



---

# Cutting Edge Computing for the Extraction of Astrophysical Parameters from GW Observations

Halston Lim

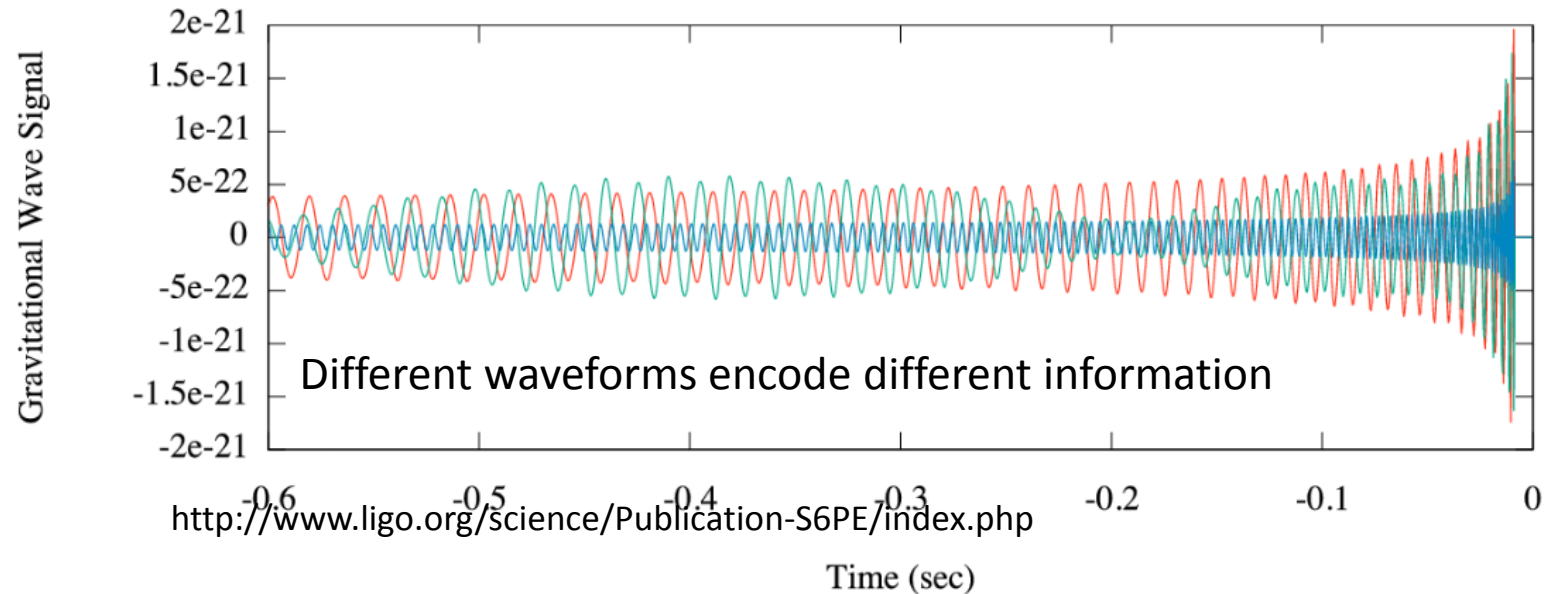
Mentors: Kent Blackburn, Vivien Raymond, Rory Smith

LIGO SURF 2014



# LIGO Parameter Estimation

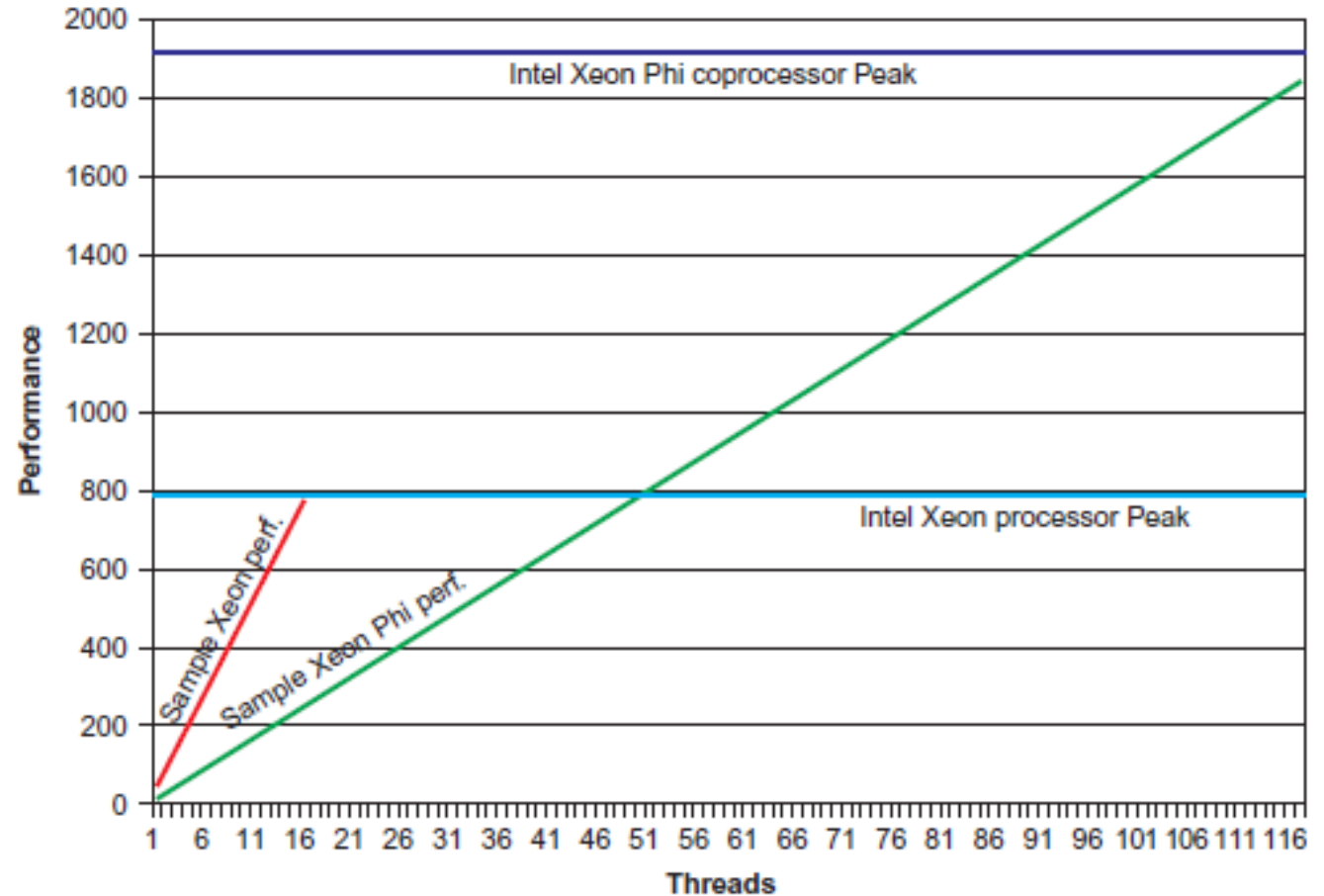
- Process of extracting source physics from GW signals
  - Goal is to estimate the **parameters** of a given waveform model (e.g. mass, orientation, time of coalescence, spin)
  - Bayes' Theorem to calculate the posterior probability density (PDF) – the probability of the parameters, given the observed data
- High computational cost for exhaustive sampling
  - Markov Chain Monte Carlo methods to (more) efficiently calculate PDFs
  - Approximations needed





# Computer Architecture and Optimization

- Faster processors
  - Intel Sandy Bridge vs. Nehalem
- Parallel architecture – many simultaneous calculations
  - Intel MIC architecture
  - Must write tailored code
- Using functions optimized for specific hardware
  - Intel Math Kernel Library (MKL)



J. Jeffers and J. Reinders, *Intel Xeon Phi Coprocessor High-Performance Programming* (Elsevier, Waltham)

LIGO-G1400931-V1

# LIGO Supercomputers

- Ran on the Caltech and Stampede (Texas Advanced Computing Center) clusters
  - High throughput – running many separate jobs (not parallel), CIT
  - High performance – running single job highly parallel, Stampede
  - Different computer usage types give different results
- Computer clusters are shared between many users – jobs are handled through job scheduler (HT Condor, SLURM)
- Stampede, ranked 7<sup>th</sup> supercomputer according to TOP500 project as of June 2014





# LIGO Research Goals

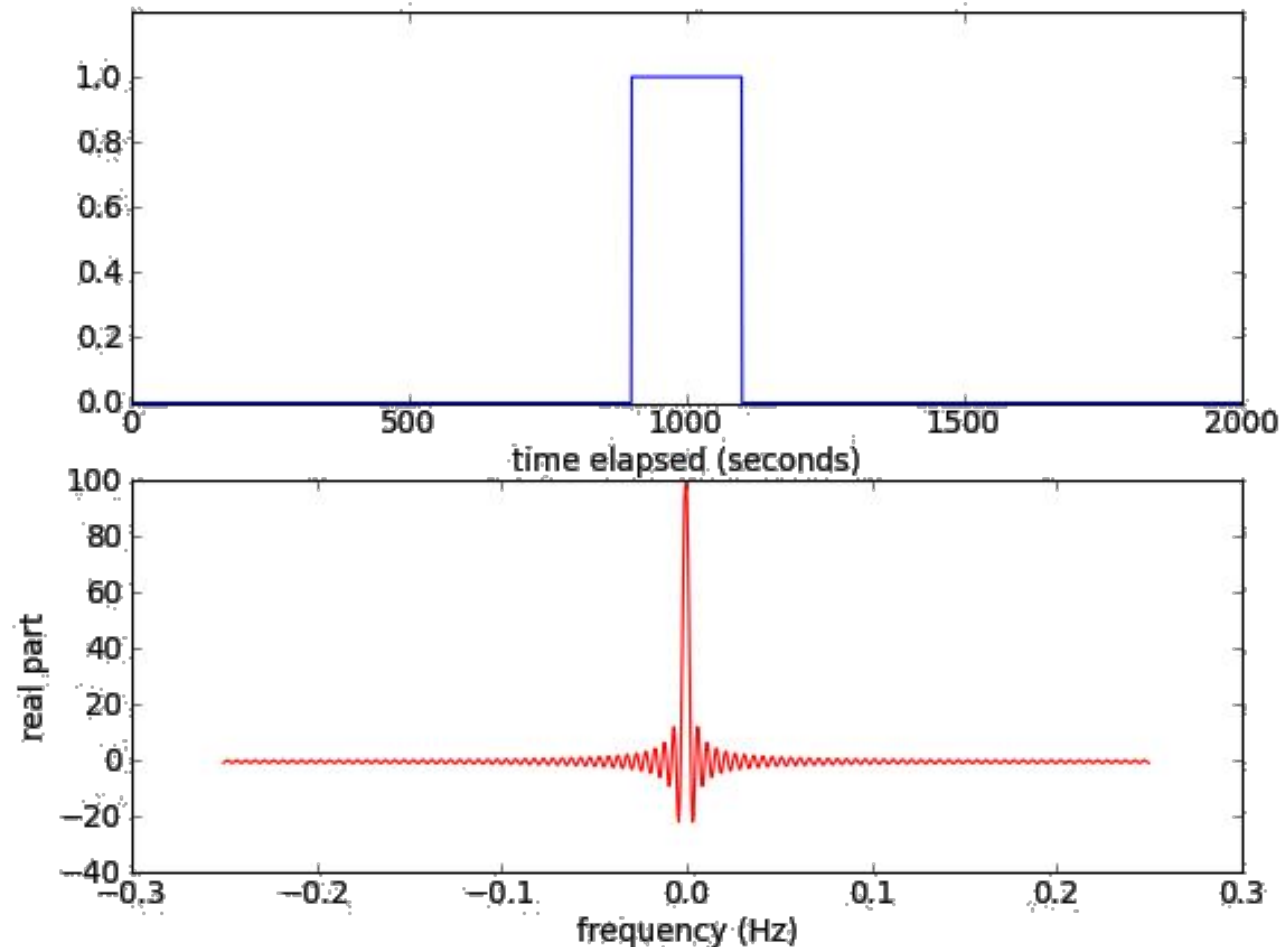
---

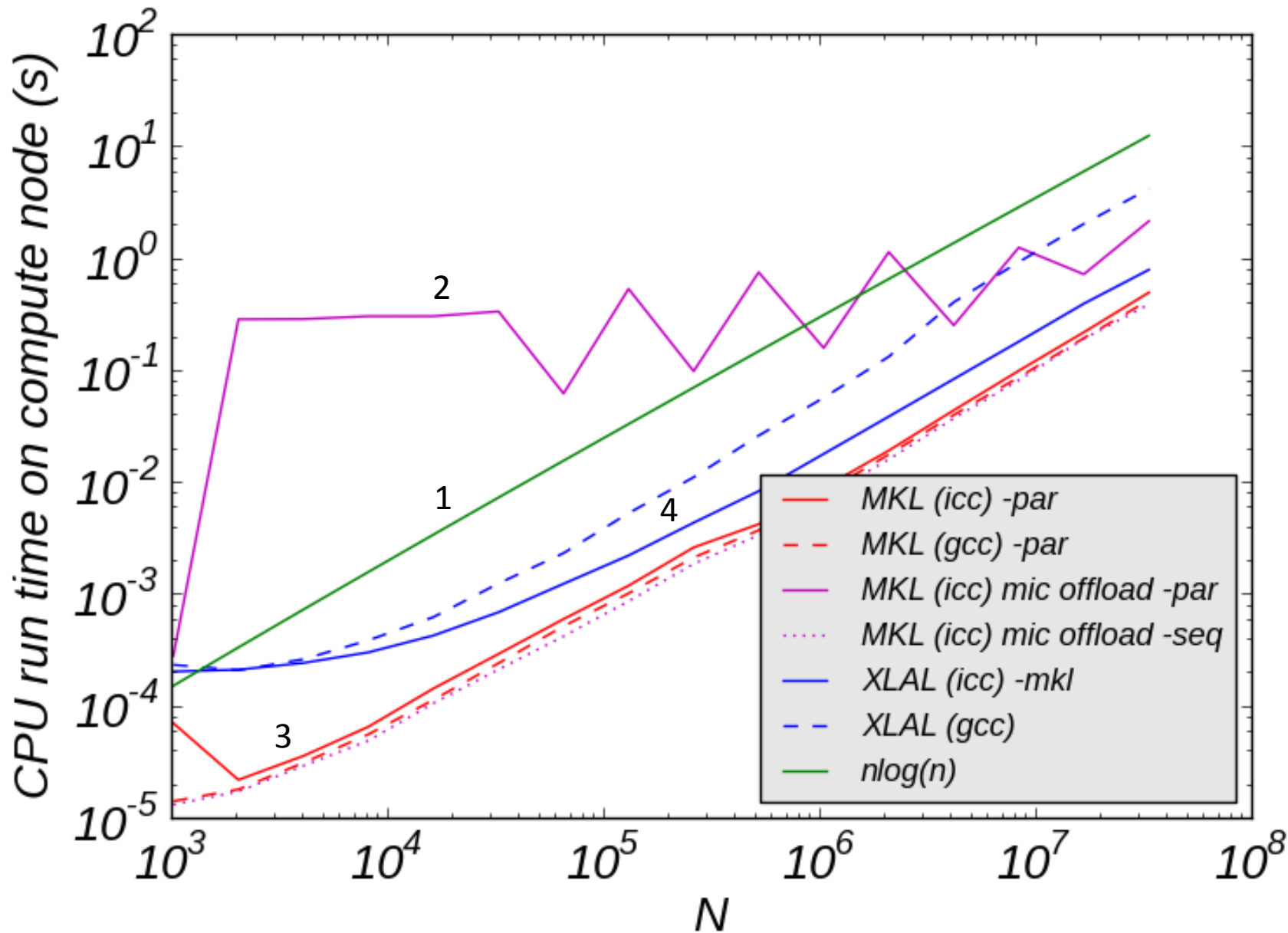
1. Create a benchmark application to test the performance of functions used in GW parameter estimation
2. Optimize the LALSuite parameter estimation code on the Stampede cluster
3. Investigate how the Reduced Order Quadrature method speeds up parameter estimation

User	Usage (SUs)
Stuart Anderson (stuart)	1491208.7330
Joseph Areeda (areeda)	0.0000
Juan Barayoga (barayoga)	0.0000
James Blackburn (kent)	74764.7820
Duncan Brown (dabrown)	172.9350
Santiago Caride (sano)	0.0000
Marco Cavaglia (cavaglia)	0.0000
Nelson Christensen (nchrise)	0.0000
Michael Coughlin (coughlin)	0.0000
Peter Couvares (pfcouvar)	2177.6240
Edward Daw (edaw)	0.0000
Benjamin Farr (bfarr)	2052.3290
Evan Goetz (egoetz)	0.0000
Chad Hanna (channa)	0.0000
Ian Harry (spxiwh)	0.0000
Paul Hopkins (hopkinsp)	0.0000
Ra Inta (inta)	0.0000
Sergey Klimenko (klimenk0)	0.0000
Dan Kozak (dkozak)	0.0000
Tjonnie Guang Feng Li (tgfli)	0.0000
Halston Lim (hblim)	26078.6330

# LIGO Benchmark Results

- Create benchmark program to find CPU for a sinusoidal forward FFT
- Benchmarked on the Stampede cluster comparing
  - Intel MKL vs. XLAL (LIGO Algorithm Library) functions
  - Intel compiler vs. generic GNU compiler
  - Threaded vs. sequential MKL
  - MIC vs. no MICs





1. Theoretical FFT algorithm should scale as  $n\log(n)$

2. MIC processors most scalable, but too much overhead

3. MKL outperformed XLAL by order of magnitude

4. Inefficiencies in the MKL wrapper



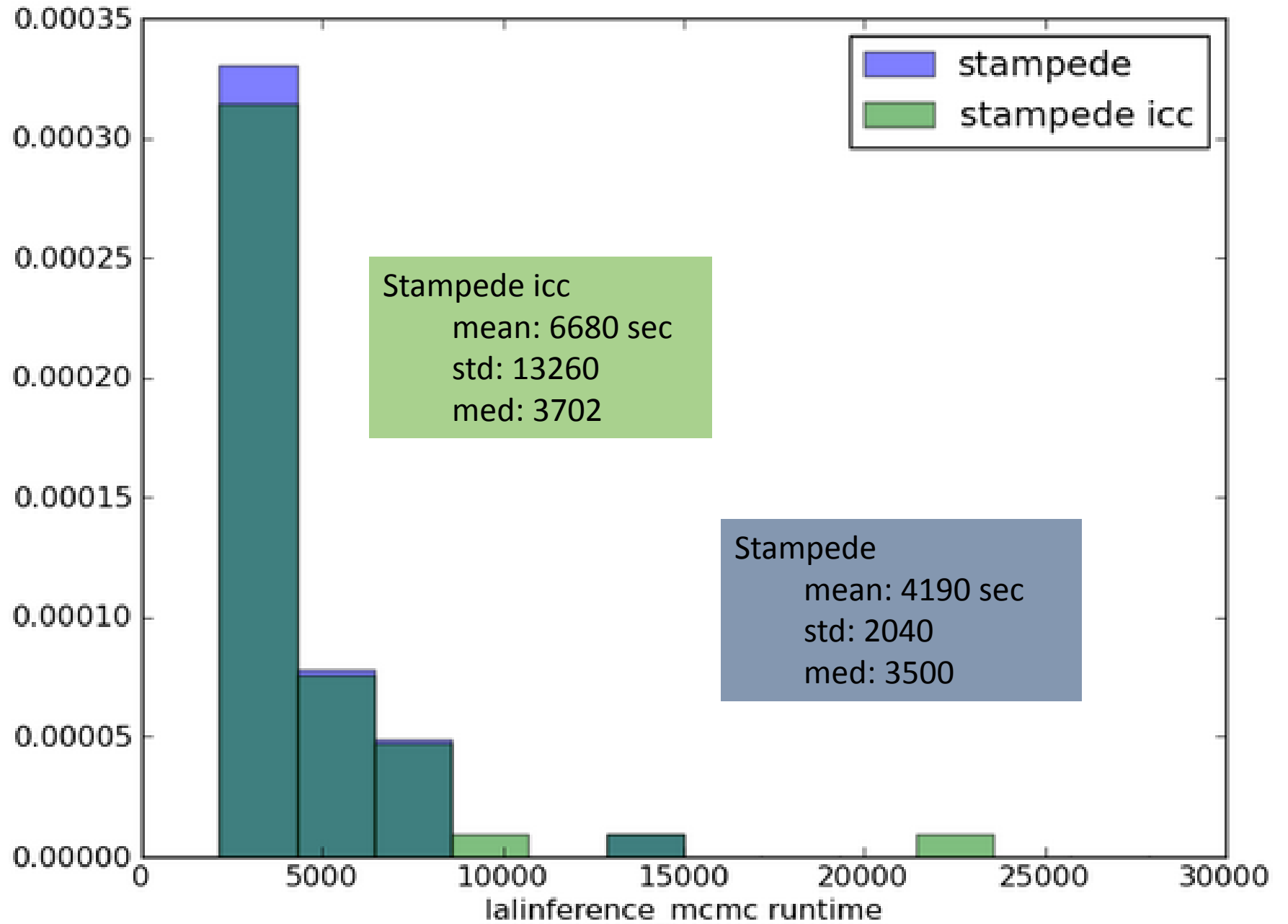
# Comparing Performance with the LALSuite Parameter Estimation Code

---

- Compared wall-clock time of parameter estimation code with GNU and Intel compiler
  - Simulated data in Advanced LIGO-Virgo detectors
  - Injected 50 binary neutron star signals with fixed component masses of  $1.4 \odot$  and uniform in volume and orientation
  - SNR from 5 to 20
  - Taylor F2 3.5 PN (no spin model)
- Parameter estimation code is more than just FFT, so runtime improvements may not be as drastic
- Runtime differences between ICC and GCC compiler



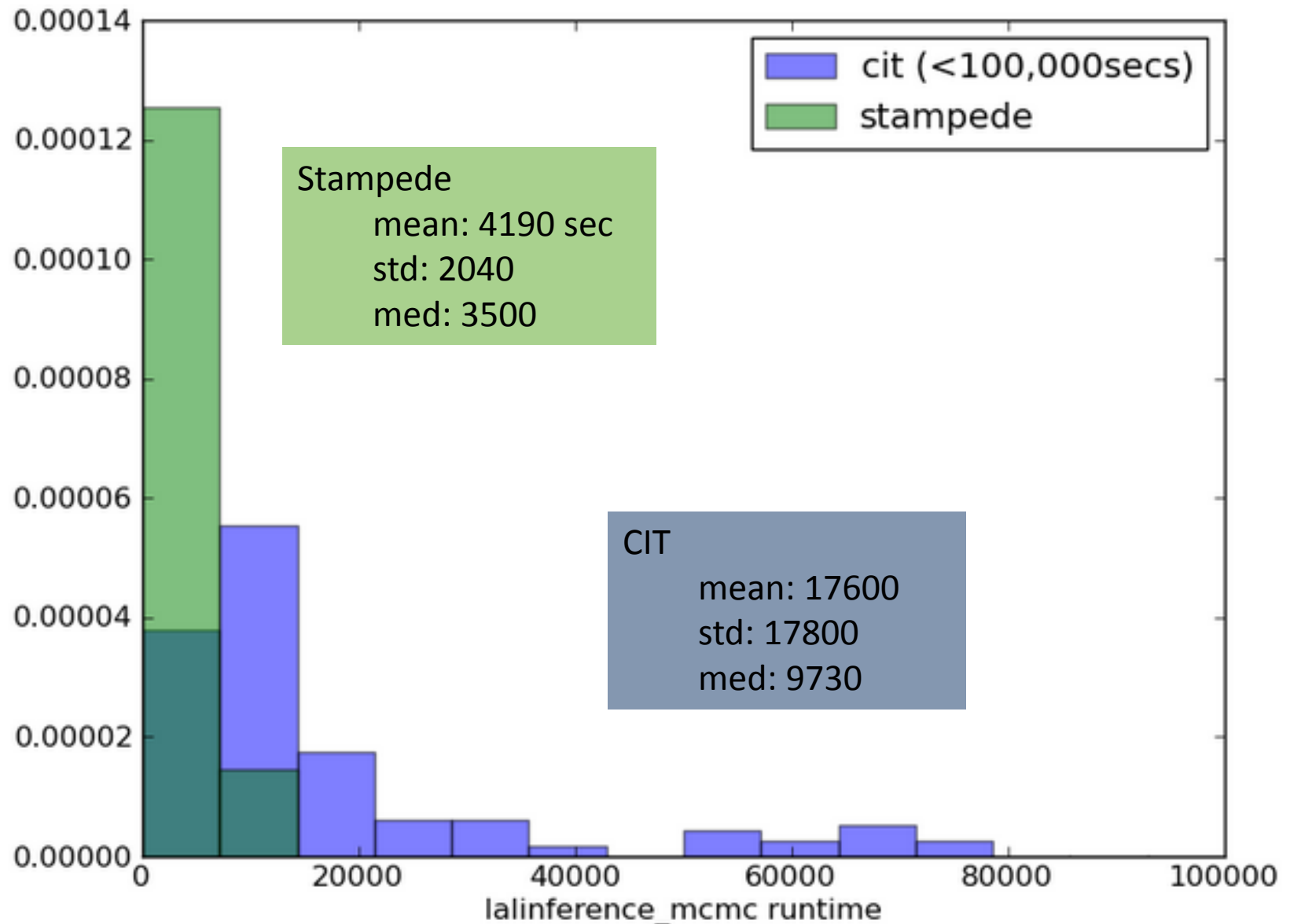
- Compiler had little impact on speed up
  - In rough agreement with the FFT results





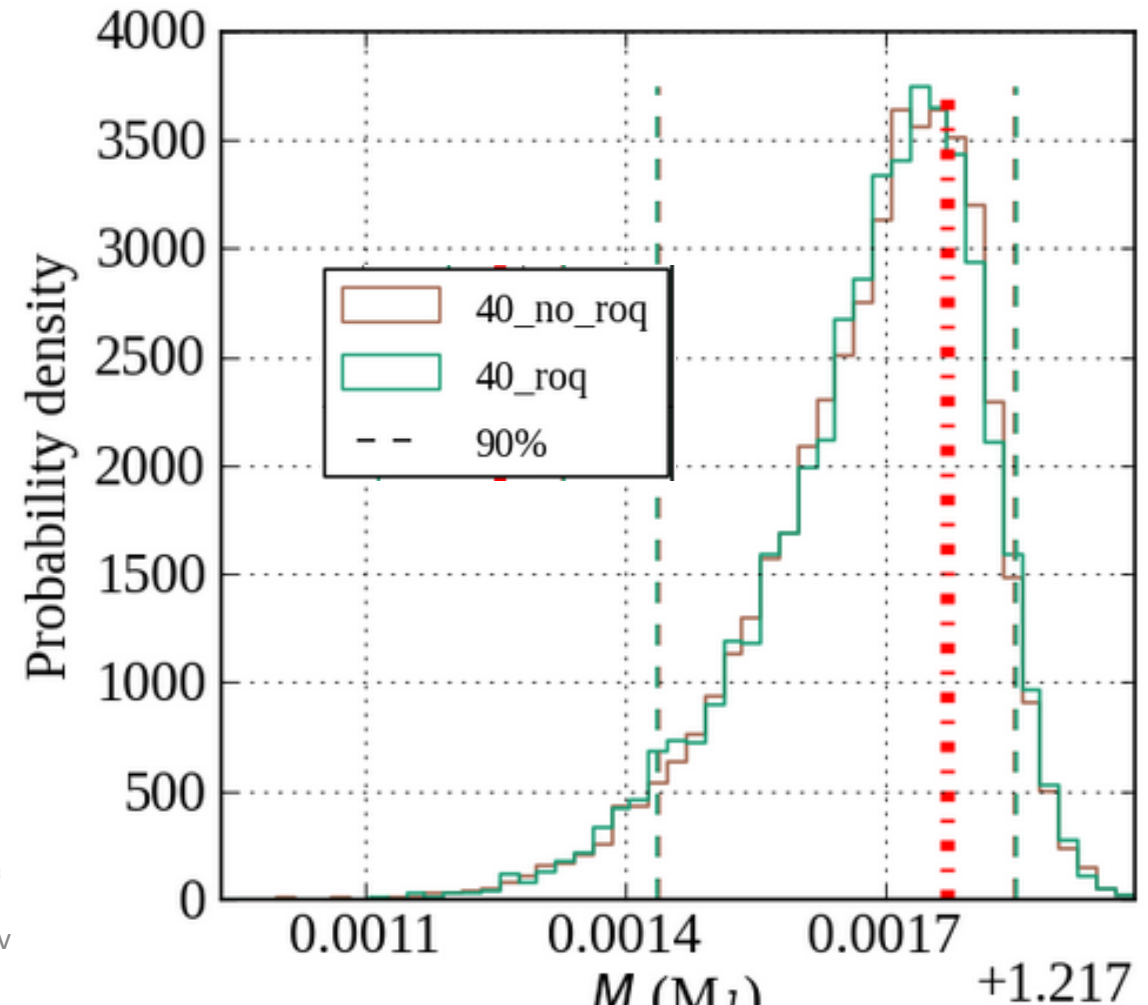
# CIT vs. Stampede performance

- GCC compiled
- Excluded 11 outliers in CIT set that had time  $> 10^5$  secs
- Stampede runs were both more consistent and efficient



# LIGO Incorporating Reduced Order Modeling (ROQ)

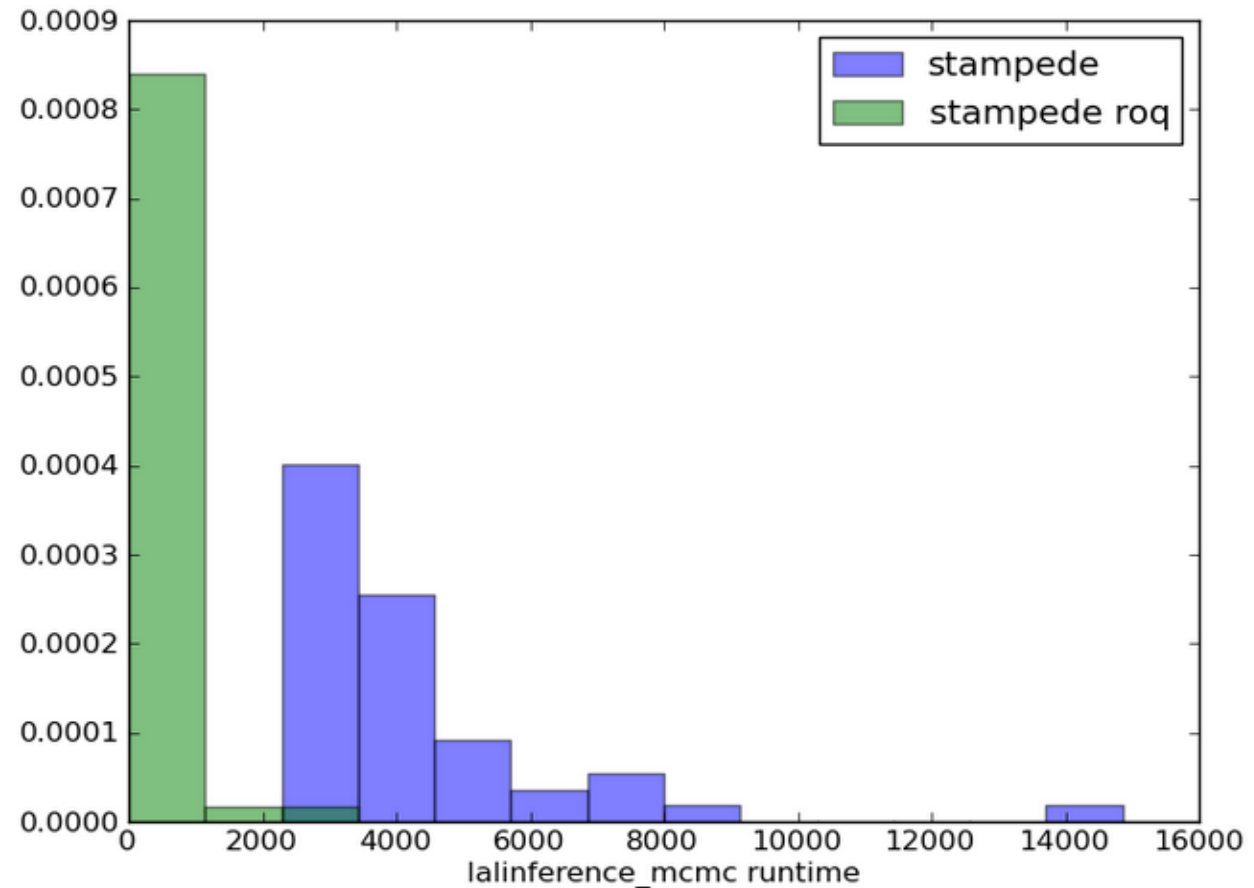
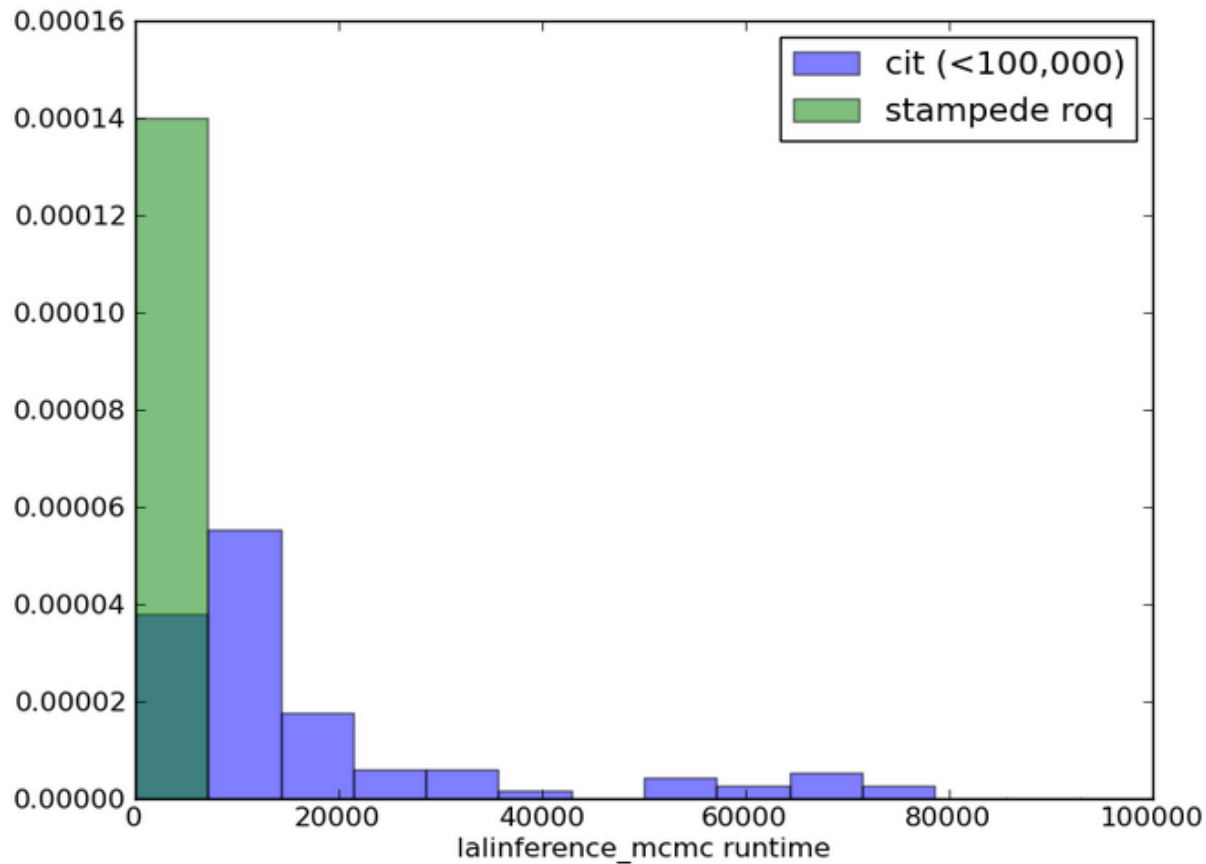
- Reduced order quadrature scheme can reduce parameter estimation times by a factor of 30\*
- Faster runtime = more physics
- Comparison for a single injection
  - SNR 17.043
  - **Non-ROQ** 90 minutes to generate
  - **ROQ** 7 minutes to generate



\* P. Canizares, S. E. Field, J. Gair, V. Raymond, R. Smith and M. Tiglio, arXiv:1404.6284 [gr-qc].



# ROQ relative performance at $f_{\text{low}} = 40$ Hz




	Stampede	Stampede ROQ	CIT
Mean CPU time (s)	4190	<b>330</b>	17600
Standard Deviation	2040	<b>449</b>	17800
Median	3500	<b>218</b>	9730



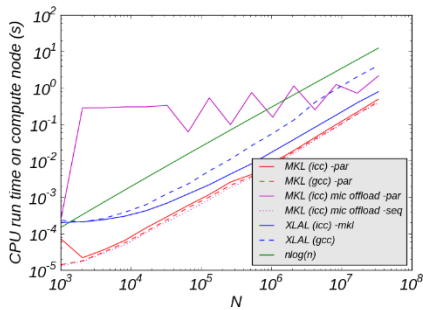
# Future Work: Lowering the detection frequency

- Lowering detection threshold frequency means more GW signal can be detected
  - Pros: Detecting more signal can give more information about source -> better parameter estimation
  - Cons: More computational intensive to analyze
- Can we justify lowering  $f_{\text{low}}$  through a parameter estimation improvement?
- How can ROQ enable us to answer this question:

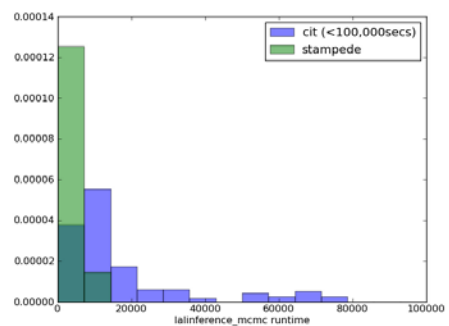
Method Name	40 Hz Analysis Time	20 Hz Analysis Time
Non-ROQ	90 minutes	~1 week
ROQ	7 minutes	60 minutes



# Conclusion



Various algorithms used in GW analysis like the FFT can be optimized on hardware



Large speed ups possible for parameter estimation code

- Running on HPC nodes such as Stampede (3-5 X faster)

ROQ drastically speeds up parameter estimation

- May make analyzing signals from 20 Hz feasible

Method Name	40 Hz Analysis Time	20 Hz Analysis Time
Non-ROQ	90 minutes	~1 week
ROQ	7 minutes	60 minutes



# LIGO Next steps

---

- Confirm results at 20 Hz with ROQ
- Work to further optimize LALSuite
  - Incorporate MKL as much as possible, as order of magnitude improvement in FFTs not seen yet in `lalinference_mcmc`
  - Make applications scalable for the MIC coprocessors
- Test speed ups with spin models at 30 Hz
- Profile the parameter estimation application to reveal inefficient hotspots in the code

# Acknowledgements

- Kent Blackburn, Vivien Raymond and Rory Smith
- Alan Weinstein
- NSF Foundation