



# Gravitational waves

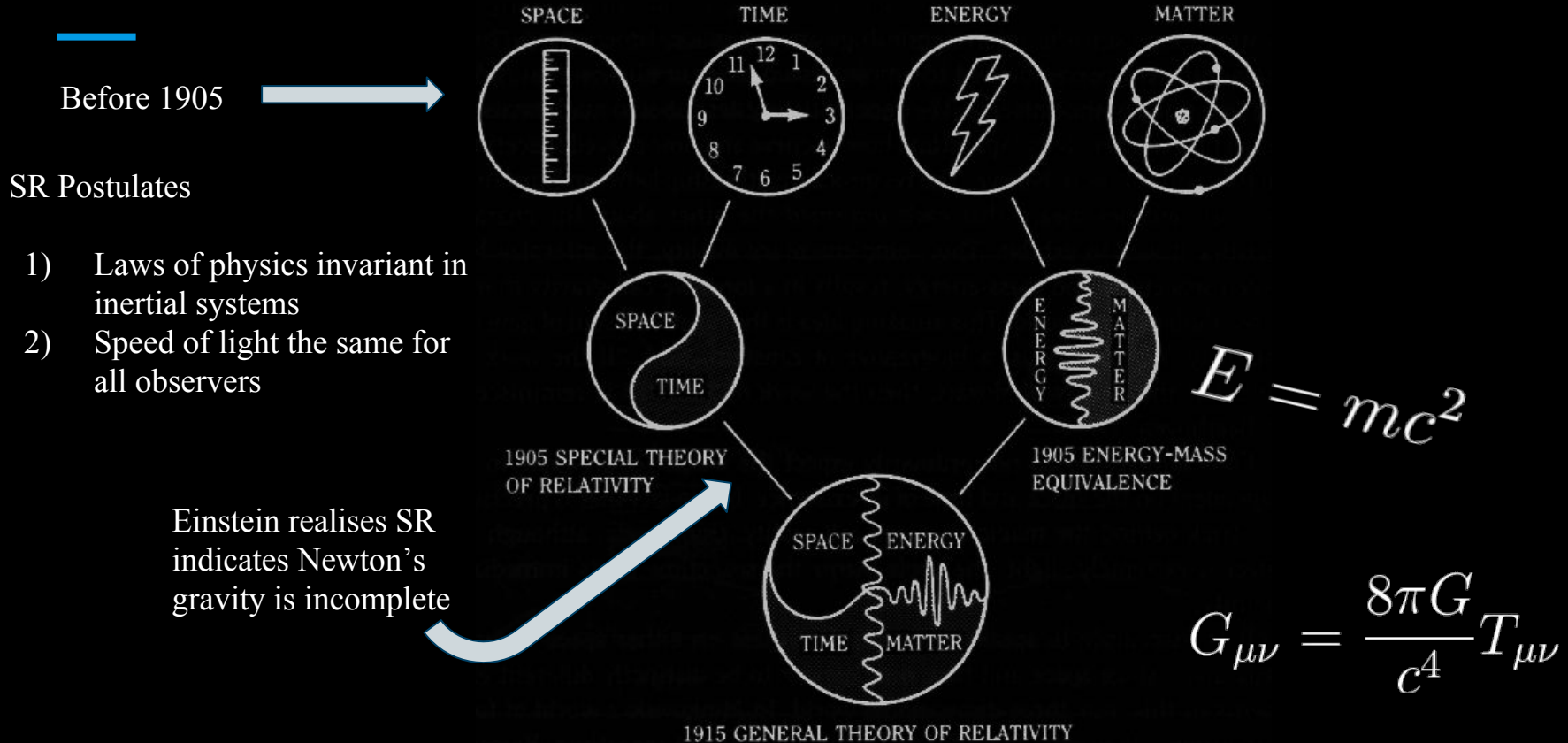
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The dawn of gravitational wave  
astronomy



# Einstein's ideas:

## Unifying space-time and energy and matter

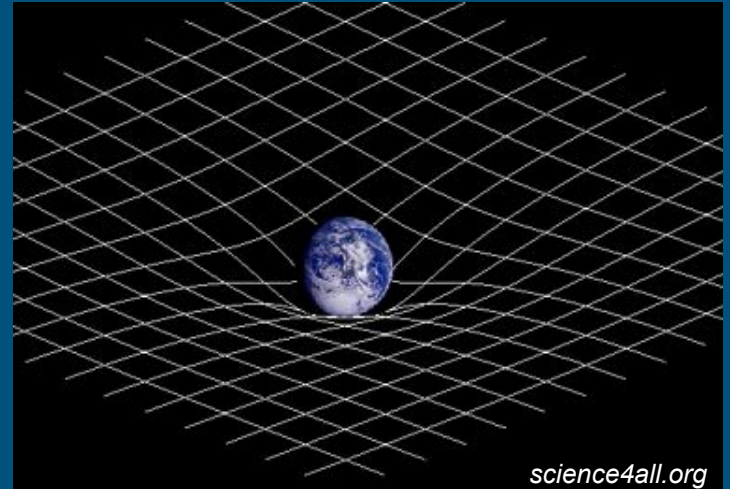


# Understanding General Relativity

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“Spacetime tells matter  
how to move; matter tells  
spacetime how to curve.”

– J. A. Wheeler



$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

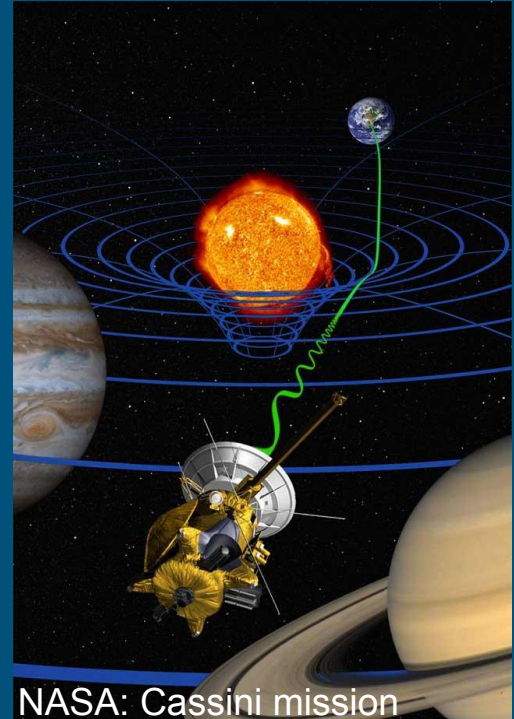
# Testing General Relativity

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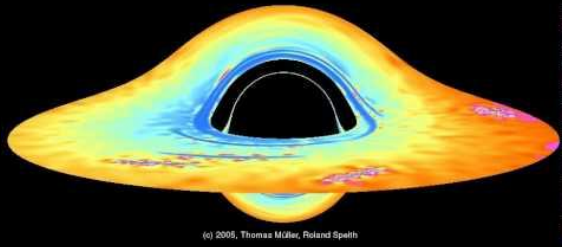
# Predictions of GR: gravity bends light



$$E = mc^2$$

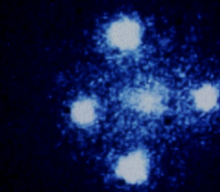


# Predictions of GR: gravity bends light



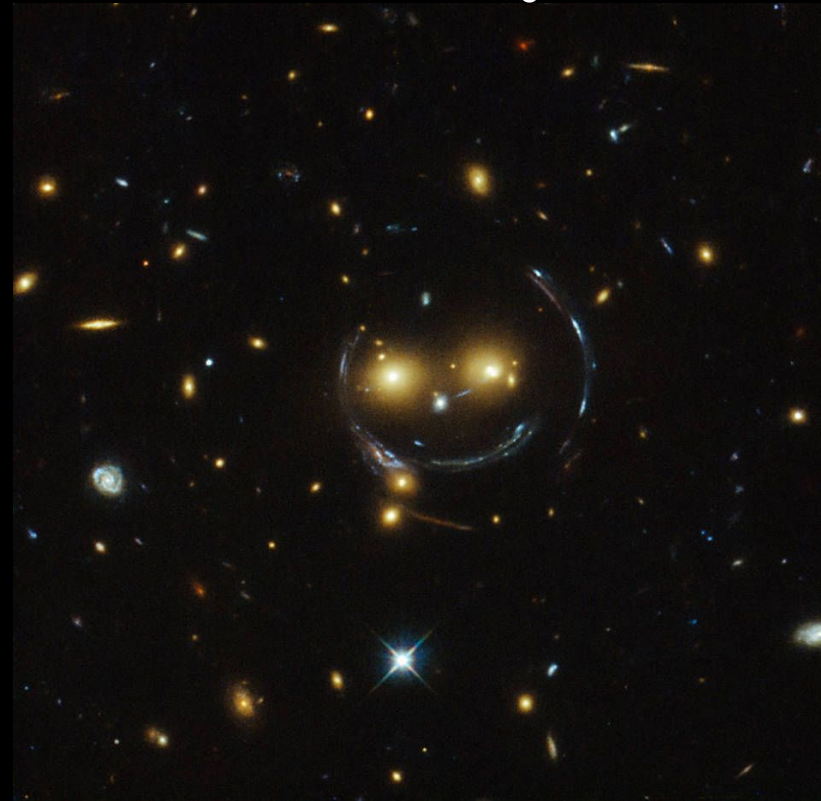
(c) 2008, Thomas Müller, Roland Speth

*Simulation of an  
accretion disk around a  
black-hole*

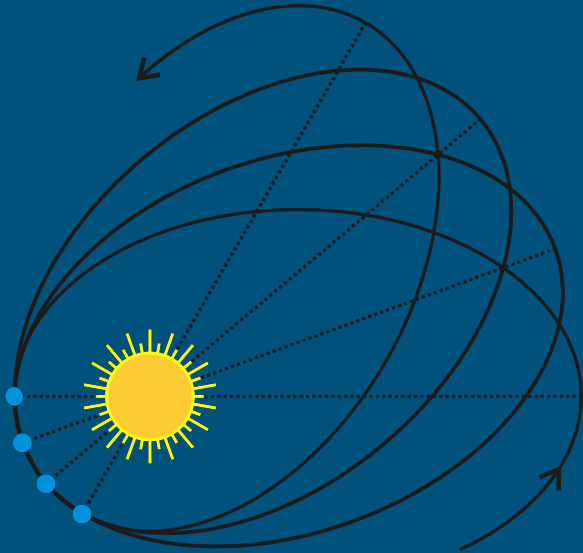


*Einstein's cross*

*The cosmic grin*



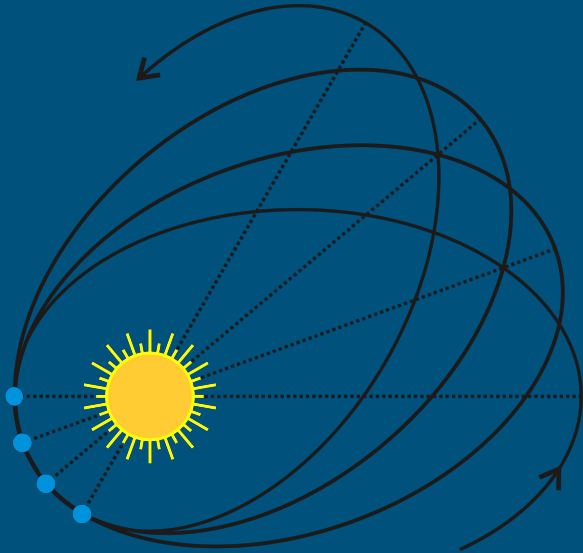
# Predictions of GR: Perihelion advance of Mercury



Amount arcsec/century	Cause
531.63 +/- 0.69	Gravitational tugs of the other planets
531.63 +/- 0.69	Total
574.10 +/- 0.65	Observed

source: [wikipedia.org/wiki/Tests\\_of\\_general\\_relativity](https://en.wikipedia.org/wiki/Tests_of_general_relativity)

# Predictions of GR: Perihelion advance of Mercury

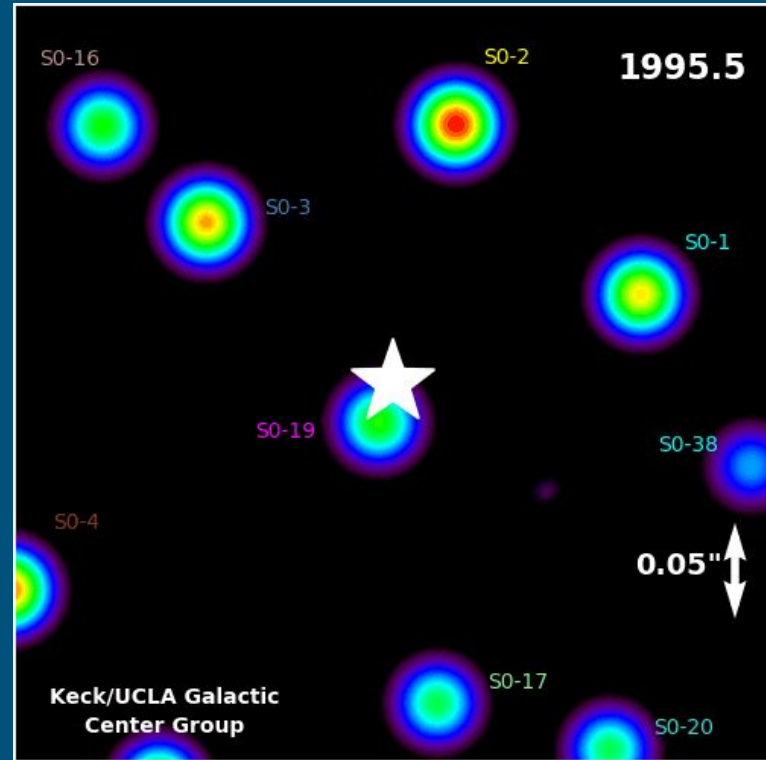
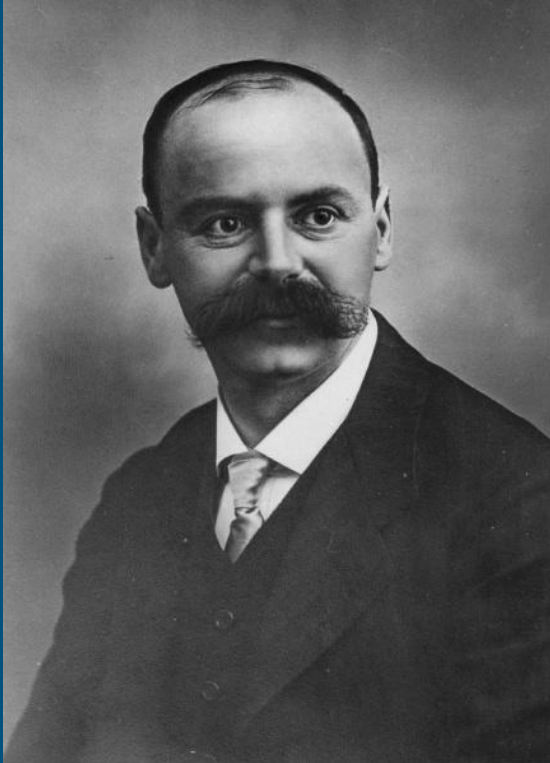


Amount arcsec/Julian century	Cause
531.63 +/- 0.69	Gravitational tugs of the other planets
42.98 +/- 0.04	General Relativity
574.64 +/- 0.69	Total
574.10 +/- 0.65	Observed

source: [wikipedia.org/wiki/Tests\\_of\\_general\\_relativity](https://en.wikipedia.org/wiki/Tests_of_general_relativity)



# Predictions of GR: Black holes



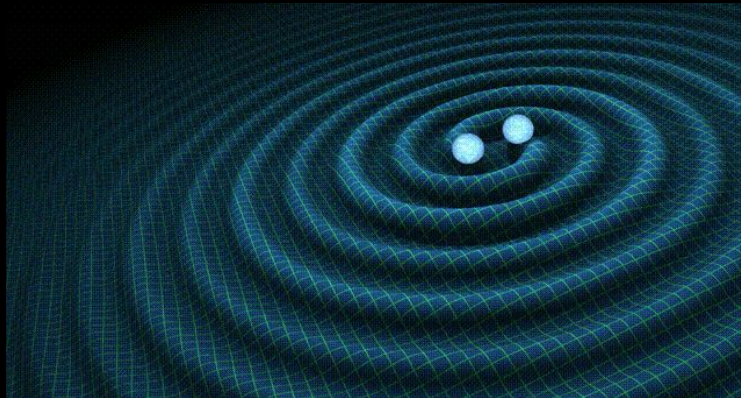
# Predictions of GR: Gravitational waves

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# Predictions of GR: Gravitational waves

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- In 1916 Einstein took the full GR equations and linearised them looking at a special case, this resulted in a so-called **wave equation**
- Gravitational waves will be produced in systems with a **time-varying quadrupole moment**
- They are understood to be **ripples in the fabric of spacetime**

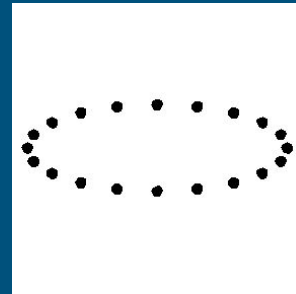
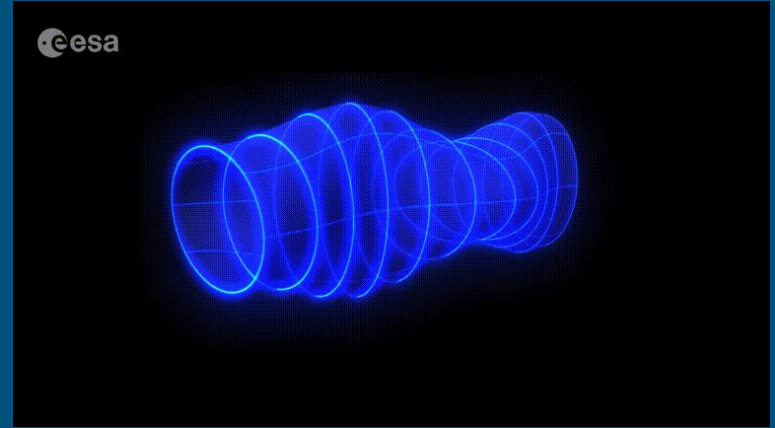


# The effect of gravitational waves

- The wave stretches and squeezes spacetime as it passes through

- Definition of strain: 
$$h = \frac{\Delta L}{L}$$

- $h \sim 10^{-21}$  corresponds to changing the radius of the Earth's orbit around the Sun by the diameter of single atom



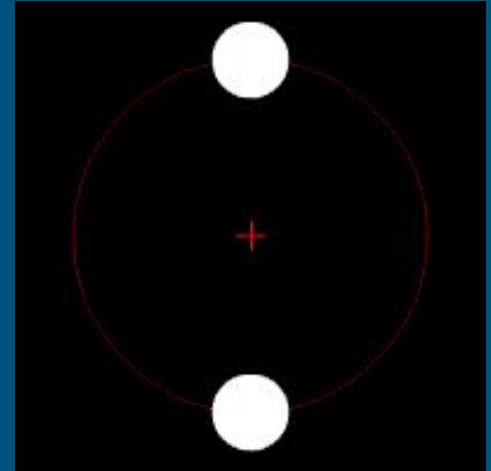
# Sources of gravitational waves

- Terrestrial sources are small:
  - Shaking your hand:  $h \sim 1e-52$
  - Battleships colliding  $h \sim 1e-46$
- So we look to space and Binary systems

- Binary strain:

$$h \sim \left( \frac{1}{R} \right) \times \left( \frac{G^2}{c^4} \right) \times \left( \frac{m_1 m_2}{r} \right)$$

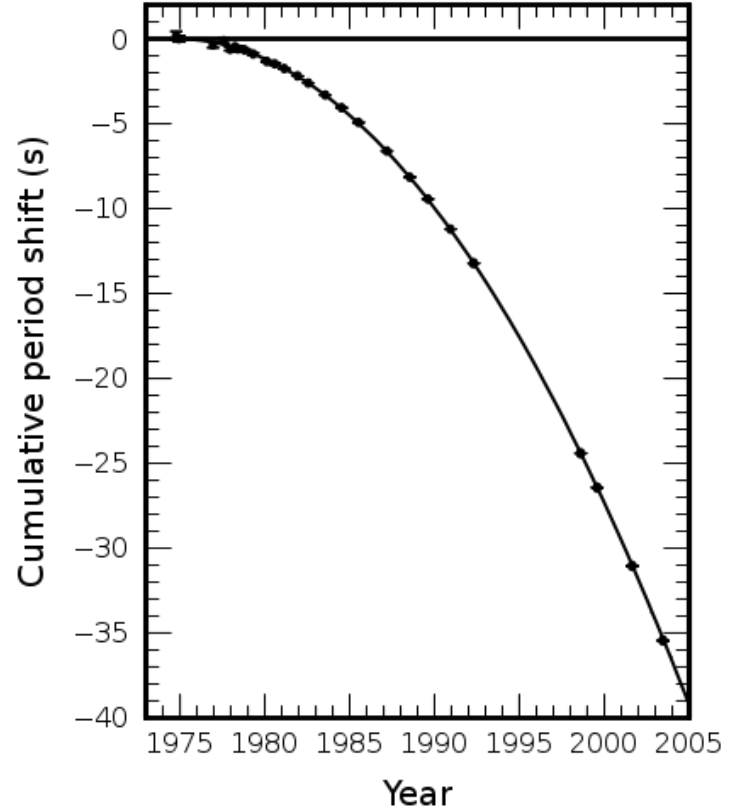
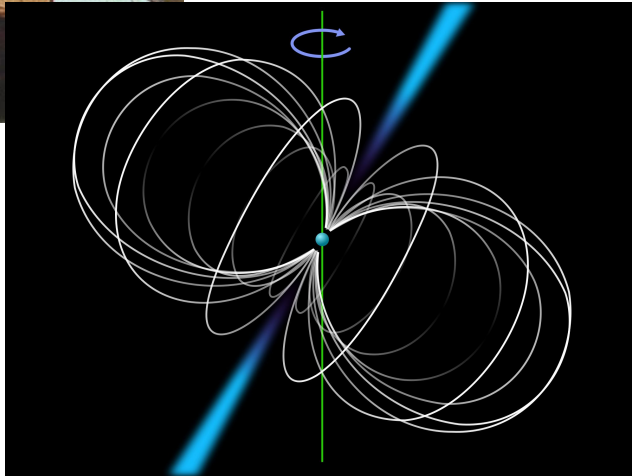
- Io orbiting Jupiter:  $h \sim 1e-25$
- Typical compact object binary *coalescence* system:  $h \sim 1e-19$  or smaller



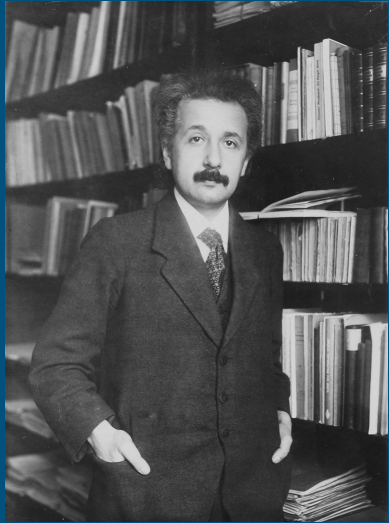
# Predictions of GR: Orbital decay due to gravitational wave emission



Hulse & Taylor  
1993 Nobel prize



# Detecting gravitational waves



1916: Einstein predicts  
GWs



1960's: Joseph Weber's resonant  
bar detectors

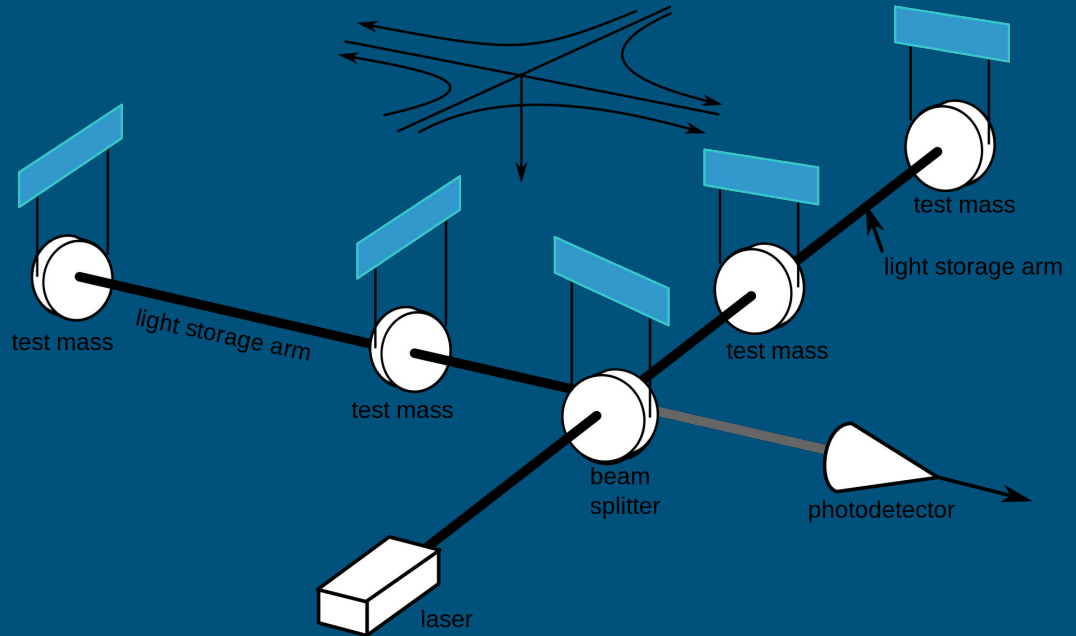


1980's: LIGO (Laser  
interferometer gravitational  
wave observatory)

# Directly detecting gravitational waves

In 1962 Michael Gertsenshtein and Vladislav Pustovoi in Moscow Russia outlined the idea for a laser interferometer to:

“monitor the relative motion of freely hanging mirrors in response to a gravitational wave”

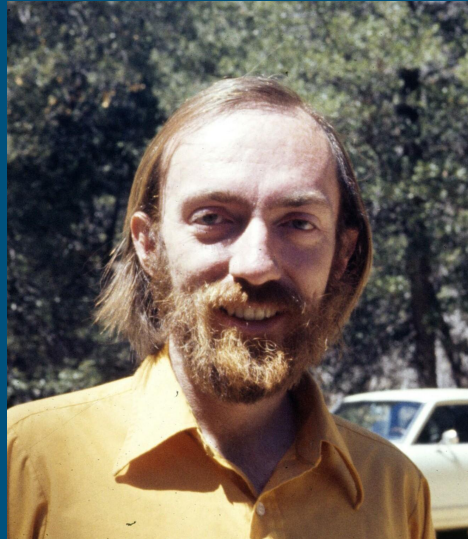




# LIGO: Laser Interferometer Gravitational-Wave Observatory

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- Founded in the 70's by Rainer Weiss, Kip Thorne and Ron Drever



# History of LIGO

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- **1972** Rai Weiss completes the invention of the detector by identifying all the fundamental noise sources and detailing how to reduce them
- **1980s** NSF funded a 40m and 1.5m prototype to demonstrate feasibility
- **1989** LIGO envisioned with a two-stage proposal for 4km detectors:
  - Initial LIGO: low probability of detection, but allows technology to be developed
  - Advanced LIGO: high probability of detection
- **1989** GEO collaboration (British and German) established with plan to build 3km detector
- **1993** Virgo detector approved by French and Italian collaboration
- **1992-2005** Building detector sites and commissioning, first runs
- **2007** Unification of LIGO, Virgo and GEO into Ligo-Virgo collaboration (LVC)
- **2006-2011** Initial detector era science runs
- **2011 - 2015** Advanced detector upgrades
- **2015** Start of Advanced detector observation runs

# The interferometer network



# LIGO-Virgo collaboration

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- Over 1000 people in the collaboration consisting of experts in
  - Theoretical physics
  - Data analysis
  - Data engineering
  - Detector characterisation + calibration
  - Detector design
- On the detection side, we are split into four groups by the source type
  - Compact binary coalescence
  - Bursts and transience (generic search)
  - Continuous wave emission
  - Stochastic

# Detecting gravitational waves

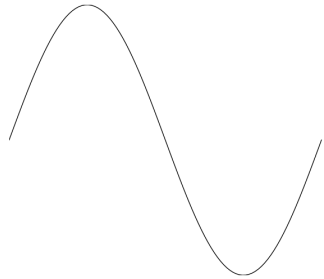
signal

+

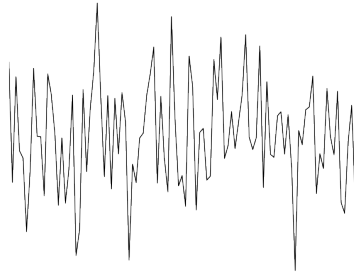
noise

=

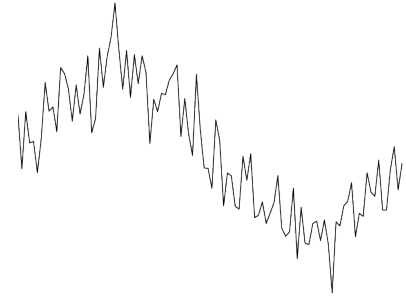
data



+



=



1. Guess what the signal looks like to make a *template*
2. Used *matched filtering* to search the data for the template

# The detection

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# September 14th 2015

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From: Marco Drago <marco.drago@aei.mpg.de>  
Subject: [lsc-all] Very interesting event on ER8  
Date: Mon, 14 Sep 2015 12:55:46 +0200

Hi all,  
cWB has put on gracedb a very interesting event in the last hour.  
[...]  
It is not flag as an hardware injection, as we understand after some  
fast investigation. Someone can confirm that is not an hardware  
injection?

Marco

# Detection process: GW140915

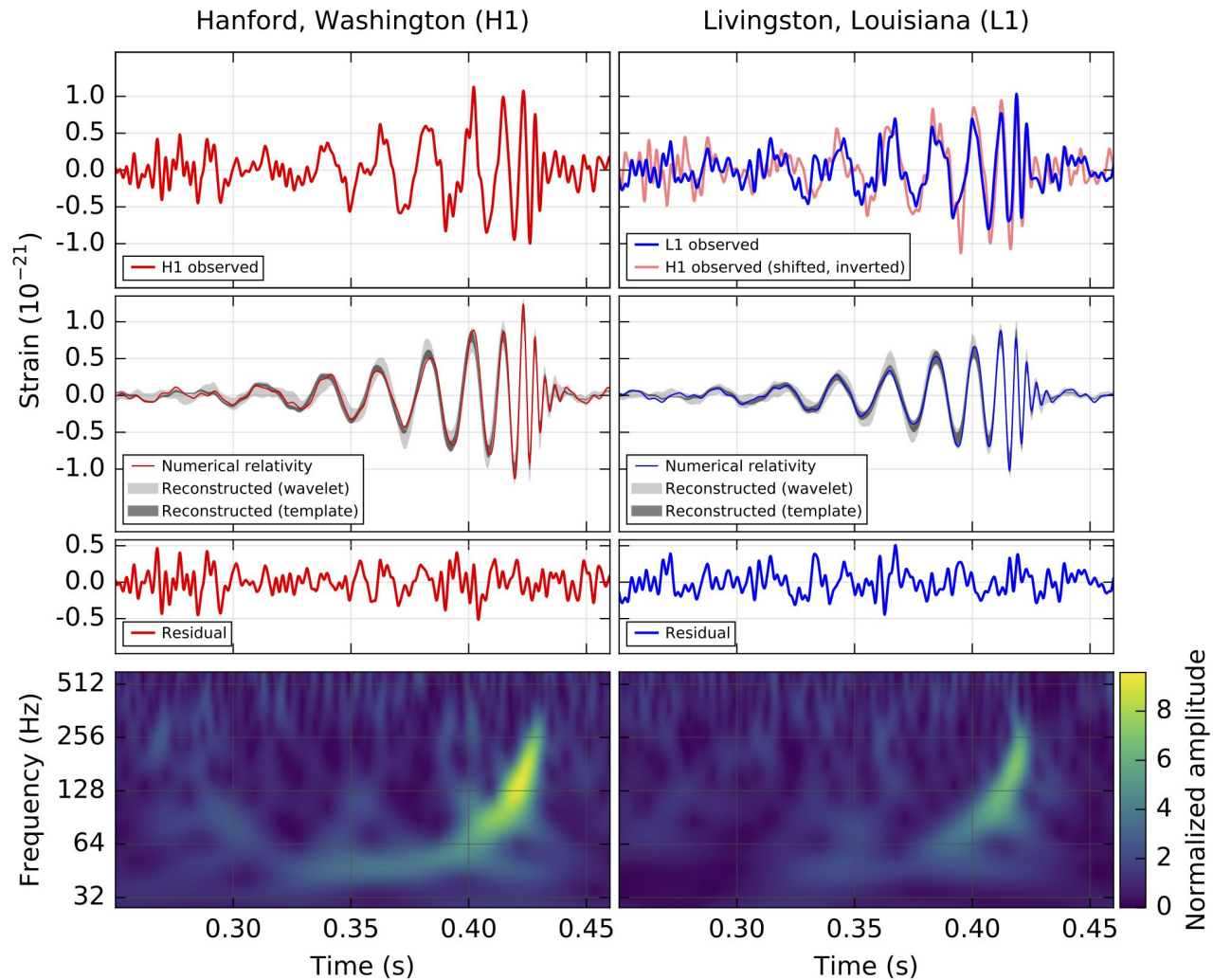
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- Signal detected at both Hanford and Livingston - no other detectors online at the time
- Detectors frozen in configuration and checked for
  - Any terrestrial cause (via on-site instruments)
  - Any malicious or accidental injection
- Signal detected by two independent pipelines (more about this later..)
- Detection committee established to 'play devil's advocate'



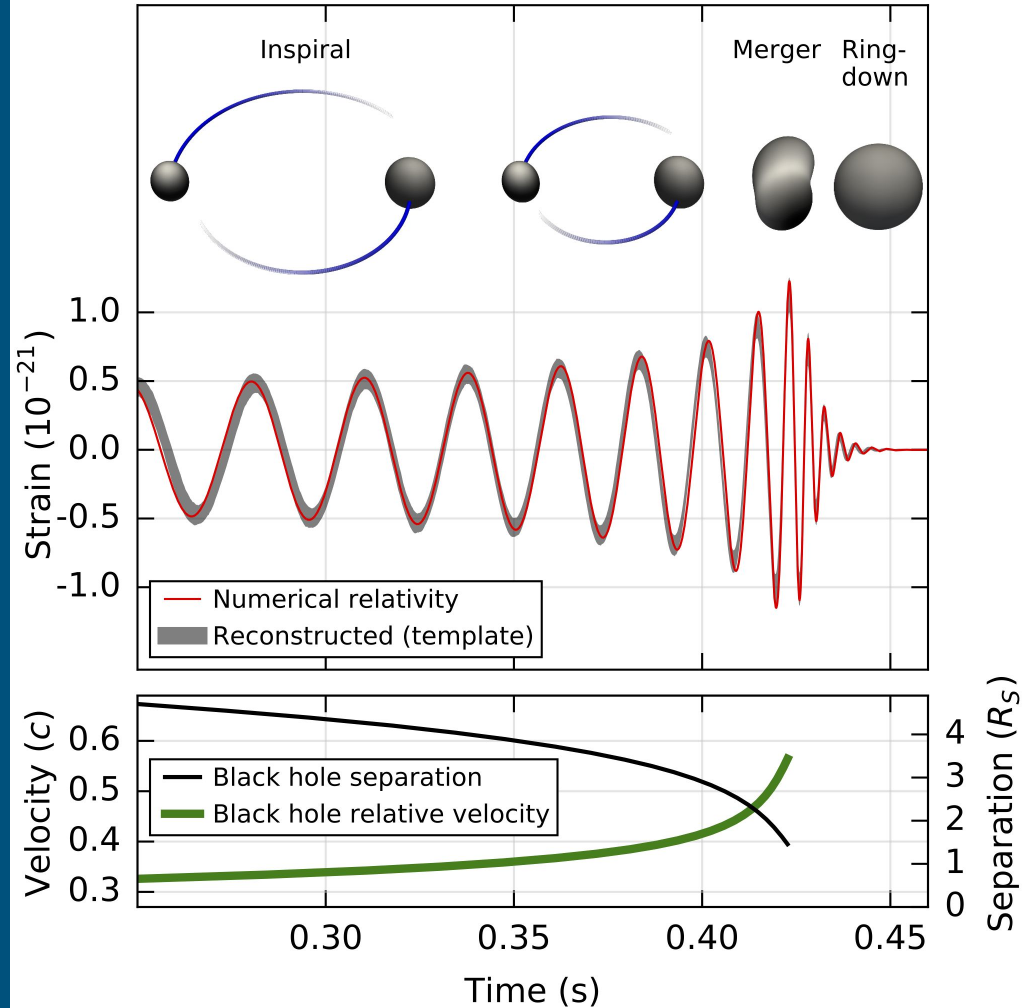
# Observed signal strain $h(t)$

Plots taken from:  
[B. P. Abbott et al. \(LIGO Scientific  
Collaboration and Virgo Collaboration\)  
\(2016\)](#)  
[Phys. Rev. Lett. 116, 061102](#)



# The source:

- Peak strain:  $h \sim 1e-21$
- $R = 410$  Mpc or 1.3 billion ly
- Initial blackholes had mass  $36 M_{\odot}$  and  $29 M_{\odot}$
- Final blackhole has mass  $62 M_{\odot}$
- $3 M_{\odot}$  radiated in GWs
- Occurred 1.3 Billion years ago prior to multicellular life on earth



# How certain are we that this isn't noise?

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- Signal to noise ratio: 24
- False alarm rate: 1 event per 202 600 years
- Next loudest event was 1 event per 10 years
- Equivalent to a significance greater than  $5.1 \sigma$
- Possible electromagnetic confirmation from Fermi gamma ray telescope  
(although BBH mergers not expected to be visible)

# The announcement

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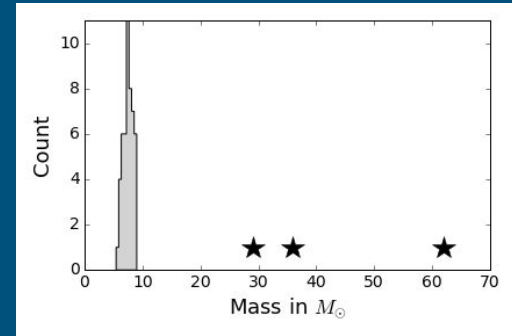
- February 11th 2016 LIGO held a press conference
- *“We have detected gravitational waves, we did it”* - David Reitze
- Simultaneously released the detection paper “Observation of Gravitational Waves from a Binary Black Hole Merger” and the data
- 10 other papers were also released covering astrophysics and detector

So, what did we learn from  
GW140915?

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# The dawn of gravitational wave astronomy

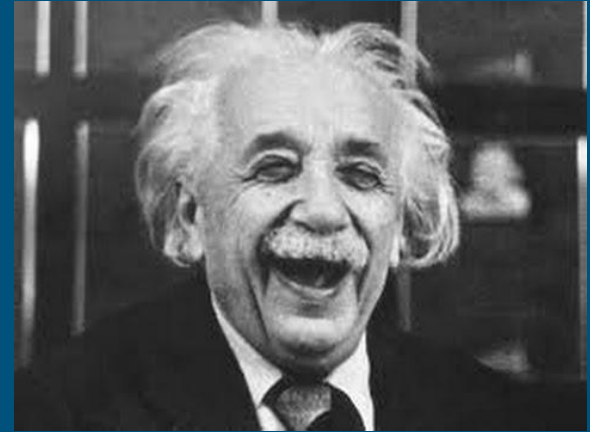
- GW140915 tells us:
  - GWs exist and can be measured
  - Tests General Relativity
  - Blackholes with masses around 30 - 60  $M_{\odot}$  exist
  - Blackholes coalesce within the age of the universe
  - We can detect other events!
- In the future we might:
  - Detect more BBH systems
  - Detect BNS and BH+NS systems
  - Multimessenger astronomy
  - Detect isolated NS
  - Have a headstart when detecting a supernova event - last one was in 1604
  - Detect the stochastic background



# Conclusions

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- First detection of gravitational waves
- Further confirms Einstein's theory of GR
- Result of 30+ years of hard-work and commitment
- We can look forward to an exciting era of gravitational wave astronomy



# Current detector status

