

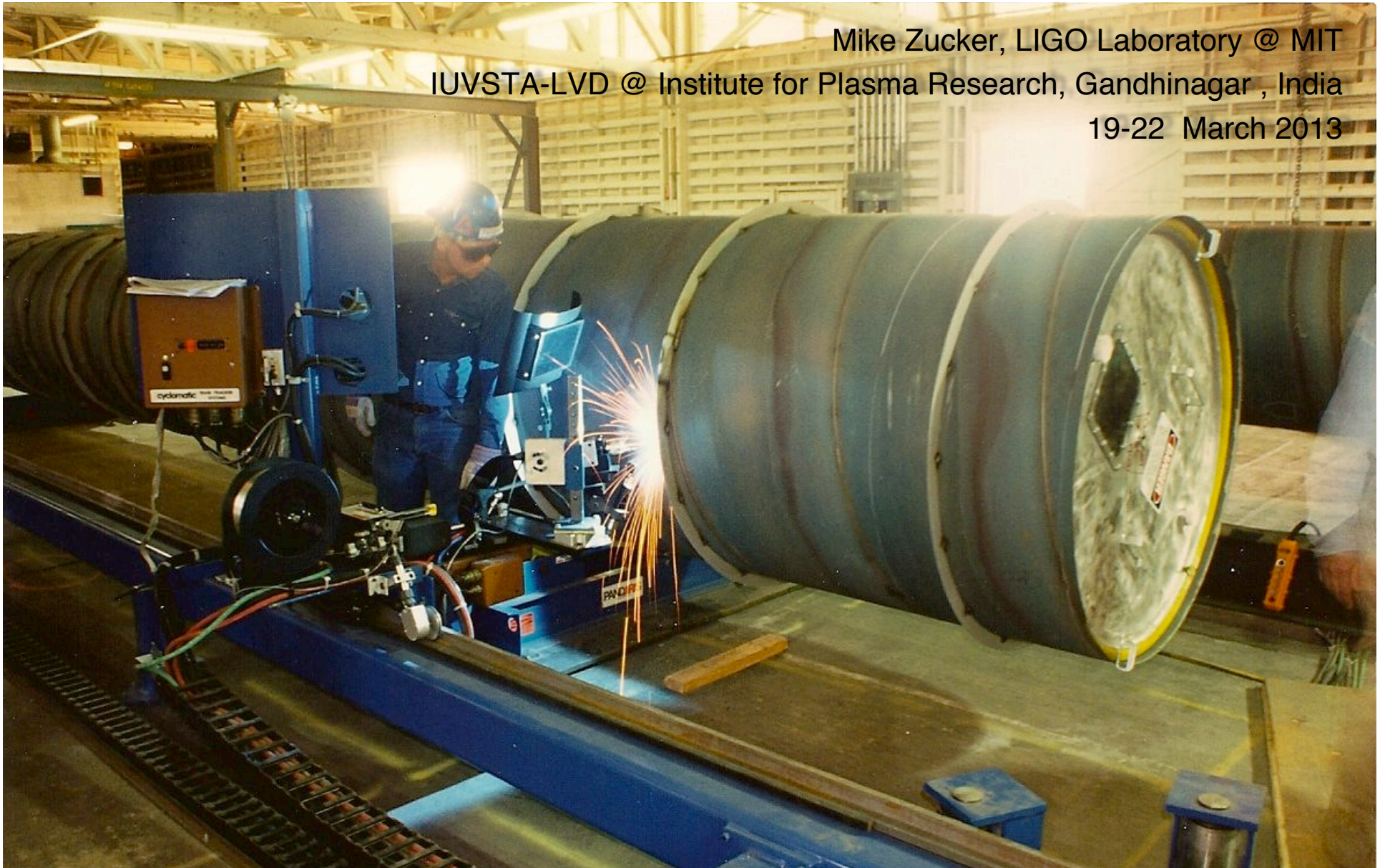


The LIGO Vacuum Equipment and Beam Tubes: Retrospective and Prospective

Mike Zucker, LIGO Laboratory @ MIT

IUVSTA-LVD @ Institute for Plasma Research, Gandhinagar, India

19-22 March 2013





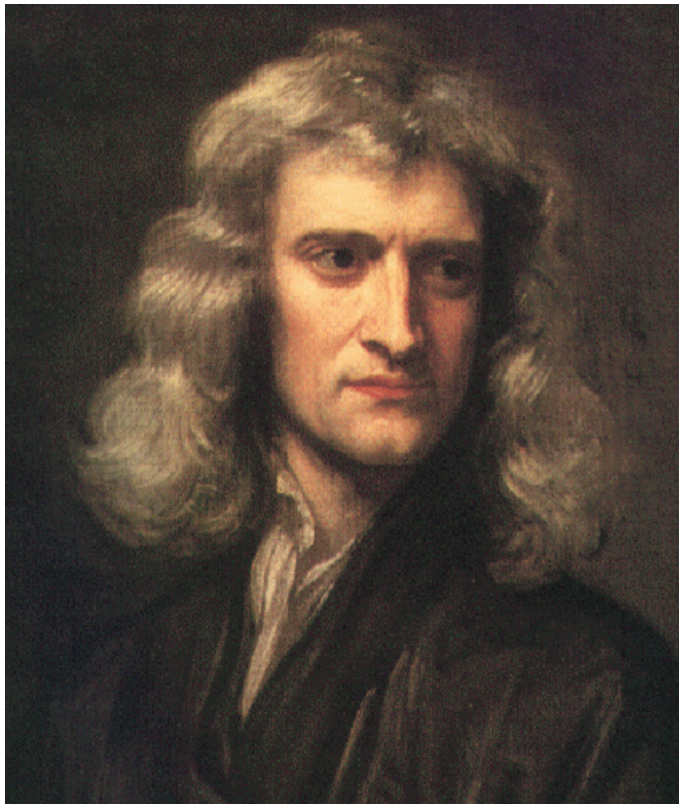
Outline

- **About LIGO and LIGO India**
- **Vacuum Requirements & Constraints**
- **Vacuum Equipment**
- **Beam Tubes**
- **Paths for Improvement**

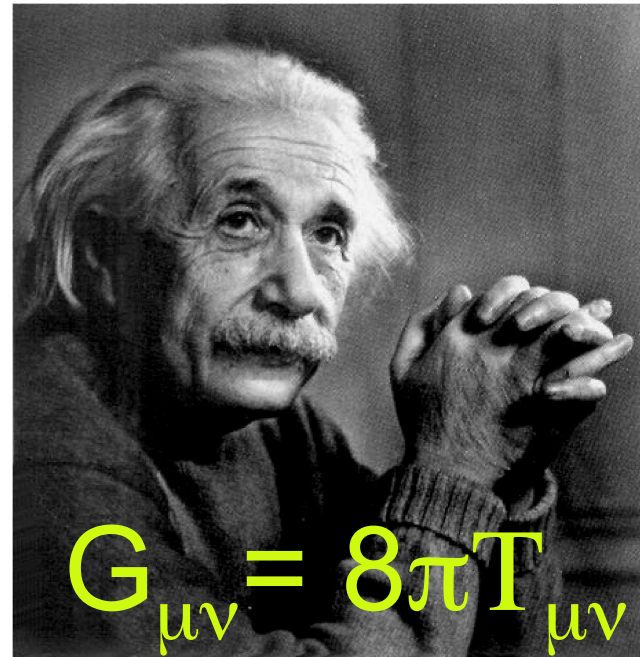


Why must there be gravitational waves?

Newton's puzzle:
"instantaneous action at a distance"



LIGO-G1500176



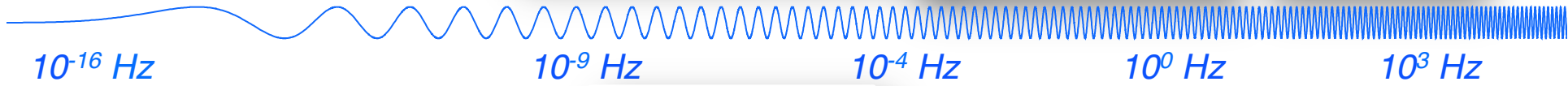
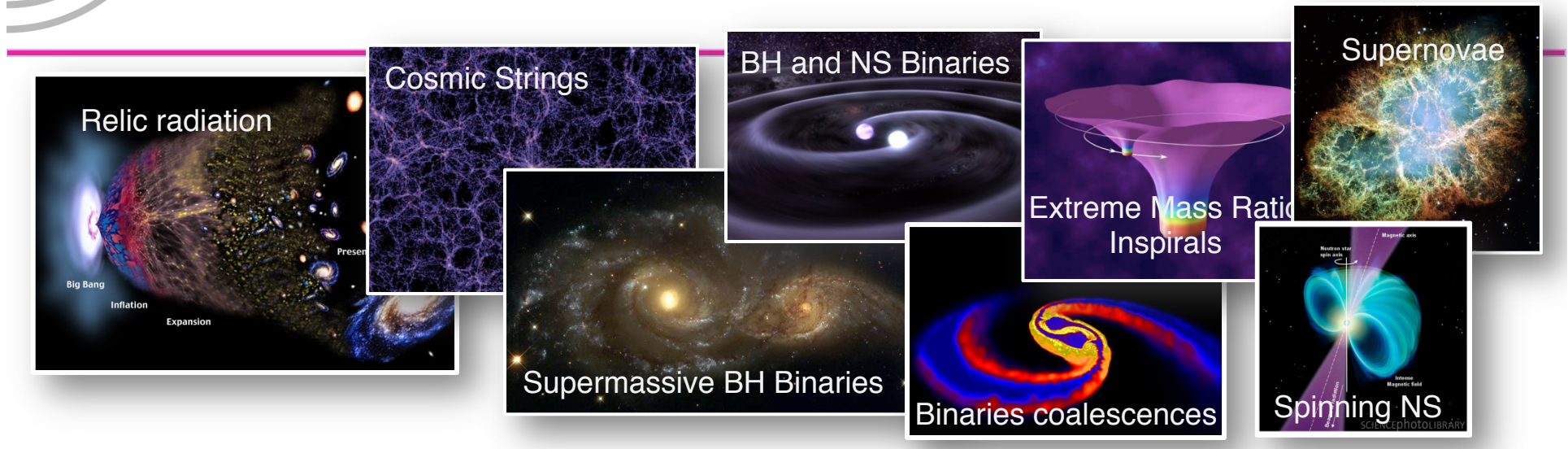
General Relativity
Spacetime itself is a medium
Geometry carries information



*Changes of matter
anywhere ripple the
geometry everywhere*

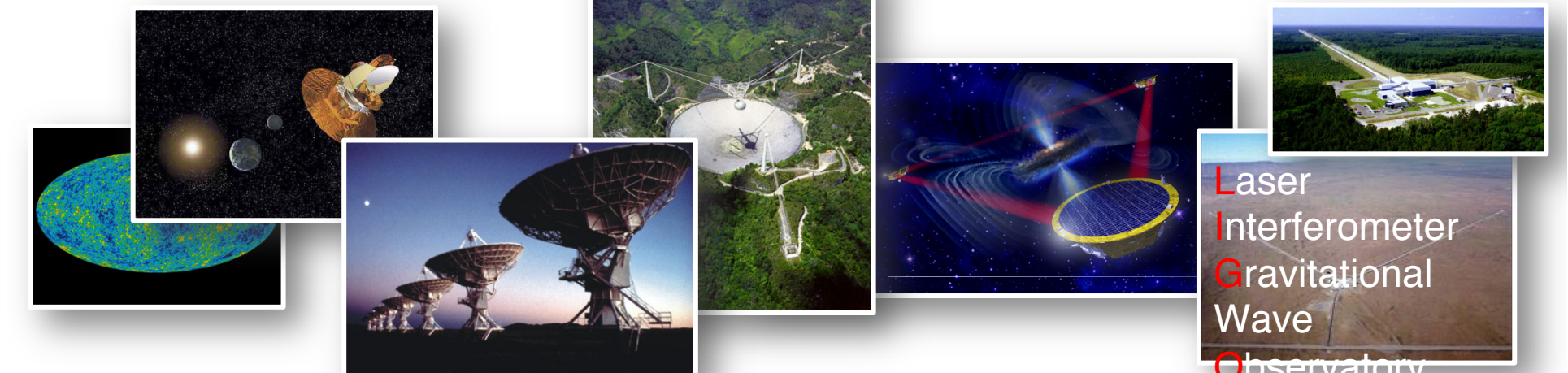


The GW Spectrum



10⁻¹⁶ Hz **10⁻⁹ Hz** **10⁻⁴ Hz** **10⁰ Hz** **10³ Hz**

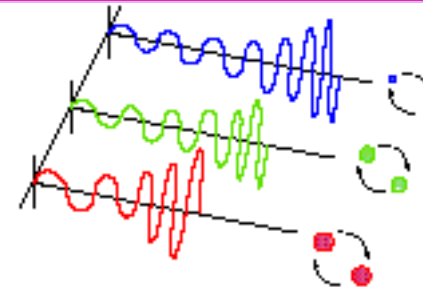
Inflation Probe **Pulsar timing** **Space detectors** **Ground interferometers**



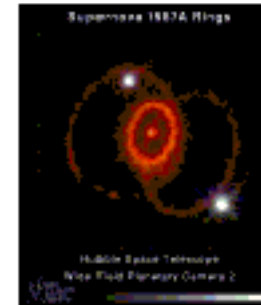
LIGO-G1300176

Some Expected Astrophysical Sources

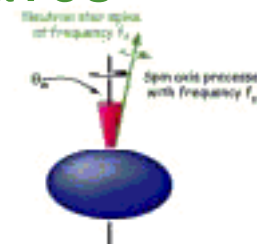
- Compact binary inspiral: “chirps”
 - » NS-NS, NS-BH, BH-BH



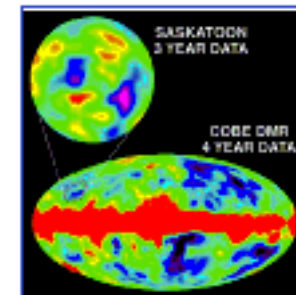
- Supernovas or GRBs: “bursts”
 - » GW signals observed in coincidence with EM or neutrino detectors



- Pulsars in our galaxy: “periodic waves”
 - » Rapidly rotating neutron stars
 - » Modes of NS vibration

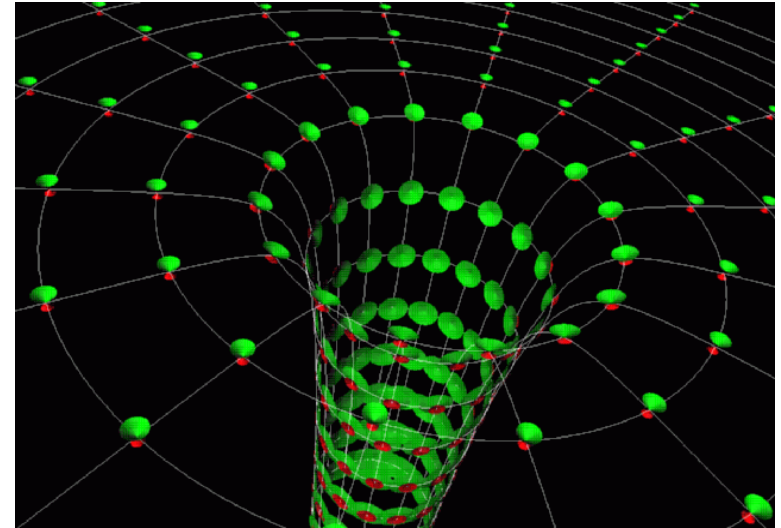
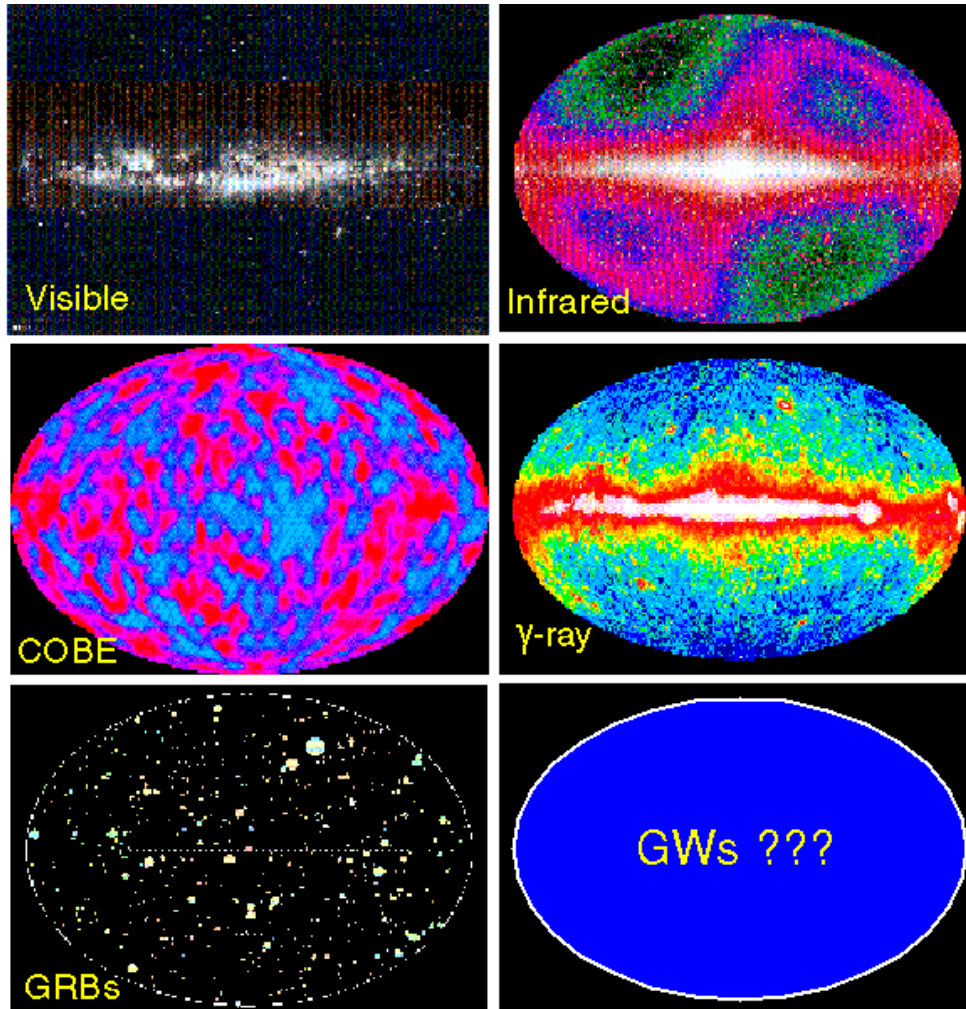


- Cosmological: “stochastic background”
 - » Probe back to the Planck time (10^{-43} s)





A New 'Sense'- A New Universe



Gravitational Waves will provide complementary information, as different from what we know as sound is from sight.



Great promise, but a great challenge...

A wave's strength is characterized by its *strain*

$$h = \Delta L / L$$

We can calculate the expected strain at Earth for, say, an orbiting binary system;

$$|h| \approx 4\pi^2 GMR^2 f_{orbit}^2 / c^4 r \approx 10^{-21} \left(\frac{R}{20\text{km}} \right)^2 \left(\frac{M}{M_{\odot}} \right) \left(\frac{f_{orbit}}{400\text{Hz}} \right)^2 \left(\frac{10\text{Mpc}}{r} \right)$$

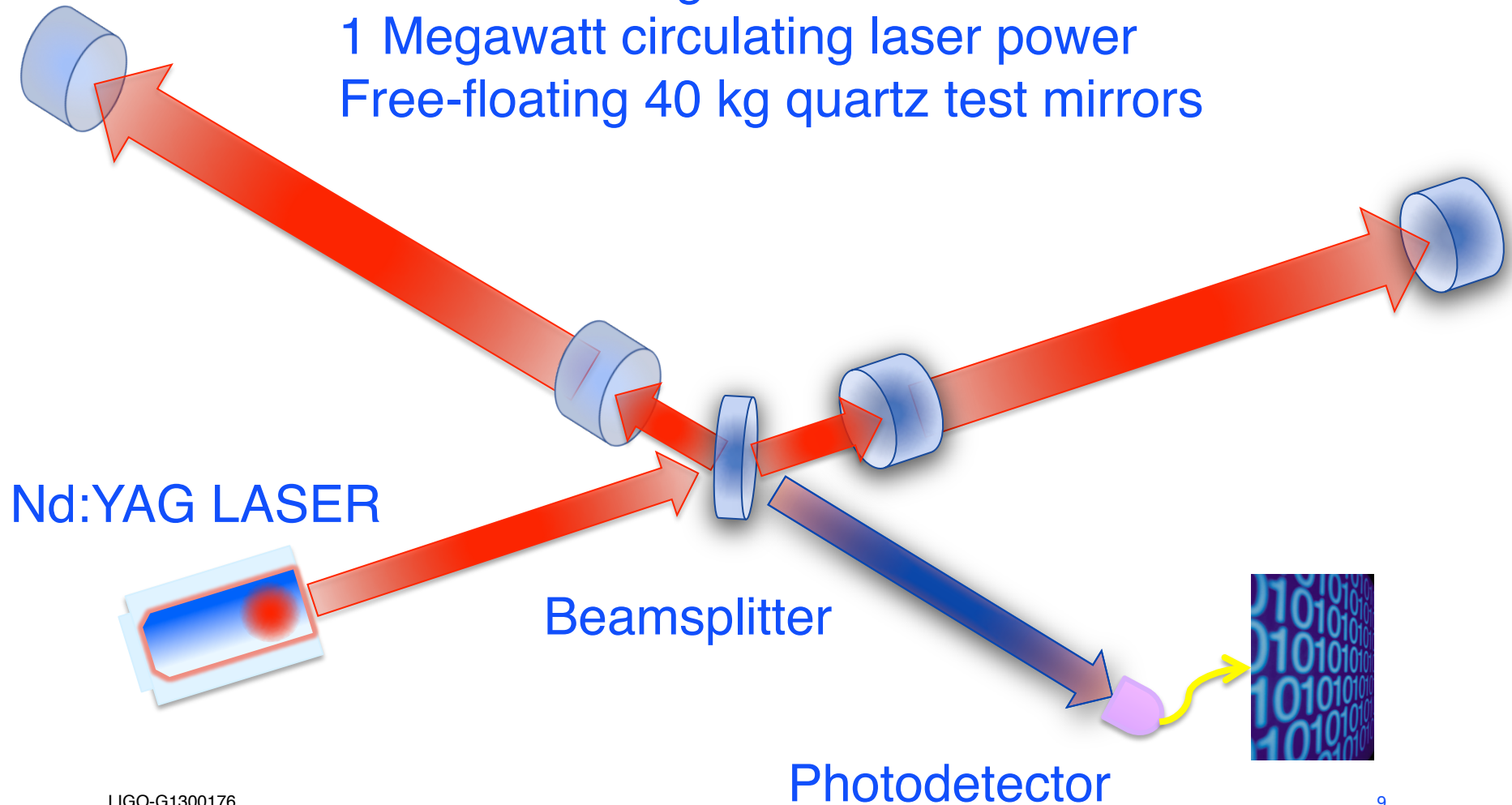
If we make our interferometer 4,000 meters long,

$$\Delta L = h \times L \approx 10^{-21} \times 4,000 \text{ m} \approx 10^{-18} \text{ m}$$

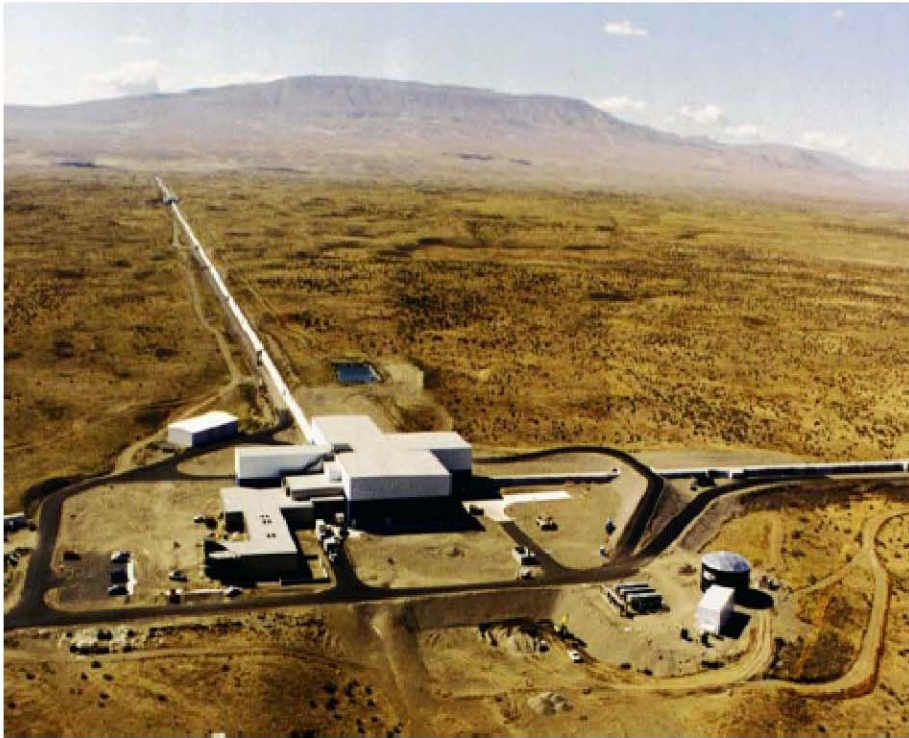


Laser Interferometer Gravitational-Wave Detector

4 kilometer long “arm” cavities
1 Megawatt circulating laser power
Free-floating 40 kg quartz test mirrors



LIGO Observatory Sites



LIGO Hanford Observatory [LHO]

26 km north of Richland, WA

2 km + 4 km interferometers in same vacuum envelope



LIGO Livingston Observatory [LLO]

42 km east of Baton Rouge, LA

Single 4 km interferometer



LIGO Scientific Collaboration



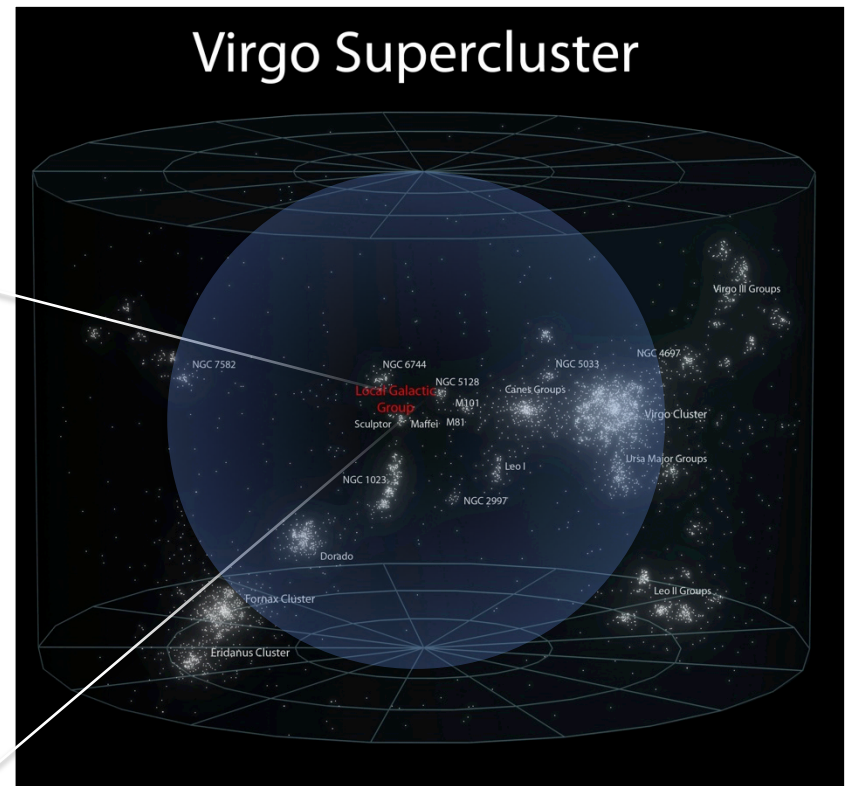
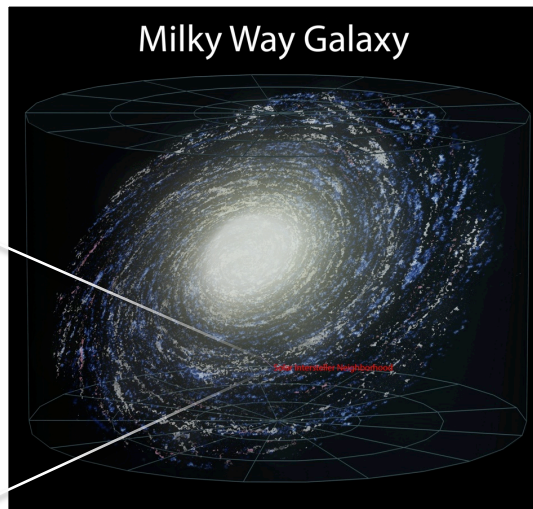


No Confirmed Detections Yet...

- First generation detectors reached about 100 galaxies
- Current predictions in range of 10^{-4} CBI yr⁻¹ galaxy⁻¹
- Need better sensitivity!

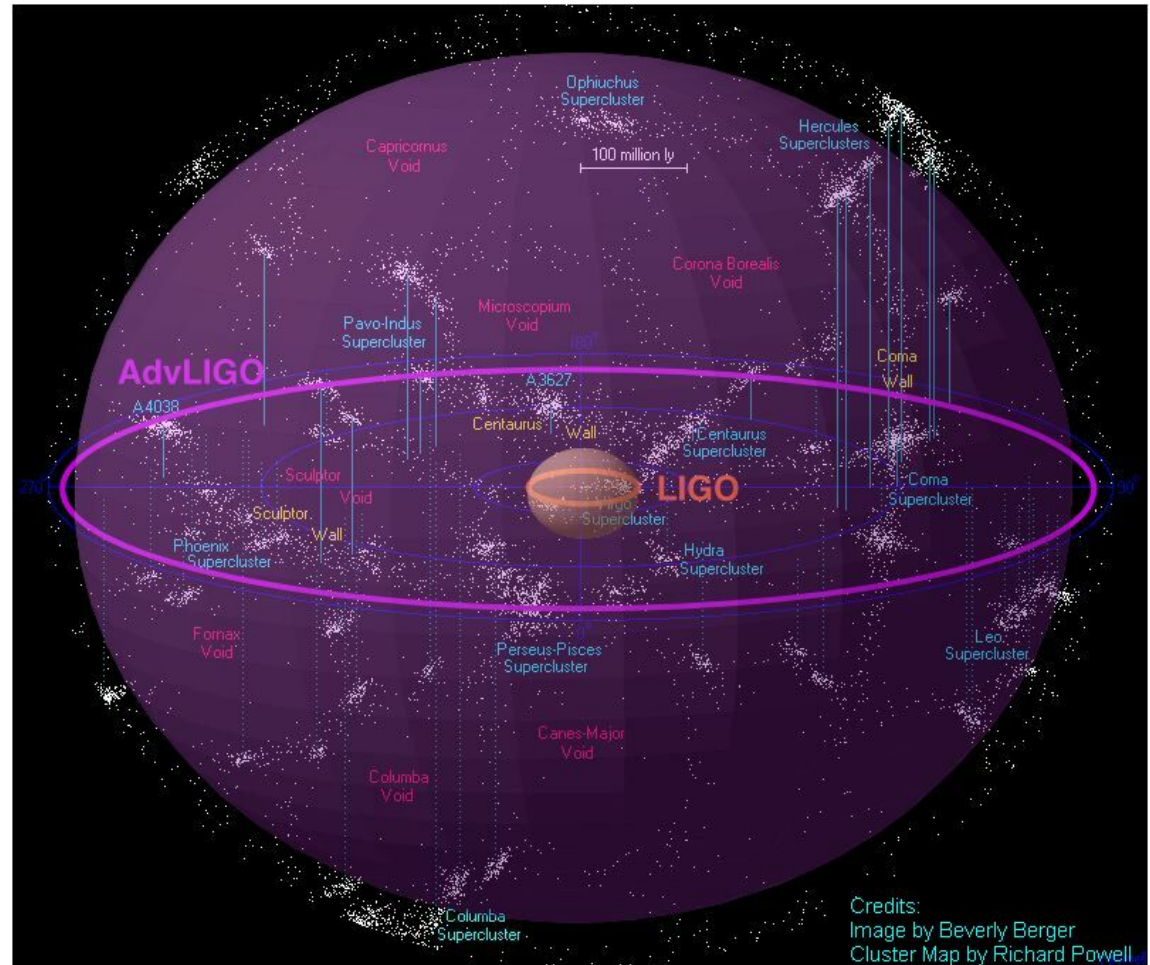


LIGO-G1300176



Advanced LIGO

- Initial detectors completed in 2000
- Design sensitivity achieved 2005
- Ran ~ 2.5 years
 - » No confirmed detection
- Facilities, vacuum system designed to be compatible with “ultimate” future interferometers
- Advanced LIGO detector upgrade funded ‘08, now being installed
 - » Design 10x more sensitive
 - » 1,000x greater observable volume (or event rate)



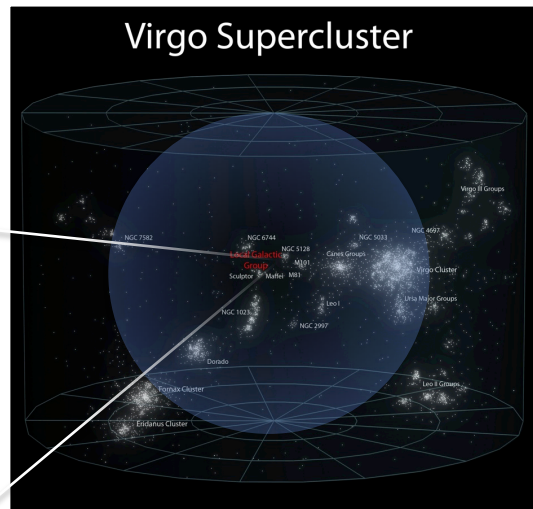


Advanced LIGO: 10x More Range

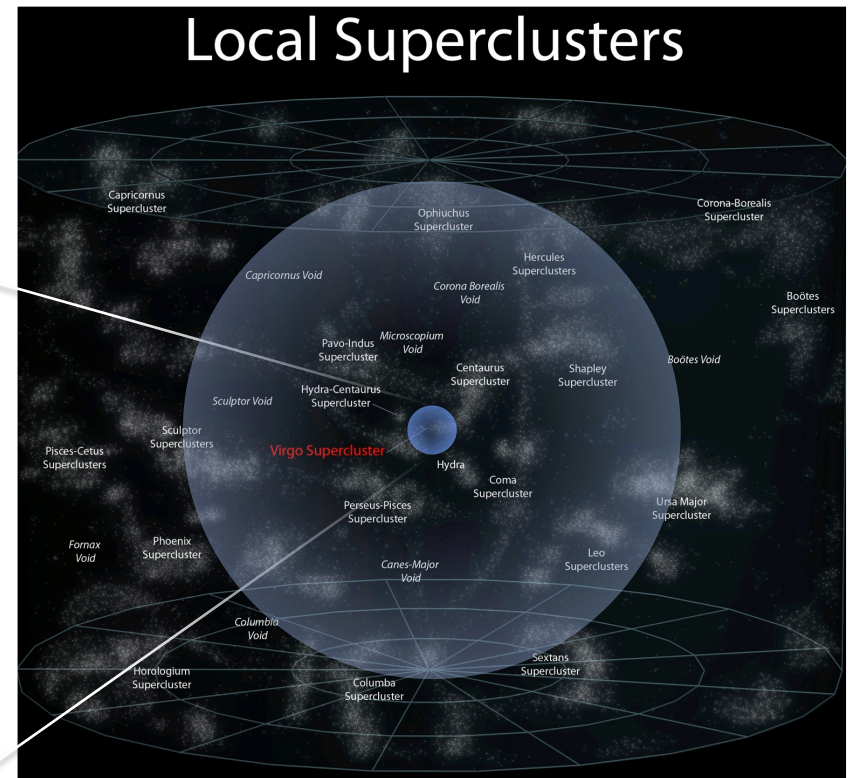
- Advanced detectors will reach about 100,000 galaxies
- Roughly 1 CBI event per month expected



LIGO-G1300176



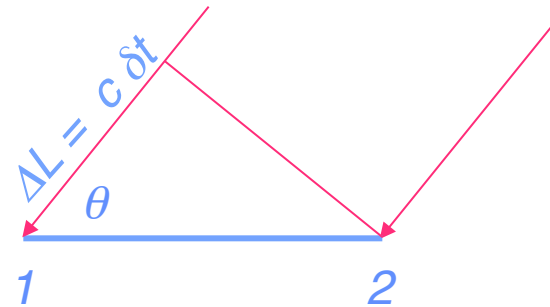
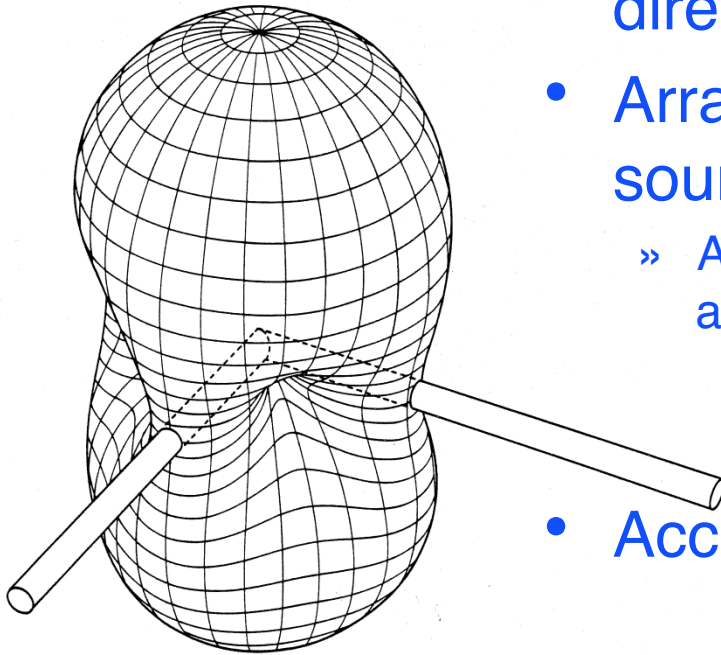
Initial LIGO Range



Advanced LIGO Range

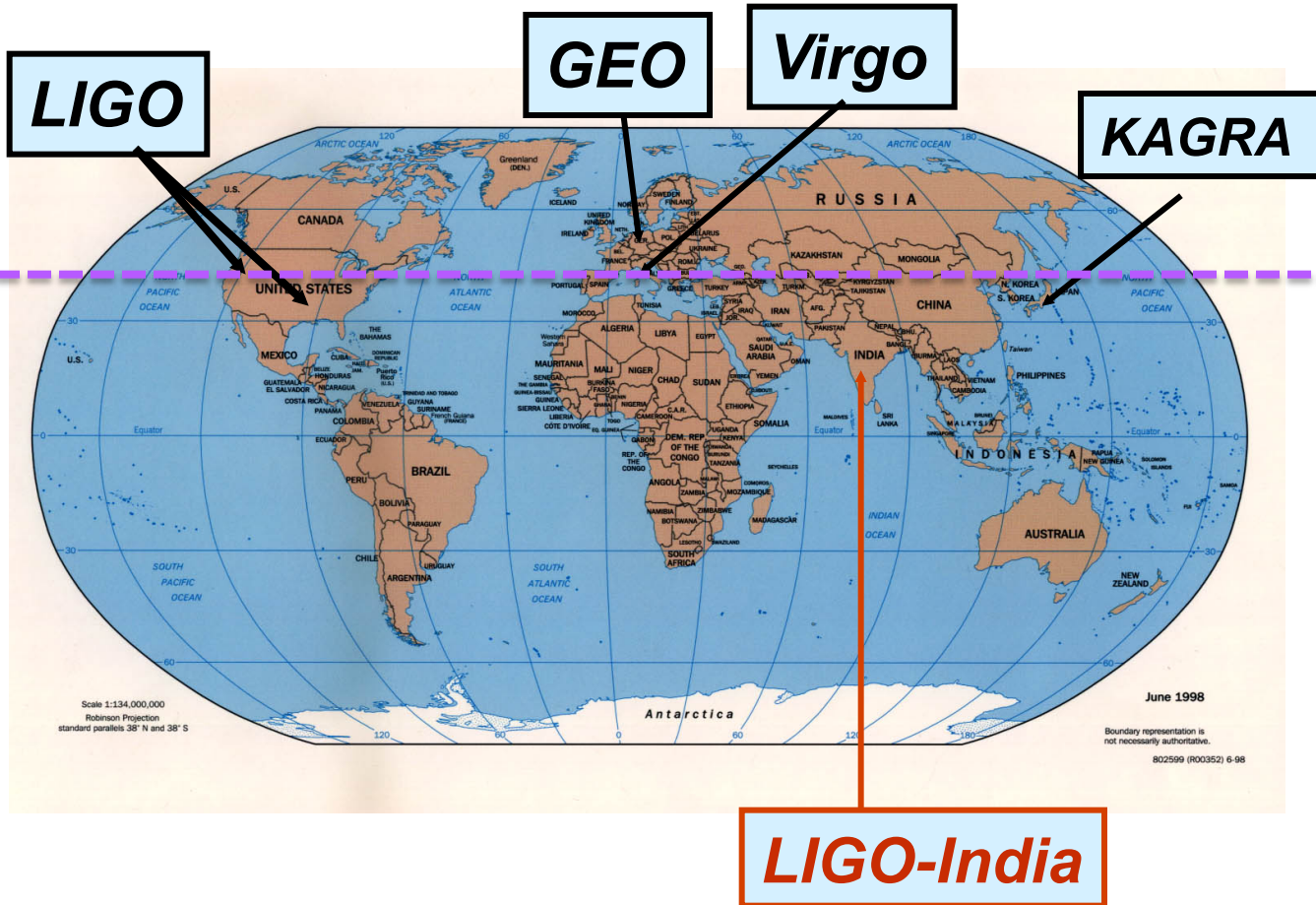
Source Localization and MultiMessenger Astrophysics

- GW detectors are nearly omnidirectional
- Array working together can determine source location
 - » Analogous to “aperture synthesis” in radio astronomy
- Accuracy tied to diffraction limit





Future Global Detector Network



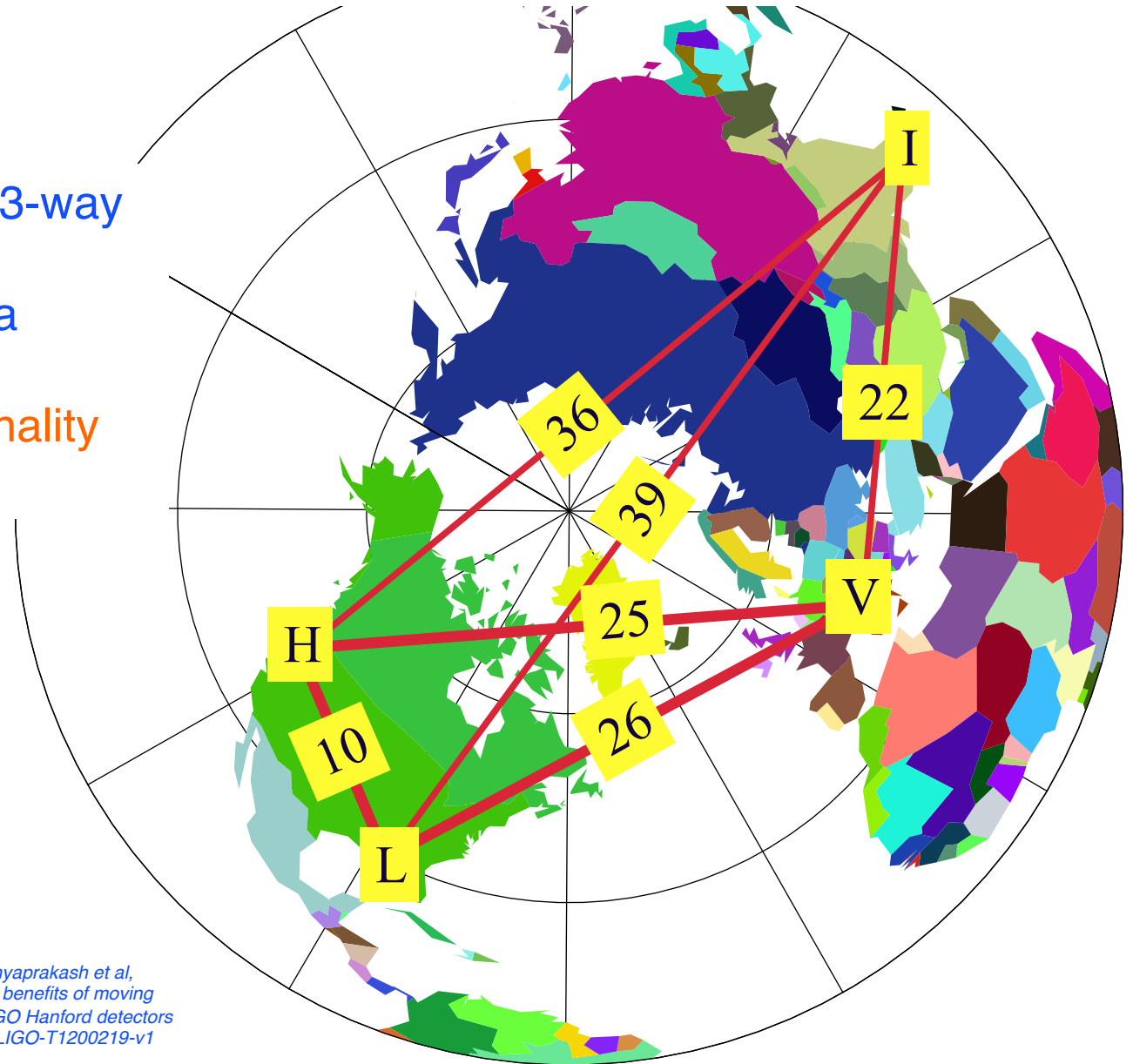
US, Europe and Japan detectors are close to co-planar— not optimal

India site out of plane breaks degeneracy, improving sky coverage



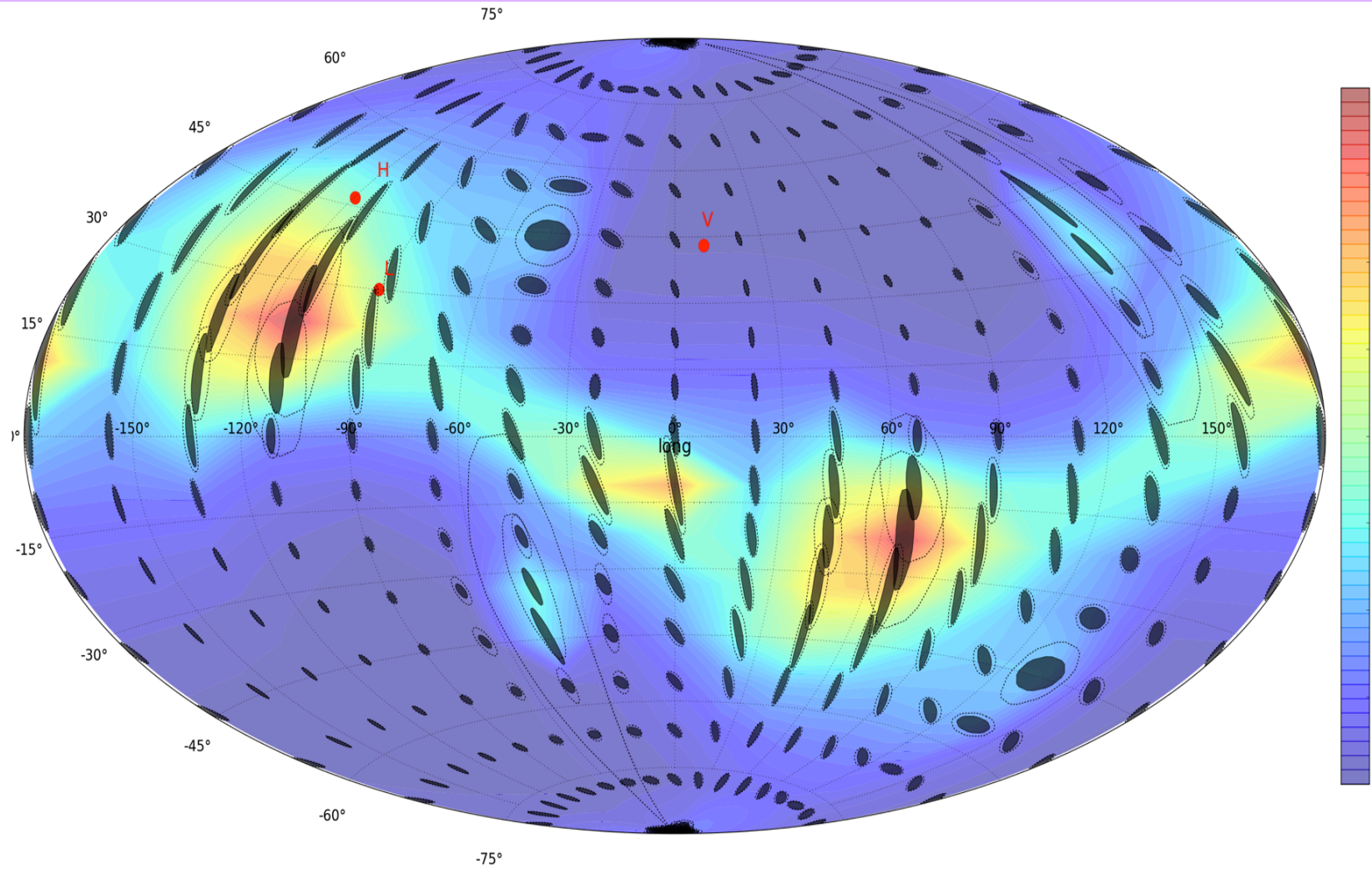
LIGO A Network with LIGO India, LIGO US and Virgo

- HLVI contains **four** 3-way networks
- Baselines with India are **1.5x** longer
- **~ 4x** better directionality (for given SNR)



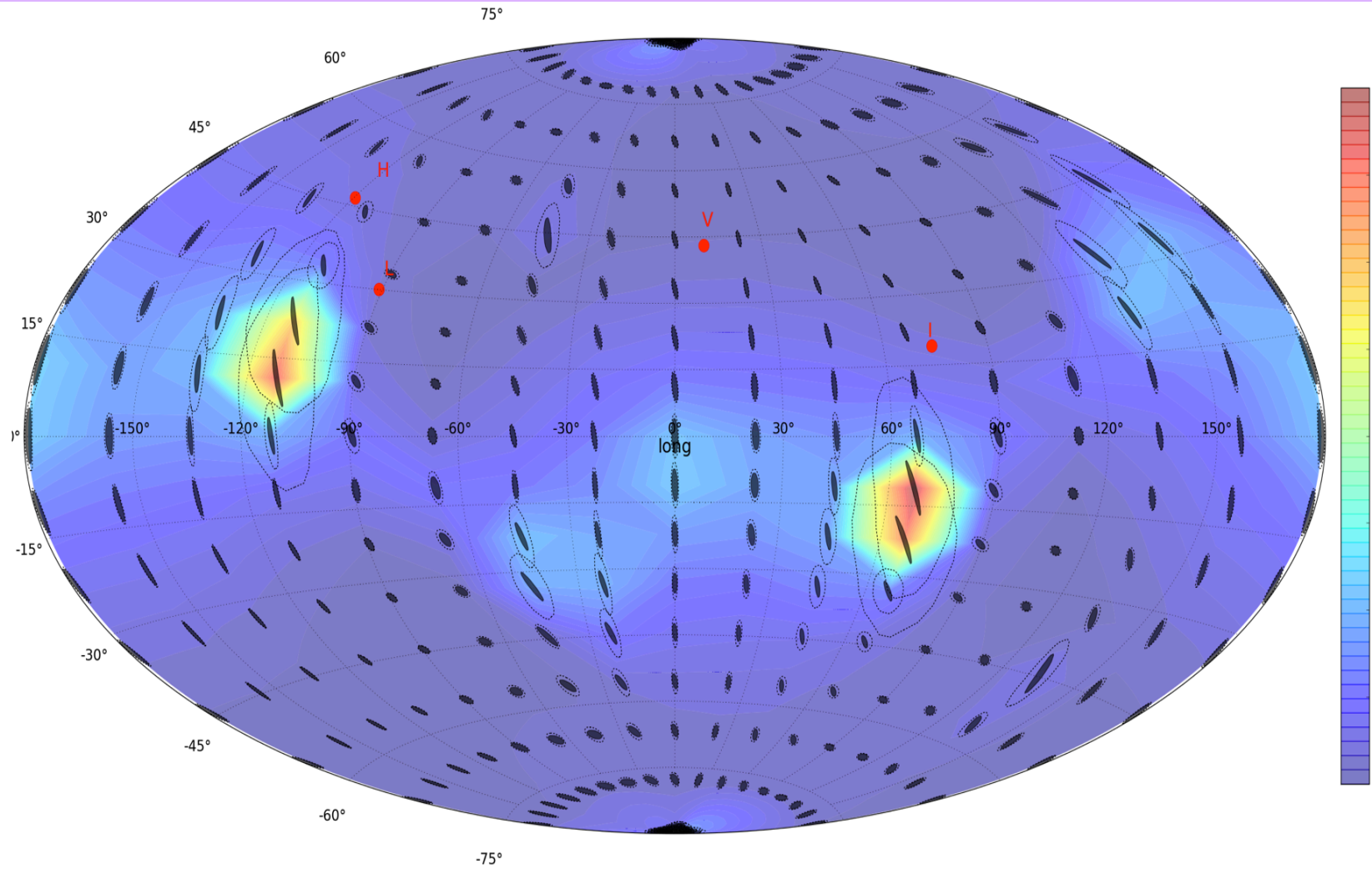
*B. S. Sathyaprakash et al,
'Scientific benefits of moving
one of LIGO Hanford detectors
to India', LIGO-T1200219-v1*

Timing Errors without LIGO-India





Timing Errors with LIGO-India

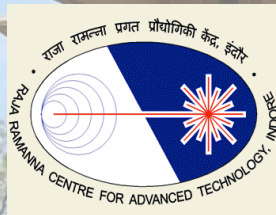




LIGO India Core Team



IPR, Gandhinagar – Facilities, Beam Tubes, Vacuum Equipment, Controls



RRCAT, Indore – Interferometer Optimization, Installation, Commissioning

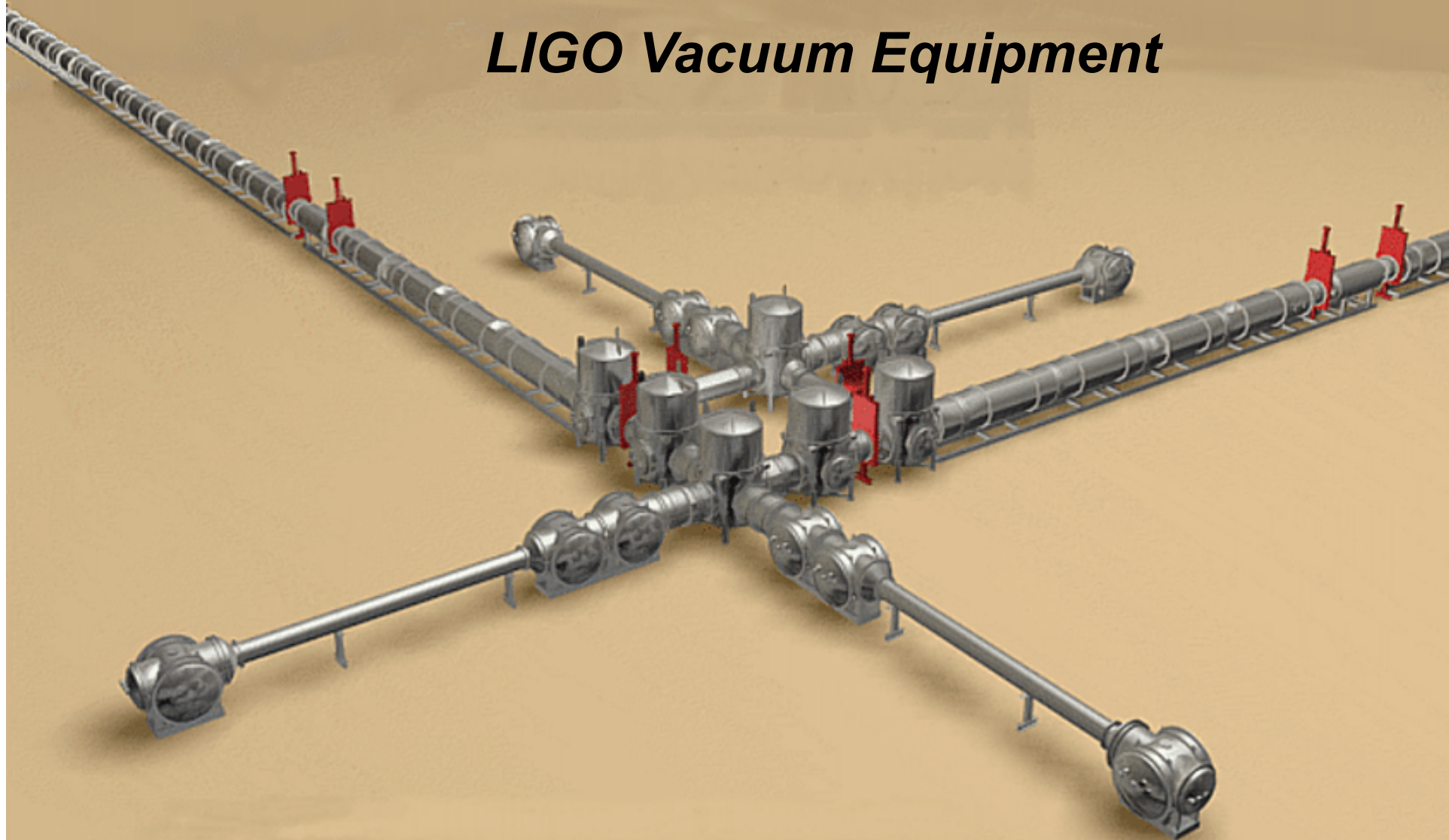


IUCAA, Pune – Site survey, Astrophysical Data Analysis & Computing



LIGO-US WILL CONTRIBUTE COMPLETE INTERFEROMETER PLUS TECHNOLOGY TRANSFER AND TECHNICAL SUPPORT

LIGO Vacuum Equipment





LIGO Really Has Two Vacuum Systems

“Vacuum Equipment:” Chambers, pumps, instruments

- Houses detector apparatus
- Isolation (valves), access (doors)
- Electrical, mechanical, optical penetrations
- Pumping & instrumentation
- Somewhat “conventional”
- $F:A \sim 10^{-2} \text{ Is}^{-1}\text{cm}^{-2}$

Beam tubes

- A long hole in the air;
Never to be vented
- Highly “unconventional”
 - 20 million liters (per site)
 - 600 million cm^2 (per site)
 - 200 l/s char. conductance
 - $F:A \sim 10^{-5} \text{ Is}^{-1}\text{cm}^{-2}$





LIGO Vacuum Requirements

(partial list)

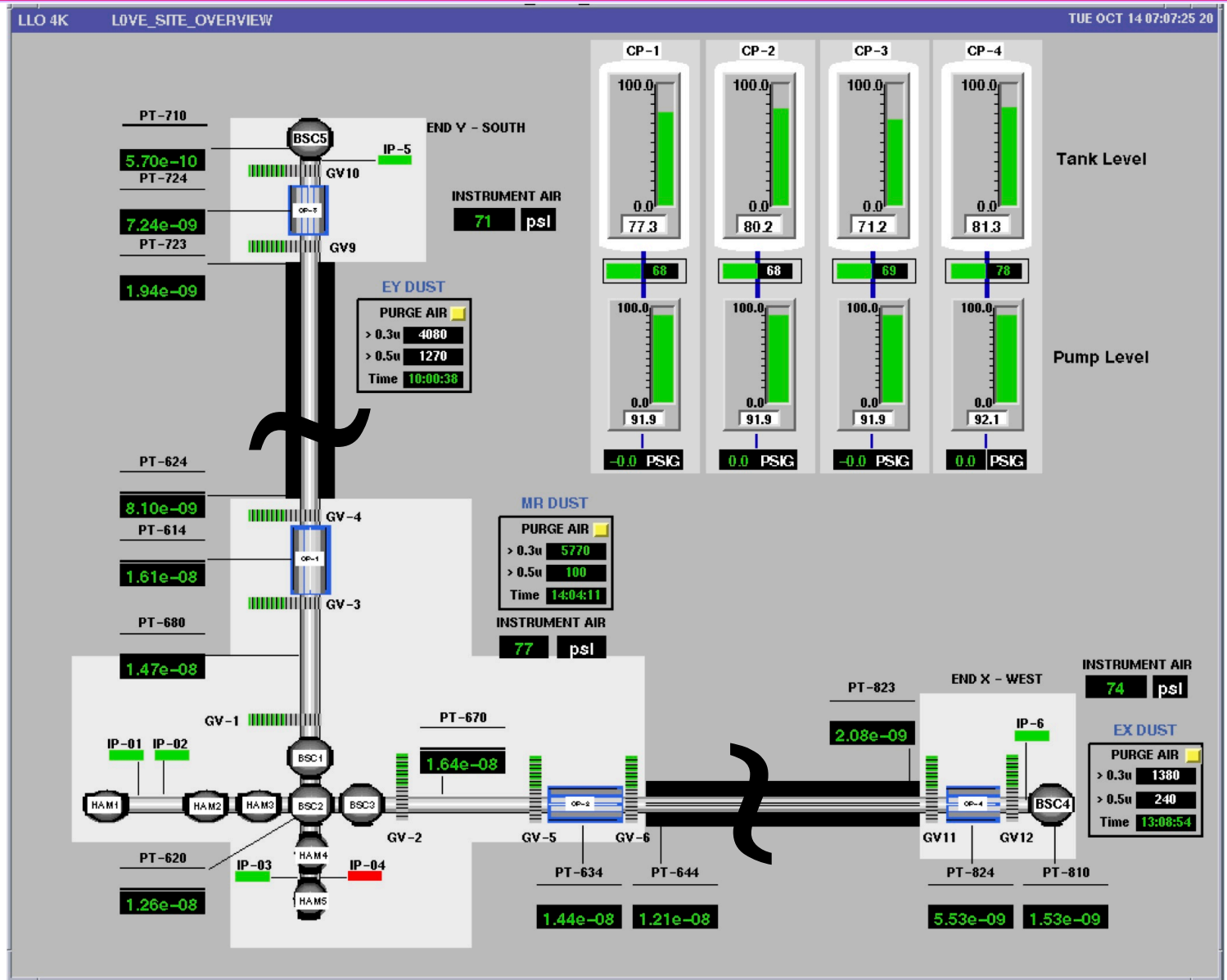
- **Light scattering phase noise from residual gas**
Function of molecular polarizability, transit speed and partial pressure
Primary goals for beam tubes:
 - $P(\text{H}_2) < 10^{-9}$ Torr
 - $P(\text{H}_2\text{O}) < 10^{-10}$ Torr
- **Contamination of optics**
Mirror absorption < 0.1 ppm
Hydrocarbons: < 1 monolayer/10 years
 - Aggressive cleaning and vacuum bake of every componentParticles: $< \text{one } 10 \mu\text{m particle}$ on any mirror
 - ISO Class 5 or better cleanroom protocol for worker access, internal components, surface exposure
- **Vibration-free environment**
 - No mechanical, turbo or closed-cycle cryo pumps in steady state operation

NB: Unlike accelerator, plasma, or aerospace applications, we have no radiation, thermal, or ion loading ; in LIGO outgassing is passive at ambient temperature



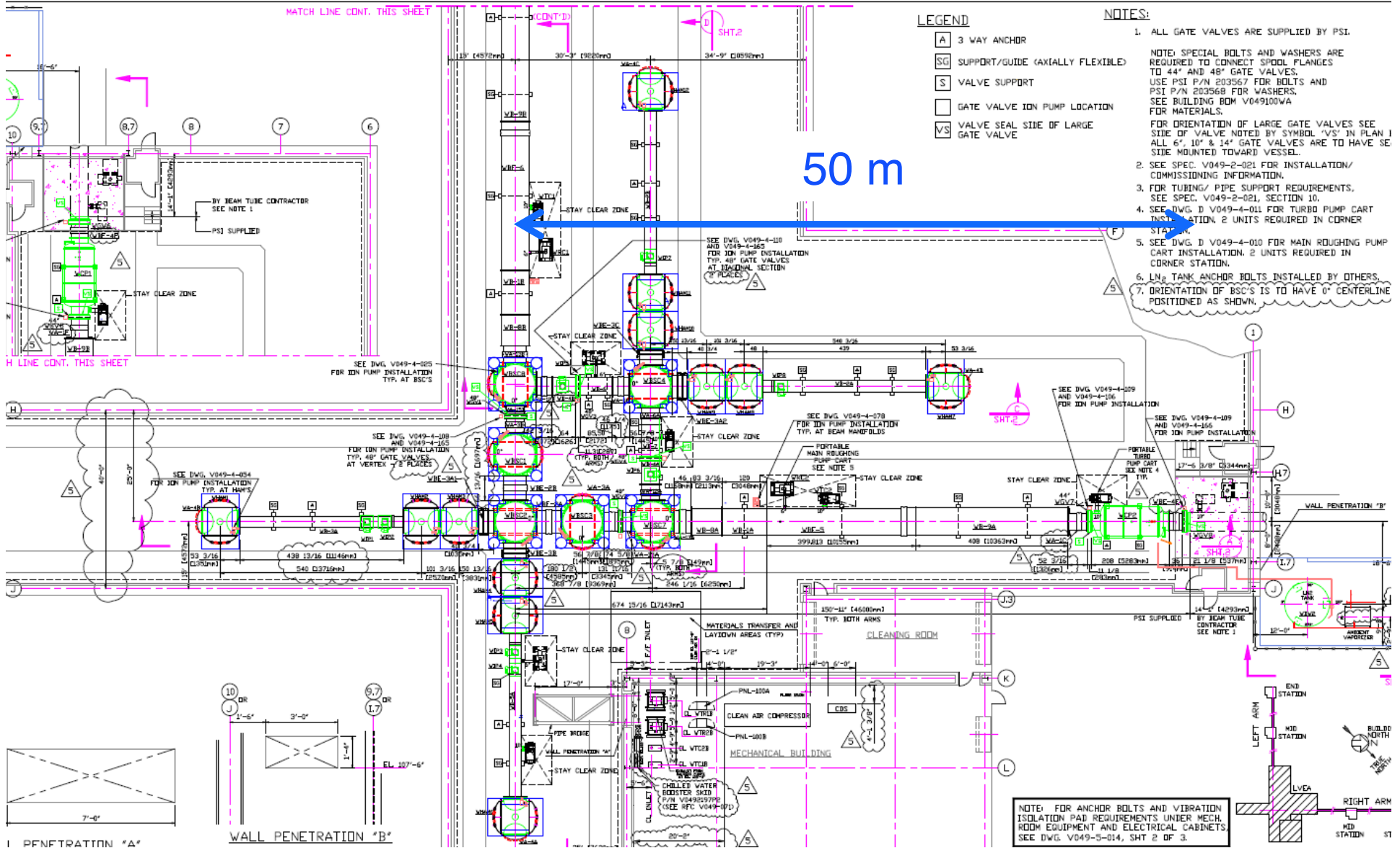
Vacuum System Schematic

10 m
4,000 m
50 m

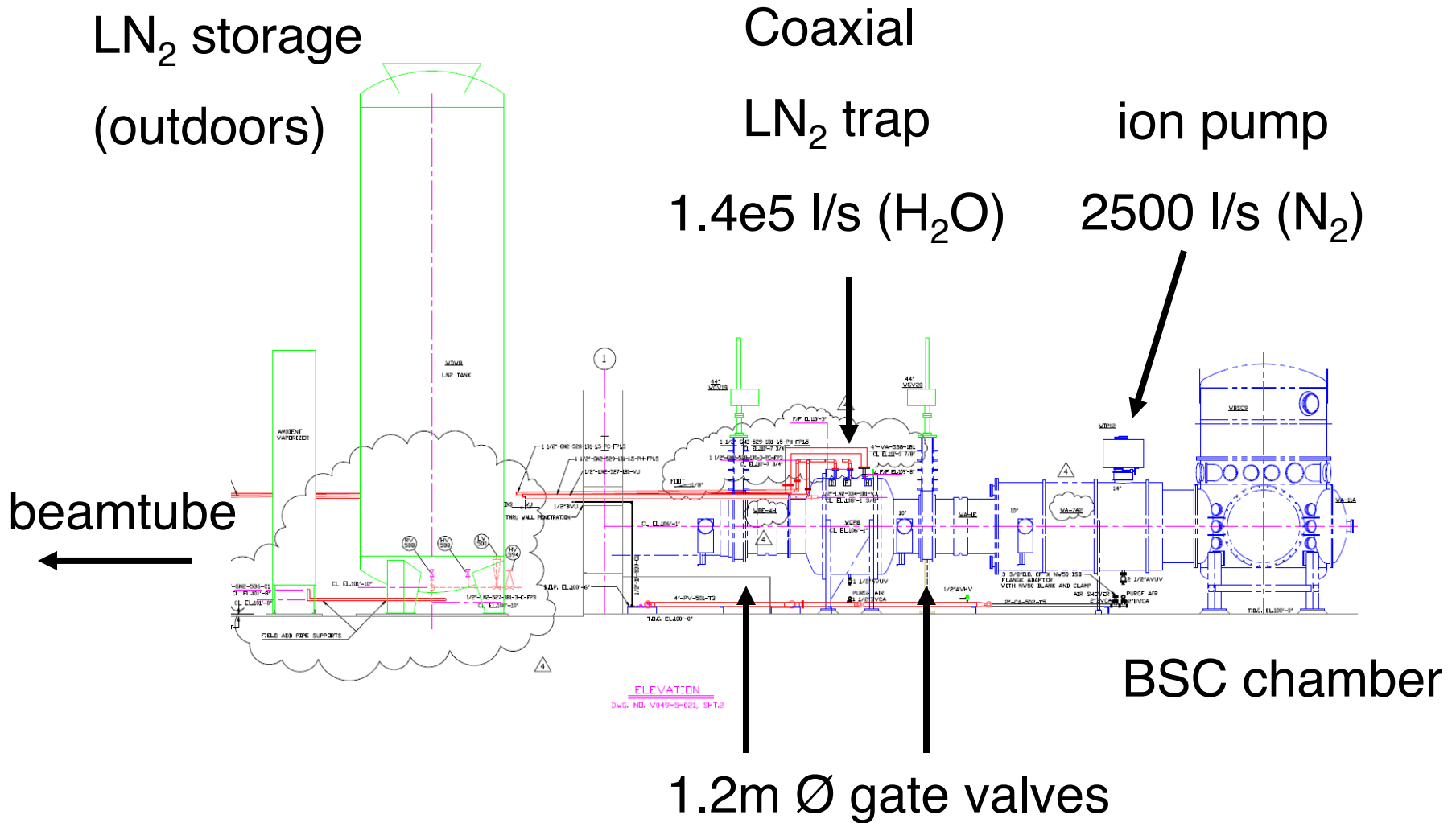




Corner Station Layout



End Station Arrangement



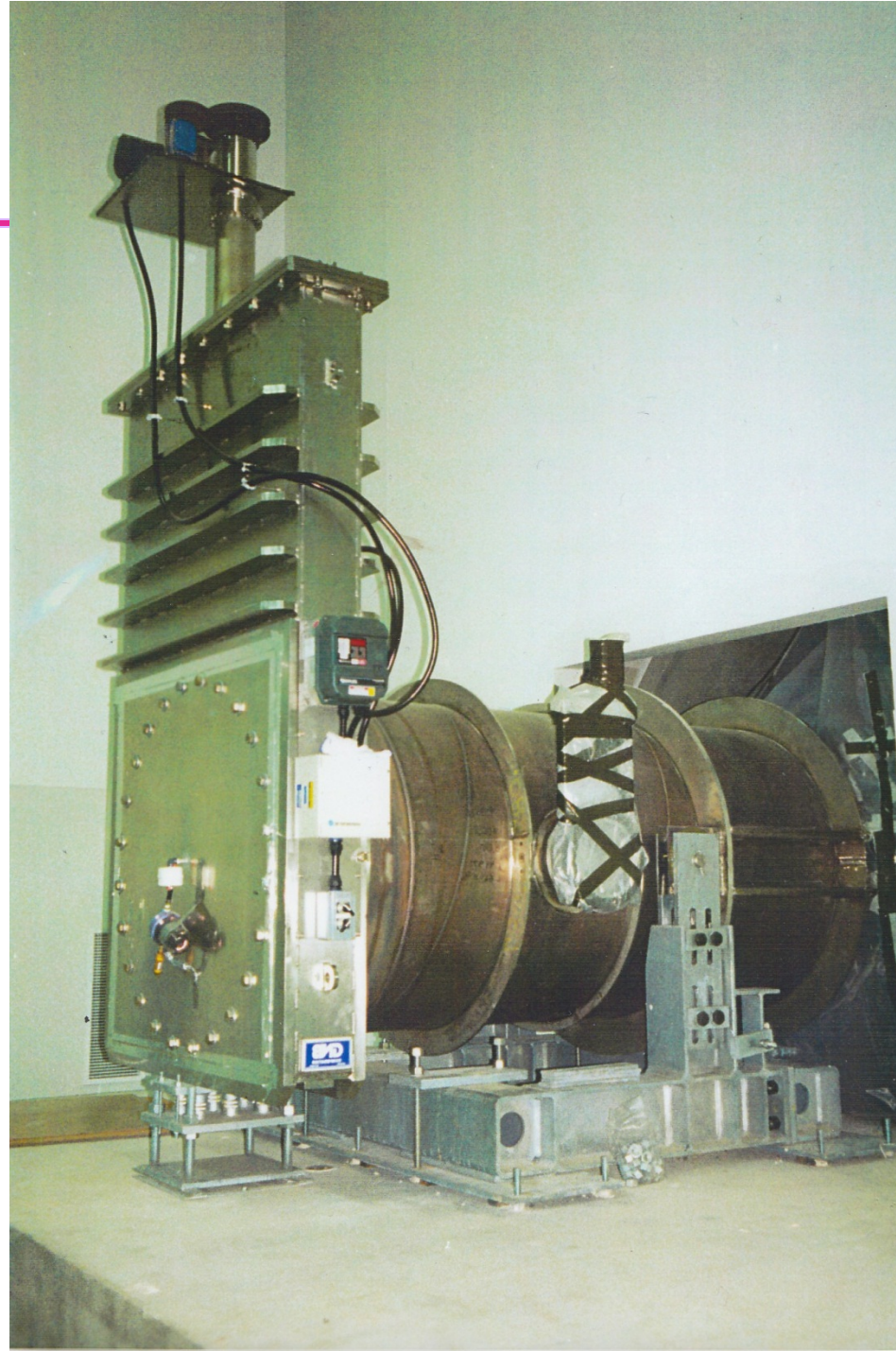
Beamtube Gate Valves

- 40" & 44" ID gate valves to isolate beamtubes, sections, LN2 traps
- Double O-ring gates & bonnet seals with pumped annulus*
- Two actuator varieties: *electric* (ballscrew) and *pneumatic* (cylinder)
- Custom design by GNB Corp.

* Principal volume is vulnerable when gate is open!



- In US, we installed beam tubes first to allow time for bakeout
- VE contractor supplied valves & pumps to tube contractor

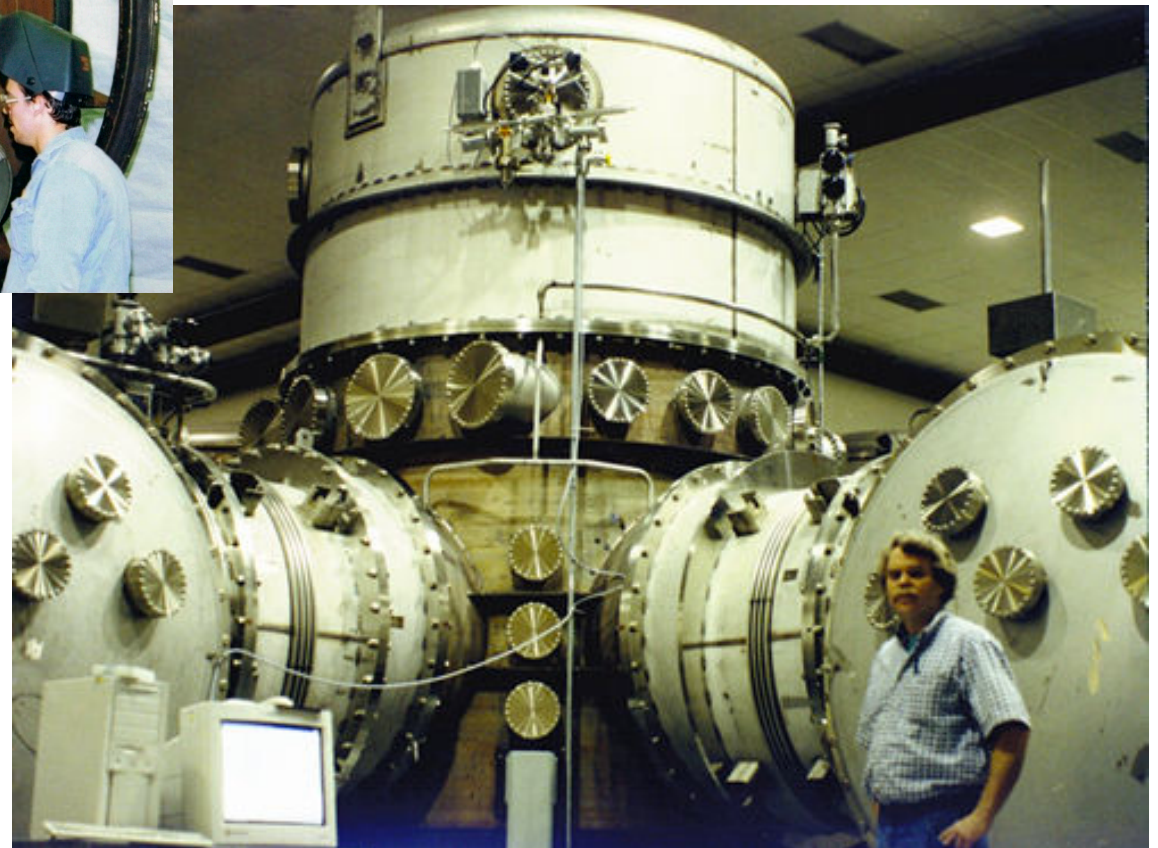
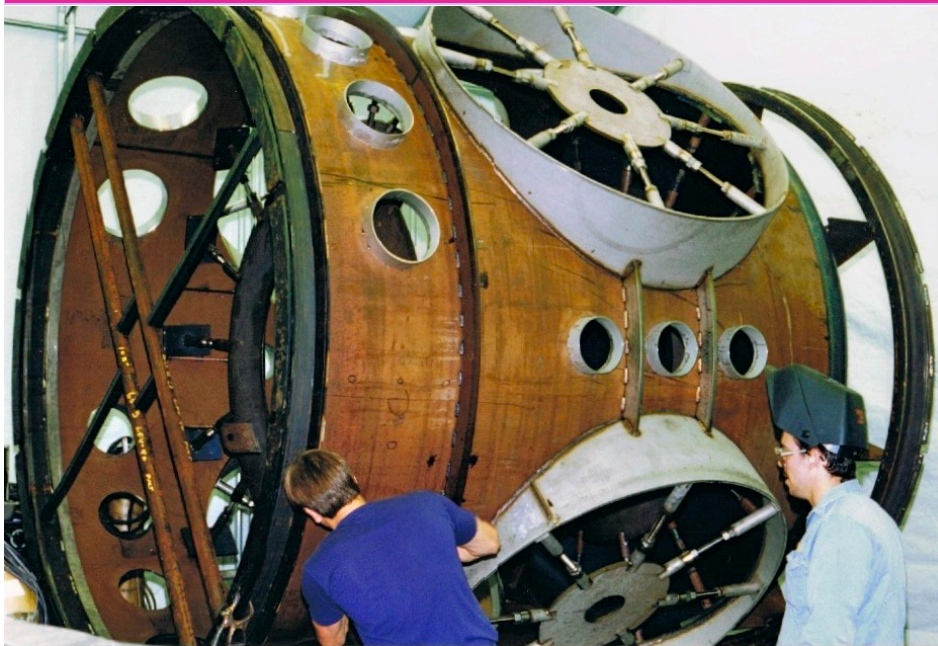


BSC chamber

(Basic Symmetric Chamber)

- 2.8m Ø x 5.5m h for large cavity optics
- Upper third is a removable dome
- Thin (10-15mm) 304L SS shell with welded stiffeners, F&D heads
- Combination of GTAW and plasma welding
- Major weldments stress-relieved*

*NOTE SURFACE FINISH!



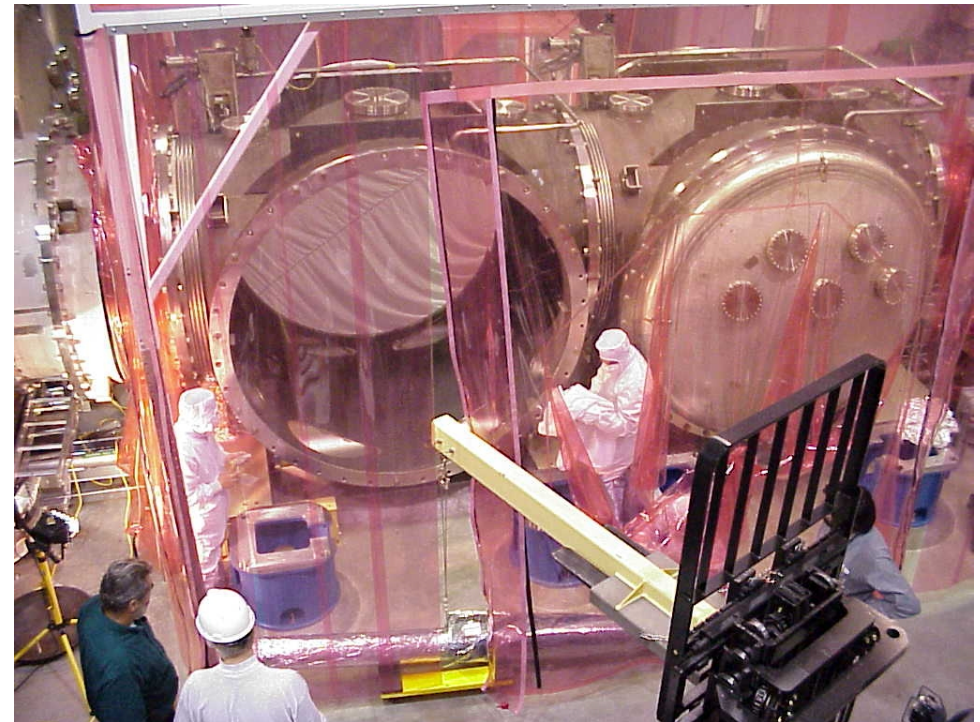
- Ports < 35cm Ø: ConFlat™
- Ports > 35cm Ø: Dual O-ring
 - Treated Viton elastomer
 - DRY (no grease)
 - Isolated pumped annulus between inner and outer seal
 - Permeation and damage tolerant

In Situ Chamber Bakeout



Particulate control: movable ISO Class 5 cleanrooms

- Part of VE contract due to special features required for chamber access



BSC Equipment Installation

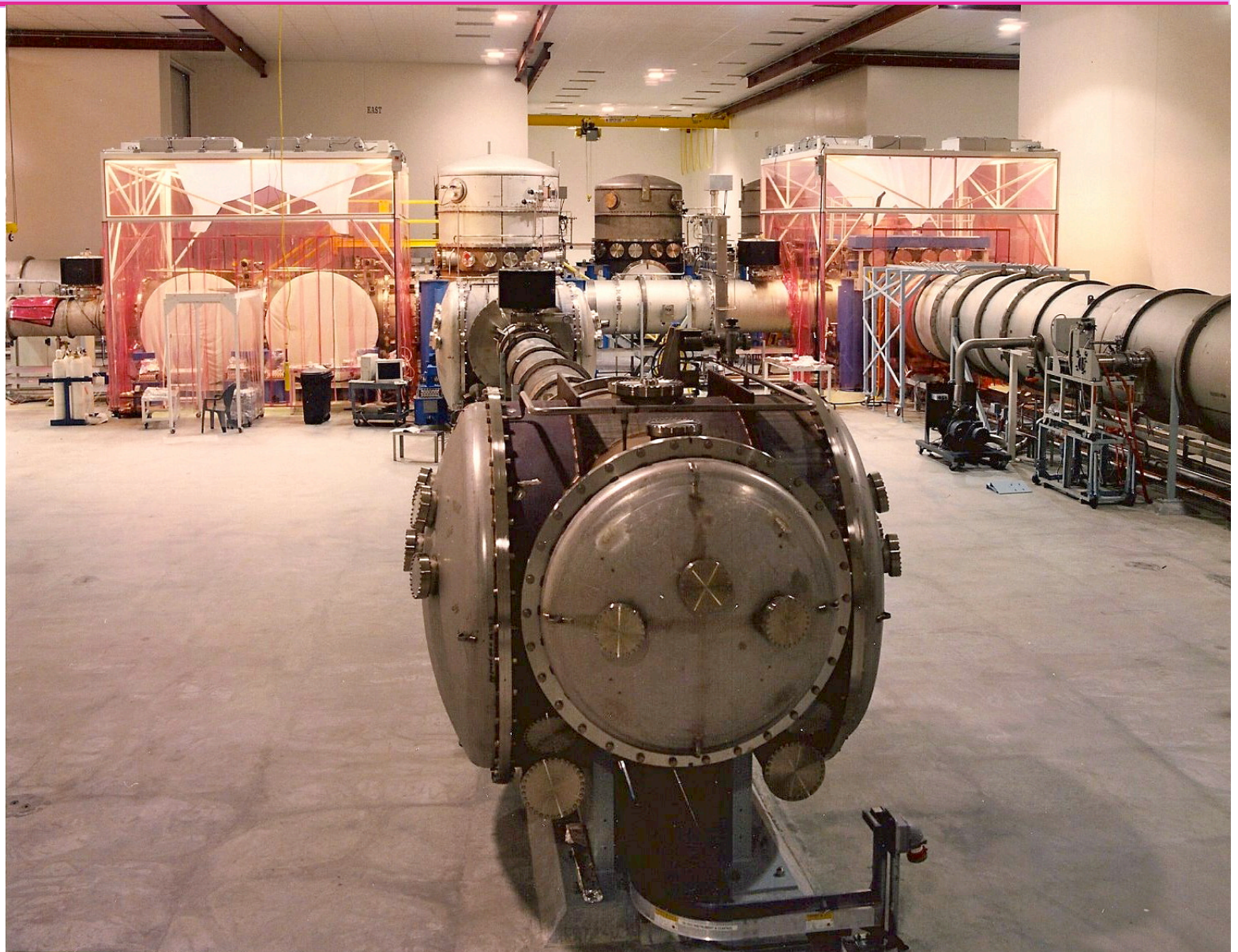


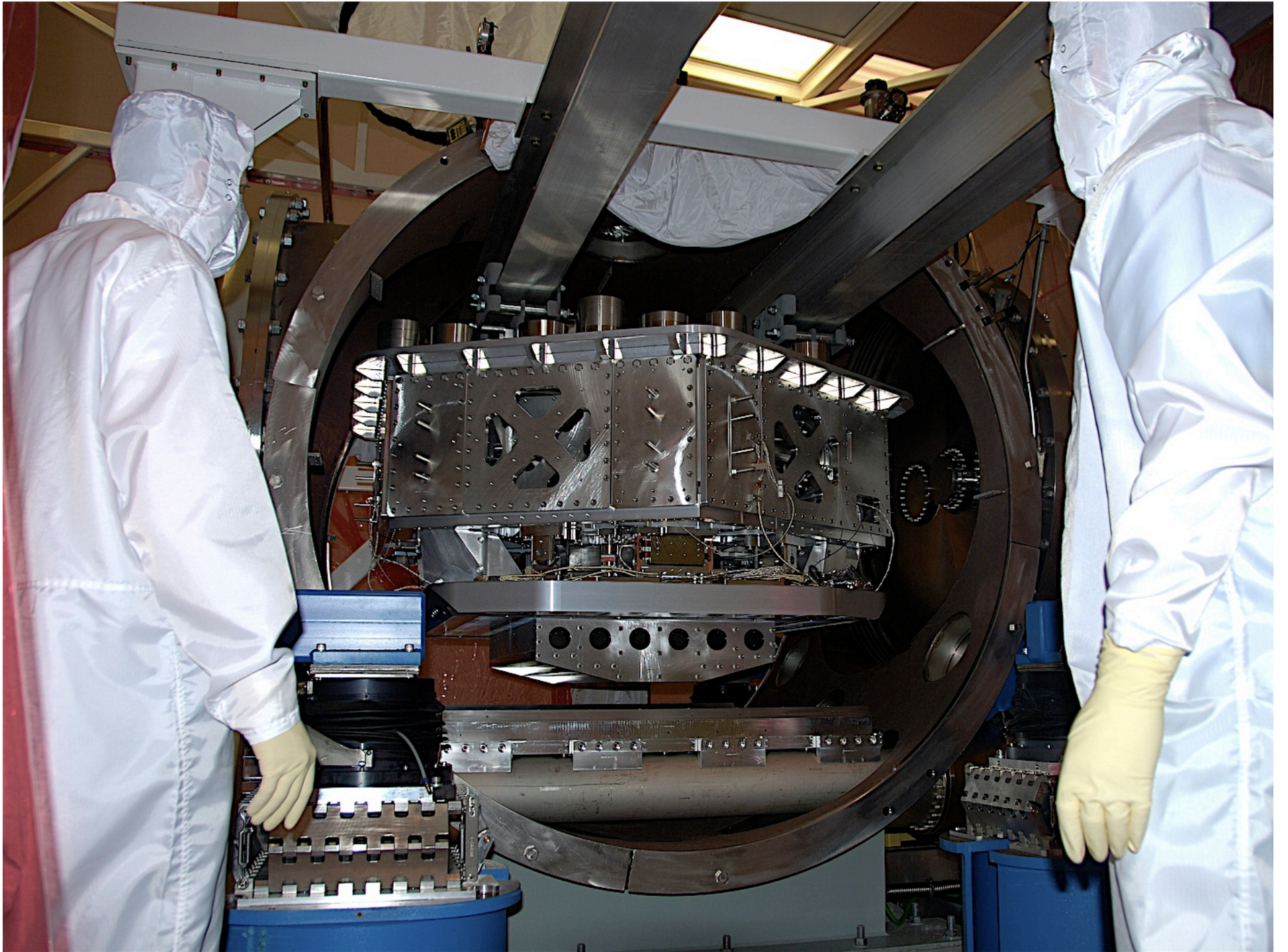


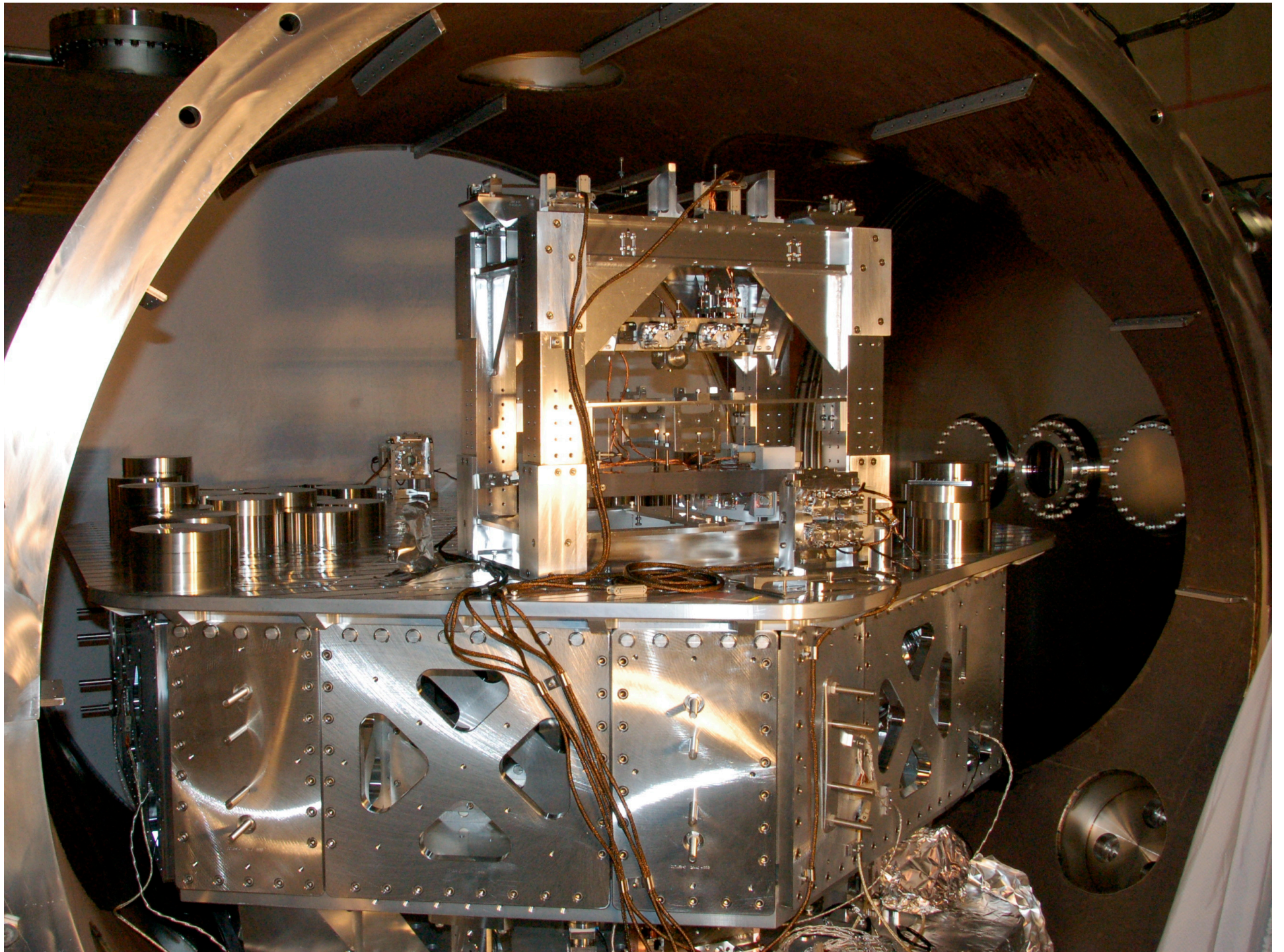
HAM chamber

(Horizontal Access Module)

- House complex input/output optics
- 2.1m \varnothing x 2m w
- More than 70% of area is removable access doors



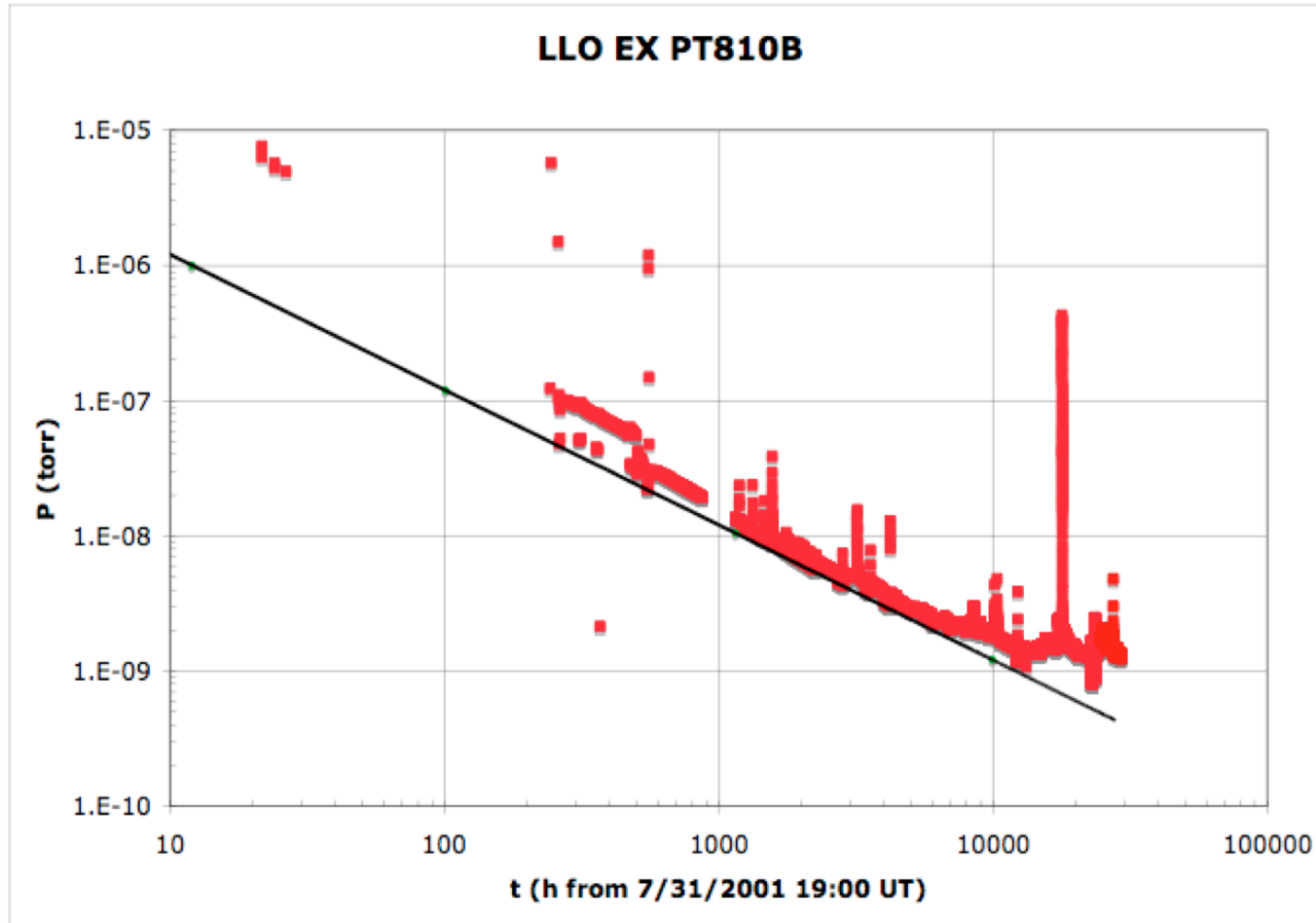








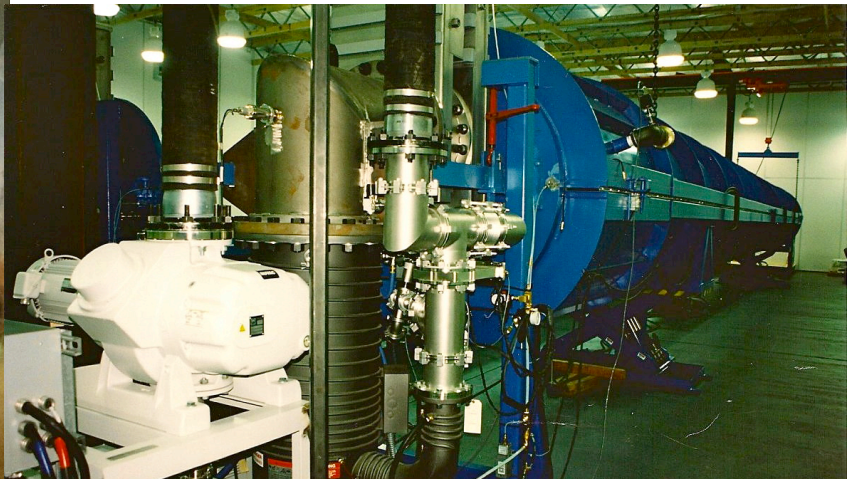
End Station Pressure Evolution after Backfill

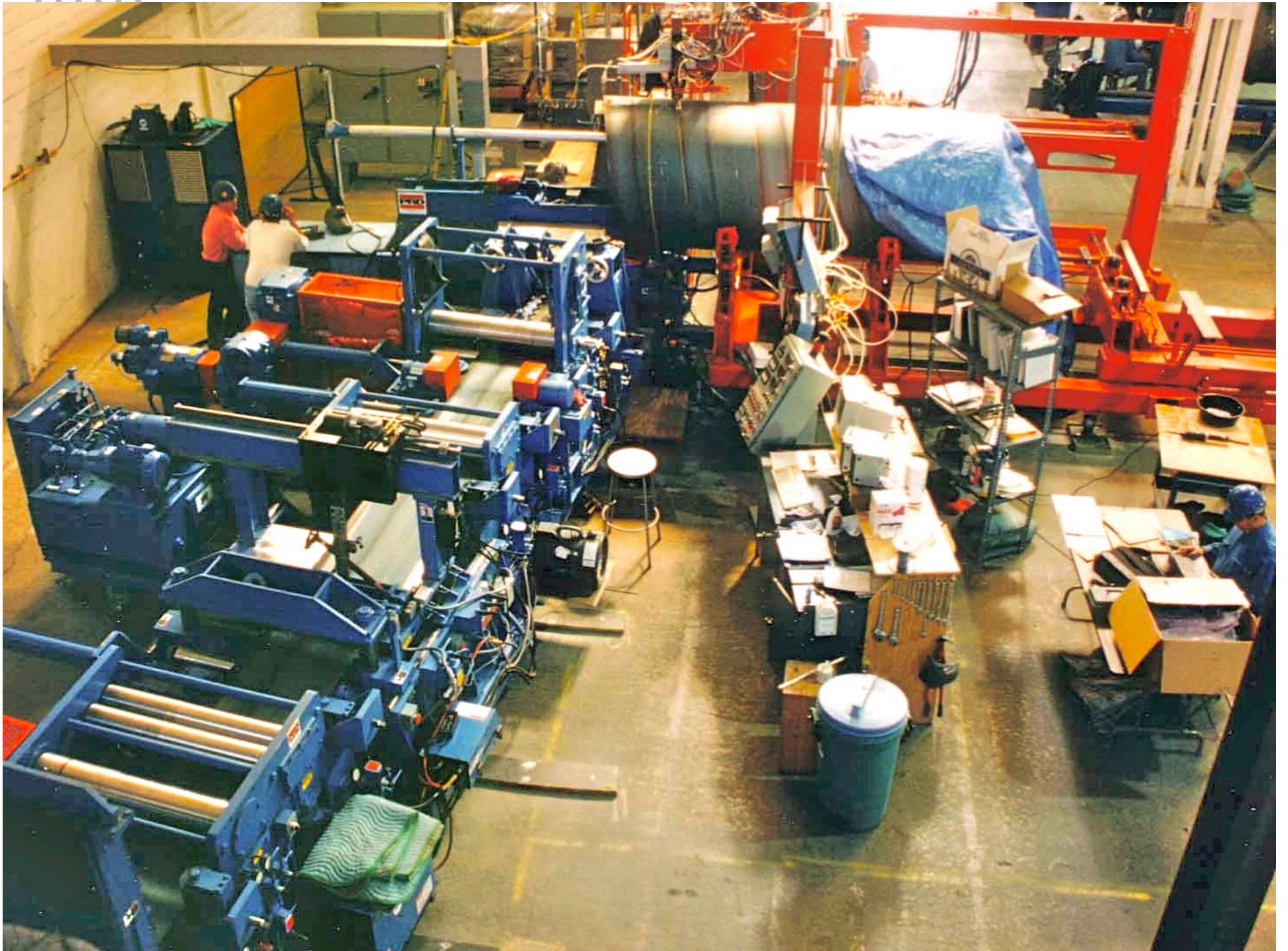




Beam Tubes

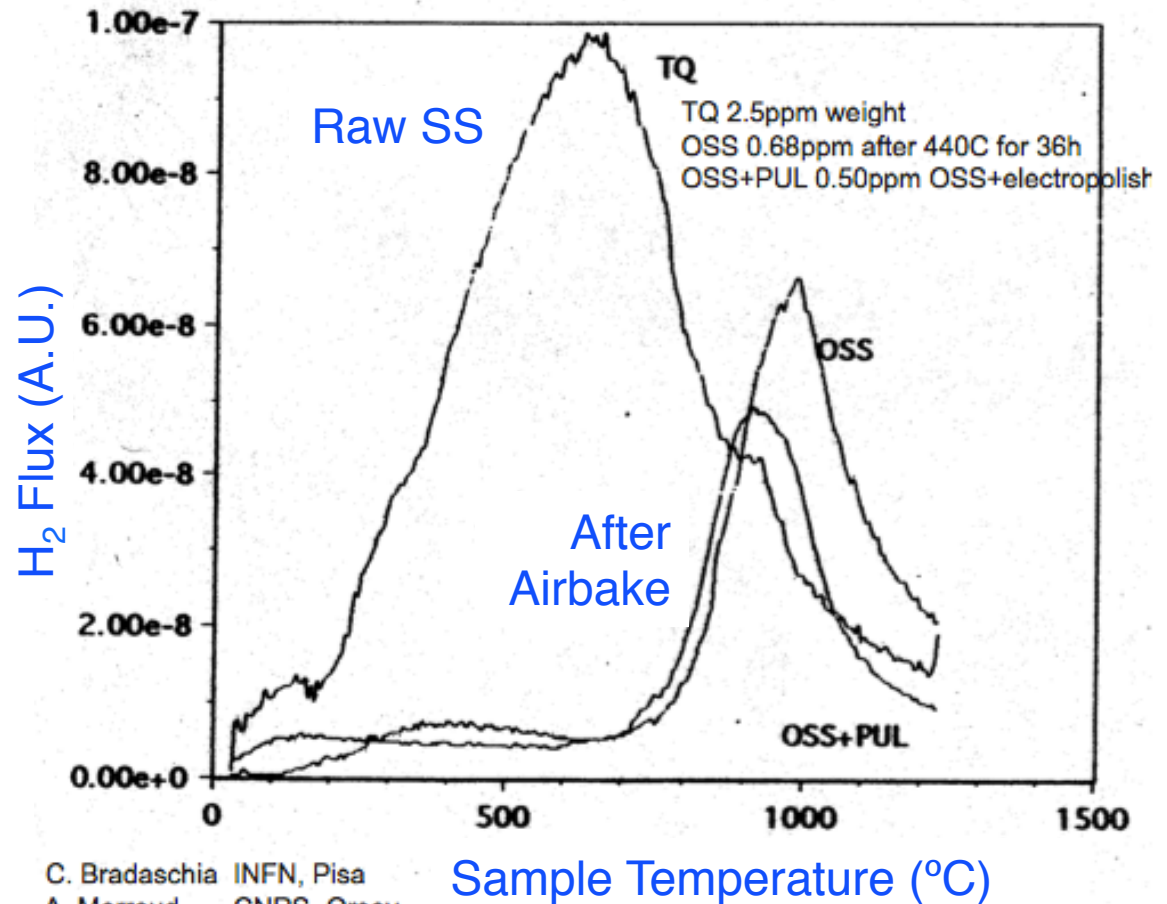
- 304L SS, 3.2 mm thick with external stiffeners
- raw coiled stock **air baked 36h @ 455C** to deplete hydrogen
 - » $J_{H_2} < 1e-13$ TI/s/cm²
 - » permits ultimate P without distributed pumps
 - » process developed by LIGO
- prepared coil spiral-welded into 1.2m tube on modified culvert mill
- 16m sections cleaned, leak checked, and capped
- FTIR analysis to confirm HC-free
- sections field butt-welded together in travelling clean room
- Over 50 linear km of weld—





Depleting H from raw SS before fabrication: An economical alternative to high T vacuum bakeout

- SS sheet from mill is baked in air 36 hours at 455 °C
- (Hotter treatment deemed inadvisable due to carbide formation)
- Total dissolved hydrogen is reduced ~ 3x
- Remaining H is tightly bound, high activation T
- Care is required in welding to avoid re-introduction of H



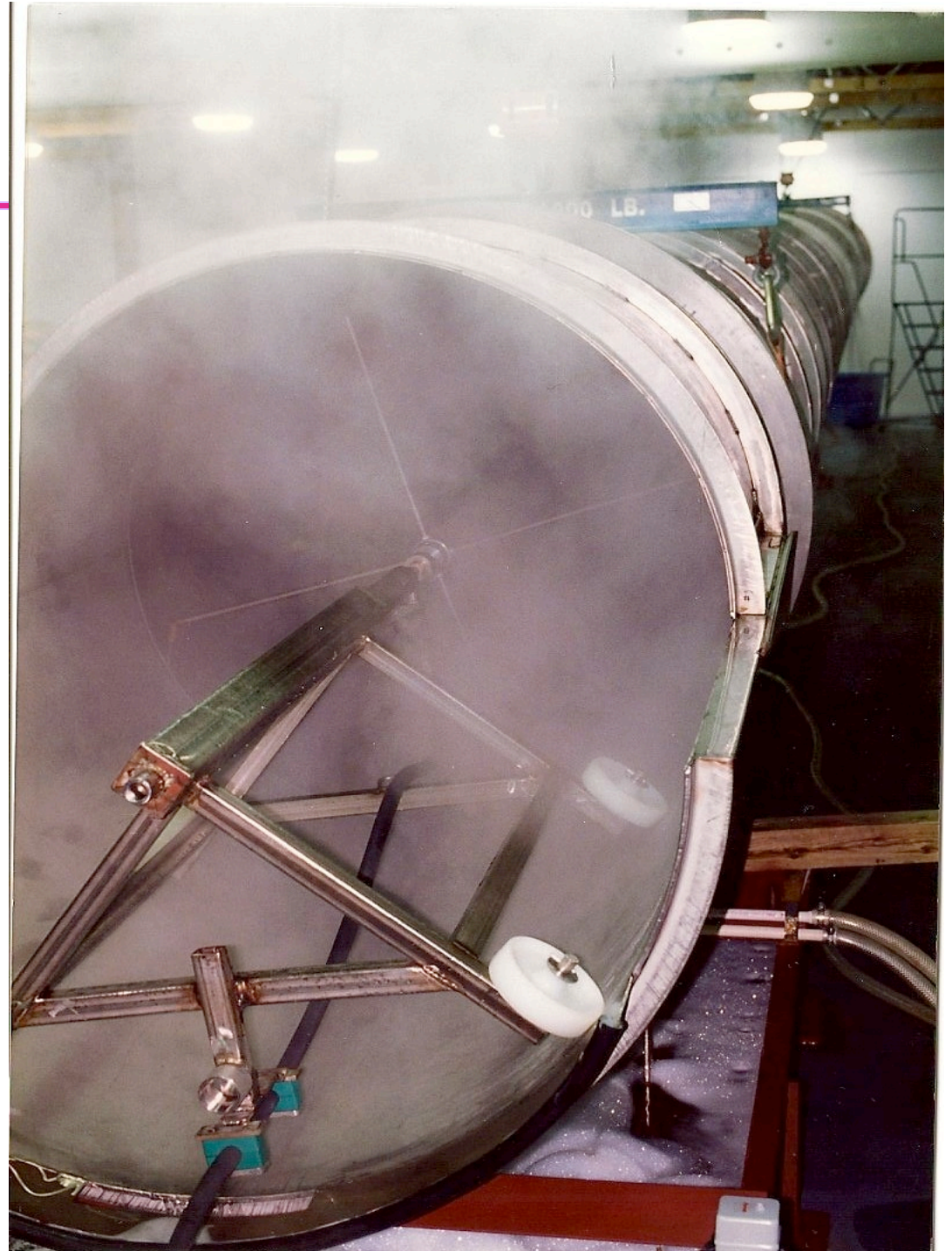
C. Bradaschia INFN, Pisa
 A. Marraud CNRS, Orsay

data courtesy of Virgo



Cleaned with
pressurized hot water
and detergent

QA by FTIR sampling



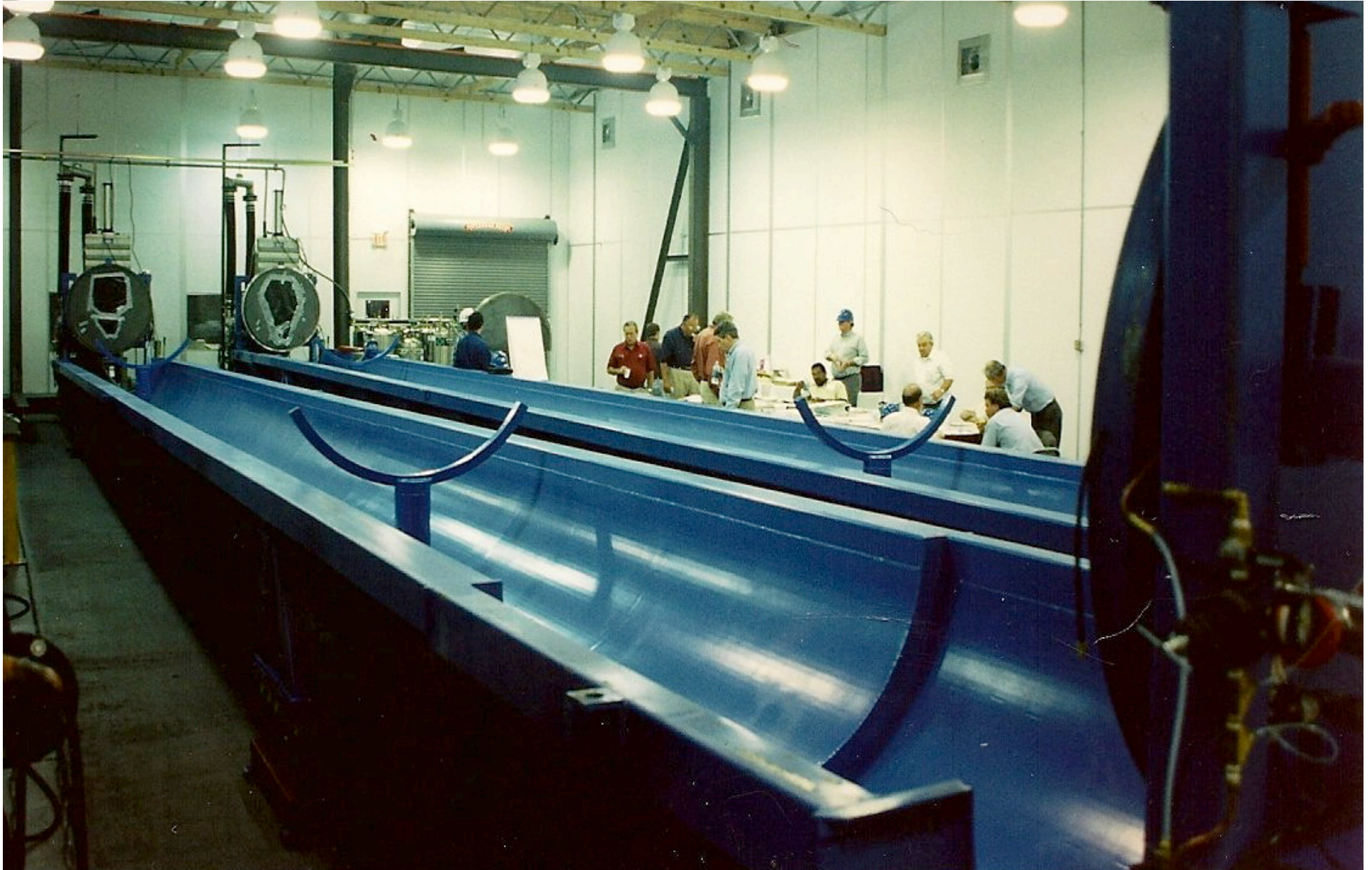


Leak Test "Coffin"





Leak Test "Coffin"



Beam Tube Field Assembly





Tube Assembly

- Field butt weld made in movable shelter
- Internal “dam” shields inside of weld with inert gas
- Dam is later filled with He for leak test
- Finally, garbed worker crawls in to remove dam & place optical baffles

LIGO-G1300170





Welding Shelter

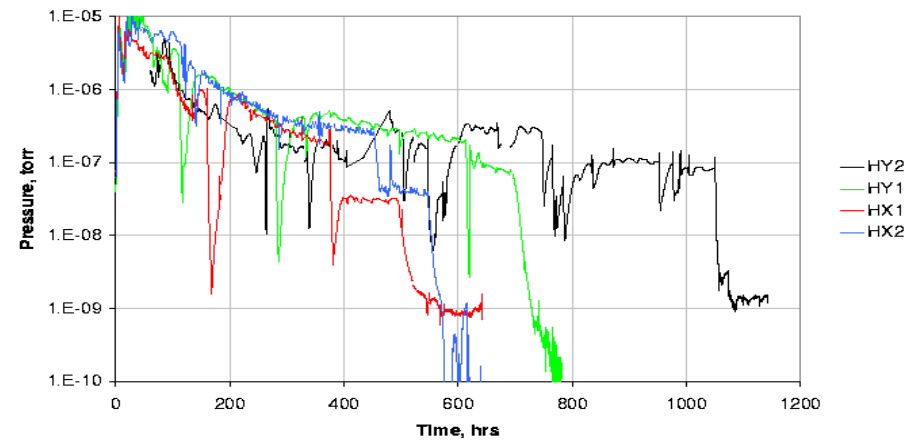
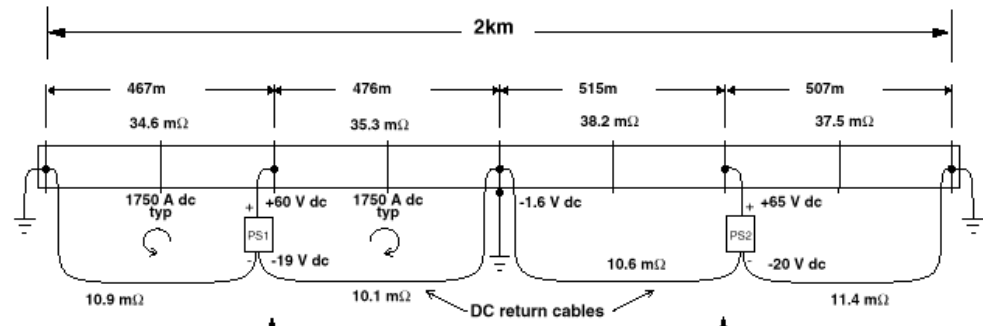




I²R Bakeout to Desorb Water



- Glass wool insulation
- $I_{DC} = 2,000 \text{ A}$
- ~ 3 weeks @ **160°C**
- Final $J_{H_2O} < 2e-17 \text{ TI/s/cm}^2$
- Tubes **never** to be vented





How Often do you get a Second Chance?? Some Updates Under Study for LIGO India

- Move to “conventional” large gate valve design ?
 - » Reduce cost with single O-ring gate seal, metallic bonnet seal?
 - » Increase reliability with “standard product”?
 - » Reduce complexity of two actuator styles
- Remove oxide from stress-relieved vessel walls?
 - » Oxide patina was good for suppressing IR scatter, good for vacuum performance, but began to flake off with age— had to be removed!
 - » Considering conventional treatments such as passivation or electropolish
- Is spiral-welding still the most economical tube construction?
- Are there alternative ways to achieve ultralow dissolved hydrogen?
- Should we upgrade monitoring & instrumentation along tubes?
- Should we consider 316L in place of 304L for higher corrosion resistance?
- Should we consider flanging or other means to reversibly isolate tube sections after installation?



Cautionary Tale: The LLO Y Beamtube Leak

- Discovered only late last year, but...
- Reconstruct $t=0$ in 2008, $F = 2.5e-4$ torr-liter/sec (!)
 - » Unnoticed due to sparse instrumentation, masked by detector outgassing
 - » Approximately $1e5$ x specification; **MUST BE REPAIRED**
 - Legacy of water vapor deposited since '08 may persist, even after repairs
- Localized near Y midpoint by gradient methods
- Followed by He MSLD test: **VERY DIFFICULT ON THIS SIZE SYSTEM !**
- At least **4 distinct leaks** discovered to date, in at least 3 zones
 - » Confirms this is no fluke or isolated defect, but a progressive problem
 - » Most likely a spectrum of sources will be found
 - » Largest breach is now sealed but about $1e-5$ torr-liter/sec remains unaccounted for
- Those discovered to date coincide with
 - » Welds (both a spiral weld and a stiffener fillet), plus
 - » Animal residues (mice or mud wasps; emit "corrosion accelerants"), plus
 - » Local history of persistent water incursion
- Team (incl. yours truly) has been assigned to find, repair, prevent recurrence
 - » Outside metallurgical & welding specialists under contract
 - » Additional diagnostics on known leaks, representative fab samples
 - » Too soon to reliably bound full \$ and schedule impact!



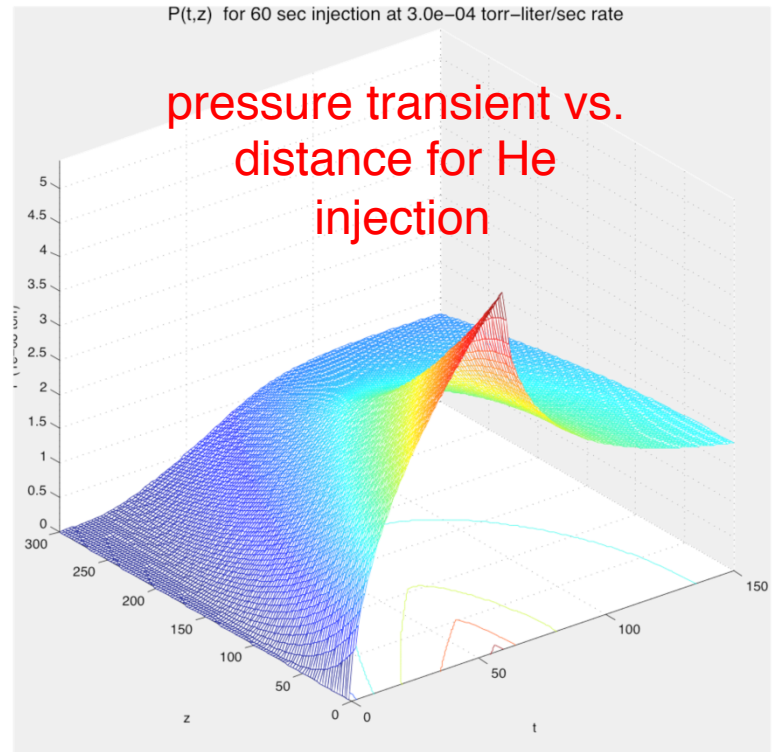
mouse latrine



evacuated temporary cover



mud dauber nest







Final Remarks

- LIGO facilities are among the largest high-vacuum systems ever built, and have stringent requirements
- Novel and cost-effective methods were developed to meet these challenges successfully in the US
- In trouble-free operation over a decade, LIGO is now installing second-generation instruments in Louisiana and Washington
- A third identical interferometer was built and is now designated for India when the site is available
- With benefit of hindsight, new technology, and Indian innovation, we believe the LIGO India vacuum system will be even better!



Thank you IUVSTA!

Thank you IPR!

--Reference Slides--



LIGO

Livingston Observatory



LIGO-G130017



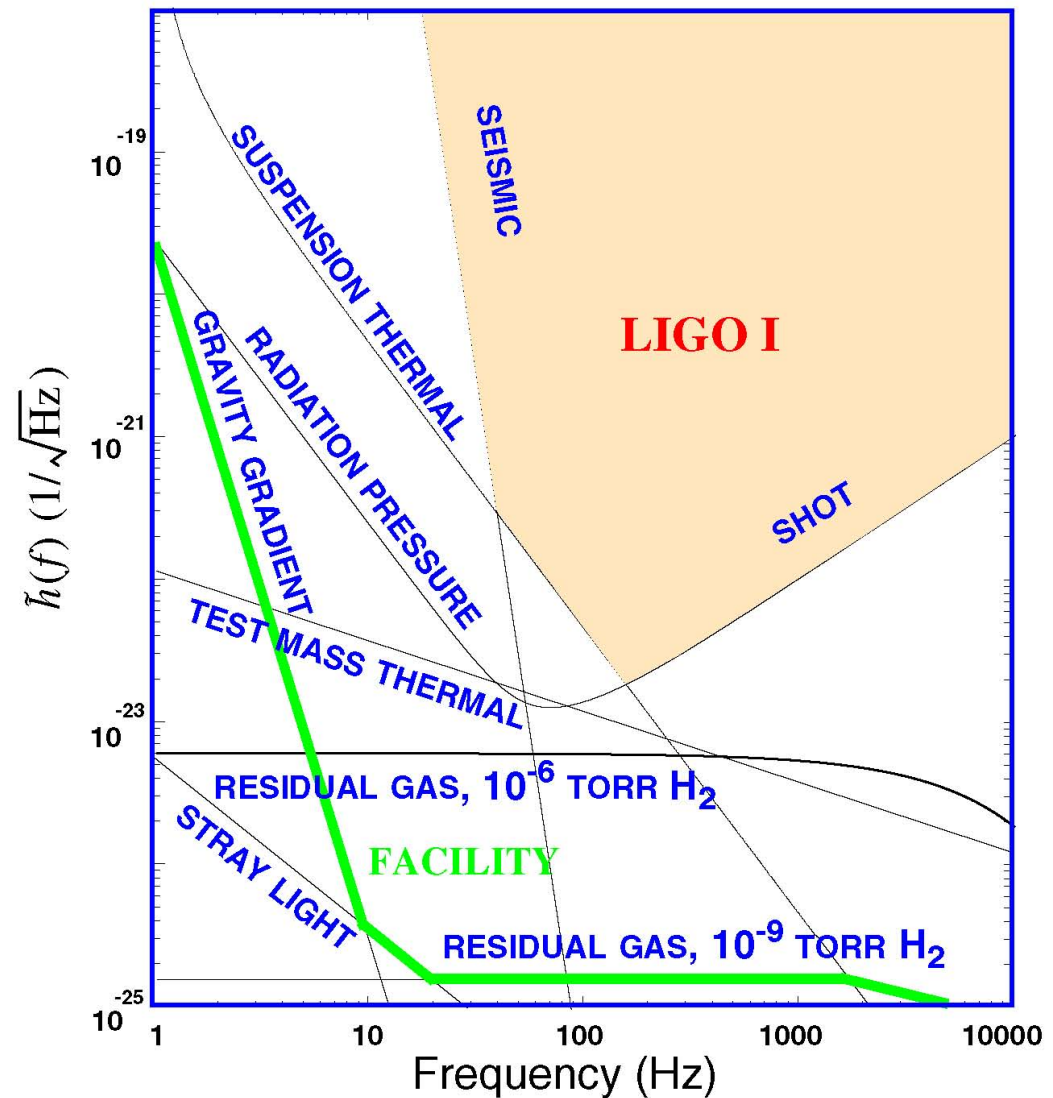
LIGO

Hanford Observatory



Limits to Sensitivity

- First detectors reached design sensitivity in 2005
- Now installing Advanced detectors
- Vacuum requirement
 $<10^{-9}$ torr H_2
 $<10^{-10}$ torr H_2O





Residual Gas Index Fluctuation Noise

$$S_L(f) = \frac{4\rho(2\pi\alpha)^2}{v_0} \int_0^{L_0} \frac{\exp[-2\pi f w(z)/v_0]}{w(z)} dz$$

$$\Delta\tilde{L}(f) \equiv \sqrt{S_{\Delta L}(f)} = \sqrt{2S_L(f)}$$

ρ = gas number density (\sim pressure)

α = optical polarizability (\sim index)

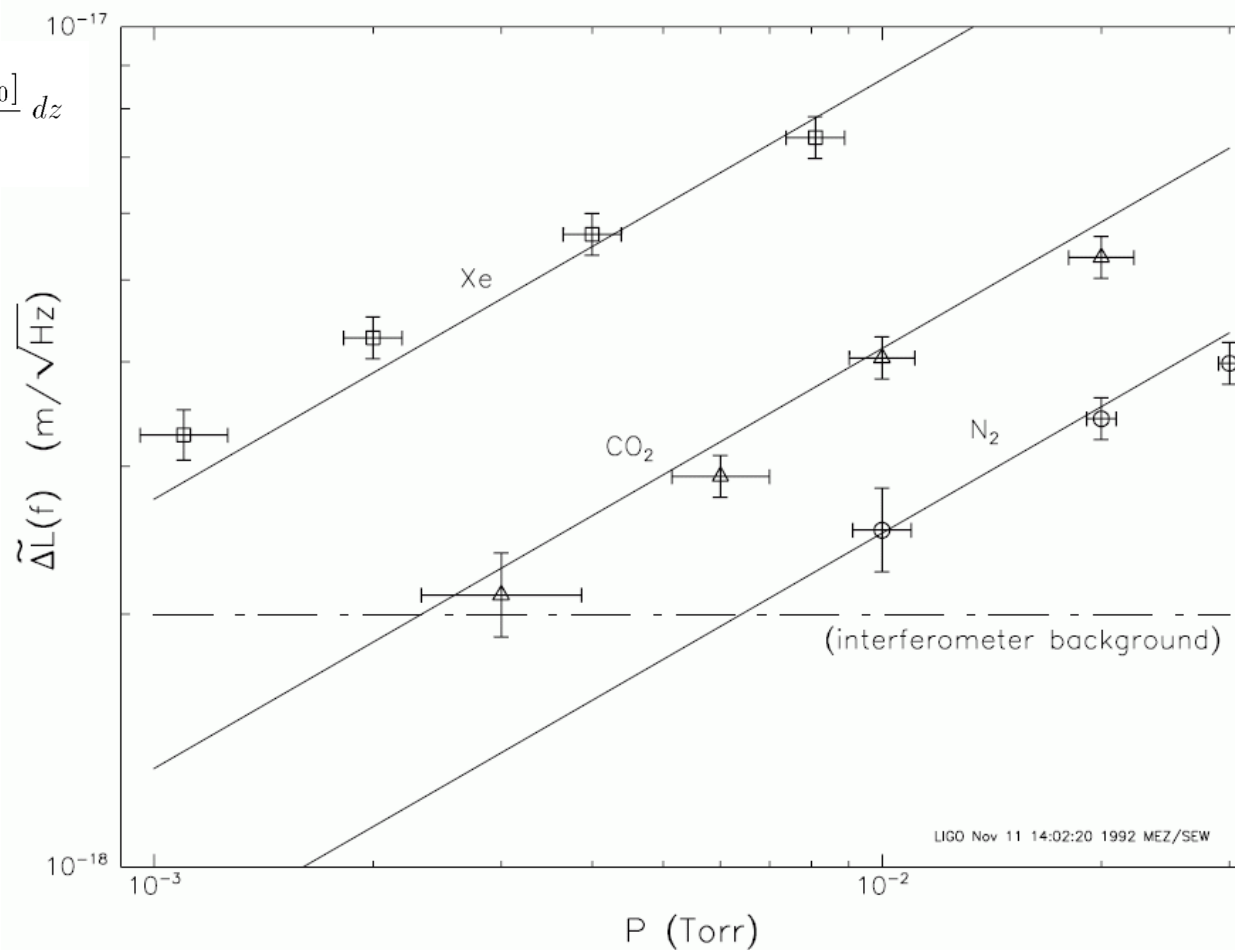
w = beam radius

v_0 = most probable thermal speed

L_0 = arm length

ΔL = arm optical path difference

Statistical model
verified by
interferometer
experiment





Residual Gas Pressure Limits in Beam Tubes

To avoid optical phase noise in laser path

$$h(f) = 4.8 \times 10^{-21} R \left(\frac{x}{H_2} \right) \sqrt{\langle P(\text{torr}) \rangle_L}$$

Table 1: Residual gas phase noise factor and average pressure

Gas Species	R(x/H ₂)	Requirement (torr)	Goal (torr)
H ₂	1.0	1×10 ⁻⁶	1×10 ⁻⁹
H ₂ O	3.3	1×10 ⁻⁷	1×10 ⁻¹⁰
N ₂	4.2	6×10 ⁻⁸	6×10 ⁻¹¹
CO	4.6	5×10 ⁻⁸	5×10 ⁻¹¹
CO ₂	7.1	2×10 ⁻⁸	2×10 ⁻¹¹
CH ₄	5.4	3×10 ⁻⁸	3×10 ⁻¹¹
AMU 100 hydrocarbon	38.4	7.3×10 ⁻¹⁰	7×10 ⁻¹³
AMU 200 hydrocarbon	88.8	1.4×10 ⁻¹⁰	1.4×10 ⁻¹³
AMU 300 hydrocarbon	146	5×10 ⁻¹¹	5×10 ⁻¹⁴
AMU 400 hydrocarbon	208	2.5×10 ⁻¹¹	2.5×10 ⁻¹⁴
AMU 500 hydrocarbon	277	1.4×10 ⁻¹¹	1.4×10 ⁻¹⁴
AMU 600 hydrocarbon	345	9.0×10 ⁻¹²	9.0×10 ⁻¹⁵

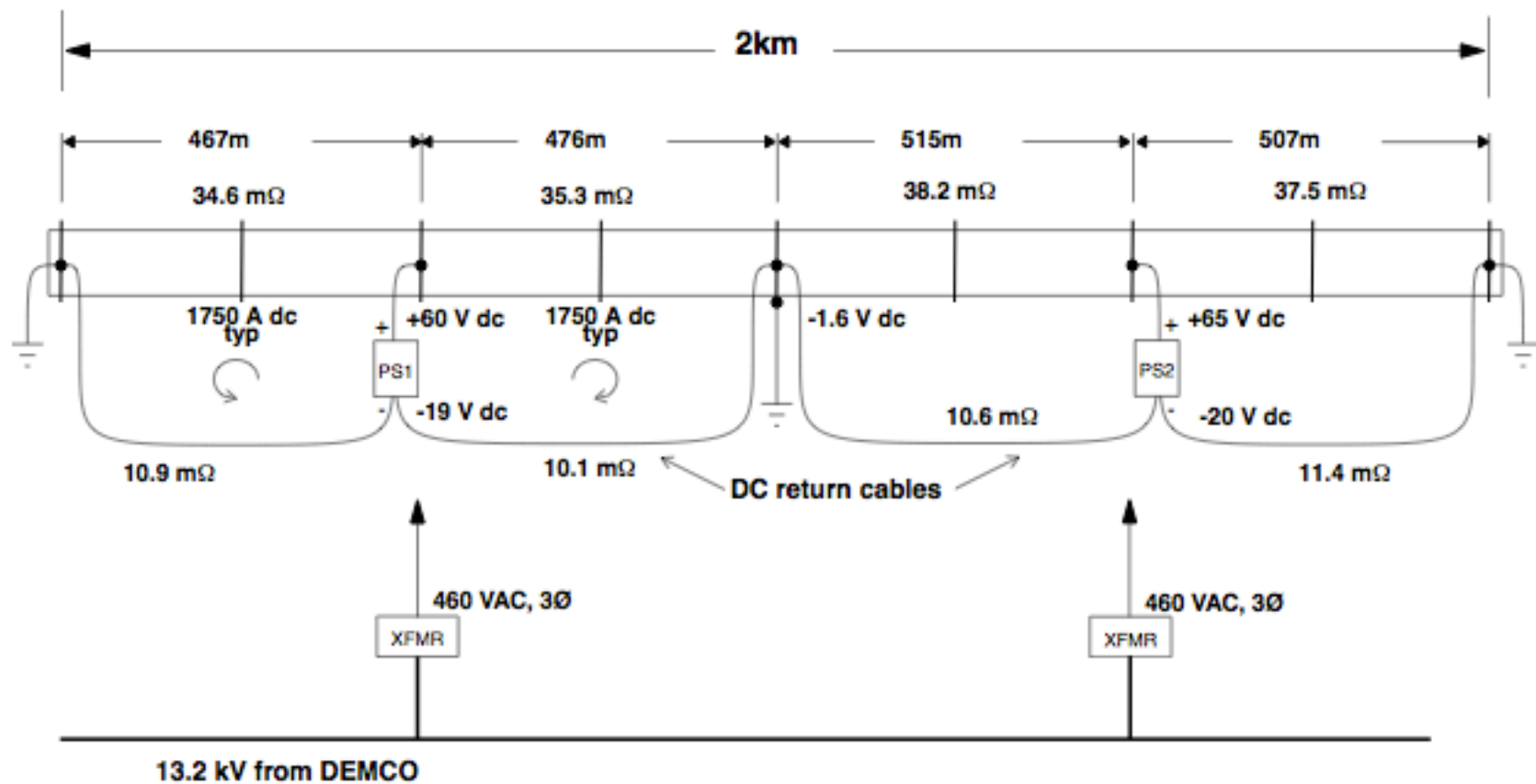


Beam Tube Properties

module length	2 km
25 cm diameter pump ports/module	9
radius of beam tube	62 cm
volume of module	4.831×10^6 liters
area of module	1.55×10^8 cm ²
initial pumping speed/surface area	1.94×10^{-5} liters/sec/cm ²
length/short section	1.90×10^3 cm
wall thickness	3.23×10^{-1} cm
stiffener ring spacing	76 cm
stiffening ring width	4.76×10^{-1} cm
stiffening ring height	4.45 cm
expansion joint wall thickness	2.67×10^{-1} cm
expansion joint convolutions	9
expansion joint longitudinal spring rate	1.5×10^9 dynes/cm



BEAM TUBE BAKEOUT ELECTRICAL HEATING POWER



Legend:

XFMR

Power Transformer

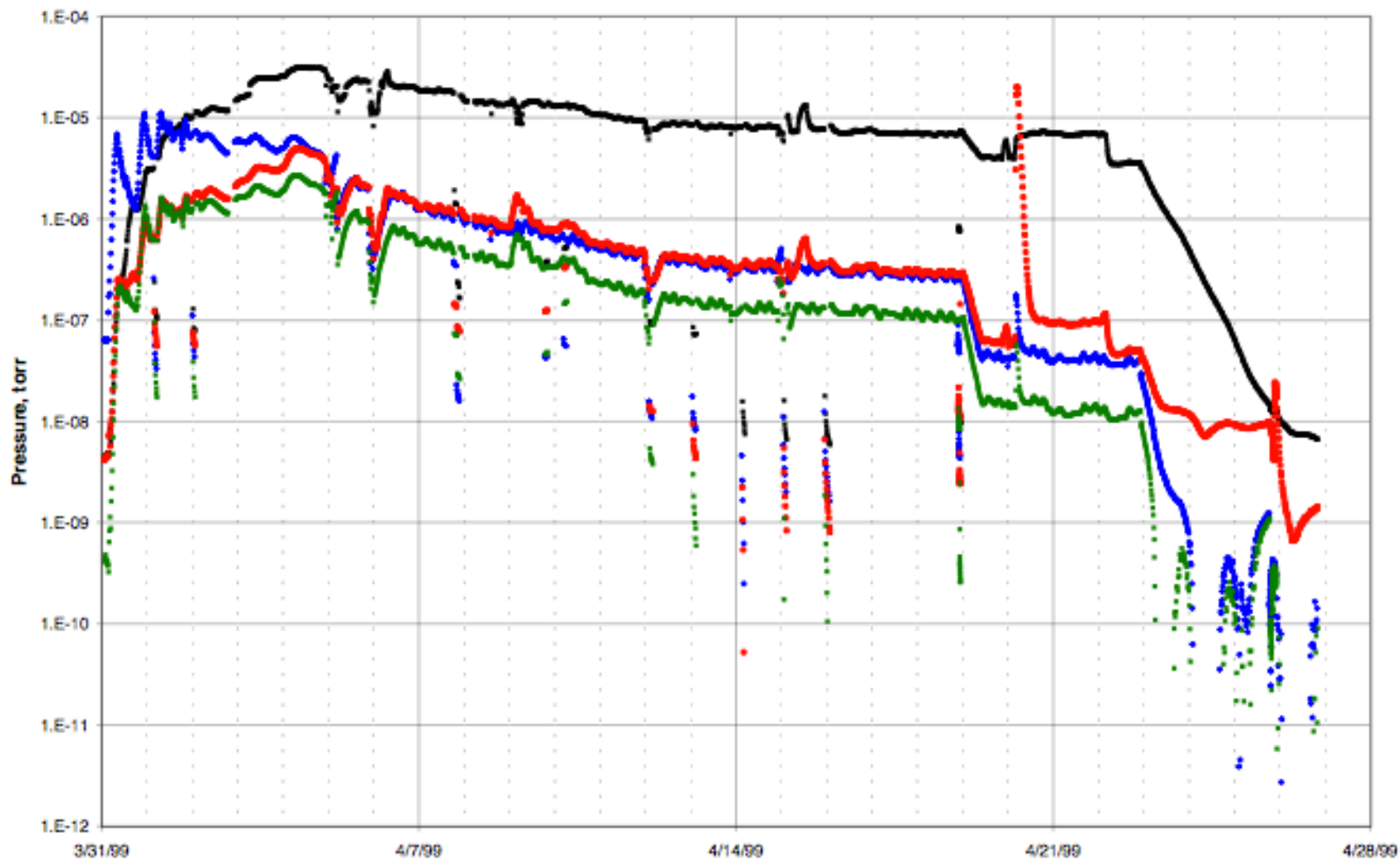
PS

Low voltage, high current
DC power supply



Pressure evolution for major species during 160°C beam tube module bakeout

HX2 RGA PRESSURE, AMU 2 (blk), AMU 18 (blu), AMU 28 (red), AMU 44 (green)





Postbake measurements of module X1 at Hanford

March 11-12, 1999

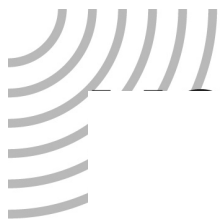
Table 1: Results from gas model solution of 16.9 hour postbake accumulation ending March 12, 1999 at 10:00AM .

molecule	Outgassing rate @ 10C	pressure@ 10C	outgassing rate @ 23C	pressure@ 23C
	torr liters/sec/cm ²	torr	torr liters/sec/cm ²	torr
H ₂	1.6 x 10 ⁻¹⁴	1.0 x 10 ⁻⁹	5.2 x 10 ⁻¹⁴	3.4 x 10 ⁻⁹
CH ₄	< 2 x 10 ⁻²⁰	< 3.4 x 10 ⁻¹³	< 8.8 x 10 ⁻²⁰	< 1.5 x 10 ⁻¹²
H ₂ O	< 3 x 10 ⁻¹⁹	< 5.2 x 10 ⁻¹³	< 1.3 x 10 ⁻¹⁸	< 2.3 x 10 ⁻¹²
N ₂	< 9 x 10 ⁻¹⁹ **	< 1.5x 10 ⁻¹³		
CO	< 1.3 x 10 ⁻¹⁸	< 1.7 x 10 ⁻¹³	< 5.7 x 10 ⁻¹⁸	< 7 x 10 ⁻¹³
O ₂	< 1.2 x 10 ⁻²⁰	< 2.3 x 10 ⁻¹⁴		
A	< 2.5x 10 ⁻²⁰	< 3.6 x 10 ⁻¹⁴		
CO ₂	< 6.5 x 10 ⁻²⁰	< 1.2x 10 ⁻¹³	< 2.9 x 10 ⁻¹⁹	< 5.2 x 10 ⁻¹³
NO+C ₂ H ₆	< 1.5 x 10 ⁻¹⁹	< 1.6 x 10 ⁻¹³	< 6.6x 10 ⁻¹⁹	< 7.2 x 10 ⁻¹³
H _n C _p O _q	∑ amu41,43,55,57 < 1.2 x 10 ⁻¹⁹	< 2.2 x 10 ⁻¹³	∑ amu41,43,55,57 < 5.3 x 10 ⁻¹⁹	< 9.7 x 10 ⁻¹³

Volume = 2.4 x 10⁶ liters and Area = 7.8 x 10⁷ cm²

** The equivalent air leak into the module Q < 3.5x 10⁻¹¹ torr liters/sec from amu 28.

Correction from 10C to 23C uses a binding temperature of 8000K for hydrogen and 10000K for all other molecules



Beam Tube Bakeout Results

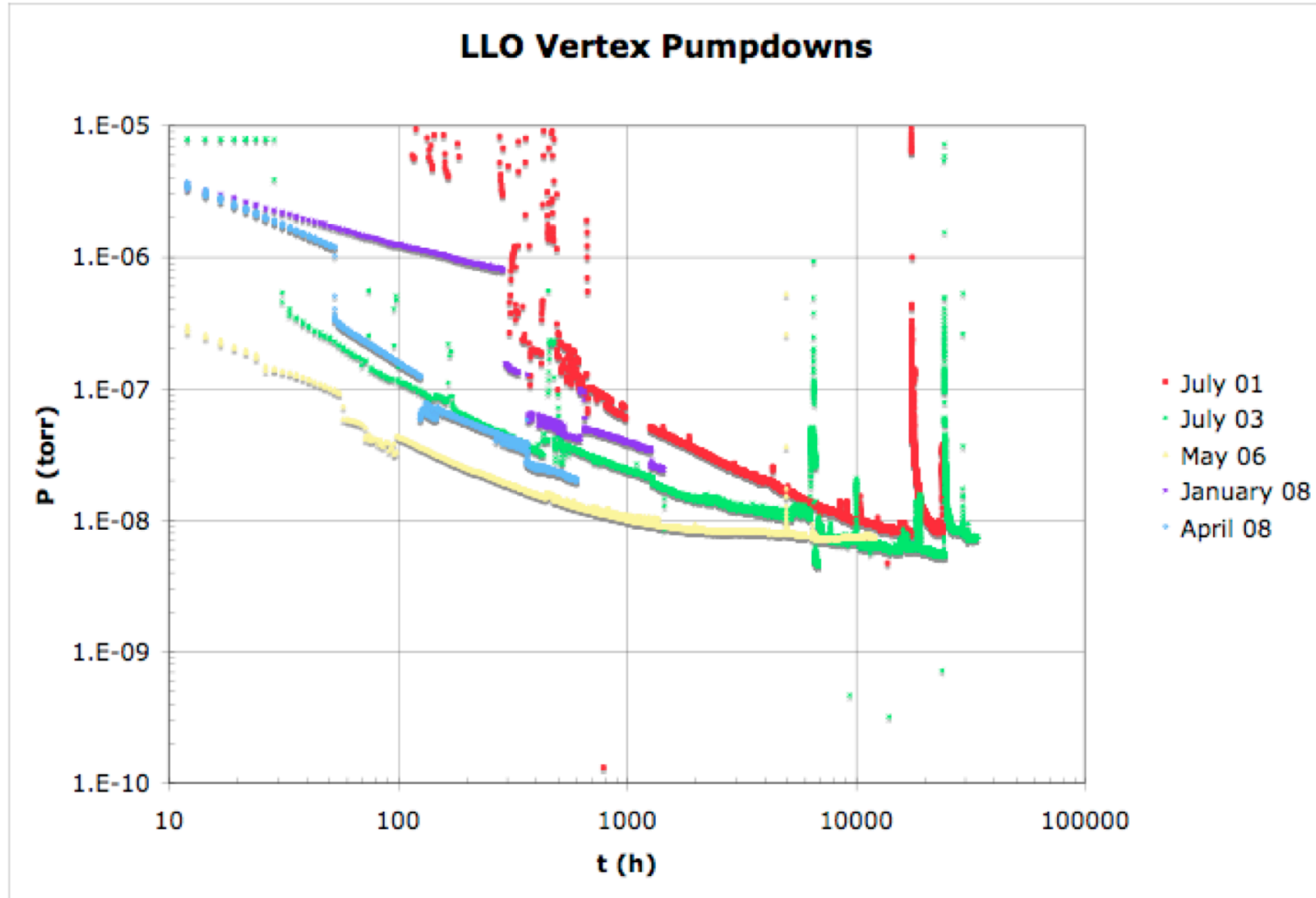
molecule	Outgassing Rate corrected to 23 °C torr liters/sec/cm ² (All except H ₂ are upper limits)					
	Goal*	HY2	HY1	HX1	HX2	
H ₂	4.7	4.8	6.3	5.2	4.6	× 10 ⁻¹⁴
CH ₄	48000	< 900	< 220	< 8.8	< 95	× 10 ⁻²⁰
H ₂ O	1500	< 4	< 20	< 1.8	< 0.8	× 10 ⁻¹⁸
CO	650	< 14	< 9	< 5.7	< 2	× 10 ⁻¹⁸
CO ₂	2200	< 40	< 18	< 2.9	< 8.5	× 10 ⁻¹⁹
NO+C ₂ H ₆	7000	< 2	< 14	< 6.6	< 1.0	× 10 ⁻¹⁹
H _n C _p O _q	50-2 [†]	< 15	< 8.5	< 5.3	< 0.4	× 10 ⁻¹⁹
air leak	1000	< 20	< 10	< 3.5	< 16	× 10 ⁻¹¹ torr liter/sec

*Goal: maximum outgassing to achieve pressure equivalent to 10⁻⁹ torr H₂ using only pumps at stations

[†]Goal for hydrocarbons depends on weight of parent molecule; range given corresponds with 100-300 AMU

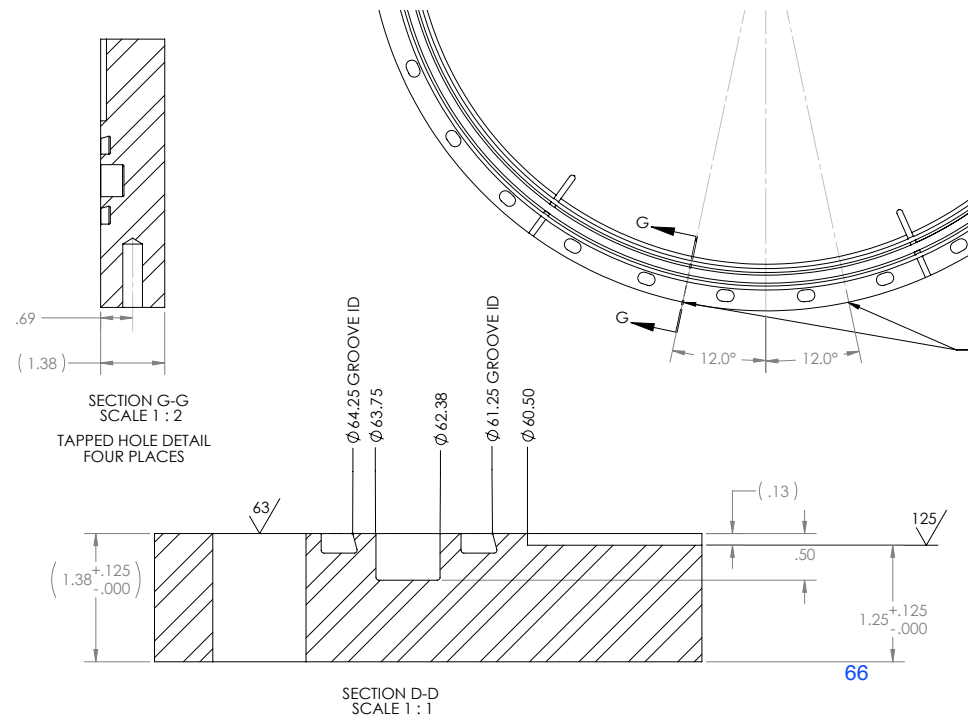
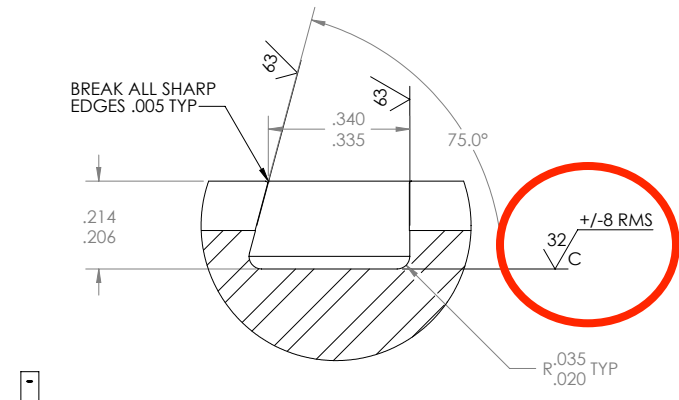


Vertex Pressure Evolution after Backfill



Large Flange Design

- Dual “dry” o-ring
- Pumped annulus between
- (independent of main volume, except in open gate valve)
- Seal faces single-point machined
- Controlled circumferential 32 μ inch “tooth” finish
- Custom Viton (Flourel) cord formulation (55 gallon min. order)
- Cleaned & baked after molding to remove volatile compounds and mold release wax



- All “dry” pumping
- Initial evacuation by blowers & maglev turbomolecular pumps
- Maintained by noble diode ion pumps and coaxial LN₂



LIGO-G1300176



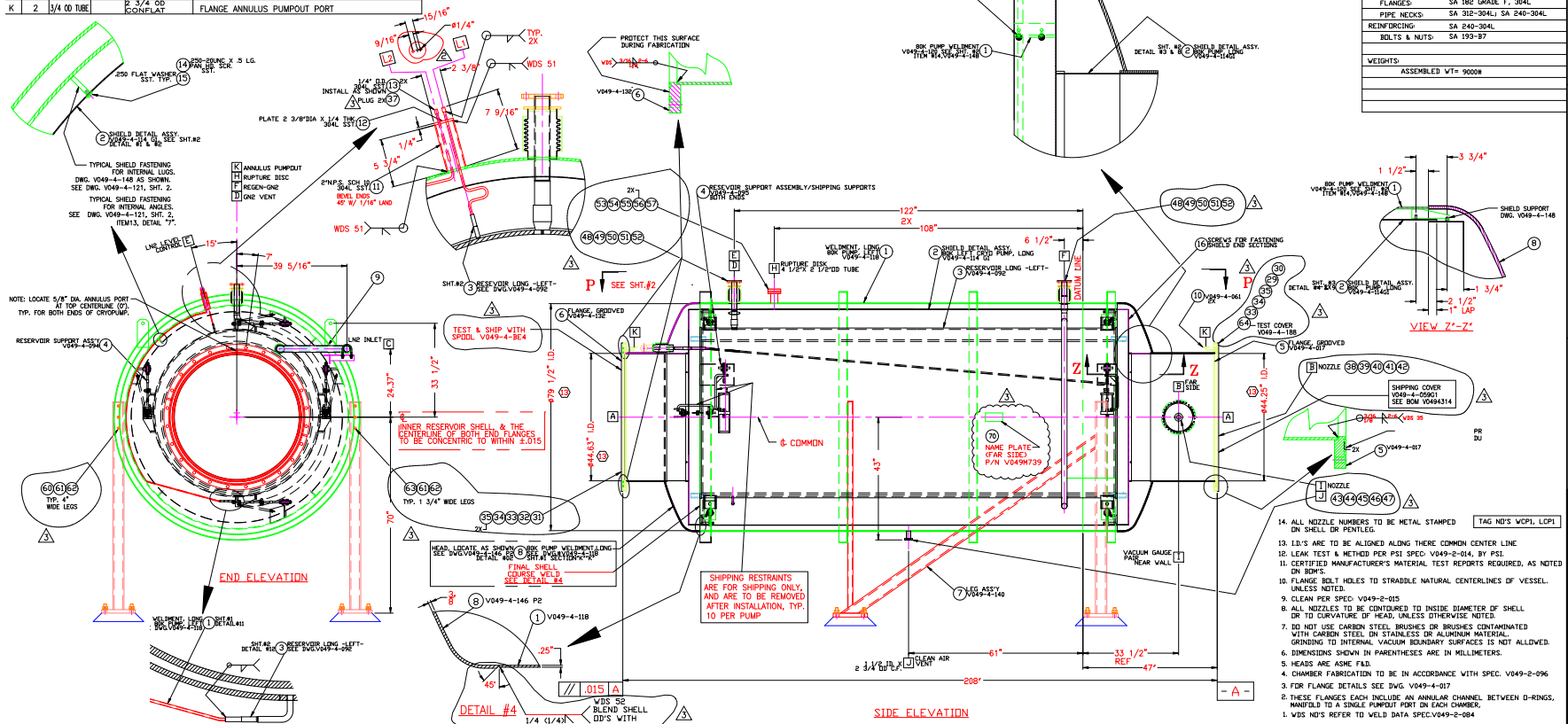


80K LN₂ Cryopump

- Special “low vibration” design
 - » LN₂ reservoir suspended by compliant springs with Flourel dampers
 - » Dual-phase liquid delivery (horizontal vacuum-jacketed lines) maintain continuous liquid and vapor flow without slugging
 - » Sloped “chute” introduces new liquid to reservoir
 - » Reservoir shape & free surface area designed to preclude boiling
 - » Continuous level control (PID with differential pressure level sensor)
- Low-emissivity aluminum cold surface
- Low-emissivity tube liners reduce thermal flux
- Outdoor storage dewars refilled periodically by truck delivery

MARK	QTY	SIZE	RATING	TYPE	DESCRIPTION	MARK	QTY	SIZE	TYPE	DESCRIPTION
A	2	44.63" ID.		SEE NOTE #1	LASER BEAM / ACCESS	L1	1	1/4"	FPT	LN ₂ LEVEL CONTROL
B	1	107.00" TUBE		1/2" OD. CONFLAT W/BLIND FLANGE	PUMPOUT PORT	L2	1	1/4"	FPT	LN ₂ LEVEL CONTROL
C	1	1/2" X 2"		W/BLIND FLANGE						
D	1	1/2" NPT		W/BLIND FLANGE	LN ₂ INLET					
E	1	1/2" NPT		W/BLIND FLANGE	GN ₂ VENT					
F	1	1/2" NPT		W/BLIND FLANGE	SEE L1 & L2 REGEN-GN ₂					
I	1	1/2" OD. TUBE		2 3/4" OD. CONFLAT W/BLIND FLANGE	VACUUM GAUGE PAIR					
J	1	1/2" OD. TUBE		2 3/4" OD. CONFLAT W/BLIND FLANGE	PURGE AIR					
K	1	2 1/2" OD. TUBE		2 3/4" OD. CONFLAT W/BLIND FLANGE	RUPTURE DISC.					
H	2	3/4" OD. TUBE		2 3/4" OD. CONFLAT W/BLIND FLANGE	FLANGE ANNULUS PUMPOUT PORT					

DESIGN DATA	
CORROSION ALLOWANCE:	0
POSTWELD HEAT TREATMENT:	NBS
FIREPROOFING:	NA
RADIOPHOGRAPHING:	NONE
MATERIALS	
HEADS:	SA 240-304L
SHELLS:	SA 240-304L 304L
FLANGES:	SA 182 GRADE F, 304L
PIPE NECKS:	SA 312-304L SA 240-304L
REINFORCING:	SA 240-304L
BOLTS & NUTS:	SA 193-B7
WEIGHTS:	
ASSEMBLED WT:	9000



DWG. NO.	DESCRIPTION
LIGO-G1300176	CRYOPUMP GENERAL ARR'G-T-LONG-LEFT

DWG. NO.	DESCRIPTION
V049-0-100	DRAWING TREE/BDM/STRUCTURE
V049-4-140	SUPPORT LEG ASSY.
V049-4-114 G1	SHIELD ASSY.
V049-4-095	RESERVOIR SUPPORTS
V049-4-118	VESSEL WELDMENT
V049-4-092	BOK PUMP RESERVOIR

SYMBOL	CHARACTERISTIC	UNLESS OTHERWISE SPECIFIED DIMENSIONS ARE IN INCHES UNLESS NOTED:
—	FLATNESS	FRACTIONAL # 11
○	CYLINDRICITY	MEASUREMENT 90-90-90 45°
∥	PARALLELISM	THREE PLACE DECIMAL, # 105
⊥	PERPENDICULARITY	THREE PLACE DECIMAL, # 105
∠	ANGULARITY	BREAK CHANGES IN OUT-RENDERING ALL DIMENSIONS
⊕	TRUE POSITION	USED D ₉₉
⊙	CONCENTRICITY	NEXT ASSY.

REV	DESCRIPTION	ISSUE DESCRIPTION
3	REVISED FOR FABRICATION AS NOTED	DHW PEF PC RDC REC PV 5/19/97 0431
2	REVISED FOR FABRICATION AS NOTED	DHW PEF PC RDC REC PV 1/17/97 0400
1	ISSUED FOR FABRICATION/FR UPDATE - MAJOR REVISIONS	DHW PEF GS RDC REC PV 12/30/96 0346
0	ISSUED FOR FR	REC DMW TDV 7/2/96 0144

PROCESS SYSTEMS INTERNATIONAL, INC.
 25 WALTON DR., WESTBOROUGH, MASSACHUSETTS 01581 USA
BOK PUMP GENERAL ARR'G-T-LONG-LEFT
LIGO VACUUM EQUIPMENT
 4940451
 V049-4-004
 SHEET 1 OF 2