

Mirror quality and interferometer performance Hiro Yamamoto - Caltech LIGO

- Introduction
- Gaussian beam and Modal model
- Cavity basic
- PSD, BRDF and loss
- Mirror polishing and coating
- Beam splitter
- Thermal effects
- Stable cavity







Gaussian beam and Modal model

Gaussian beam : stationary state in a two mirror cavity (FP)

$$R_{00}(x,y,z,t) = G_{00}(x,y,z) \exp[i(\omega \cdot t - k \cdot z)]$$

$$G_{00}(x,y,z,t) = \sqrt{\frac{2}{\pi}} \frac{1}{w(z)} \exp(-r^{2}(\frac{1}{w(z)^{2}} + i\frac{k}{2R(z)}) + i \cdot \eta(z))$$

$$w(z)^{2} = w_{0}^{2}(1 + \frac{z^{2}}{z_{0}^{2}}), \quad R(z) = z + \frac{z_{0}^{2}}{z}, \quad \eta(z) = a \tan(\frac{z}{z_{0}})$$

$$HG_{mn} = G_{00}(x, y, z, t) \sqrt{\frac{1}{2^{m+n}m!n!}} H_m(\frac{\sqrt{2}x}{w(z)}) H_n(\frac{\sqrt{2}y}{w(z)}) \exp[i(m+n)\eta(z)]$$
$$LG_{pm} = G_{00}(x, y, z, t) \sqrt{\frac{p!}{(p+|m|)!}} \exp(im\varphi) L_p^{|m|}(\frac{2r^2}{w(z)^2}) \exp[i(2p+|m|)\eta(z)]$$

LIGO-G1300120-v2 JGW-G1301555-v2

Gaussian beam and Modal model

Tilt

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L is the distance to waist from the reflection point

$$E_{ref}(\theta) = G_{00}(\theta = 0) \cdot \exp(i\omega t - i(x \cdot k_x + z \cdot k_z))$$

$$\sim G_{00} \cdot (1 - i x \cdot k \cdot \theta)$$

$$= G_{00} \cdot (1 - i \cdot \frac{1}{\sqrt{2}} H_1(\frac{\sqrt{2}x}{w(z)}) \cdot \frac{\theta}{\Theta(L)} exp[i(\eta(z) - \eta(L)$$

$$= G_{00} - i \frac{\theta}{\Theta(L)} \cdot G_{10}$$
$$\Theta(z) = \frac{1}{\pi} \frac{\lambda}{w(z)}, \quad H_1(x) = 2x$$

 curvature mismatch



$$E_{ref}(\delta R) = G_{00}(R = R_{in}) \cdot Exp[-ikr^{2}(-\frac{2}{2R_{m}})]$$

$$= G_{00}(R = \infty) \cdot Exp[ikr^{2}(\frac{1}{2R_{in}} - (\frac{1}{R_{in}} - \frac{1}{R_{m}}))]$$

$$\approx G_{00}(R = \infty) \cdot Exp[ir^{2}(\frac{1}{2R_{in}})] \cdot (1 - ikr^{2}\frac{\delta R}{R_{in}^{2}})$$

$$L))]) = G_{00}(R = -R_{in})(1 - ik\frac{w^{2}}{2}\frac{\delta R}{R_{in}^{2}}(1 - L_{1}^{0}(\frac{2r^{2}}{w^{2}})))$$

$$\approx G_{00}(R = -R_{in}) + i\pi \frac{w^2}{\lambda R_{in}} \frac{\delta R}{R_{in}} LG_1^0(R = -R_{in})$$

$$\delta R = R_m - R_{in}, \ L_1^0(r) = 1 - r$$

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Gaussian beam and Modal model
• Cavity field and Gouy phase

$$E_{1} = E_{0} \cdot r_{1}r_{2} \cdot \exp[i2\phi]$$

$$\phi = \begin{cases} m+n+1\\ 2p+1m+1 \end{cases} \Delta \eta - kL$$

$$\Delta \eta = \eta(z_{2}) - \eta(-z_{1}) = a\cos(1 - \frac{L}{R}) \text{ for } R_{1} = R_{2} = R$$

- Resonance condition
 - » $\phi = n\pi$ for main mode
 - » non resonant for other modes



Cavity basic Resonant vs non-resonant





JGW-G1301555-v2





LIGO Surface structure with different spatial distribution D160 71-6 Removed (ဥာက္က) LASTI, LEM-1 1 Rms=0.056nmPTV=0.364nm 60 40 30 20 -1 10 -2 -5 X (mm) D 160mm 5mm 10mm Fizeau IFO Phase Measuring Microscope **Integrating Sphere** ،۸۸۸۸ MMMMMM, WWWWWW

LIGO-G1300120-v2 Get-together meeting of JGW March 1st, 2013 JGW-G1301555-v2



Scattering by aberration

$$\begin{split} E_{ref} &= E_{ref}^{0} \cdot \exp(i2k\delta(x,y)) & dP = \iint dx \, dy \left| E_{ref}^{0} \right|^{2} 4k^{2} \delta(x,y)^{2} \\ &= E_{ref}^{0} \cdot (1+i2k\delta-2(k\delta)^{2}) \\ &= E_{ref}^{0} \cdot (1-2(k\delta)^{2}) + E_{ref}^{0} \cdot i2k\delta \\ &= \int dx \, dy \delta(x,y)^{2} / S \\ &= \int df \, PSD_{1D}(f) \\ for \, \delta(x,y) &= \delta_{0} \sin(k_{x}x) \\ &= E_{ref}^{0} \cdot i2k\delta_{0} \sin(k_{x}x) \\ &= E_{ref}^{0} \cdot i2k\delta_{0} \sin(k_{x}x) \\ &= E_{ref}^{0} \cdot (k\delta_{0}(\exp(ik_{x}x) - \exp(-ik_{x}x))) \\ &= E^{0} \frac{2\pi\delta_{0}}{\lambda} [\exp(-i(kz - k_{x}x)) - \exp(-i(kz + k_{x}x))] \\ &= E_{ref}^{0} \frac{2\pi\delta_{0}}{\lambda} [\exp(-i(kz - k_{x}x)) - \exp(-i(kz + k_{x}x))] \end{split}$$

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Peeking at LIGO mirror profile





15

100

 Composite - Poly Fit 18

0.0001

Polishing and coating ETM04 : coating is tough

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Polishing by Coastline and ASML Requirement and result of ITM04

Surface Specification **Specification Value Actual Value** Pass/Fail Location Parameter Spherical, CC, RoC 1 Central 160 mm 1934 m -1938.61 m PASS 5m/+15m Radius Difference Central 160 mm 1938.53 m ± 3 m 0.08 m PASS from all ITMs Astigmatism Central 160 mm 0.12 nm PASS $\sigma_{RMS} < 3 \text{ nm}$ Amplitude (Z₂₂) Figure Error (LSF) Central 300 mm PASS $\sigma_{RMS} < 2.5 \text{ nm}$ 0.37 nm <1mm-1 0.15 nm PASS Central 160mm Z_{0.0} · Z_{1.1} · Z_{2.0} · Z_{2.2} Fit $\sigma_{RMS} < 0.3 \text{ nm}$ PASS 0.137 nm Error (HSF) Center, Ø60 σ_{RMS}≤ 0.16 nm mm, Ø120 mm 1-750mm-1



Coating by LMA ITM04 and ITM08







LMA ETM01 coating accepting test short wavelength spiral pattern



- Any other effects
 - » Field aberration due to this pattern
 - Field in FP with this map Field in idealistic FP[®]
 - Very fine grid sizes to make sure FFT is OK
 - » Mode analysis if any mode could dominate
 - No dominant mode for LGpm (2p+m<25) and HGmn (m+n<25)
 - » If ITM has similar pattern, can they interfere
 - ITM = MAPPING (DATAFILE("ETM01pattern.dat"), "-x","y") * 0.5
 - Loss = loss by ETM + loss by ITM no additional by interference



spiral pattern on ET

20



Hiro Yamamoto LCGT F2F mtg @ ICRR on August 4, 2011

LIGO-G1100857 JGW-G1100517



LMA ETM01 coating accepting test long wavelength central plateau

Old coating system, one at a time

- » The beam size on ETM is larger than that on ITM and the plateau size on ETM needs to be 20% wider, when coating to coating variation is taken into account
- New coating using the planetary system, a pair at a time
 - » Higher order mode, mostly LG20, in the FP cavity is ~100ppm
 - Better than old, 120ppm, and two ETMs will be "identical", but is this good enough?
 - The plateau size is around the same as the old one
 - Astigmatism uncertainty due to the substrate is not a major issue
 - Asymmetry in the far outside is better (smaller) in the new coating
 - » Coupled cavity simulation
 - LG20 in SRC shows no increase of LG20 by the mode healing
 - Stable signal recycling cavity kills LG20 in SRC
 - LG20 in PRC is ~2000ppm increase by the ETM coating aberration

Thermal distortion test mass



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Not so nice looking mirrors





Higher order mode fraction on SRM

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