

Understanding and Improving the Accuracy of Advanced LIGO Calibration

LIGO-T1600260

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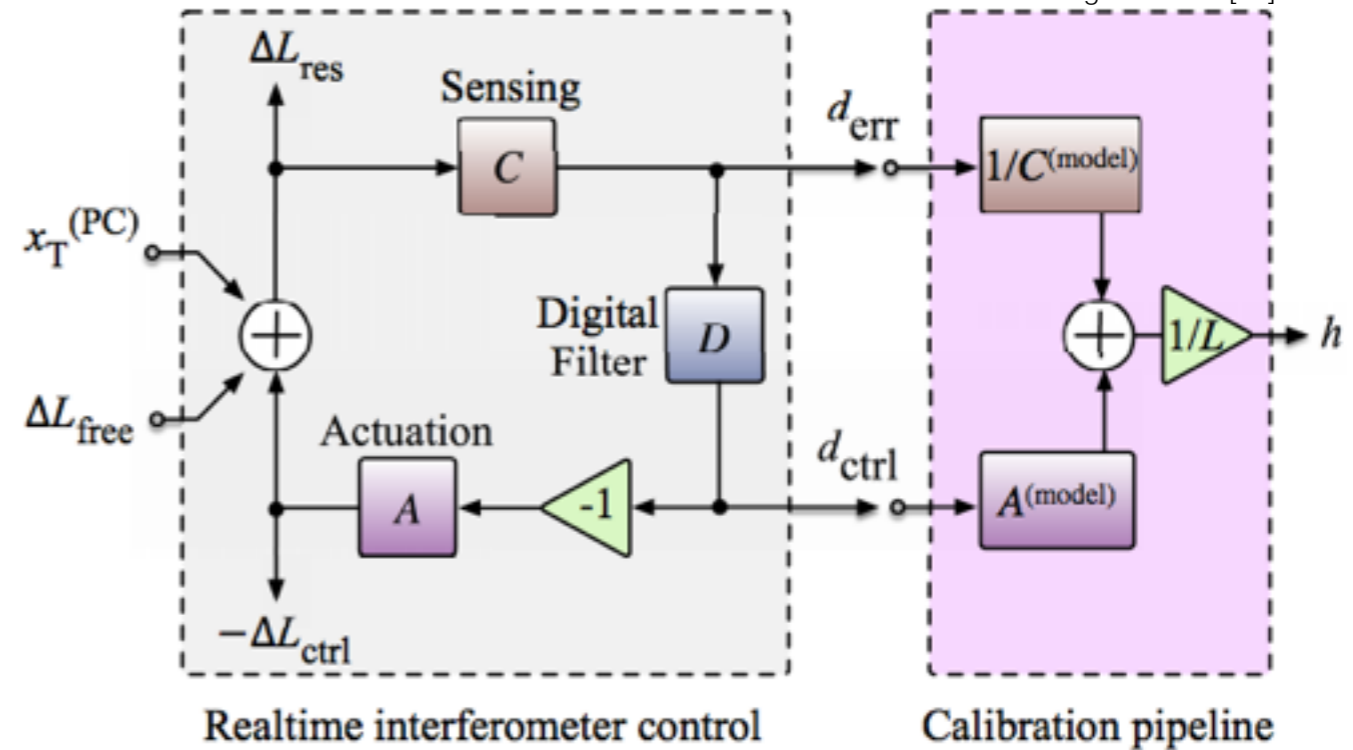
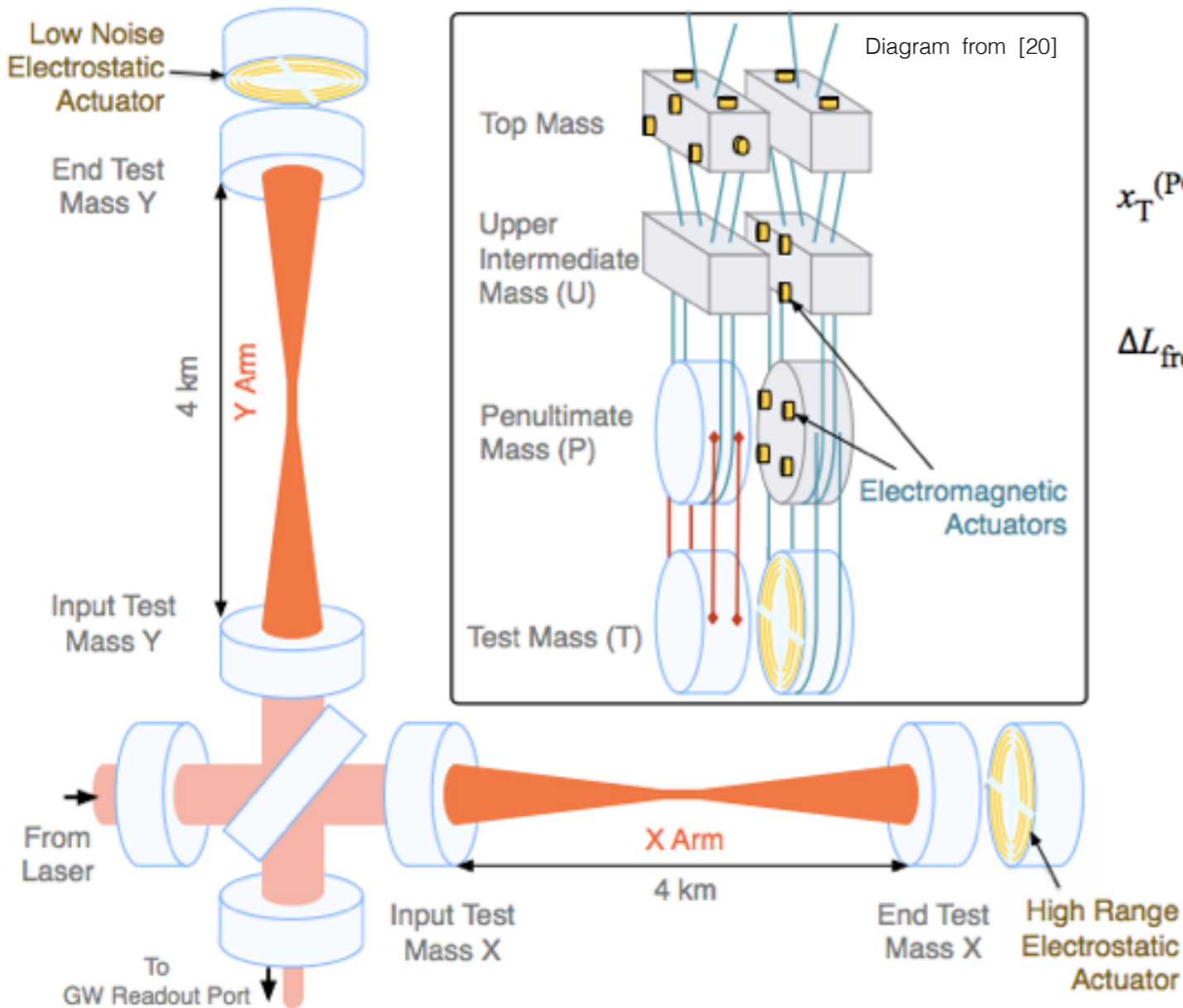
Objective

Estimate the impact of aLIGO calibration uncertainty on precision tests of GR using observations of GWs from BBH mergers

Optical Gain Calibration
&
Ringdown deviation

Calibration methods

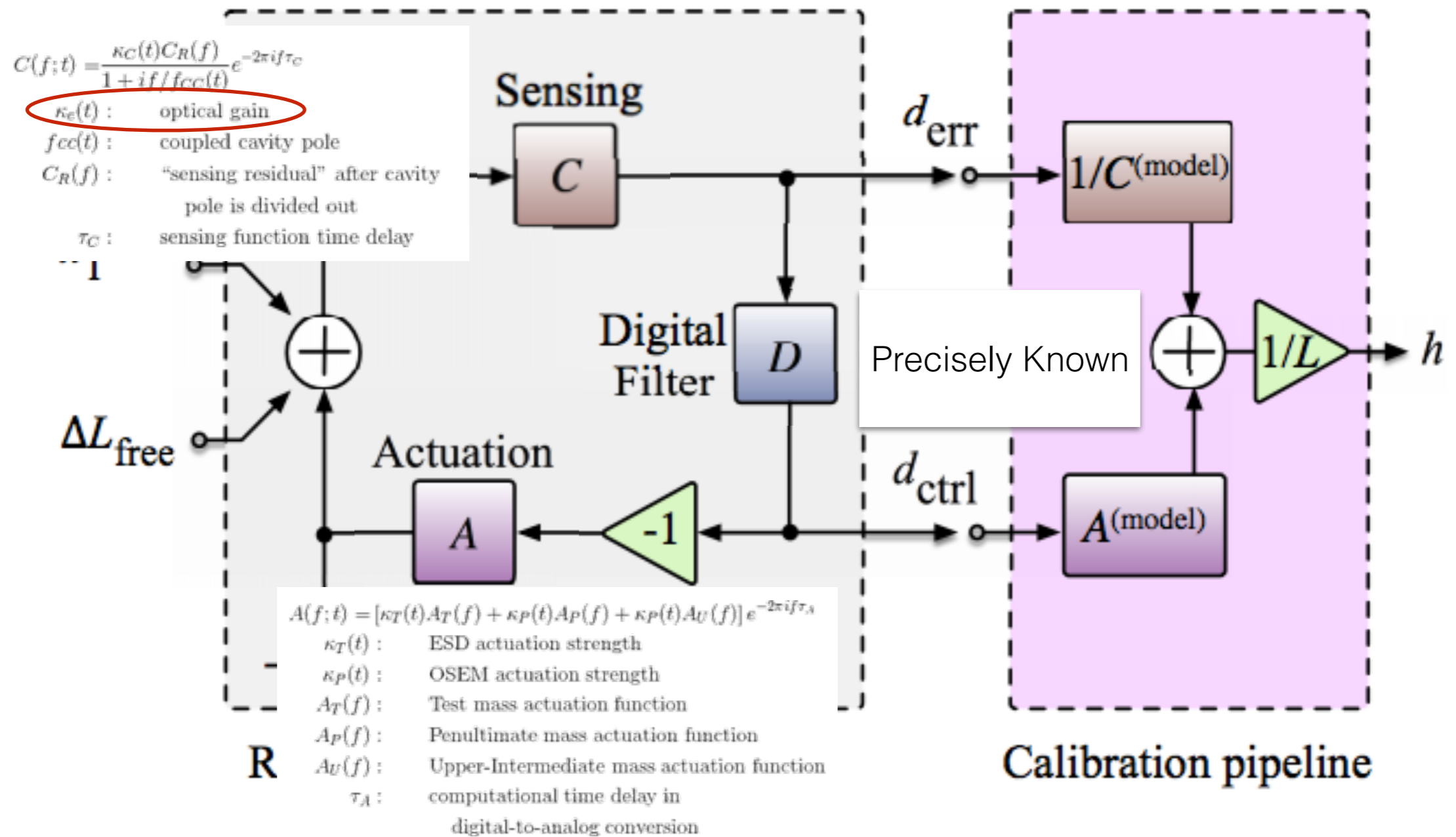
Diagram from [20]



$$R_e(f; t) = \frac{1 + A(f; t)D(f)C(f; t)}{C(f; t)}$$

$$= \frac{1 + G(f; t)}{C(f; t)}$$

Calibration methods

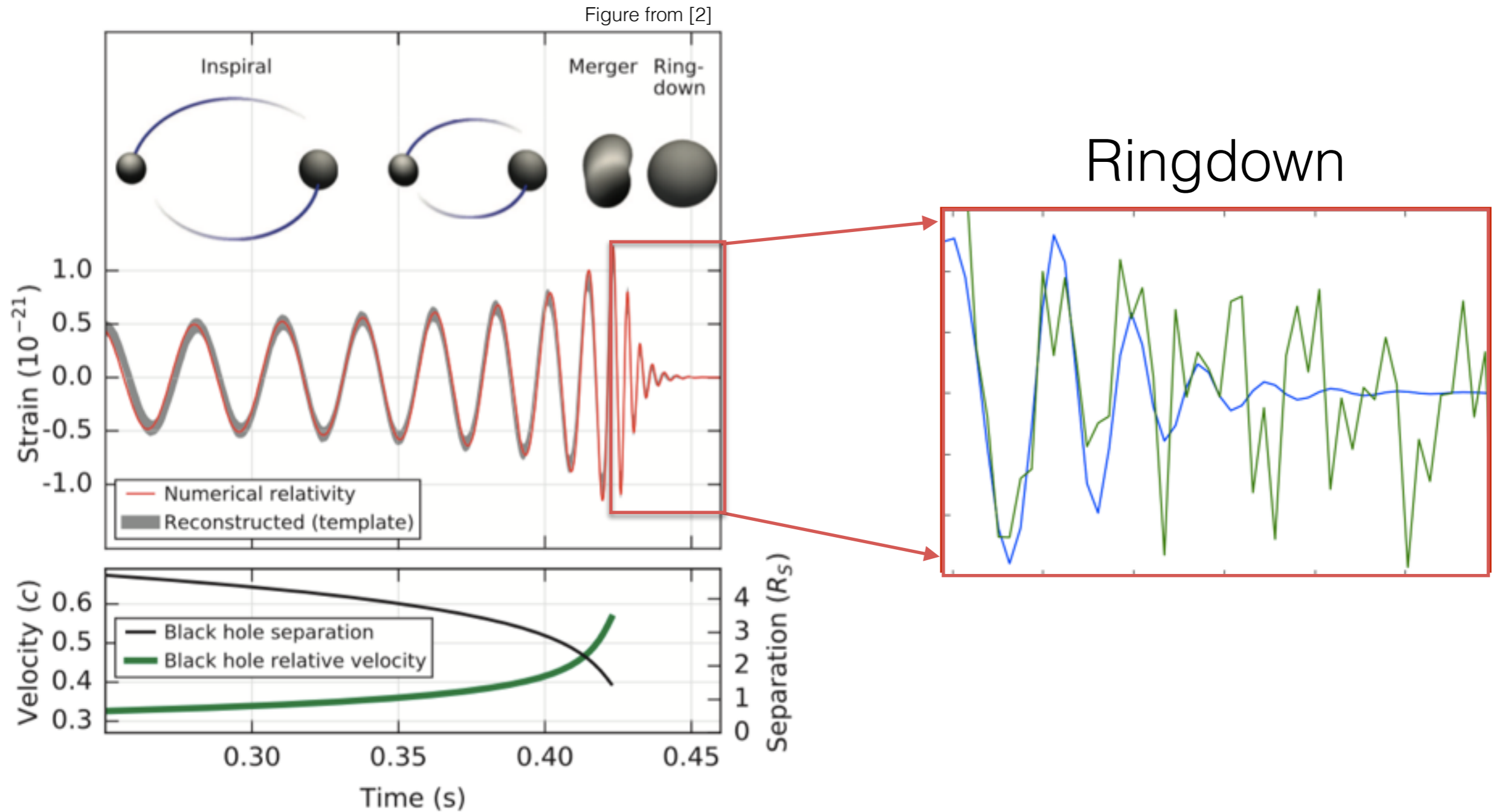


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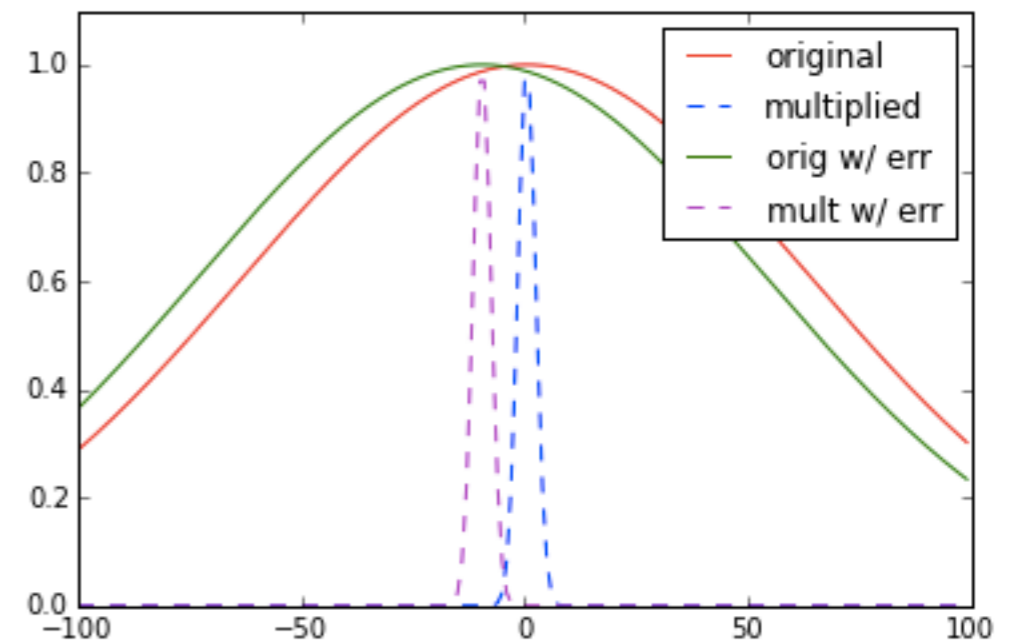
Calibration uncertainty & precision tests of GR



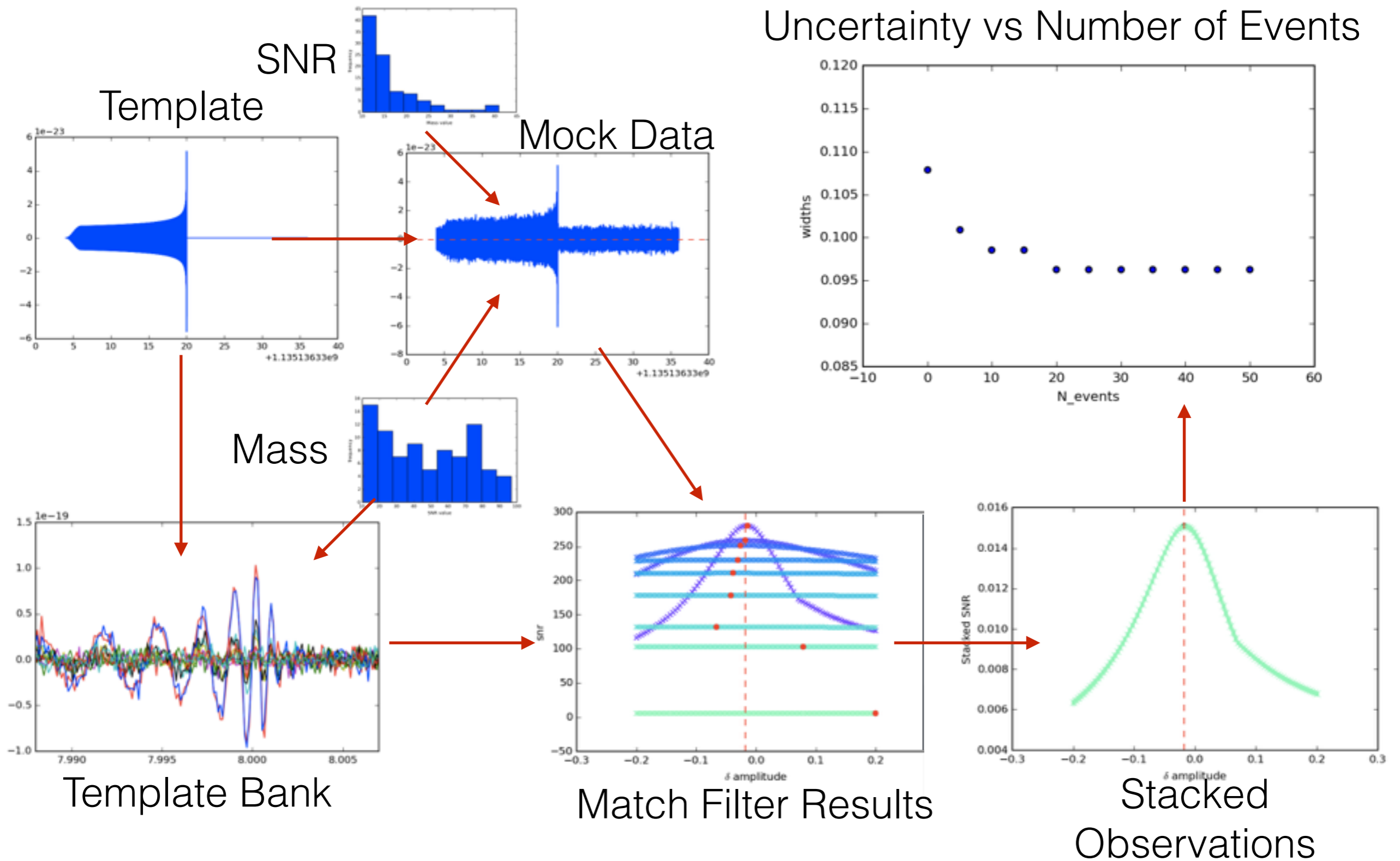
Calibration uncertainty & precision tests of GR

Stacking Observations

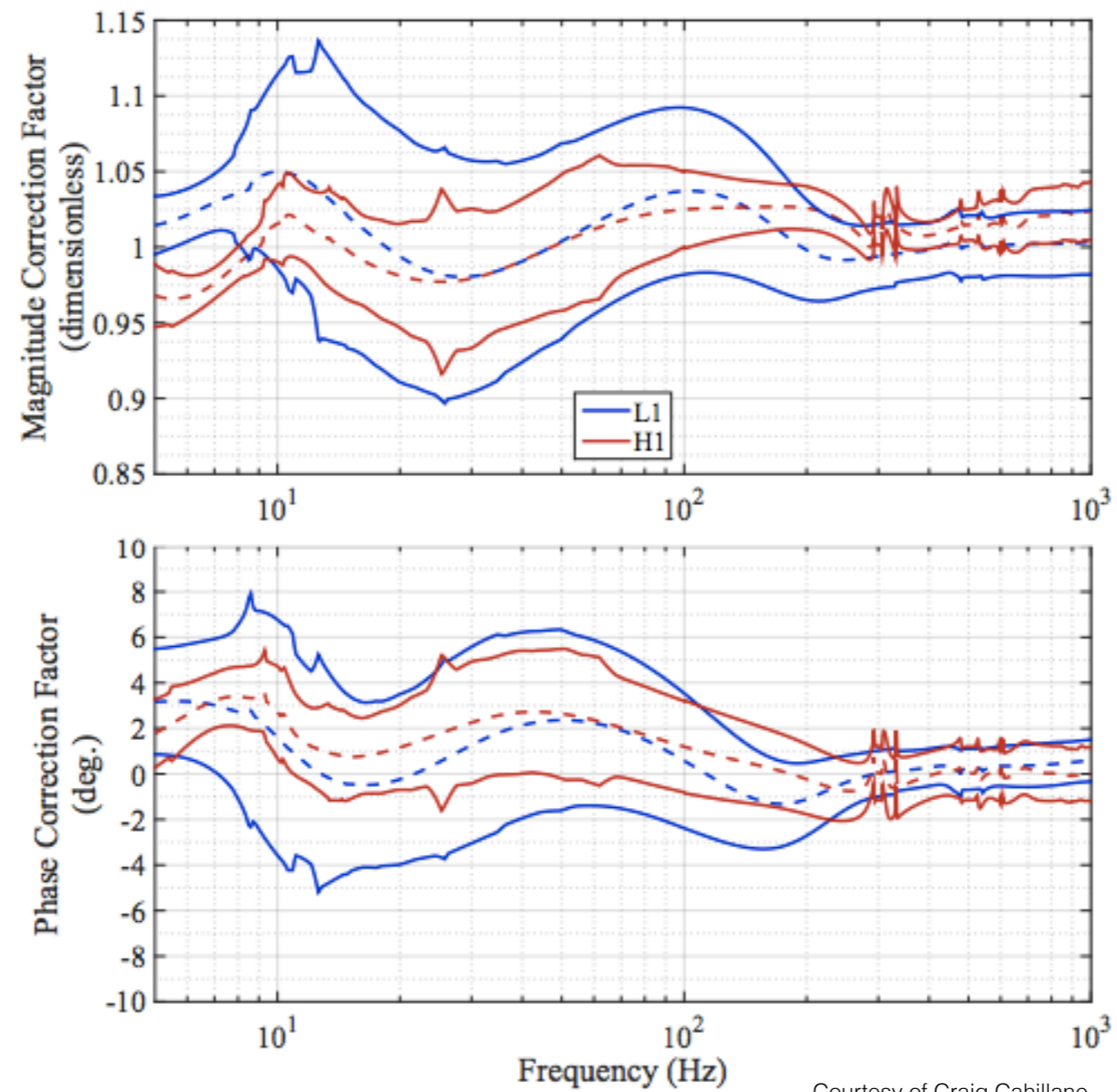
$$P(\theta|x_i) \propto \left[\prod_i P(x_i|\theta) \right] \times P(\theta)$$



A Simple Model



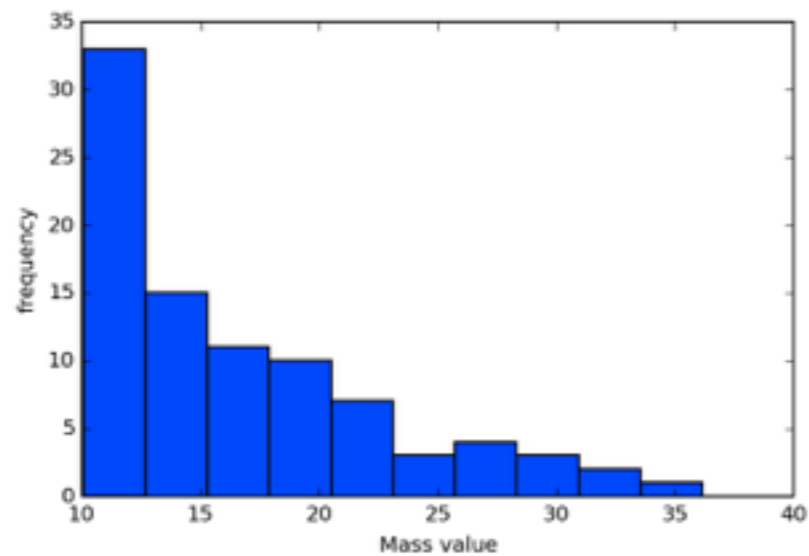
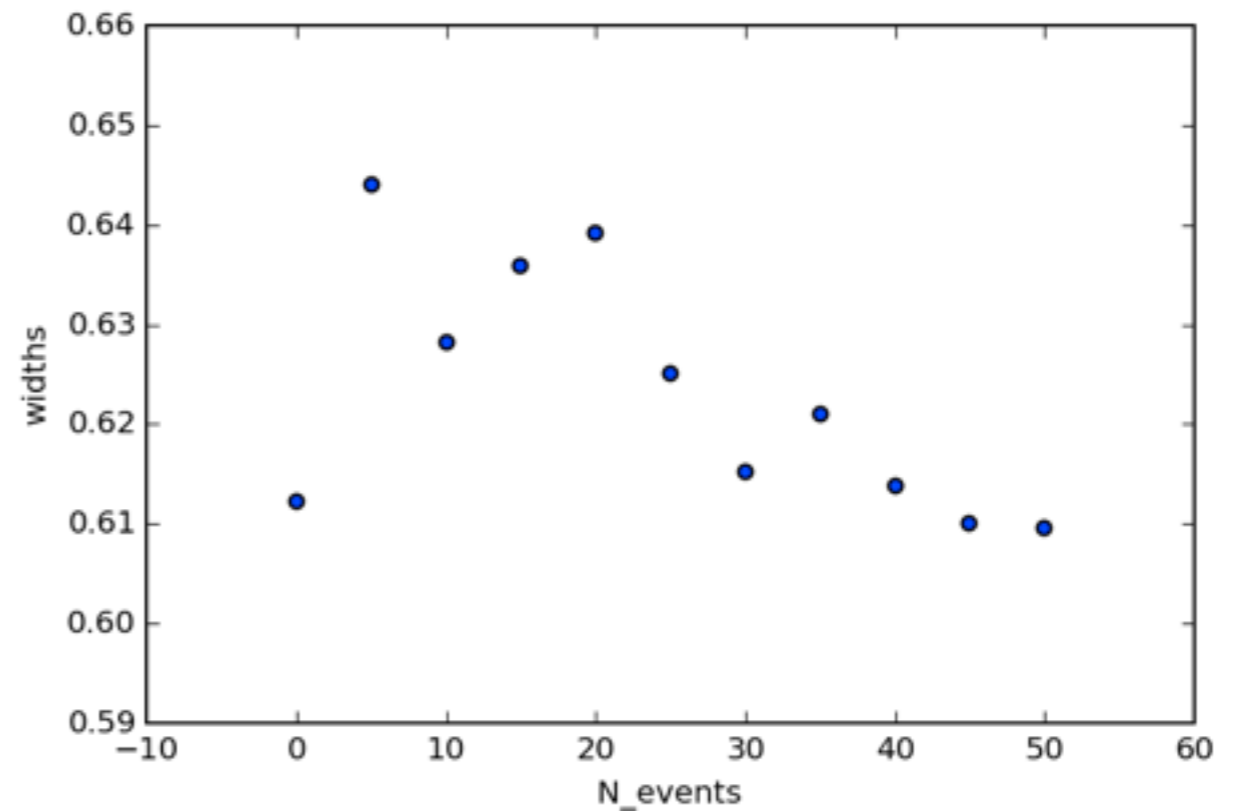
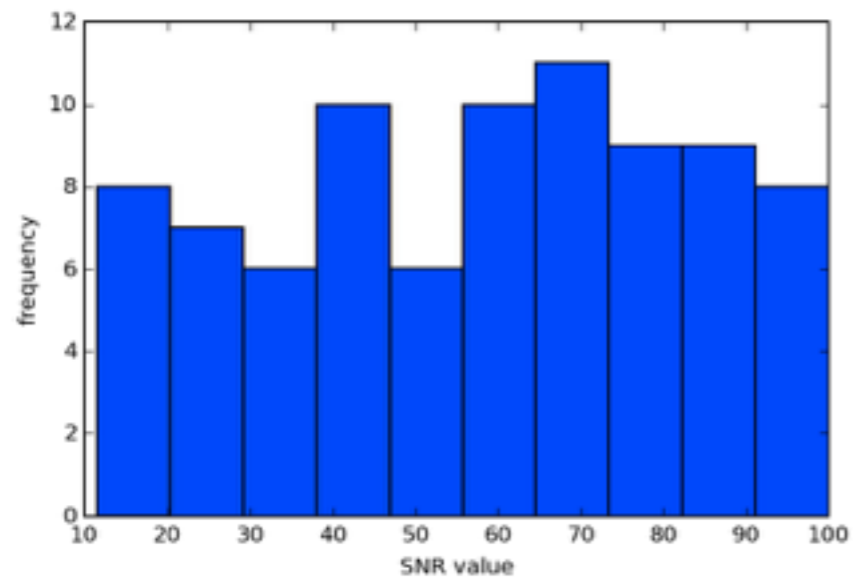
Calibration methods



Courtesy of Craig Cahillane

Results

50 events
550 templates



Next Steps

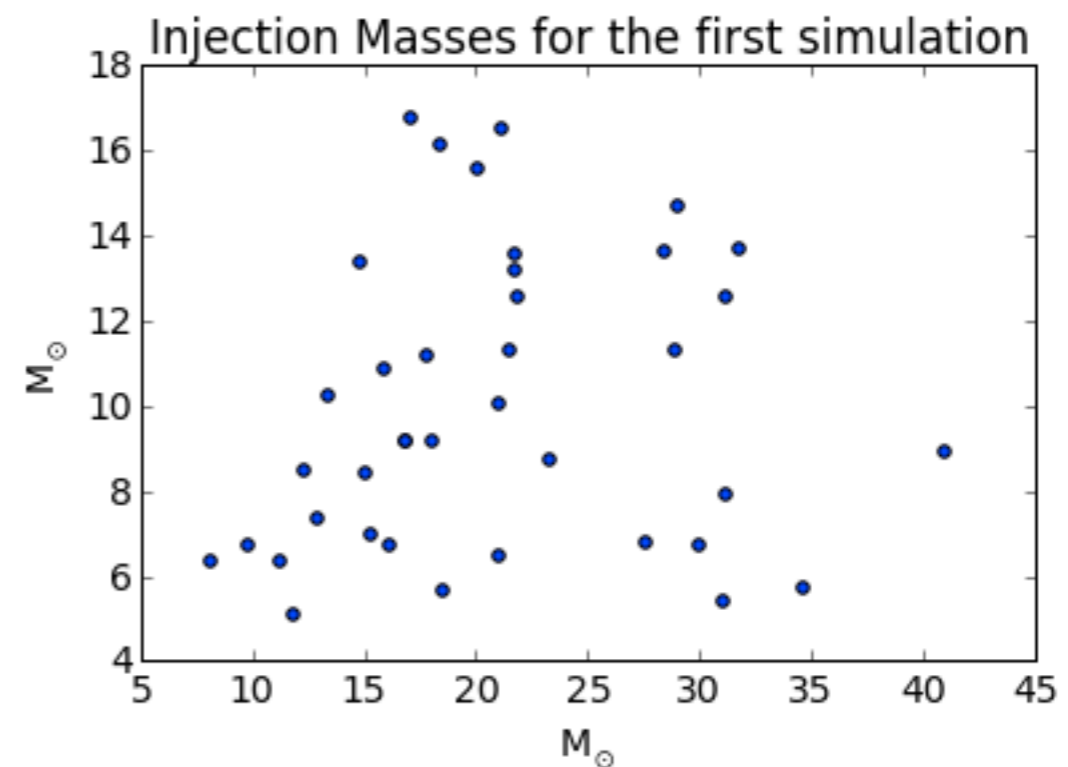
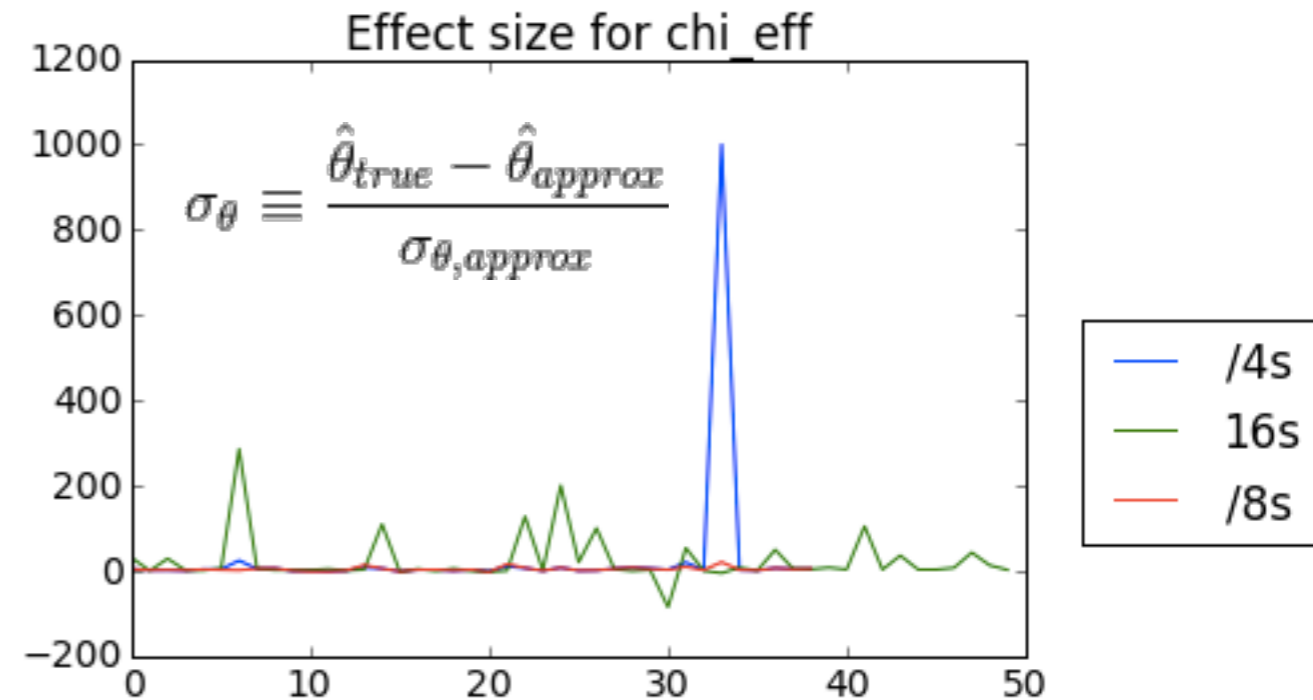
- Repeat with calibration errors & error models
- Repeat for low SNR
- Increase sophistication of model: LALSuite

LALInference & TIGER

Test Infrastructure for
General Relativity

Odd's Ratio

$$O_{GR}^{modGR} \equiv \frac{P(H_{modGR}|d, I)}{P(H_{GR}|d, I)}$$



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- Kent Blackburn

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