

aLIGO Design  
Presentation at UF  
October 5th, 2015  
GM

Source Material:

- T0900043
- Optics Express, Vol 16(14), 10018 (2008)
- Appl. Opt. 42(7) 2003, 1244-1256
- Appl. Opt. 42(7) 2003, 1257-1268



# Understanding aLIGO

Generation of signal field:

Phase modulation of cavity internal field





$$E_{Cav} = E_0 e^{i(kL - \omega t)} e^{ik\delta L \sin \Omega t}$$

$$\approx E_0 e^{i(kL - \omega t)} \left[ 1 + \frac{k\delta L}{2} (e^{i\Omega t} - e^{-i\Omega t}) \right]$$



# Understanding aLIGO

Carrier In  
  
  
 SB Out



Signal (outside cavity):

- Amplitude: 
$$E_{Out}^{SB} = t_{Cav}(\Omega_{GW}) \frac{k\delta L}{2} E_0$$

$$t_{Cav} = \frac{it_I}{1 - r_I r_E e^{i\phi_{RT}}} \quad \phi_{RT} = \frac{\Omega}{FSR}$$

- Max-Signal ( $\Omega_{GW}$  small): Impedance matched cavity

$$T_{ITM} = Losses$$



# Understanding aLIGO

Carrier In  
← SB Out



$$T_{ITM} = Losses$$

State of the art cavity losses:

- < 1ppm coating absorption
- ~5-10 ppm transmission
- ~50-100 ppm scatter across beam
- ➔  $T_{ITM} \sim 100\text{ppm}$

Cavity with  $T_{ITM} \sim 100\text{ppm}$ :

- Finesse =  $3 \times 10^4$
- Line width (HWHM)  $\sim 1\text{Hz}$
- ➔ would average out all relevant GW signals



# Advanced LIGO Design

Carrier In  
SB Out



- $T_{ITM}$  is compromise between signal amplitude (gain) and detector band width
- but also reduces carrier inside cavity

$$E_0 = \frac{t_I}{1 - r_I r_E} E_{in}$$



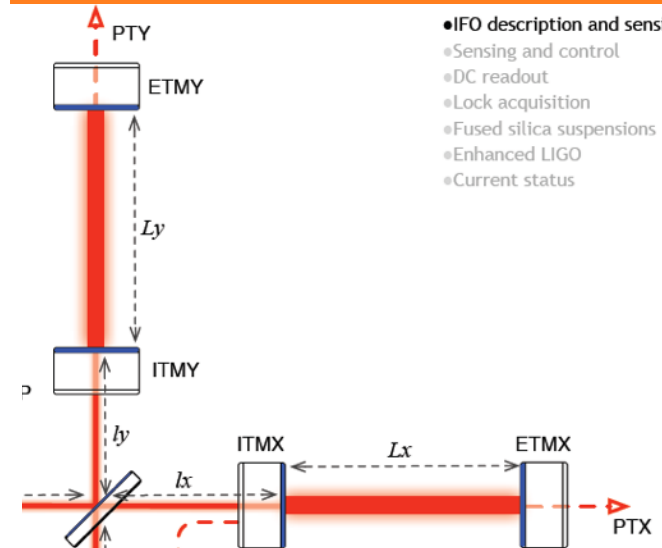
# Advanced LIGO

Take advantage of

- **Michelson Interferometer**
- **Quadrupole nature of GW**
  - **SB generated 180deg out of phase in both arms**

➔ **Tune MI to dark**

- **Carrier goes back to laser**
  - **all laser noise suppressed in dark port**
- **Signal field goes to dark port**
  - **not affected by anything between BS and laser**



- IFO description and sensitivity goals
  - Sensing and control
  - DC readout
  - Lock acquisition
  - Fused silica suspensions
  - Enhanced LIGO
  - Current status

$$E_{refl}^0 = E_{in} \left[ r_{BS}^2 r_{Cav1} e^{ikl_1} + t_{BS}^2 r_{Cav2} e^{ikl_2} \right]$$

$$E_{DP} = E_{signal} \left[ t_{BS} e^{ikl_1} + r_{BS} e^{ikl_2} \right]$$



# Advanced LIGO

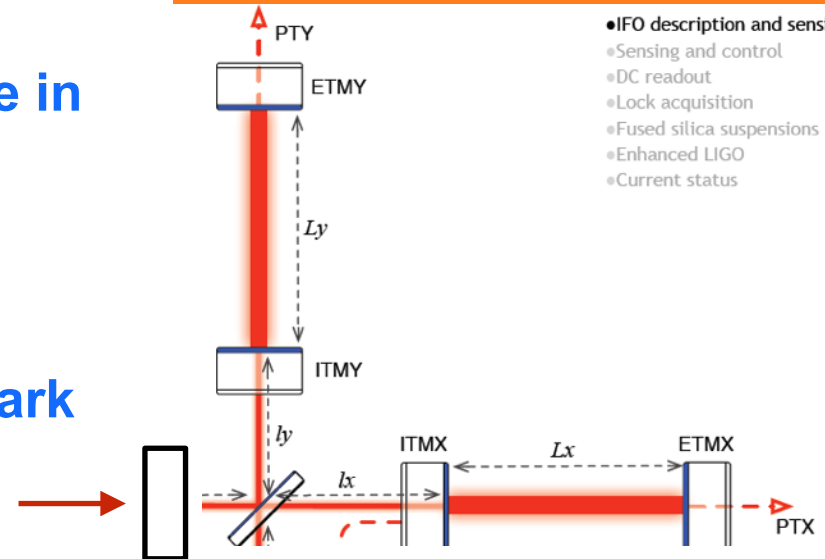
Take advantage of

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➔ **Allows power recycling. Best gain: Impedance matching**



- IFO description and sensitivity goals
  - Sensing and control
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$$R_{MI} \approx R_{Cav} \approx 1 - \frac{4L}{T_I} + \text{MI losses}$$



# Advanced LIGO

Lets put this together:

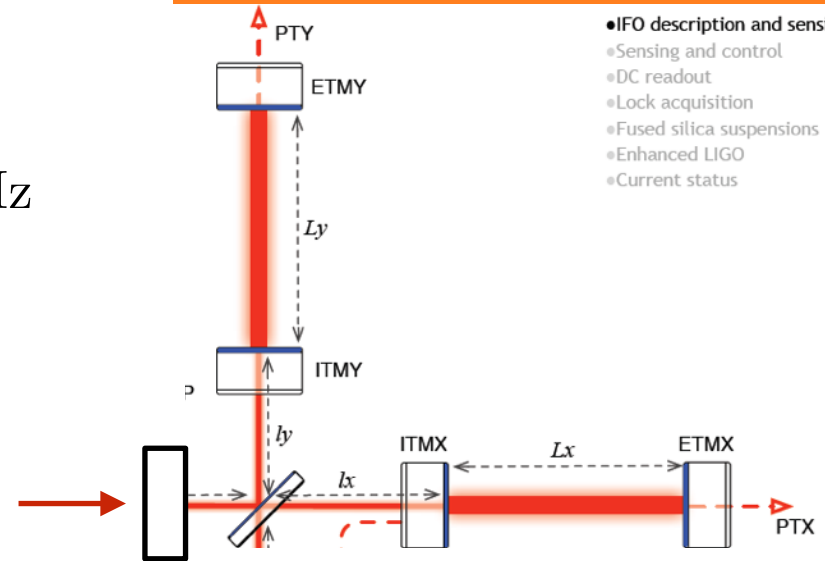
- GW bandwidth sets  $T_{ITM}$

$$HWHM = \frac{FSR}{2F} = \frac{FSR}{2\pi} T_I \approx 100\text{Hz}$$

$$T_I \approx 1.5\%$$

- Anticipated Losses determine  $T_{PR}$

$$T_{PR} \approx \frac{4L}{T_I} \approx \frac{450\text{ppm}}{1.5\%} = 3\%$$





# Advanced LIGO

## Signal Recycling:

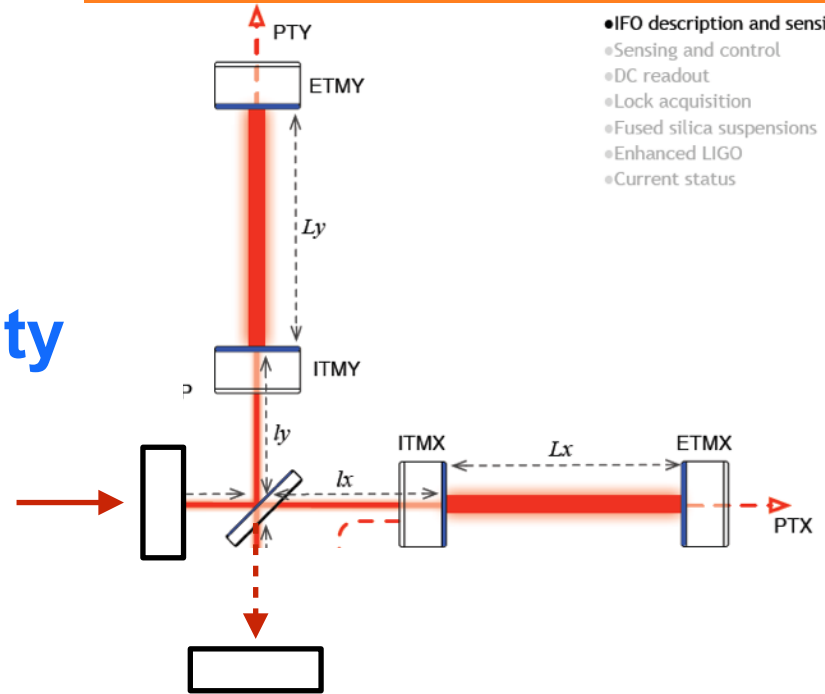
- Coherent amplification or
- Coherent extraction of signal field

## IFO for signal ~ 3 mirror cavity



ITM/SR: Compound mirror

$$r_{CM} e^{i\phi_{CM}} = \frac{r_{ITM} - r_{SR} e^{i\phi_{SR}}}{1 - r_{ITM} r_{SR} e^{i\phi_{SR}}}$$



- IFO description and sensitivity goals
  - Sensing and control
  - DC readout
  - Lock acquisition
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# Advanced LIGO

## Signal Recycling:

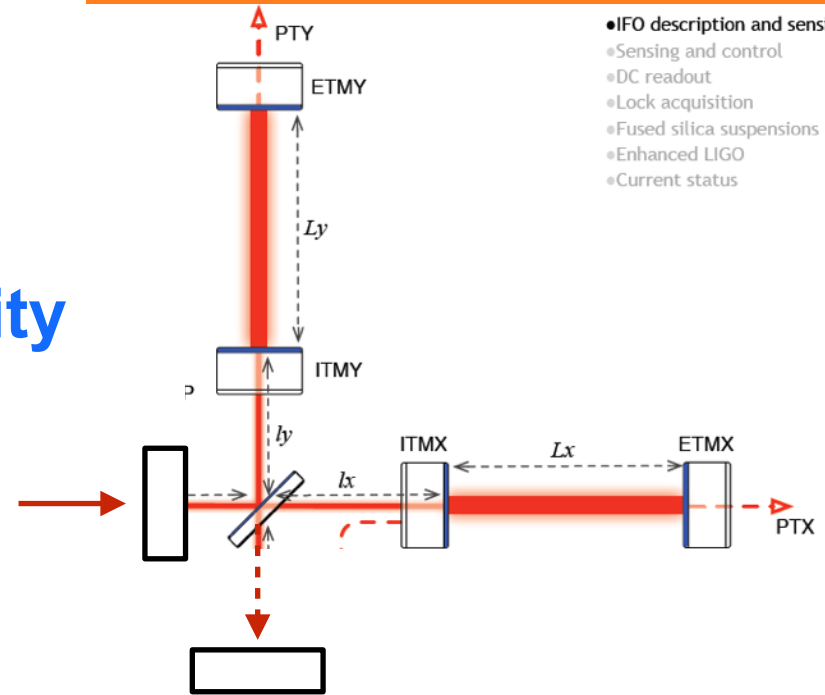
- Coherent amplification or
- Coherent extraction of signal field

## IFO for signal ~ 3 mirror cavity

$$r_{CM} e^{i\phi_{CM}} = \frac{r_{ITM} - r_{SRE} e^{i\phi_{SR}}}{1 - r_{ITM} r_{SRE} e^{i\phi_{SR}}}$$

used to tailor frequency response of aLIGO

$$T_{SR} = 20\% \quad \text{today}$$



- IFO description and sensitivity goals
  - Sensing and control
  - DC readout
  - Lock acquisition
  - Fused silica suspensions
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taking into account all known

- noise sources
- potential signals
- locking issues

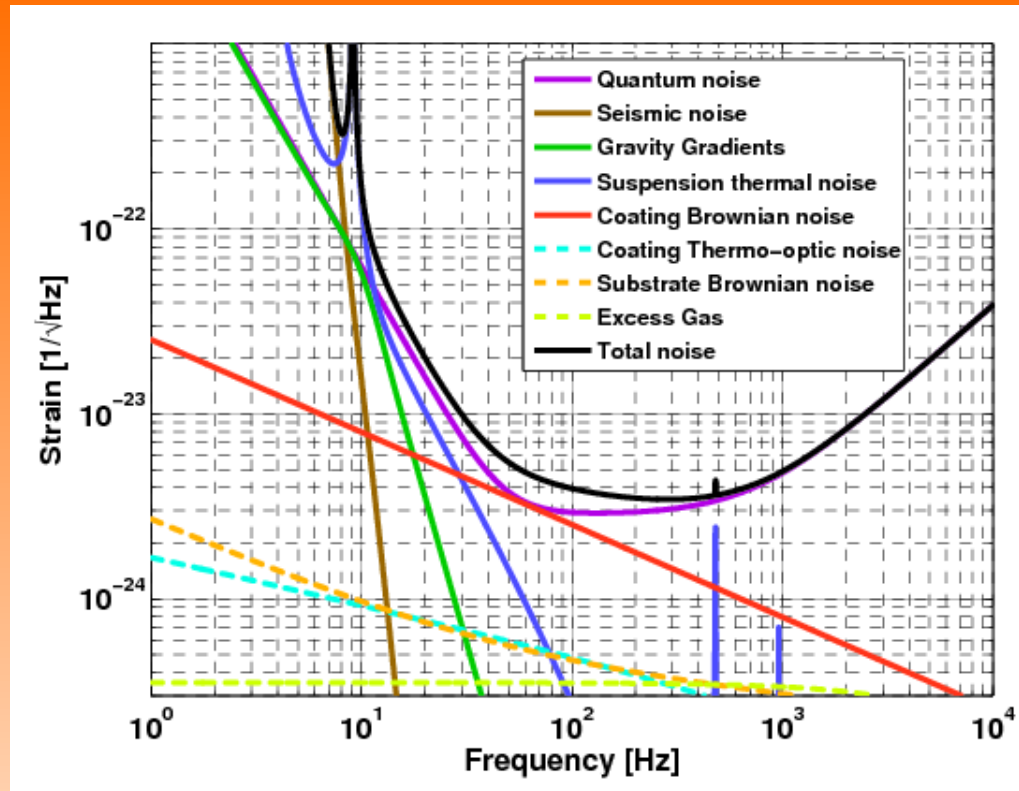


# Advanced LIGO Design

Important mirror parameters:

✓ Transmissivity

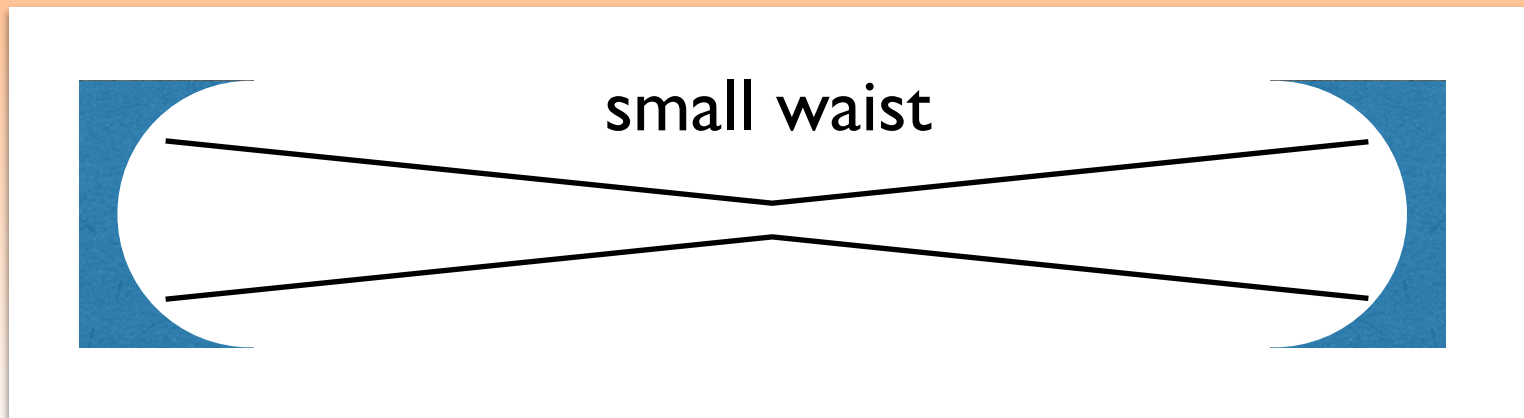
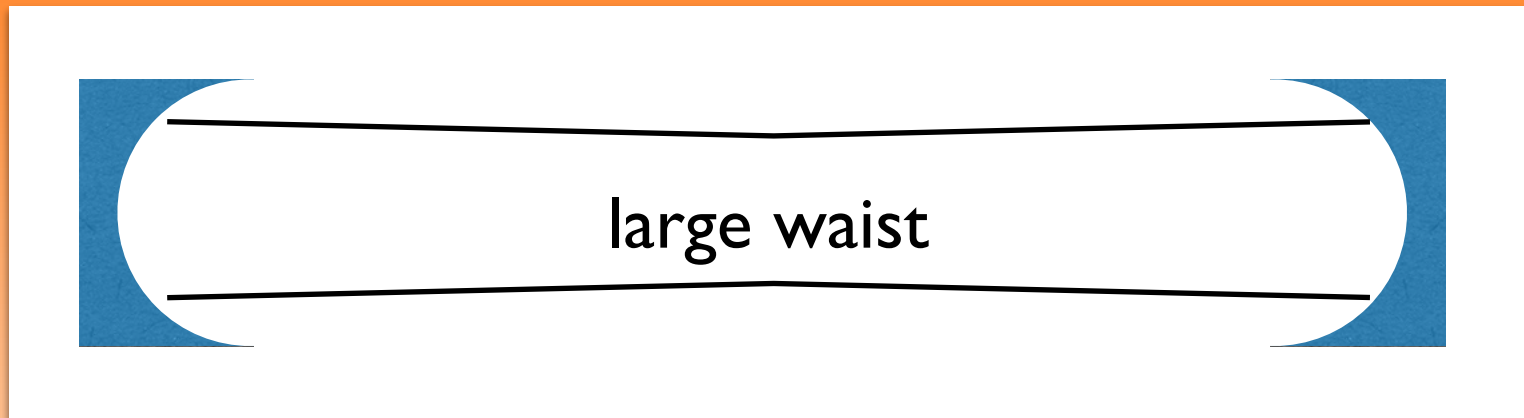
- Mass = 40kg
  - as heavy as necessary
  - as light as possible
  - RPN  $\sim$  Suspension thermal
    - Limits useful mass
- Radius of curvature
  - Increase beam size
    - reduce coating TN
  - limited by
    - diffraction losses to  $\sim 6\text{cm}$
    - cavity stability
    - stability against thermal deformation



# Advanced LIGO Design

Beam size on test masses:

- Two options



# Advanced LIGO Design

Beam size on test masses:

- Two options

- Large:

- Smaller beam size at BS
- ROCs difficult to measure  $\sim 52\text{km}$        $\delta s \approx 36 \text{ nm}$

- Small:

- Stable against ROC increases due to heating
- ROCs easier to measure  $\sim 2080\text{m}$        $\delta s \approx 1 \mu\text{m}$

More coating layers on ETM (increases thermal noise)

Diffraction larger issue at beam splitter:

- Increase beam size at ETM to  $\sim 6.3\text{cm}$
- Decrease beam size at ITM to  $\sim 5.7\text{cm}$

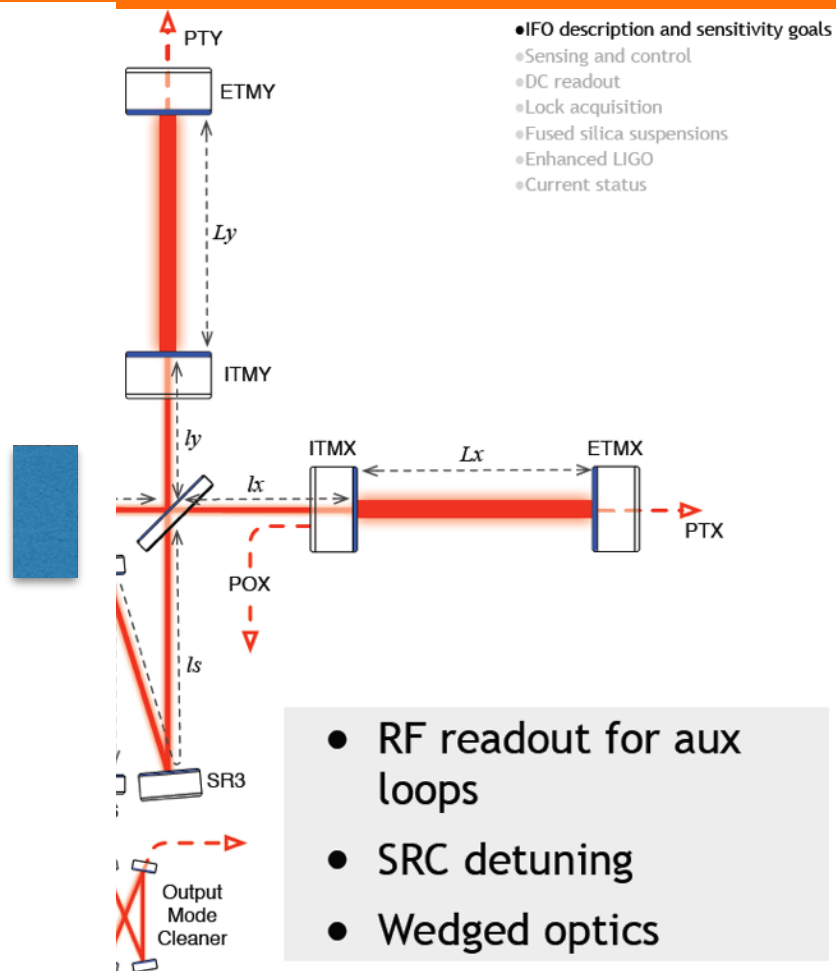
Final design:     $R_{\text{ETM}} = 2245\text{m}$        $R_{\text{ITM}} = 1934\text{m}$



# Advanced LIGO Design

## Recycling cavity design:

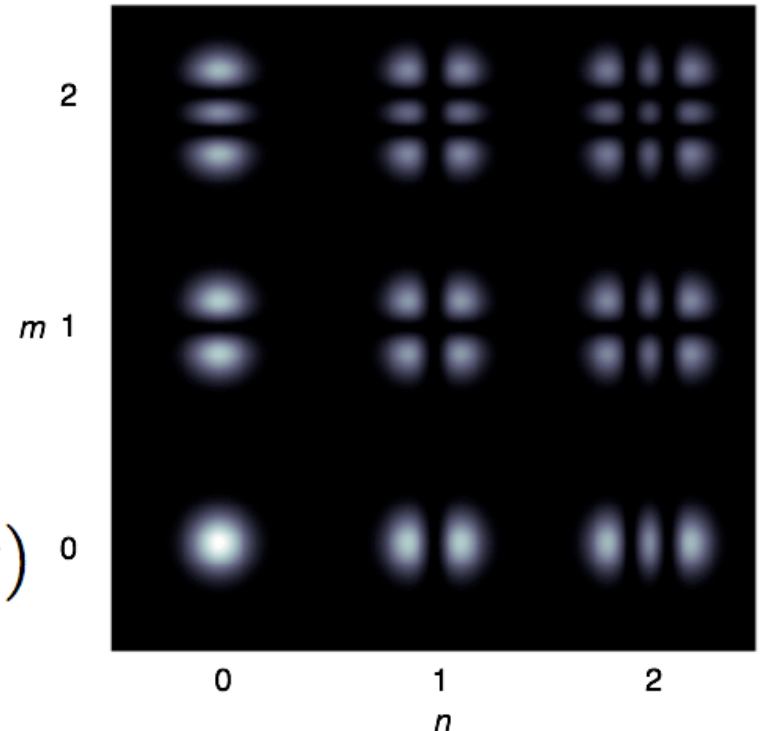
- Two options
- Simple, but barely stable:
  - $R \sim 1400\text{m}$
  - $g = 1 - 20/4000 \sim 1$  (nearly unstable)
    - higher order modes near resonant
    - bad for alignment sensing
    - very power dependent eigenmodes
    - bad experience in iLIGO
- Folded recycling cavity
  - $g$  selectable
  - HOMs non-resonant
  - nearly power independent eigenmodes



# Higher-order HG modes

- Spatial field distributions which reproduce after cavity roundtrip = spatial eigenmodes
- Have different resonance frequencies (except for  $g \sim 1$ )
- Separable in  $x$  and  $y =$  HG-modes

$$u_{nm}(x, y, z) = u_n(x, z)u_m(y, z)$$



$$u_{nm}(x, y, z) = \left(2^{n+m-1} n! m! \pi\right)^{-1/2} \frac{1}{w(z)} \exp(i(n+m+1)\Psi(z)) \times \\ H_n\left(\frac{\sqrt{2}x}{w(z)}\right) H_m\left(\frac{\sqrt{2}y}{w(z)}\right) \exp\left(-i\frac{k(x^2+y^2)}{2R_C(z)} - \frac{x^2+y^2}{w^2(z)}\right)$$

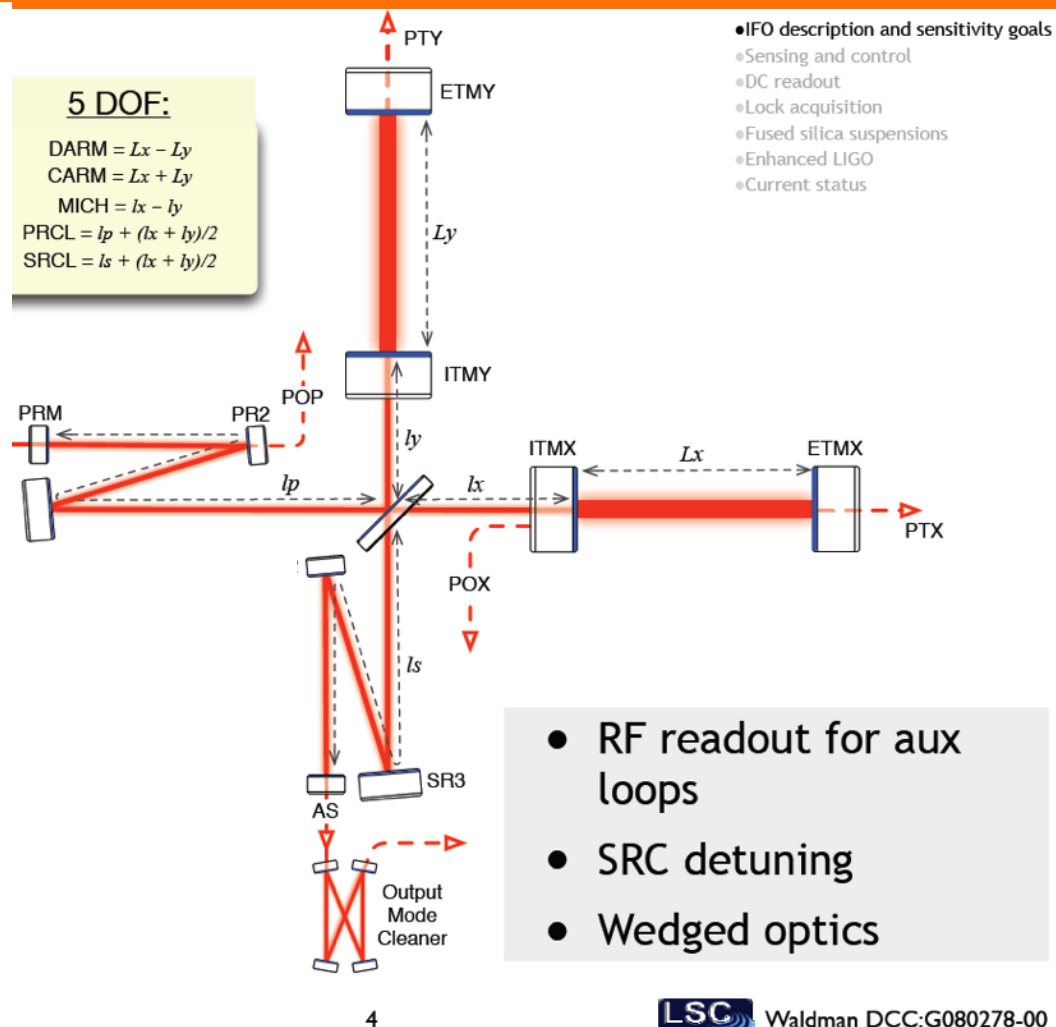
$$\begin{aligned} H_0(x) &= 1 & H_1(x) &= 2x \\ H_2(x) &= 4x^2 - 2 & H_3(x) &= 8x^3 - 12x \end{aligned}$$



# Advanced LIGO Design

## Recycling cavity design:

- Two options
- Folded recycling cavity
  - g selectable
  - HOMs non-resonant
  - nearly power independent eigenmodes
- PR3 focuses ~6cm beam
  - waist after PR2
- Stable mode between PR2 and PRM (sets Gouy phase)
- MM very sensitive to PR2/PR3 distance/ROCs
- SR-cavity similar





# Advanced LIGO Design

Length sensing basics:

- Paul will discuss in more detail

Important for design:

Each degree of freedom has to be sensed by some field

- Requires resonant SBs inside recycling cavities
- $f_1$  for power recycling
- $f_2$  for signal recycling but has to pass through power recycling and MI



# Advanced LIGO Design

## Power Rec. cavity design:

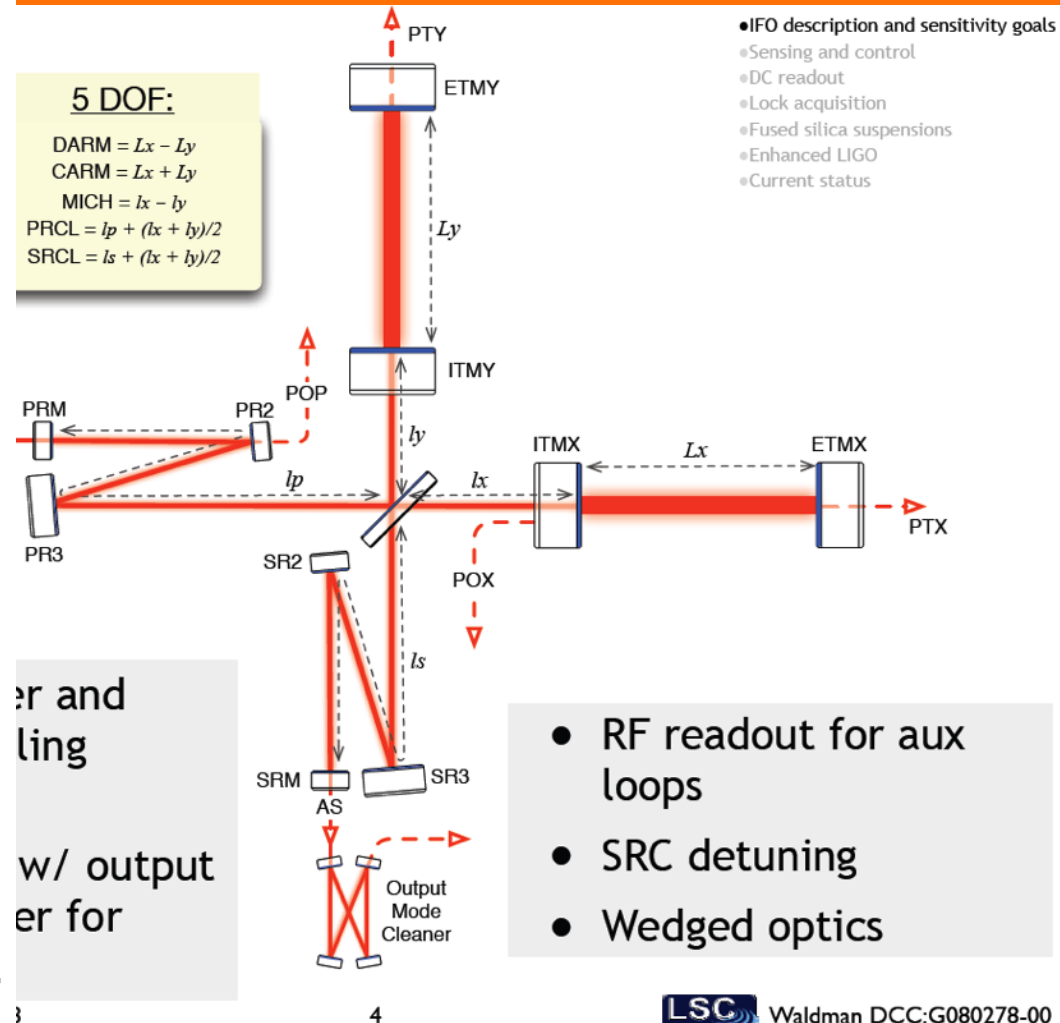
- Length ~55m depends on Vacuum envelope
  - ~16m between PRM/PR2
    - HAM2/3 distance
    - times 3
  - + HAM3-ITM distance
  - incl. BS/ITM substrates

gives modulation frequencies:

$$f = (N + 0.5) \times \frac{c}{2L}$$

$$N = 3 \quad f \approx 9.1 \text{ MHz}$$

0.5: because carrier is resonant in arm cavities



# Advanced LIGO Design

## Signal Rec. cavity design:

- Length  $\sim 55\text{m}$
- SB has to pass through PR

## Two ways:

- Make MI bright for 2nd SB
- Couple both cavities



$r_{PR}$



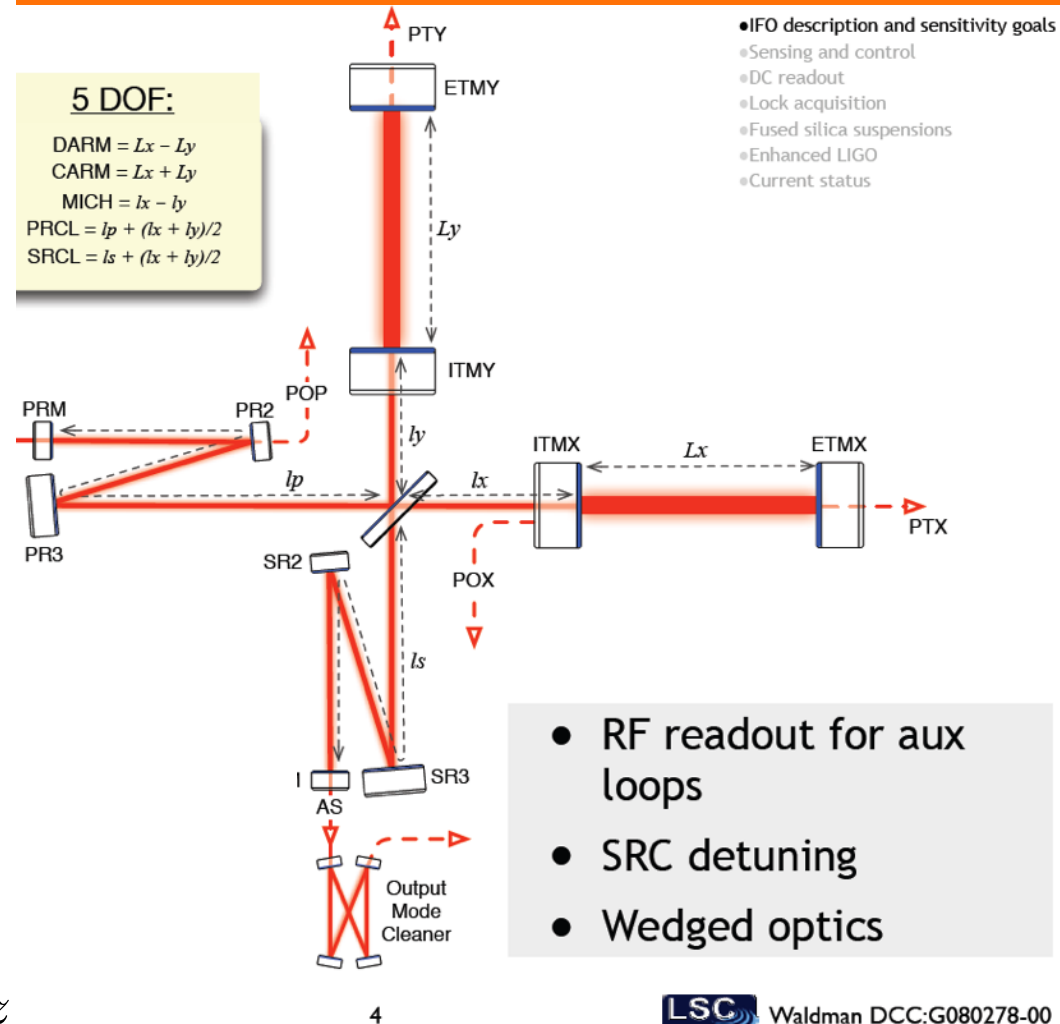
$r_{BS}$



$r_{SR}$

$$t_{BS} = 2\pi f_{Mod} \frac{2\Delta L}{c}$$

$$\Delta L \approx 4 \text{ cm} \quad f_{Mod} = 45.5 \text{ MHz}$$



# Advanced LIGO: Design

## Summary:

- Transmission of ITM set by detector bandwidth
- Power recycling gain set by losses in arm cavities and MI
- Signal recycling mirror forms compound mirror with ITM to
  - tailor frequency response. Choices between:
    - High gain/low BW
    - low gain/high BW
      - ampl. displacement noise, only useful where shot noise limited
- Stable recycling cavities to
  - contain spatial mode
  - reduce scatter into other spatial modes
- Length sensing:
  - Need sidebands to probe recycling cavities



