

# Scaling up coating Brownian noise measurements

M. Abernathy

July 17, 2015

## 1 Intro

The purpose of this document is to calculate levels of coating Brownian thermal noise given laboratory direct measurements at smaller scales. We will do this by applying some scaling rules derived from coating Brownian noise theory. By the end of this document, we should have calculated the level of coating Brownian thermal noise expected in the aLIGO interferometers if they were coated with the same coatings as measured in the TNI [1, 2, 3] and in the Caltech rigid cavity measurement (TNC) [4, 5], but with a film thickness scaled to be comparable with the HR films on the ETMs and ITMs in aLIGO. We can then use these values to compare with the noise levels seen in the aLIGO interferometers to determine if they have reached coating thermal noise limitations.

## 2 Lab Measurements

### 2.1 TNI

The bulk of the information on the TNI can be found in reference [1]. However, I had to dig up the mirror dimensions from the DCC [6]. The mirrors are 4" in diameter and 4" in height. The beam waist,  $w_0$ , was about 160 *um* for all measurements. Four coatings were measured in the TNI:

1. Quarter-wave stack: silica/tantala
2. Thermal noise optimized stack: silica/tantala
3. Quarter-wave stack: silica/Ti:tantala
4. Thermal noise optimized dichromic stack: silica/Ti:tantala.

These are listed as TNI(1-4) in table 1. All are on silica substrates

The values for  $S_x(100\text{Hz})$  were calculated by taking the values for  $\phi_{SiO_2}$ ,  $\phi_{Ta_2O_5}$ , and  $\phi_{TiO_2:Ta_2O_5}$  given in tables IV and V of [1], and plugging them back into equations (1) and (2) of that same paper. This should give a good approximation of the actual value of  $S_x$  they would have measured.

Optic	$w_0$ [ $\mu\text{m}$ ]	$a$ [cm]	$d$ [ $\mu\text{m}$ ]	$H$ [cm]	$R$	$S_x(100\text{Hz})$ [ $\text{m}^2 \text{Hz}^{-1}$ ]
TNI(1)	160	5.08	4.55	10.16	??	1.06e-35
TNI(2)	160	5.08	5.41	10.16	??	8.82e-36
TNI(3)	160	5.08	4.21	10.16	??	7.79e-36
TNI(4)	160	5.08	3.81	10.16	??	6.84e-36
TNC	182	1.27	4.53	0.635	??	1.1e-35
ITM	53000	17	2.8	20	??	??
ETM	62000	17	5.9	20	??	??

Table 1: Values useful for converting coating Brownian noise. TNI(1-4) refers to the four different coatings measured in the Thermal Noise Interferometer at Caltech [1, 2, 3]. TNC refers to the reference cavity measurements made at Caltech by Tara Chalermongsak [4, 5]. Values for ITM and ETM are for the LIGO Input Test Mass and End Test Mass, respectively [7, 8]

## 2.2 TNC

These measurements are nicely covered in [4], with the exception of the substrate thickness,  $H$ , which had to be found in [5]. The silica substrates were 1" in diameter and 1/4" in thickness. The film measured here was an old quarter-wave stack produced by REO, made of silica/tantala. To get the value of  $S_x(100\text{Hz})$ , I plugged their measured value of  $\phi_c = 4.43 \times 10^{-4}$  into their equation (8). Again, this should give an approximation of the coating Brownian noise they would have measured at 100 Hz.

## 3 Scaling Rules

I think we need to apply two scaling rules, first for the beam-spot and substrate size, and then for the film thickness. For the beam-spot and substrate dependence, we rely on the work by Somiya and Yamamoto [9]. In that paper, they calculate the Brownian noise of a coating on a finite-sized substrate, which we can write as:

$$S_x^{\text{FIN}}(\Omega) = \frac{8k_B T}{\Omega} \Phi_c U^{\text{FIN}} \quad (1)$$

(from their equation (3)). Here,  $S_x^{\text{FIN}}(\Omega)$  is the coating Brownian noise for a finite size mirror at frequency  $\Omega$ ,  $k_B$  is Boltzmann's constant,  $T$  is the temperature of the optic,  $\Phi_c$  is the mechanical loss of the coating material, and  $U^{\text{FIN}}$  is the energy stored in the film from an imaginary force used to probe the fluctuations (in the shape of the beam intensity). We can then take the ratio of this value to that of a coating with the same properties on an infinite substrate  $S_x^{\text{INF}}(\Omega)$ :

$$R = \frac{S_x^{\text{FIN}}(\Omega)}{S_x^{\text{INF}}(\Omega)}, \quad (2)$$

where  $S_x^{\text{INF}}(\Omega)$  is given by their equation (29):

$$S_x^{\text{INF}}(\Omega) = \frac{4k_B T}{\Omega} \frac{d}{\pi w_0^2} \Phi_c \frac{Y_c^2(1 + \nu_s)^2(1 - 2\nu_s)^2 + Y_s^2(1 + \nu_c)^2(1 - 2\nu_c)}{Y_s^2 Y_c(1 - \nu_c^2)}. \quad (3)$$

Here,  $Y$  is the Young's modulus and  $\nu$  is the Poisson ratio, subscripts c and s indicate properties of the coating and substrate, respectively, and  $d$  indicates the film thickness. Combining equations 1, 2, and 3, we come to the relation:

$$R = 2U^{\text{FIN}} \frac{\pi w_0^2}{d} \left[ \frac{Y_c^2(1 + \nu_s)^2(1 - 2\nu_s)^2 + Y_s^2(1 + \nu_c)^2(1 - 2\nu_c)}{Y_s^2 Y_c(1 - \nu_c^2)} \right]^{-1}. \quad (4)$$

*The attached Mathematica notebook is written to calculate this ratio using the equations in [9], but for some reason, it isn't giving reasonable values*

Once we are able to calculate  $R$ , we should make them for each of the laboratory measurements, as well as for the ITM and ETM mirrors, but assuming the same thickness as the laboratory measurements. For example, we can calculate  $R_{\text{TNI}(1)}$ , the ratio of the TNI(1) measurement to it's imaginary infinite substrate case, and  $R_{\text{ITM}}(d = d_{\text{TNI}(1)})$ , which will be the same ratio for the noise we expect from the ITM mirror if it had a coating the same thickness as the TNI(1) measurement. Then we have the following:

$$R_{\text{TNI}(1)} = \frac{S_{\text{TNI}(1)}^{\text{measured}}}{S^{\text{INF}}(d = d_{\text{TNI}(1)})} \quad (5)$$

$$R_{\text{ITM}}(d = d_{\text{TNI}(1)}) = \frac{S_{\text{ITM}}^{\text{predicted}}(d = d_{\text{TNI}(1)})}{S^{\text{INF}}(d = d_{\text{TNI}(1)})} \quad (6)$$

$$S_{\text{ITM}}^{\text{predicted}}(d = d_{\text{TNI}(1)}) = S_{\text{TNI}(1)}^{\text{measured}} \frac{R_{\text{ITM}}(d = d_{\text{TNI}(1)})}{R_{\text{TNI}(1)}} \quad (7)$$

This line of reasoning only gets us halfway there, as we're really interested in  $S_{\text{ITM}}^{\text{predicted}}(d = d_{\text{ITM}})$ . This is easy to scale to, since the noise is directly proportional to the coating thickness, as in equation 3. Therefore, we can make the final scaling using the relation:

$$S_{\text{ITM}}^{\text{predicted}} = \frac{d_{\text{ITM}}}{d_{\text{TNI}(1)}} S_{\text{ITM}}^{\text{predicted}}(d = d_{\text{TNI}(1)}), \quad (8)$$

which is exactly the value we are trying to calculate.

Looking back, we can actually skip the calculation of  $S^{\text{INF}}$ , as it cancels in all of the equations, and the ratio of  $R$ s in equation 7 can be replaced with the ratios of  $U^{\text{FIN}}$ s. We can also combine equation 8 with equation 7 to get a final equation:

$$S_{\text{ITM}}^{\text{predicted}} = \frac{d_{\text{ITM}}}{d_{\text{TNI}(1)}} \frac{U_{\text{ITM}}(d = d_{\text{TNI}(1)})}{U_{\text{TNI}(1)}} S_{\text{TNI}(1)}^{\text{measured}}. \quad (9)$$

## References

- [1] Maria Principe, Innocenzo M. Pinto, Vincenzo Pierro, Riccardo DeSalvo, Ilaria Taurasi, Akira E. Villar, Eric D. Black, Kenneth G. Libbrecht, Christophe Michel, Nazario Morgado, and Laurent Pinard. Material loss angles from direct measurements of broadband thermal noise. *Phys. Rev. D*, 91:022005, Jan 2015.
- [2] Akira E. Villar, Eric D. Black, Riccardo DeSalvo, Kenneth G. Libbrecht, Christophe Michel, Nazario Morgado, Laurent Pinard, Innocenzo M. Pinto, Vincenzo Pierro, Vincenzo Galdi, Maria Principe, and Ilaria Taurasi. Measurement of thermal noise in multilayer coatings with optimized layer thickness. *Phys. Rev. D*, 81:122001, Jun 2010.
- [3] Eric D. Black, Akira Villar, Kyle Barbary, Adam Bushmaker, Jay Heefner, Seiji Kawamura, Fumiko Kawazoe, Luca Matone, Sharon Meidt, Shanti R. Rao, Kevin Schulz, Michael Zhang, and Kenneth G. Libbrecht. Direct observation of broadband coating thermal noise in a suspended interferometer. *Physics Letters A*, 328(1):1 – 5, 2004.
- [4] Tara Chalermongsak, Frank Seifert, Evan D Hall, Koji Arai, Eric K Gustafson, and Rana X Adhikari. Broadband measurement of coating thermal noise in rigid fabryprot cavities. *Metrologia*, 52(1):17, 2015.
- [5] Tara Chalermongsak. *High fidelity probe and mitigation of mirror thermal fluctuations*. PhD thesis, California Institute of Technology, 2014.
- [6] E. Black. TNI design outline. LIGO DCC:T980052.
- [7] G. Billingsly. personal communication.
- [8] W. Kells G. Billingsley, G. Harry. Core optics components design requirements document. LIGO DCC:T000127.
- [9] Kentaro Somiya and Kazuhiro Yamamoto. Coating thermal noise of a finite-size cylindrical mirror. *Phys. Rev. D*, 79:102004, May 2009.