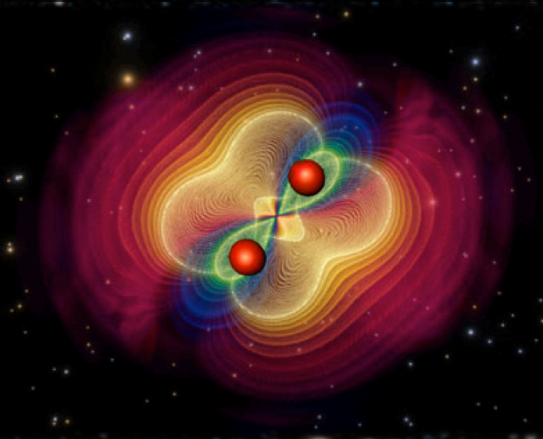
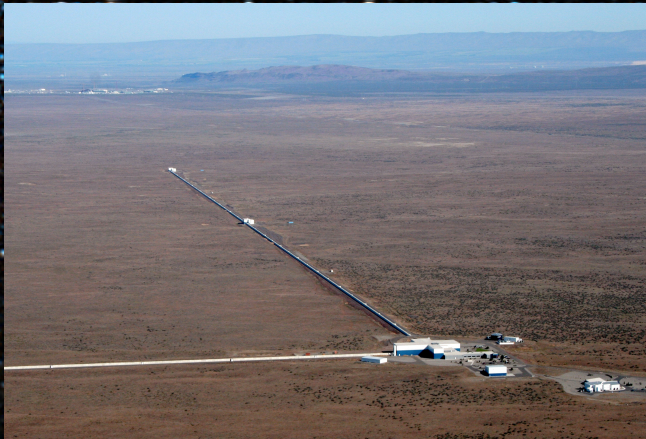


Laying a foundation for confident transient gravitational wave observations

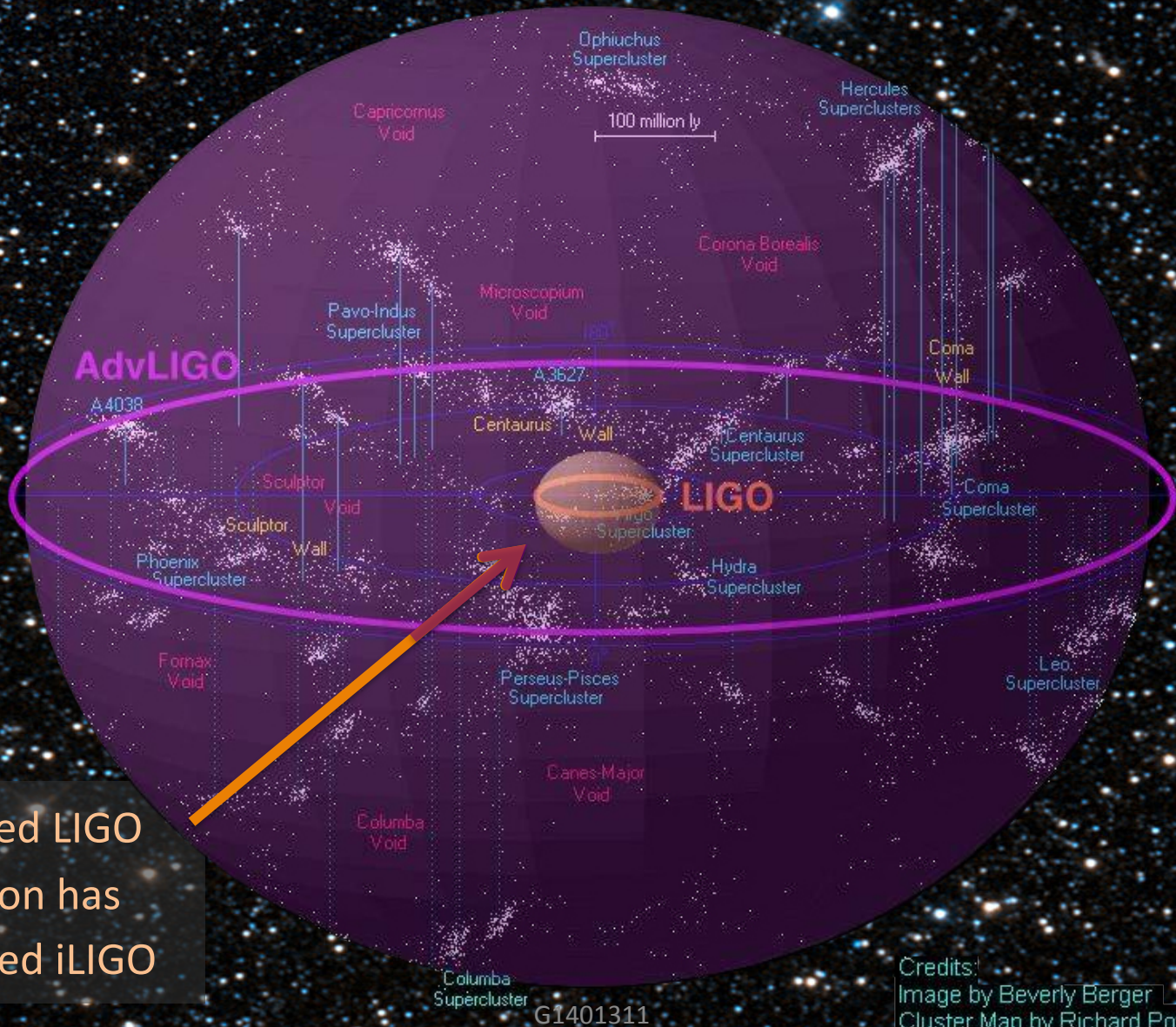


Characterizing the Advanced LIGO instruments

Jess McIver for the
LIGO Detector Characterization Group
January 20, 2015



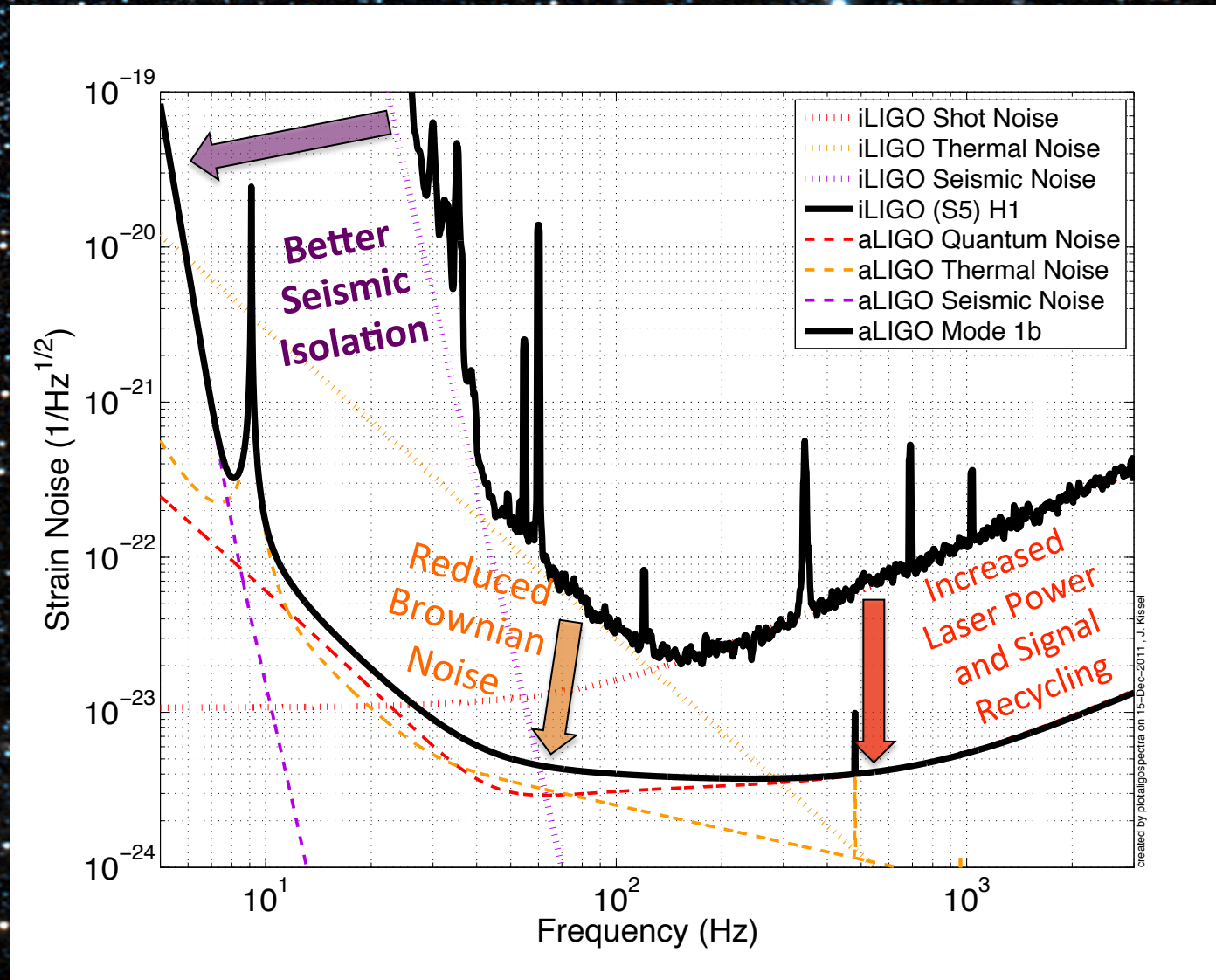
Advanced LIGO



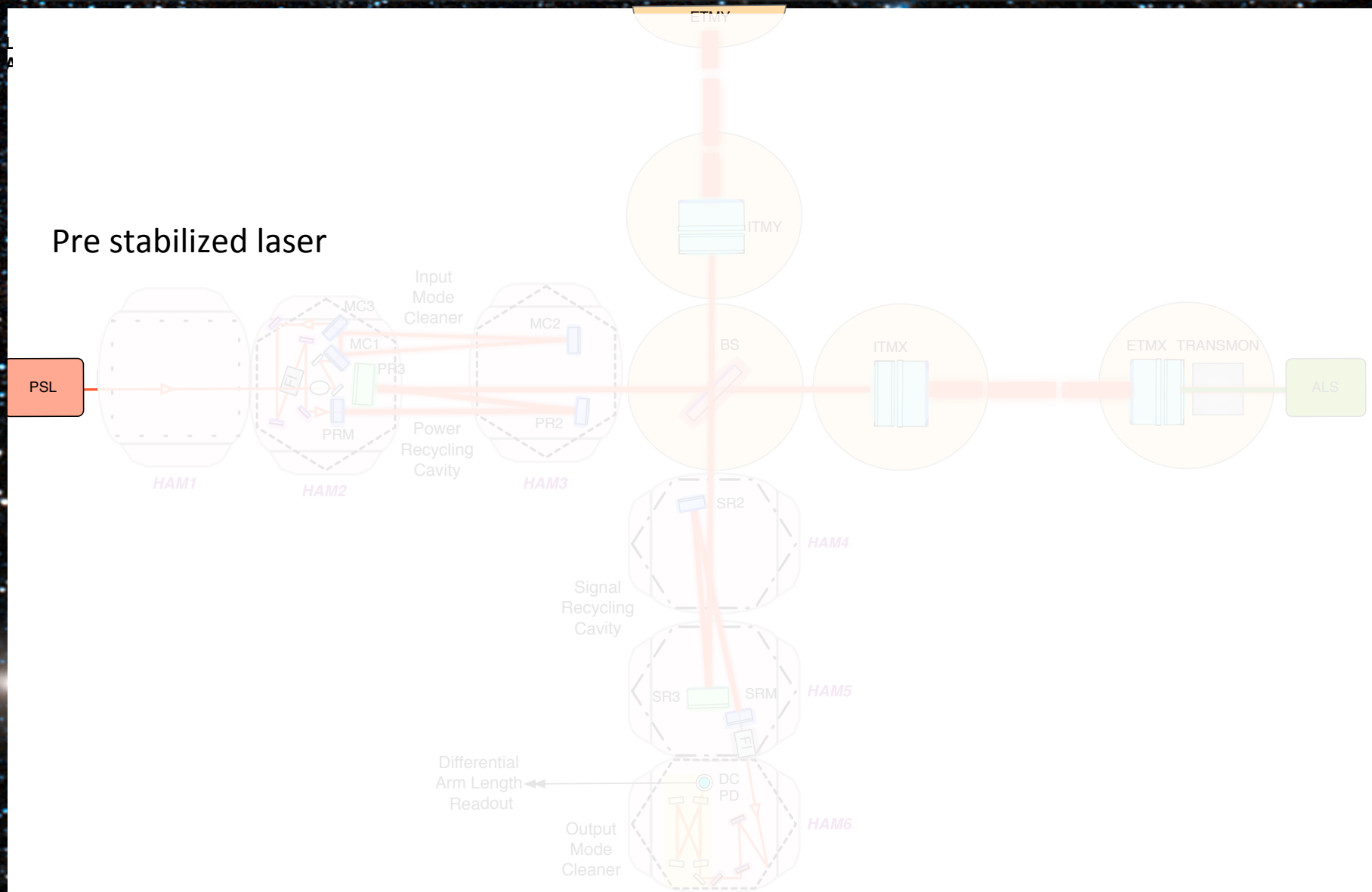
Advanced LIGO
Livingston has
surpassed iLIGO

Credits:
Image by Beverly Berger
Cluster Map by Richard Powell

aLIGO instrumental improvements

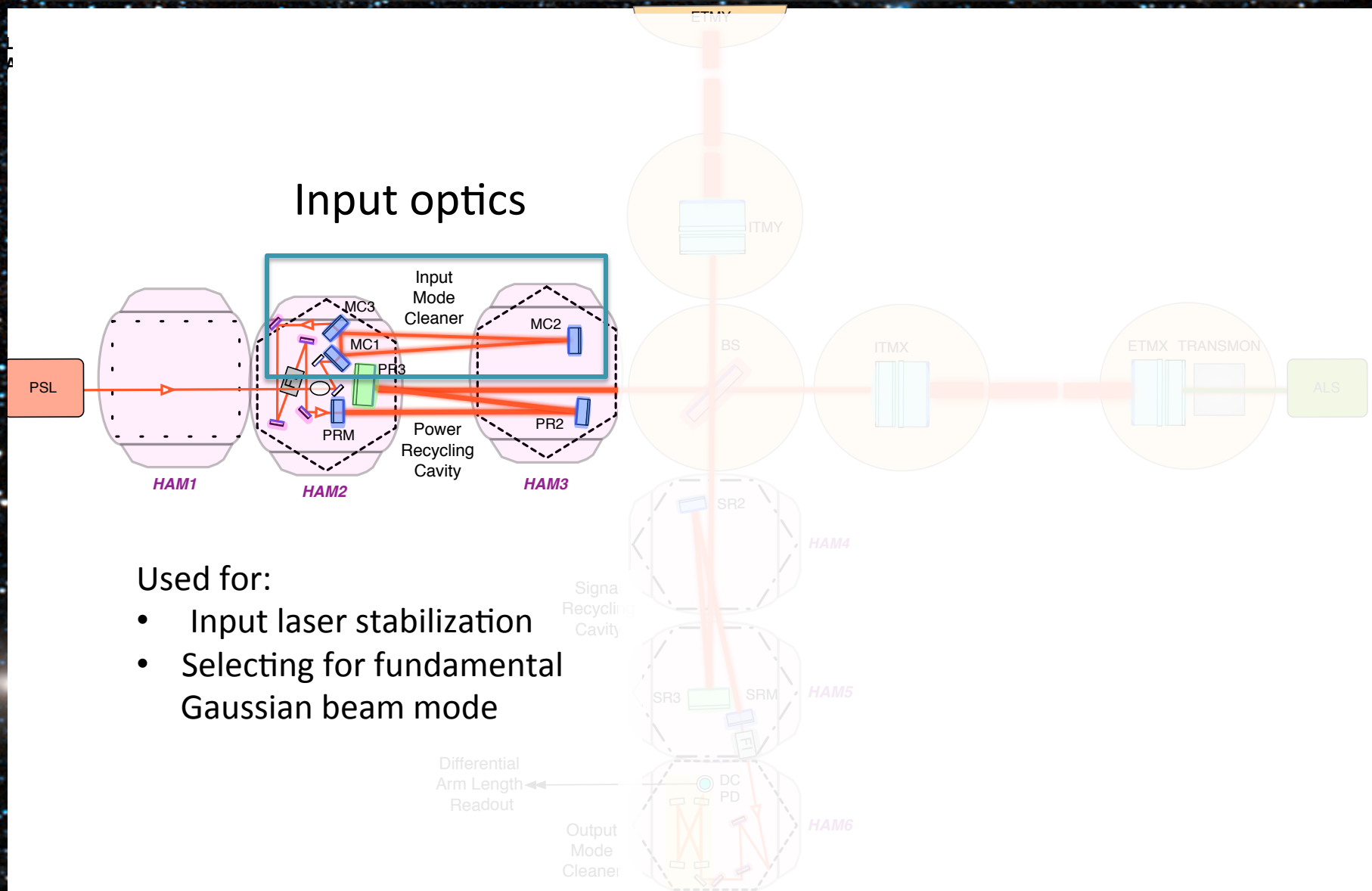


aLIGO installation and testing

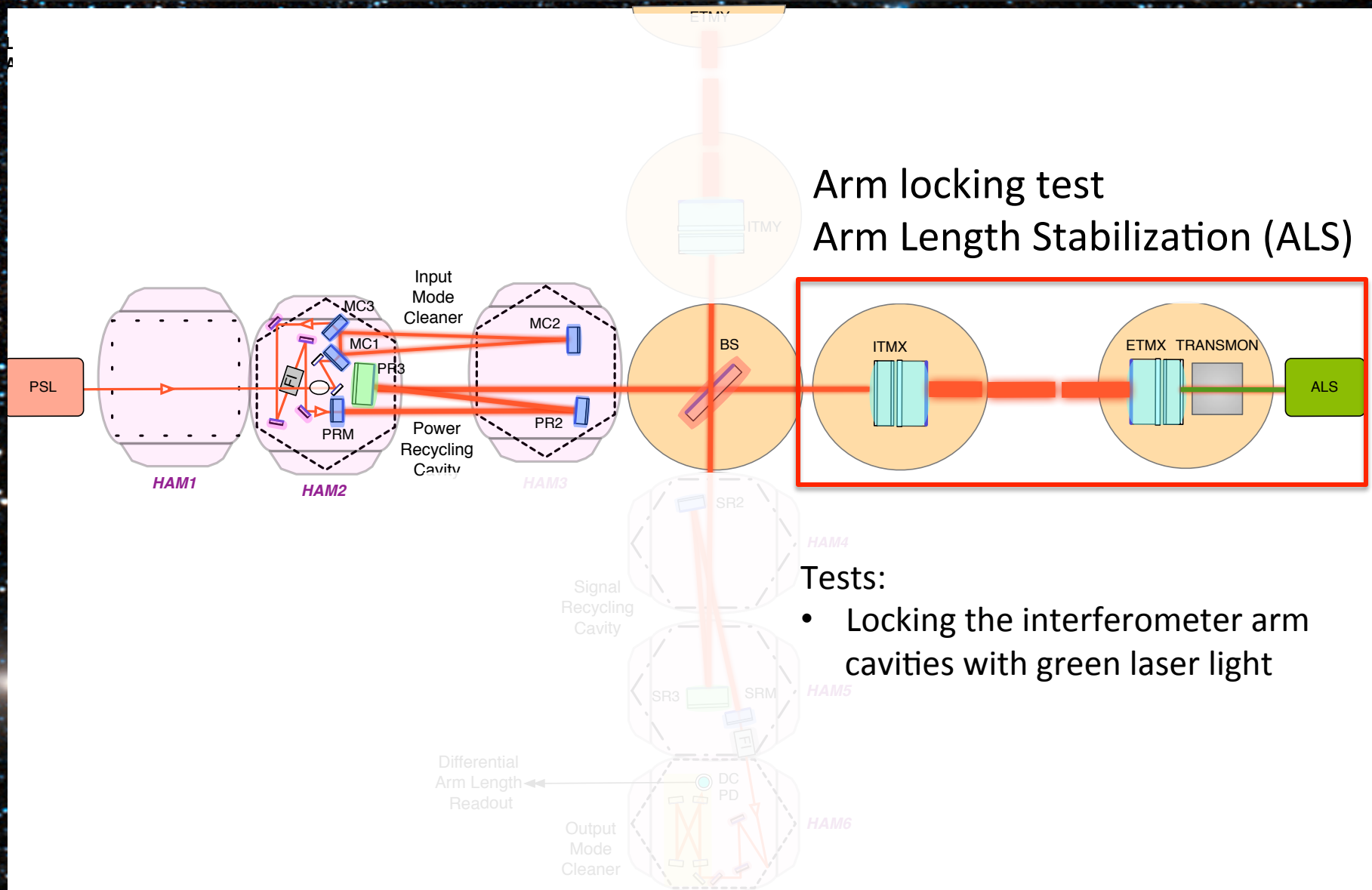


aLIGO installation and testing

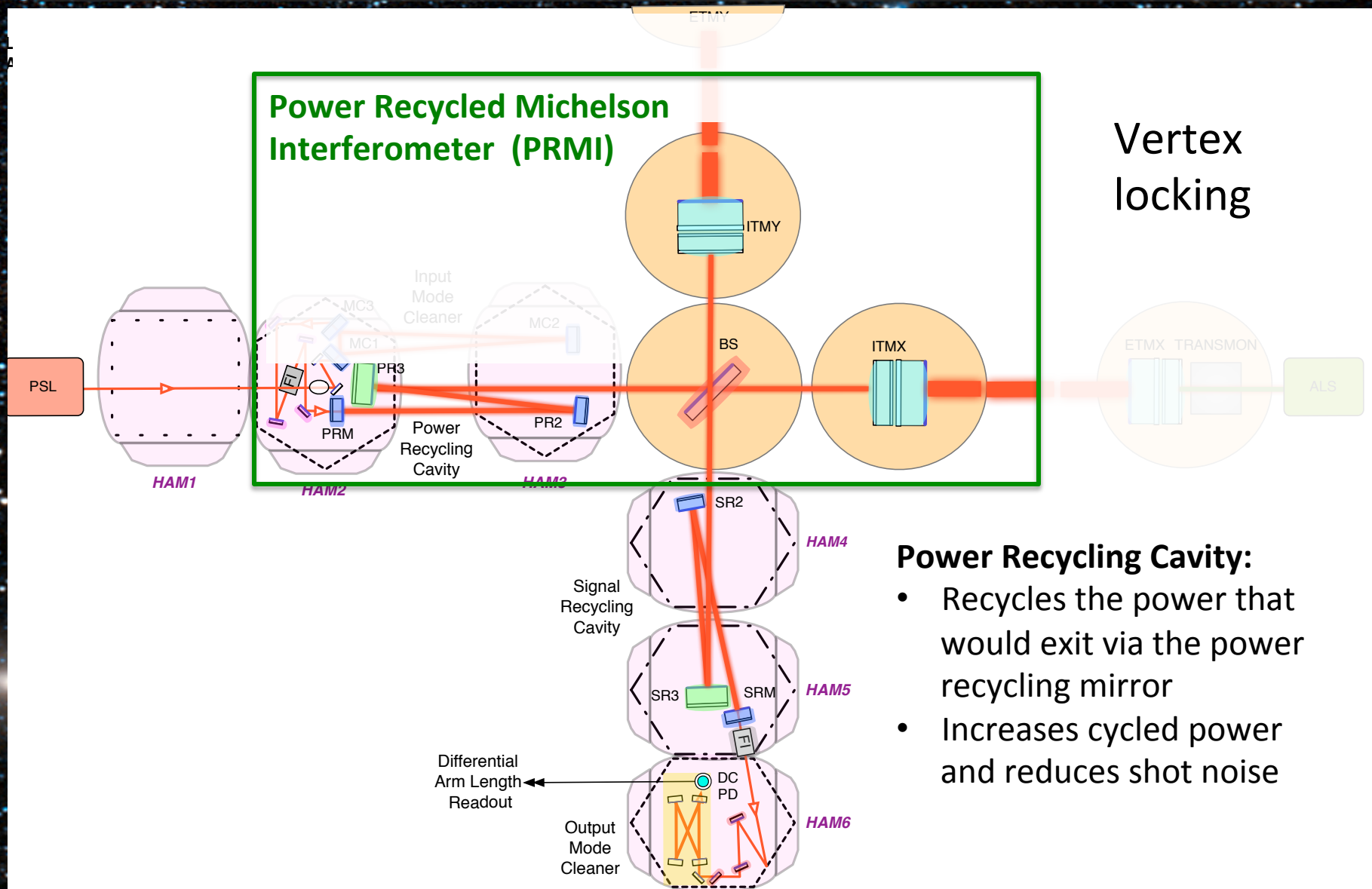
Input optics



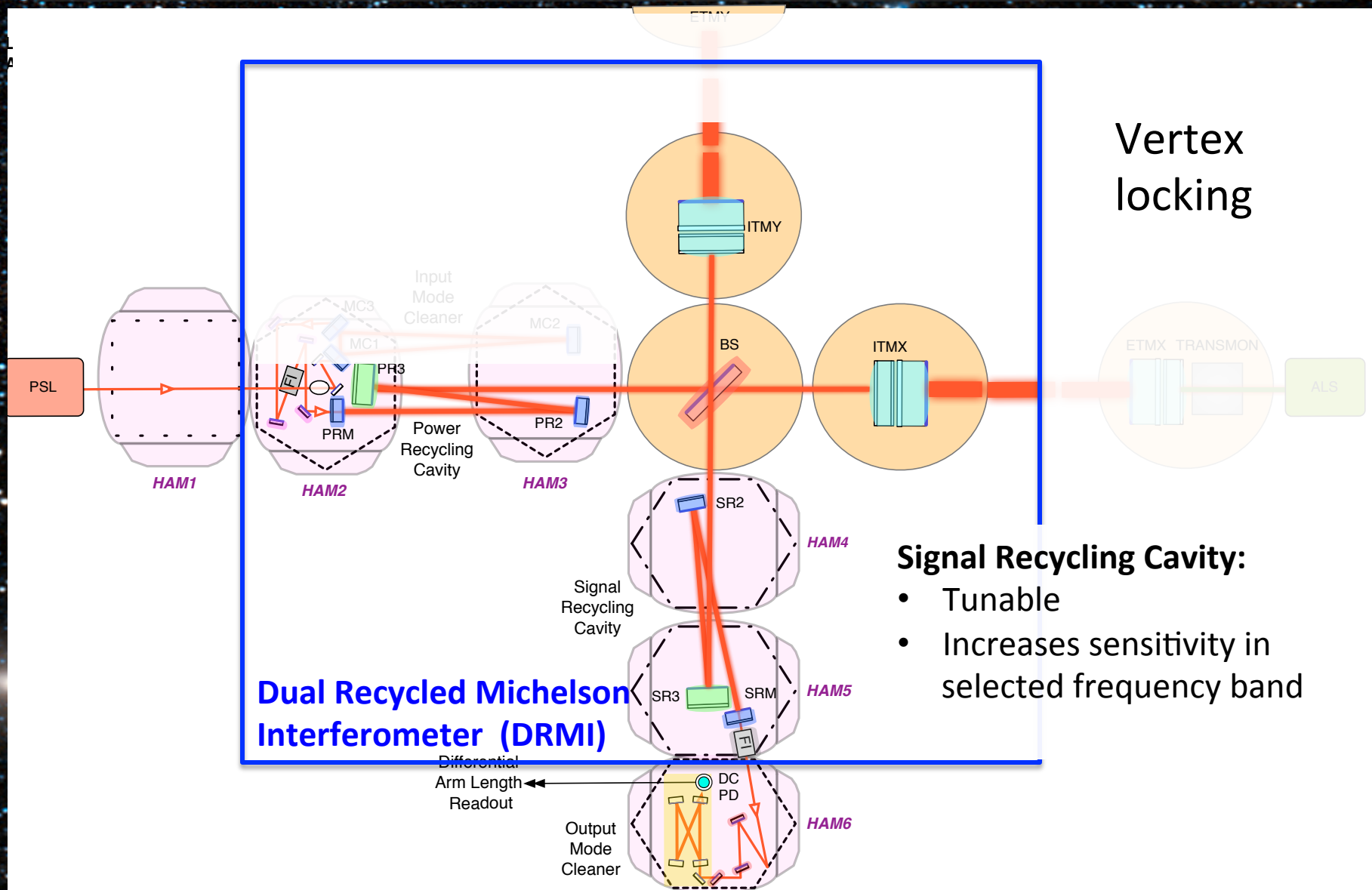
aLIGO installation and testing



aLIGO installation and testing

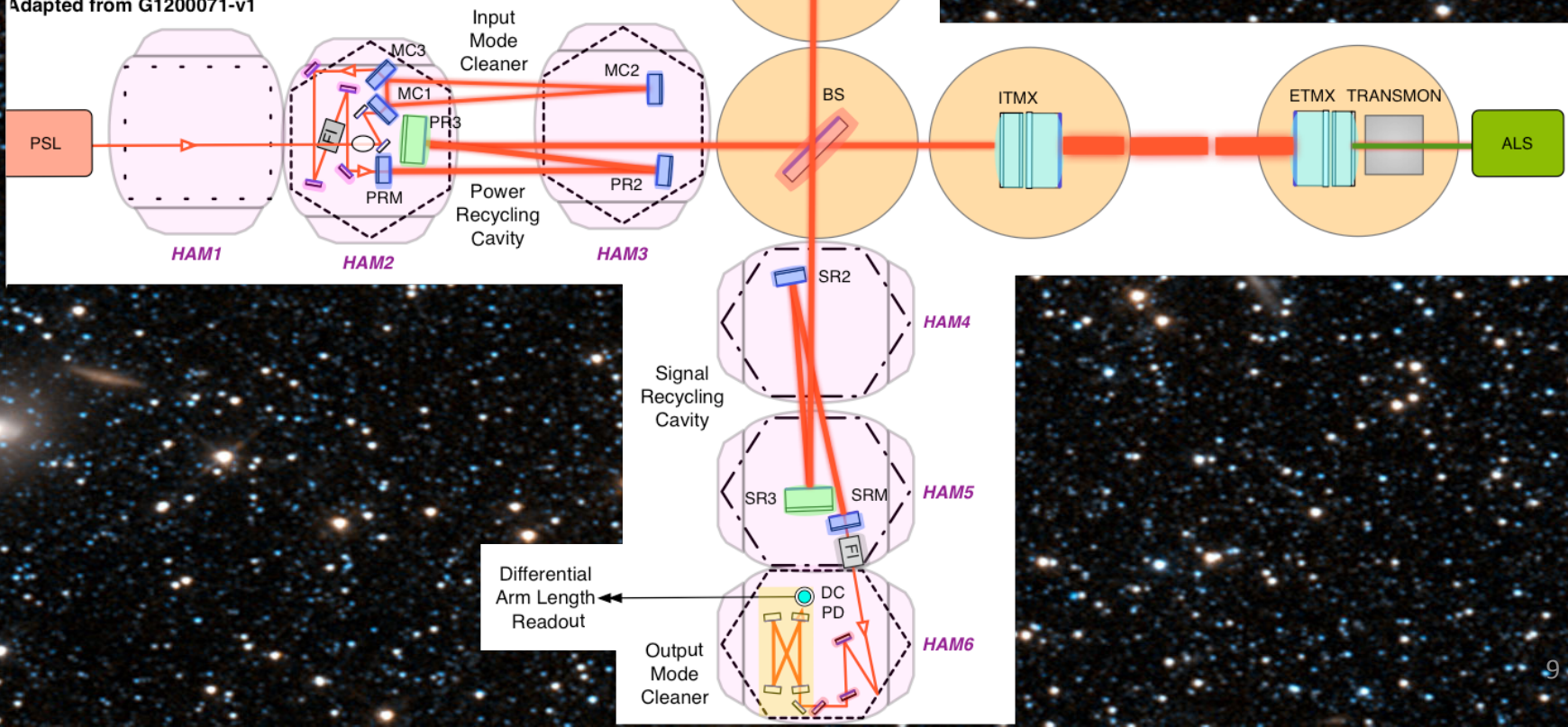


aLIGO installation and testing



aLIGO installation and testing

L. Barsotti - March 9, 2012
Adapted from G1200071-v1



Full interferometer!

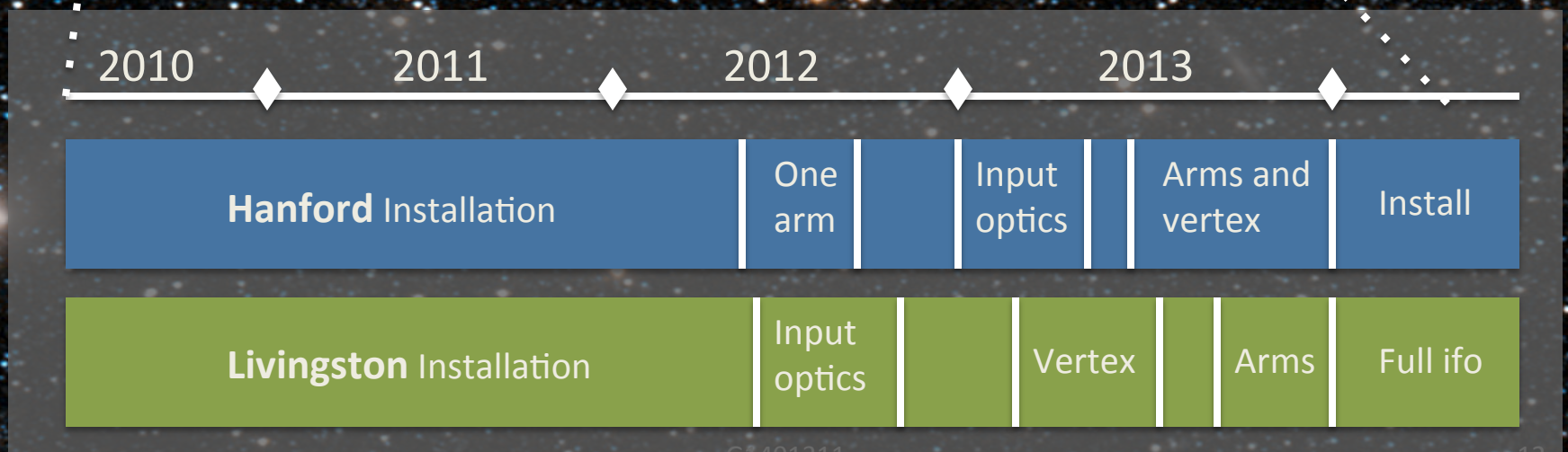
Timeline: from eLIGO to aLIGO



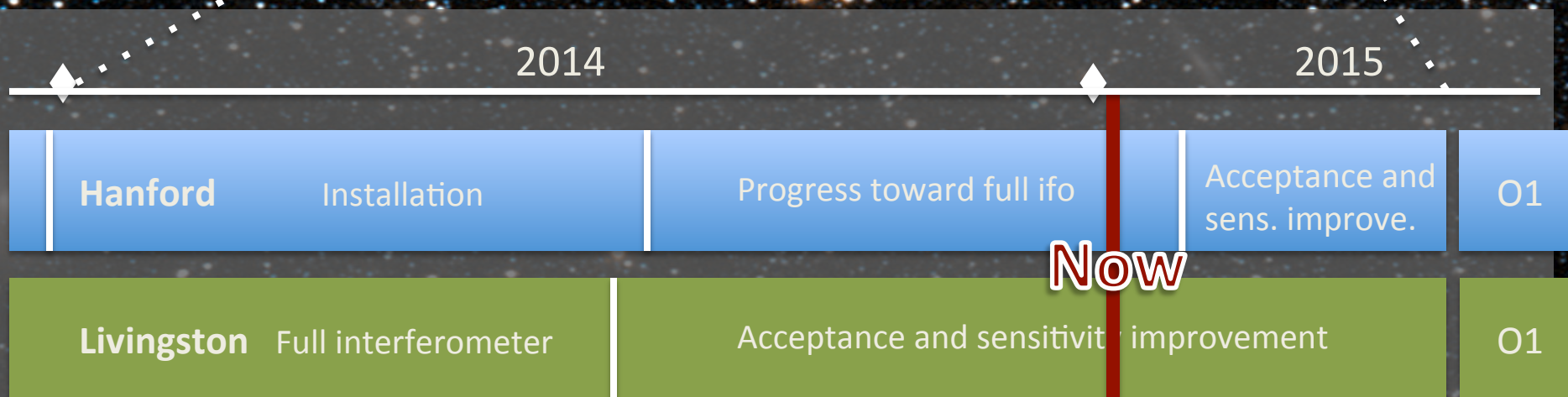
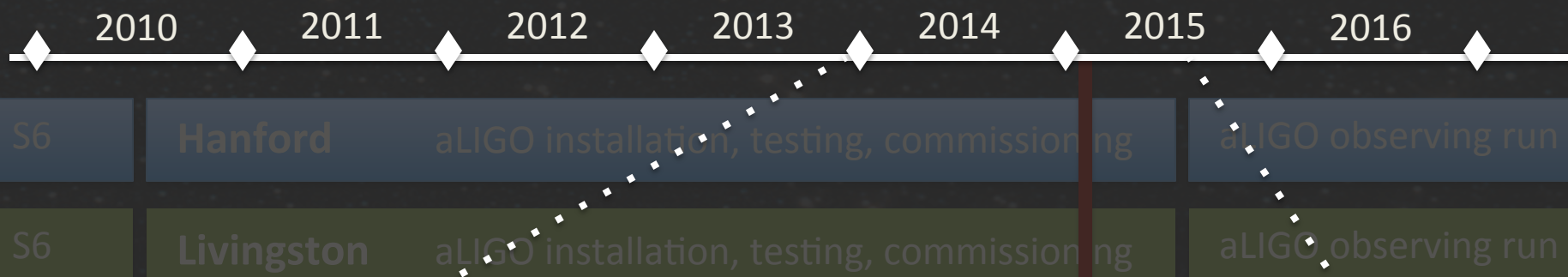
Timeline: from eLIGO to aLIGO



Timeline: installation and testing

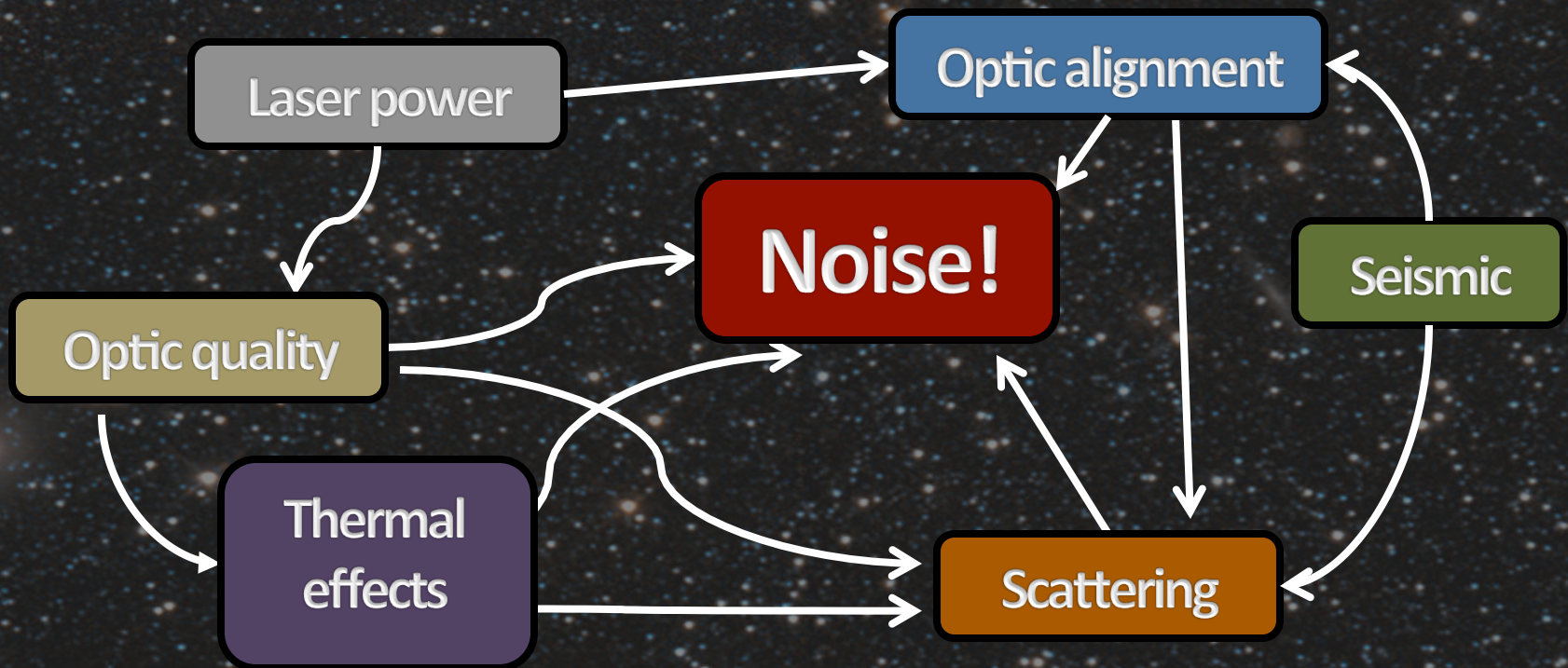


Timeline: the lead up to the first observing run



The challenges of commissioning

- Many effects cannot be tested prior to large scale implementation
- Often noise sources stem from the interaction of different subsystems and cavities



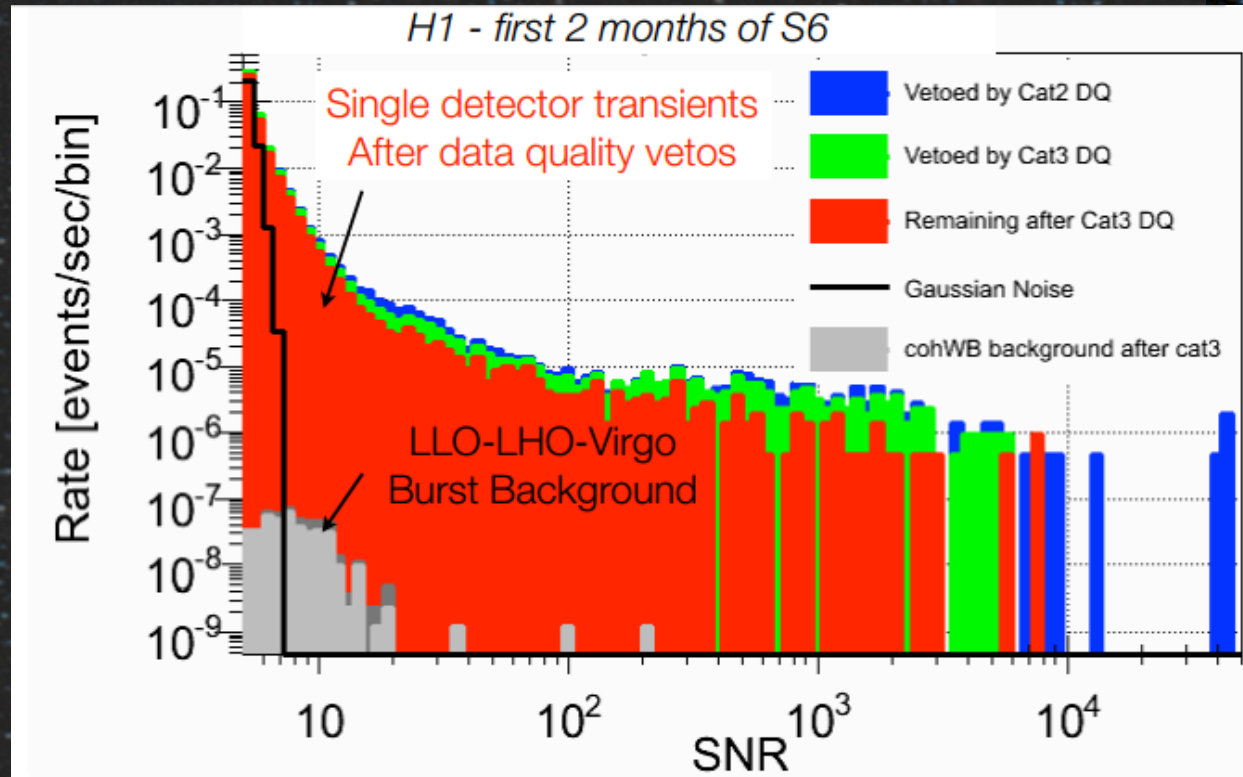


Lessons from the past

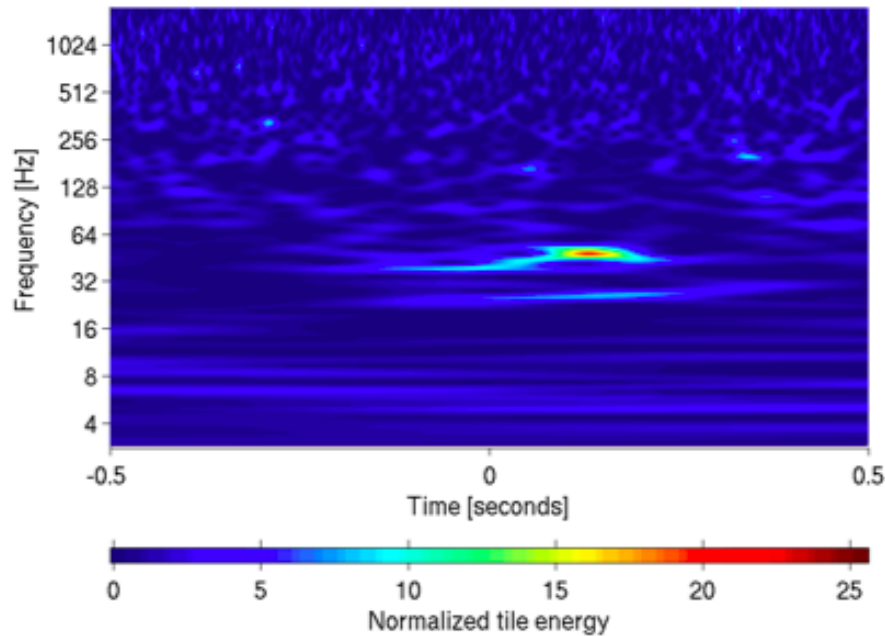
GW search pipelines are adversely affected by non-Gaussian data!

Long tails (outliers) in all-sky GW burst search background triggers greatly restrict achievable false alarm rate.

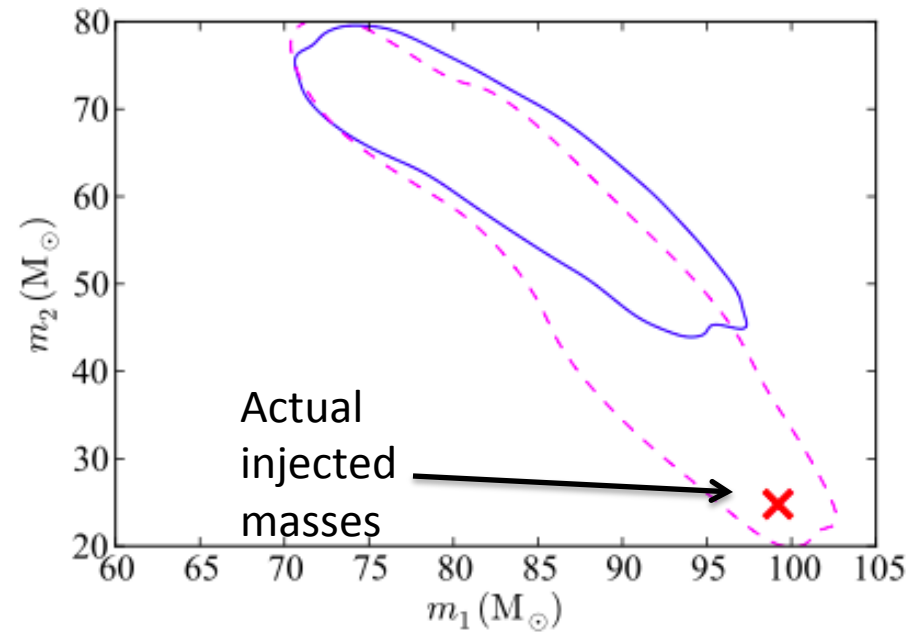
Non-Gaussian noise confuses parameter estimation for all transient searches.



Example: NINJA2 search results



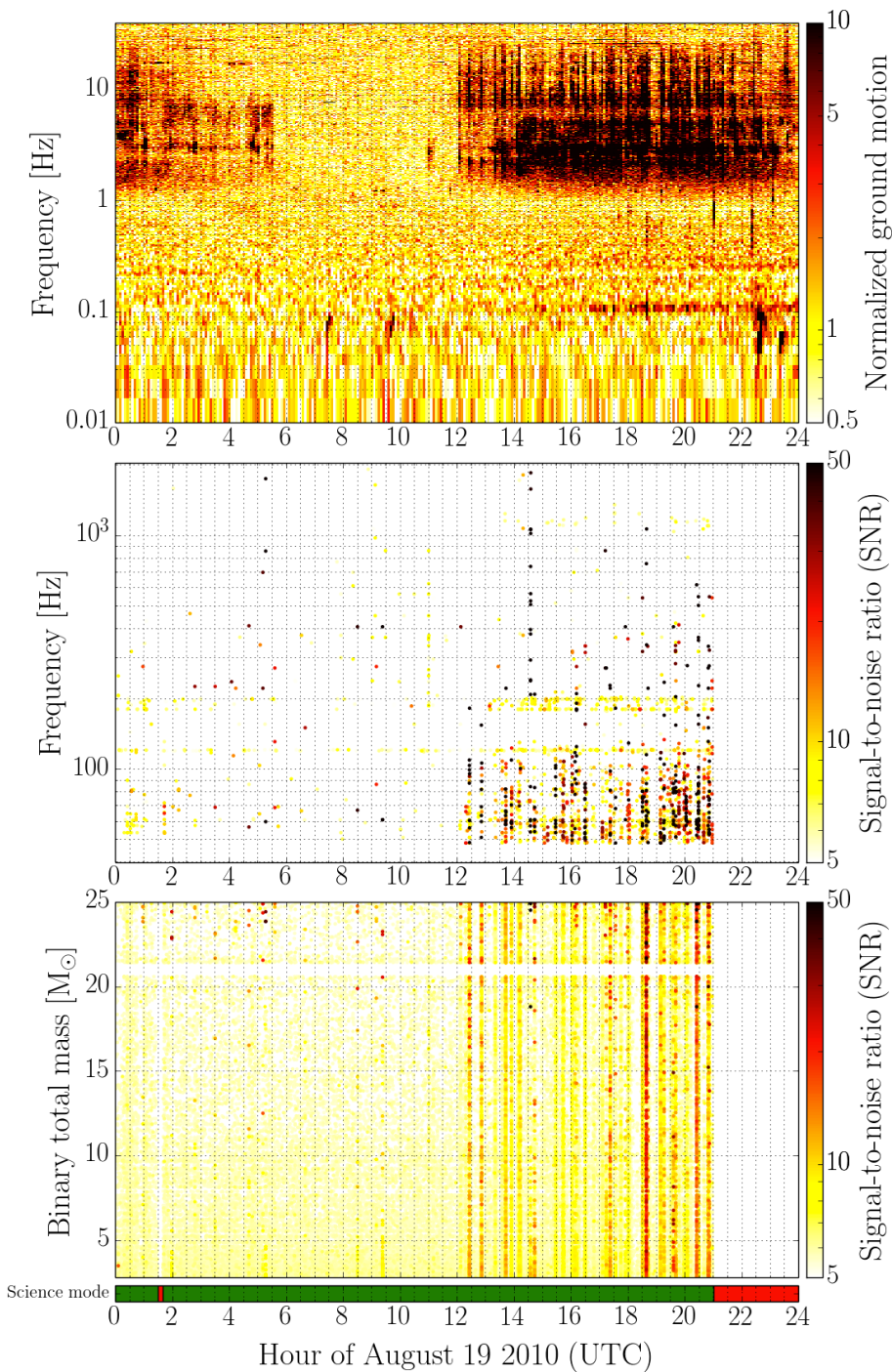
(a)



(b)

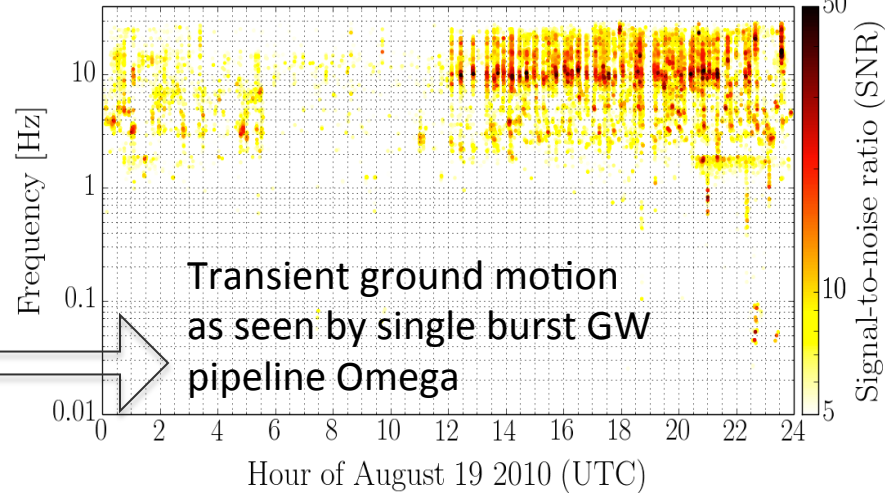
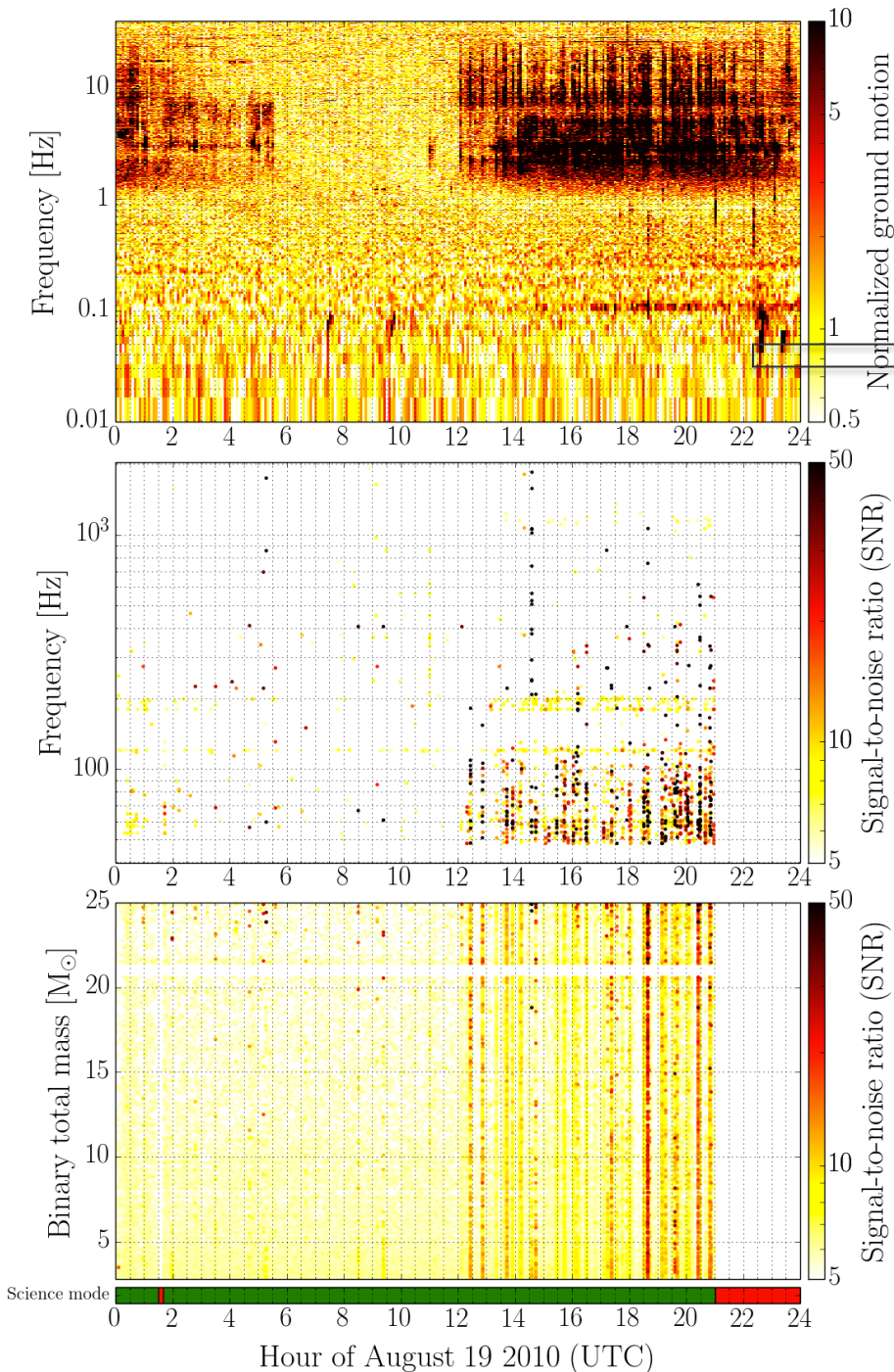
A normalized spectrogram of **Hanford recolored noise only** showing a transient event, or *glitch*, that happens to occur at the time of the injection.

Solid blue – the 95% credible region for mass estimation based on EOBNRv2 analysis using realistic recolored noise. Dashed pink – in Gaussian noise.



Example:
Transient seismic motion
 – a known and
 troublesome problem
 during S6.

arXiv preprint 1410.7764



Transient seismic motion
 – a known and
 troublesome problem
 during S6.

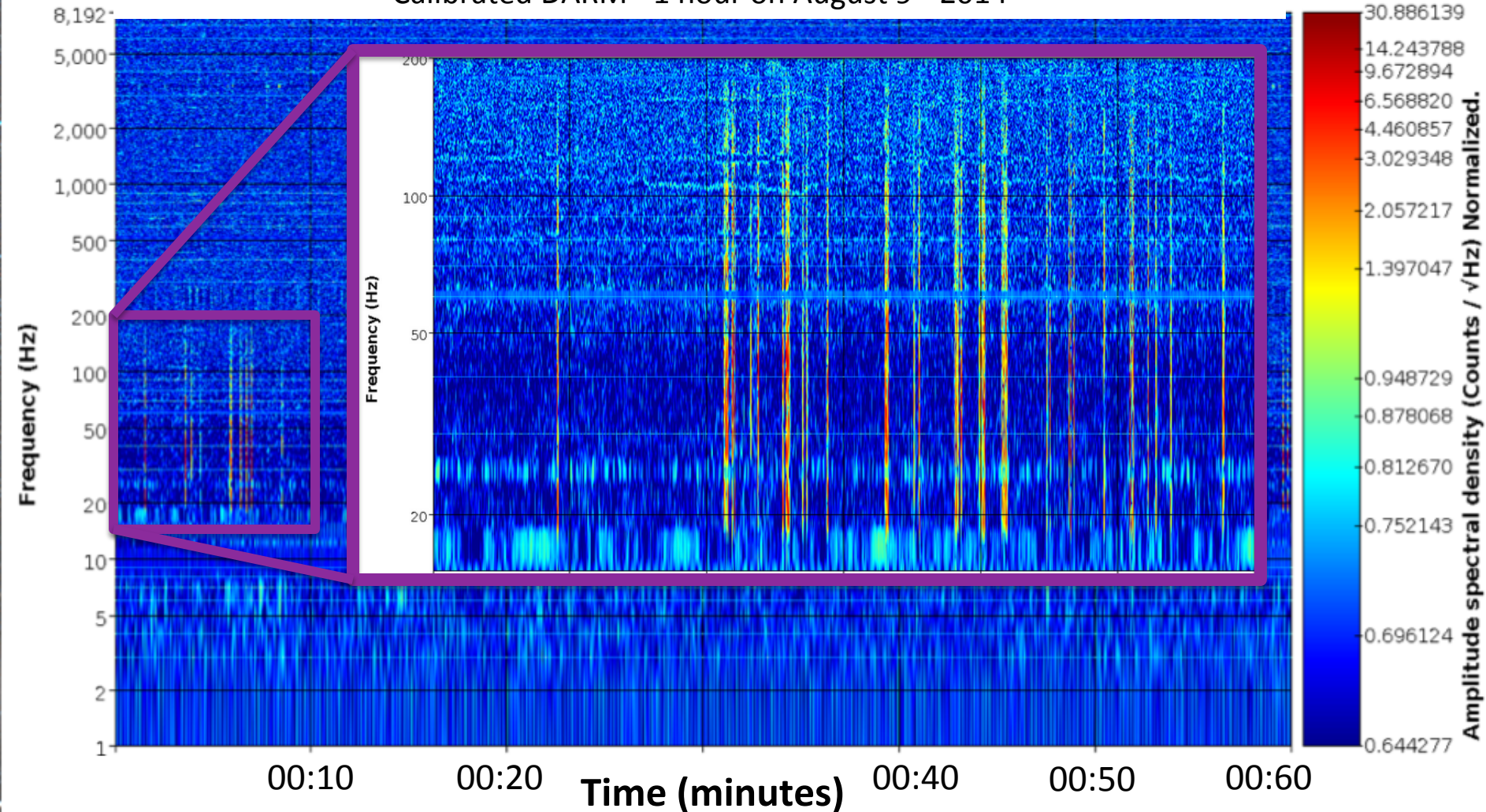
arXiv preprint 1410.7764

Recent data from Advanced LIGO

Highlights of DQ features most
troublesome for the transient searches

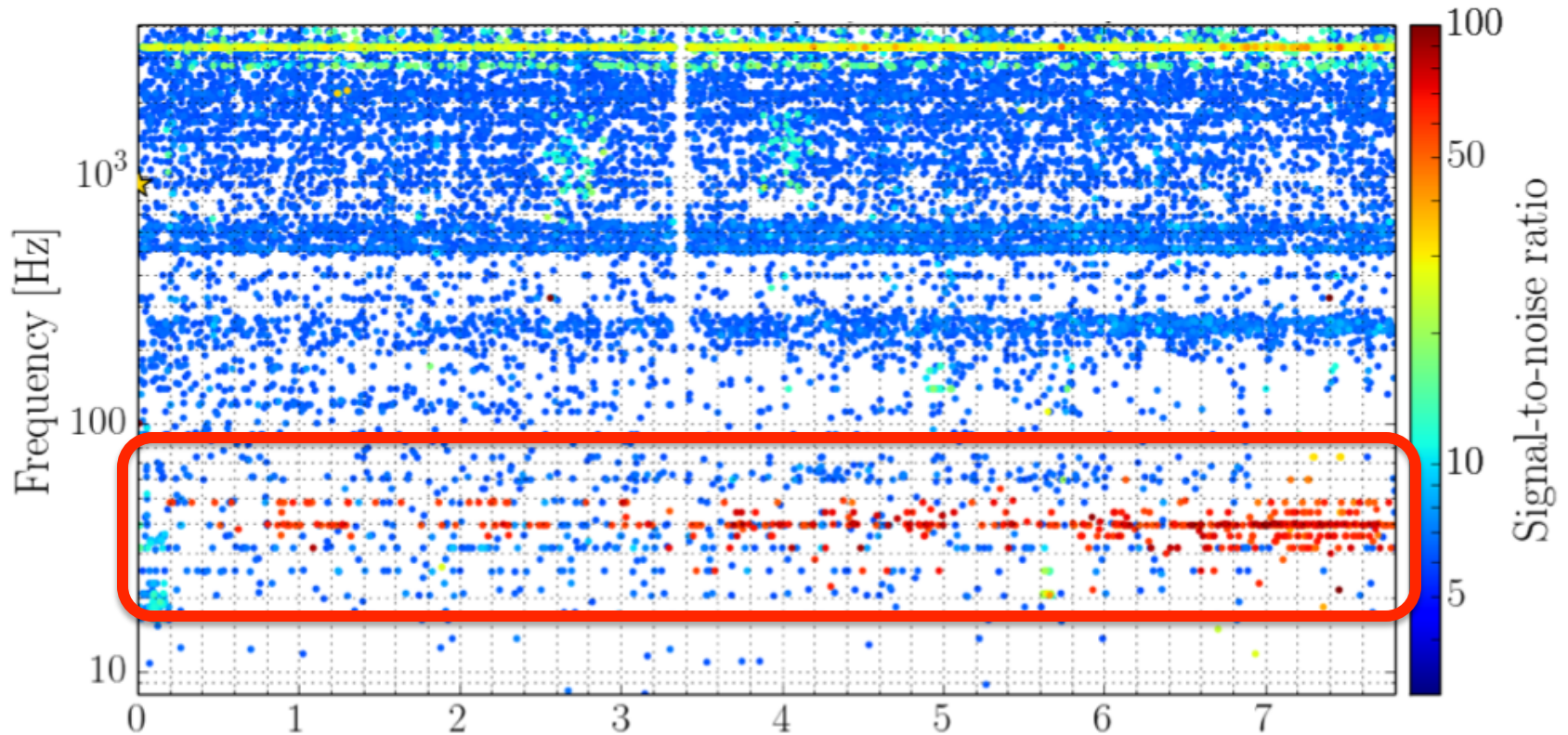
An early DQ issue diagnosis

Calibrated DARM - 1 hour on August 9th 2014



Through the lens of a single ifo burst GW pipeline

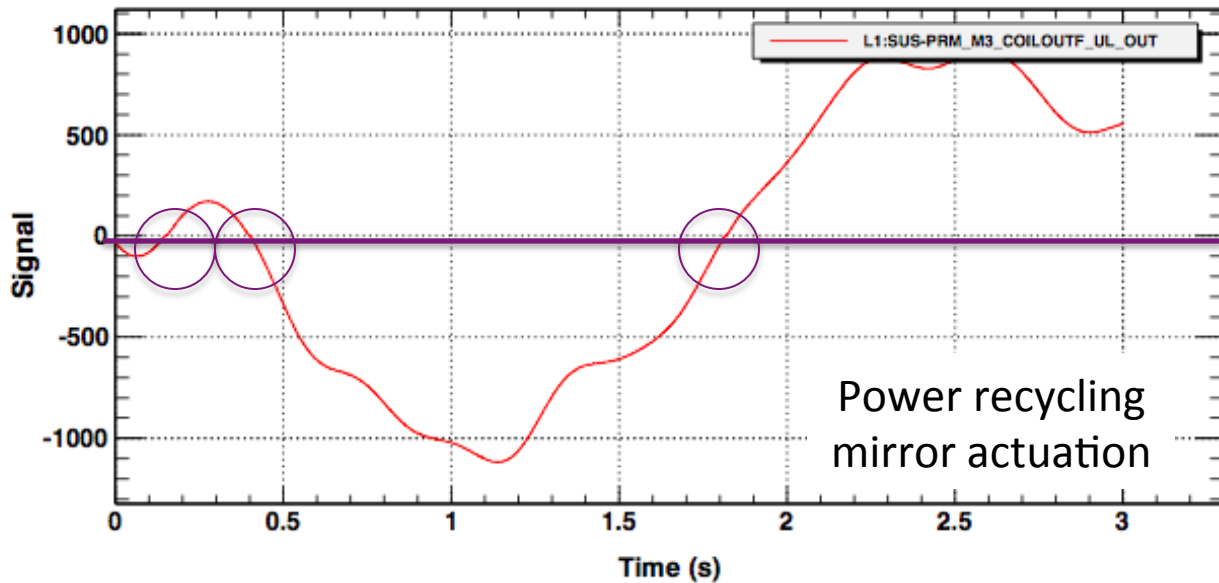
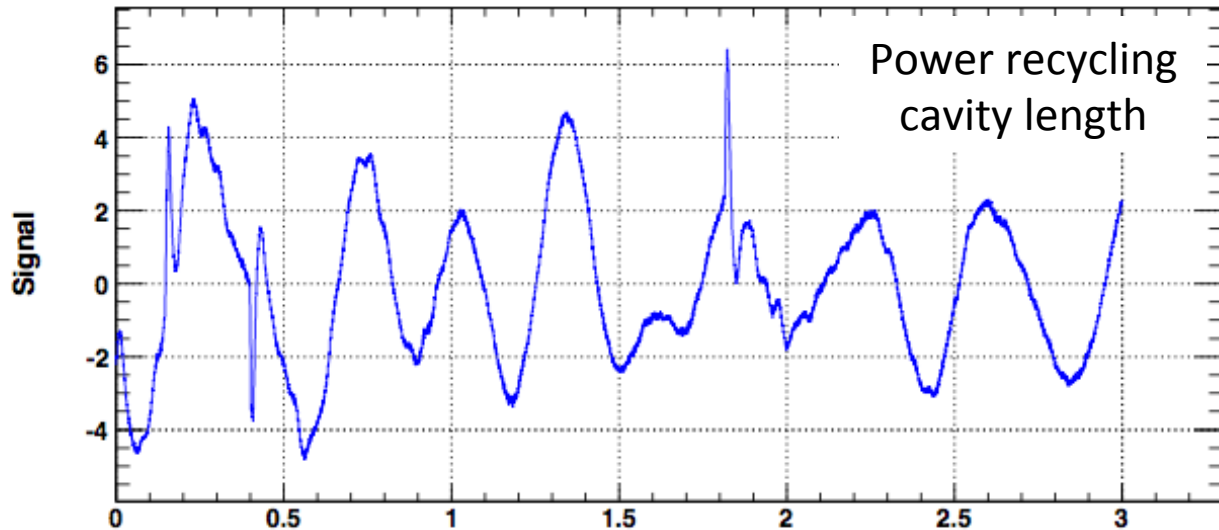
Calibrated DARM (differential arm length)



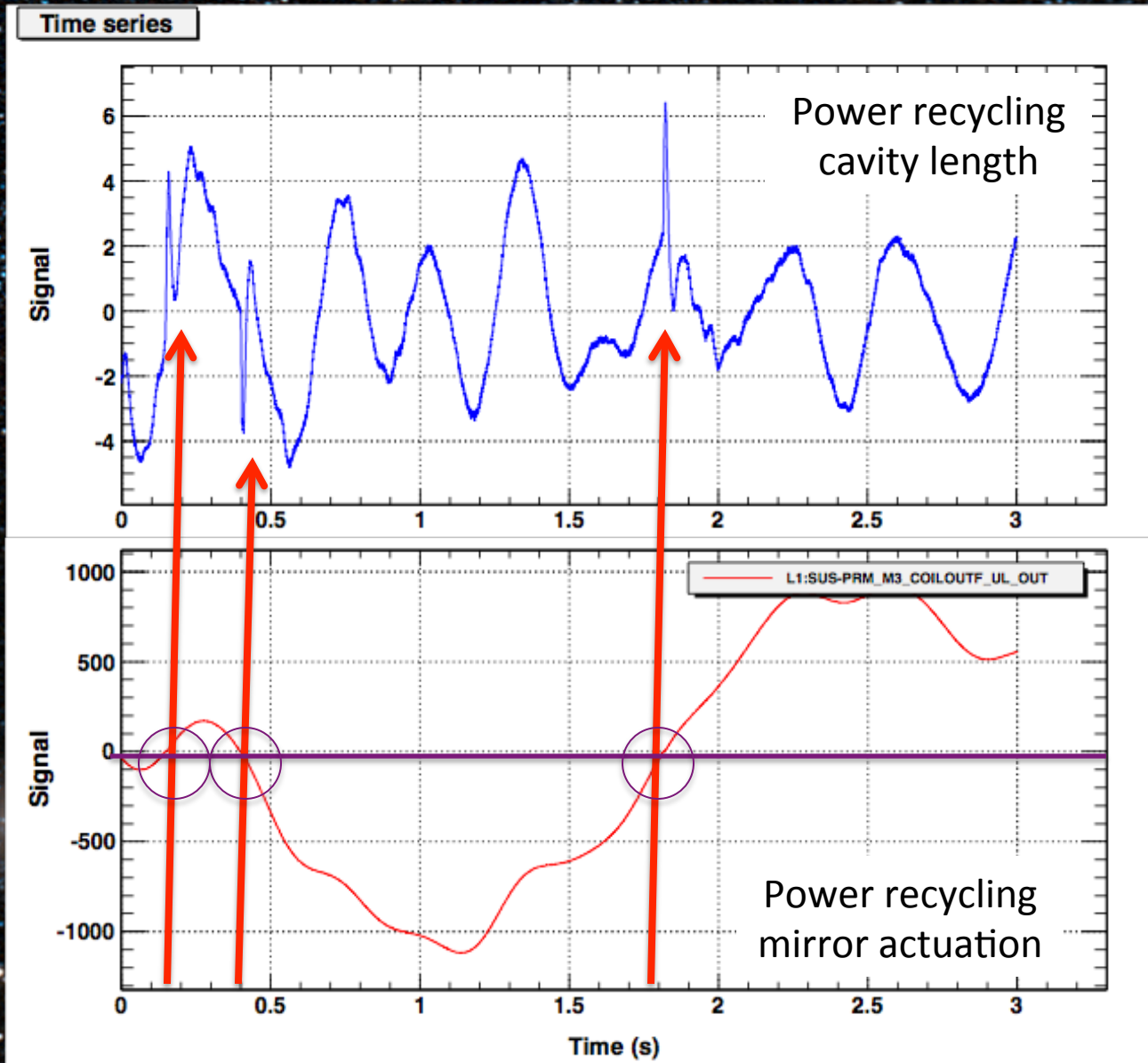
8 hours on August 9th 2014

The mechanism : major carry transition DAC glitching

Time series



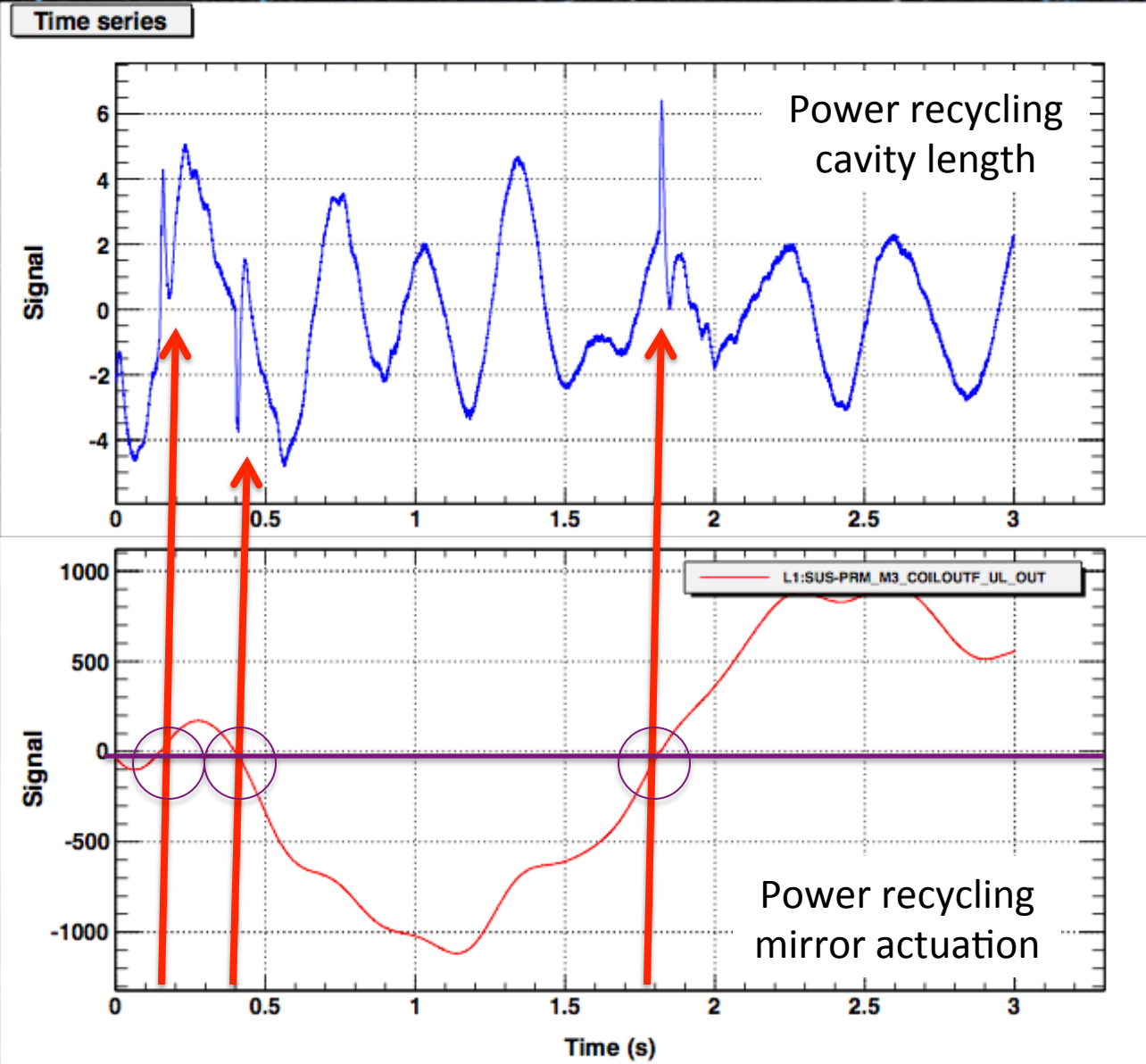
The mechanism : major carry transition DAC glitching



First dubbed “zero crossing” glitches – identified when vertex cavity actuation signals crossed zero.

Actually a subset of a broader known issue with these DACs.

The mechanism : major carry transition DAC glitching



Major-carry transitions (MCTs): value transitions that cause a most significant bit(s) to change

Examples:

0111



1000

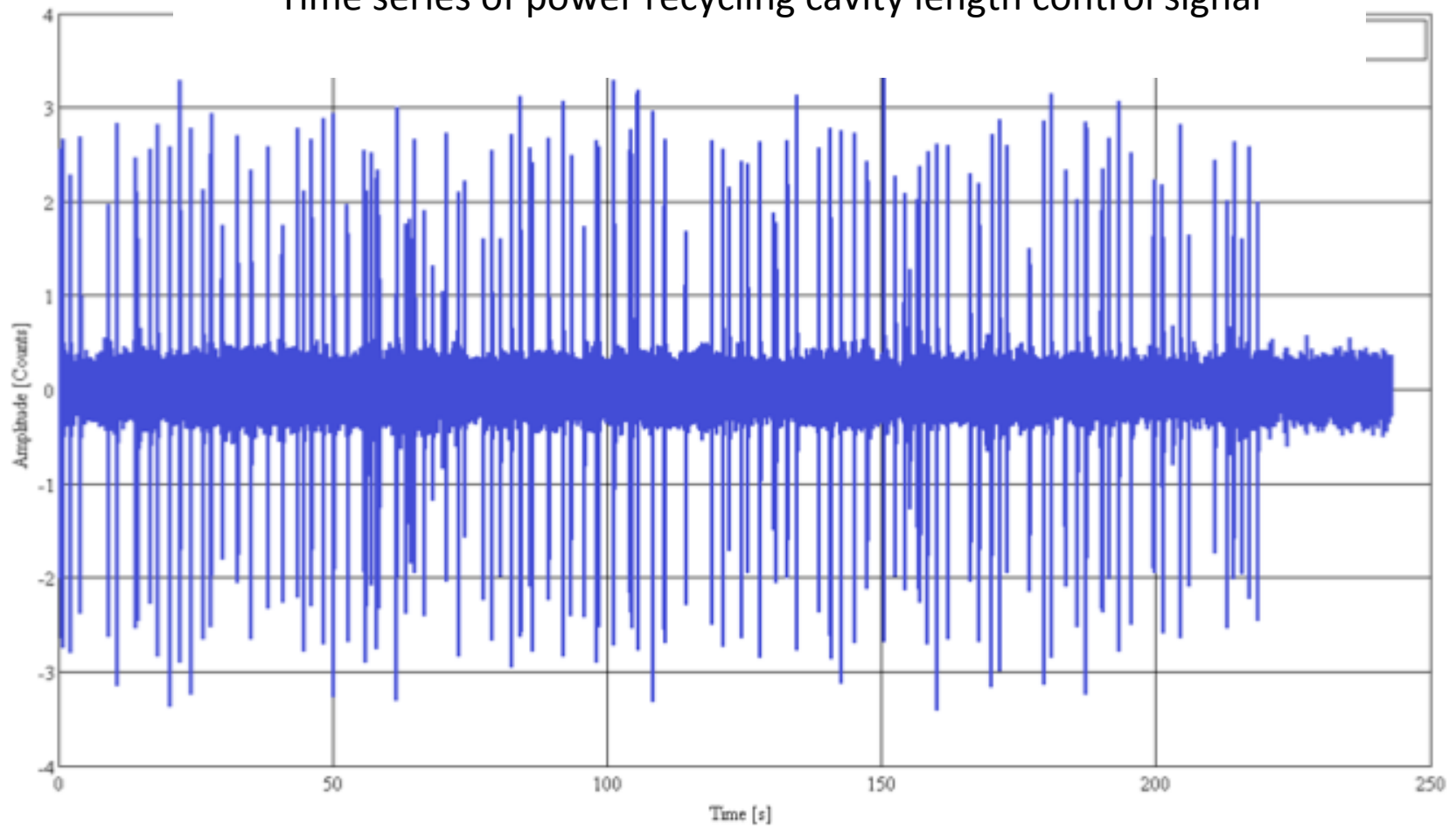
1000



0111

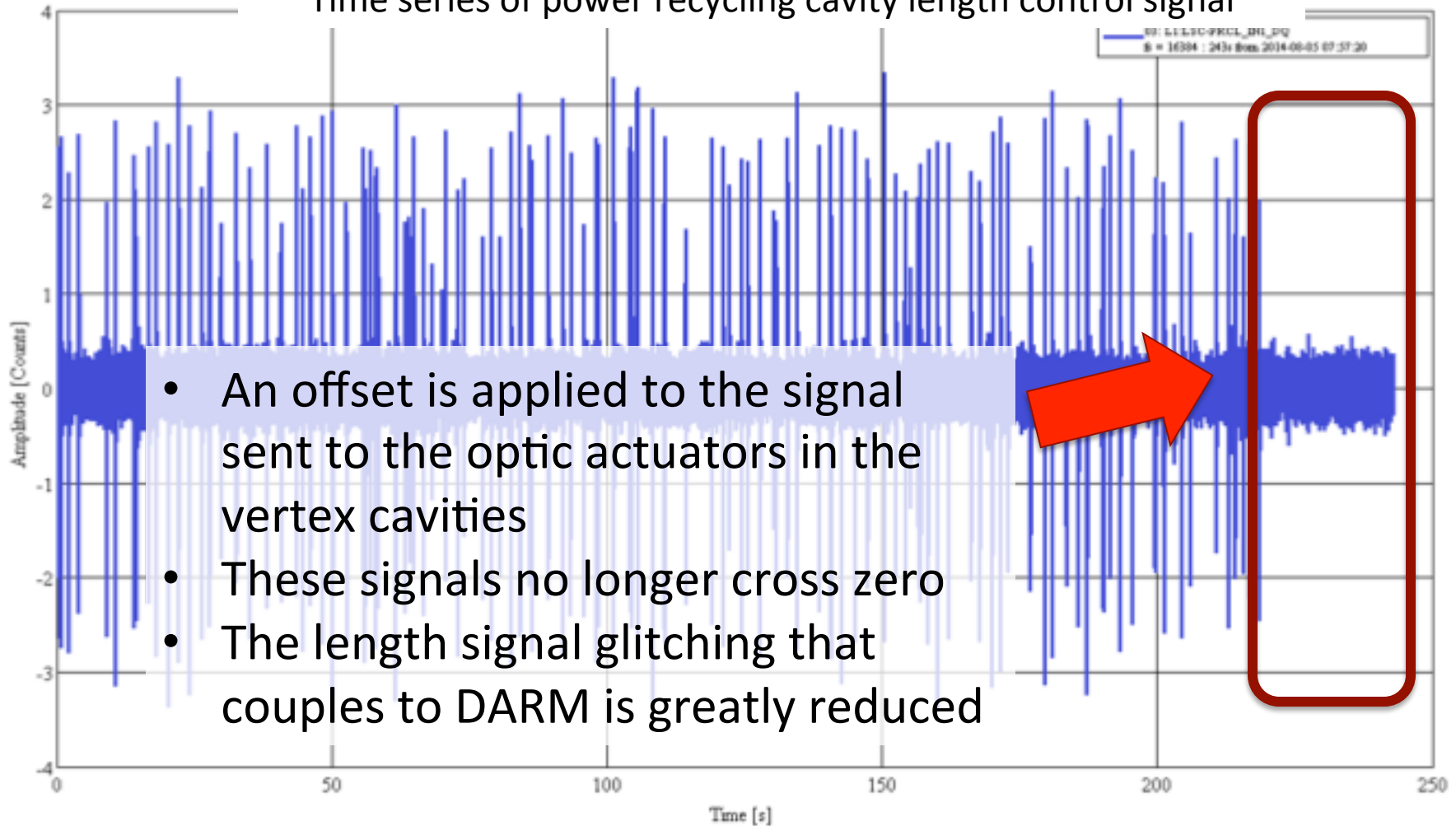
Offset applied to vertex cavity actuation

Time series of power recycling cavity length control signal



Offset applied to vertex cavity actuation

Time series of power recycling cavity length control signal

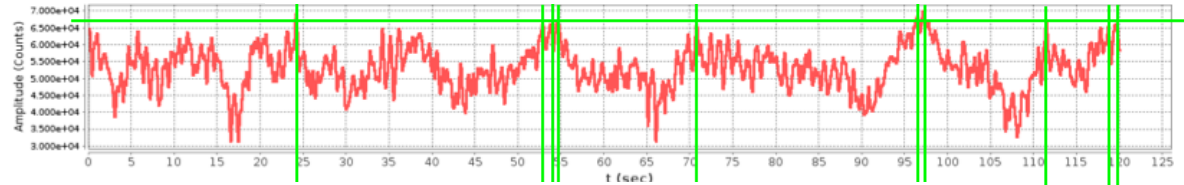


End station actuator glitches identified

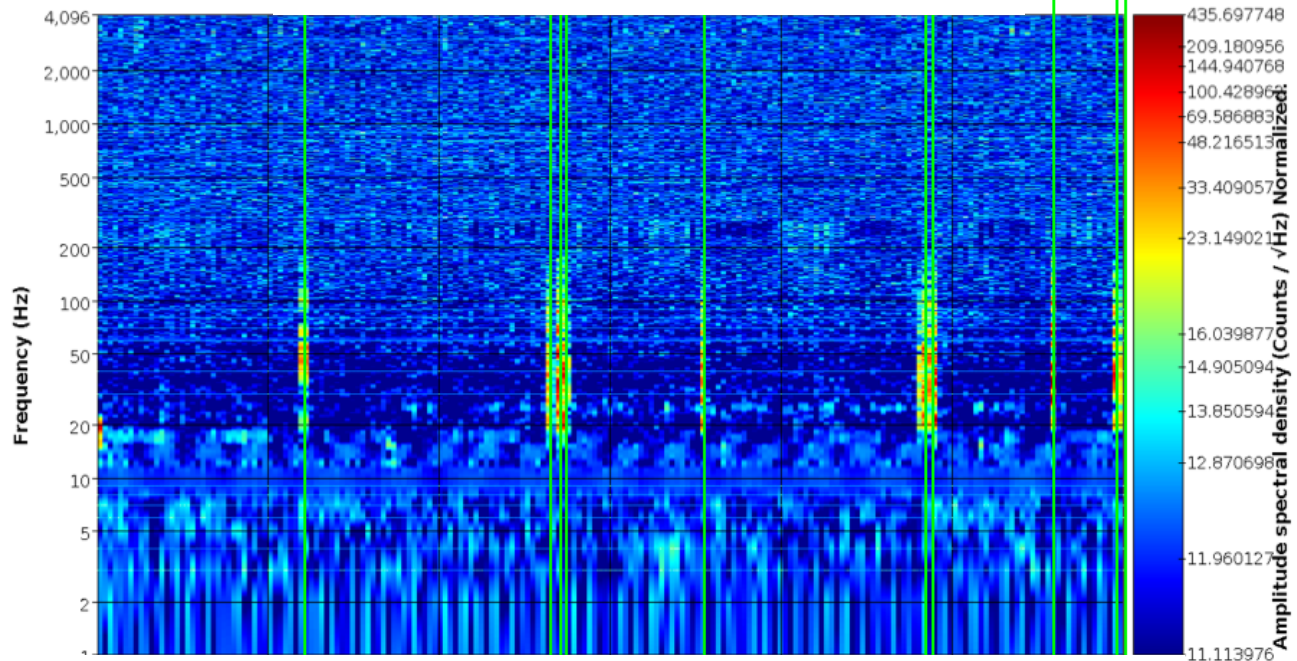
The Detector Characterization group also identified this behavior in one of the end stations at the 2[^]16 MCT

MCT glitching was later addressed with calibration of DACs – the offsets were removed in mid December

End X optic suspension actuation signal

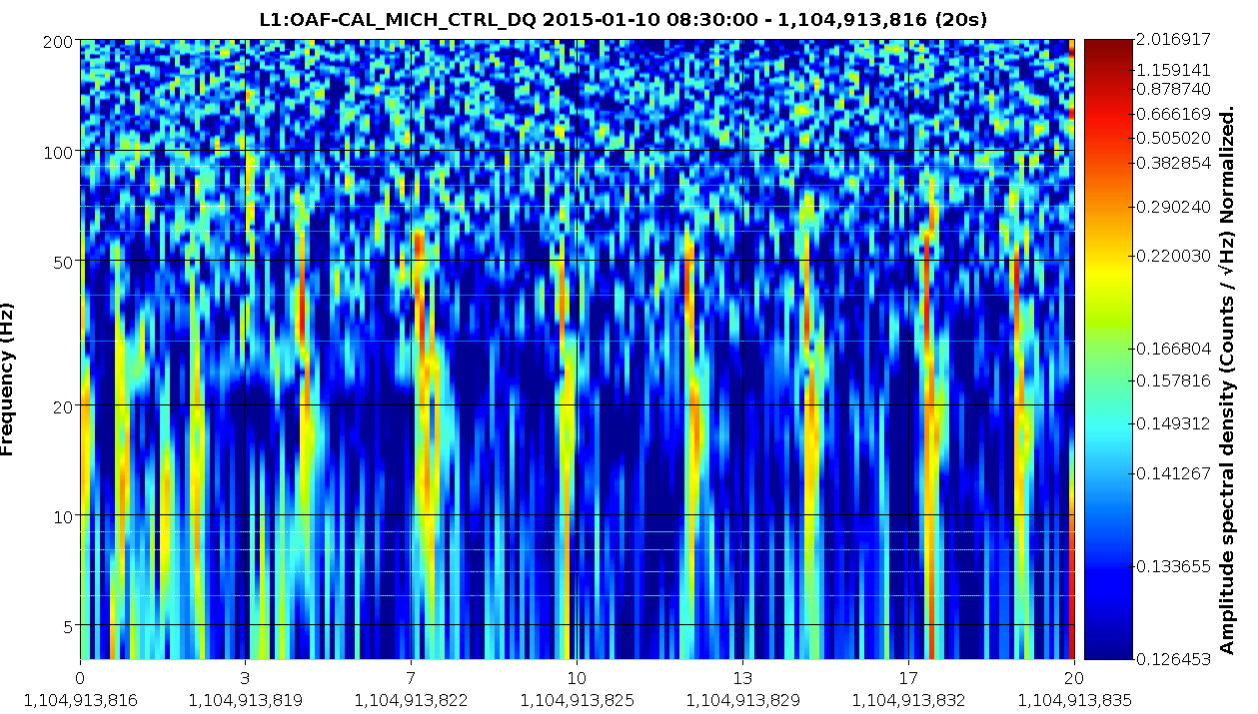
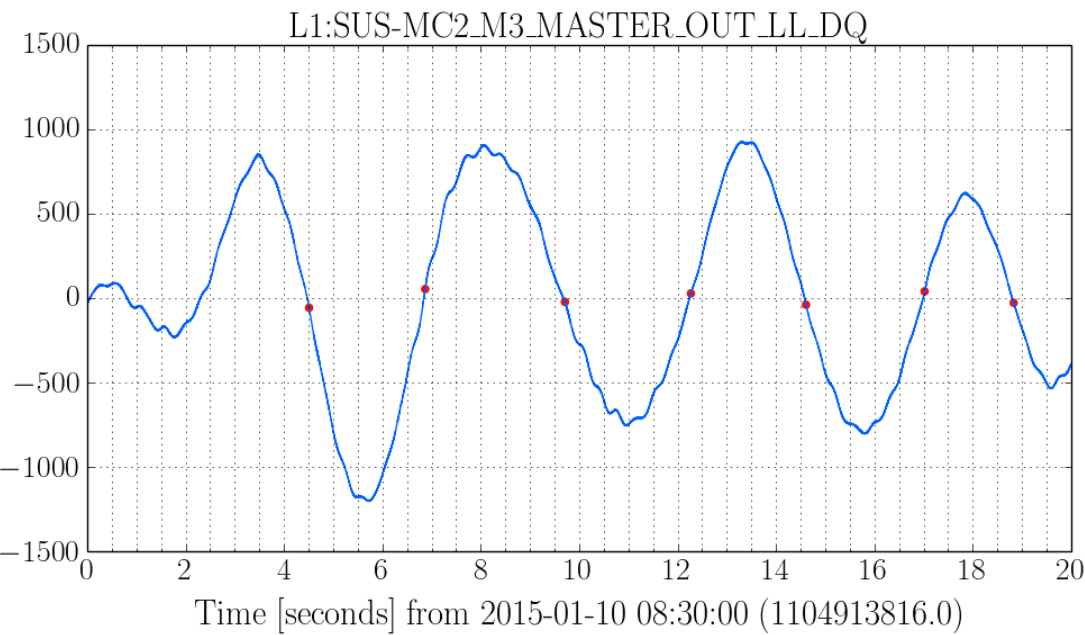


Calibrated DARM



Time – 2 hours

Fs=16,384Hz, sec/fft = 1.00, overlap = 0.40, fft length=16,384, #-FFT = 199, bw = 1, in samples = 1,966K, low = 0.20



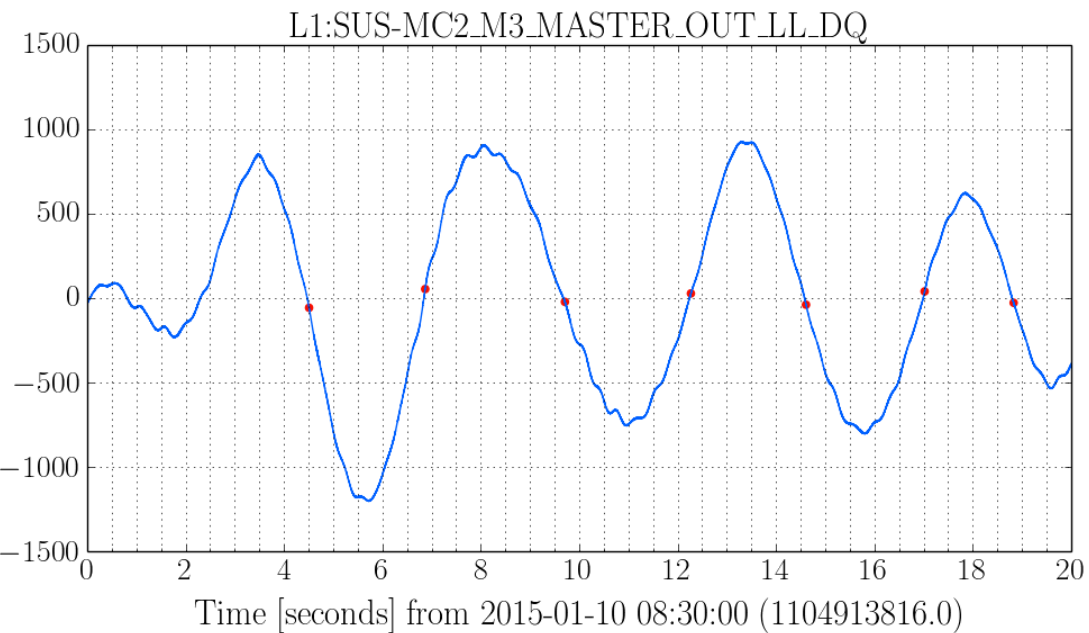
Fs=16,384Hz, sec/fft = 0.25, overlap = 0.60, fft length=4,096, #-FFT = 199, bw = 4, in samples = 328K, low = 0.20

Return of DAC
glitches

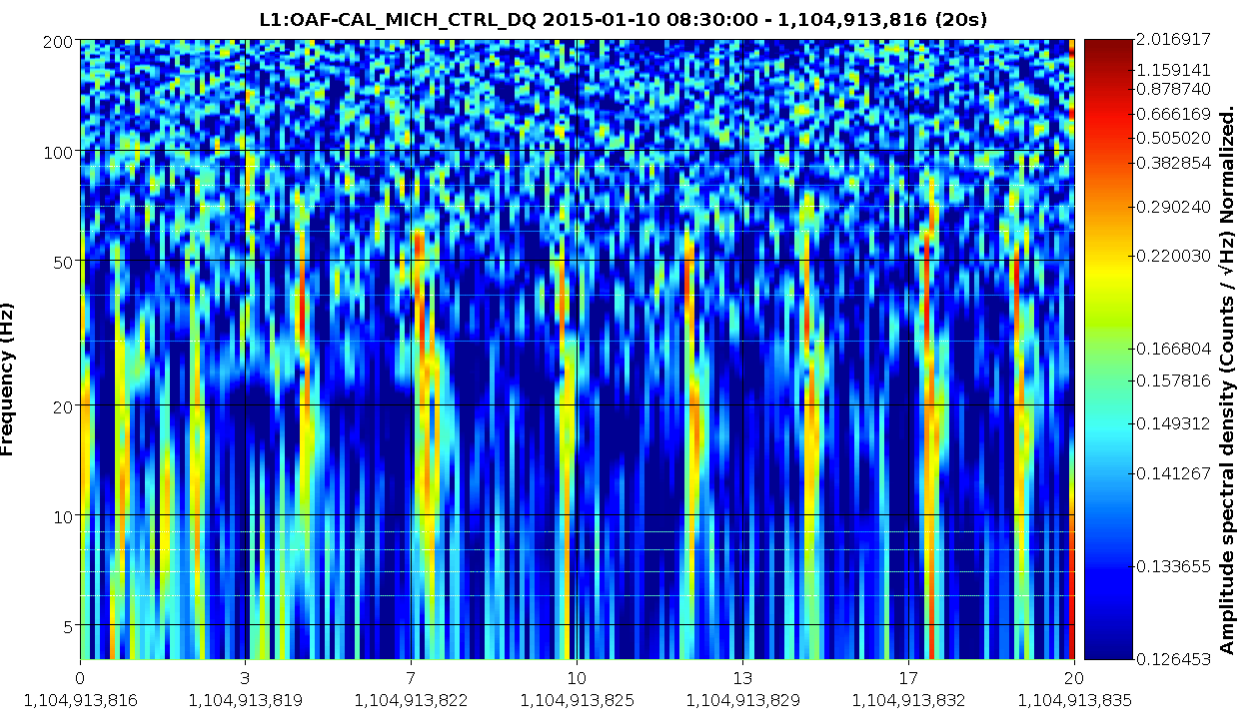
After ER6, loud
glitches were
observed in DARM

that coupled
with the
vertex
cavities (SRCL,
PRCL, MICH)

LLO alog
16354

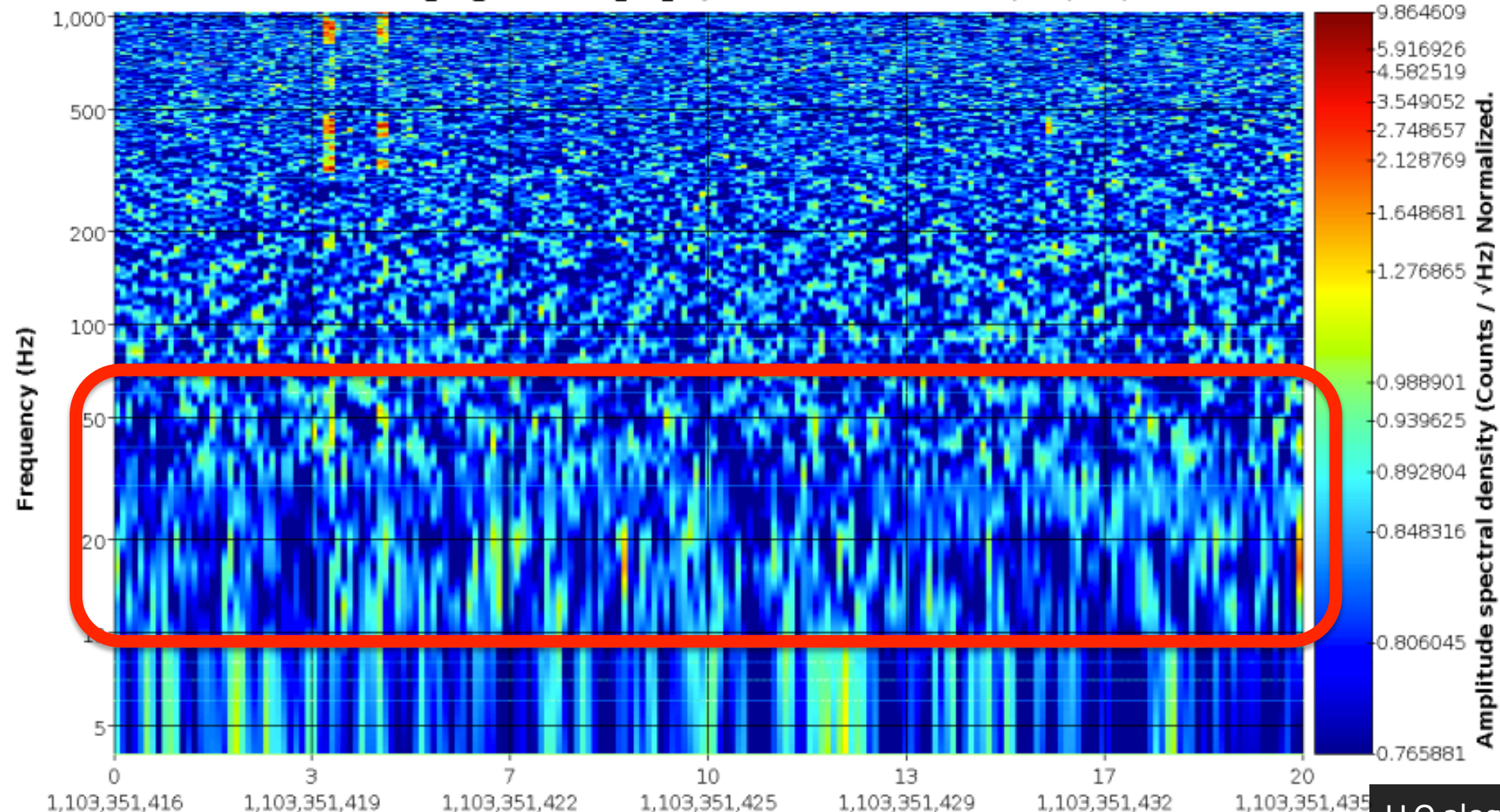


Wasn't DAC glitching fixed?



ER6 DAC calibration drift – week 1

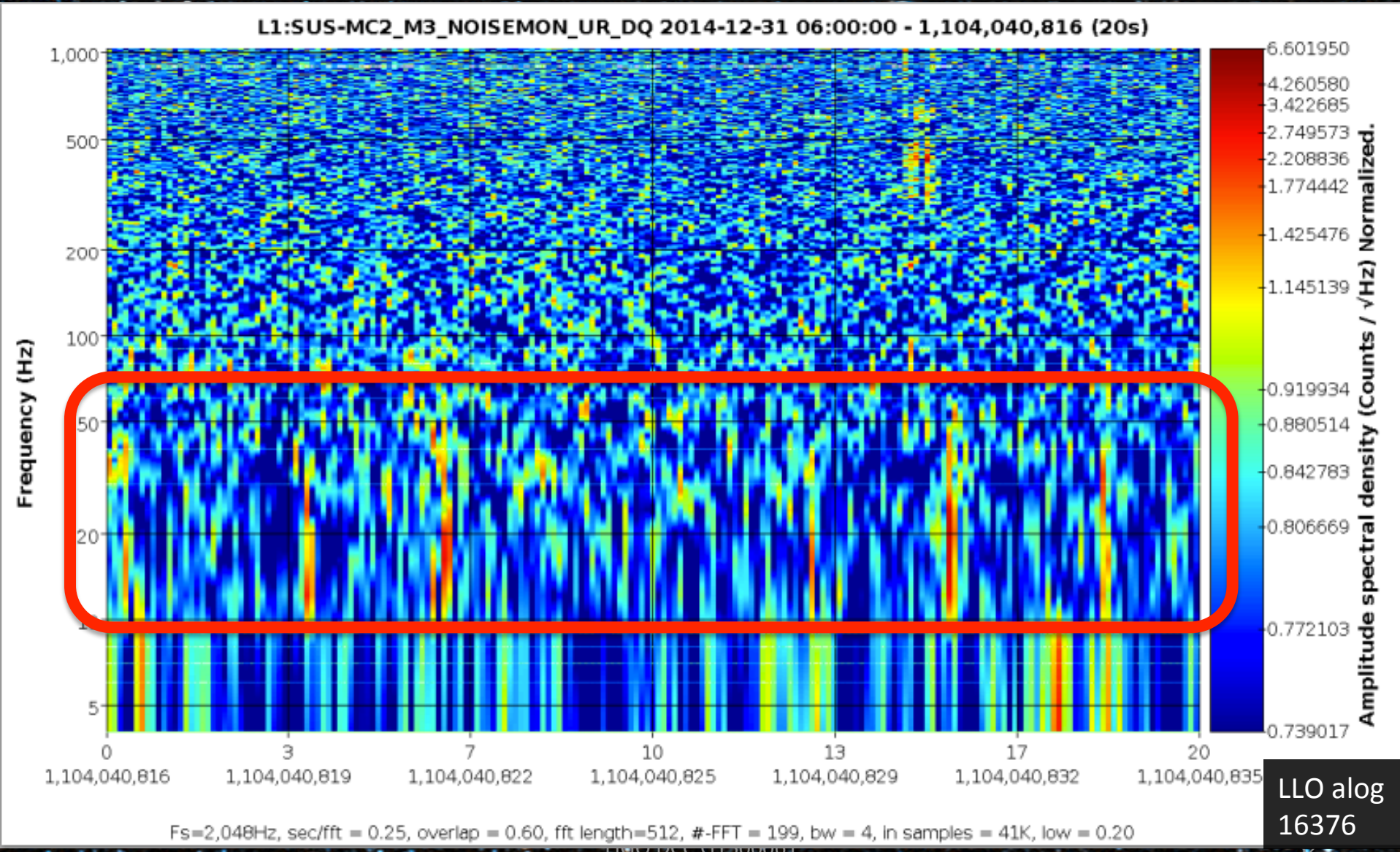
L1:SUS-MC2_M3_NOISEMON_UR_DQ 2014-12-23 06:30:00 - 1,103,351,416 (20s)



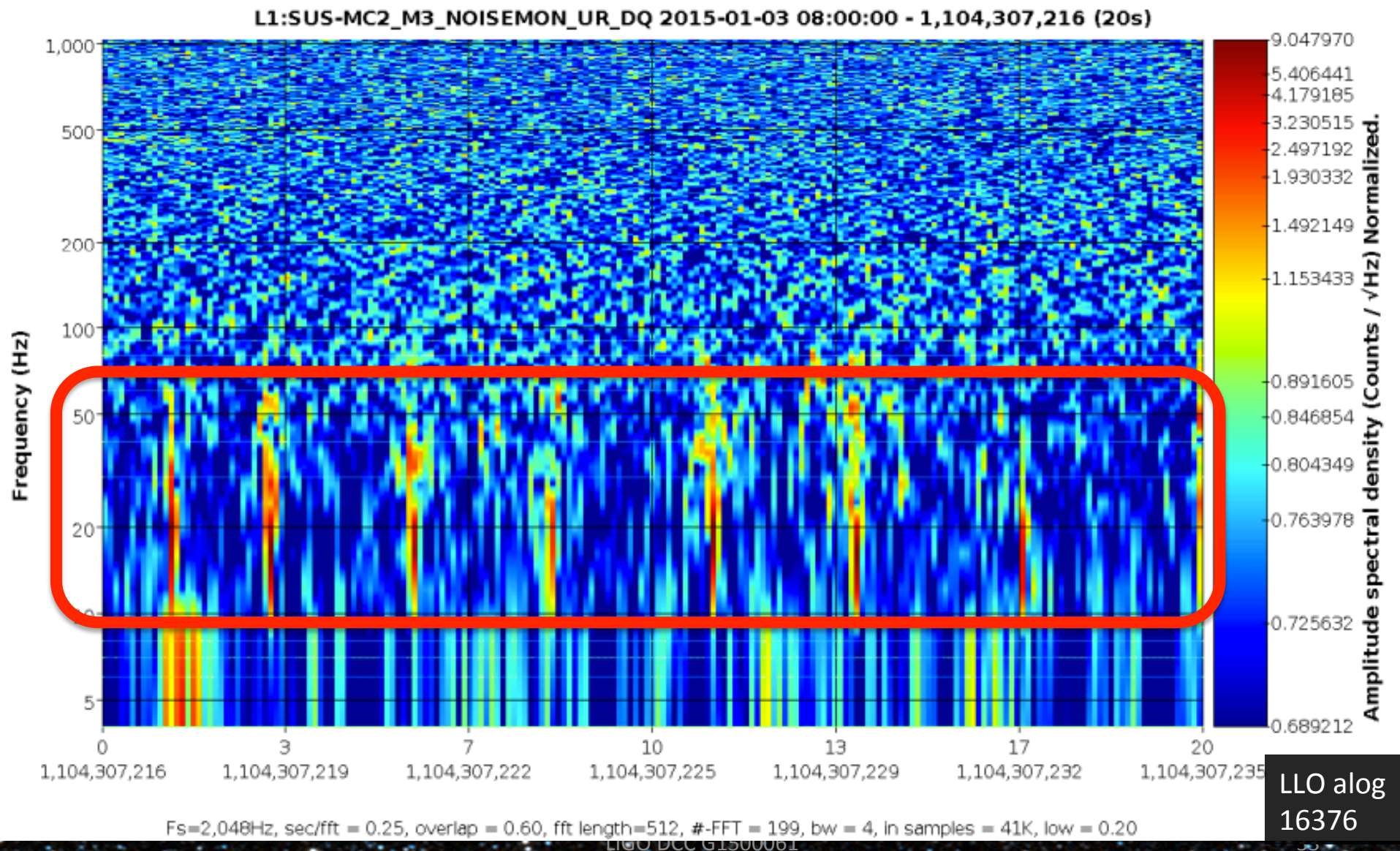
Fs=2,048Hz, sec/fft = 0.25, overlap = 0.60, fft length=512, #-FFT = 199, bw = 4, in samples = 41K, low = 0.20

LLO alog
16376

ER6 DAC calibration drift – week 2



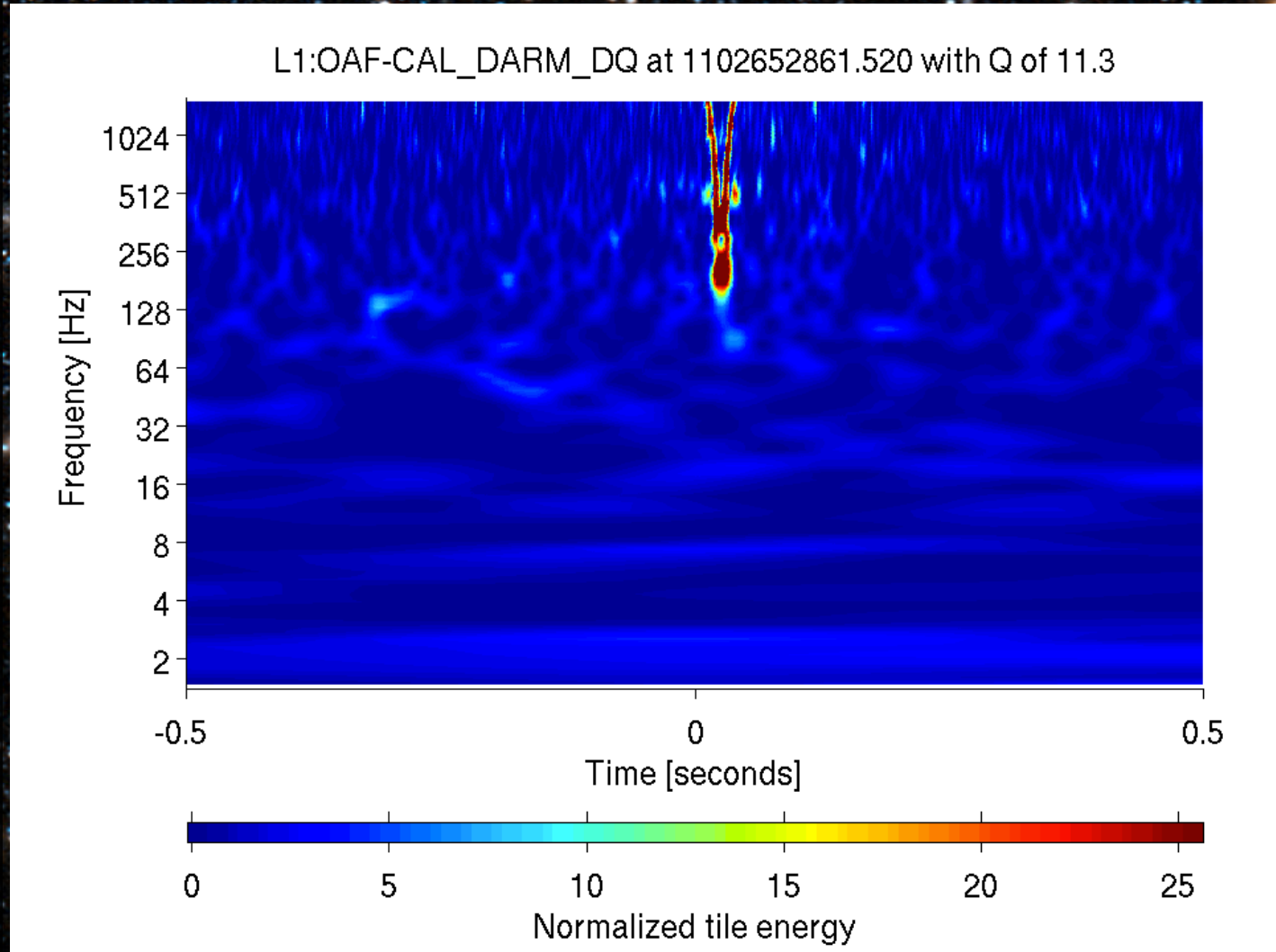
ER6 DAC calibration drift – week 3



IMC beat frequency

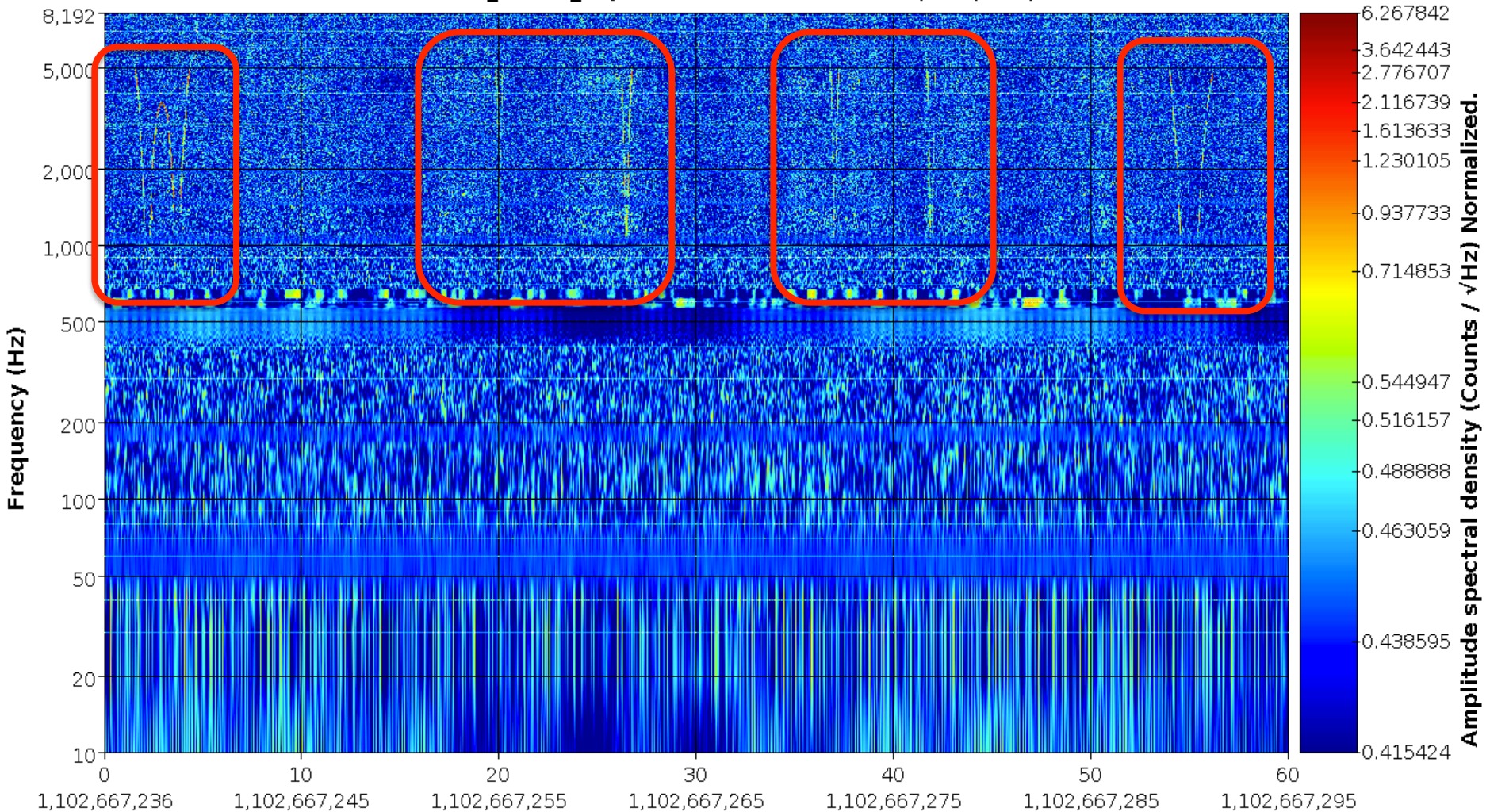
Distinct
“whistling”
feature
stemming from
a beat
frequency
(WRT -817kHz)
in the IMC that
adversely
affected the
BBH search in
ER6

Summary: LLO alog
16298



IMC beat frequency

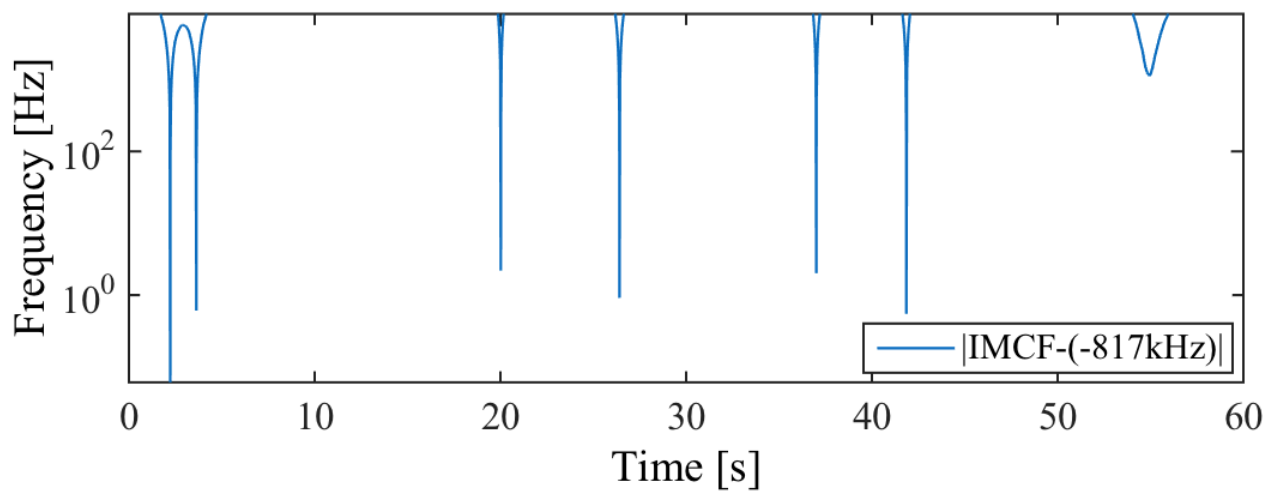
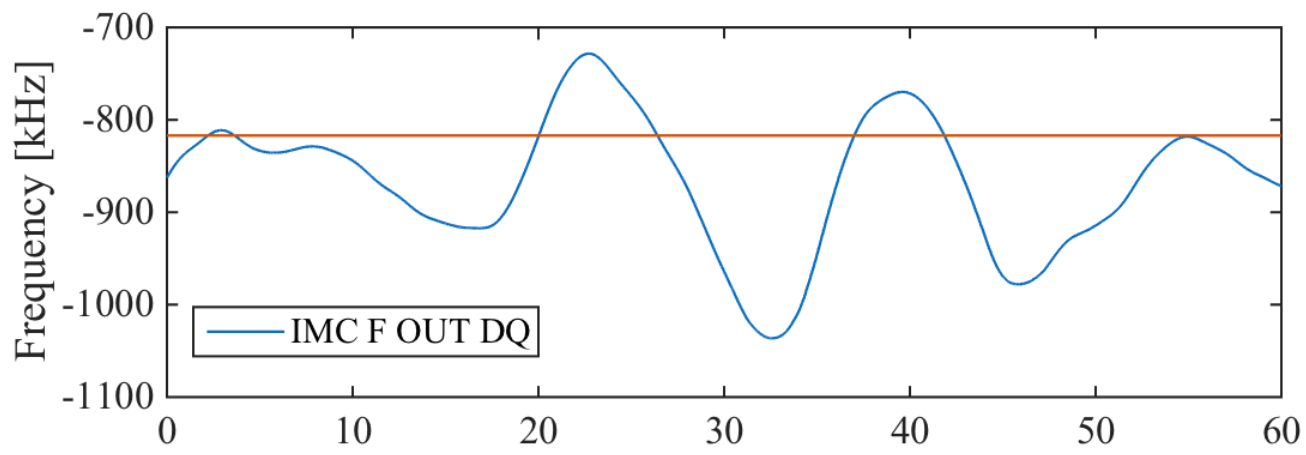
L1:OAF-CAL_DARM_DQ 2014-12-15 08:27:00 - 1,102,667,236 (60s)



Fs=16,384Hz, sec/fft = 0.10, overlap = 0.50, fft length=1,638, #-FFT = 1199, bw = 10, in samples = 983K, low = 0.20

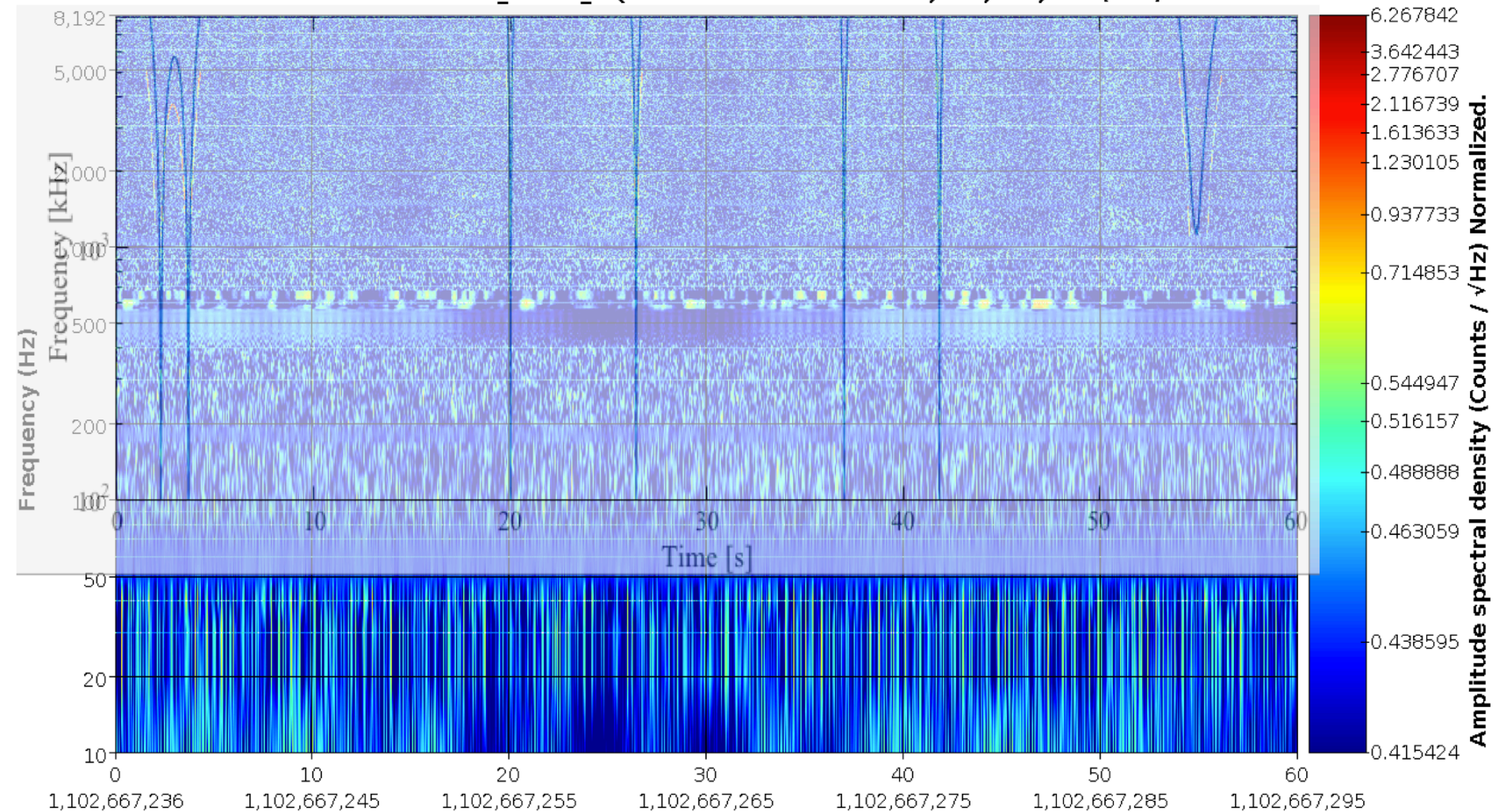
IMC beat frequency

Summary:
LLO alog
16298



IMC beat frequency

L1:OAF-CAL_DARM_DQ 2014-12-15 08:27:00 - 1,102,667,236 (60s)

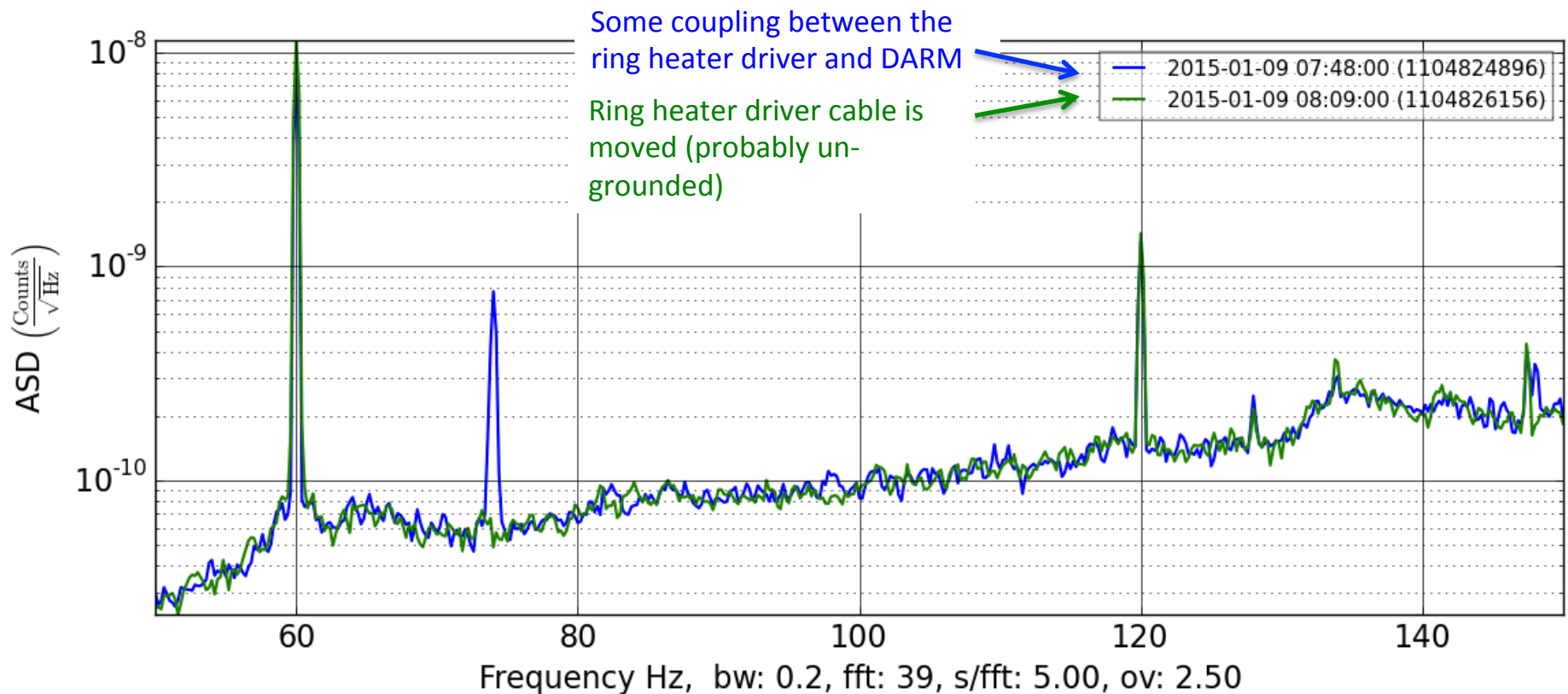


Fs=16,384Hz, sec/fft = 0.10, overlap = 0.50, fft length=1,638, #-FFT = 1199, bw = 10, in samples = 983K, low = 0.20

74 (72) Hz line – related to EY ring heater driver

LLO alogs
16316, 16291

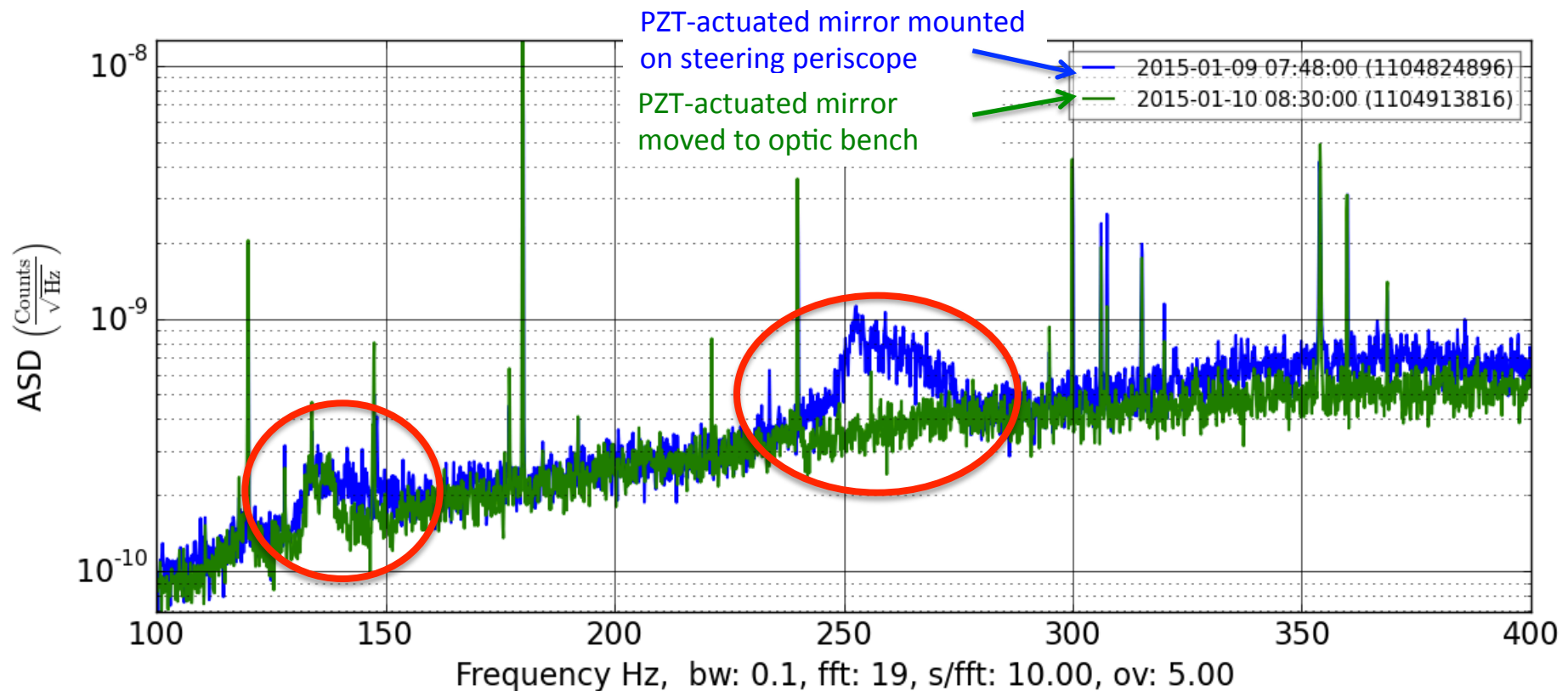
L1:OAF-CAL_DARM_DQ at 16384 Hz, t=100 s



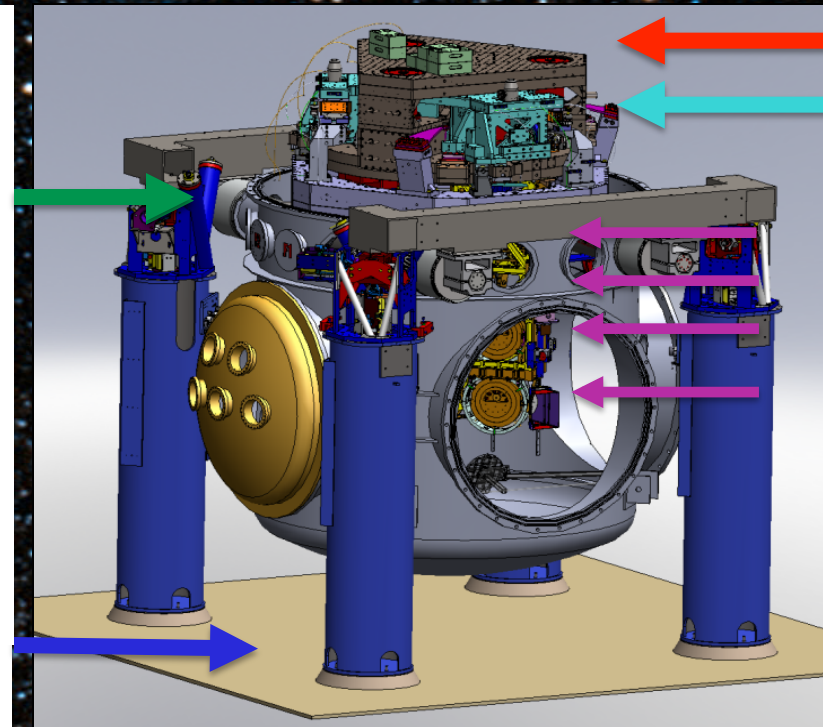
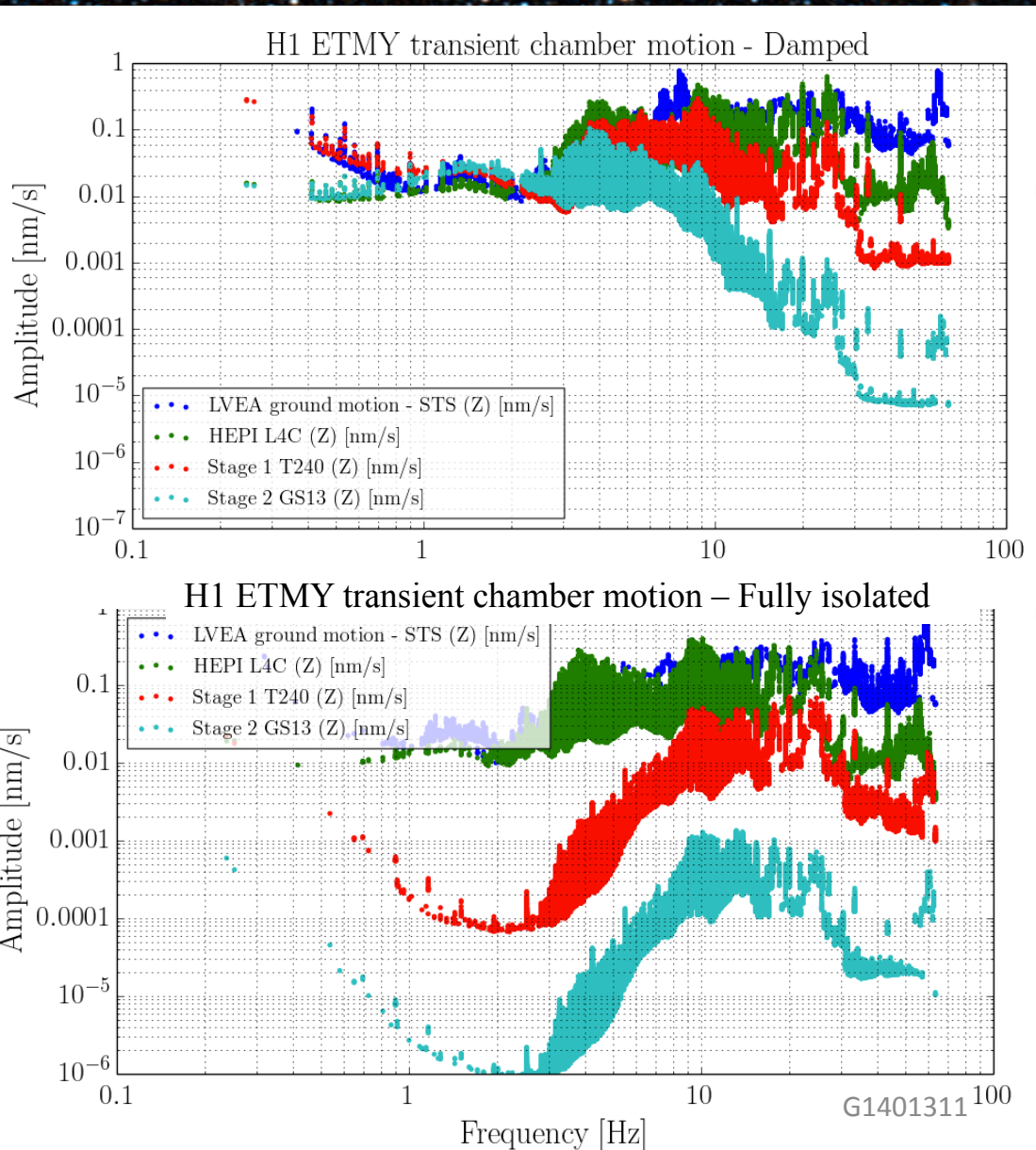
PZT-actuated mirror placement - 250Hz

LLO alog 16331

L1:OAF-CAL_DARM_DQ at 16384 Hz, t=100 s



SEI transient propagation

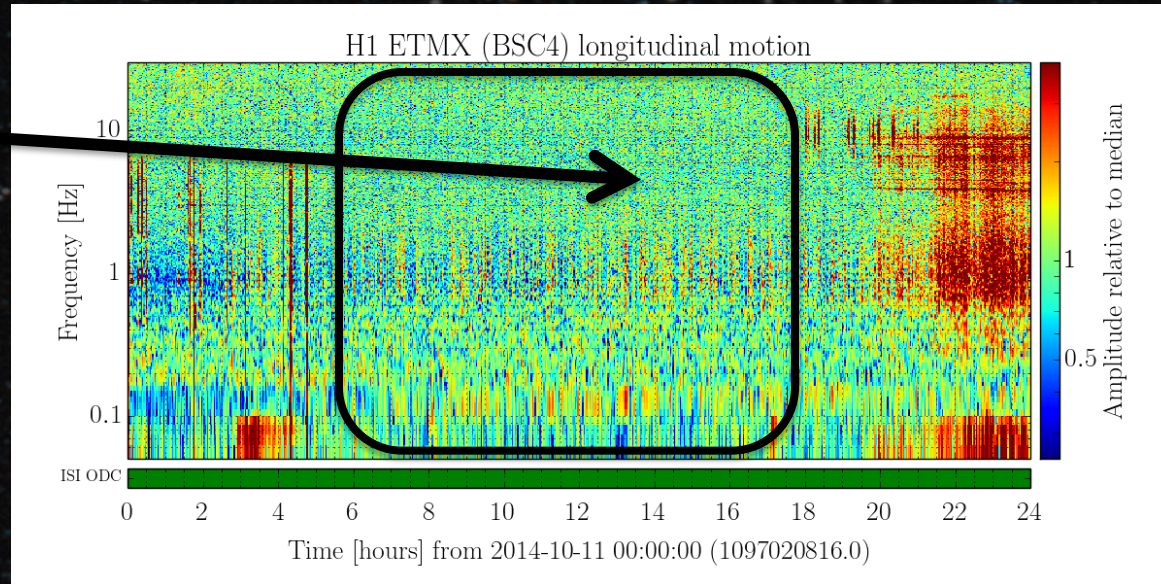


Tracing the transient motion from the ground to the optic table in various states of isolation loop aggression.

More aggressive isolation mitigates transients well < ~15Hz

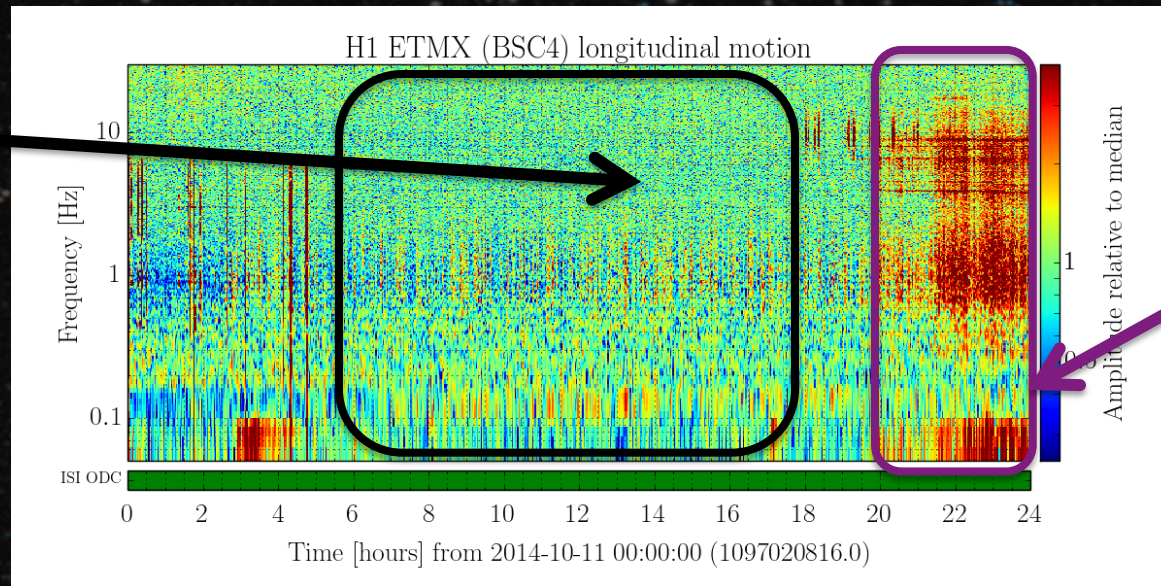
Windy vs. Quiet time transient SEI study at LHO

Most of the day,
Oct 11, high
microseism, low
wind (~5MPH)



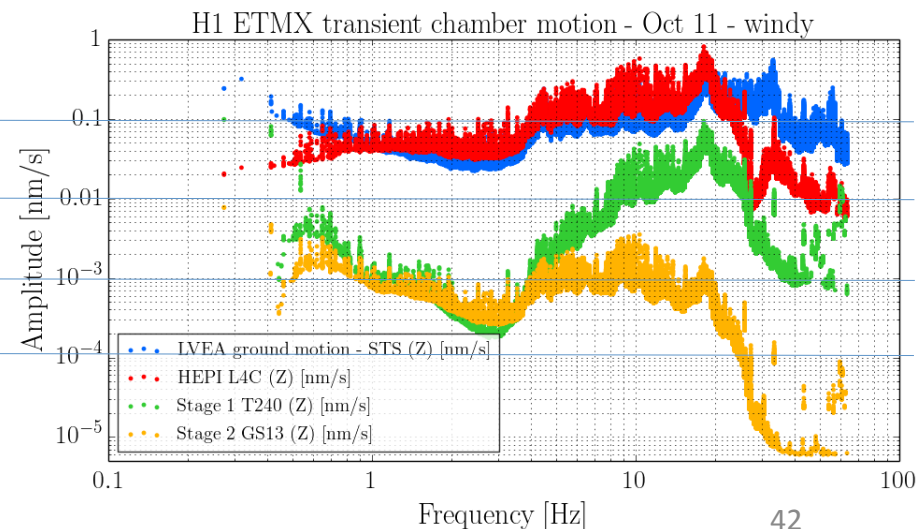
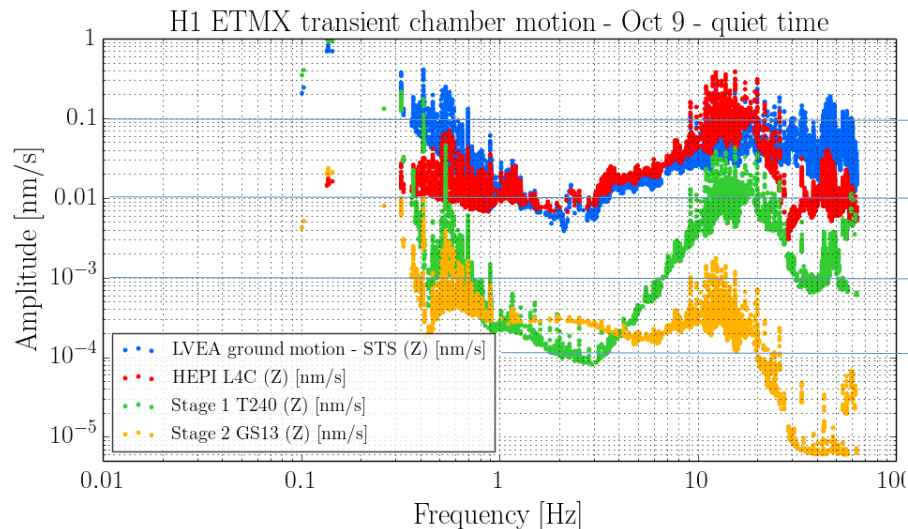
Windy vs. Quiet time transient SEI study at LHO

Most of the day, Oct 11, high microseism, low wind (~5MPH)



After ~20:00 UTC, high wind (~30MPH)

Below are Omicron triggers of two hours of “quiet” time (left) and “windy” time (right). Each dot is a transient event. Transient motion amplitude is very elevated during high wind for events of freq < ~30Hz.



Note: these plots show ETMX local ground motion, not LVEA

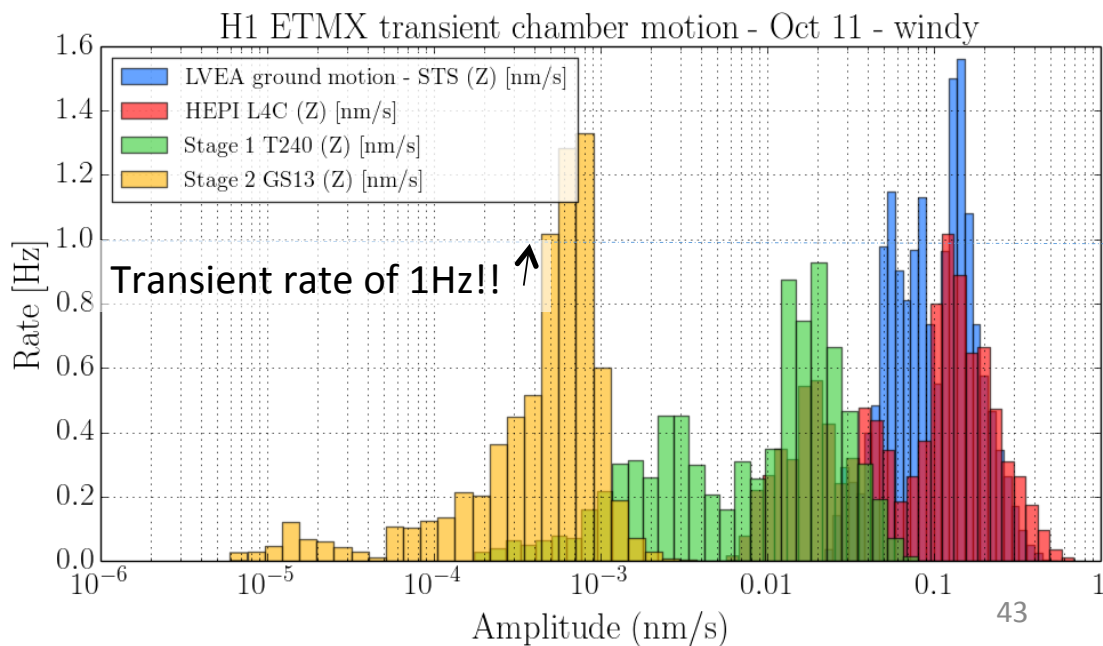
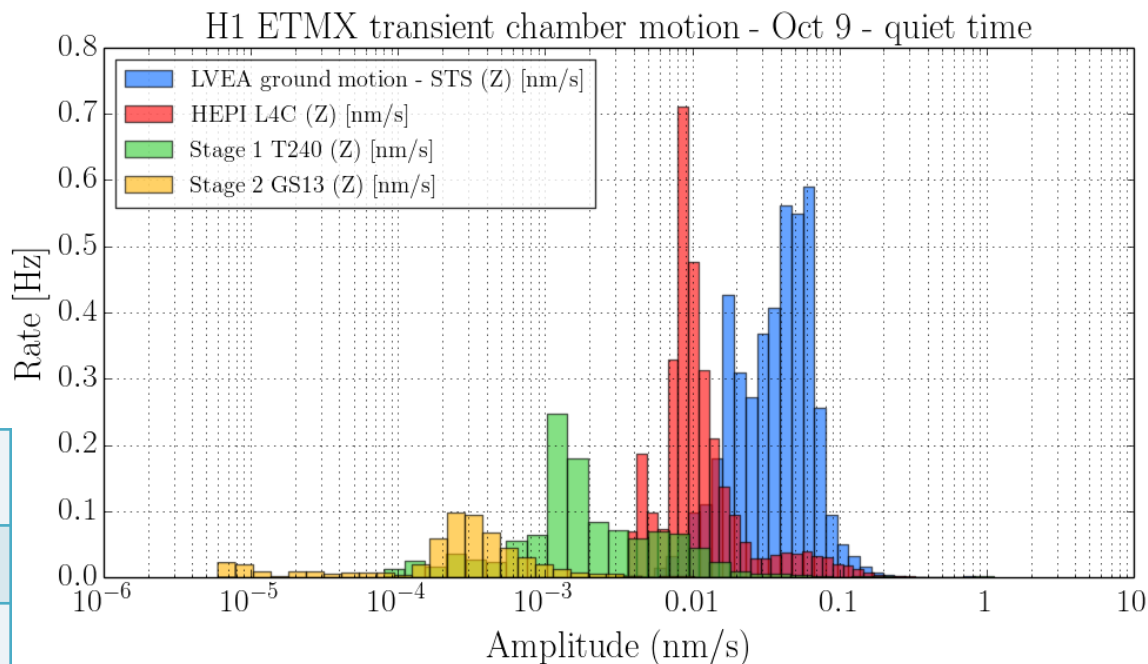
The rate of transient motion events also increases dramatically during windy time – by over a factor of 10 in optic table motion at end X.

Stage	Quiet* (# trigs)	Windy* (# trigs)	Factor increase
Ground motion	30,755	116,601	3.8
HEPI (L4C)	21,317	74,624	3.5
ISI ST1 (T240)	7,791	57,948	7.4
ISI ST2 (GS13)	3,924	49,562	12.6

* For a two hour period of relatively quiet or windy time

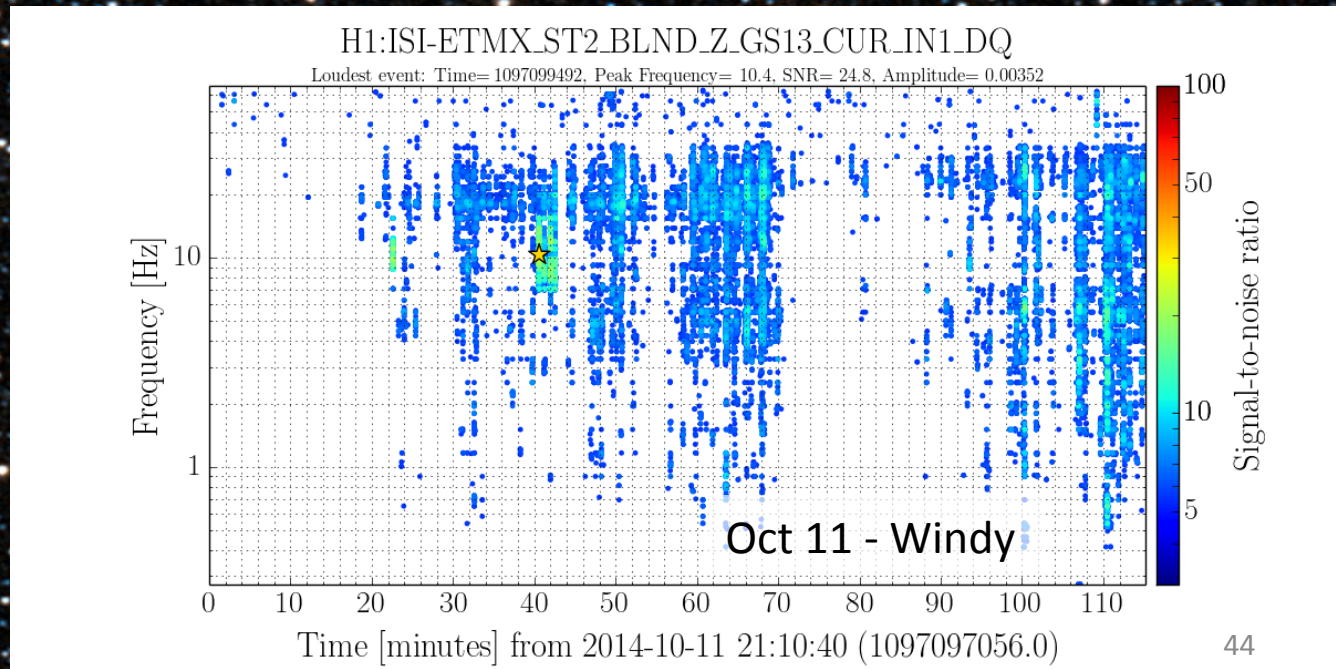
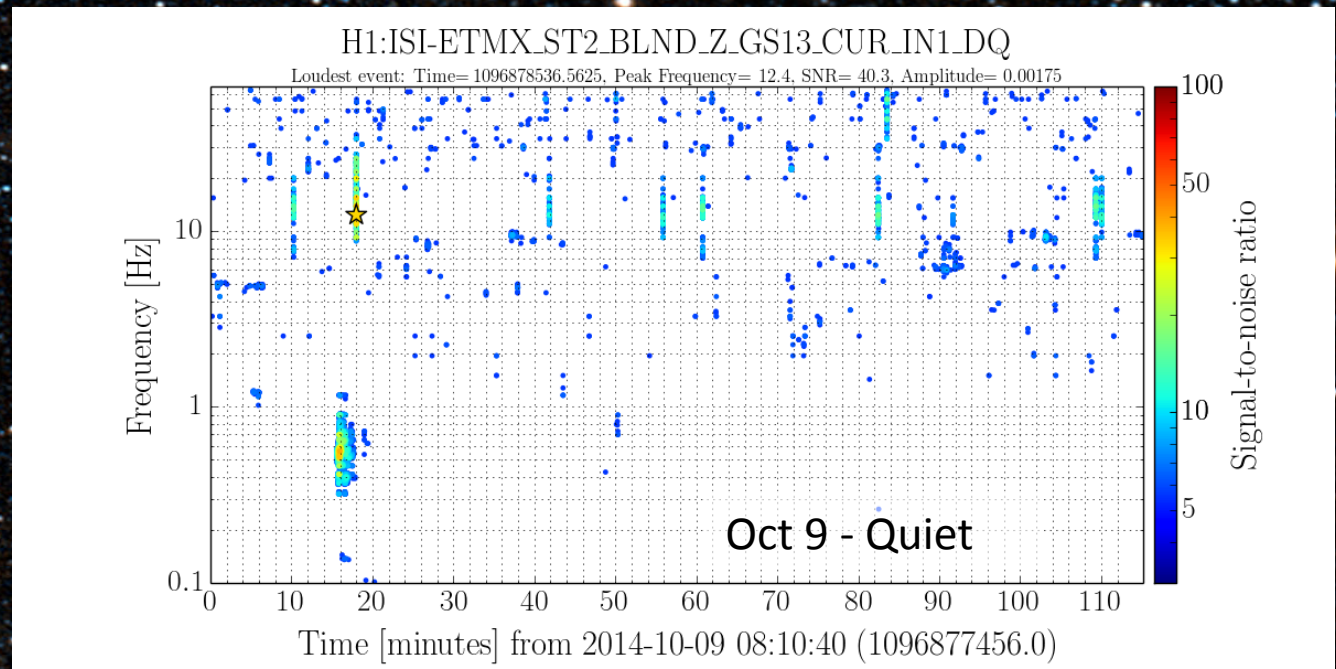
Isolated stages see a much greater increase in the rate of transients than ground motion during windy time.

Note: these plots show ETMX local ground motion, not LVEA



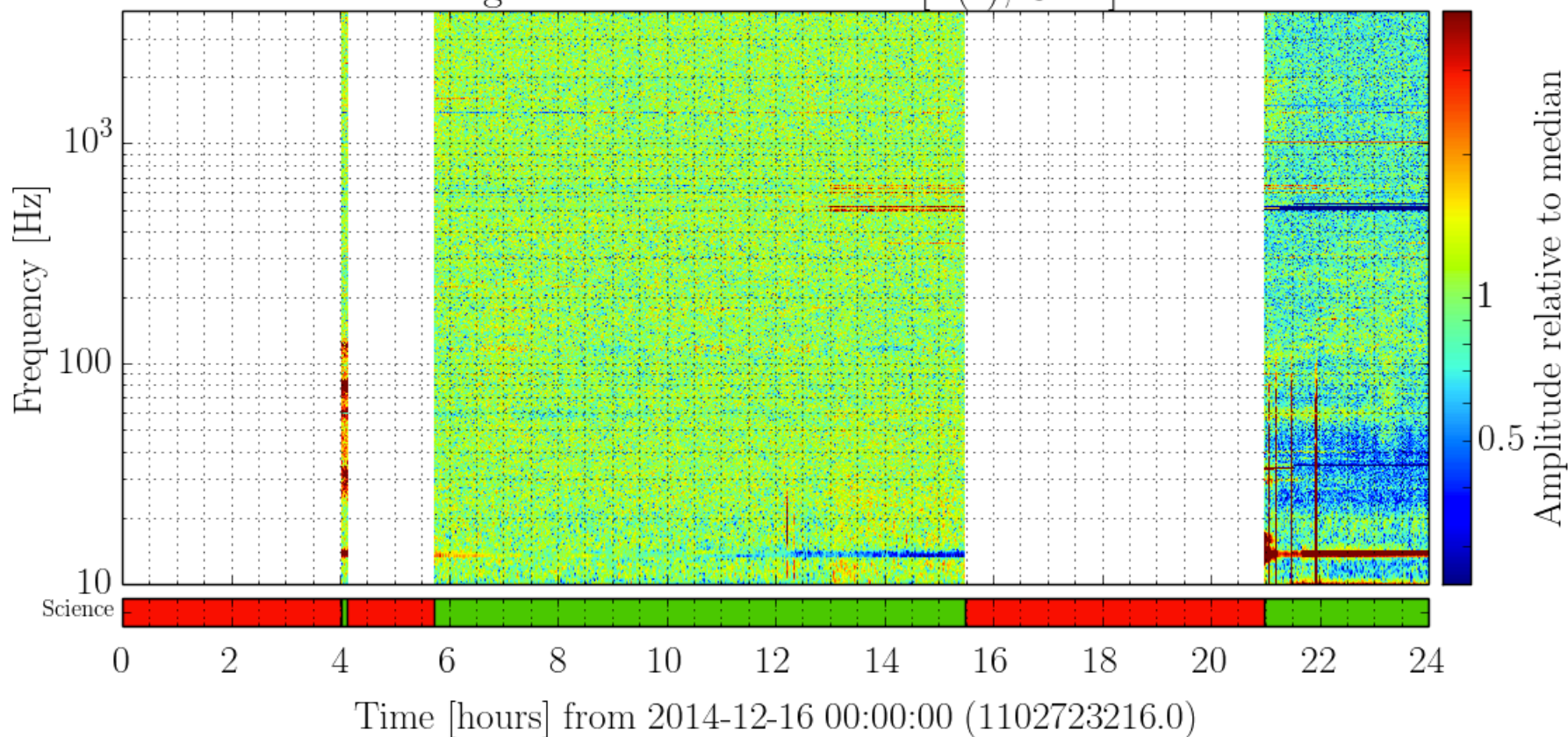
Conclusions for windy vs. quiet SEI transient study:

Ultimately, at the optic table the transient motion amplitude per event isn't significantly increased above 10-15Hz, but the **rate** of transients is greatly increased

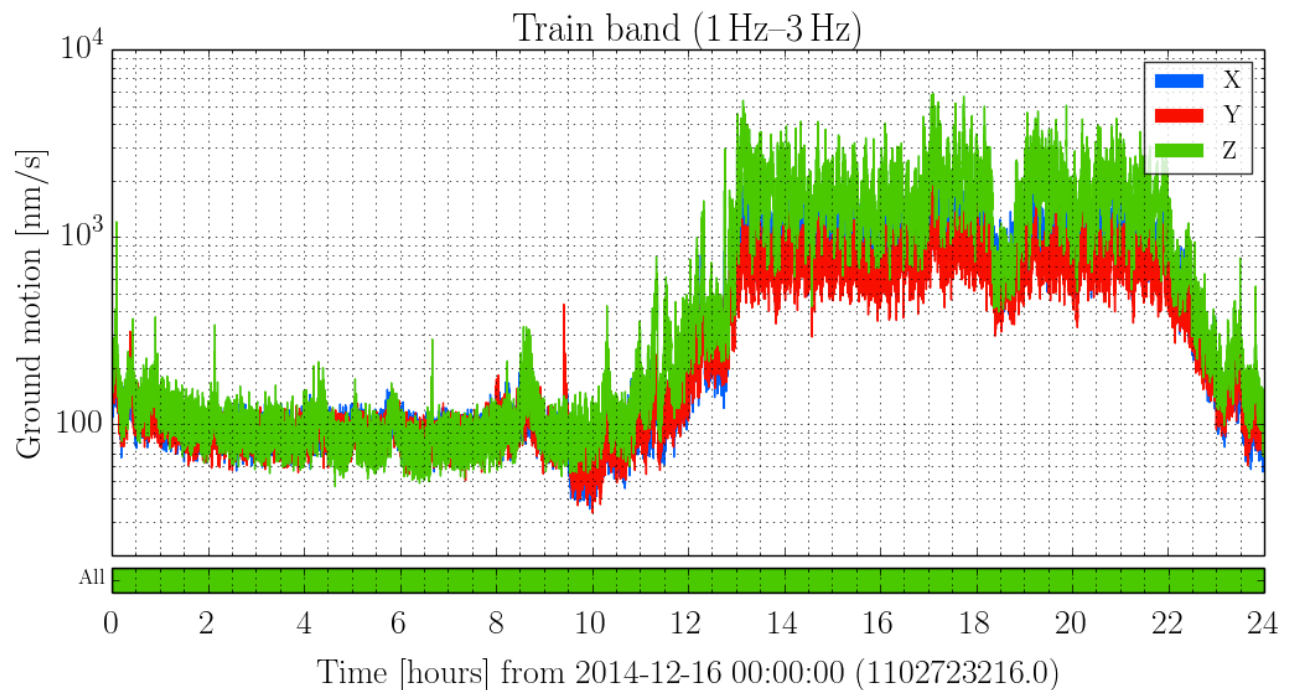
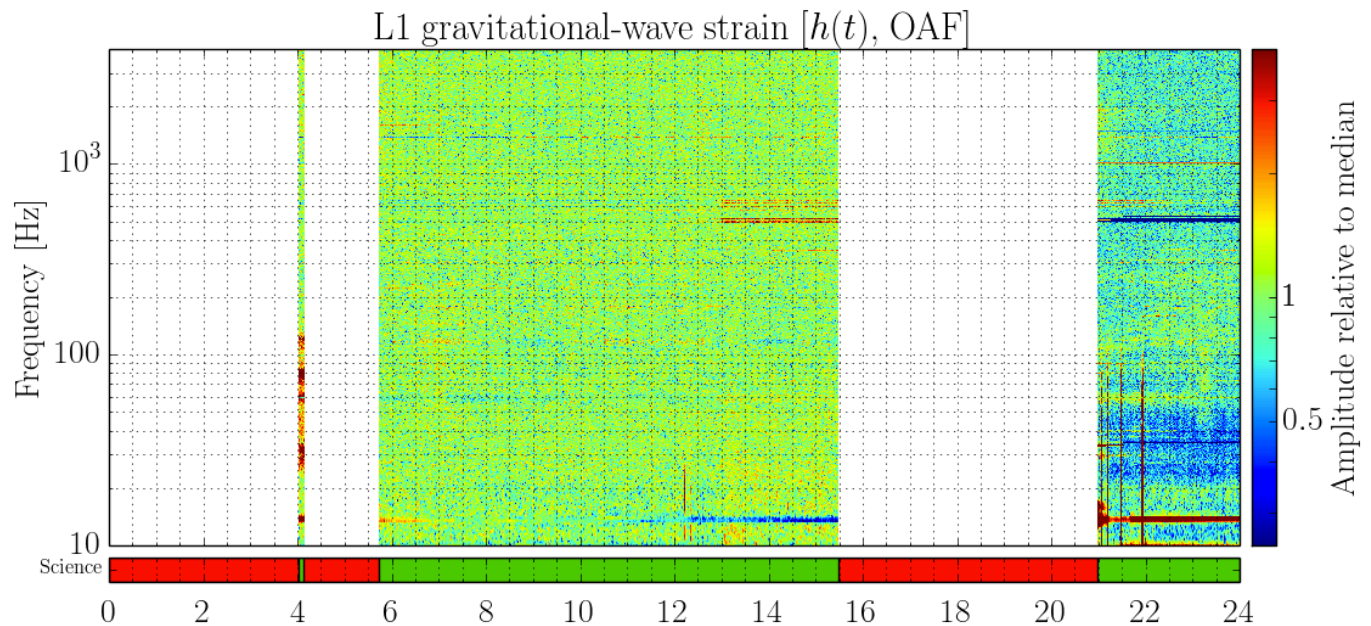


Livingston ER6 -Dec 16 lock

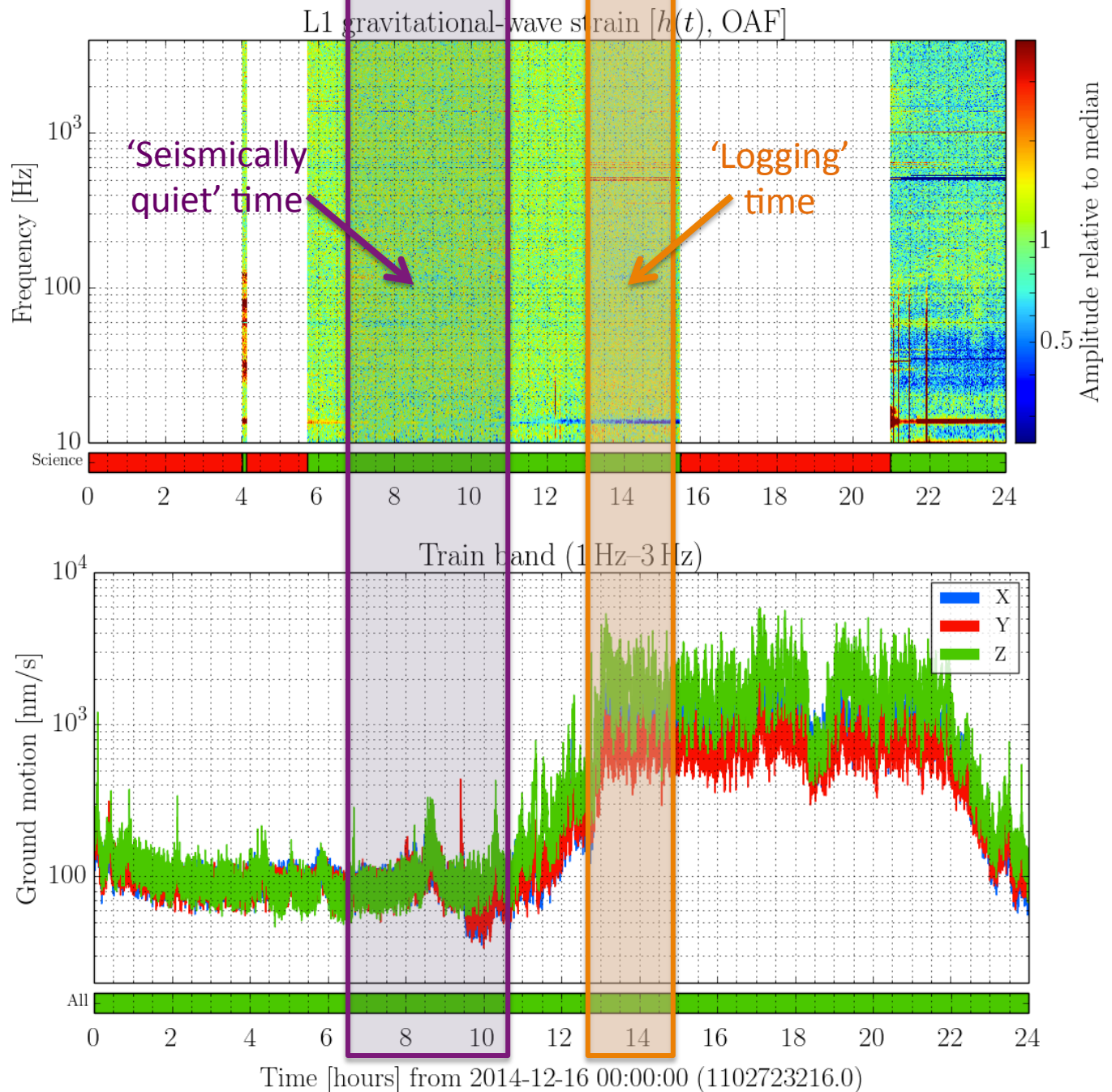
L1 gravitational-wave strain $[h(t), \text{OAF}]$



LLO alog
16101:
“violin
modes
appear to
be causing
high trigger
rate in CBC
search” (als
o reported
by cWB)

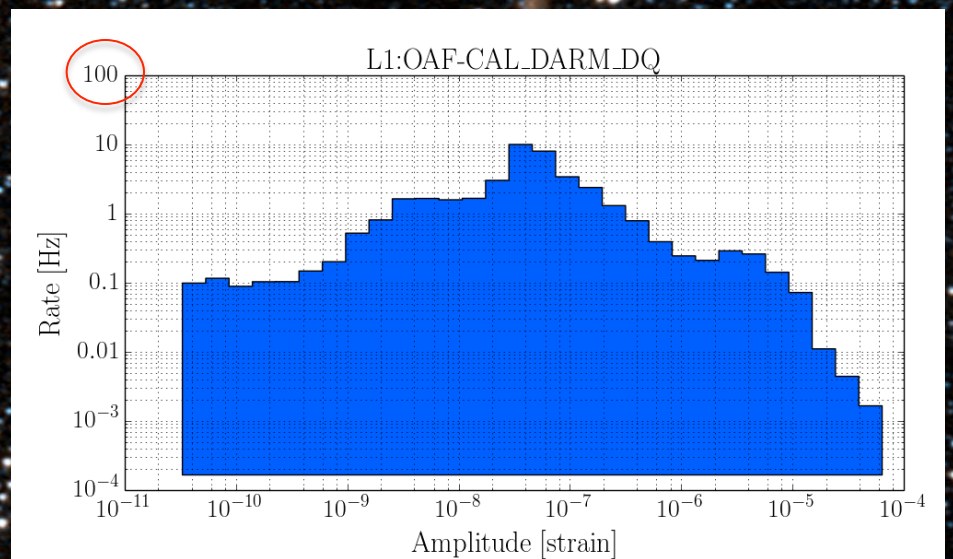
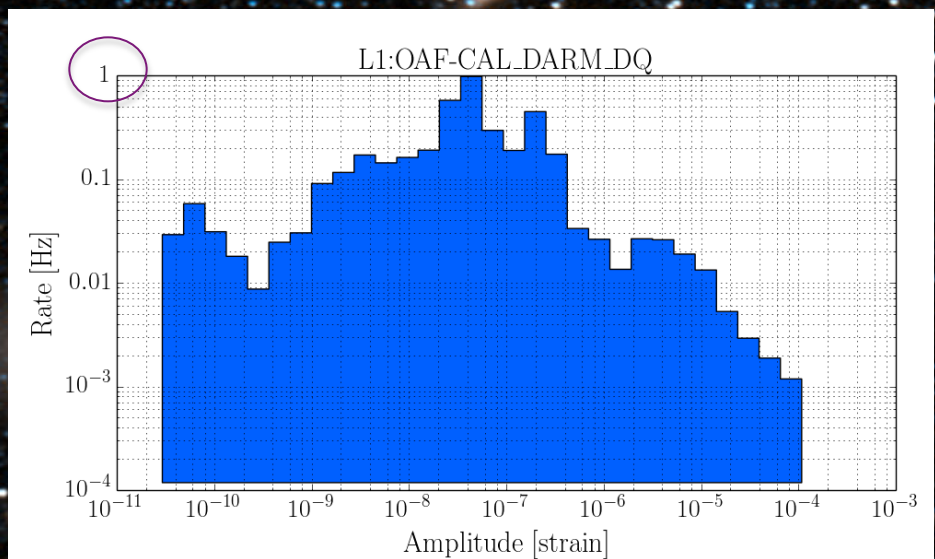
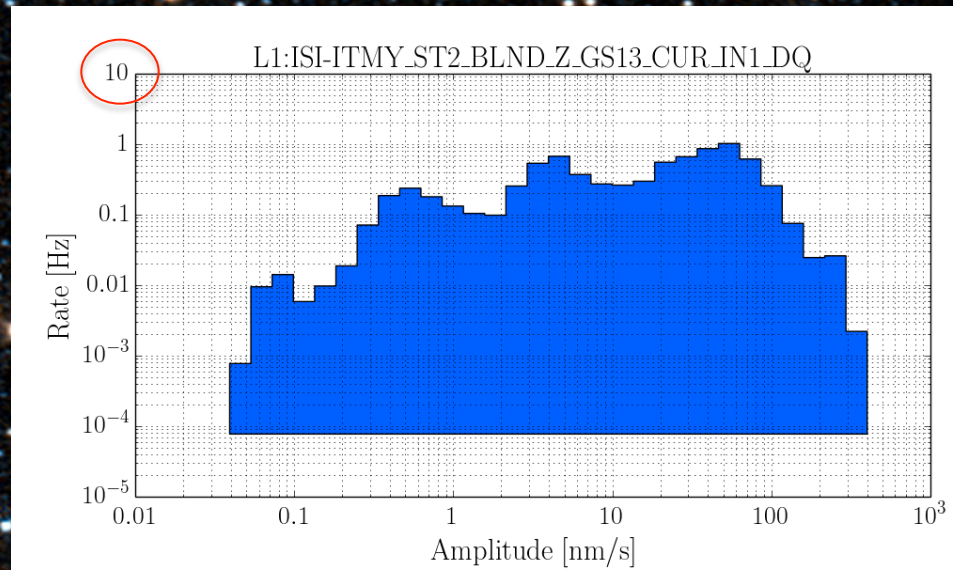
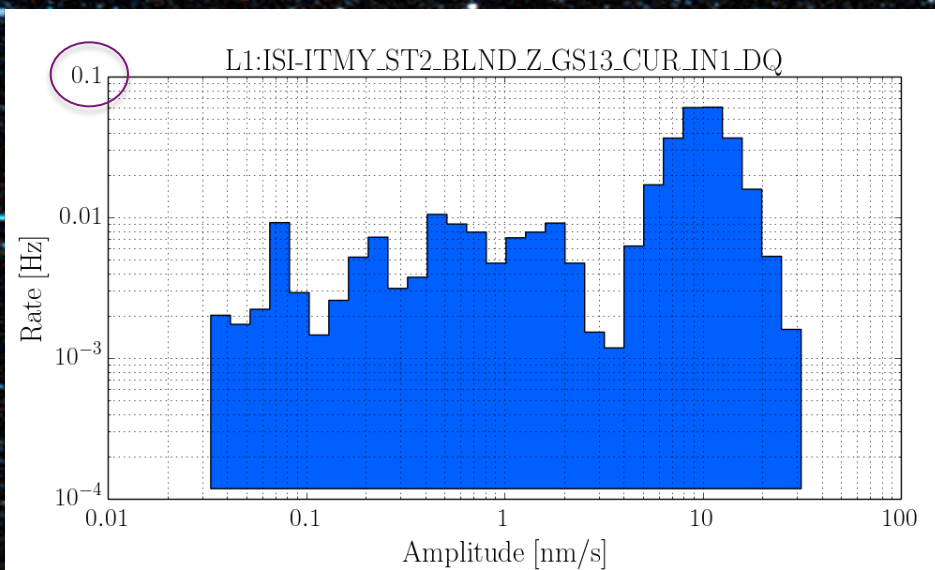


LLO alog
16101:
“violin
modes
appear to
be causing
high trigger
rate in CBC
search” (als
o reported
by cWB)



Seismically quiet part of lock

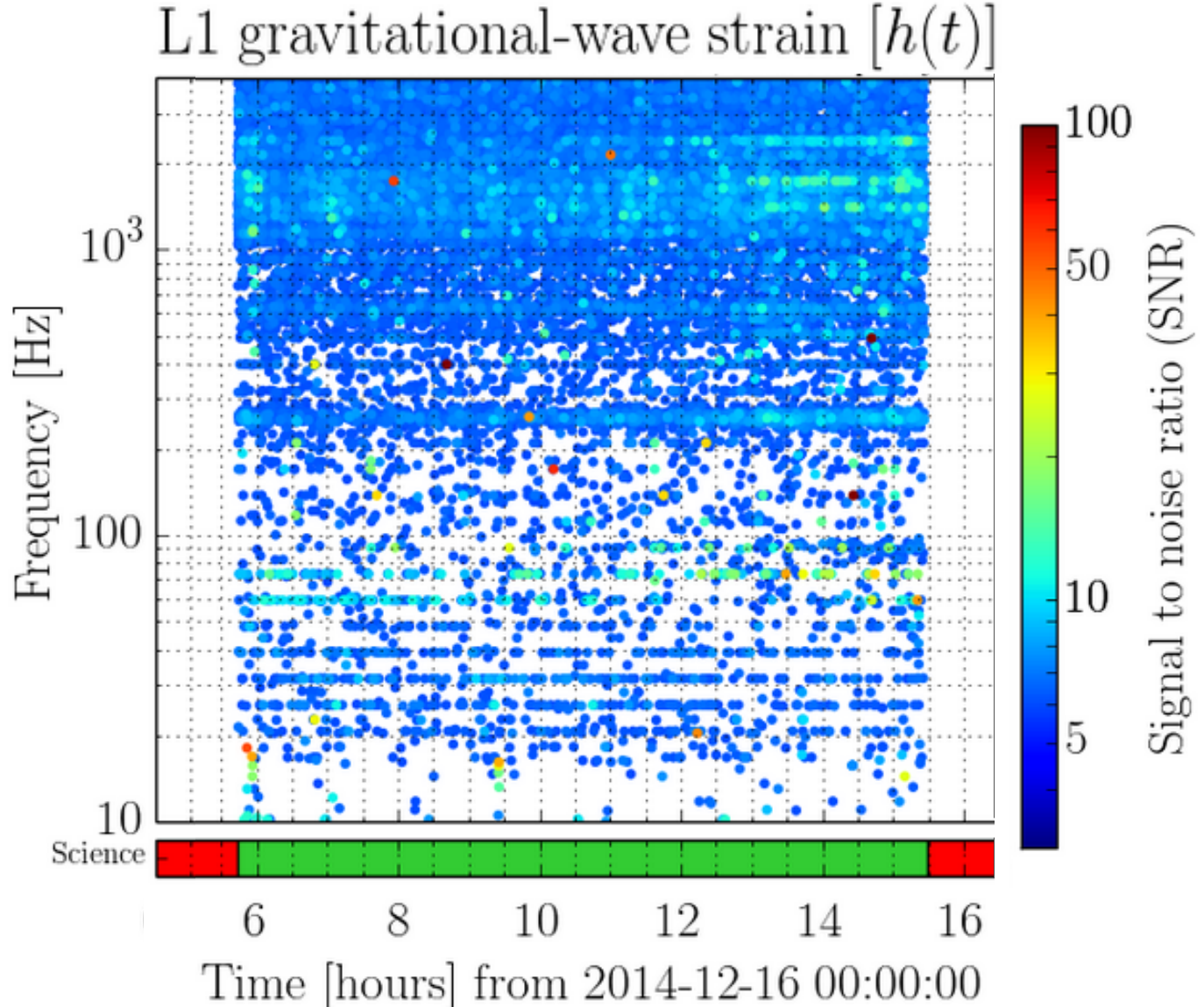
Logging stretch of lock



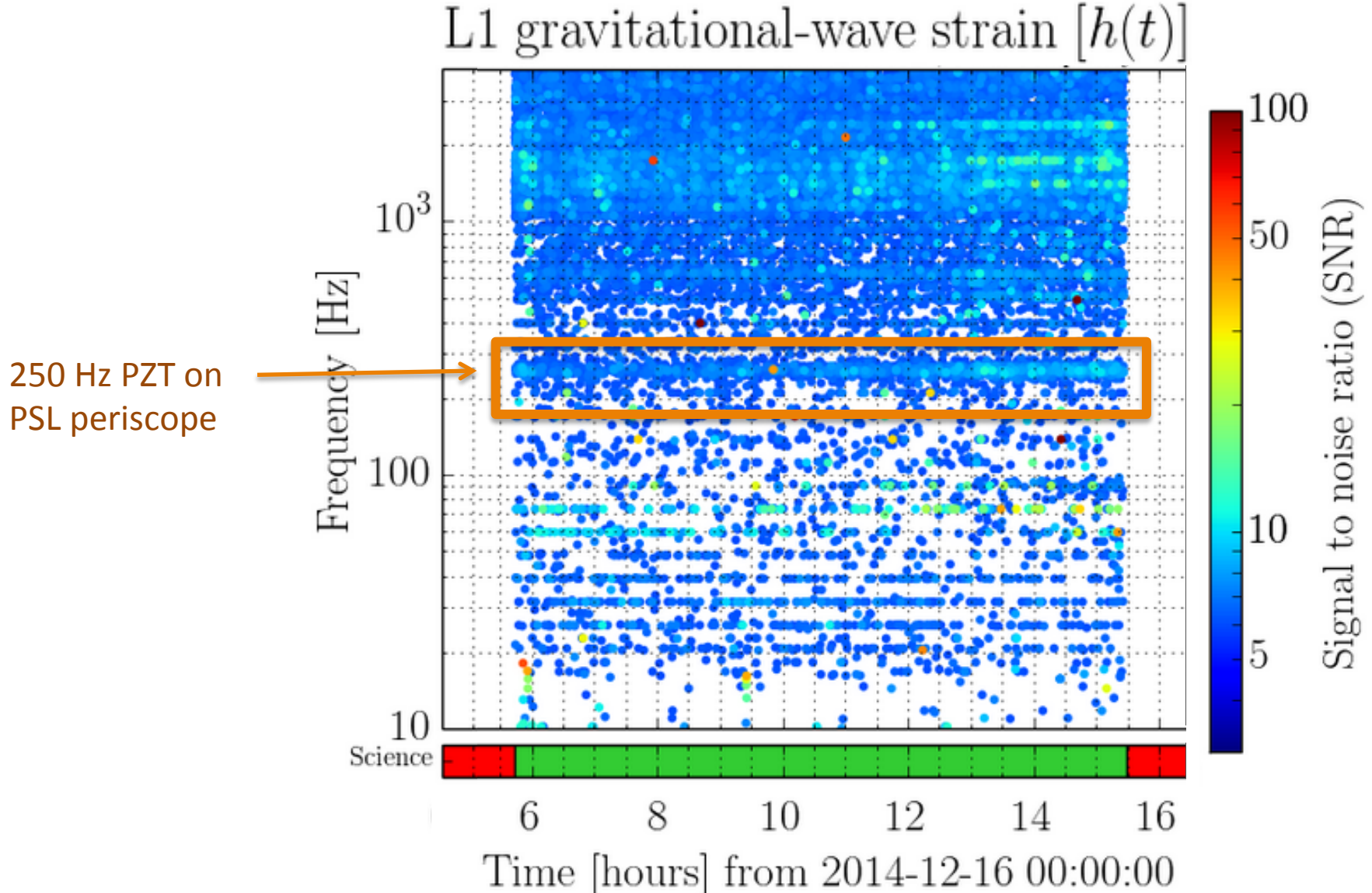
Summary - Dec 16 lock

ER6 glitch rate
looked very
clean on the
whole!

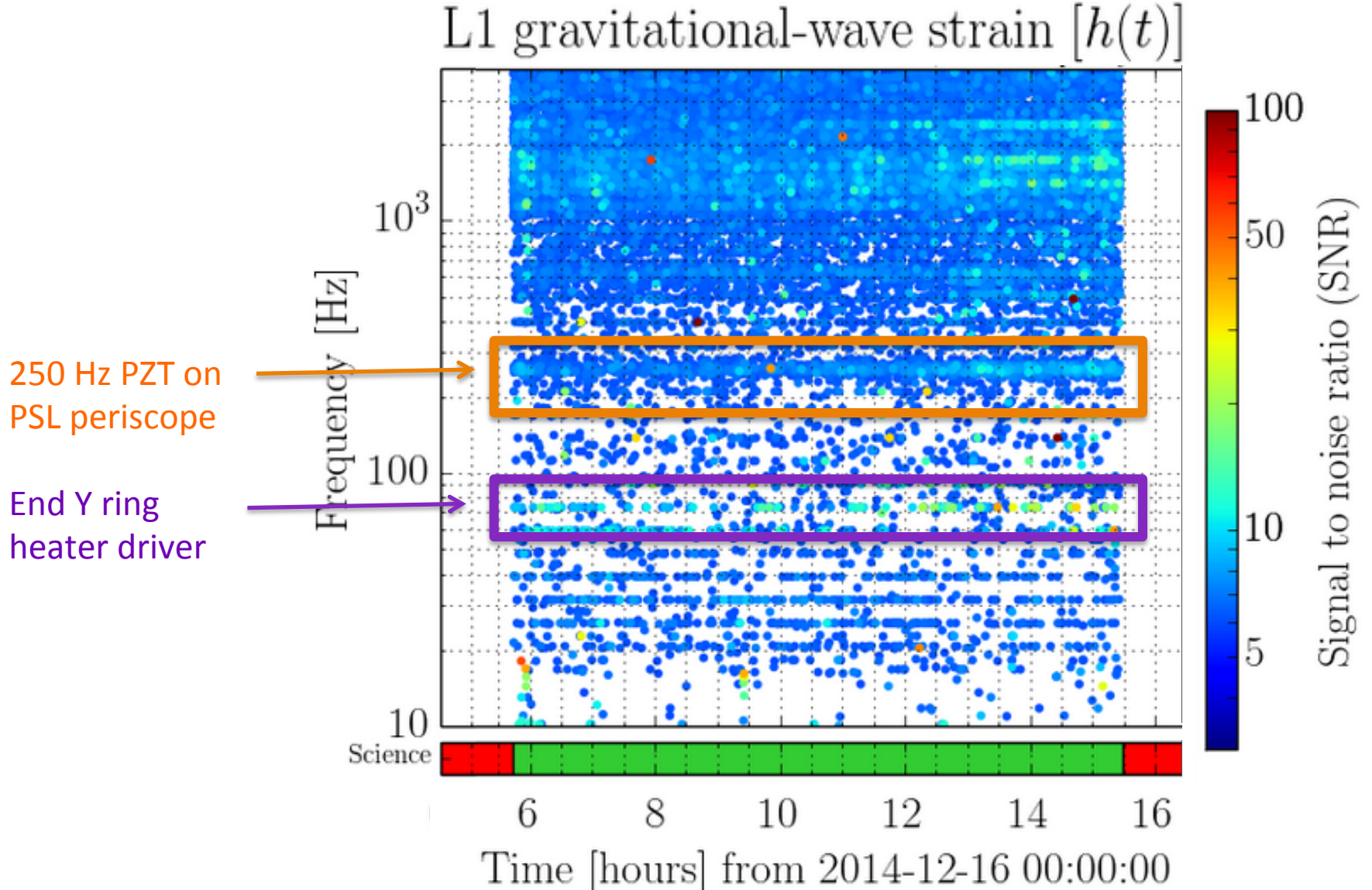
Most features
that would
affect the
transient
searches are
now
understood/
resolved



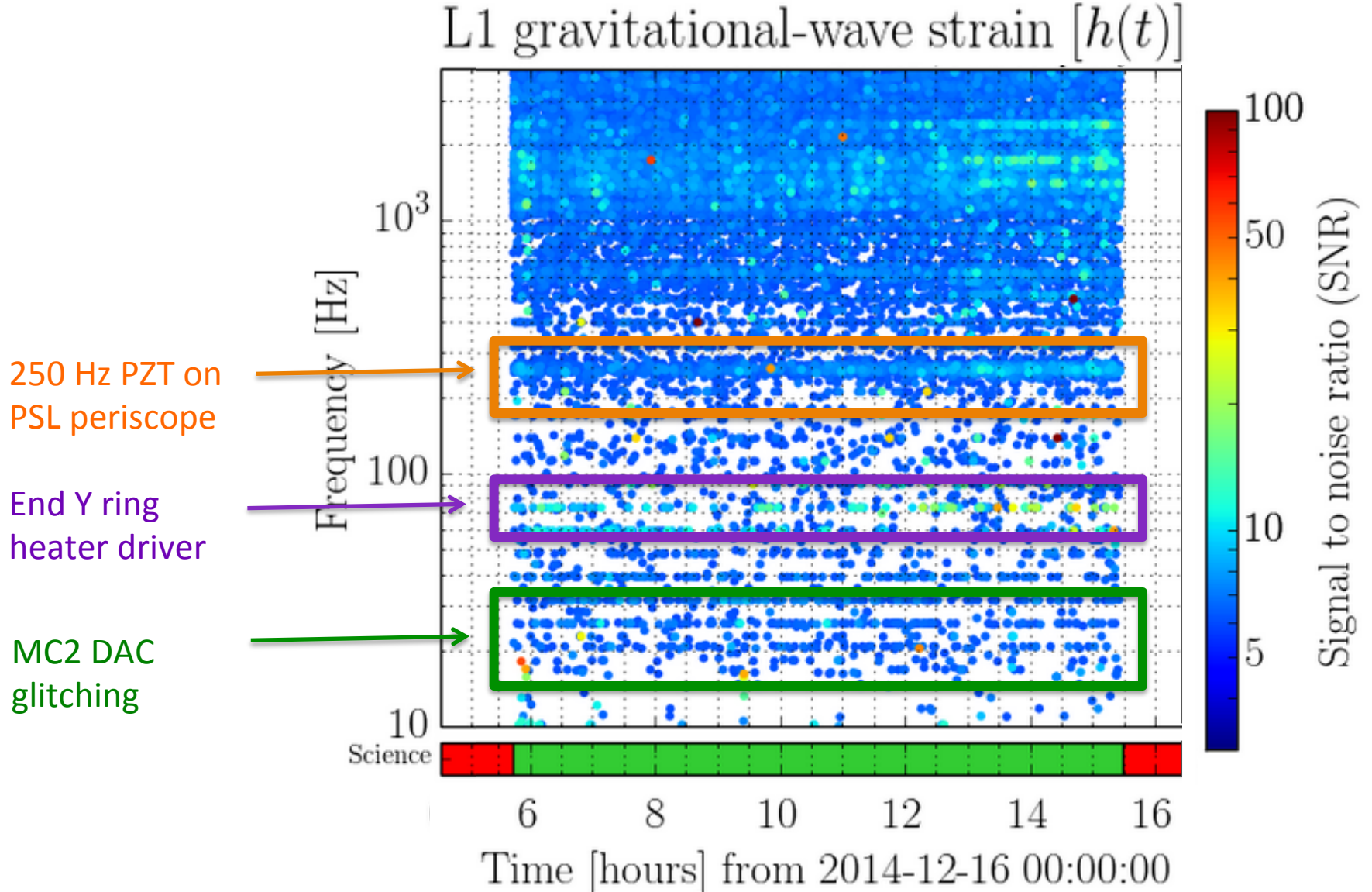
Summary - Dec 16 lock



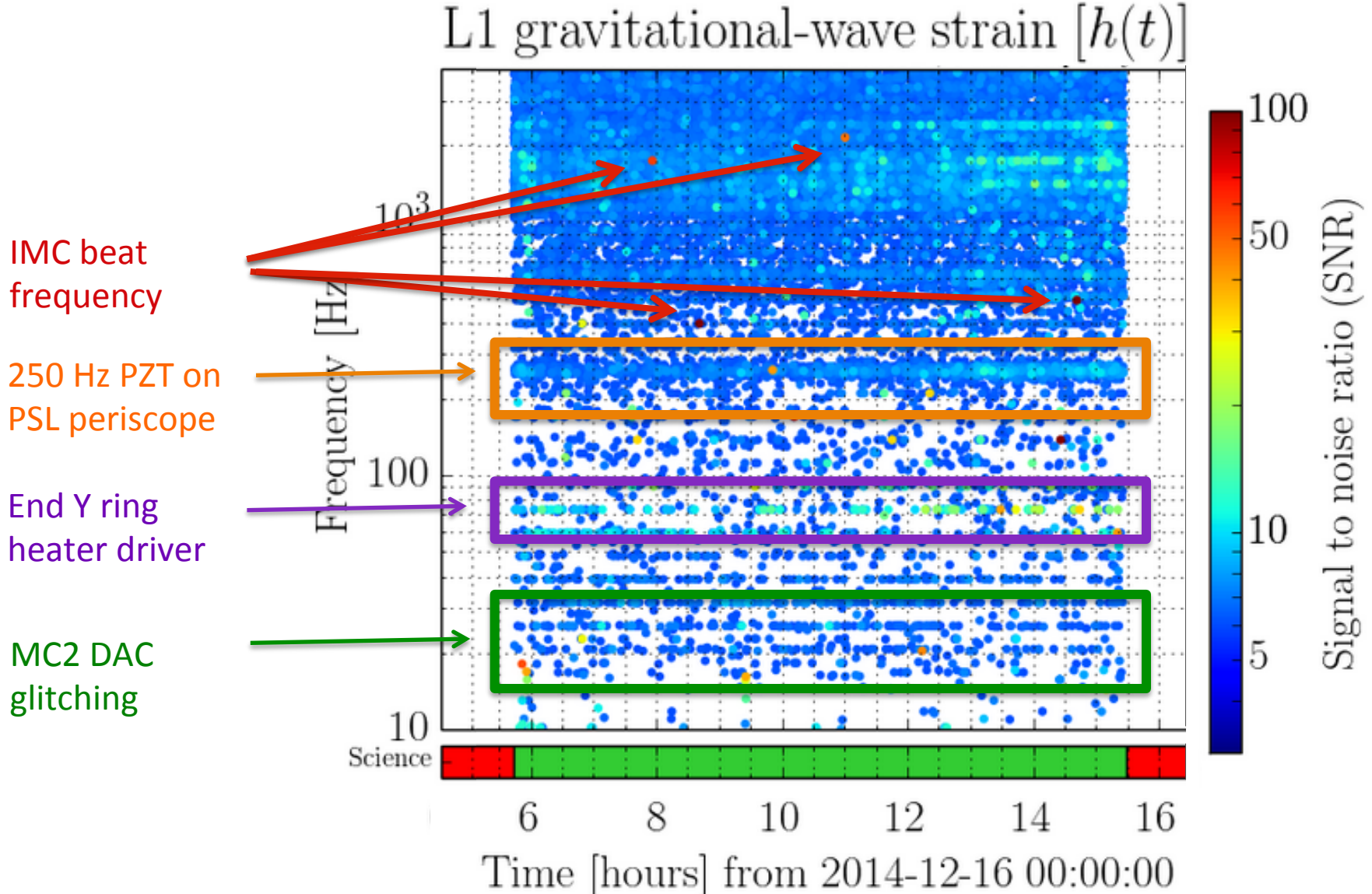
Summary - Dec 16 lock



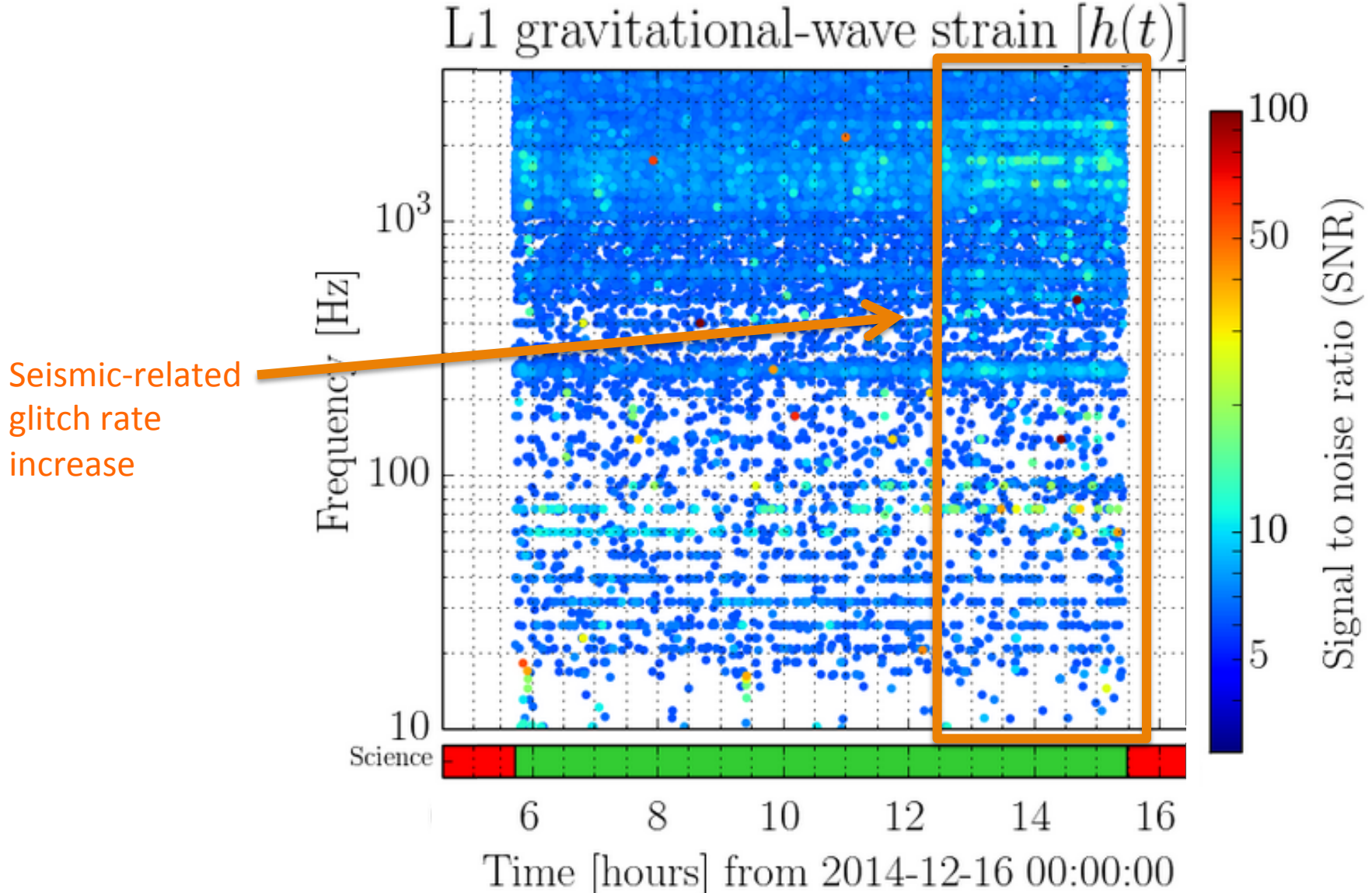
Summary - Dec 16 lock



Summary - Dec 16 lock



Summary - Dec 16 lock



Conclusions

- Thanks largely to commissioning efforts, LLO data looks GREAT! Greatly improved glitch rate compared to August.
- Already characterizing Hanford at a subsystem level.
- Other noise features that are troublesome to the transient searches will undoubtedly surface as the ifos evolve and commissioners dig into the noise floor.
- Detchar has very good handle on potential DQ features that would most affect the search backgrounds.