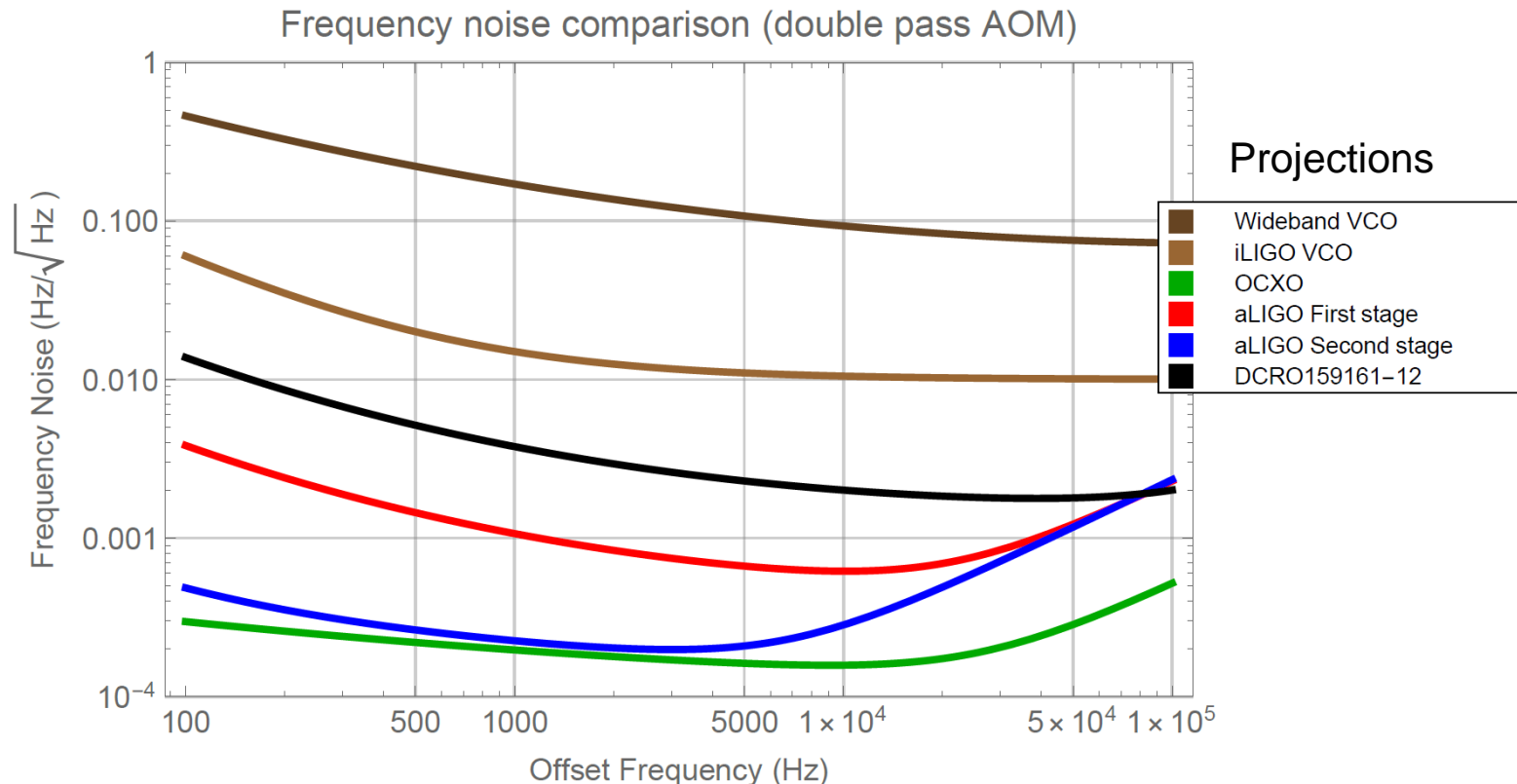
Avg=30/Bin=10

BW=0.187465

G1901316-v1

Available Options for VCOs (SAW)

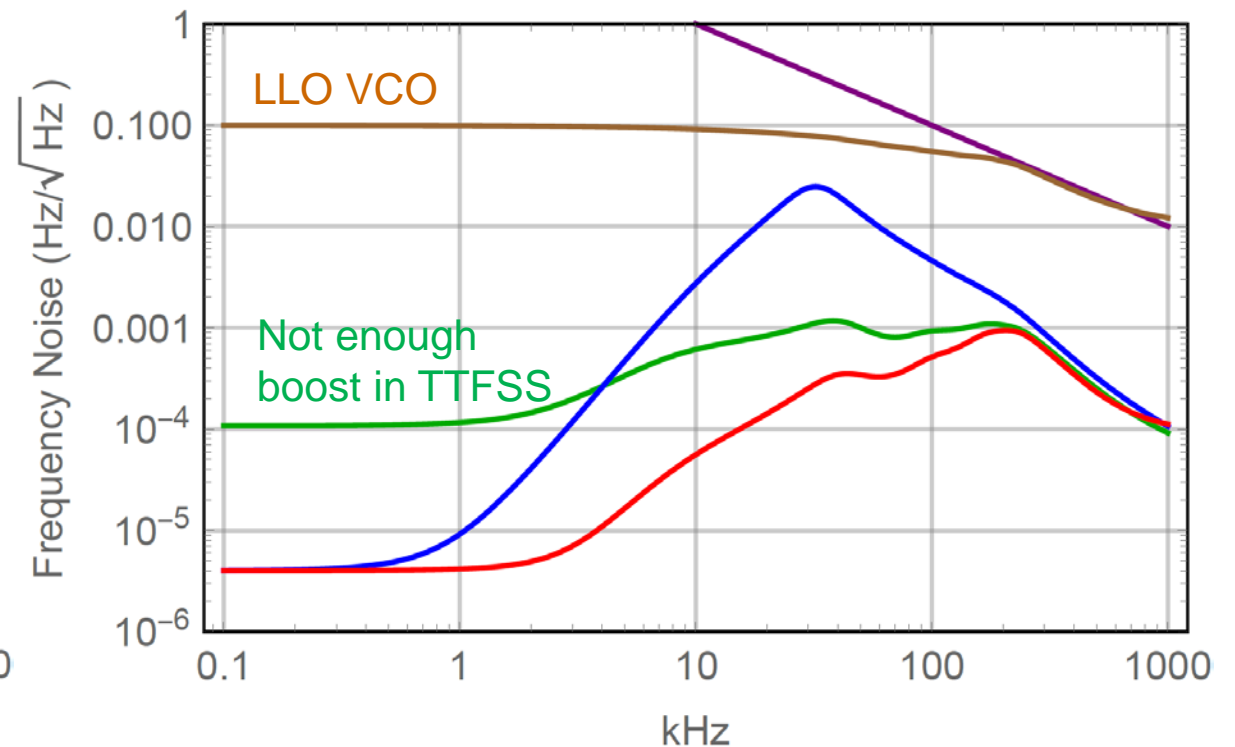
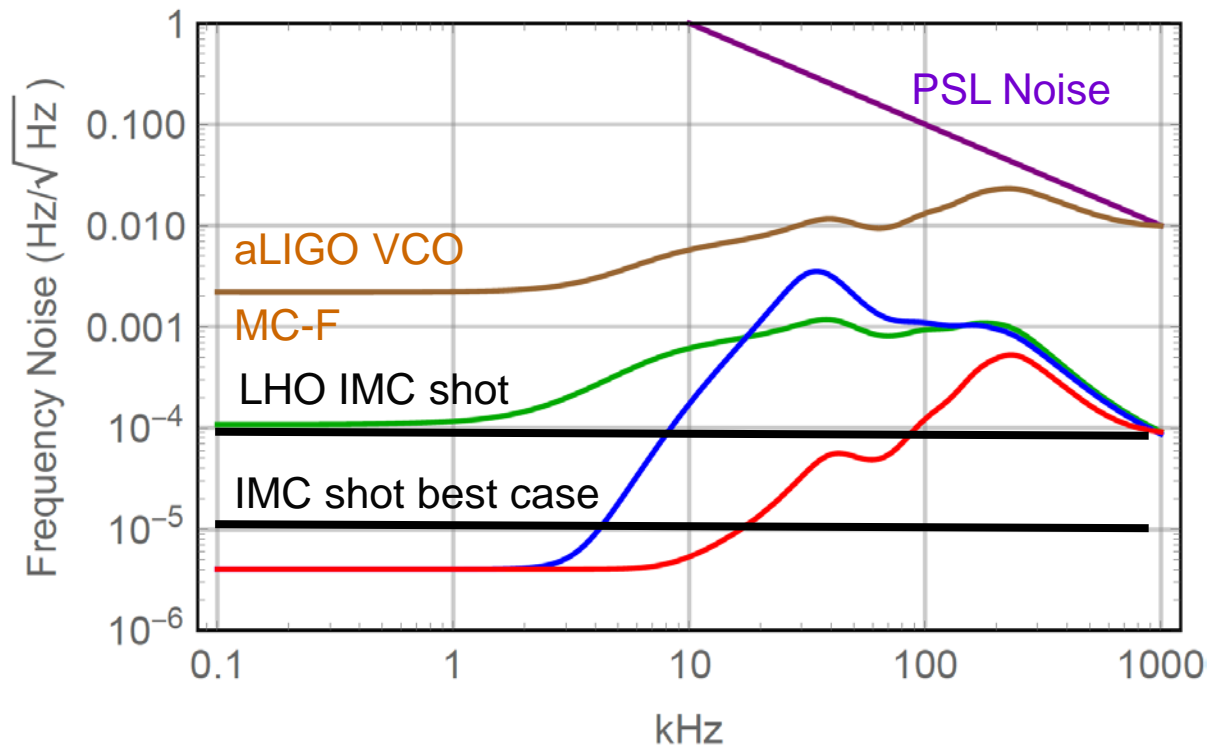
- Advanced LIGO VCO: 79.4 ± 1.2 MHz
1.05 \pm 0.14 GHz VCO (MCF91119-10) divide by 128 & mix with 71 MHz OCXO
- Initial LIGO VCO: 80 ± 5 MHz
800 \pm 50 MHz VCO (CRO750SA) divide by 10
- Best available SAW VCO: 79.6 ± 1.2 MHz
1.595 \pm 0.025 GHz VCO (DCRO159161-12) divided by 20
- 80 MHz VCO wideband (e.g., DCMO514-5)



Are we gain limited at high frequencies?

- ❑ Hierarchical approach: PSL \sim 500kHz BW, IMC \sim 100kHz BW
- ❑ No reference cavity: IMC \sim 500kHz BW, current TTFSS
- ❑ Serial approach: PSL \sim 500kHz BW, IMC \sim 500kHz BW
Requires IMC TTFSS and separate EOM

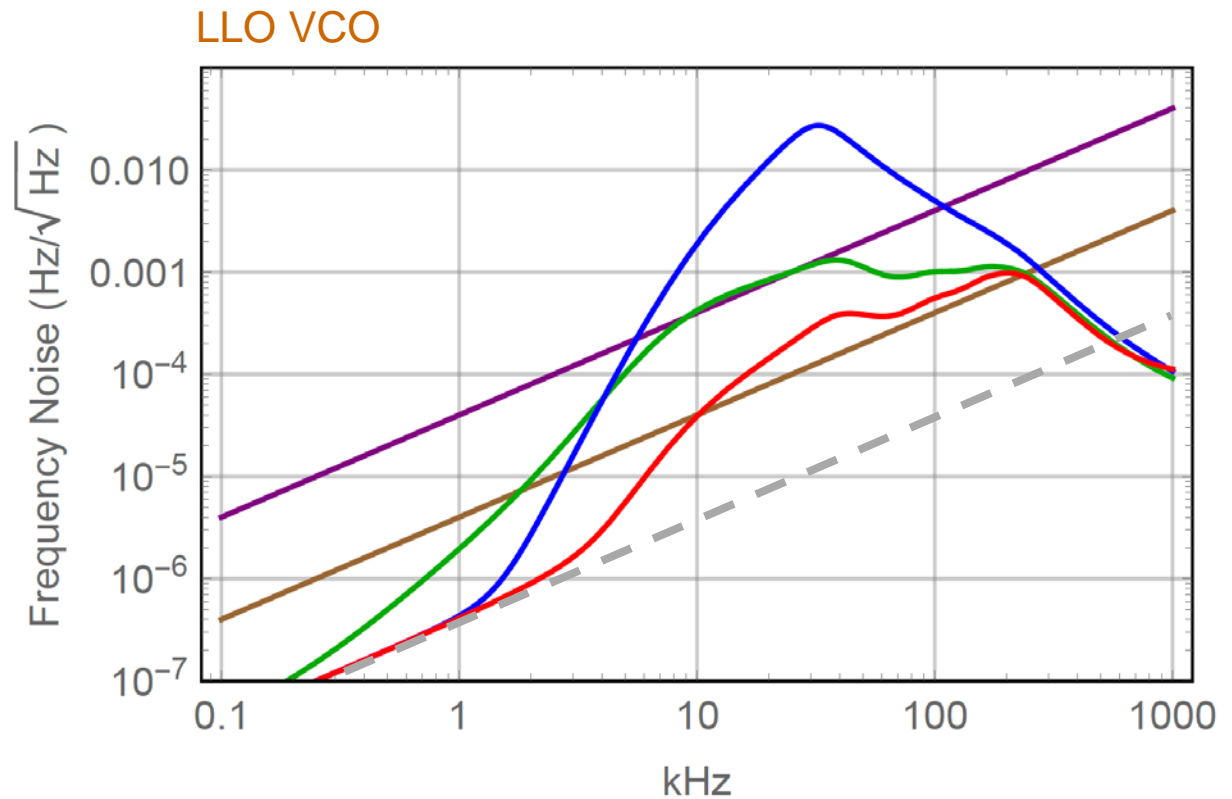
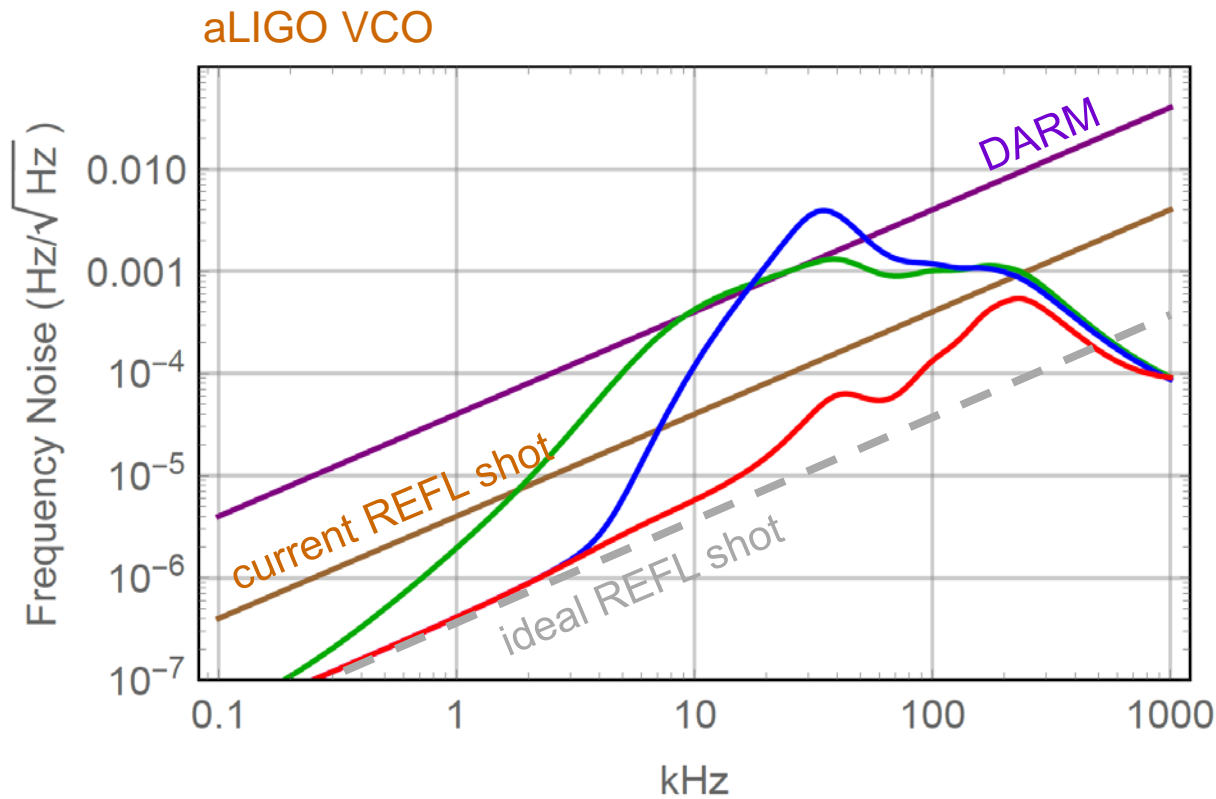
Conclusion: Serial approach & increasing modulation index will help >1 kHz and allow for increased VCO noise



Interferometer Projection (assuming ideal IMC sensing noise)

- ❑ Hierarchical approach: PSL \sim 500kHz BW, IMC \sim 100kHz BW, REFL \sim 15kHz BW
- ❑ No reference cavity: IMC \sim 500kHz BW, current TTFSS, REFL \sim 15kHz BW
- ❑ Serial approach: PSL \sim 500kHz BW, IMC \sim 500kHz BW, REFL \sim 15kHz BW
Requires IMC TTFSS and separate EOM

Below a 3-5 kHz: REFL sensing noise dominates

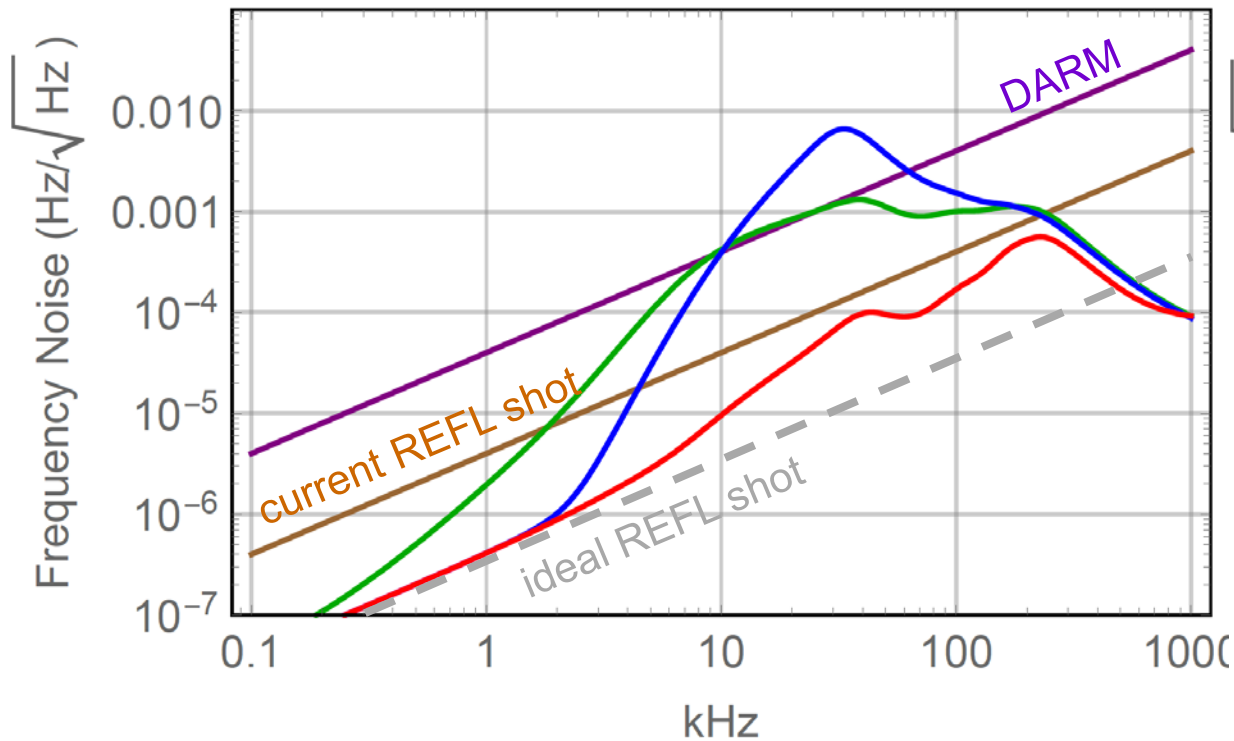


Interferometer Projection (Comparing IMC sensing noise)

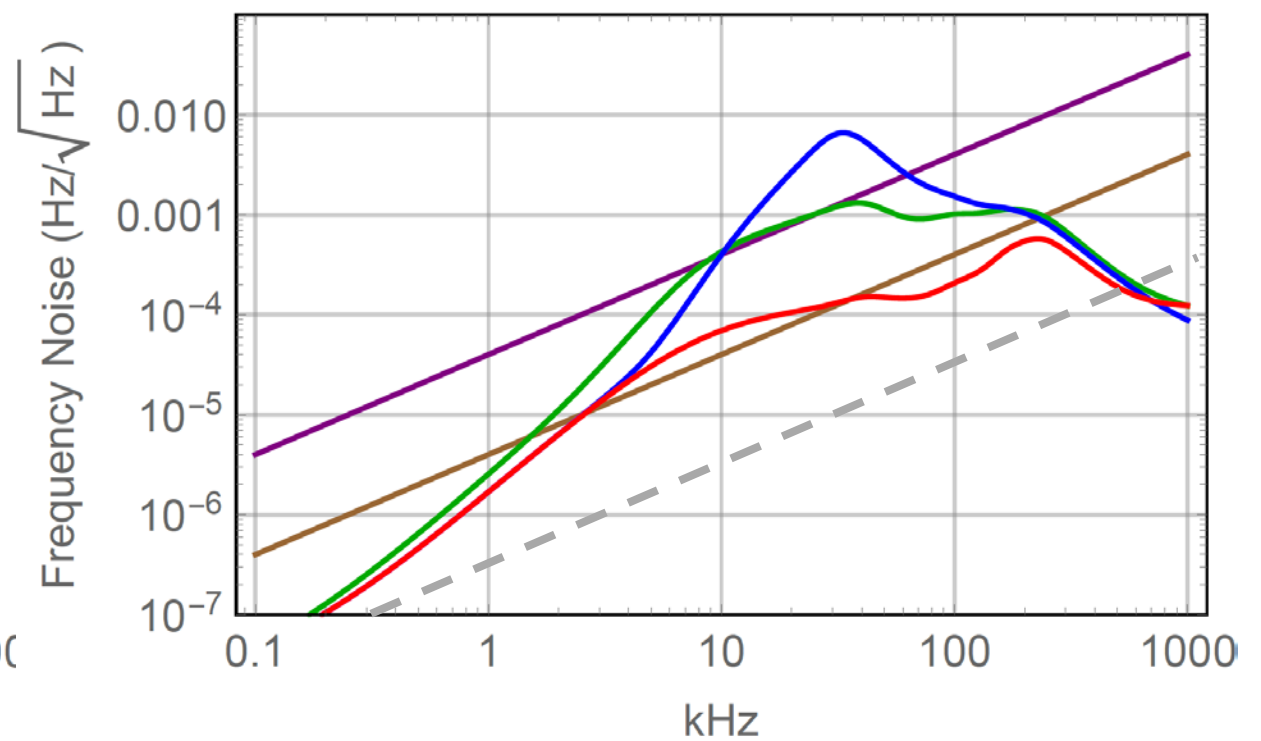
- ❑ Hierarchical approach: PSL ~ 500kHz BW, IMC ~100kHz BW, REFL ~15kHz BW
- ❑ No reference cavity: IMC ~500kHz BW, current TTFSS, REFL ~15kHz BW
- ❑ Serial approach: PSL ~500kHz BW, IMC ~500kHz BW, REFL ~15kHz BW
Requires IMC TTFSS and separate EOM

Above 3-5kHz: Better IMC sensing noise is required to take advantage of any REFL improvement

Ideal IMC Sensing Noise / iLIGO VCO



Current IMC Sensing Noise / iLIGO VCO



REFL Sensing (LHO)

□ REFL shot is high (estimate 10x over ideal): carrier dominated

	Input		AS (contrast?)		POP	REFL		Expect
Carrier	97.6%	37W	19.4%	=25mW	98.4%	82.5%	8.5mW	
9.1 MHz	0.9%	0.33W	6.3%	8mW	1.3%	3.9%	0.4mW	~4mW
45.5 MHz	1.5%	0.56W	74.3%	96mW	0.3%	13.6%	1.4mW	

Total AS port power is 240-270mW; numbers need to be multiplied by 2 to account for OMC.

REFL split ratio 1.25%, unlocked ~400mW.

40-50% of 45.5MHz sideband power is not accounted for.

70-80% of 9.1MHz sideband power is not accounted for: Loss where/why?

Recommendation:

- Investigate the 9.1MHz vanishing act
- Check carrier mode matching and high order mode content in reflection
- Maybe consider new RM with better optimized coating
- Increase REFL power (maybe limited by RF signals due to sideband imbalance)

Possible Actions

- ❑ Unknown prospects to reduce the frequency noise coupling
 - Higher power and improved squeezing will make frequency noise coupling more significant
- ❑ Investigate 9MHz sideband mystery
 - Is the poor sideband power related to optical distortions by the ITMs which produce HOMs?
 - Does this correlate with the high frequency noise coupling?
- ❑ Increase power in REFL
 - We are using 2 PDs at LHO already
 - Power is only 5mW/PD at 37W input, but RF signals tend to be large
- ❑ Reduce carrier in reflection by adjusting RM reflectivity
 - Is the REFL power dominated by carrier TEM00?
- ❑ New VCO: Propose to try 1.6GHz device
 - What's the excess noise seen by IMC-F above 3kHz?
- ❑ IMC sensing noise reduction (higher modulation index)
 - Dedicated EOM for IMC or dedicated electrode on existing EOM?
 - TTFSS for IMC (better performance by keeping reference cavity)
 - Neither will help <3kHz unless we also improve the REFL sensing noise