

Implications of dedicated seismometer measurements on Newtonian-noise cancellation for Advanced LIGO

M. Coughlin

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Newtonian Noise

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Newtonian noise or "gravity-gradient noise" arises from fluctuating seismic fields and atmospheric disturbances such as pressure and temperature fluctuations leading to a direct gravitational force on the test masses.

- LIGO Hanford \rightarrow Caltech/Hanford/GSSI.
- Virgo WEB \rightarrow Poland/GSSI.
- Homestake (3D array) \rightarrow Caltech/Minnesota.
- Seismic NN modeling \rightarrow Nikhef/GSSI.
- Atmospheric NN modeling \rightarrow APC/GSSI.

I am only worried about the seismic field portion in the following.









LHO Corner Station Array

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Seismic Velocity Histogram

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Coherence of the array (all possible pairs)

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Tiltmeters?!?

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- Single horizontal sensor beneath test mass: Seismic sensors are sensitive to Rayleigh waves... and Love (shear) waves (non-NN contributing).
- Multiple vertical sensors: Requires an array some distance from test-mass with relatively weak correction with test-mass acceleration
- Single tiltmeter beneath test mass: Not sensitive to Love waves so... no problem (theoretically).



Optimal Tiltmeter Subtraction

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Actual Tiltmeter Subtraction

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Residual Spectrum / Original Spectrum





Wiener Filter Bode Plots

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Tiltmeter-DARM Transfer Function

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Other to-dos and conclusions

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Highlights:

- We used dedicated measurements at the LIGO Hanford site to predict NN cancellation levels.
- We showed how we were able to achieve significant subtraction in line with expectations based on correlation measurements.
- We showed that significant subtraction is achievable with only a few seismometers.

Future extensions:

- Calculation of optimized arrays in inhomogeneous seismic fields without constraints on seismometer locations.
- Calculate the best sensor locations from correlation measurements (to reach ultimate cancellation limits for a given number of sensors)

• Devise the strategy to optimally pick sensors in a large underground array.