

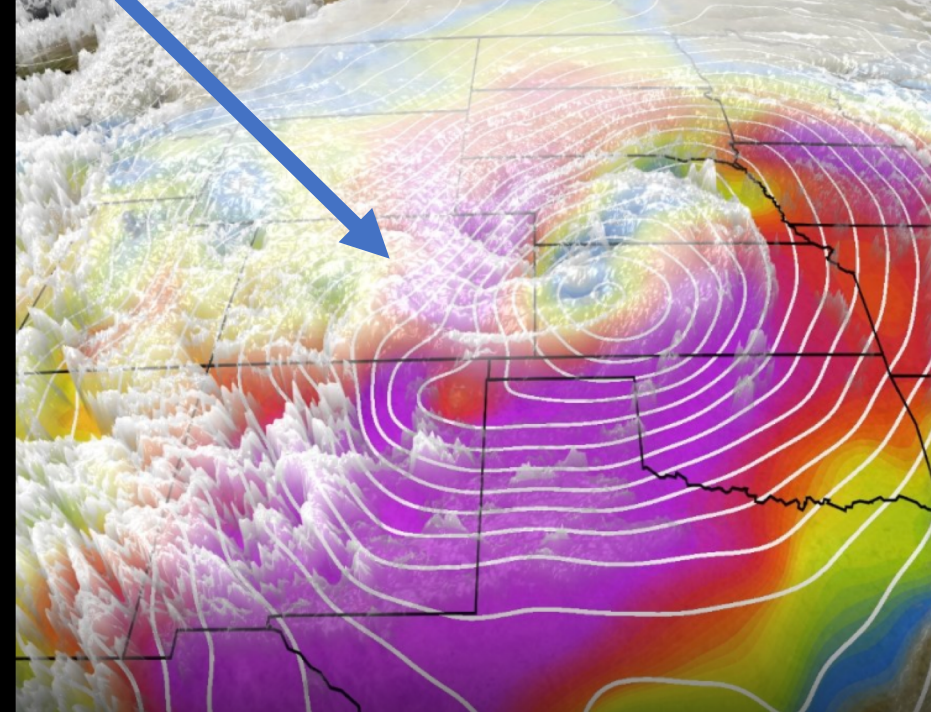
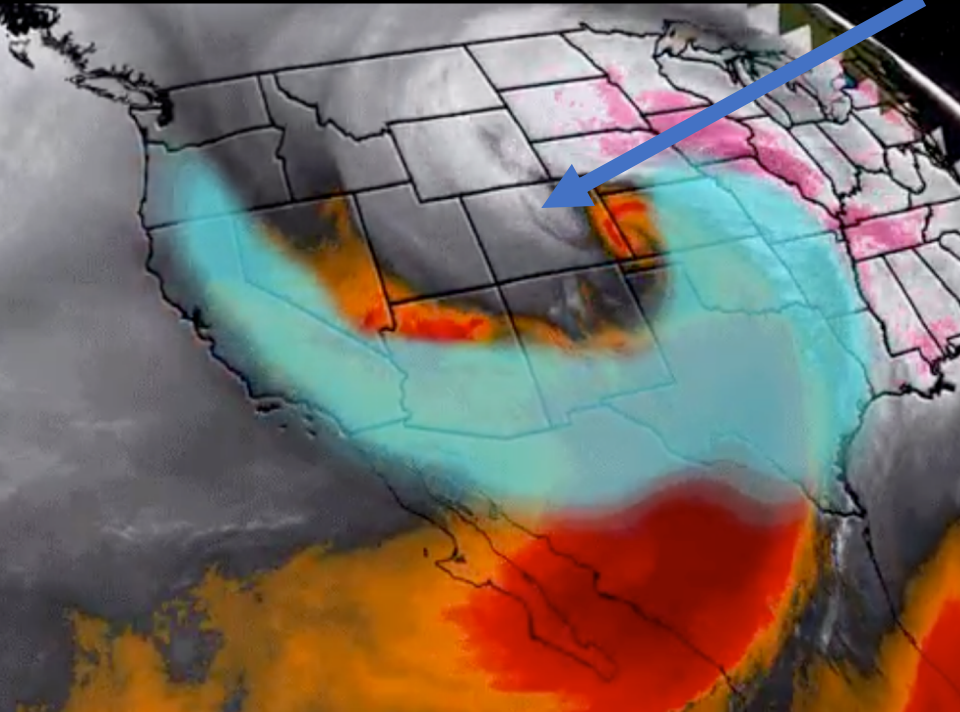


Absolute Uncertainty in the GW Detector Network:

Status Update and Debrief from
NIST Workshop

J. Kissel, for the World-wide collection of GW Detector
Calibration Groups, the LVK, and NIST

Denver / Boulder



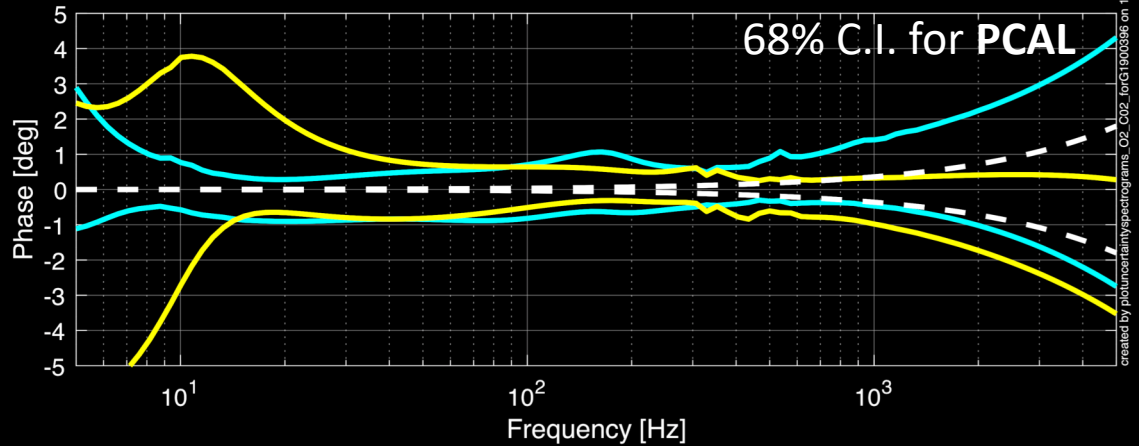
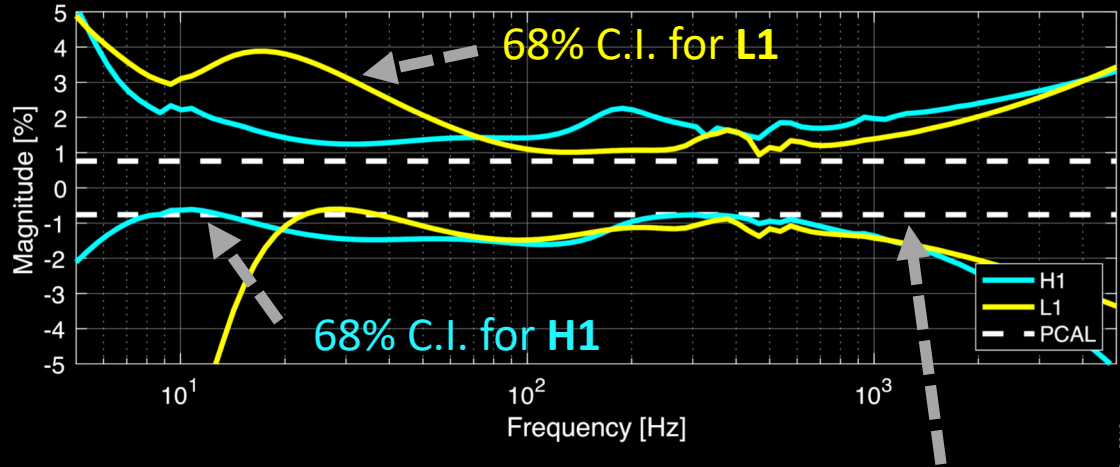
Winter Storm Ulmer destroyed all plans for attendees to fly into Denver

- Many attendees delayed
- Several attendees had cancel all together
- **Workshop reduced to one day (last Friday, 2019-03-18)**

The combination of

- **statistical uncertainty**
- **remaining systematic error**
- **unknown systematic error**

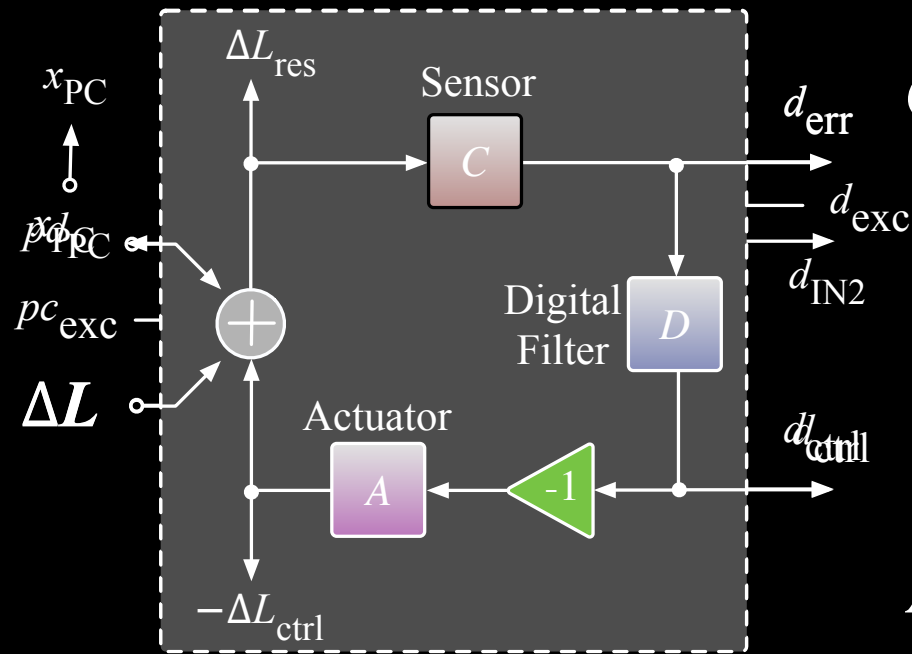
is numerically evaluated to form a 68% confidence interval, but roughly,



$$\partial h^2 \approx \partial R^2 \approx \left(\frac{1}{1+G} \right)^2 \left(\frac{\partial C}{C} \right)^2 + \left(\frac{G}{1+G} \right)^2 \left(\frac{\partial A}{A} \right)^2 + \text{Systematic Errors}$$

created by plotuncertaintyspectrums_O2_C02_forG190379 on 12-Mar-2019

With all loops closed and the detector running at its best sensitivity, we request a series of **in-loop** excitations to obtain direct measurements of the sensor and actuator



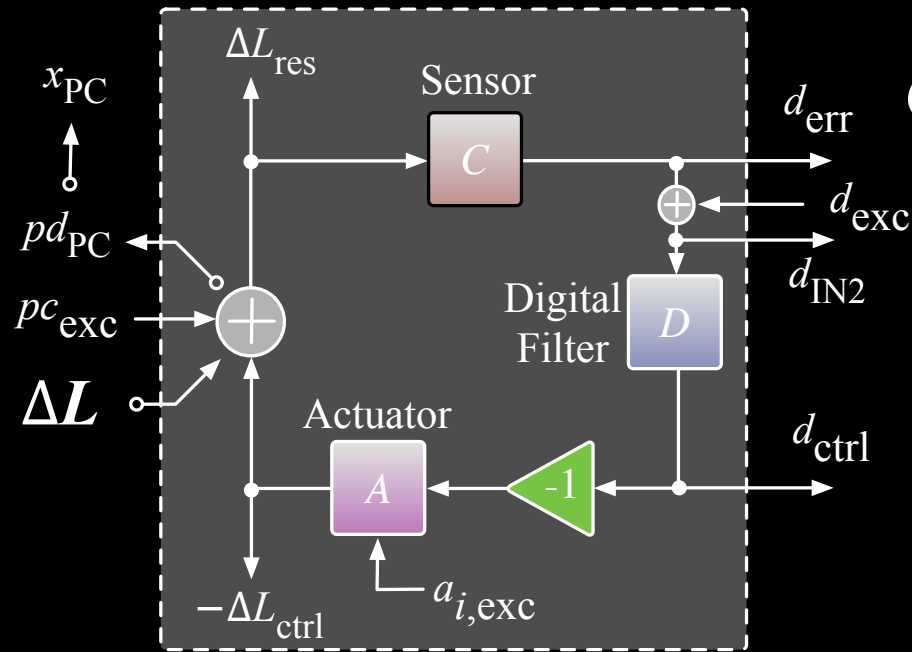
$$C = \left(\frac{1 d_{exc} G}{d_{IN2}} \right) \times \left(\frac{d_{err} C}{1 p_{c+exc} G} \right) \times \frac{x_{pc}}{p d_{pc}}$$

Complex Transfer Functions w/ Full IFO Absolute Scale Factor

$$A_i = \left(\frac{d_{err} C}{1 a_{t,exc} G} \right) \times \left(\frac{1 p_{c+exc} G}{d_{err}} \right) \times \frac{p d_{pc}}{x_{pc}}$$

Our measurements of C and A are ratios of complex transfer functions

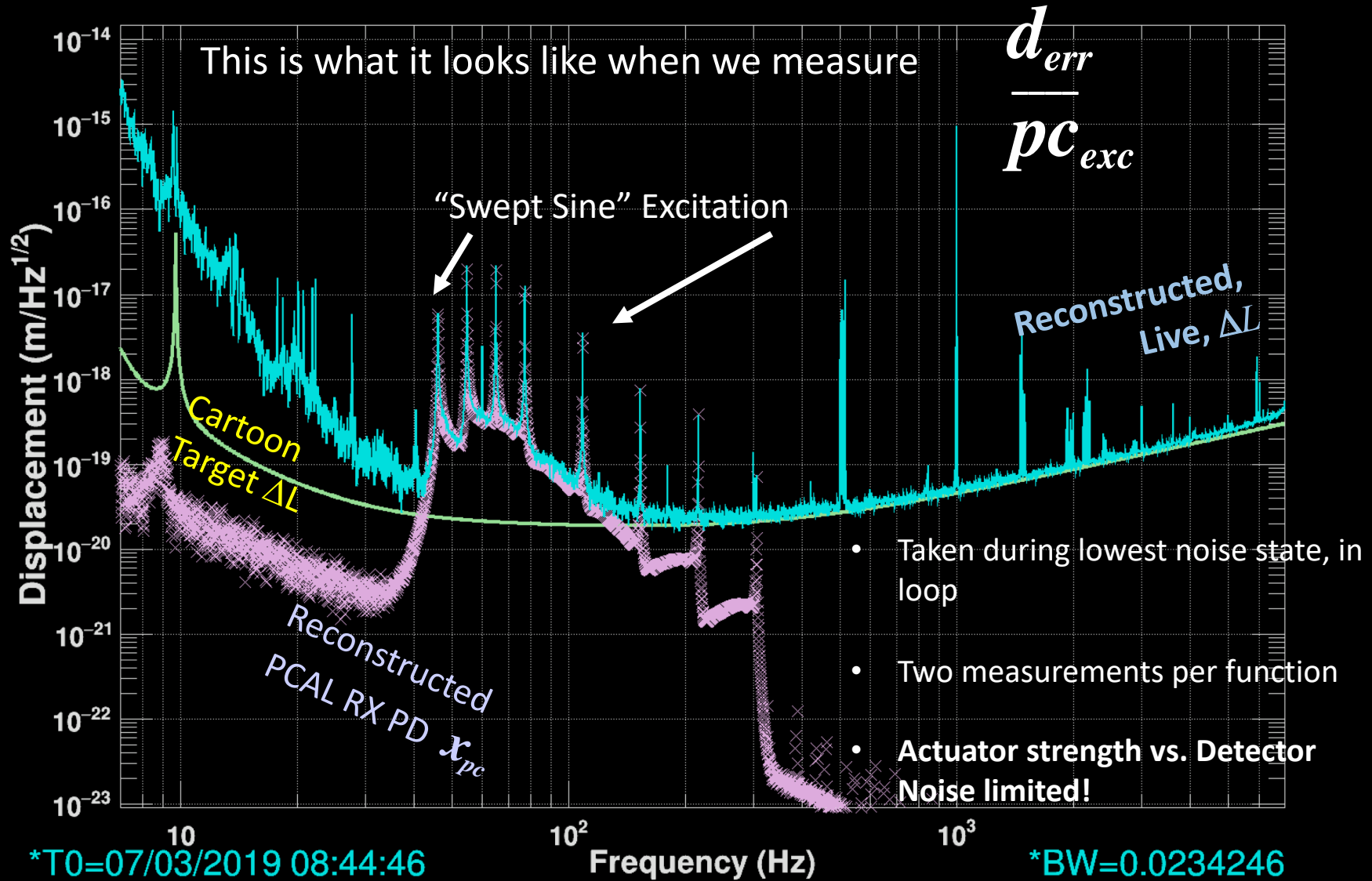
>> the frequency-dependent, magnitude and phase are important for uncertainty in A and C



$$C = \frac{d_{exc}}{d_{IN2}} \times \frac{d_{err}}{pc_{exc}} \times \frac{x_{pc}}{pd_{pc}}$$

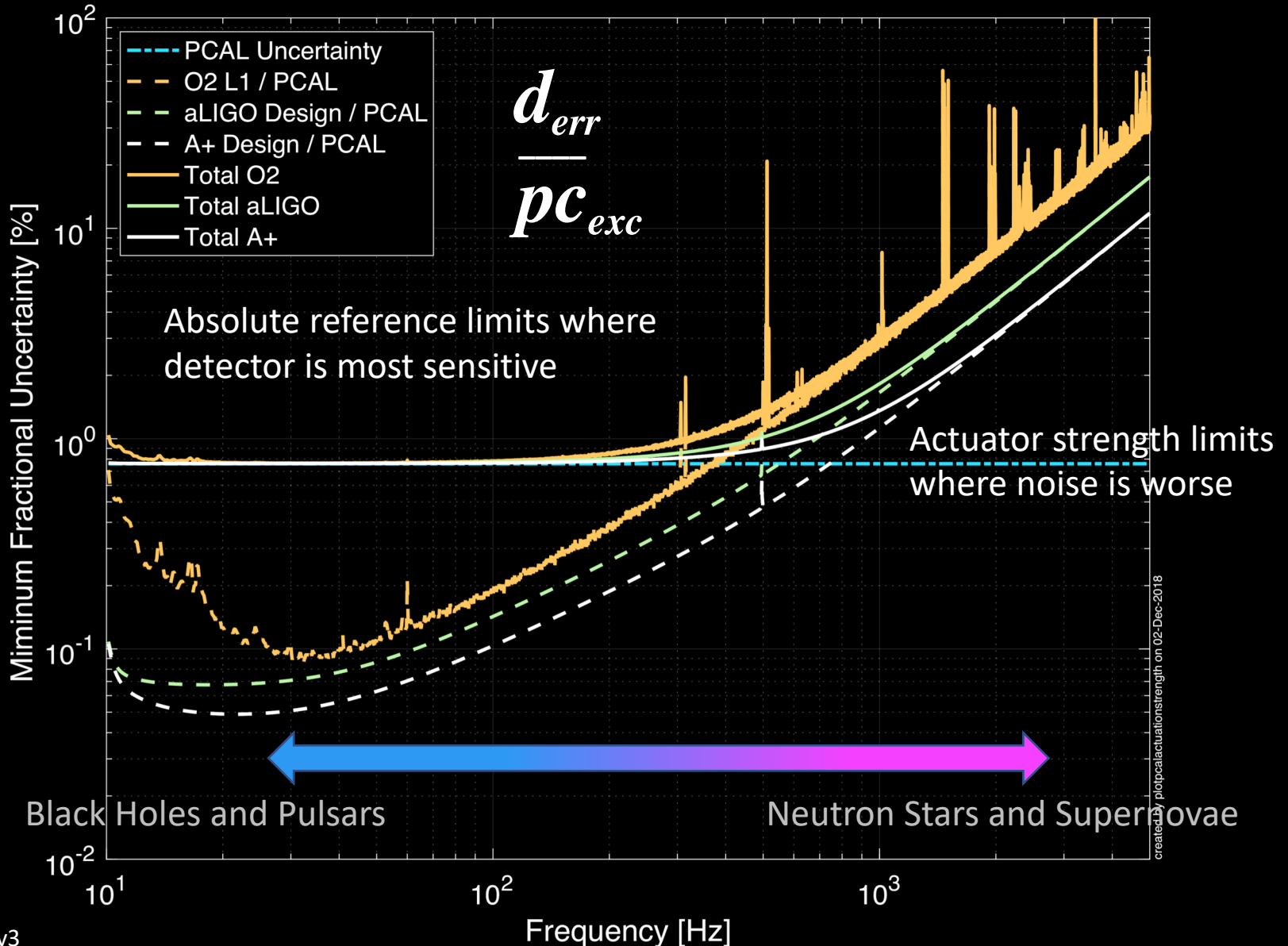
Complex Transfer Functions w/ Full IFO Absolute Scale Factor

$$A = \frac{d_{err}}{a_{i,exc}} \times \frac{pc_{exc}}{d_{err}} \times \frac{pd_{pc}}{x_{pc}}$$





Limitations (using O2 PCAL Uncertainty)





But We're Improving...



NIST



Reduces intra-network systematic bias

Improves sky-localization (for example)

We now have built and compared

FOUR
Working
Standards

to be held at each site, all referenced to the **Gold Standard** and thus **Standard** and thus **NIST**



But We're Improving...



LIGO PCAL Parameter	O2 [1]	O3 [2]
Laser Power (NIST - GS)	0.51	0.35
Laser Power (GS - WS)	0.03	0.10
Laser Power (WS - RX)	0.05	0.10
Optical Efficiency	0.37	0.10
Angle of Incidence	0.07	0.07
Mass of test mass	0.005	0.005
Rotation	0.40	0.20
Total	0.75	0.50



Overall improved between O2 & O3!

[1] Karki, S., et al. *RSI* 87.11 (2016): 114503.

[2] Sudarshan Karki's talk from the Workshop



NIST Workshop Take-away: NMIs are *Excited!*



- Our discoveries have stimulated excitement in the *global* power standards community.
 - “New” acronym: National Metrology Institutes (NMIs)
 - EUROMET study revisited – see slides 11 and 12
- NIST reduced their standard uncertainty *for us* for O3 (special signature required!)
 - see slide 9
- On their own, current NIST reference system for laser power is old and difficult to improve; they are already planning on upgrading their systems
 - see slide 13 and 14

Response of a Photodetector

Percent Difference from
Weighted Mean Response of a
Photodetector

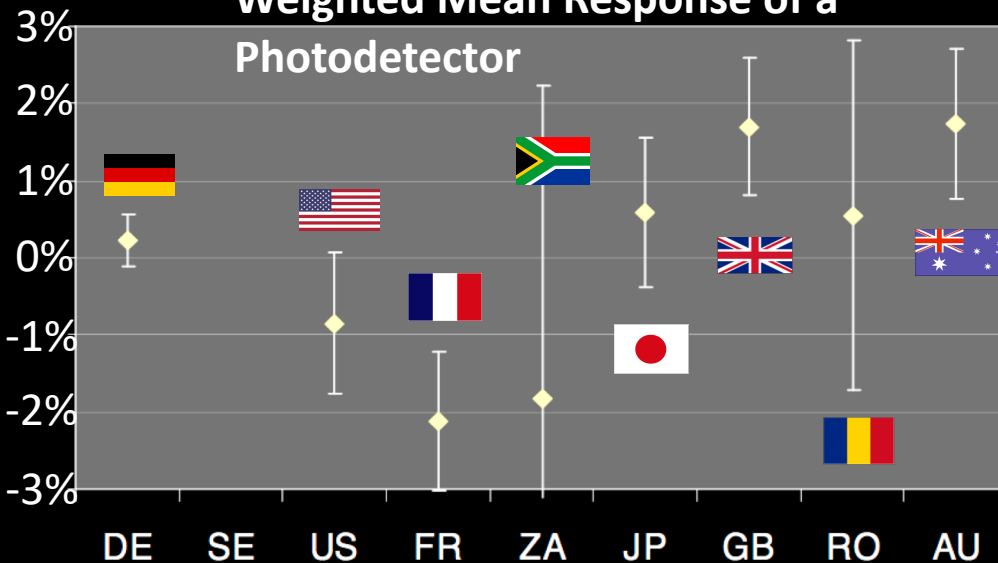


Figure 9

Kück *Metrologia* 47.1A (2010):
02003.

- NIST didn't and doesn't take it in to account
- Consensus – Initial concerns from GW community about “the plot” has been initially overblown.
- Proper statistical interpretation [3,4]:
 - NIST value and all others consistent with Consensus Estimate C.I.
 - Except France, who has since found ~2% systematic error (in the right direction)

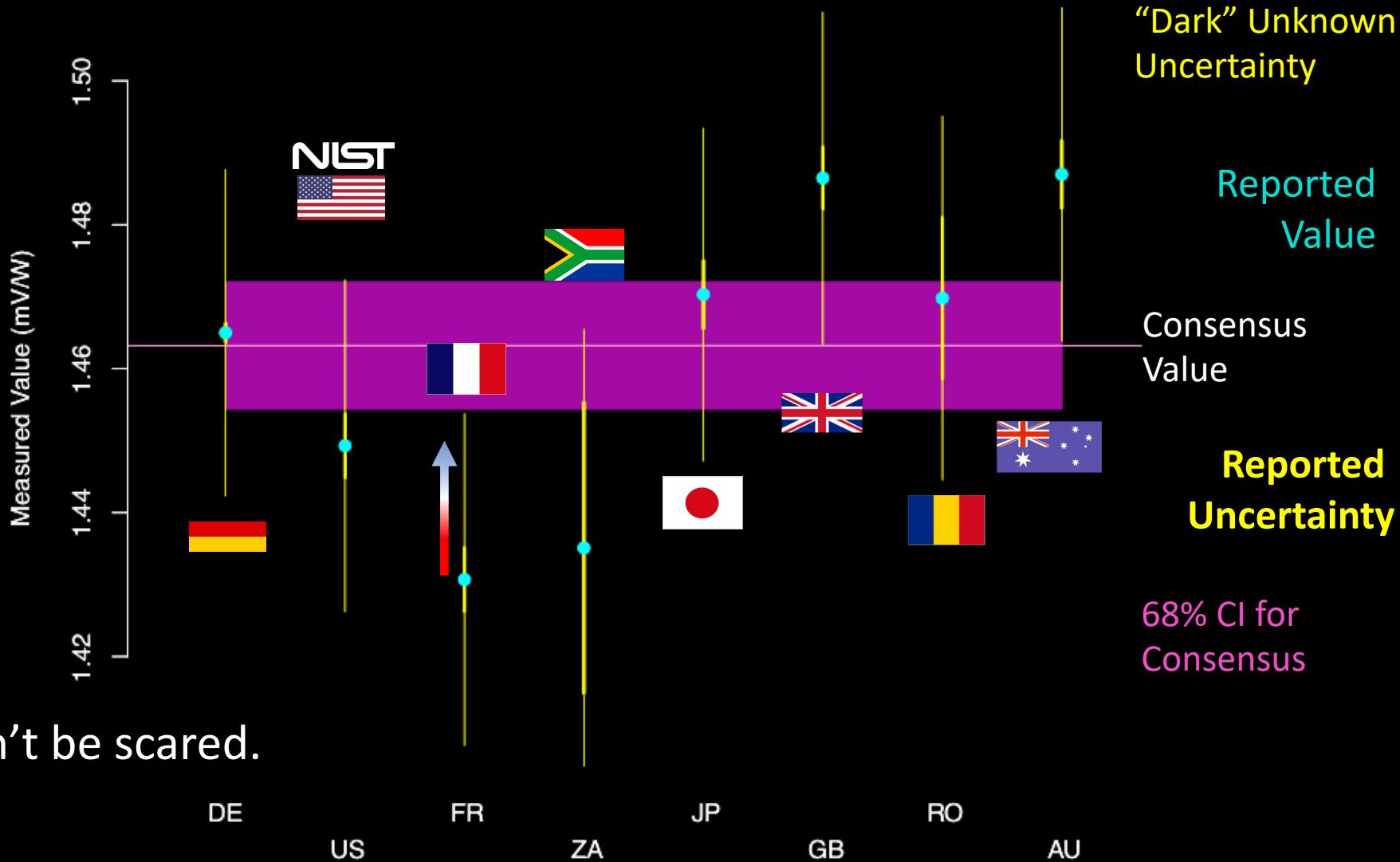
Remember this scary plot?

[3] Amanda Koepke's Talk at NIST Workshop

[4] Jimmy Dubard's Talk at NIST Workshop



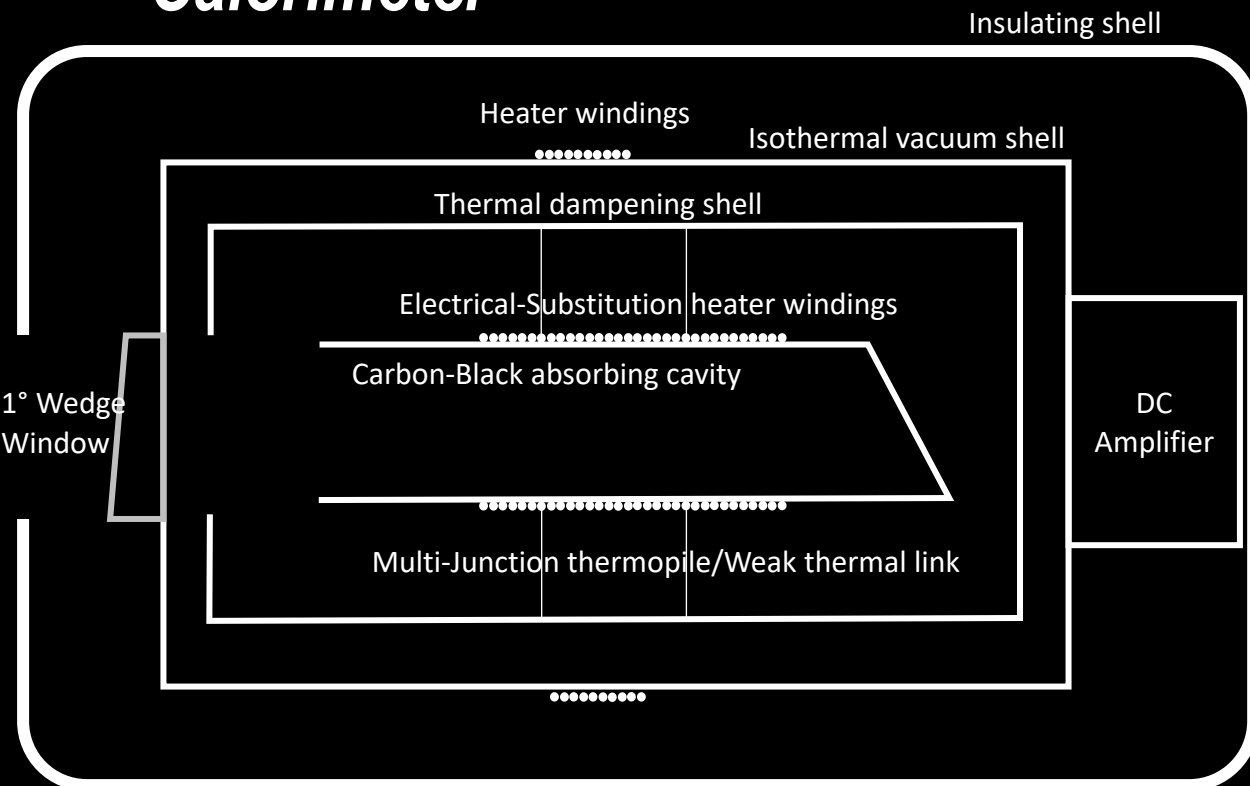
EUROMET Study Concerns Settled



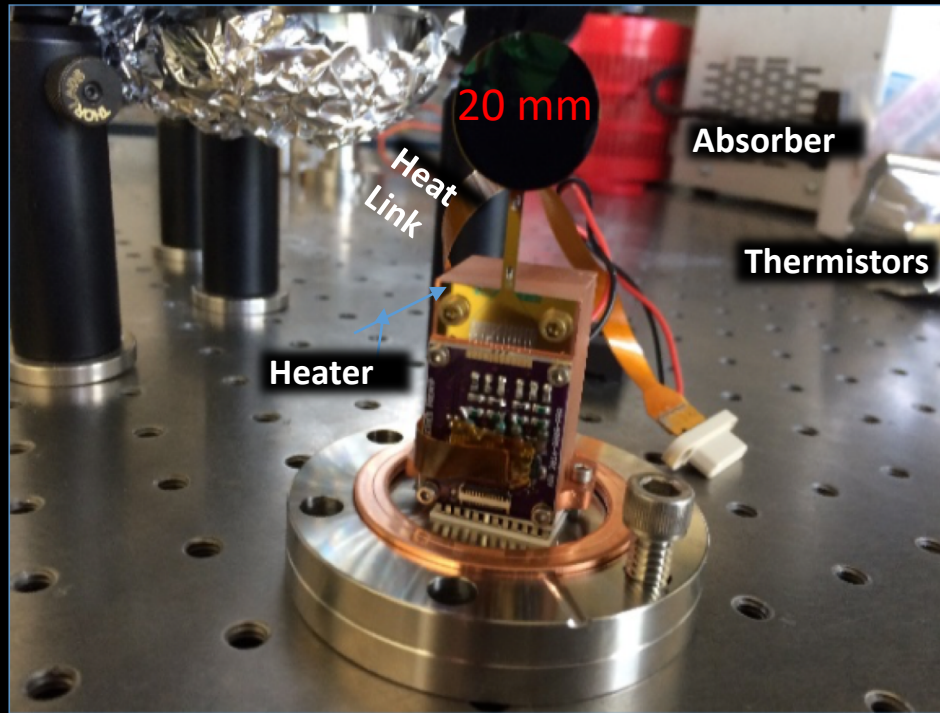
Data from Table 12, “Nd:YAG 1W” Kück *Metrologia* 47.1A (2010): 02003.

Results from consensus.nist.gov, encouraged by Amanda Koepke's Talk

Calorimeter



- Transfer method: beam splitter compares two calorimeters, then DUT replaces one of them
- Primary calorimeter traceable to *the* kilogram (now Planck's Constant) via electrical substitution
- Some recent improvements to optical systematics of method, external to calorimeter
- Further improvements to transfer method (and thus its uncertainty) require characterization of parts internal to calorimeter



Room Temperature Bolometer

Requirements:

- 2 cm beam diameter
- 100 μ W – 100 mW
- 325 nm – 1.93 μ m
- 2 minute time constant
- Mostly off-the-shelf

0.05% = 68% C.I.
time-scale = “next few years”

- Si substrate
- Peltier heater around perimeter
- Vertically aligned carbon nanotube absorber
- Transition edge sensor or commercial off-the-shelf thermistor or thin-film VOx for temperature measurement



The Message: Continue As Normal



- **NIST will be implementing a new primary reference**
 - on pace with our observing runs and improved sensitivity
 - The challenge is now on us >> preserve that absolute uncertainty all the way to the end stations and through the interferometers
- **All GW interferometers now have PCAL system and all are traceable to NIST.**
- [Bonus Slides] Observatories **will continue to research new/different standards** in parallel on a “best effort” basis
 - VIRGO will continue to with Laser Wavelength reference, and move forward with improved NCAL system
 - LIGO will re-visit Laser Wavelength and RF Oscillator references, and move forward with NCAL system
 - KAGRA is pushing hard on the NCAL effort, likely will try other methods

Thank you!





BONUS SLIDES



Workshop Event on DCC



- NIST, Labsphere and other NMI attendees graciously agreed to share their talks

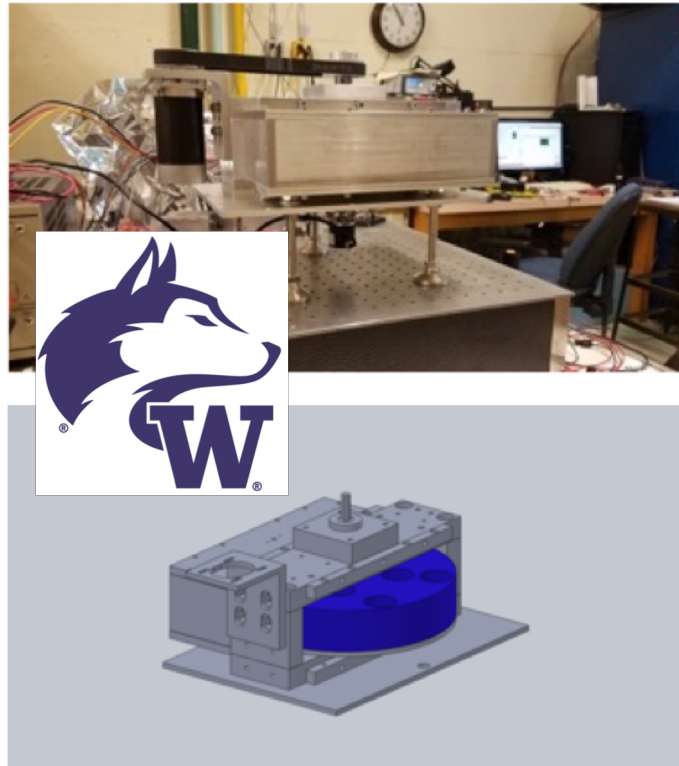
- All talks are available on DCC event page:

<https://dcc.ligo.org/cgi-bin/DocDB/DisplayMeeting?conferenceid=1029>

Virgo



LIGO



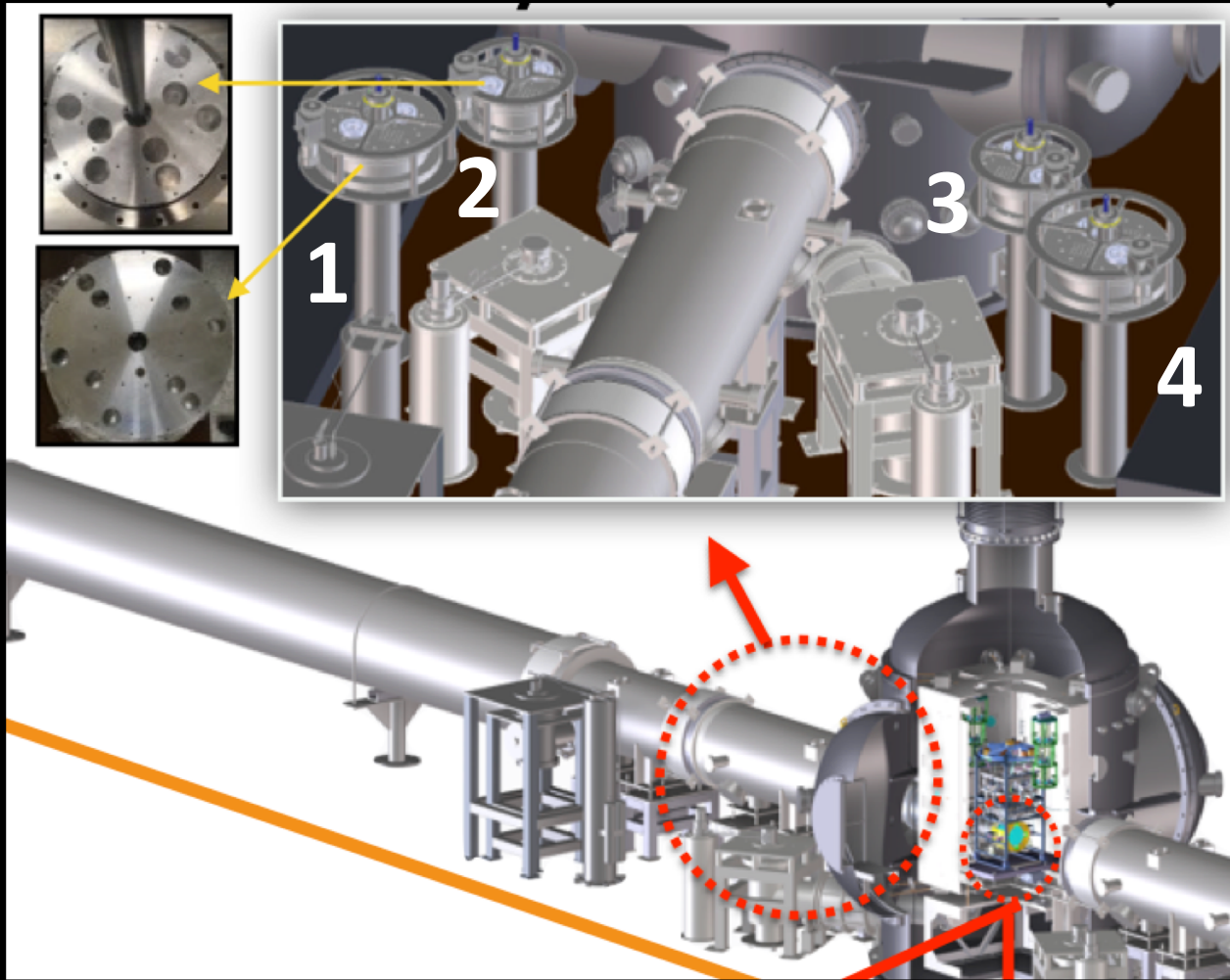
KAGRA



- World-wide excitement continues about Gravitational or Newtonian calibrators
- Develop on a “best effort” basis

<https://wiki.ligo.org/Calibration/NewtonianCalibrator>
ncal@ligo.org

Search “NCAL” or “GCAL” on DCC



- FOUR GCALs surrounding one test mass
- In concert with photon calibrator with 20 W laser
- An almost 3G calibration setup!

Inoue, Yuki, et al. *Phys Rev D* 98.2 (2018): 022005.

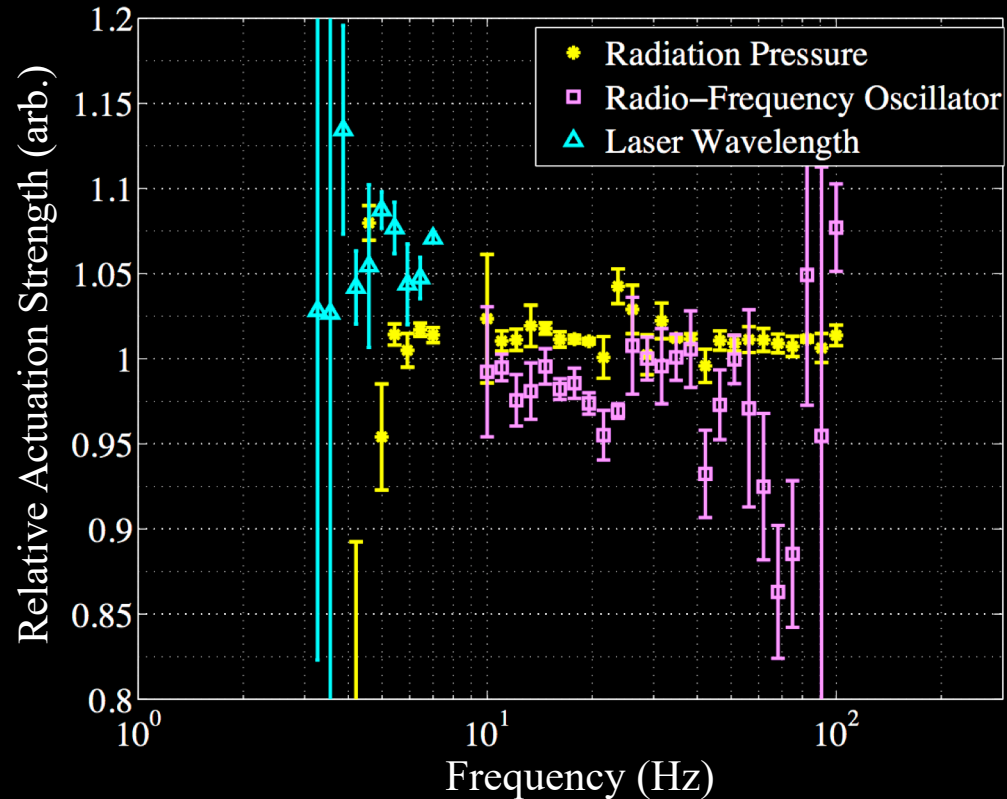
See Many more details in Yuki's Workshop Talk



LSC Fellows Revive other Methods



- In GW150914 paper, LIGO published a comparison of 3 methods
 - Laser Wavelength (old standard)
 - Radio Frequency Oscillator (re-invented using ALS)
 - Radiation Pressure (PCAL)
- Now detectors are better understood and controlled
- LSC Fellows Dripta Bhattacharjee (LHO) and Rachel Gray (LLO) will support PCAL, NCAL, and resurrect the pre-O1 study, and do it more justice



Always in seek of "The Truth," ...