

RECOVERING HARDWARE INJECTIONS IN LIGO S5 DATA

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Outline

- LOSC Open Science Data Release
- Hardware Injections of compact binary coalescence signals
 - What are these?
 - What do we expect to find in the data?
- How to generate a
- Template matching and signal recovery
- Recovery of Hanford 2 hardware injections
 - Was the match successful?
 - Do we see what we expect?
- Summary of Final Results for All Detectors

LIGO Open Data Release

- LIGO archival data will be released to public as open source data
 - S5 science run 2005-2007
 - H1 and H2 at LHO, L1 at LLO

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 - Provides access to data
- In preparation for the release:
 - Software, cookbooks, wikis, tutorials, and teaching materials
 - Bring 8 year old book-keeping up to date
 - Recover and document hardware injection signals

Hardware Injections

- Inject Compact Binary Coalescence signal into data
 - Move ETMs (mirrors) using magnetic actuators
 - Important for instrument calibration and evaluating the efficiency of searches for signals

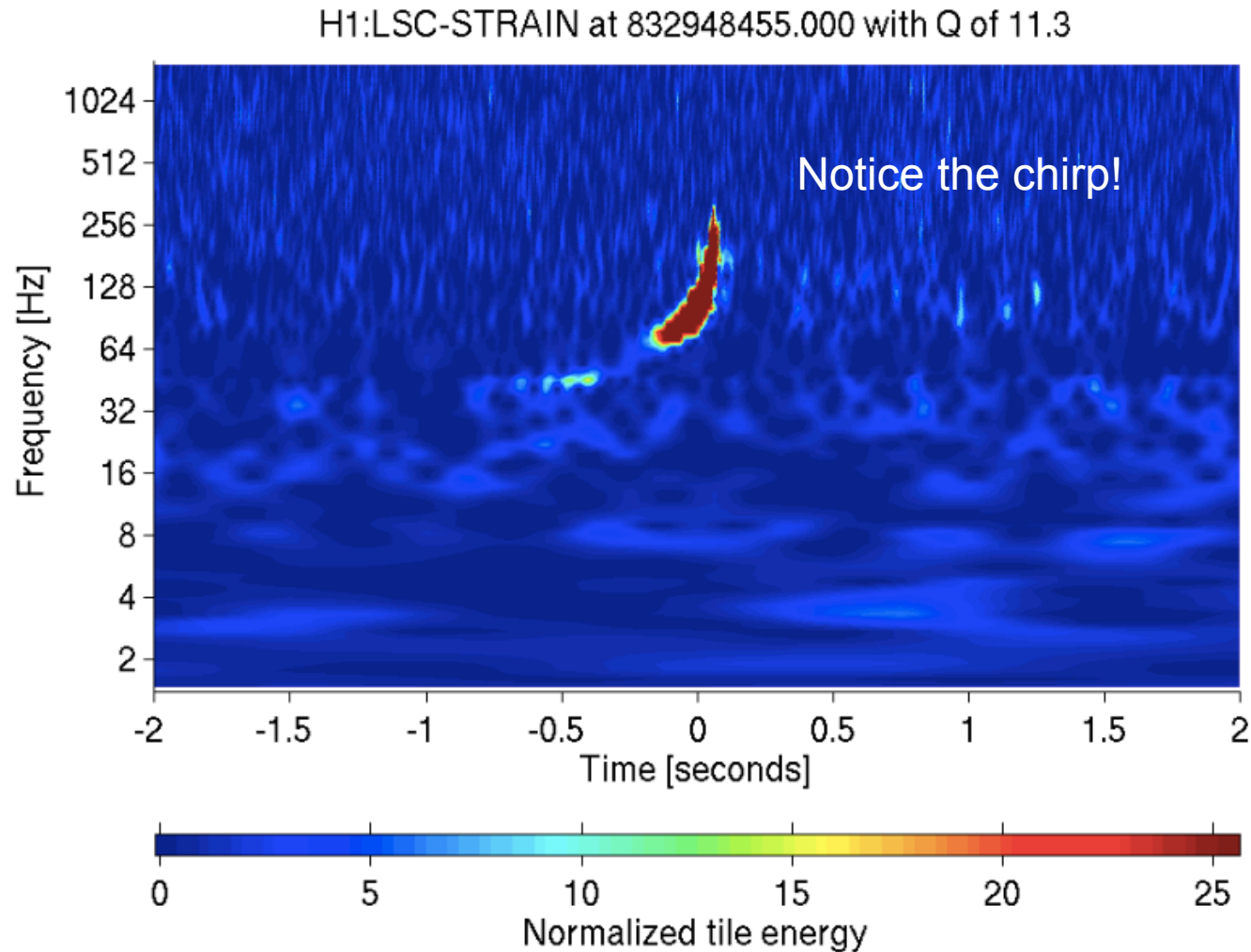
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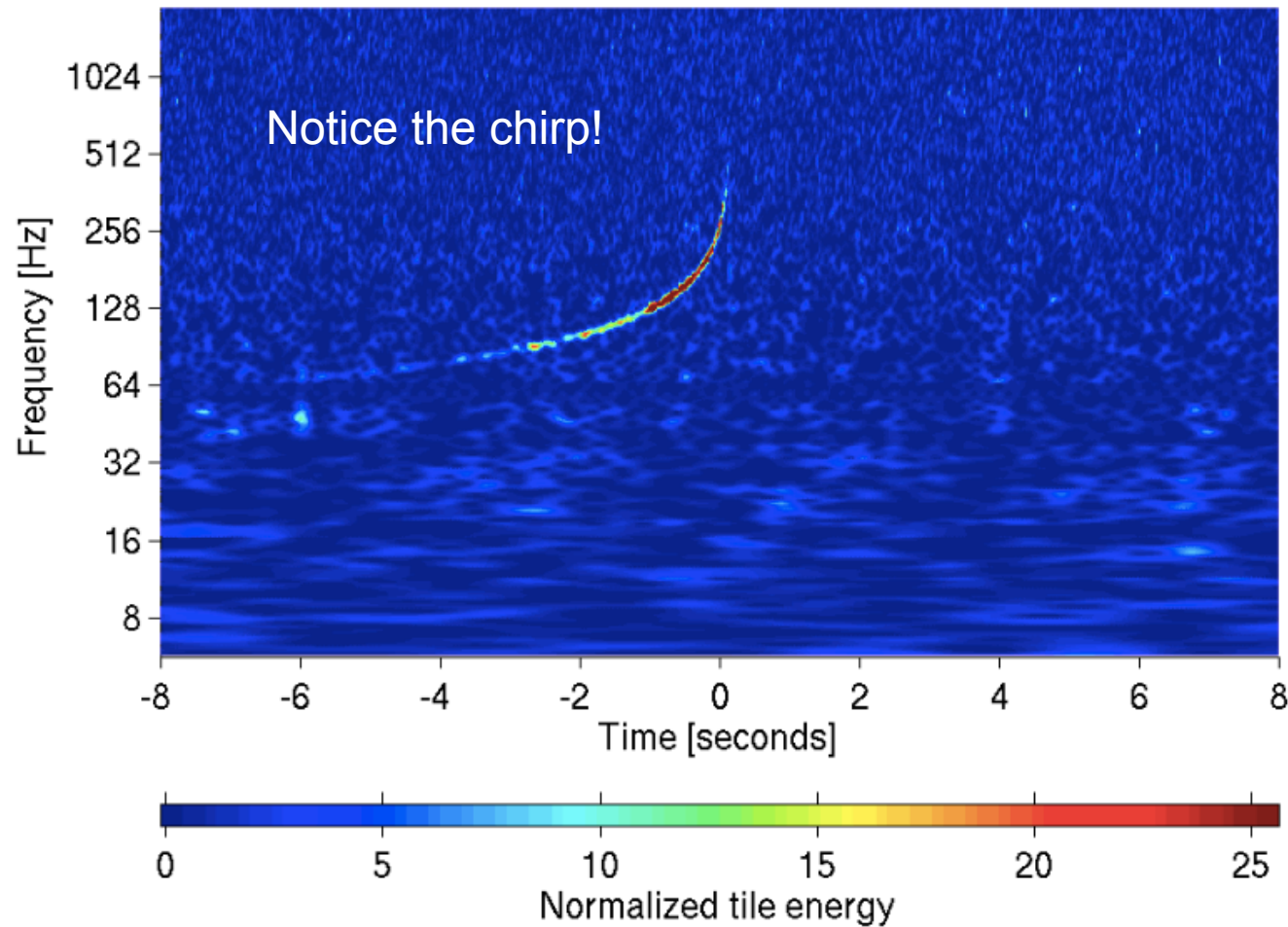
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- } Unsuccessful

10 – 10 Solar Mass Hardware Injection



1.4 – 1.4 Solar Mass Hardware Injection

H1:LSC-STRAIN at 817645695.000 with Q of 45.3

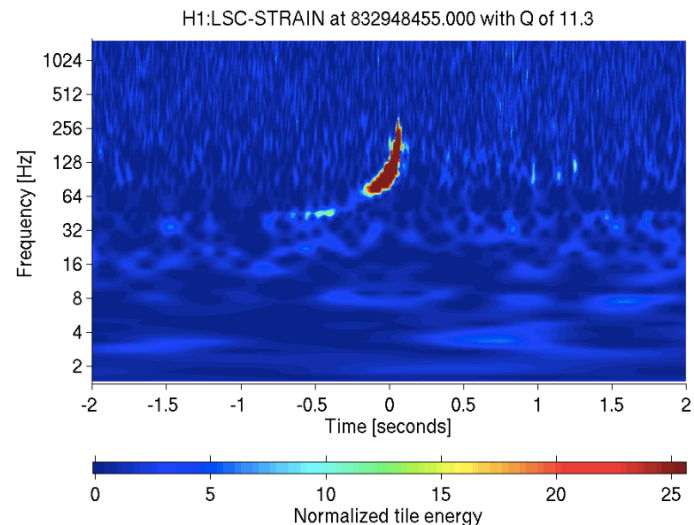


Generate Template

- Create Compact Binary Coalescence templates
 - 1.4 – 1.4 Solar mass binary
 - 3 – 3 Solar mass binary
 - 10 – 10 Solar mass binary
 - 1.4 – 10 Solar mass binary

$$\tilde{h}(f) = \left(\frac{1 \text{ Mpc}}{D_{\text{eff}}} \right) \mathcal{A}_{1 \text{ Mpc}}(M, \mu) f^{-7/6} e^{-i\Psi(f; M, \mu)}$$

- $\tilde{h}(f)$ – Strain/Hz
- \mathcal{A} – Mass dependent amplitude
- f – frequency
- $i\Psi(f; M, \mu)$ – Phase of source



Determining the Amplitude of the Template

$$\begin{aligned} \mathcal{A}_{1 \text{ Mpc}}(M, \mu) &= - \left(\frac{5\pi}{24} \right)^{1/2} \left(\frac{GM}{c^3} \right) \left(\frac{GM}{c^2 D_{\text{eff}}} \right) \left(\frac{GM}{c^3} \pi f \right)^{-7/6} \\ &= - \left(\frac{5}{24\pi} \right)^{1/2} \left(\frac{GM_{\odot}/c^2}{1 \text{ Mpc}} \right) \left(\frac{\pi GM_{\odot}}{c^3} \right)^{-1/6} \left(\frac{\mathcal{M}}{M_{\odot}} \right)^{-5/6} \end{aligned}$$

\mathcal{M} – Chirp mass, units of solar mass

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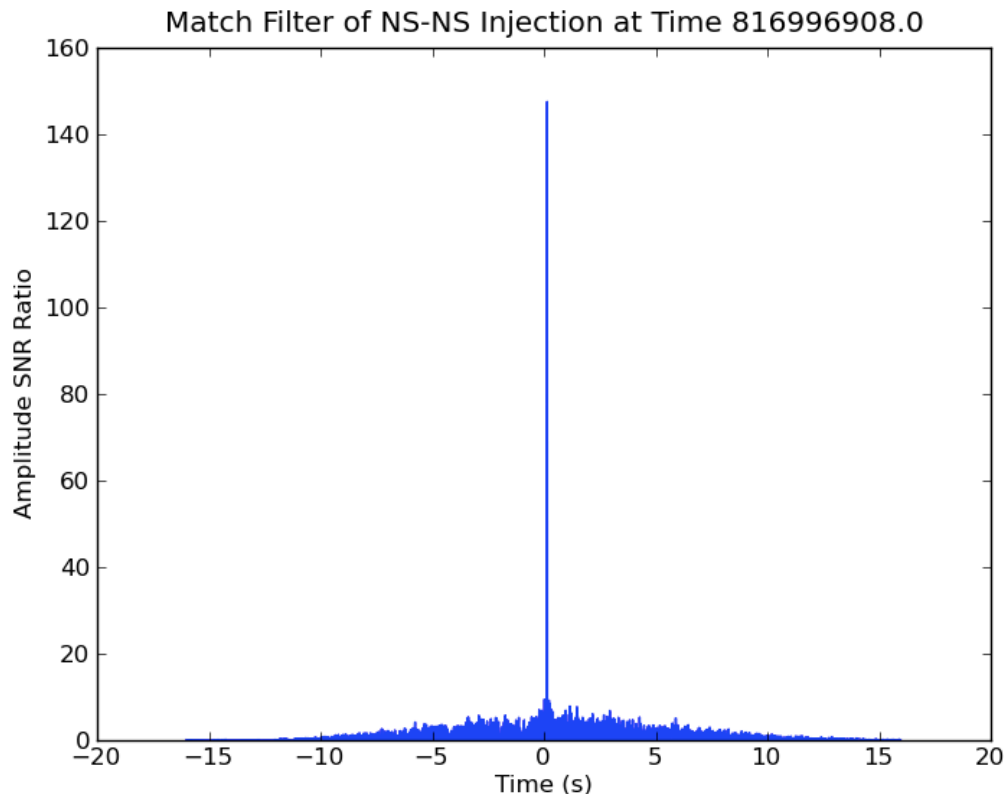
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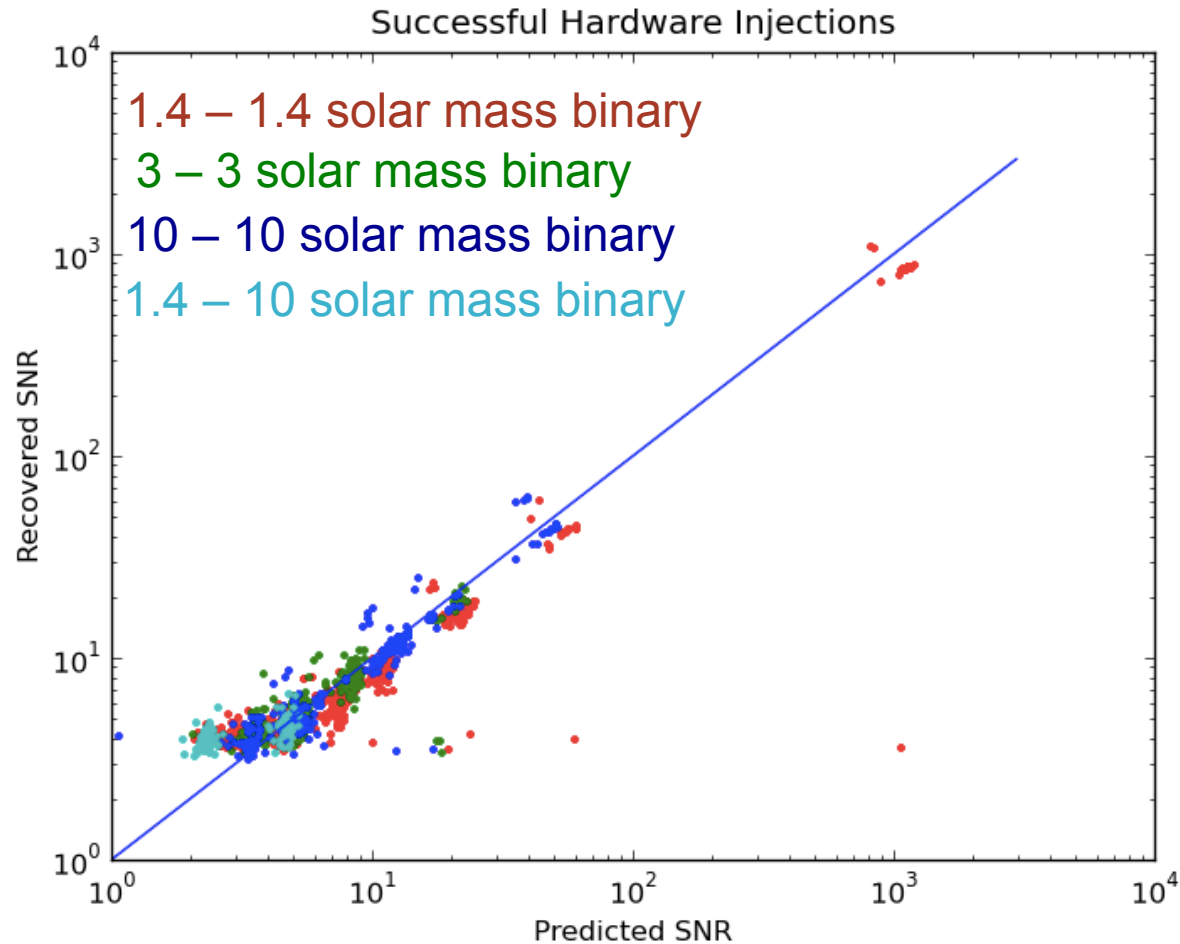
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Finding an Injection

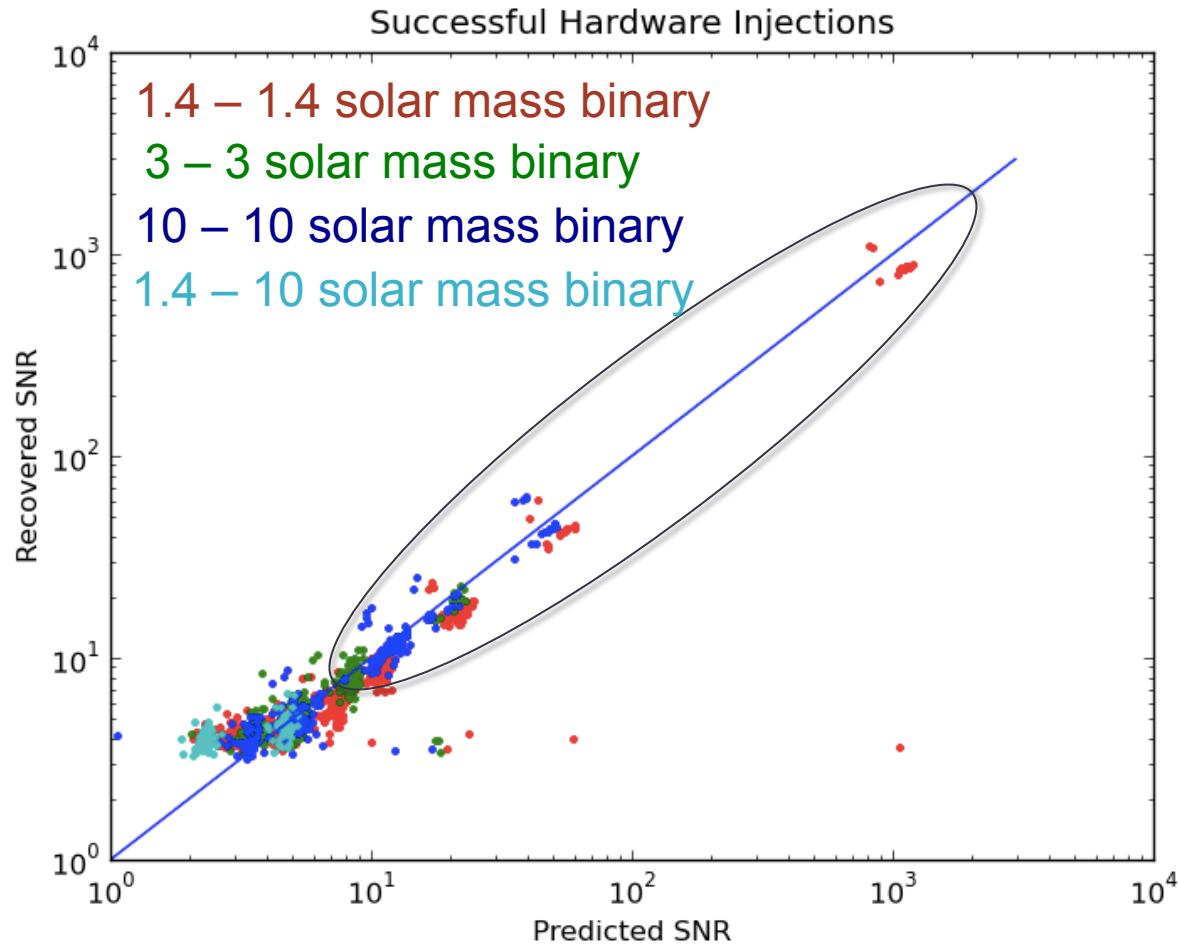
- Cross-correlate template against the data
 - Perform correlation with template starting at different times
- Look for the time shift when the cross-correlation between the template and data is high



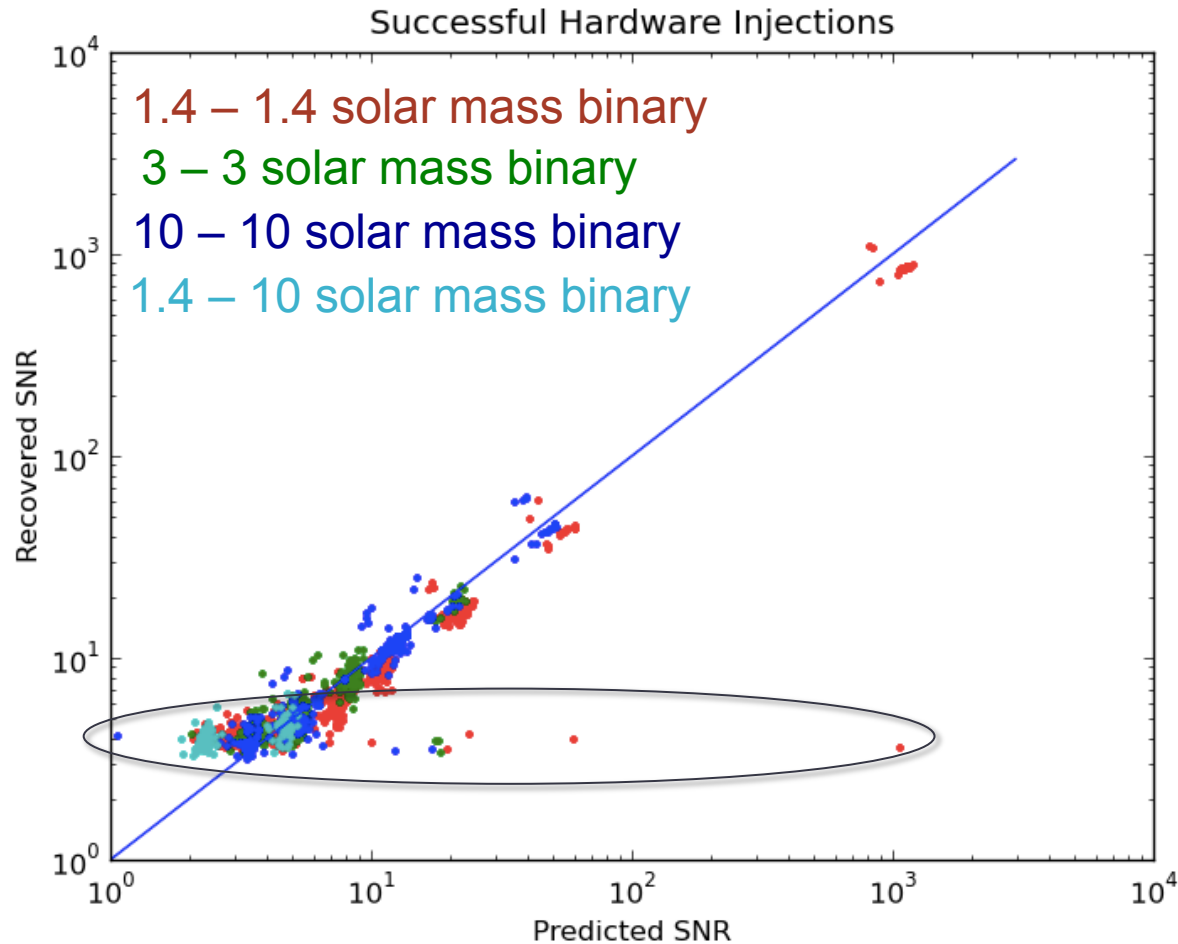
H2 Successful Injections



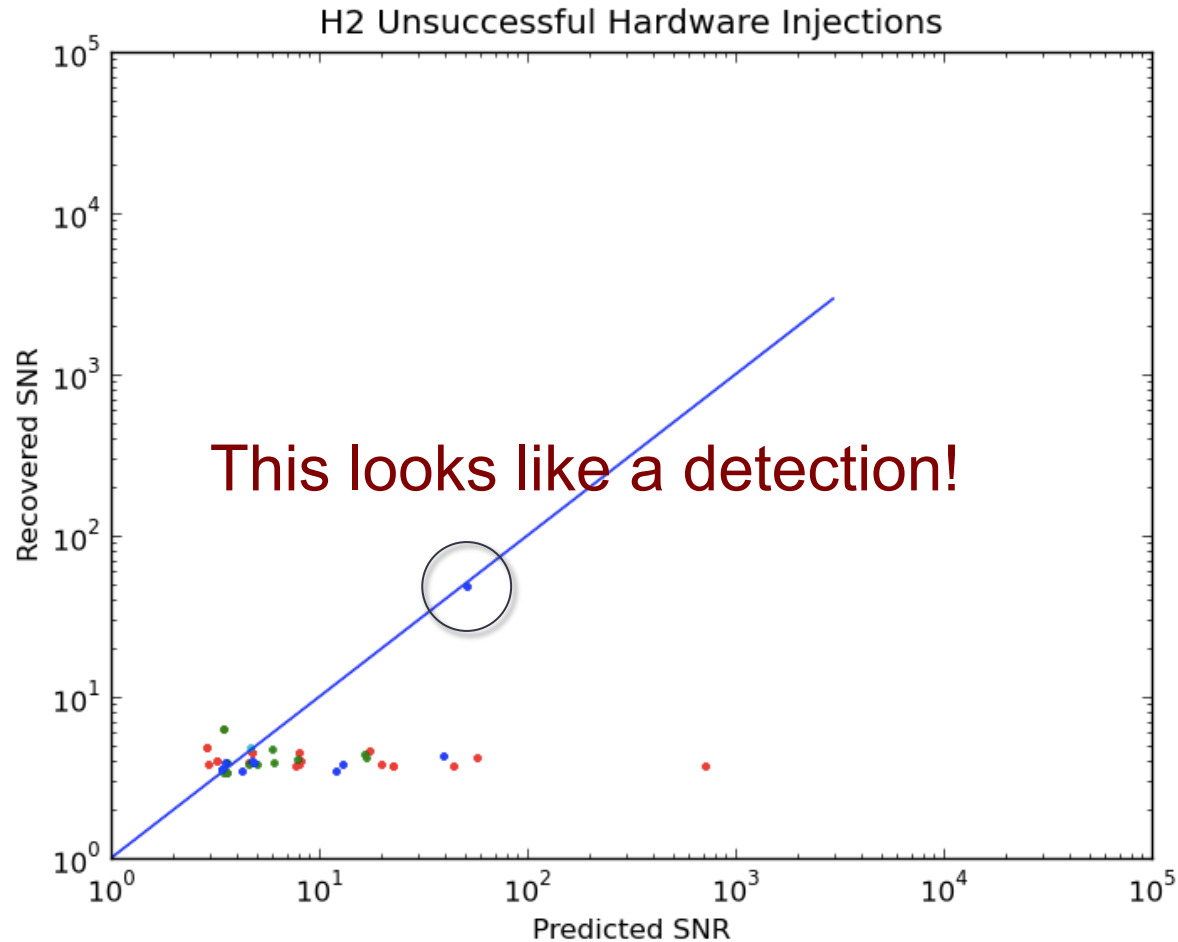
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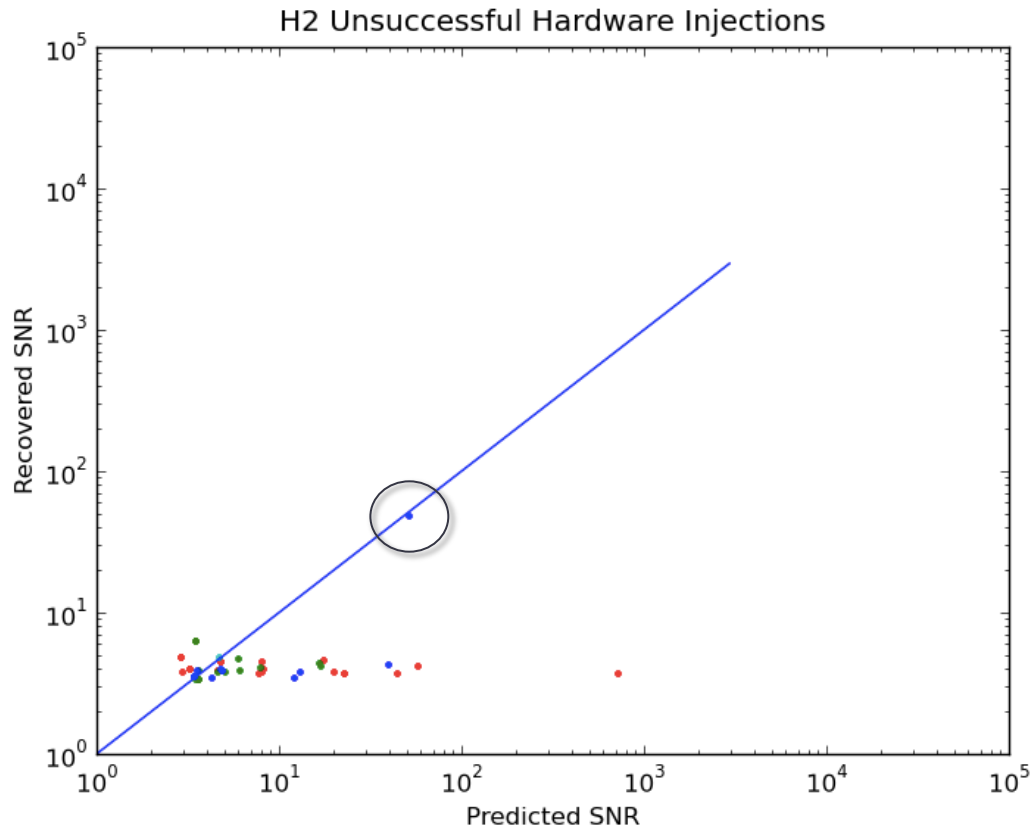


H2 Unsuccessful Injections



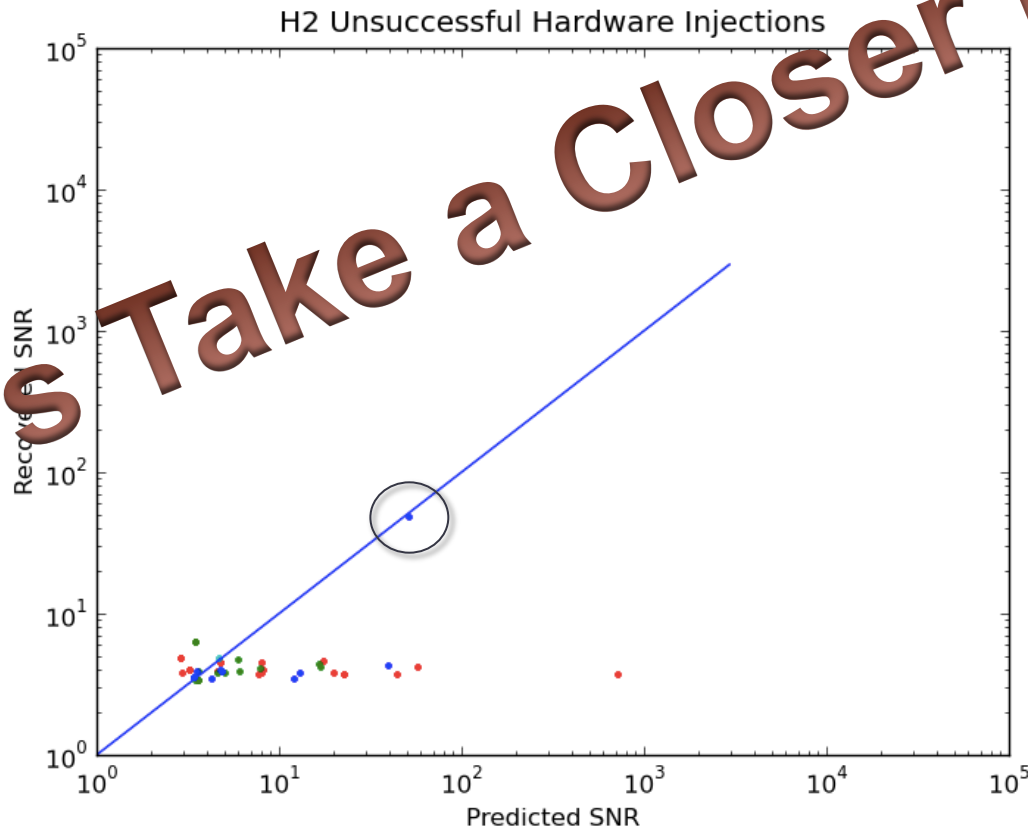
Where Did This Match Come From?

- 10 – 10 solar mass binary located 10 Mpc from Earth
- Marked Injection Compromised



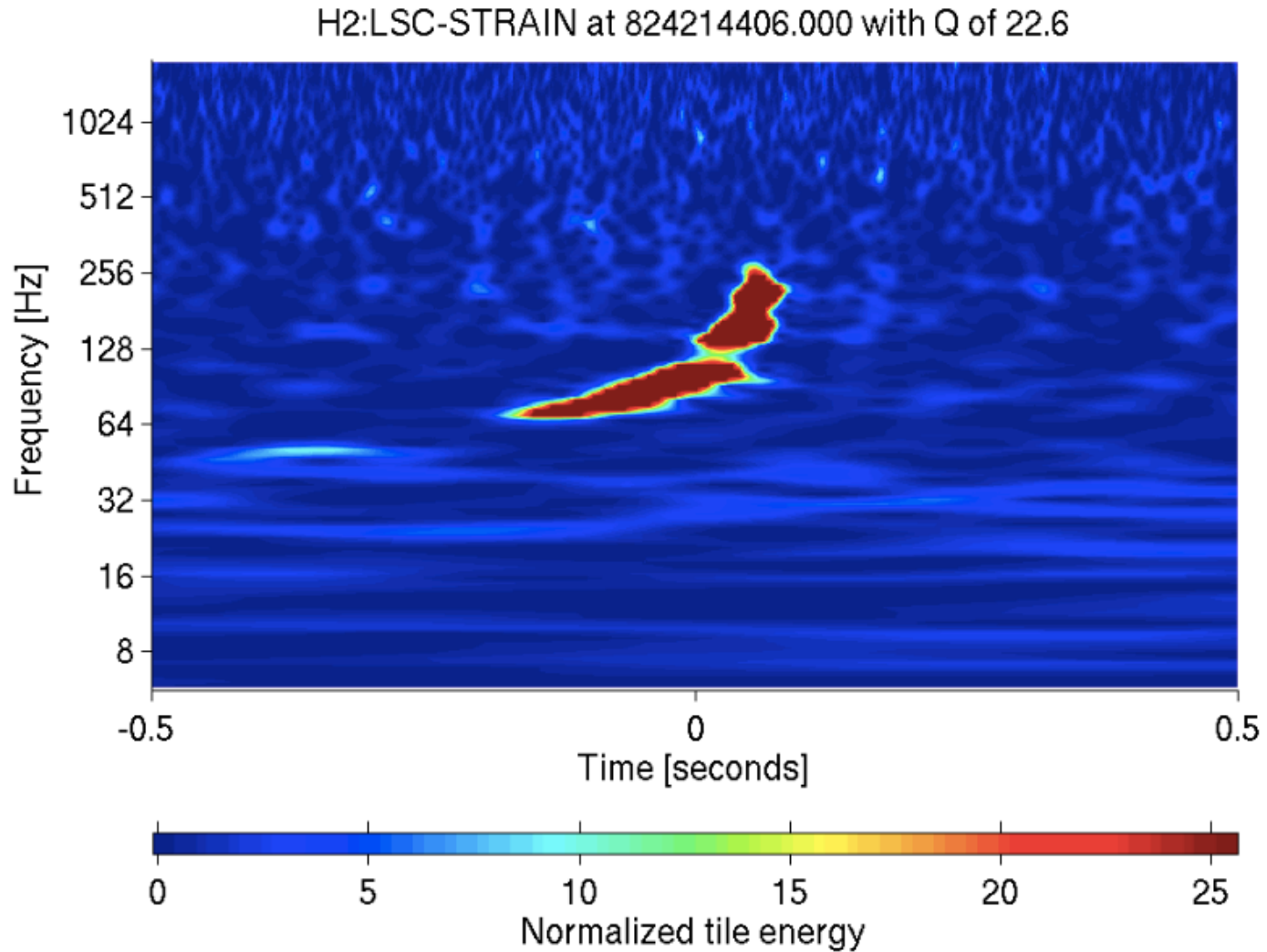
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Let's Take a Closer Look!

Spectrogram of The Injection



Summary Table of Final Results

Detector	H1	H2	L1
Total # Injections	1200	1282	1271
Successful Injections	870	929	770
Successful Injections, Predicted SNR > 8	614	333	545
For Injections with Predicted SNR > 8, Injections with Recovered SNR > 6	608	322	538
Successful Injections, Data Unavailable	21	19	14
Unsuccessful Injections	46	45	51
Unsuccessful Injections with Recovered SNR > 6	1	3	2
Unsuccessful Injections, Data Unavailable	263	289	436

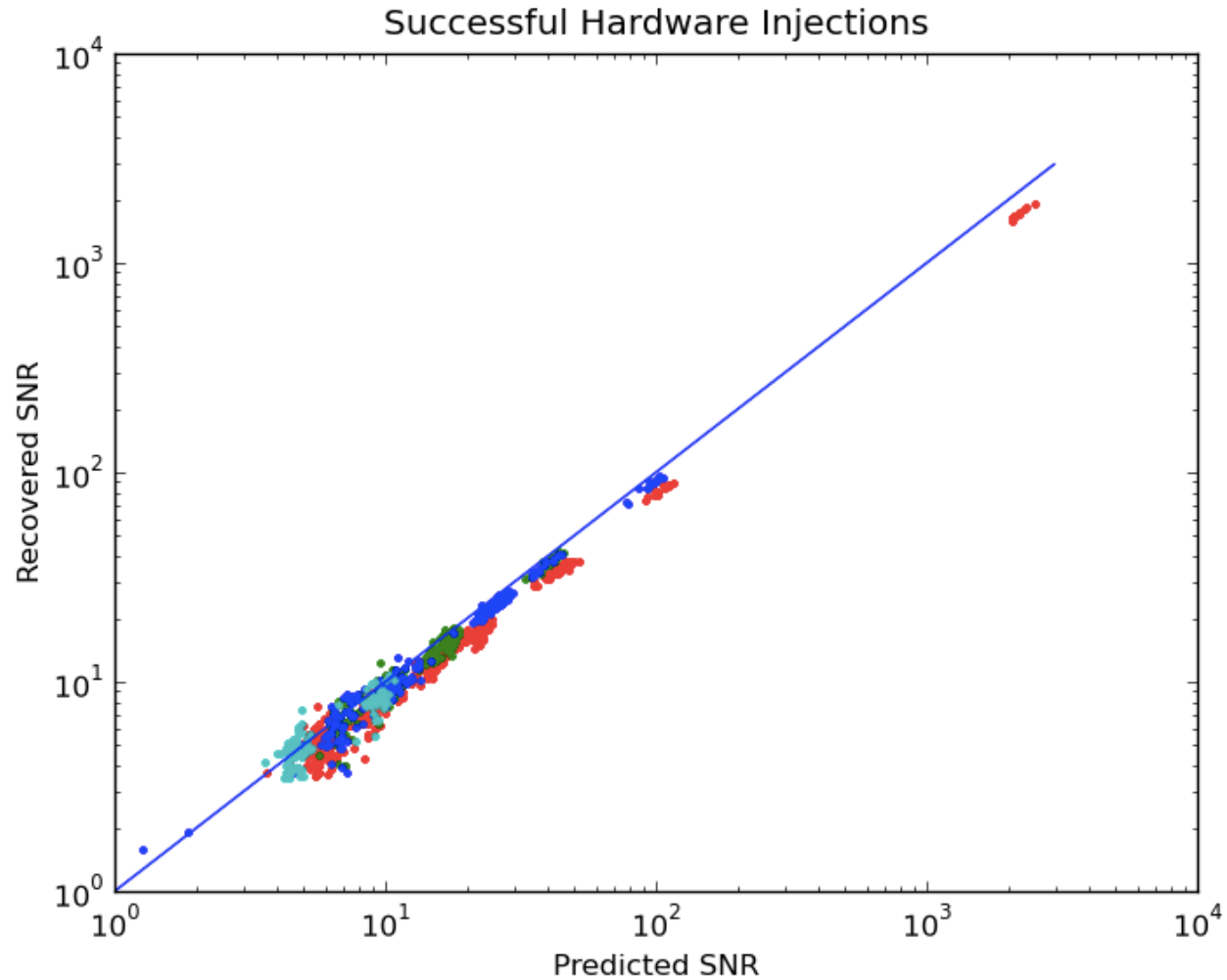
Conclusions

- LOSC will release S5 data to the public
- We search the data for hardware injections
- Our search is successfully identifies whether an injection is successful or unsuccessful
- We find some injections where we do not expect to, referencing past documentation
 - i.e. the detection we discussed
- We will continue to explain these unexpected points and summarize them in the final paper

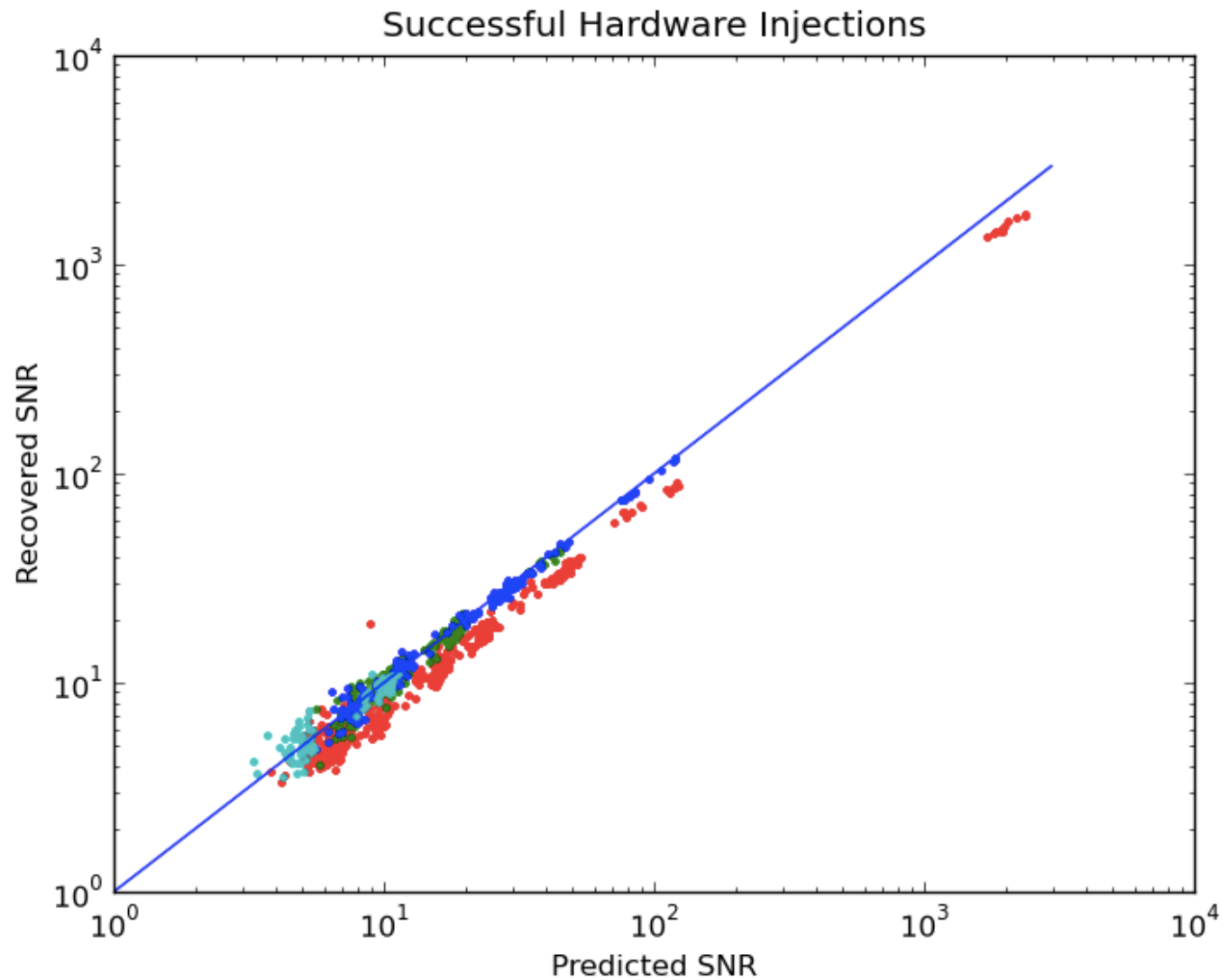
Acknowledgements

- Mentors: Jonah Kanner, Roy Williams, and Michele Vallisneri
- Collaborators: Alan Weinstein and LOSC
- LIGO and National Science Foundation
- Caltech

H1 Successful Injections



L1 Successful Injections



Template Matching: Signal-to-Noise

$$\rho_m(t) = \frac{|z_m(t)|}{\sigma_m}$$

- $\rho_m(t)$ – Amplitude signal to noise ratio of matched filter output
- $z(t)$ – Matched filter output
- σ_m – A measure of the sensitivity of the instrument

$$\rho_{Predicted} = \frac{\sigma_m}{D_{eff}}$$

- $\rho_{Predicted}$ – Predicted signal-to-noise ratio
- D_{eff} – Effective distance from source to Earth

Template Matching: Matched Filter Math

$$z(t) = 4 \int_0^{\infty} \frac{\tilde{s}(f) \tilde{h}_{\text{template}}^*(f)}{S_n(f)} e^{2\pi i f t} df.$$

- $z(t)$ – Matched filter output
- $\tilde{s}(f)$ - Data in frequency domain
- $\tilde{h}_{\text{template}}^*(f)$ - Complex conjugate of template
- $S_n(f)$ - Power Spectral Density of noise