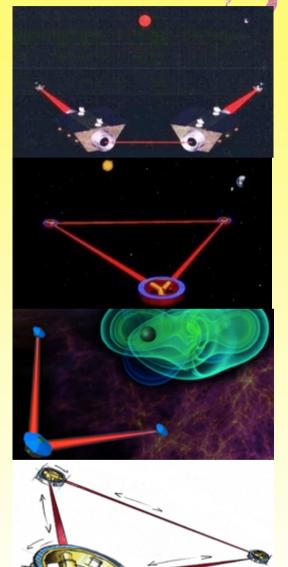


21 Years after the First LISA Symposium at RAL 1996



LISA: A Mature Concept

- M3 proposal for 4 S/C ESA/NASA collaborative mission in 1993
- LISA selected as ESA Cornerstone in 1995
- 3 S/C ESA/NASA LISA appears in 1997
- Joint ESA-NASA Mission Formulation study 2005-2011
- Reformulation 2012-13 as
 ESA-led eLISA (evolving LISA)
- Now back to 3-arm LISA with NASA



But then in March 2011... go > Advanced search nature.com Search This site + Publications A-Z index Browse by subject ORIGINAL My account CONTINUOUS RESEARCH NEW FROM ONLINE ONLINE NATURE PUBLISHING GROUP: \sim E-alert sign up SCIENTIFIC REPORTS 2 **RSS** feed OPEN MULTI Now Accepting Submissions! PEER ACCESS DISCIPLINARY REVIEWED Subscribe

Published online 22 March 2011 | Nature 471, 421 (2011) | doi:10.1038/471421a

News

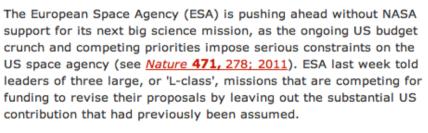
Europe makes do without NASA

US budget crisis forces European Space Agency to abandon plans for joint mission.

Stories by Reywords

- European Space Agency
- L-Class missions
- <u>LISA</u>
- <u>IXO</u>
- ESJM-Laplace

This article elsewhere



"The decision was made very reluctantly," says David Southwood, director of science and robotic exploration at ESA. "NASA could not meet our timetable to launch."

22 April 2011

 China hopes research centre can quell food-safety fears 22 April 2011

Related stories

- US Mars mission takes pole position
 08 March 2011
- ESA on countdown to flagship mission
 selection



LISA Redefinition Study

- Redesign for ESA-led mission
- Cost-cap for ESA cost at 850 M€

 Build on LISA Pathfinder hardware
 Shorter arms, smaller telescopes, simpler orbits, less mass
 Use cheaper launcher

Cut one arm off !

\rightarrow Descoping is an art!

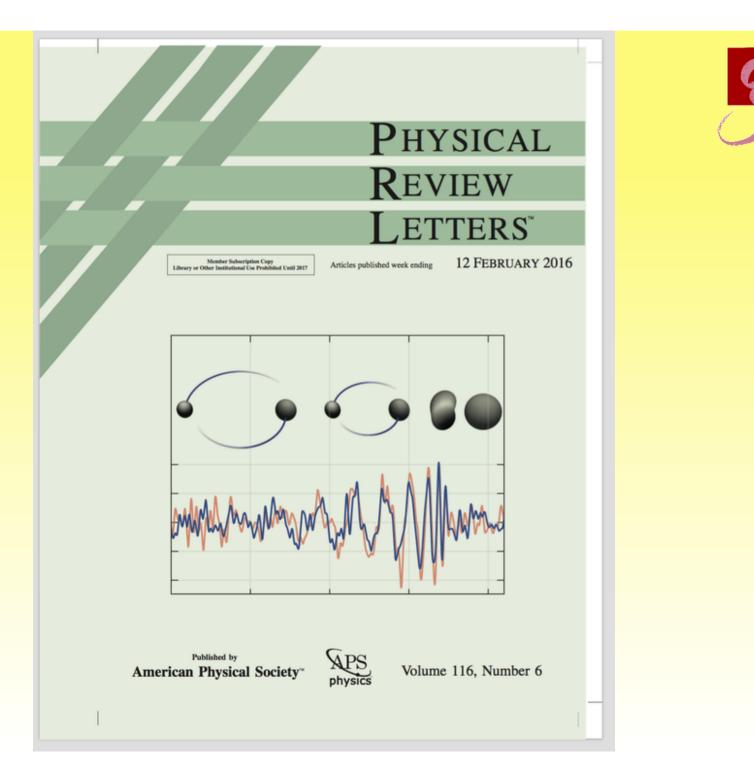




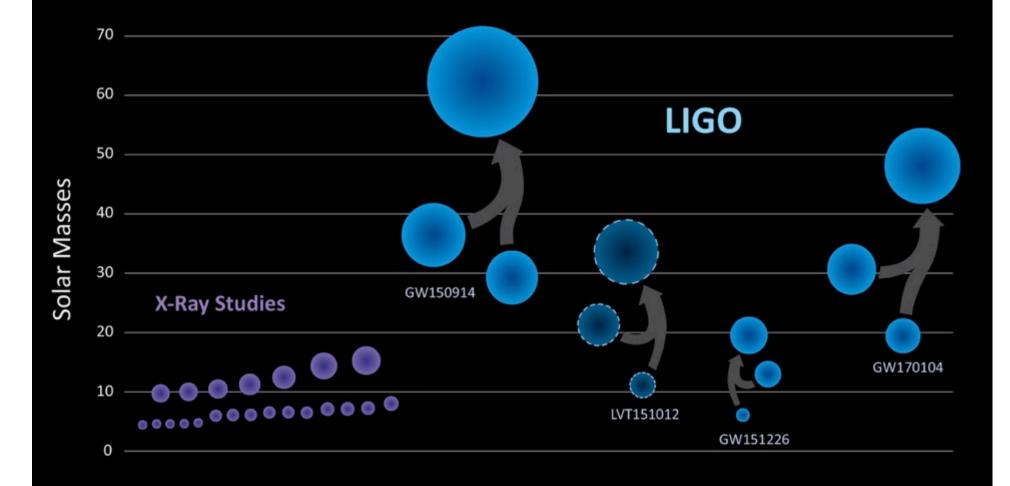
Then came 2015/2016....



• And two things happened!



Black Holes of Known Mass



Credit: LIGO

LISA Pathfinder



• Testing LISA technology in space!

Cesa

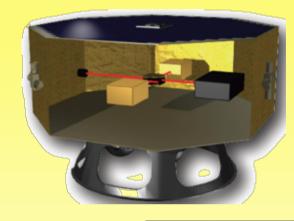
15Anfinder



First Proposed in 1998 as ELITE

G

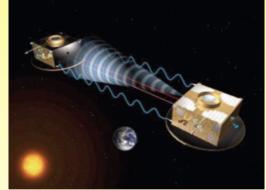
• European Llsa TEchnology Satellite

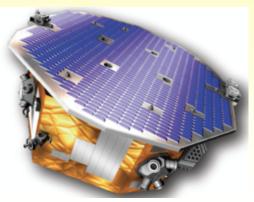


Renamed to SMART-2 in 2000

 Tech demo for LISA and Darwin
 Launch date 2006

Descoped to LISA Pathfinder
 – Darwin demo cancelled







100 Years since GR Publication: Dec. 2, 2015

Countdown to LPF Launch

LPF has launched!

LISA Pathfinder Mission Timeline

LPF begins Apogee Raising Manouevers LPF reaches Lagrange Point L1 Operations begin with IOCR on 03

LPF journeys to Lagrange Point L1

LPF separates from Launcher

LPF launch on 02-Dec-2015 at 04:15 UTC Propulsion Module Separation

LPF Power Up for Launch Countdown

Test Mass 1 Release 16-Feb-2016 at 12:00 UTC

Test Mass 2 Release 15-Feb-2016 at 12:00 UTC

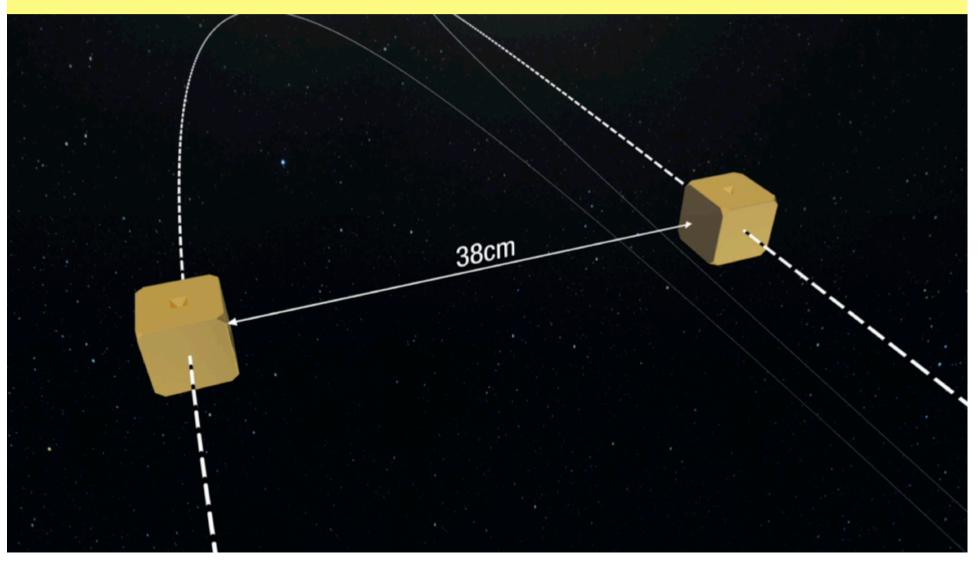
Dez Feb Mrz Apr Feb Dez Лrz



The Stillest Place in the Universe!



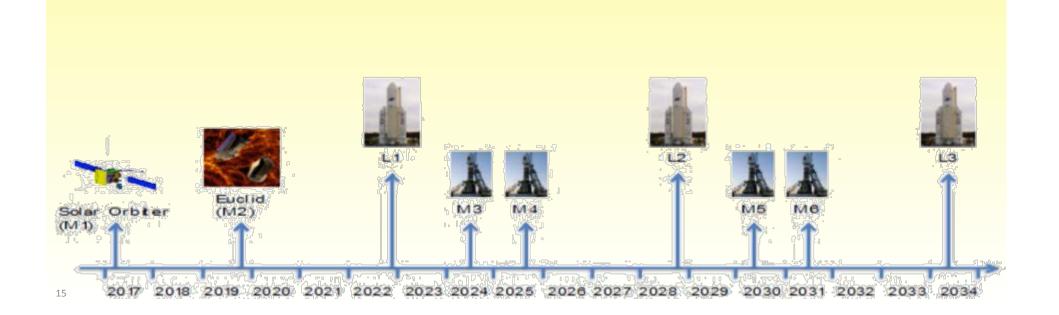
• More sensitive than the weight of a virus!



ESA L2 and L3 Missions



• Call for Mission Concepts fall 2016



NASA is back in LISA!



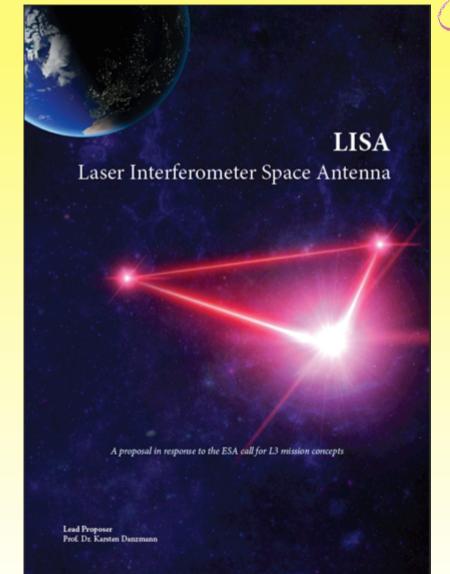
-

A Midterm Assessment

The National Academies of SCIENCES • ENGINEERING • MEDICINE

LISA Mission Concept Document

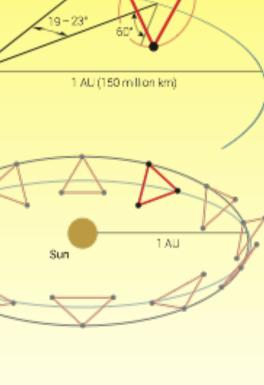
- Submitted on January 13th, 2017
- The LISA Consortium: 12 EU Member States plus the US !



https://www.lisamission.org/proposal/LISA.pdf

Mission Profile and Orbit

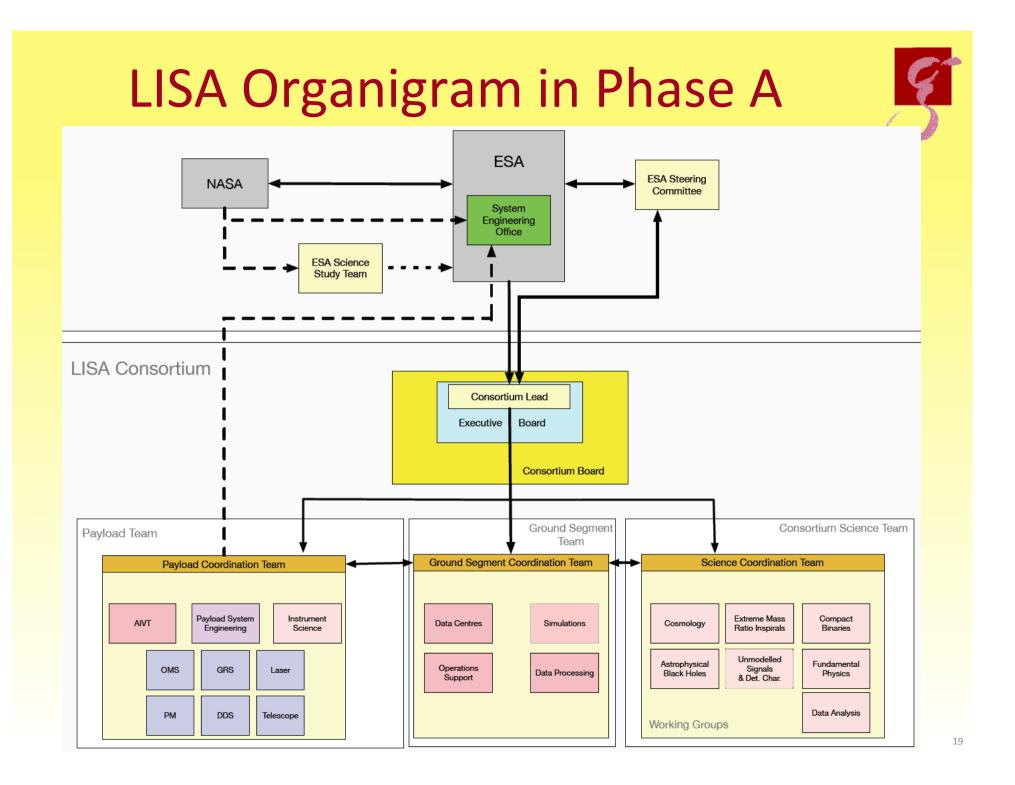
- Three arms of 2.5 Million km
- 2W lasers
- 30 cm telescopes
- Breathing angles ± 1 deg
- Doppler shifts ± 5 MHz
- Launch on dedicated Ariane 6.4
 - Transfer time ~400 days
 - Direct escape V_{∞} = 260 m/s
 - Propulsion module and S/C composite

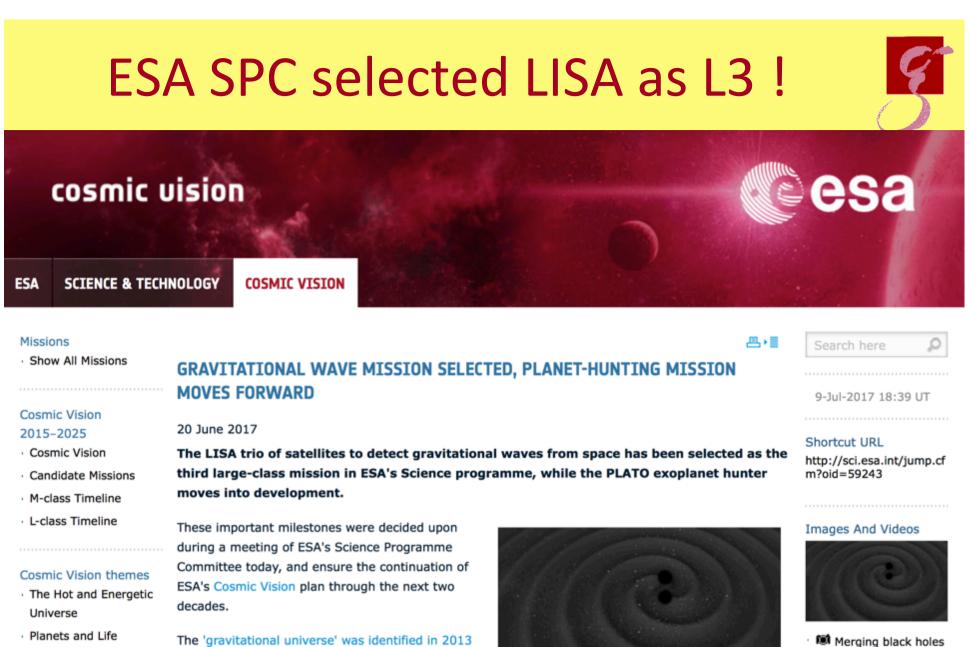


2.5 million

Earth

Sun





- as the theme for the third large-class mission, L3, searching for ripples in the fabric of spacetime created by celestial objects with very strong gravity,

The Solar System

Fundamental Laws

The Universe

Searching for

exoplanetary systems



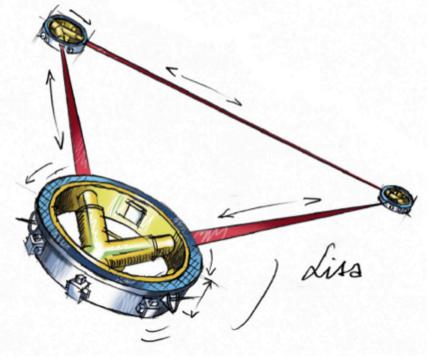


LISA Study Introduction

Systems

Session 1 ESTEC, 08-03-2017

Prepared by the CDF* Team



(*) ESTEC Concurrent Design Facility



Presentation of First Study Phase



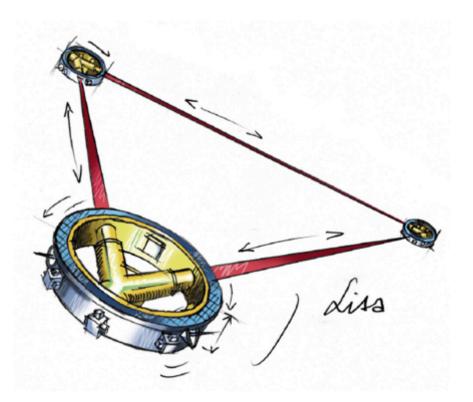
LISA

Welcome

Internal Final Presentation ESTEC, 05-05-2017

Prepared by the CDF* Team

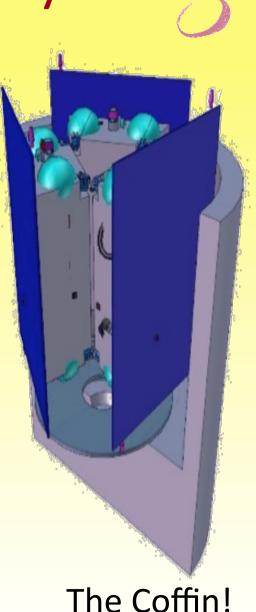
(*) ESTEC Concurrent Design Facility





LISA Phase 0 System Study

- Feasible technical baseline for LISA
- Dedicated launch with Ariane 6.4
 - Transfer 15 mo + commissioning 9 mo
 - 4 yrs of science operations, extended mission to 10 yrs
- Electric propulsion for orbit transfer, cold gas for micropropulsion
- Payload derived from previous studies, some open trade-offs
- S/C mass 1860 kg, incl. 480 kg P/L, no jettisonable propulsion module
- Power 2.5 kW
- Healthy mass margins (> 1000 kg growth potential)



Launch and transfer



Separation of the stack right after launch Separate trajectory for each S/C to final orbit

Sun

1 Alf (190 million land



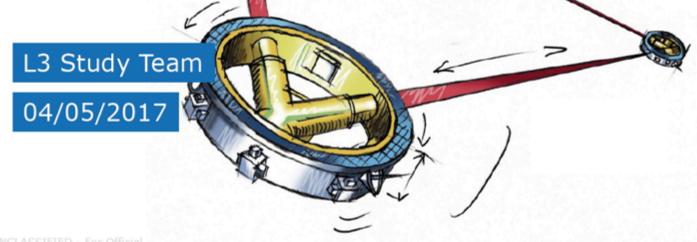
Launch in stacked configuration Direct injection into escape trajectory







L3/LISA ESA-NASA Technical Interface Meeting #1



A UNCEASSIFIED - FOI UIIGIA Sea

Bangreen Speen Ageneri



L3/LISA Phase 0 Planning #1

ESOC

24 May 2017

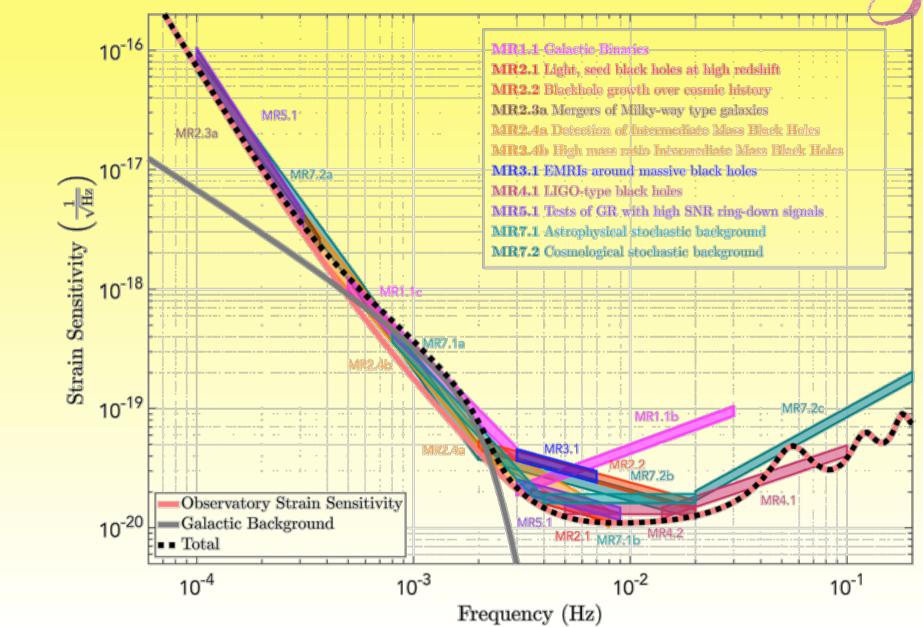
ESA Study Team

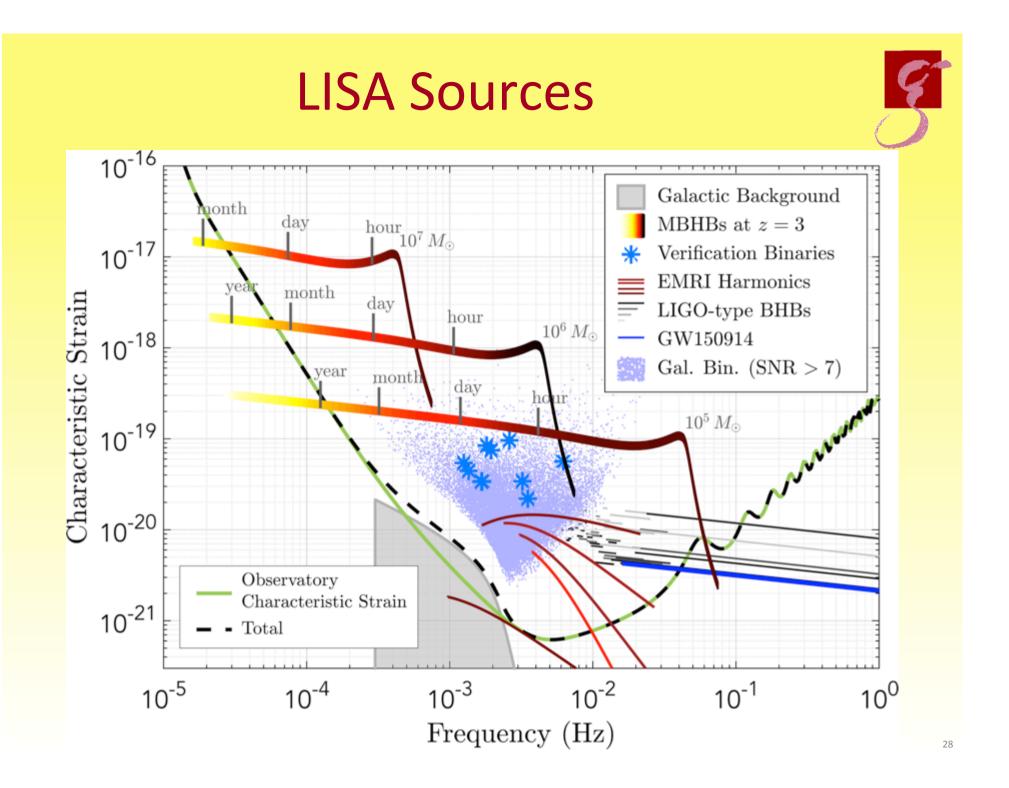
ESA-L3-EST-PL-MIN-001

European Space Agency

LISA Requirements







LISA for Astrophysics, Cosmology, and Fundamental Physics



Massive Black Holes (10^4 to 10^8 M_{\odot})

- When did the first Black Holes appear in pre-galactic halos and what is their mass and spin?
- How did Black Holes form, assemble and evolve from cosmic dawn to present time, due to accretion and mergers?
- What role did Black Holes play in re-ionisation, galaxy evolution and structure formation?
- What is the precise luminosity distance to loud standard siren black hole binaries?
- What is the distance redshift relation and the evolution history of the universe?
- Does the Graviton have mass?

Extreme Mass Ratio Inspirals, EMRIs (1 to 10 M $_{\odot}$ into 10⁴ to 5 x 10⁶ M $_{\odot}$)

- How is the stellar dynamics in dense galactic nuclei?
- How does dynamical relaxation and mass segregation work in dense galactic nuclei?
- What is the occupation fraction of black holes in low-mass galaxies?
- How large are deviations from Kerr Metric, and what new physics causes them?
- Are there horizonless objects like boson stars or gravastars?
- Are alternatives to GR viable, like Chern-Simons or scalar tensor theories or braneworld scenarios?

Ultra-Compact Binaries in Milky Way

- What is the explosion mechanism of type Ia supernovae?
- What is the formation and merger rate of compact binaries?
- What is the endpoint of stellar evolution?

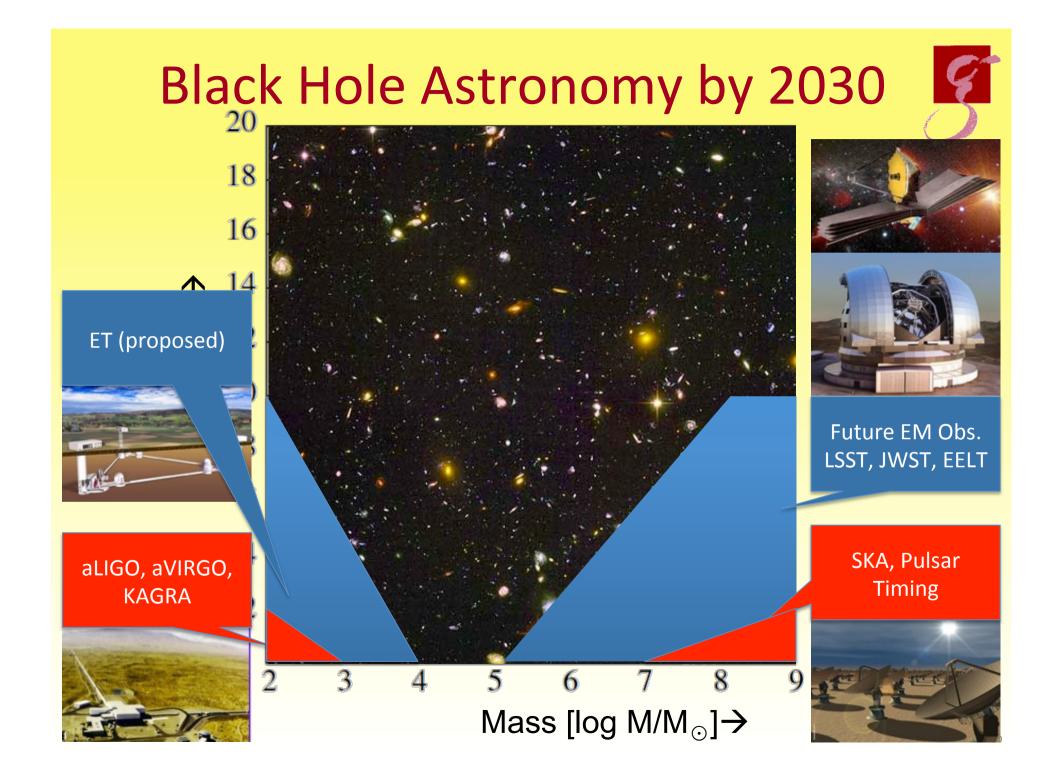
Stochastic Signals

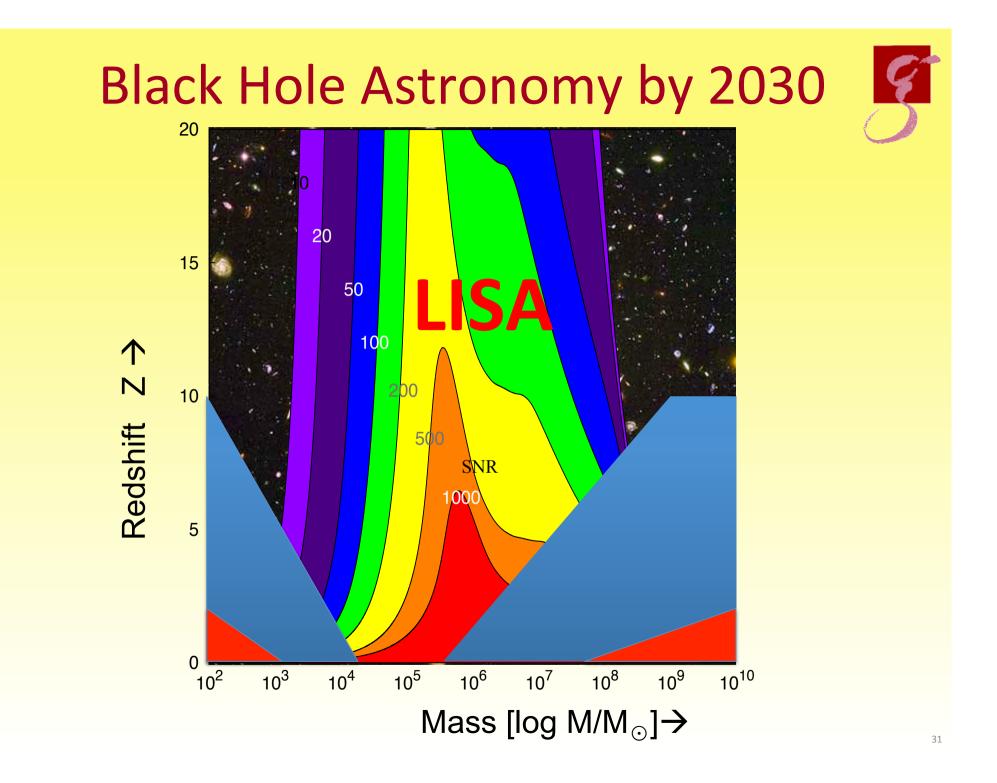
- Directly probe Planck scale epoch at 1 TeV to 1000 TeV before decoupling of microwave background
- Were there phase transitions and of which order?
- Probe Higgs field self coupling and potential, and search for supersymmetry.
- Are there warped sub-millimetre extra-dimensions?
- Can we see braneworld scenarios with reheating temperatures in the TeV range?
- Do topological defects like Cosmic Strings exist?



The Unknown !

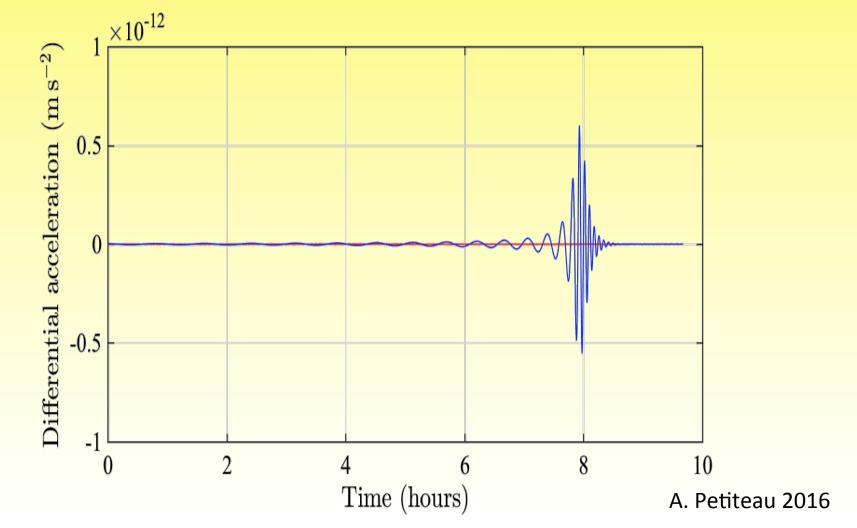






Black Hole Mergers far above Noise

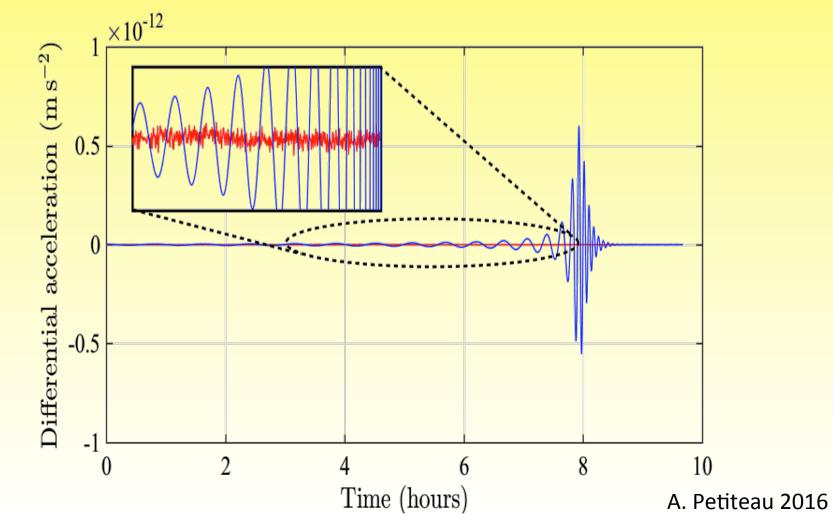
- $10^5 M_{\odot} BH$ binary merger at z=5
- In Red: Pathfinder instrumental noise



32

Black Hole Merger far above Noise

- $10^5 M_{\odot}$ BH binary merger at z=5
- In Red: Pathfinder instrumental noise



33

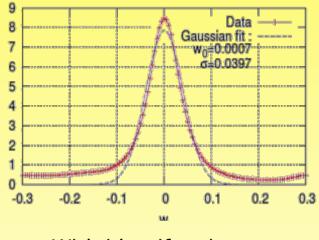
All Binary Black Holes cross LISA band: C **Trace Galaxy Mergers** 20 black hole - black hole mergers Ο 18 space based 16 gravitational wave observatory 14 -Redshift (z) 12 10 future EM probes 8 6 4 -10 20 100 200 2 1000 0 3 7 8 2 5 6 9 4 log(M/M_☉) 34

Cosmology with Standard Sirens

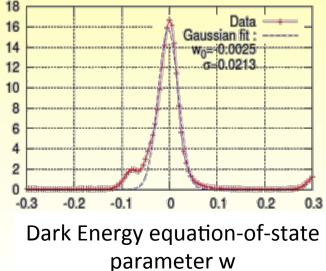


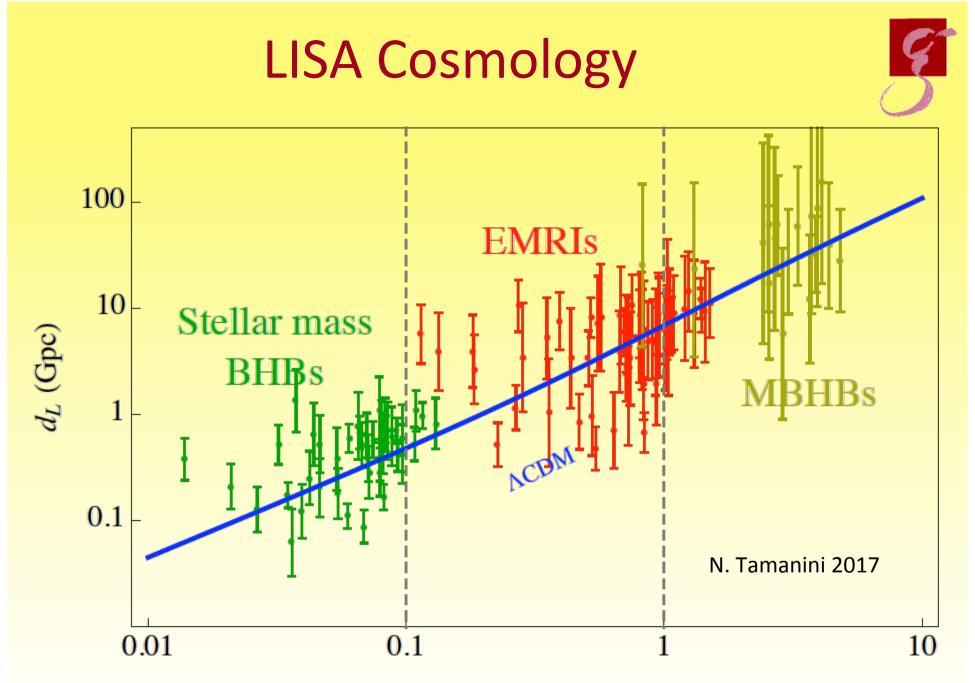
- With luminosity distances, LISA gives accurate and independent measurements of H₀ and w.
- EMRIs, without EM counterparts:
- Hubble const. H₀ to ±0.4% = ±0.3 km s⁻¹ Mpc⁻¹ after 20 EMRI detections: ~3 months LISA (MacLeod & Hogan, PRD, 2008; SDSS)
- Compare WMAP: ±1.2 km s⁻¹ Mpc⁻¹.
- MBH mergers out to z = 3, no EM counterparts:
- Dark energy equation of state parameter w to ±2-4% in 3 years (Petiteau et al, ApJ, 2011; Millennium).
- Compare EUCLID: ±2%.

No identifications (b) without electromagnetic counterpart



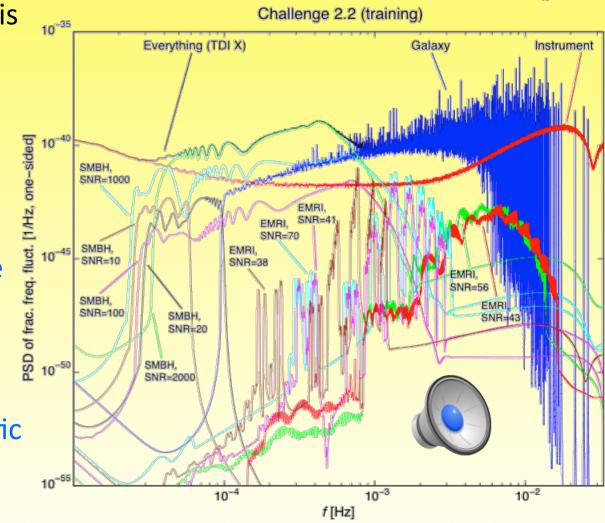






LISA Mock Data Challenge (< 2010)

- Practicing data analysis on synthetic data
- Blind international challenge
- Full LISA data stream
 - Instrumental noise
 - 4 MBH events
 - 5 EMRI events
 - 26.1 million Galactic binaries



Effective data analysis algorithms are needed!

The new LISA Data Challenge (LDC)

- Resurrecting data challenges
- Project hosted under git-lab:
 - https://gitlab.in2p3.fr/stas/MLDC (sign up is required)
- Project oriented:
 - Each data set aims at particular data analysis problem
- Ultimate goal:
 - Build a robust data analysis pipeline for the LISA mission.
- Web-page will be open shortly for everyone to sign up for the challenge and download the simulated data set.



000

Welcome on the MLDC website

In support of the Laser Interferometer Space Antenna (LISA) gravitational wave observatory, a Web site and application for the Mock LISA Data Challenge is provided.

You can find information of the MLDC on the MLDC Round page. Release data of the LISACode simulation code can be find at Object list. A request form can be find at Query. Contact information are at Contact.

Data Processing Center (DPC)

MLDC Web application is part of the DPC of LISA. The goal of the DPC is to provide tools for code developpement and execution for the data analysis. The platform is reachable at LISA DPC.

News

New dataset is provided...

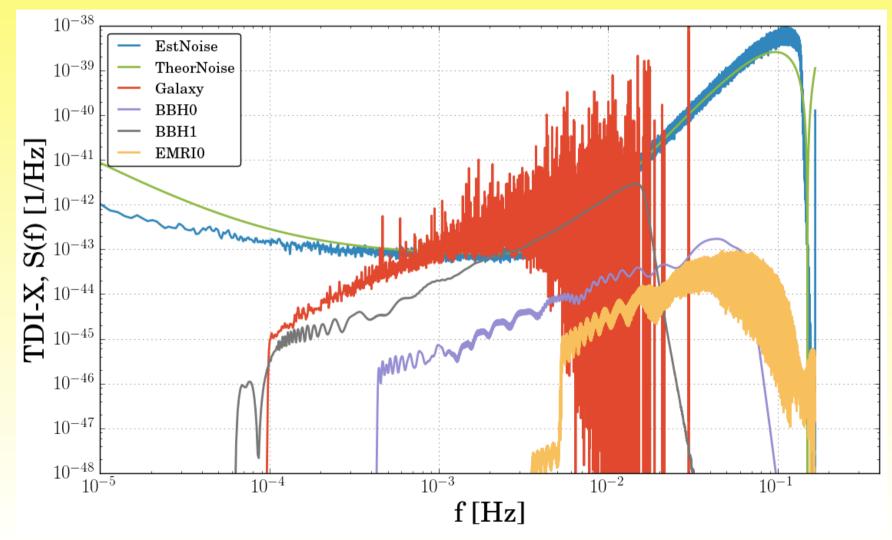
Website 1.0 is ready

Work in progress: missing query form, formulary, login registration.

First Data Release

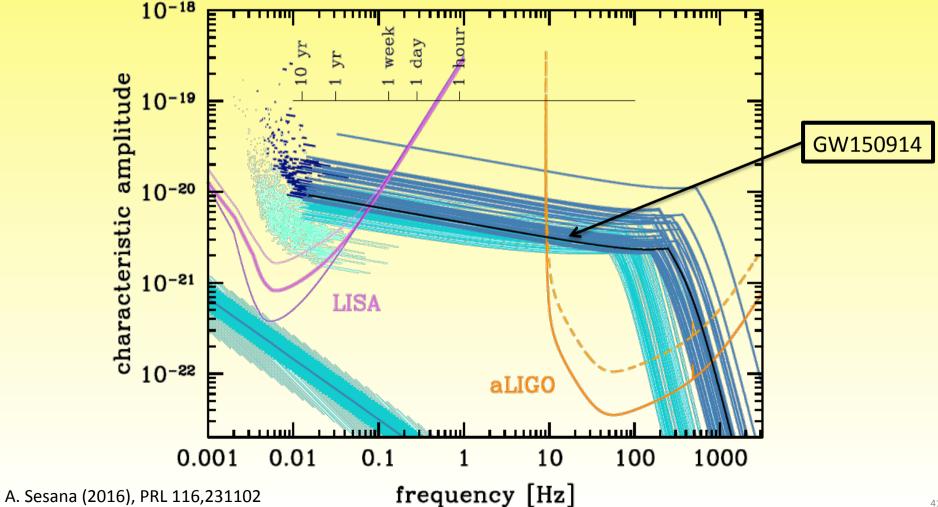


• First data release at the end of November



LISA: LIGO Event Predicted 10 Years in Advance!





ESA L2 and L3 Missions

- Call for Mission Concepts fall 2016
- Decision on L3 Adoption 2021
- Launch of L2 in 2028
- Launch of L3 in 2034
- LISA shall be ready for an early launch!

