

Thermal State of Advanced LIGO Test Masses: Implementation of a Real-Time Mirror Degradation Monitor

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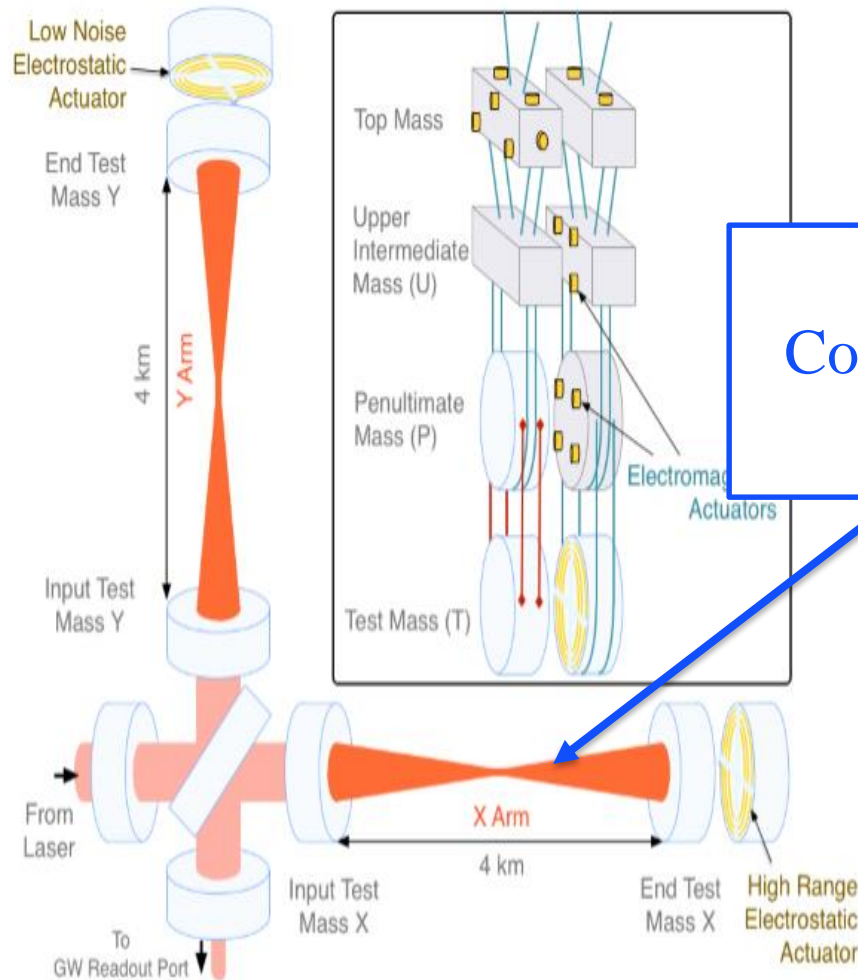
Objectives:

I will focus on the following topics:

- ❖ Background – Thermal Compensation System
- ❖ Finite Element Modelling
- ❖ Parameterization
- ❖ Kalman Filter Implementation
- ❖ Results
- ❖ Future Work

Advanced LIGO (aLIGO)

- ❖ Michelson interferometer with Fabry-Perot optical cavities
- ❖ Utilizes high-reflectivity fused silica mirrors
- ❖ Optical cavity with 800 kW ultimate optical power
- ❖ Apply Heat → Thermal transient forms in the mirrors



Thermal Compensation System

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Thermal Transient Effects

❖ Two Main Effects:

Thermal Lensing

- Aberrations in the beam
- Mode matching problems between cavities

Reduces sensitivity of detector

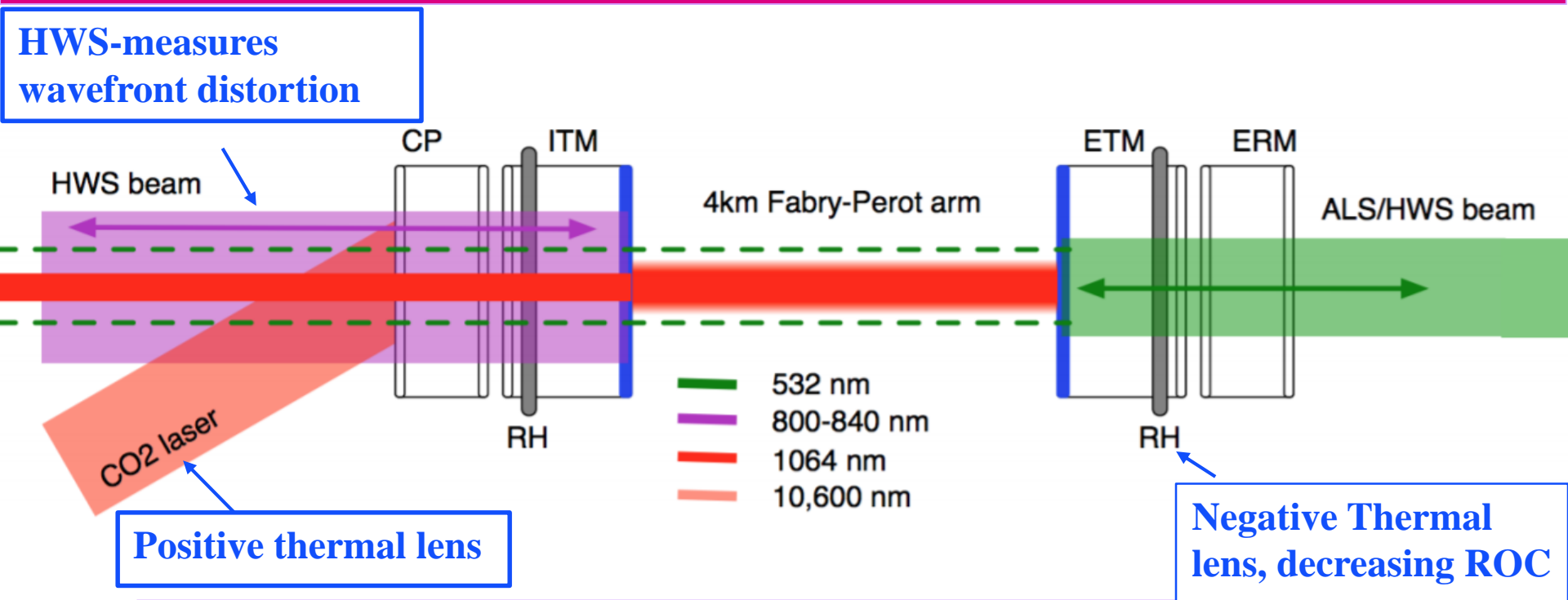
- Change of radius of curvature of mirrors (ROC)
- Shifts frequency of transverse optical modes (TEM)

Parametric instabilities
-Freq. between TEM & Fundamental mode = MM

Warming shifts mechanical mode (MM) frequencies

Thermal Expansion

Thermal Compensation System

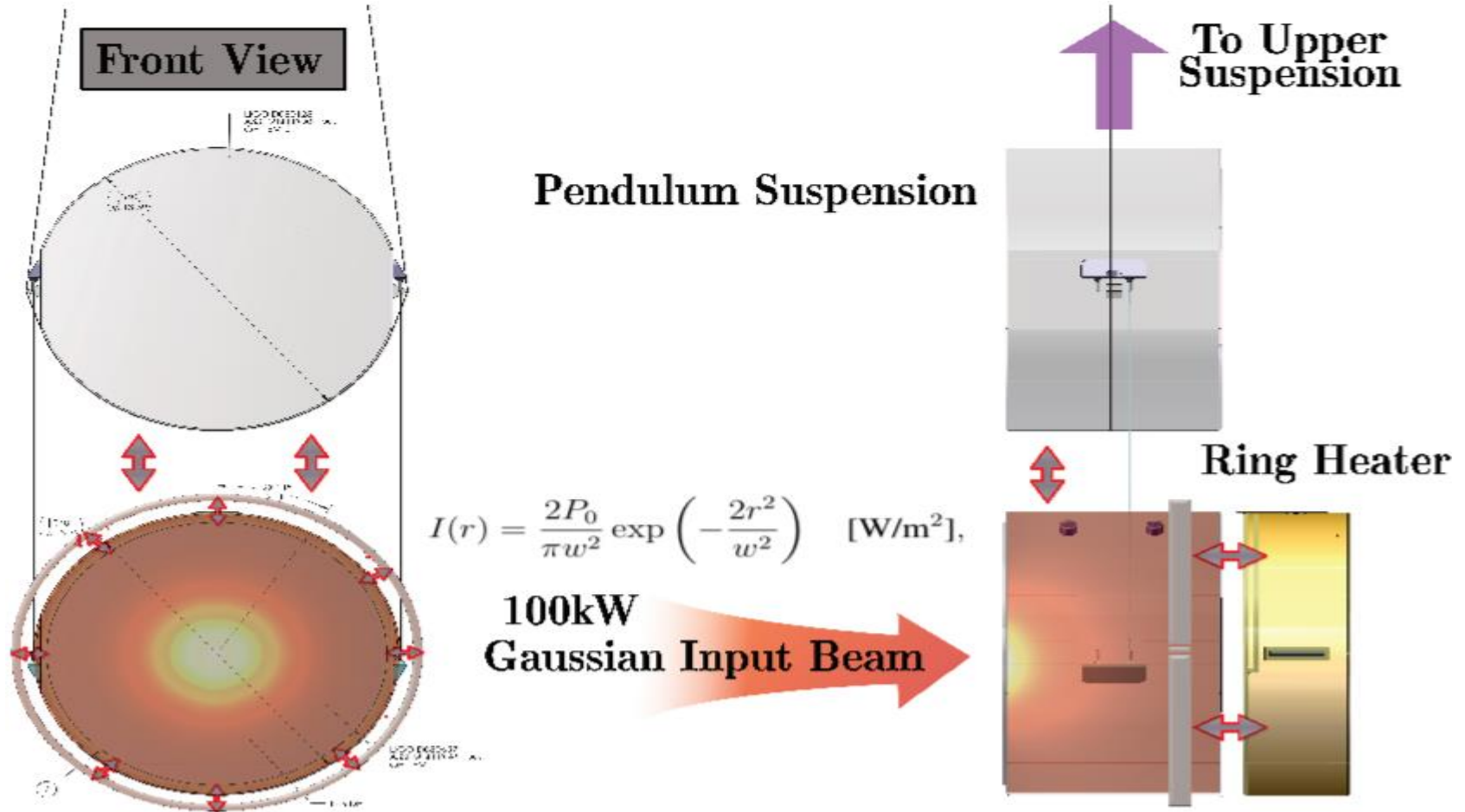


- ❖ Purpose: compensate for laser power absorbed in test masses
- ❖ Helps mitigate thermal lensing optical distortion effects

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[1]

Shift in Mechanical Mode Frequencies



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[1]

Finite Element Model: Method

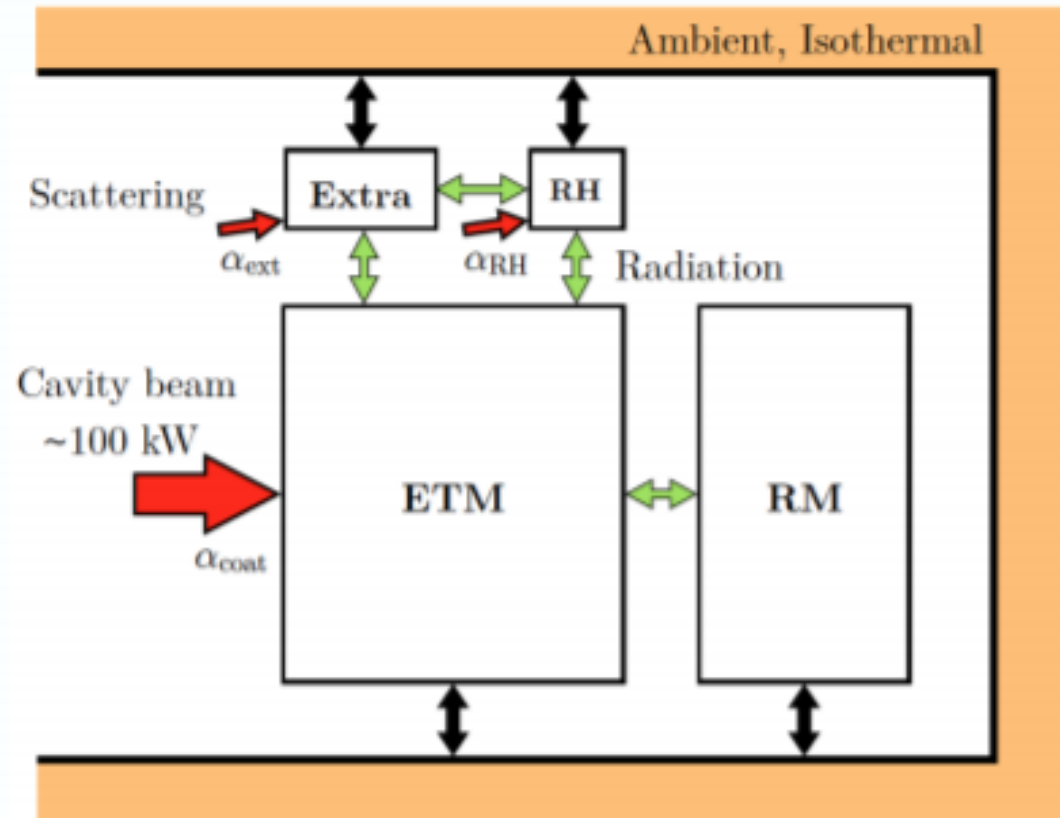
- ❖ Mechanical Mode Frequencies = Test Mass Thermometers
- ❖ Depend on:
 - ❖ Dimensions of the Test Mass (Mirror)
 - ❖ Elastic Constants:
 - ❖ Young's modulus $Y(T)_{bulk}$ - relation between stress and strain-uniaxial deformation
 - ❖ Poisson's Ratio ν - ratio between transverse strain to axial strain

$$\omega_m = \beta_m \sqrt{\frac{Y(T)_{bulk}}{\rho(1+\nu)}} \quad [1]$$

- Mechanical Mode Frequency: ω_m
- Constant Dependent on the geometry of the cylinder: β_m

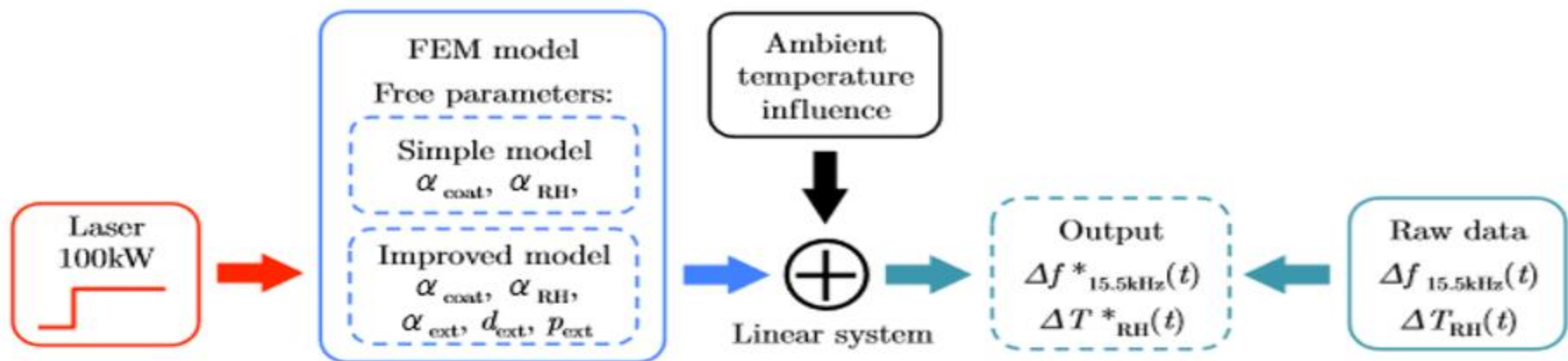
- ❖ Heat Transfer model between ETM & surrounding elements
- ❖ Transfer of heat when arm cavity is locked
- ❖ Mechanisms involved
 - ❖ RH
 - ❖ RM
 - ❖ Extra Term: Complex Structures surrounding it

Radiating Modelled Bodies



[2]

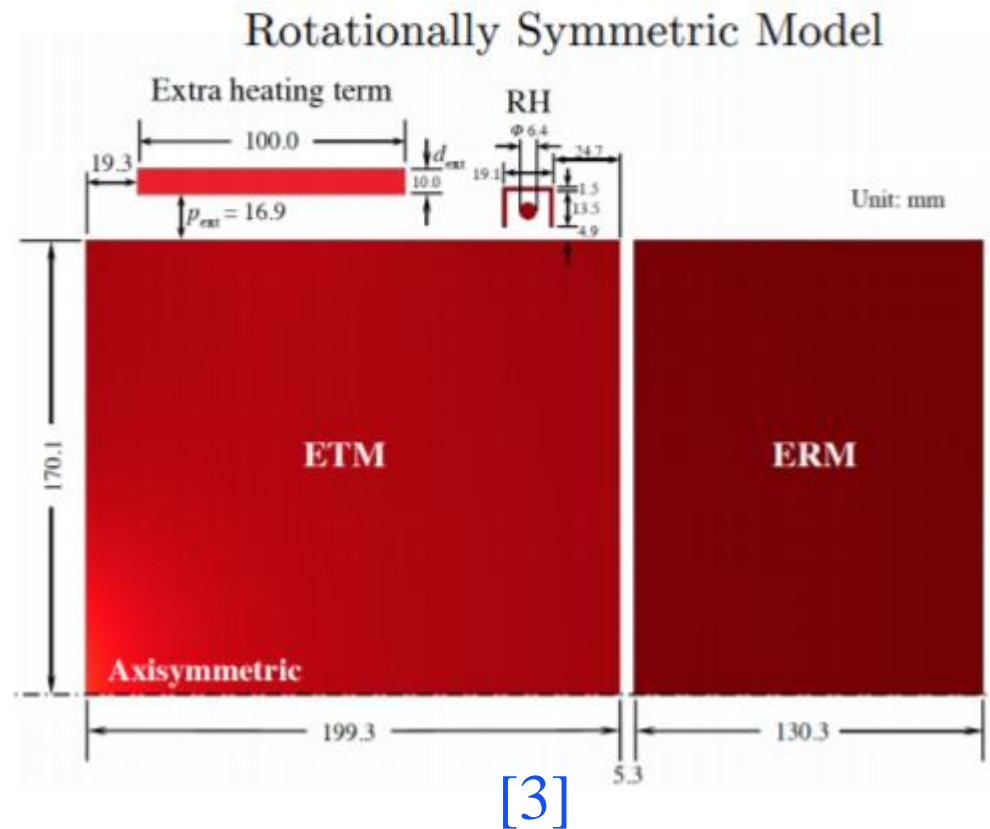
- ❖ aLIGO test mass
 - ❖ Cylinder: 170 mm radius, 200 mm thickness
 - ❖ Heraeus Suprasil 3001 fused silica
 - ❖ 100 kW laser beam
 - ❖ Coating absorption of 1 ppm corresponds to total absorbed energy 0.1 W
- ❖ Inputs, outputs, and the free parameters involved when modelling a test mass



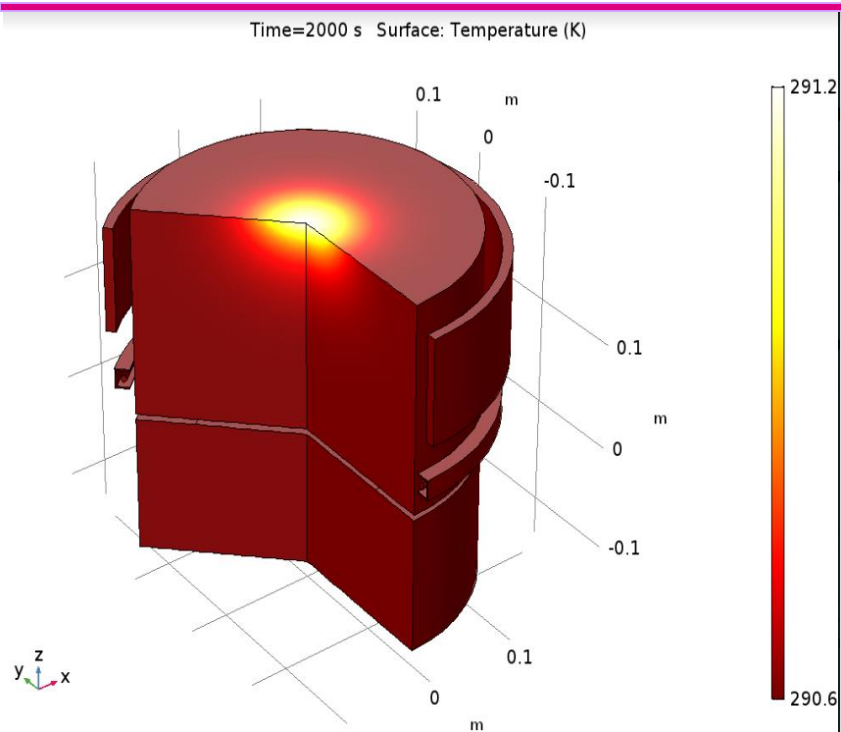
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Finite Element Model: COMSOL

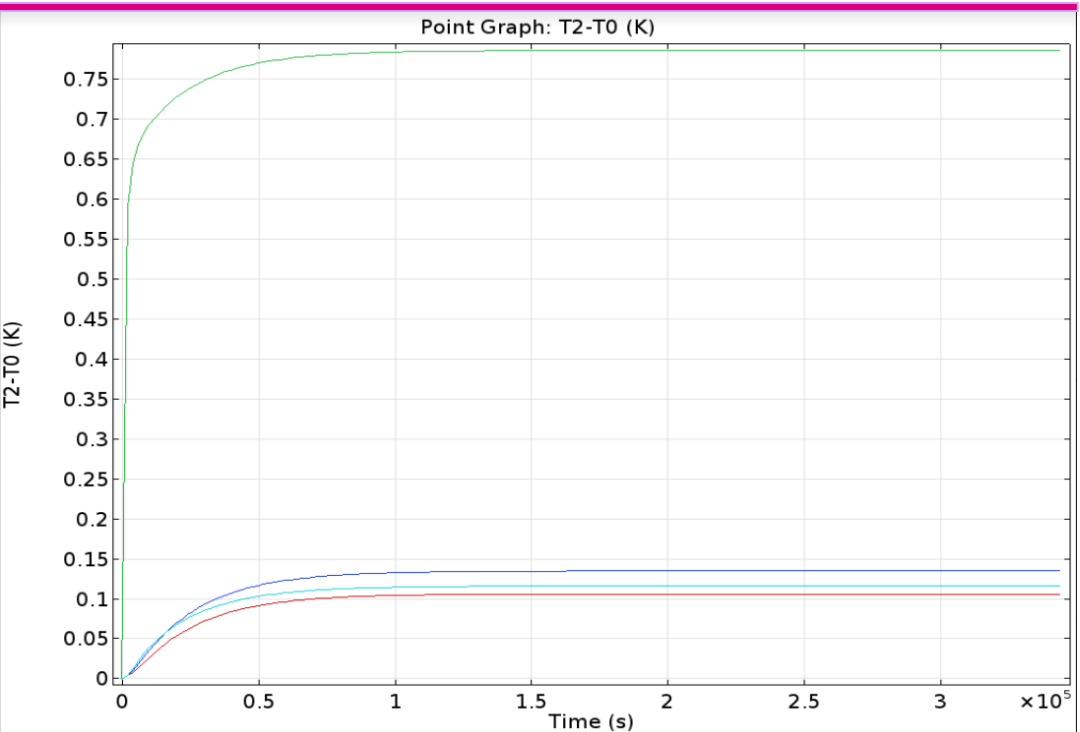
- ❖ 2-dimensional Axis-symmetric representation of the system
- ❖ Input laser beam heating load
- ❖ Restricted to only monitor circularly symmetric mechanical eigenmodes
- ❖ LIGO historic data for 5.9, 6.0 and 8 kHz modes
 - ❖ Only the 8 kHz mode is axis-symmetric



Finite Element Model: COMSOL



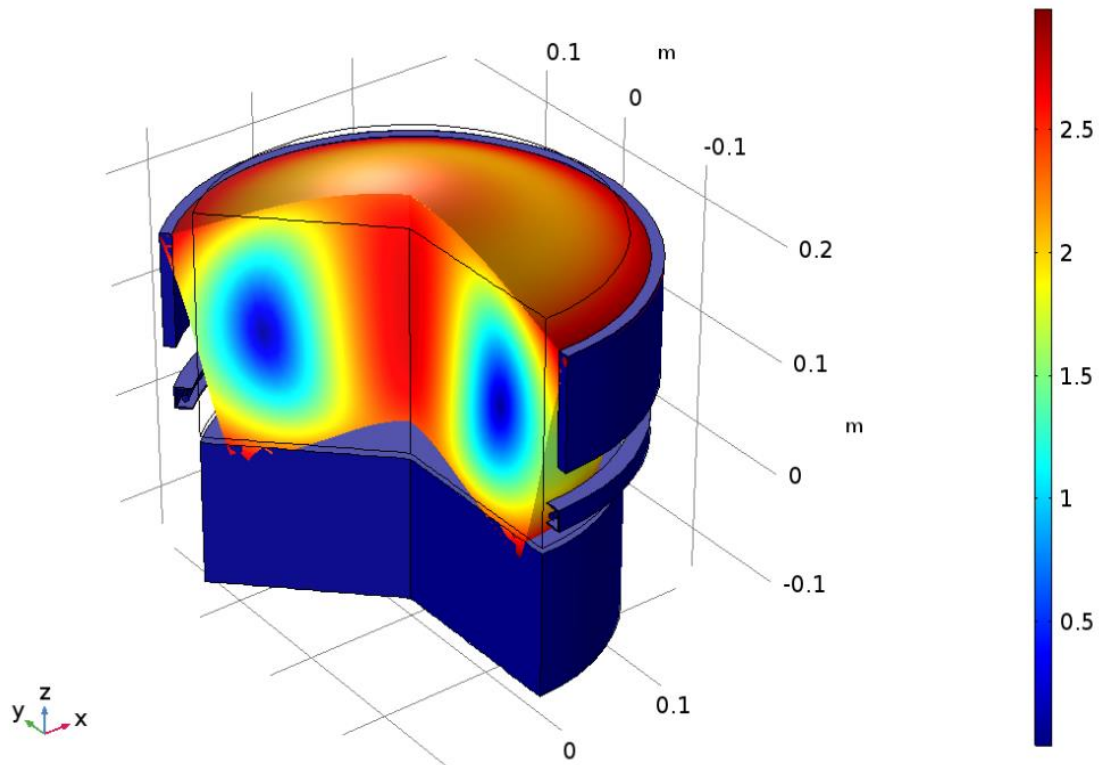
COMSOL depiction of ETMX



Temperature Change with Self Heating

Applying Heat Equation in a system with a fixed laser beam

Eigenfrequency=8126.161904 Hz Surface: Total displacement (m)



- ❖ Modelled ETM mode shape for the 8 kHz eigenmode

Model Parameterization

- ❖ Take COMSOL numeric simulation model of 8 kHz eigenfrequency shift and fit a model

First- Order Exponential Model

$$A(1 - e^{-b_1 t}) + c_1$$

A: Total change in frequency

b_1 : Model time constant

c_1 : Frequency at room temperature

Second-Order Exponential Model

$$A(1 - 2e^{-b_2 t} + e^{-2c_2 t}) + d$$

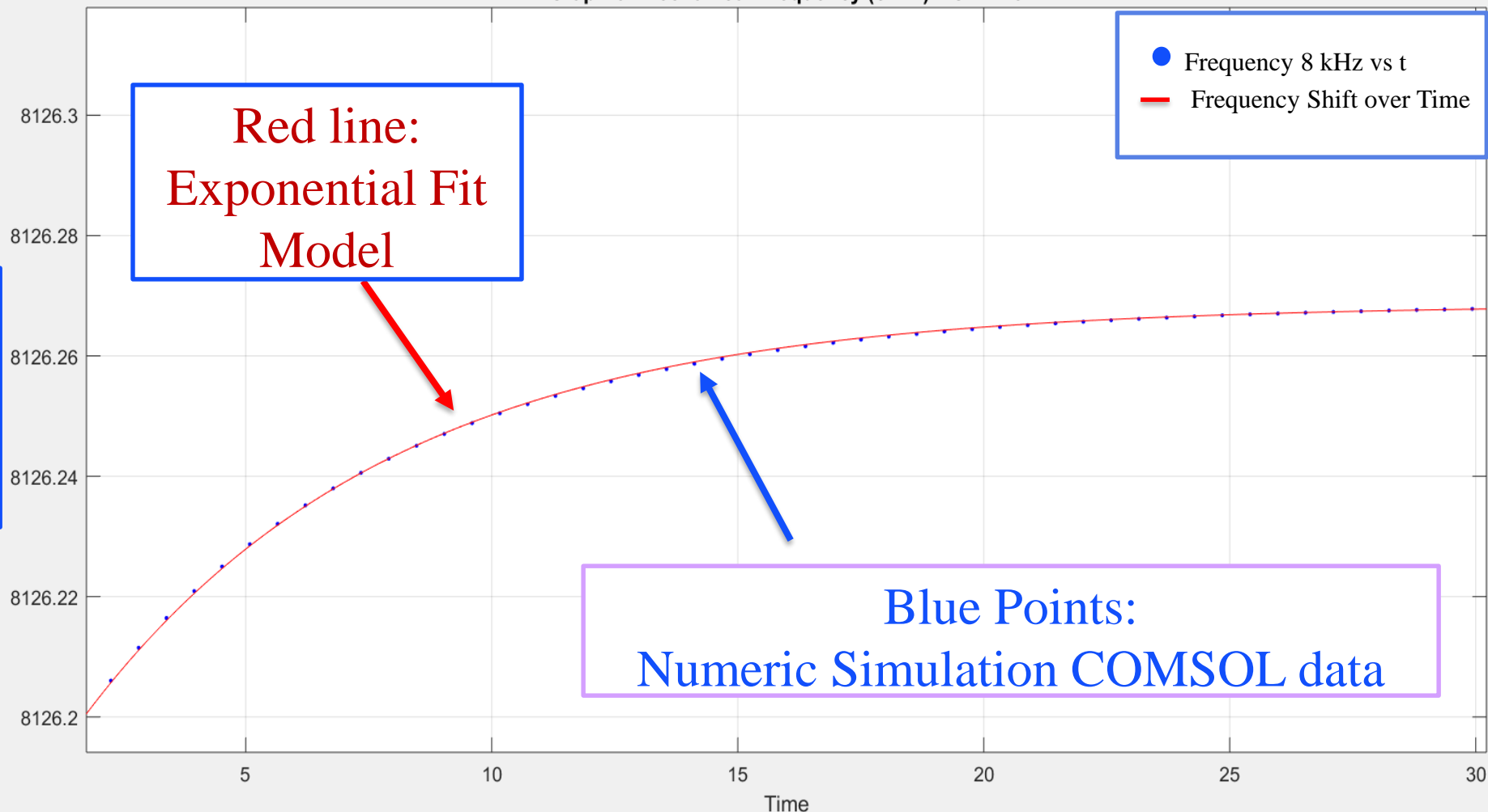
A: Total change in frequency

b_2 & c_2 : Time constants

d: Frequency at room temperature

Model Parameterization

Graph of Mechanical Frequency (8kHz) v.s. Time



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Simon Tait's Technique

Applying the exponential model and tracking eigenfrequency shift data

Experimental frequency tracking to extract coating absorption

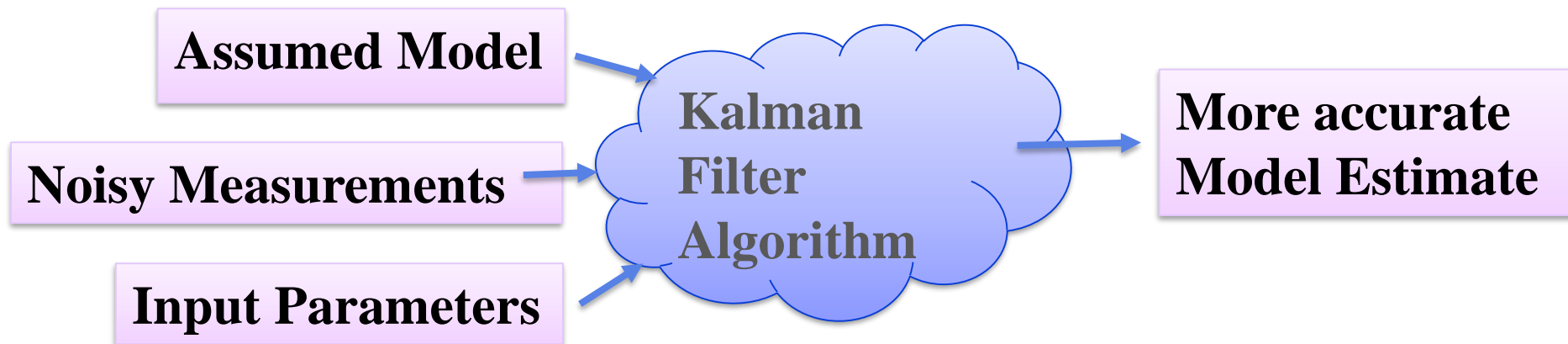
This Project's Technique

Applying the exponential model, eigenfrequency measurements, control parameters

Implement a **Kalman Filter** to extract coating absorption

Kalman Filter Theory

- ❖ Recursive algorithm
- ❖ Input: A linear model and noisy measurements
- ❖ Output: Less noisy and more accurate estimates
- ❖ Only requires current state to propagate to next time step
- ❖ Error (variances) are used to optimize estimates



- ❖ Combines inputs and measurements into model of the system \longrightarrow **minimize uncertainty in the model**

Kalman Filter Theory

- ❖ Requires state space representation of system
- ❖ Present state is dependent on the previous state:

$$\mathbf{x}_k = \mathbf{A}_k \mathbf{x}_{k-1} + \mathbf{B}_k \mathbf{u}_k + \mathbf{w}_k$$

\mathbf{A}_k : State Transition Model

\mathbf{x}_{k-1} : Previous state

\mathbf{B}_k : Input Control Model

\mathbf{u}_k : Control Vector

\mathbf{w}_k : Process noise with \mathbf{Q}_k covariance

- ❖ Observation is taken representing the true state \mathbf{x}_k :

$$\mathbf{z}_k = \mathbf{C}_k \mathbf{x}_k + \mathbf{v}_k$$

\mathbf{z}_k : Observation

\mathbf{C}_k : Observation Model

\mathbf{v}_k : Measurement noise with \mathbf{R}_k covariance

Kalman Filter Theory

Prediction:

- Predict state ahead
 $x_{k|k-1}$
- Predict the error covariance ahead
 $P_{k|k-1}$

Correction:

- Calculate Kalman Gain (Minimum Mean Square Error: minimize trace of the state error)
 K_k
- Update estimate with measurement
 y_k & $x_{k|k}$
- Update error covariance
 $P_{k|k}$

Initiate with x_{k-1} and P_{k-1}

Building a Kalman Filter

1. Understand the situation
2. Model the state process
3. Model the measurement process
4. Model the noise
5. Test the Filter
6. Refine the Filter

→ **Inputs:** Exponential Model,
Noisy Eigenfrequency
Measurements
Input Control Parameter:
Laser Power

Advantages

- Recursive nature
 - Does not depend on the history to determine the next state

Disadvantages

- Relies on an accurate model
- Depends on linearity of system

Project Initial Kalman Filter

Approach:

Normalized exponential model

$$f(t) = \mu(1 - e^{-\frac{t}{\tau}})$$

State-Space Representation:

$$f(s) = \mathcal{L}\left\{\mu(1 - e^{-\frac{t}{\tau}})\right\} = \frac{N}{s^2 + Ds} \quad N = \frac{\mu}{\tau} \quad D = \frac{1}{\tau}$$

Transfer function of the system: $\frac{f(s)}{P(s)} = sf(s) = \frac{N}{s+D}$

Inverse Laplace Transform \longrightarrow Differential Equation

$$\mathcal{L}^{-1}\{f(s)(s + D) = P(s)N\} \quad \dot{f}(t) + Df(t) = Np(t)$$

$$[f_k] = [1 - D\Delta k][f_{k-1}] + [N\Delta k][P_k]$$

State-Matrix: $A = [1 - D\Delta k]$

Input-Control Matrix: $B = [N\Delta k]$

Measurement-Matrix: $C = [1]$

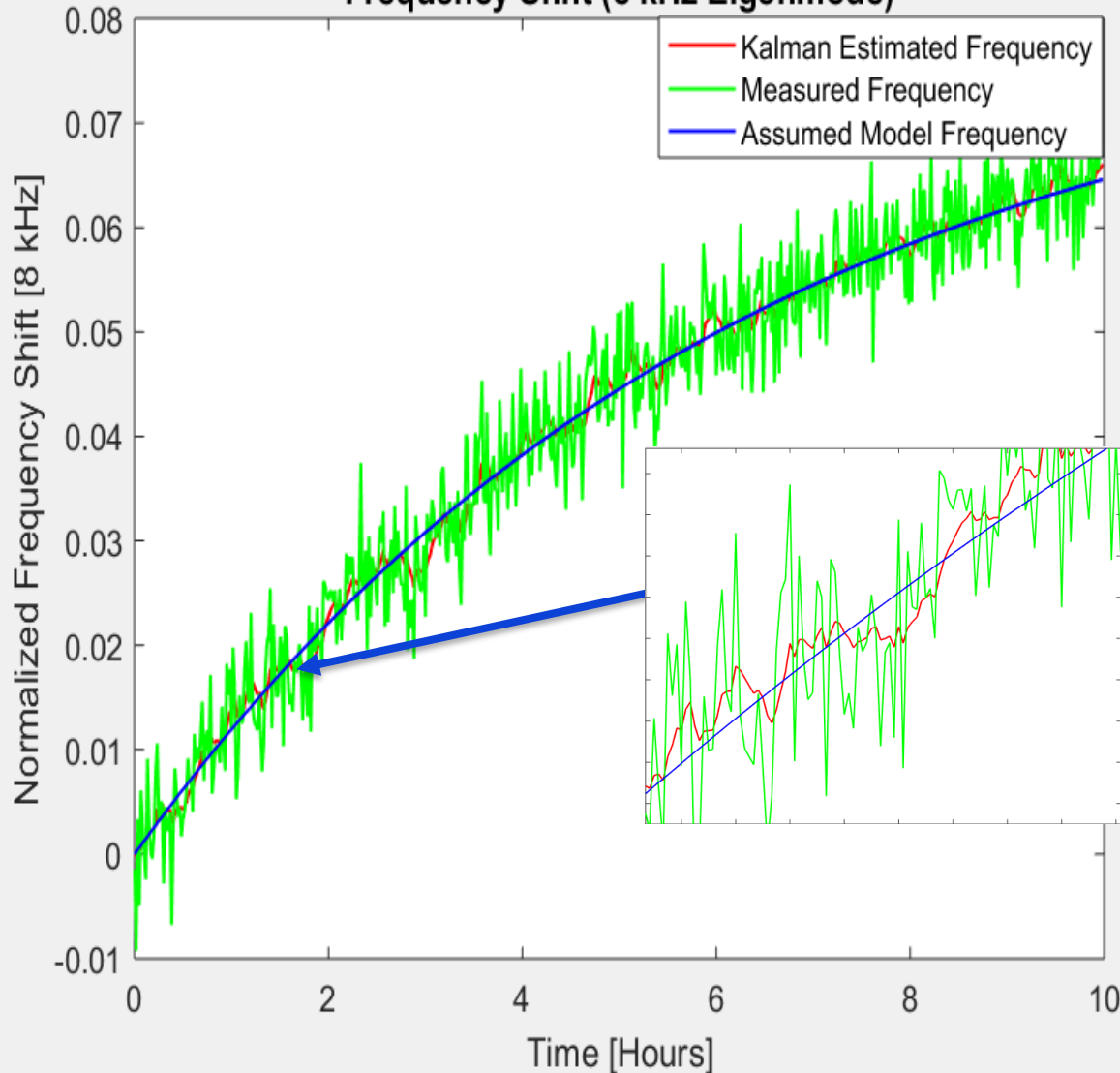
Observation Model: $z_k = f_k$

Parameters:

Gain (μ)	Time Constant (τ)
0.8114 Hz	6.289 Hrs.

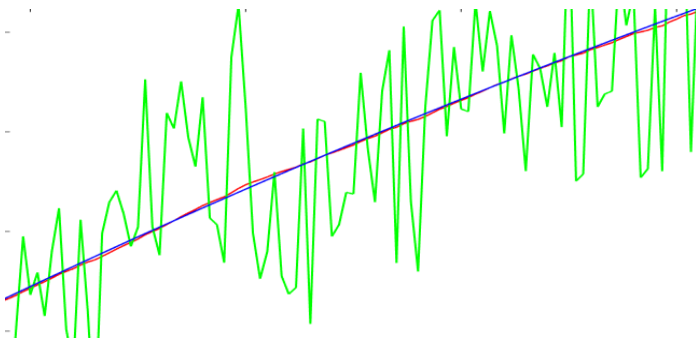
Initial Kalman Filter: Simulated Results

Frequency Shift (8 kHz Eigenmode)



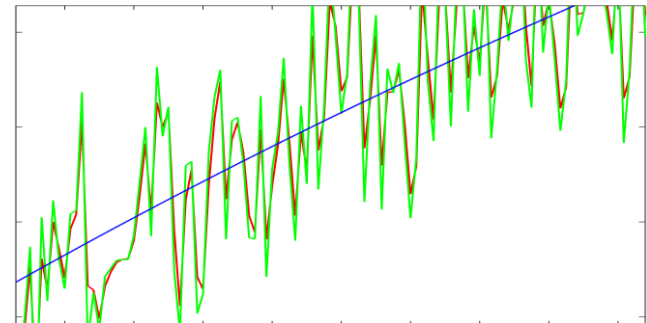
Underfit:

- $R_k \gg Q_k$ variance
- Trusts model over measurements



Overfit:

- $Q_k \gg R_k$ variance
- Trusts measurements over model



- ❖ Dependent on the change of the gain parameter:

$$P_{\alpha} = \frac{2(P_{in}k_{PRC}k_{AC})}{\pi\omega^2} e\left(-\frac{2r}{\omega^2}\right) \frac{1}{\alpha_c}$$

P_{α} : Power absorbed by the optic

ω : beam radius of the incident Gaussian light source (6.2 cm)

r : distance from the center of the beam

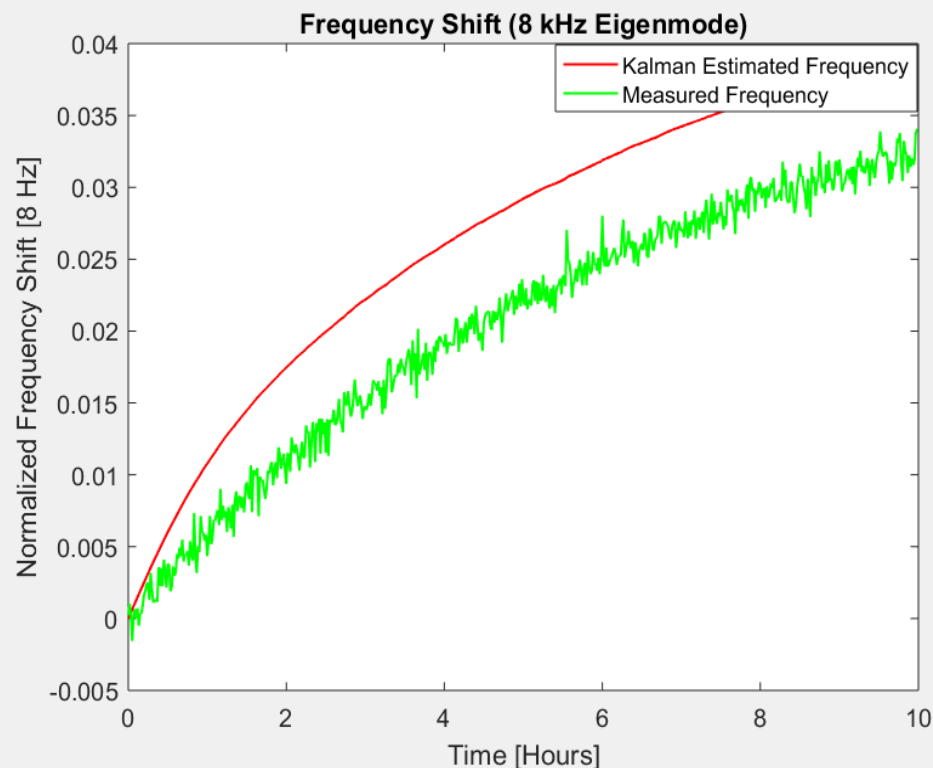
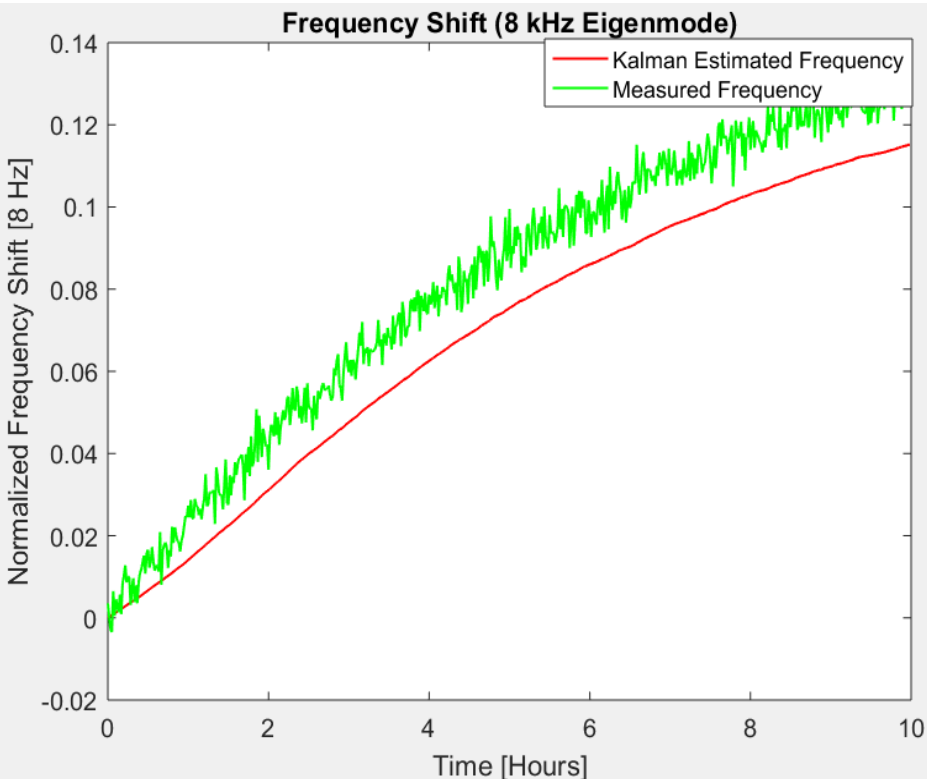
P_{in} : power input into the interferometer

k_{PRC} : gain of the power recycling cavity

k_{AC} : gain from the arm cavity

α_c : coating absorption

Problem Extracting Coating Absorption



Simulated Lock state where the noisy measurements have factor of 2 multiplied to input laser power

Simulated Lock state where the noisy measurements have factor of 0.5 multiplied to input laser power

- ❖ Coating absorption $\alpha_{coating}$ is proportional to the gain parameter in the state space model
- ❖ The gain is not linearly related to the system
- ❖ Kalman Filters function with linear systems
- ❖ Options:
 - ❖ Linearize the parameter to the system
 - ❖ Create a nested Kalman Filter that updates the change in gain

Nested Kalman Filter Approach

- ❖ Gain directly related to the input-control model B
- ❖ Update B and the process covariance at the end of each lock state

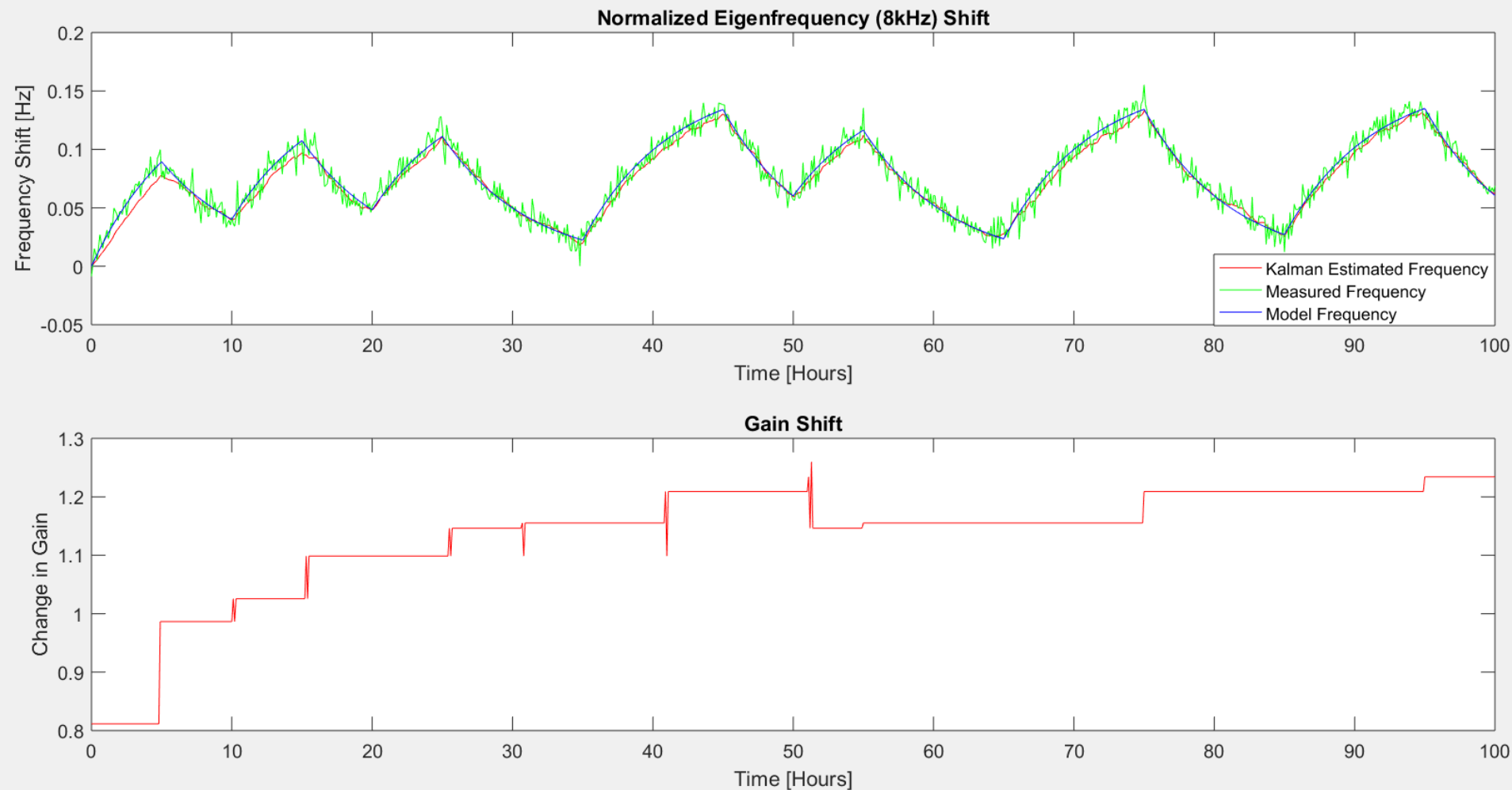
$$\text{Process Covariance: } Q_k = Bw^2B'$$

- ❖ Measure average residuals during lock time frame between measurement data and Kalman Estimate

$$B_{updated} = B_{initial} + (\textit{average residual}) * B_{initial}$$



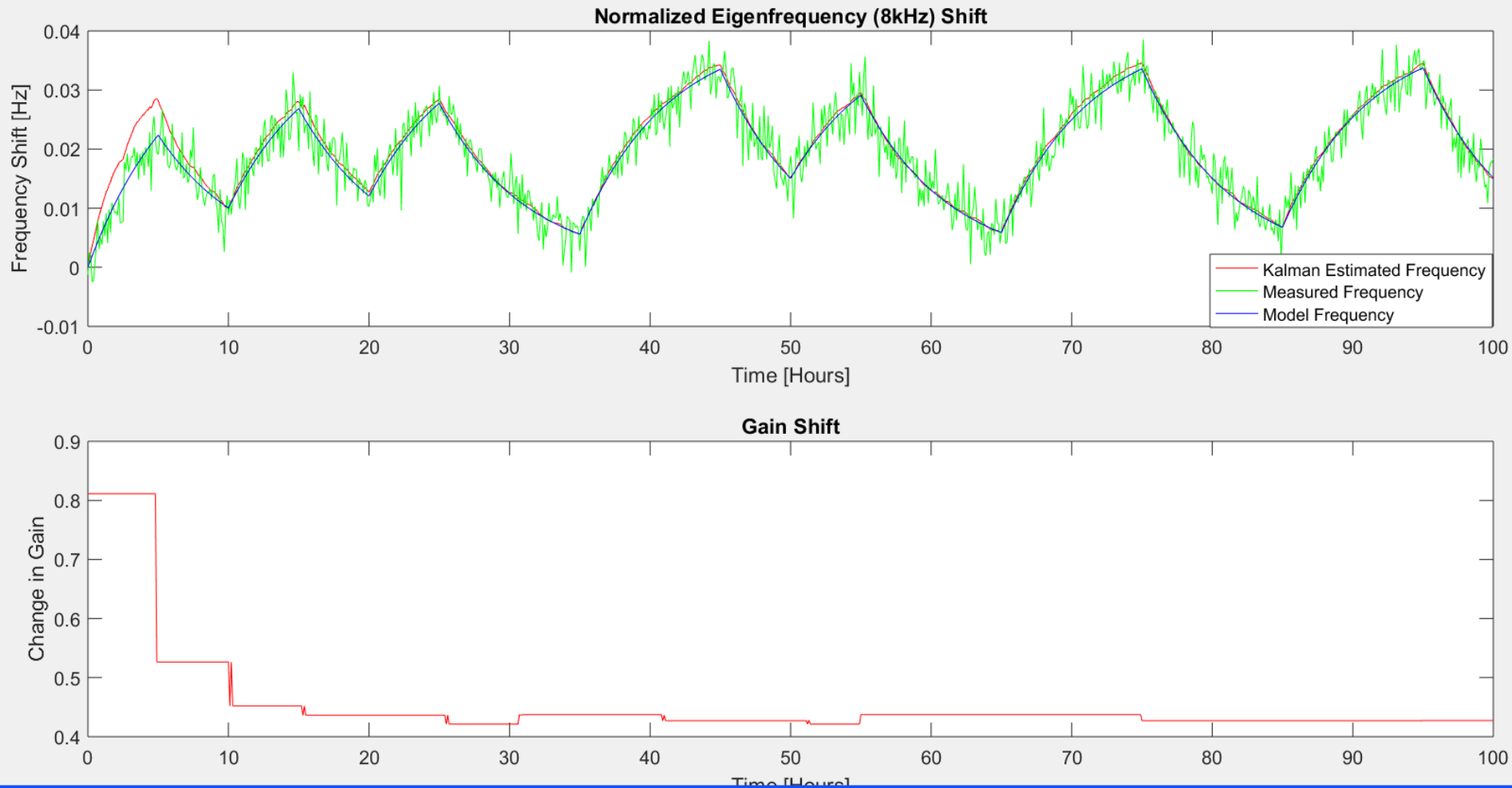
Nested Kalman Filter: Overfit Simulation



Simulated period of locked and unlocked states where the noisy measurements have factor >1 applied to the input laser power and the nested Kalman Filter updates in a bias towards the noisy behavior



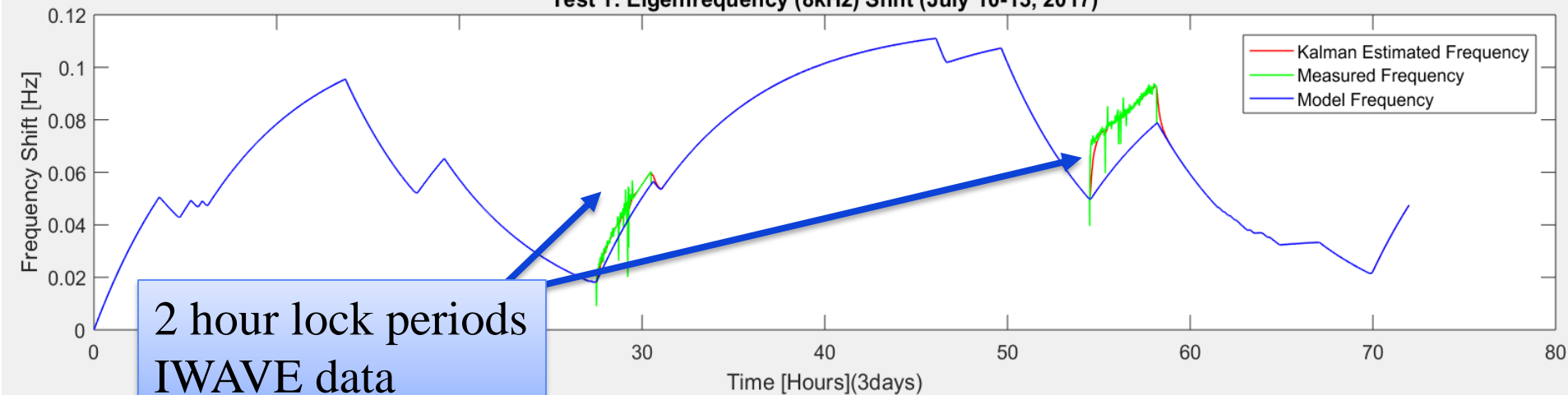
Nested Kalman Filter: Underfit Simulation



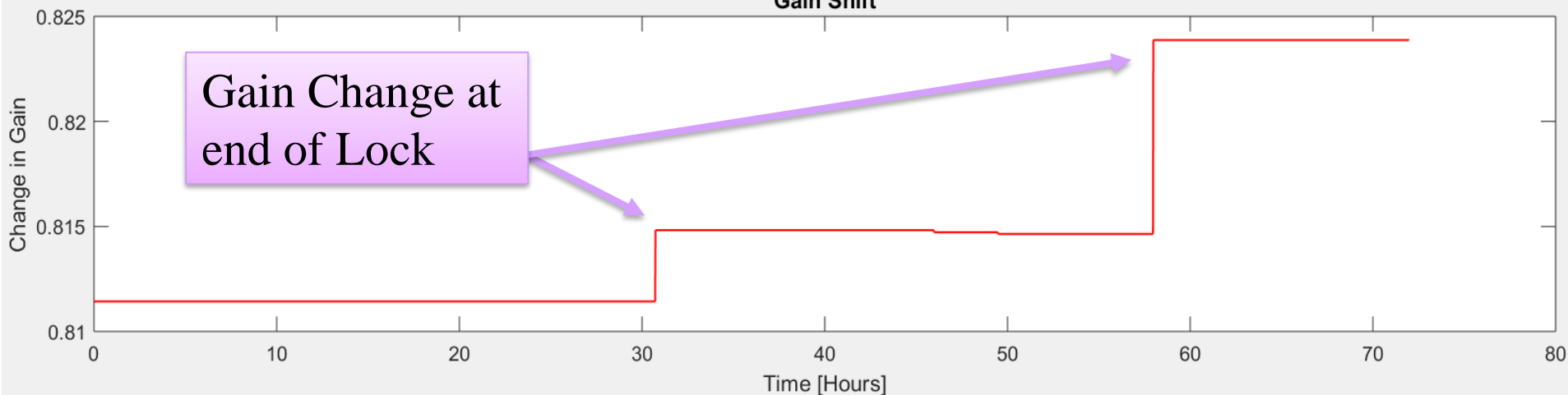
Simulated period of locked and unlocked states where noisy measurements have a factor < 1 applied to the input laser power and nested Kalman Filter updates itself in a bias towards the noise's behavior

Results: Testing Data from July 2017

Test 1: Eigenfrequency (8kHz) Shift (July 10-13, 2017)

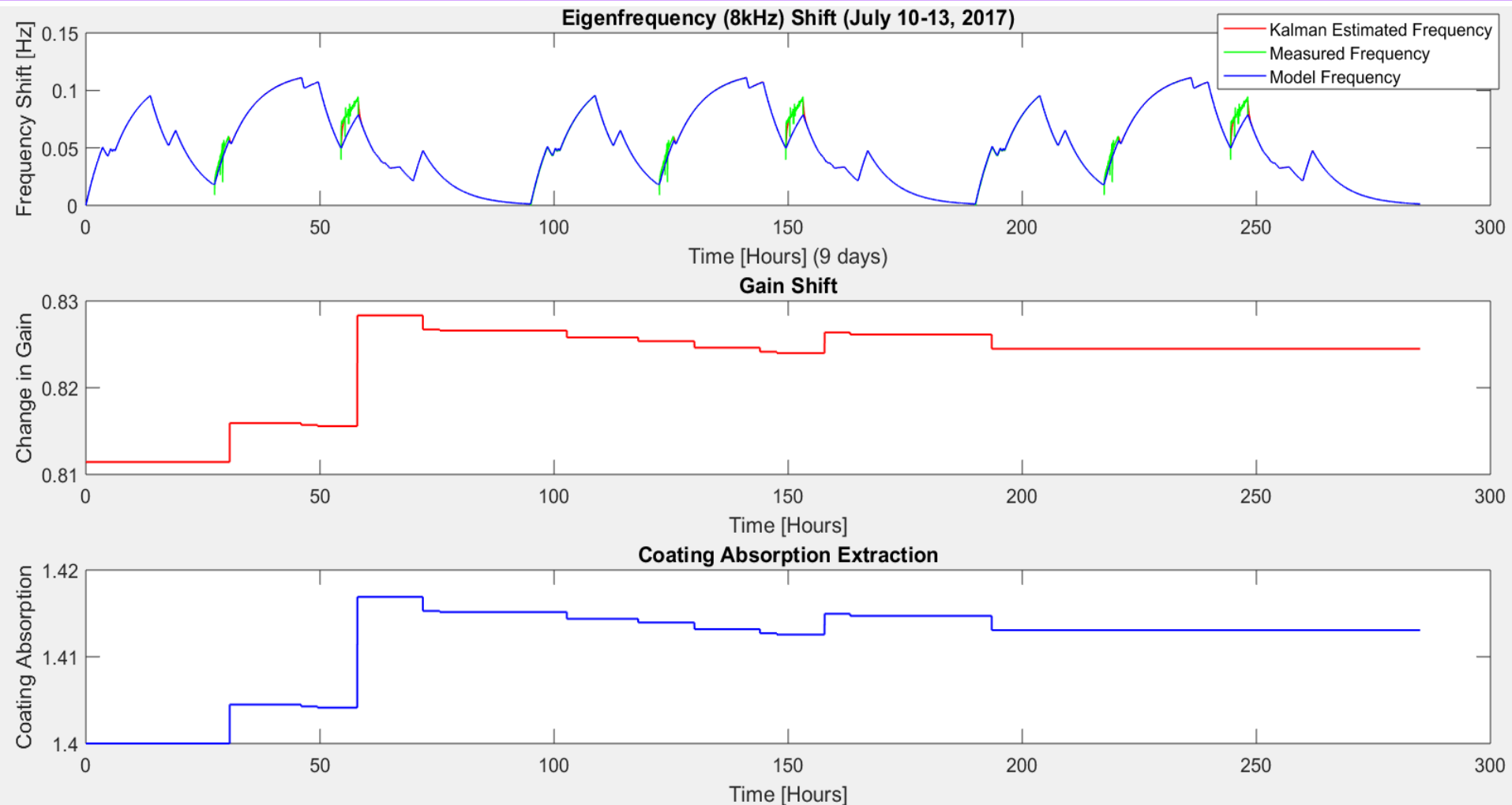


Gain Shift



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Results: Testing Data from July 2017



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Conclusions

Takeaways:

- Several other factors including ambient temperature need to be incorporated to improve the model of the system
- Kalman Filters provide useful monitors and more accurate models of a system

Future Work:

- Run the filter over longer periods of time with IWAVE data to improve absorption estimate
- Improve and change the model to incorporate other parameters, remove outliers from noisy IWAVE frequency data
- Implement the Kalman Filter as a real-time LLO monitoring system to further improve absorption estimation and other parameter estimations
- Combine this time evolution (1 eigenmode over time) behavior with spatial evolution (several eigenmodes) behavior to create a stronger mirror degradation monitor

Acknowledgements

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References

- [1] S. Tait, *An Instantaneous Absorption Estimate of aLIGO Test Masses*. pgs. 1-26 (2018).
- [2] H. Wang, C. Blair, M. Dovale Alvarez, A. Brooks, M. F. Kasprzack, J. Ramette, P. M. Meyers, S. Kaufer, B. Oreilly, C. M. Mow-Lowry, A. Freise, *Thermal modelling of Advanced LIGO test masses*. Class LIGO Document. 26 Apr (2017).
- [3] S.C. Tait, I.W. Martin, C. Blair, R. Jones Z. Tornasi, A. Bell, J. Steinlechner, J. Hough S. Rowan. *Optical Absorption of Ion Plated Coatings and Instantaneous absorption at LLO*.
[https://dcc.ligo.org/DocDB/0150/G1800531/001/LVC2018.pdf\(2018\)](https://dcc.ligo.org/DocDB/0150/G1800531/001/LVC2018.pdf(2018)).
- [4] S. Tait, *Thermal Modelling*. LIGO COMSOL models.
- [5] G. Valdes. *Data Analysis Techniques for LIGO Detector Characterization*. University of Texas at San Antonio (2017).
- [6] A. Brooks, *Seminar on Kalman Filters*. (2014).

Any Questions?

In-Depth Kalman Filter Analysis

A_k represents the state transition model used to the previous state $\mathbf{x}_{k|k-1}$ and B_k represents the input-control model. The input-control model is applied to the control-vector \mathbf{u}_k and the state matrix is applied to the state-vector $\mathbf{x}_{k|k-1}$. The process noise is represented by which in this case is a univariate normal distribution with covariance Q_k .

$$\mathbf{x}_{k|k-1} = A_k \mathbf{x}_{k|k-1} + B_k \mathbf{u}_k$$

$$P_{k|k-1} = A_k P_{k-1|k-1} A_k^T + Q_k$$

Calculating Kalman Gain:

$$S_k = C_k P_{k|k-1} C_k^T + R_k$$

$$K_k = P_{k|k-1} C_k^T S_k^{-1}$$

Updating Estimate with Measurement and updating error covariance

$$\mathbf{y}_k = \mathbf{z}_k - C_k^T \mathbf{x}_{k|k-1}$$

$$\mathbf{x}_{k|k} = \mathbf{x}_{k|k-1} + K_k \mathbf{y}_k$$

$$P_{k|k} = (I - K_k C_k) P_{k|k-1}$$