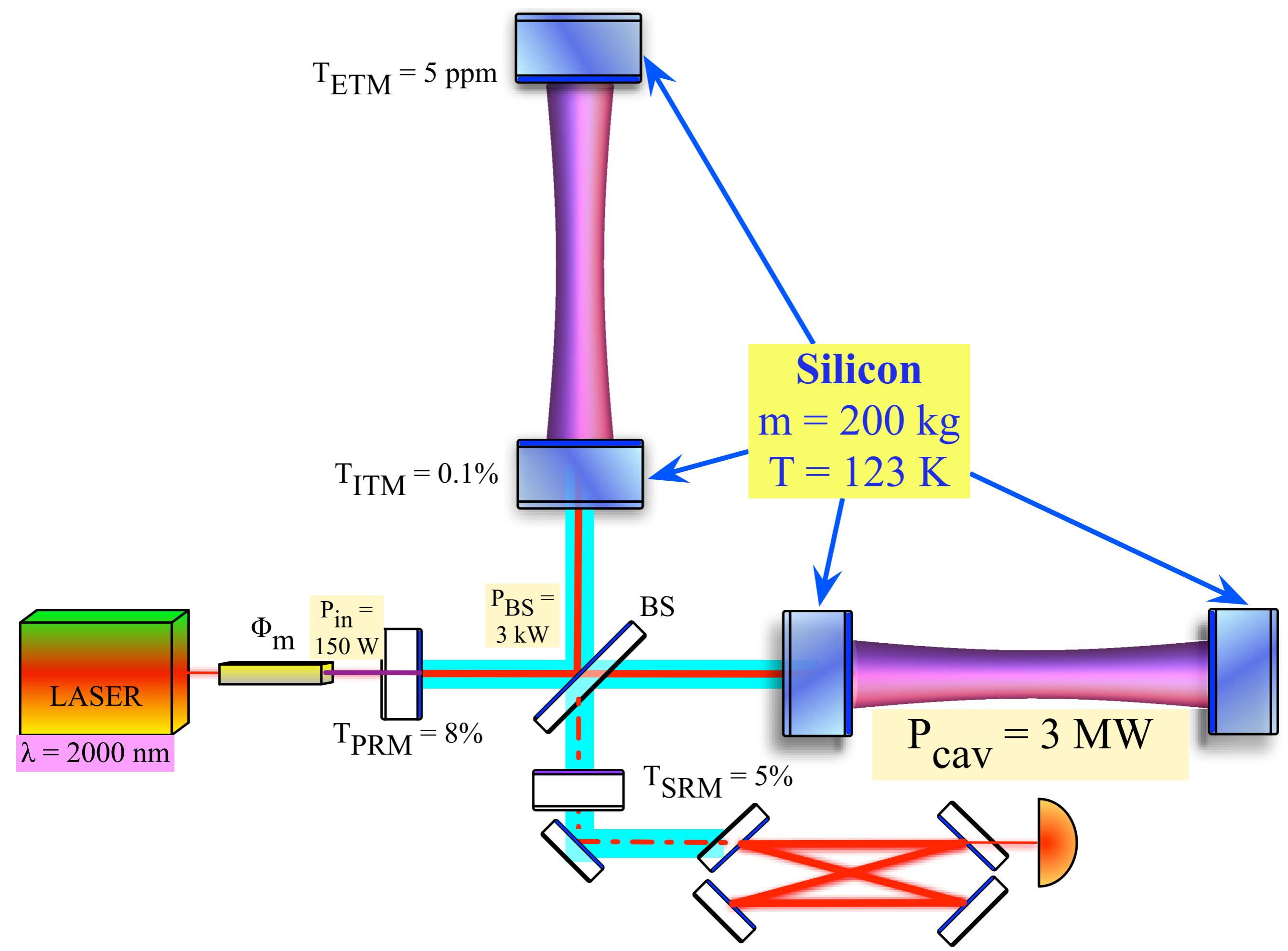
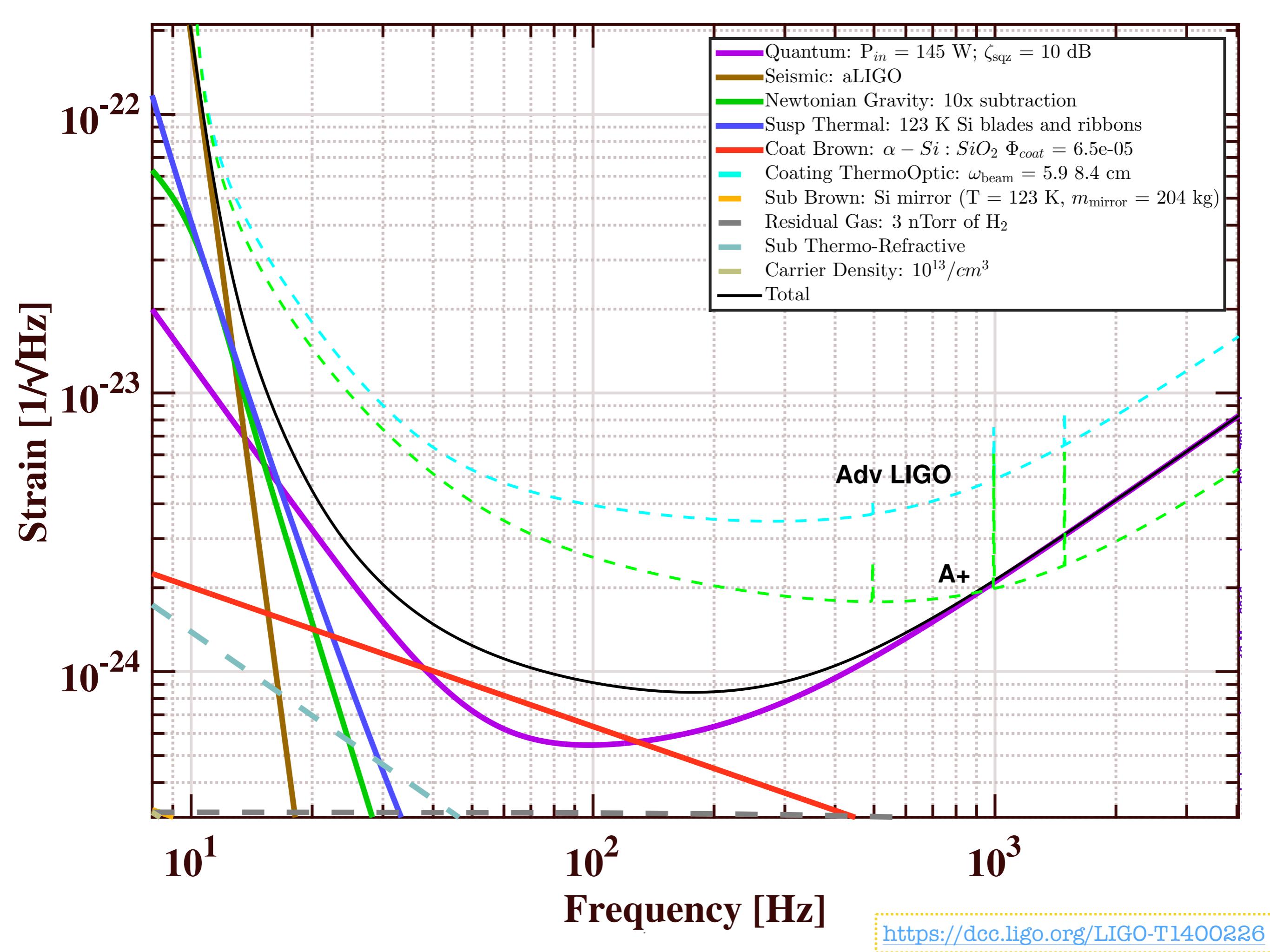


LIGO Voyager

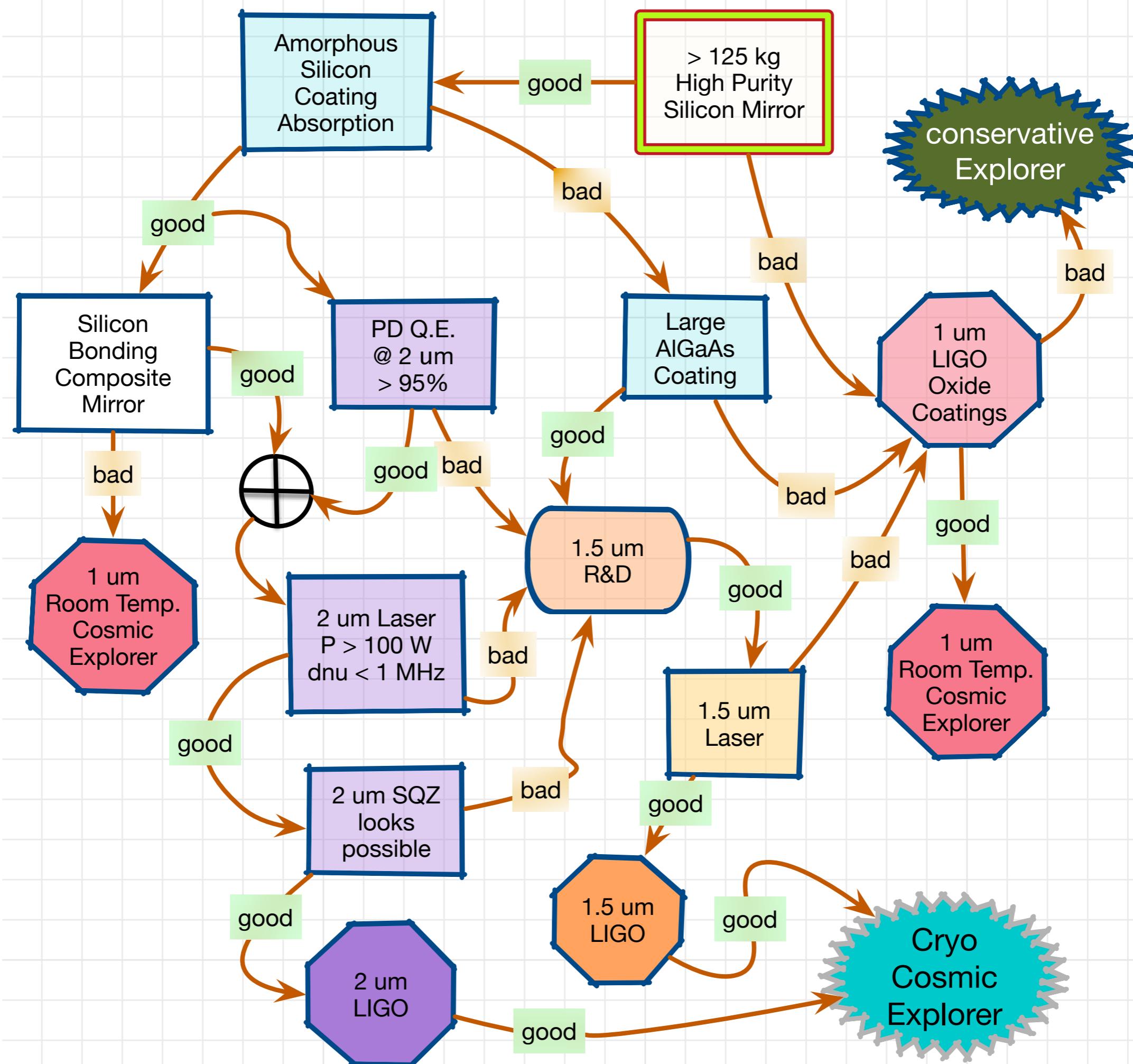
Rana Adhikari
Caltech





Sub-Systems

- Silicon Mirrors: 140-**200 kg**, mCZ
- Coatings: a-Si/SiO₂ or others
- *Wavelength Choice: 1.55 - 2.1 microns*
- Cryogenics: 123 K (for zero CTE), radiative (non contact) cooling
- Lasers (~2 micron): P_{PRM} ~ 140 W, P_{ARM} ~ 3 MW
- Thermal Compensation: Silica compensation plates only (CO₂ lasers, no ring heaters; no heating of test mass)
- Photodiode Quantum Efficiency: 80 -> 99% for 2 micron



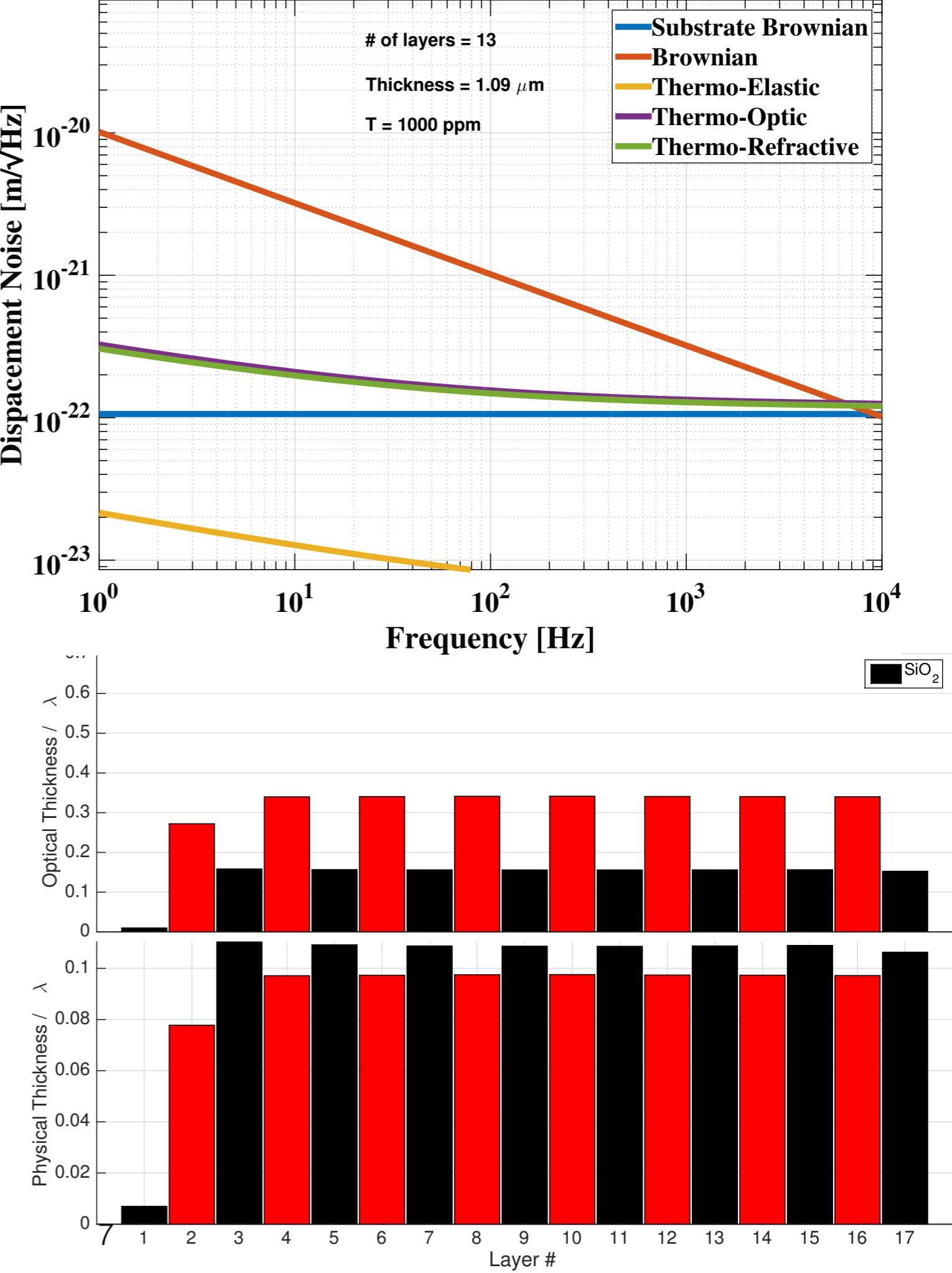
200 kg Silicon Mirror

- $P_{\text{abs}} < 5 \text{ W}$ (goal); $P_{\text{coat}} = \alpha * 3\text{MW}$; $P_{\text{sub}} = d_{\text{thick}} * \alpha * P_{\text{BS}}$ ($1\text{W} \sim > \underline{10 \text{ ppm/cm}}$)
- 3 ppm/cm (FZ): FZ max diameter $\sim 20 \text{ cm}$
- mCZ from SEH can get 10-20 kOhm in wafers after high T annealing (to trap oxygen)
- samples acquired, absorption measurements in progress (< 2 ppm)
- SEH Japan will make 45 cm diameter mCZ
- how to sequence all of the annealing? Different processes for substrates, coatings.



Coatings

- a-Si / SiO₂ baseline
- Pohl, Hellman data
- Glasgow IBS results
- Evidence of high T deposition leading to low friction due to high surface mobility*
- high T deposition with IBS this year
- lower absorption in a-Si (1-5-2 microns) (Glasgow)



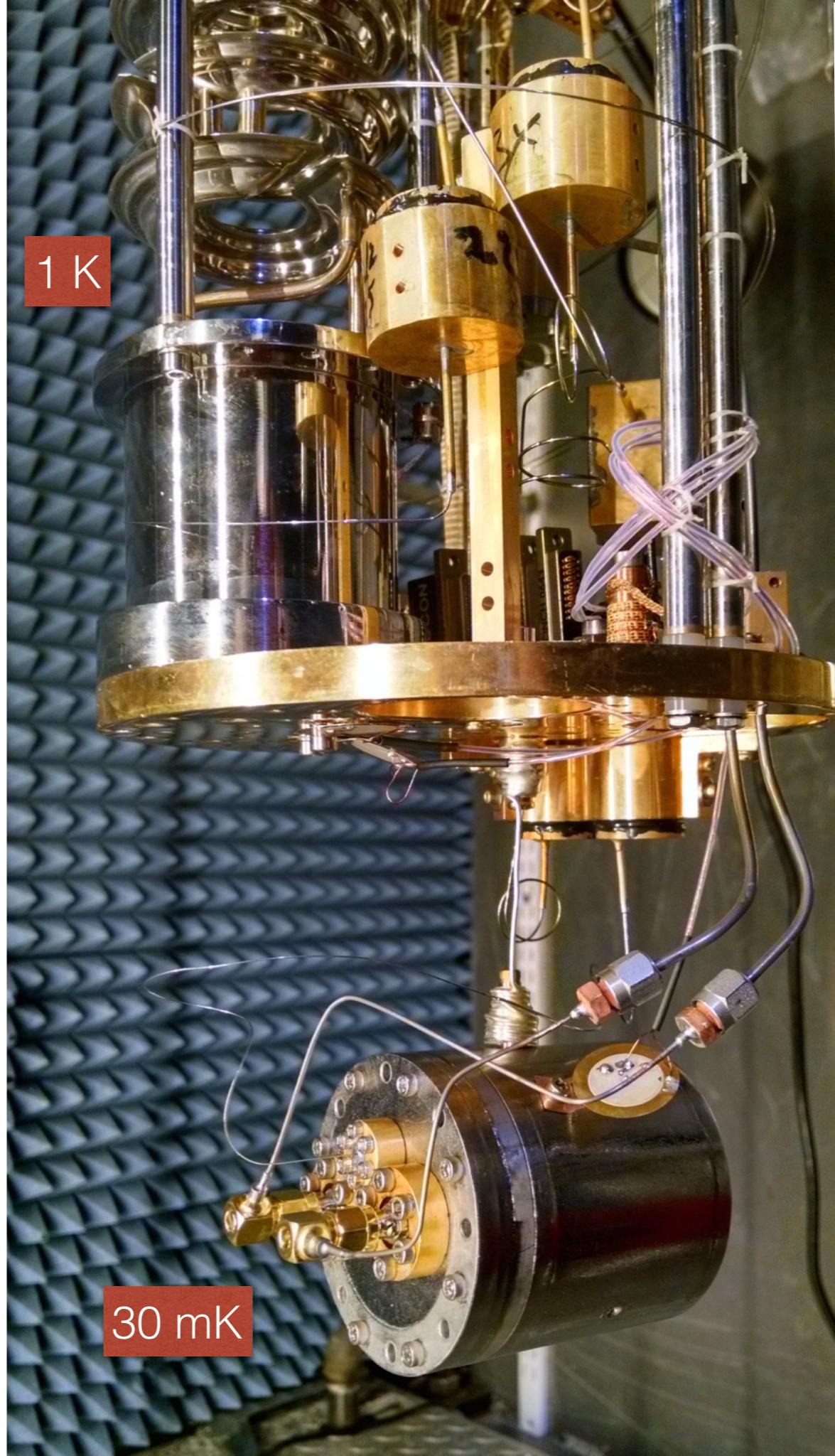
*Physics Today (Jan 2016):

<http://arxiv.org/abs/1512.03540>

serious Cryogenics

Liquid Helium
Resonant 'Bar'
at **Caltech** (Schwab)

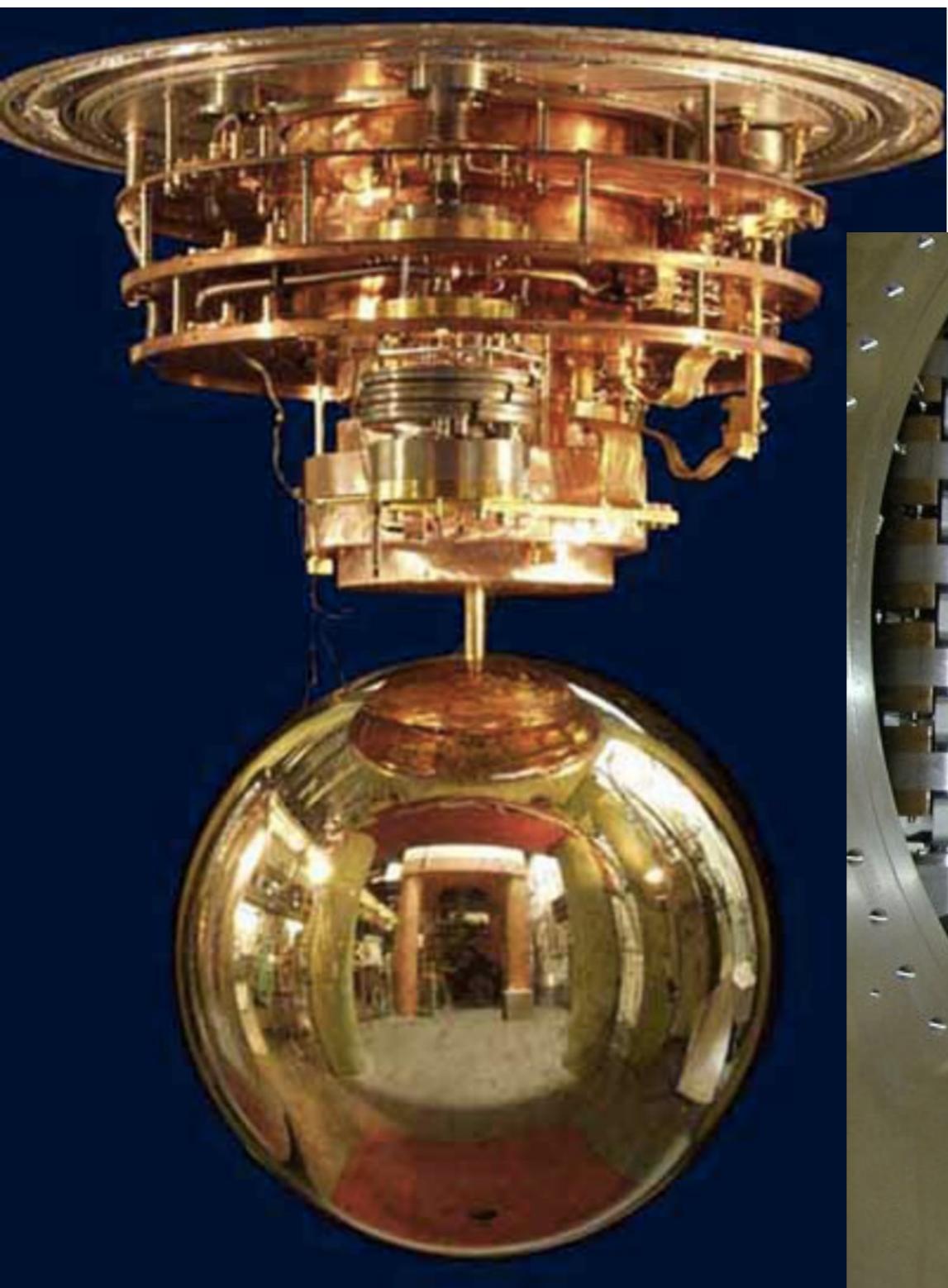
4 cm Niobium cavity
filled with
Superfluid ^4He
 $Q \sim 100$ million



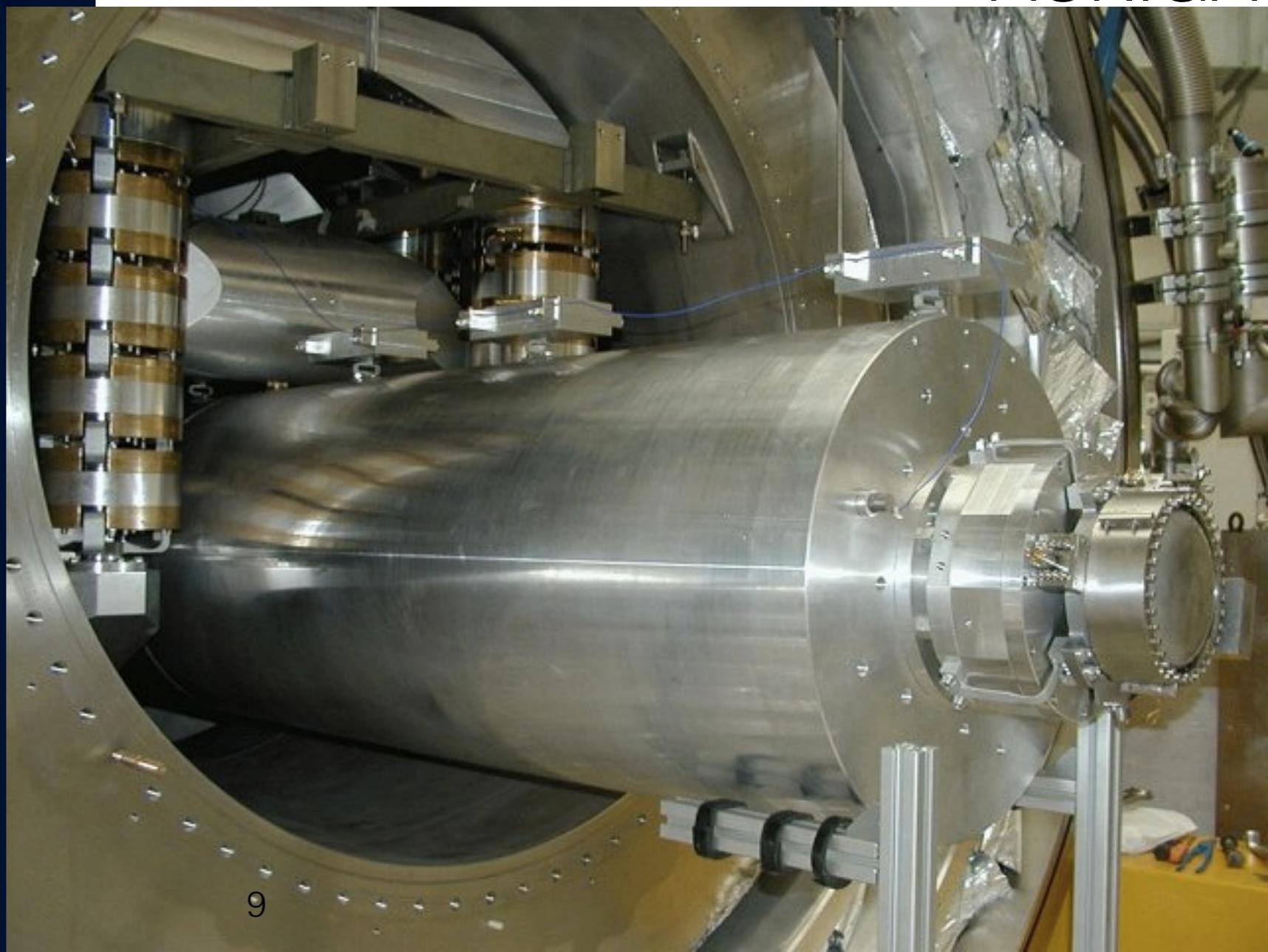
from Laura de Lorenzo

serious Cryogenics

Large GW Detectors < 10 K
care taken to filter seismic
& minimize upconversion



miniGRAIL
@ Leiden



AURIGA

- No serious issues here; this is **NOT** like CERN or KAGRA or dilution fridges

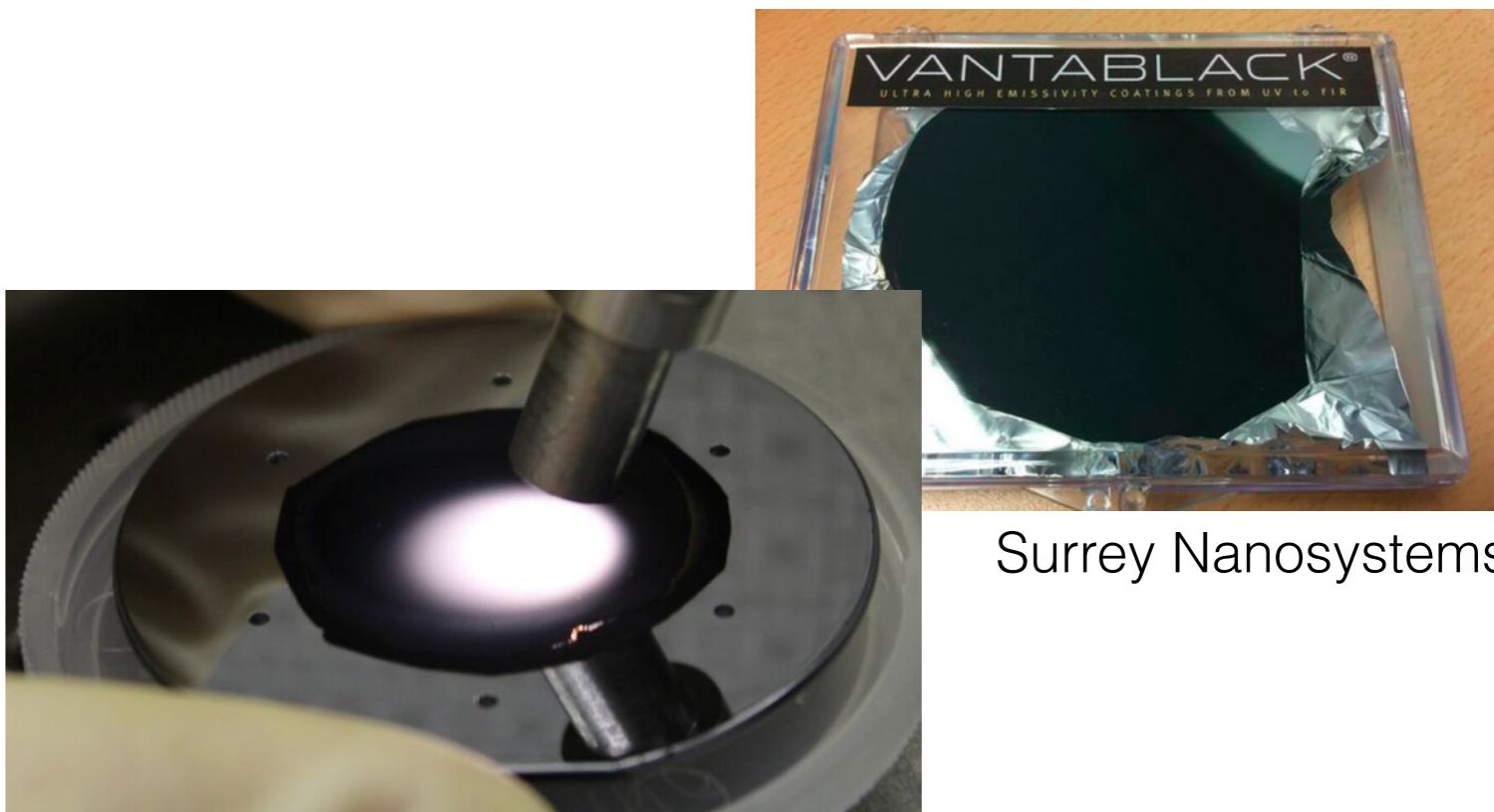
- ~10 W heat extraction capability in steady state

- Prelim mech drawing & backscatter analysis done (Stanford/CIT engineers).

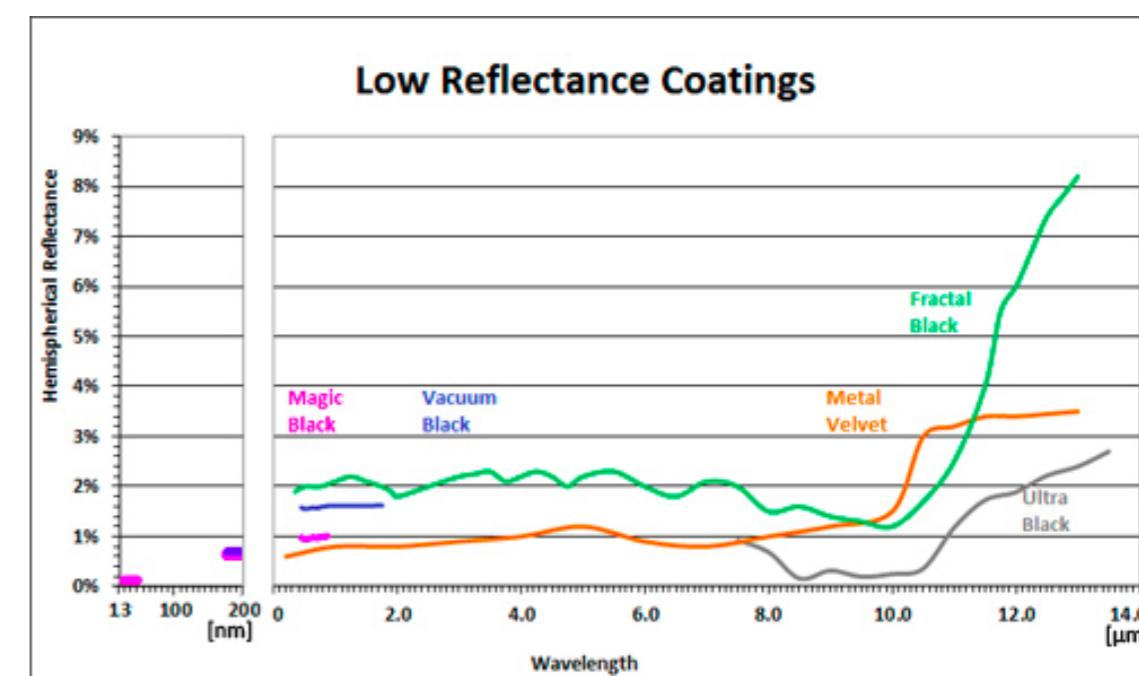
- Vibration from cryogenics no worse than existing cryo pumps.

- How to do initial cool down?
Heat switches?

Cryogenics



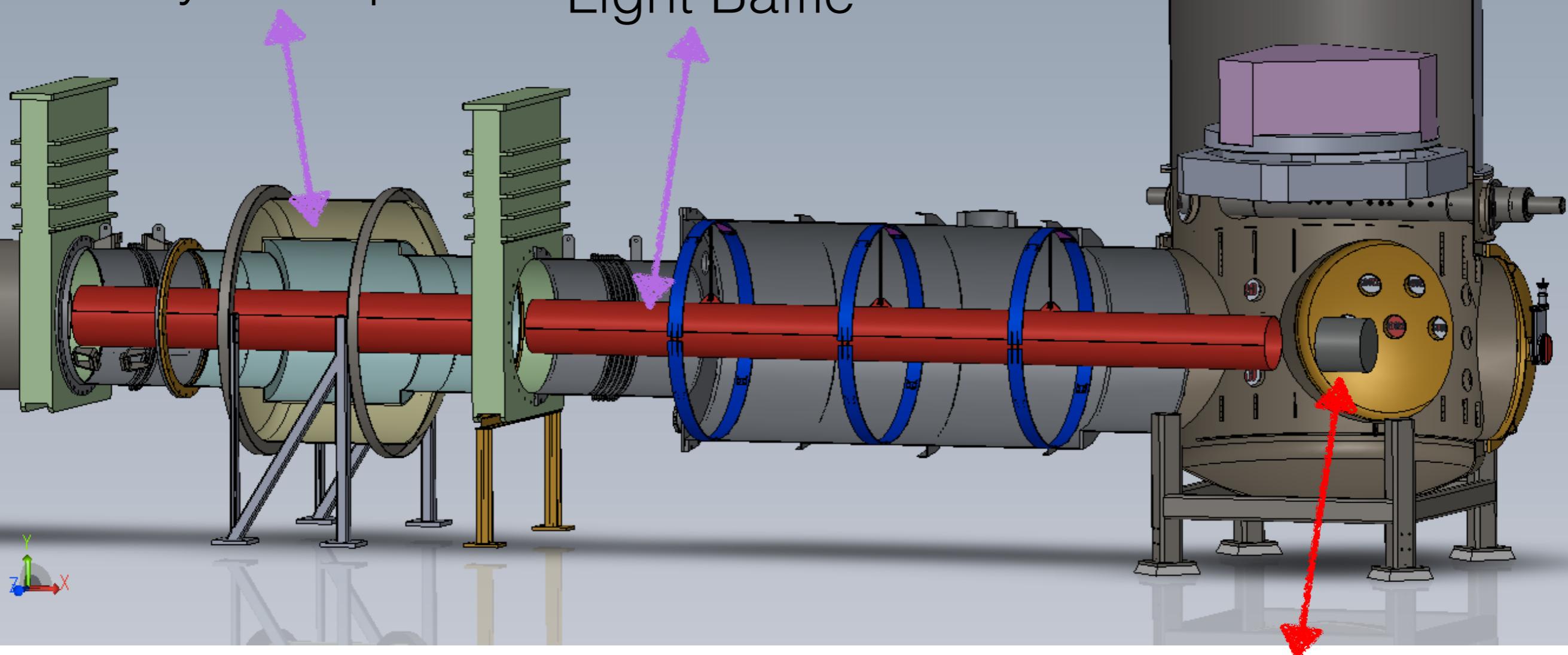
John Hagopian / NASA Goddard



Existing
77 K
Cryo Pump

80 K
Radiation /
Light Baffle

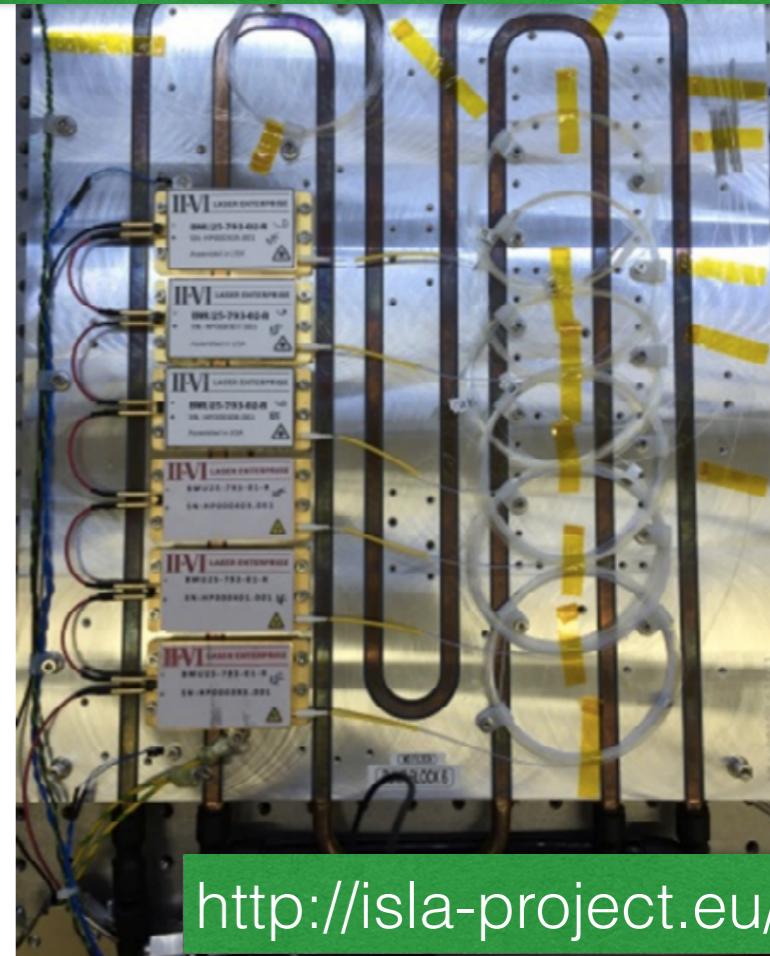
Invisible
Suspension



123 K
Test Mass

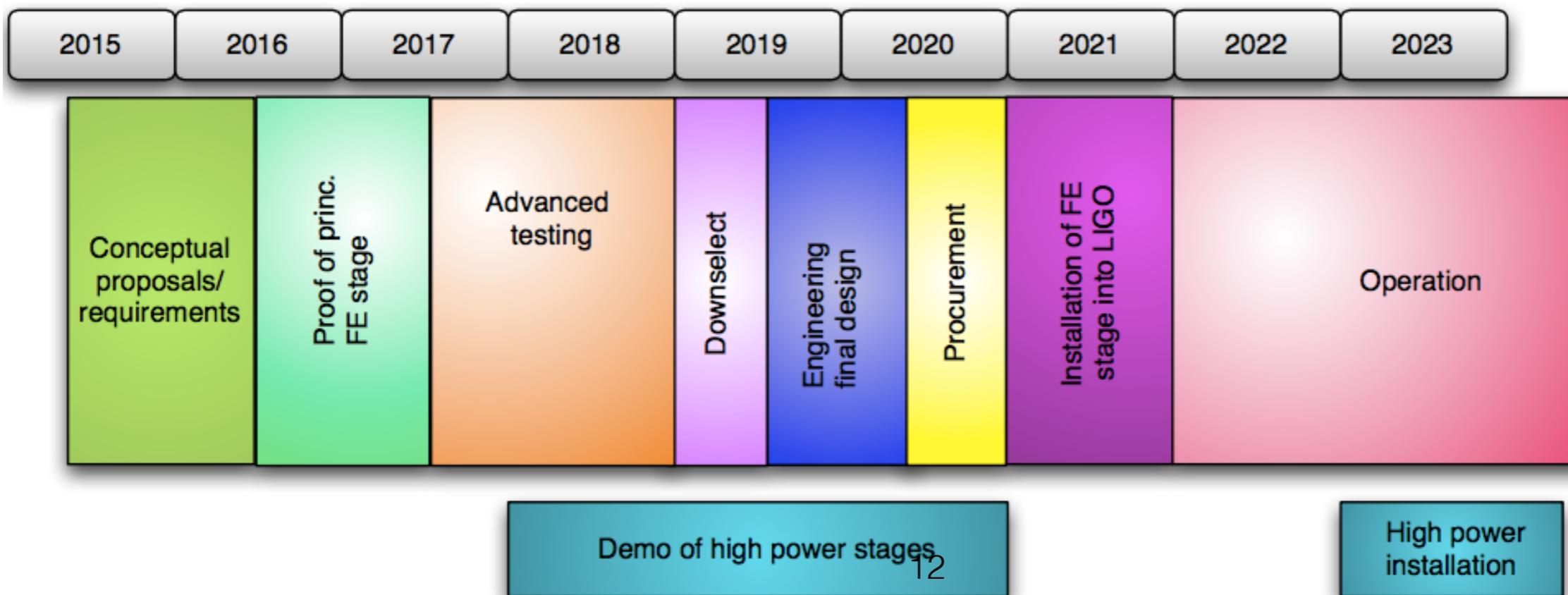
2 micron lasers

- Tm:YAG, Ho:YAG commercial lasers exist (low power, low noise, or high power, high noise)
- Adelaide lasers (Veitch LVC talk)
- Testing at CIT this summer



<http://isla-project.eu/>

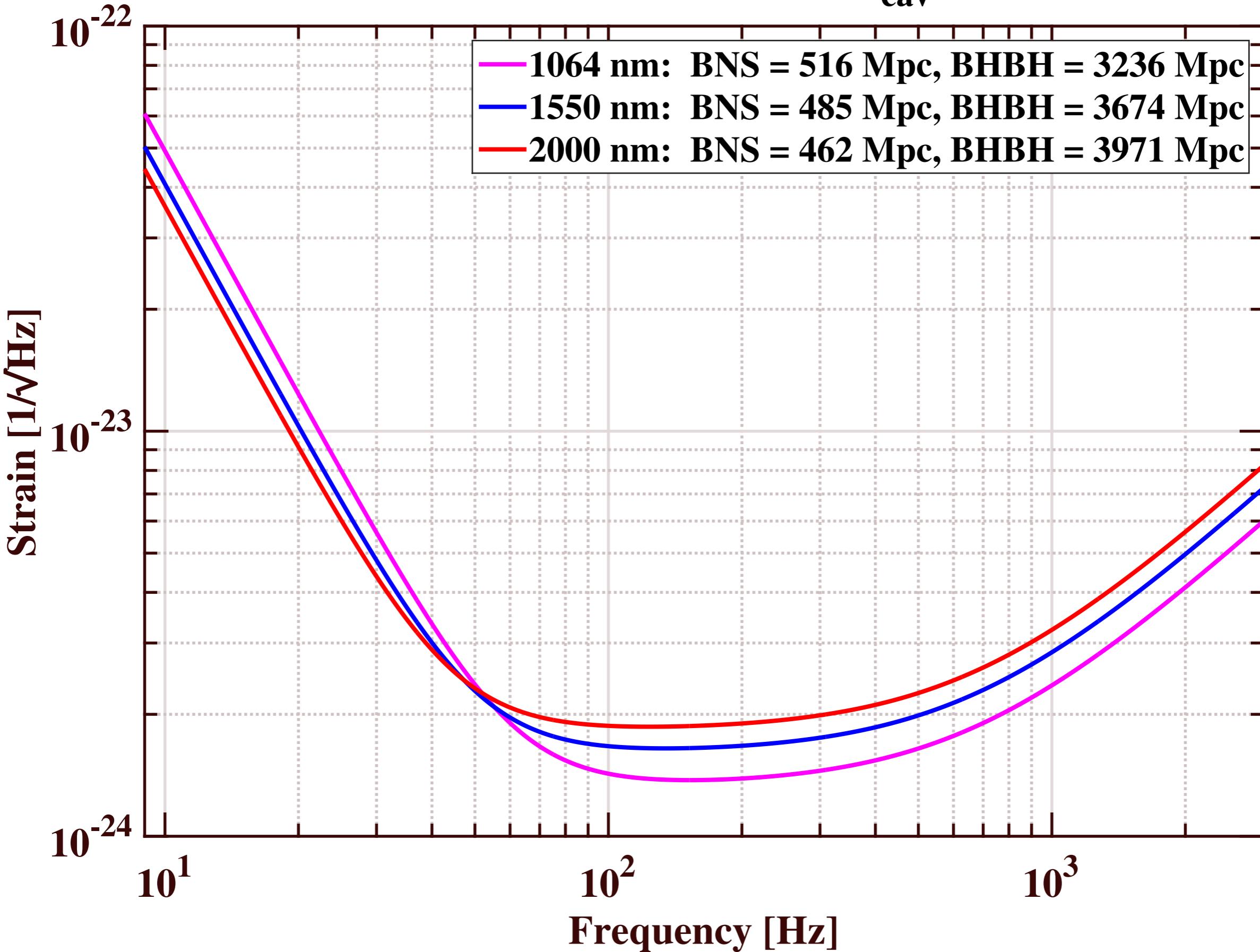
Laser dev Timeline



Wavelength Choice

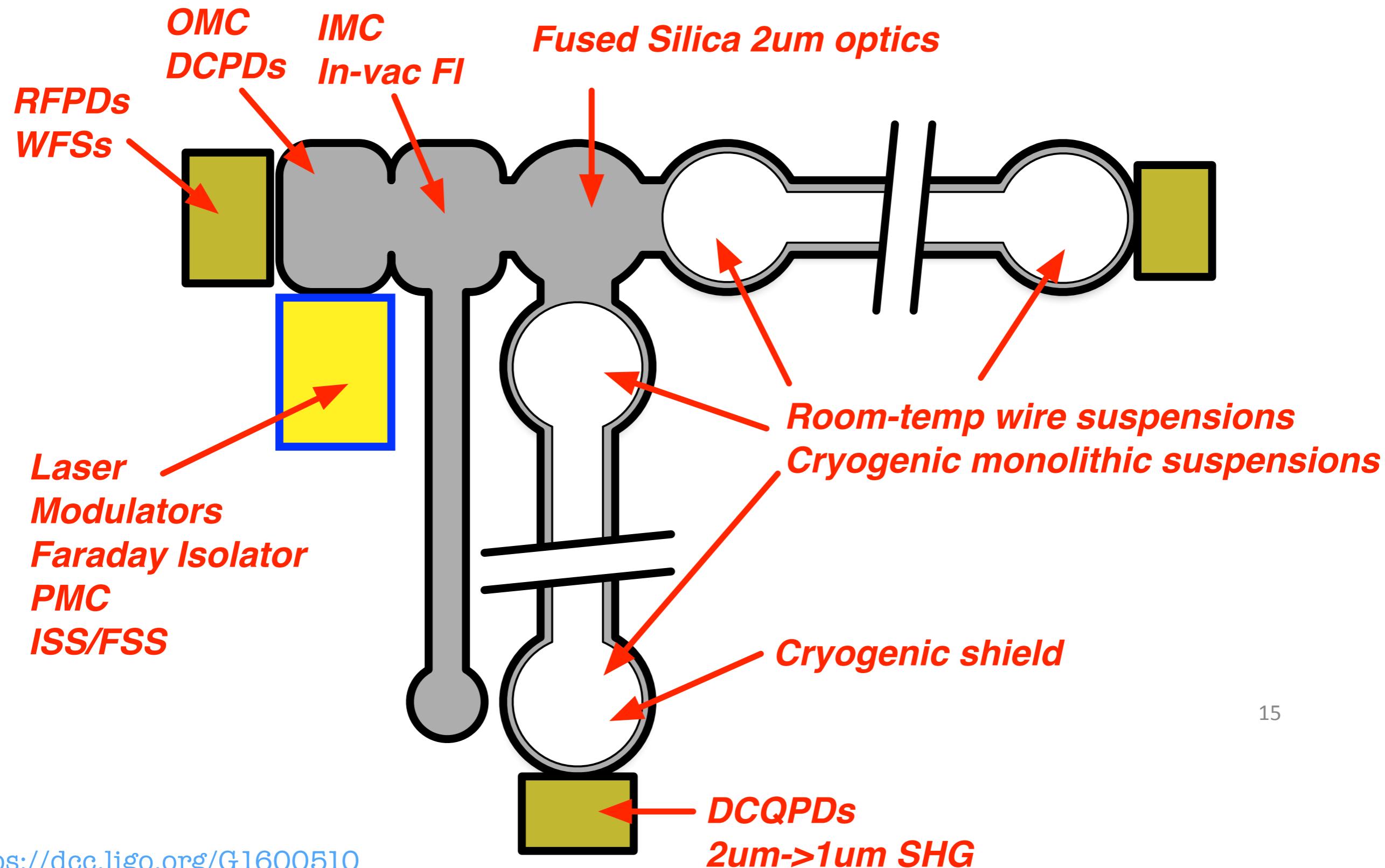
- We know and like 1064 nm. Lots of experience.
- Many new issues with 1.5 - 2.1 microns.
- ~200 W lasers feasible with 1.8 - 2.1 microns
- PD QE > 80% today. No showstoppers yet.
- Scatter loss decreases with wavelength; quantum noise improvement. Increases ultimate reach assuming we solve “*nuisance*” losses (FI, OMC, MM, PBS, clipping, alignment, viewports, etc.)

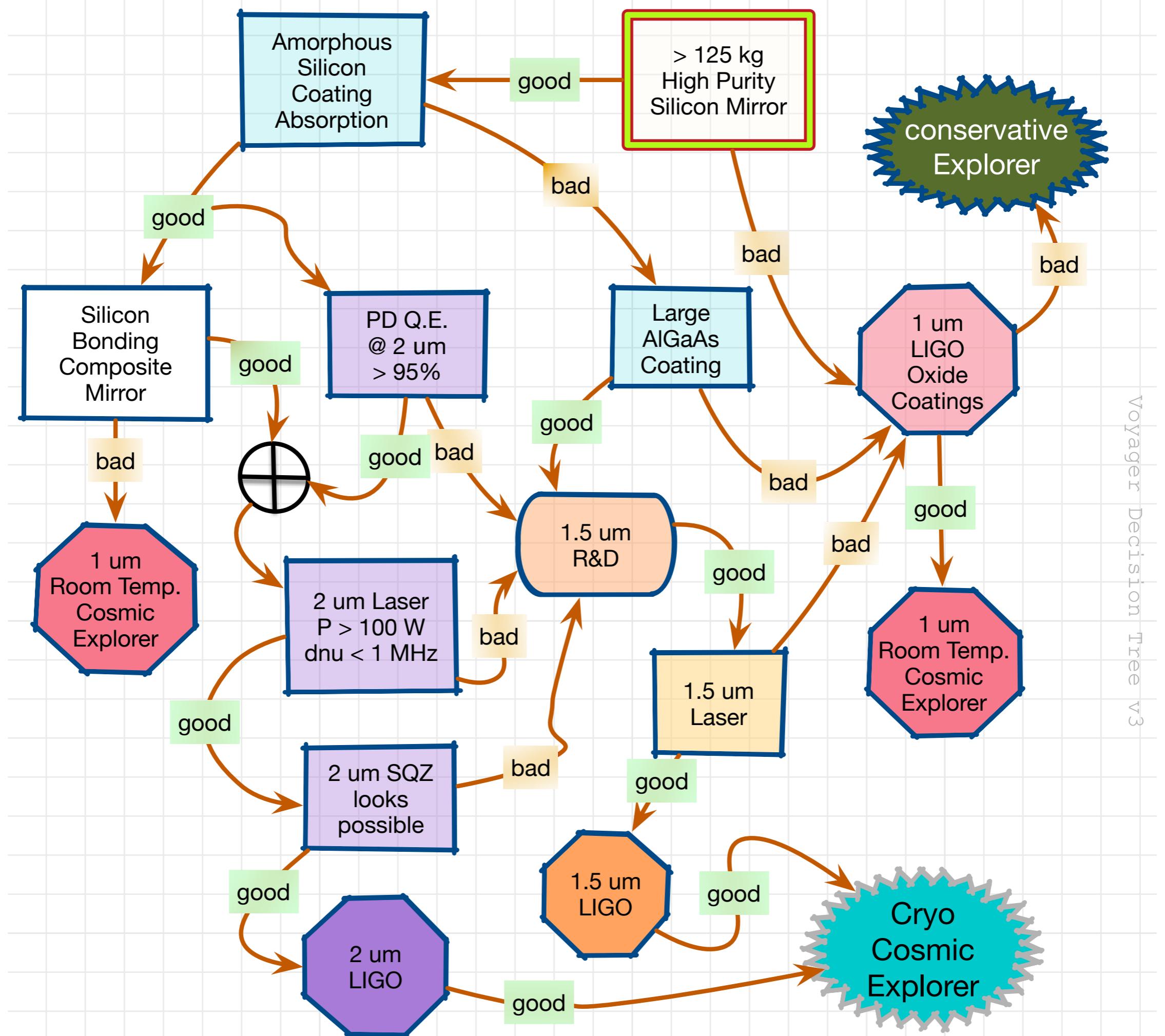
Quantum Noise: $m = 100 \text{ kg}$, $P_{\text{cav}} = 3 \text{ MW}$



1 MW @ 1064 nm ~ 2 MW @ 2128 nm

Prototype Facility





Voyager Decision Tree v3