

Results from TOBAs

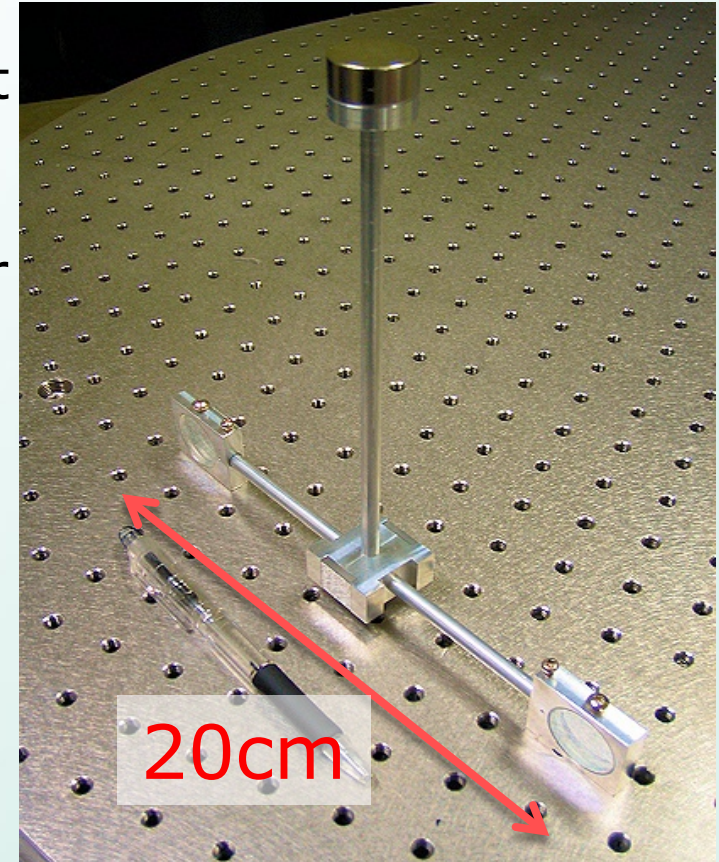
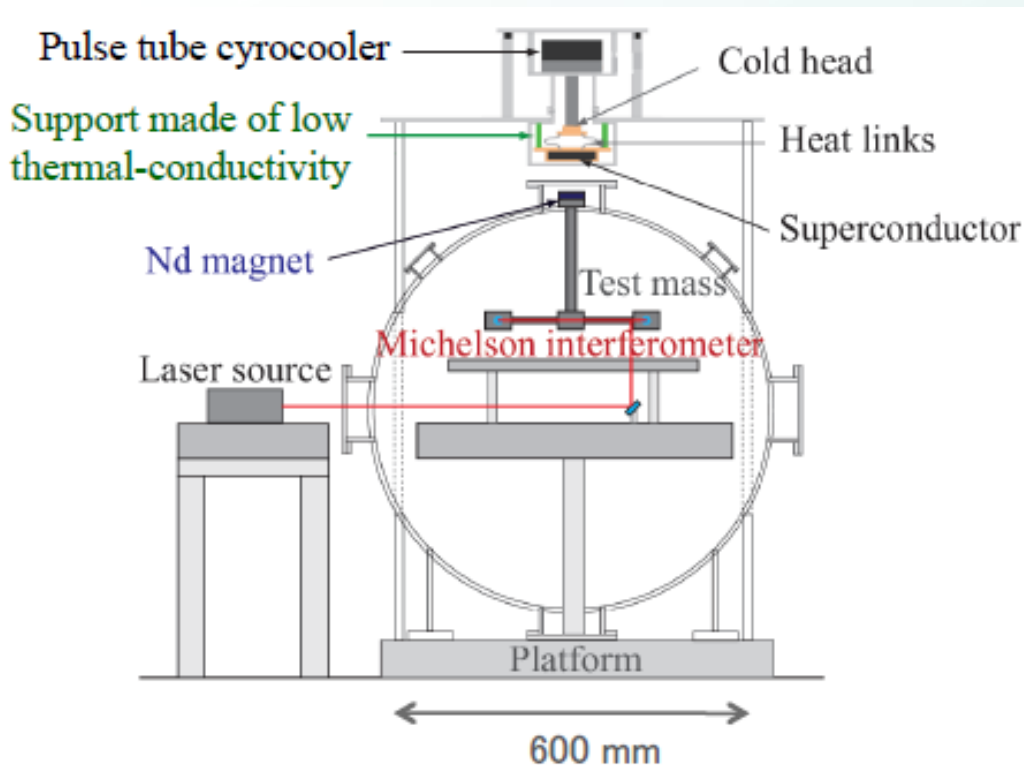
Cross correlation analysis to search for
a Stochastic Gravitational Wave Background

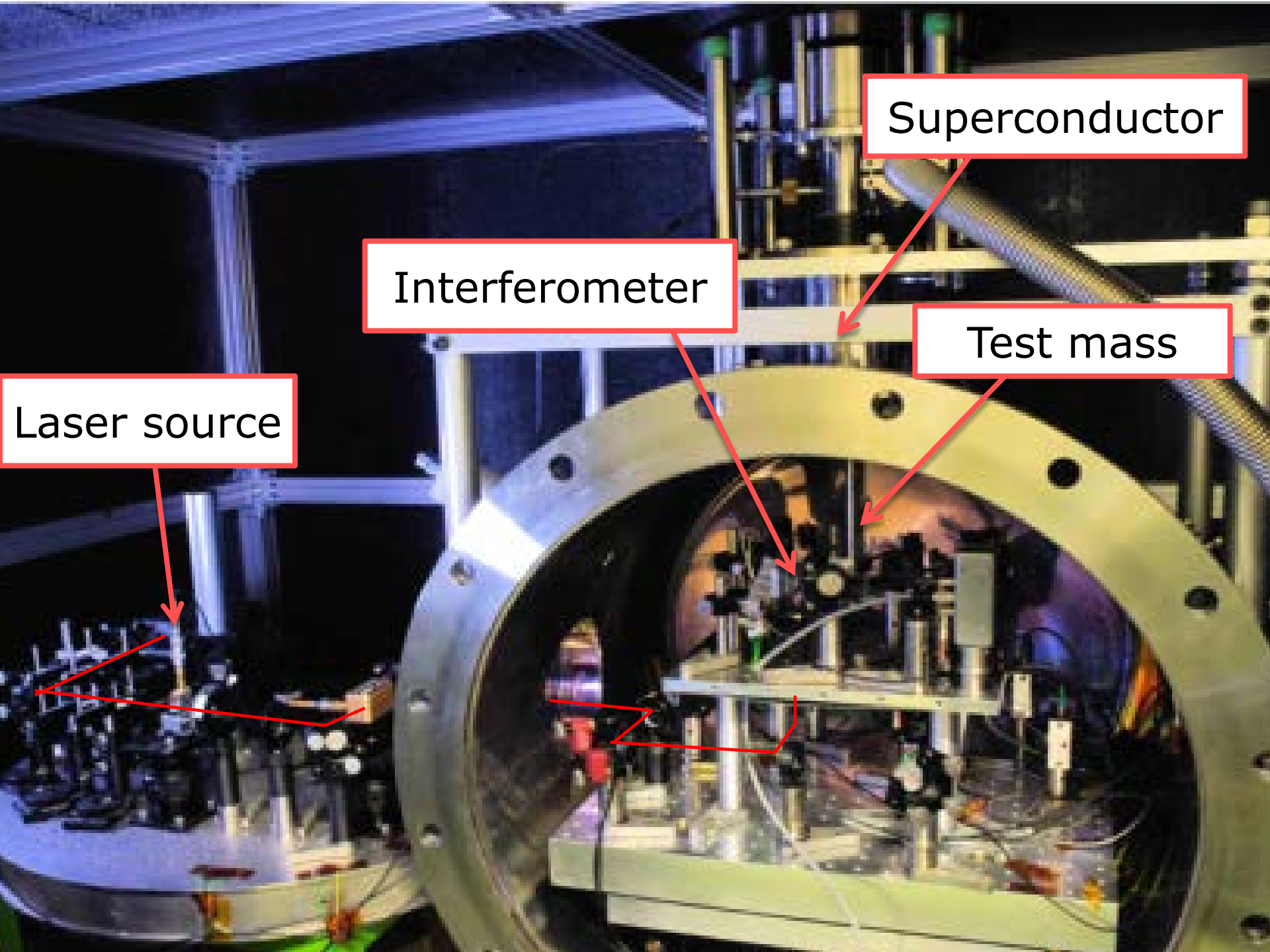
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Prototype TOBA

- 20-cm small torsion bar
- Suspended by the flux pinning effect of the superconductor
- Rotation monitor:
laser Michelson interferometer
- Actuator: coil-magnet actuator





Superconductor

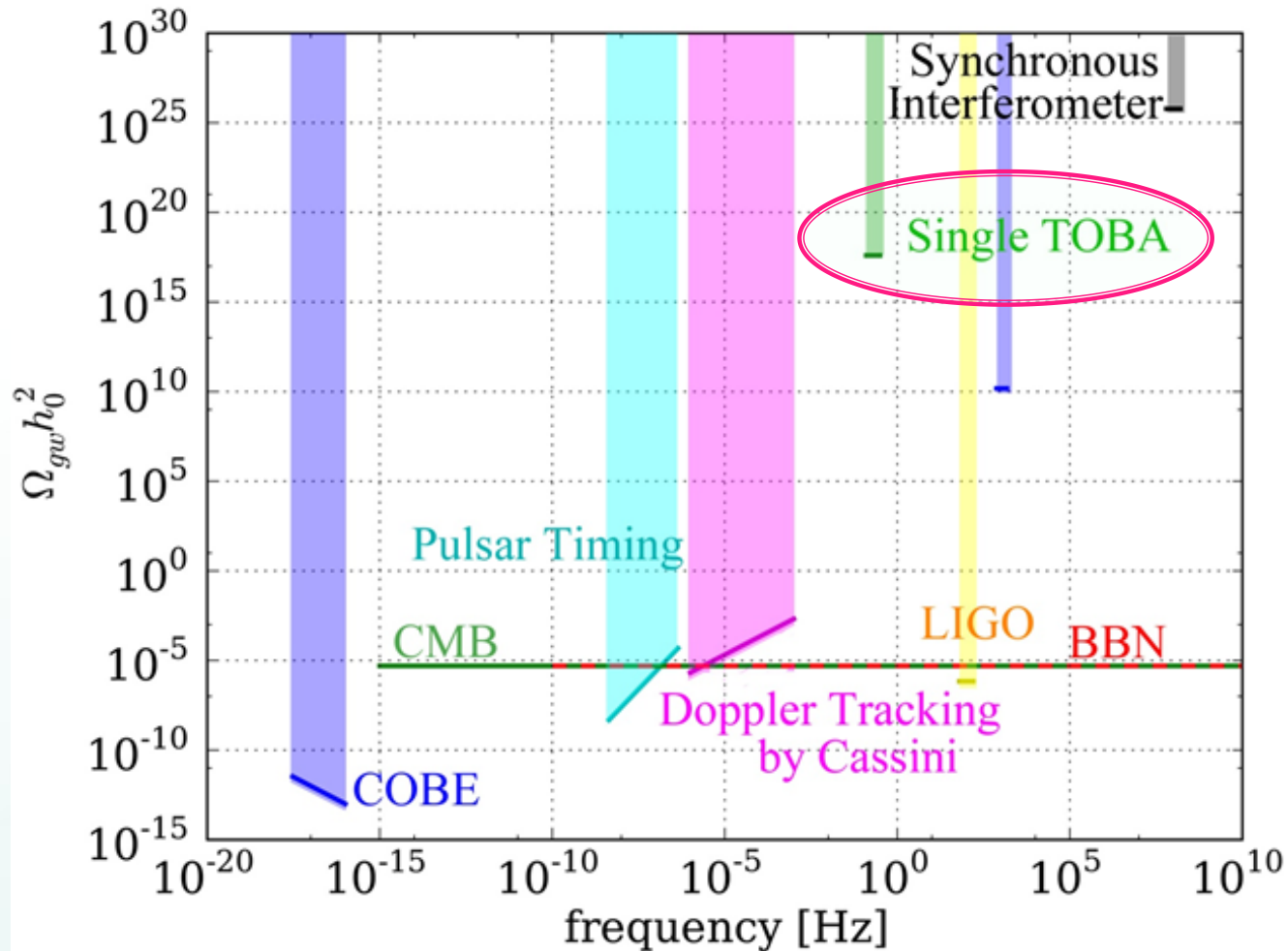
Interferometer

Test mass

Laser source

Previous Result

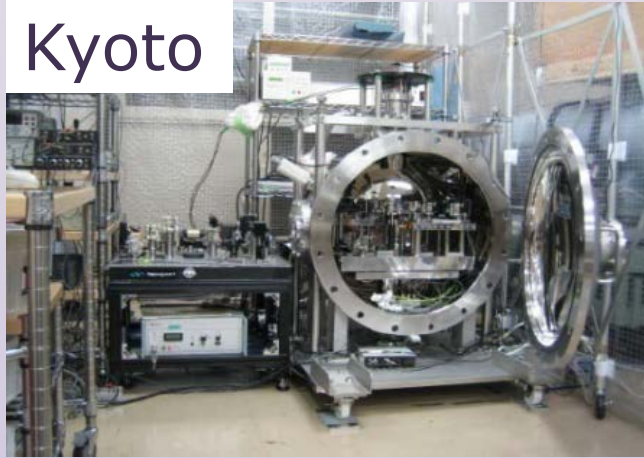
Upper limit on a stochastic GW background



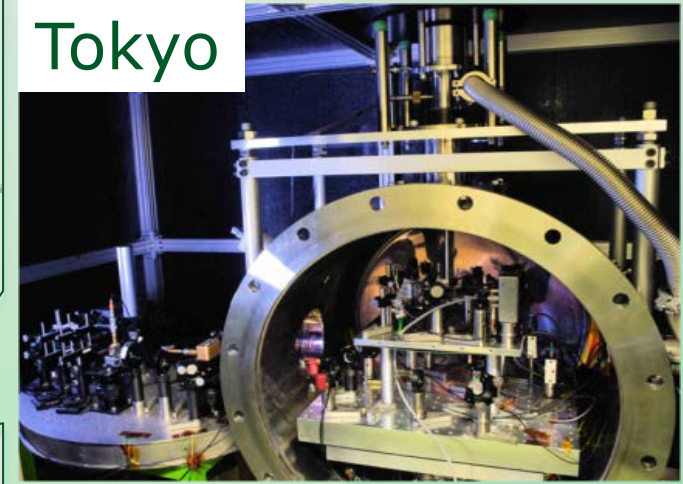
For the detection of a stochastic gravitational-wave background, simultaneous observation is necessary.

Simultaneous Observation

Kyoto



Tokyo



Kyoto

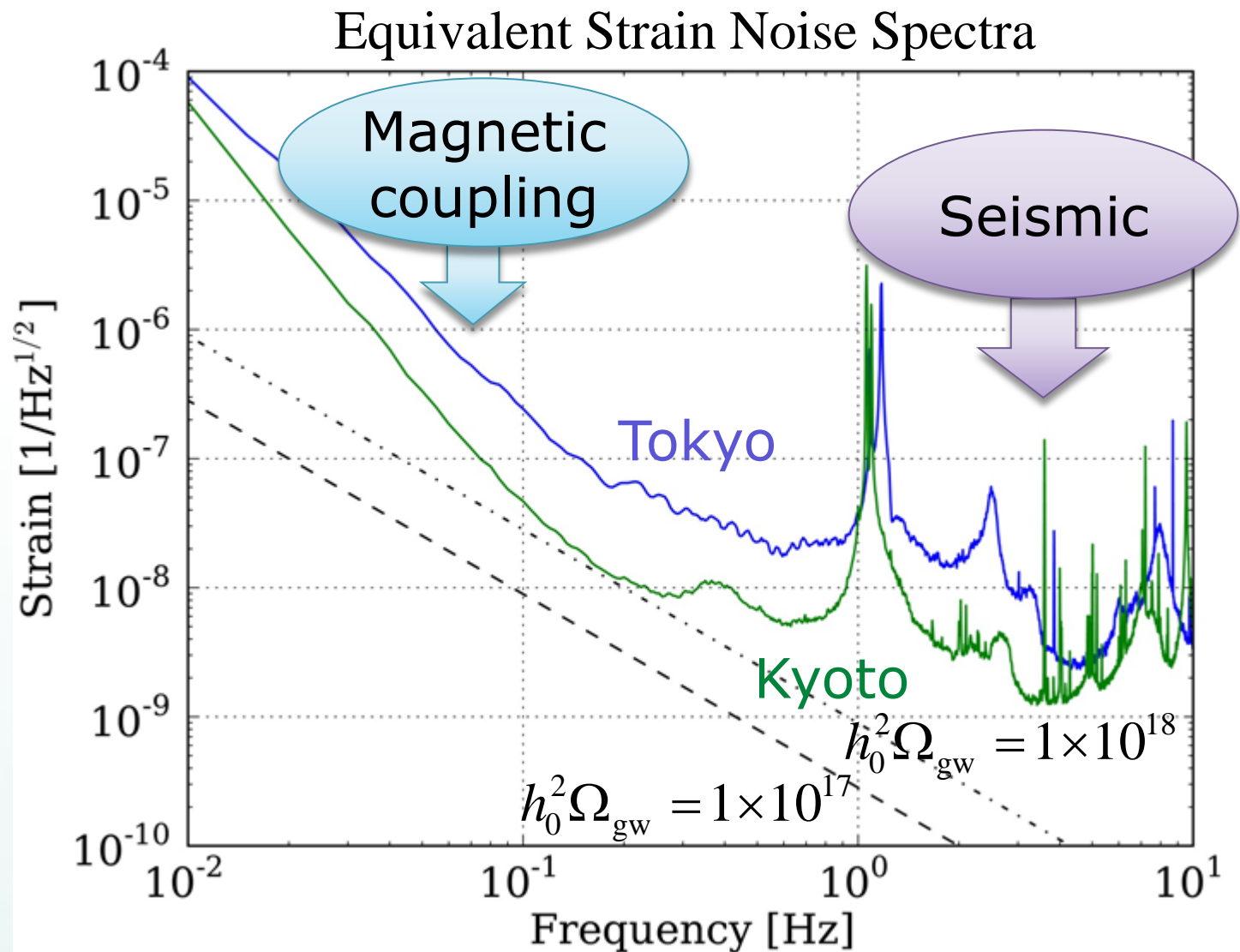
Tokyo

~370km

DATE: 21:30 – 7:30, Oct. 29, 2011

Sampling frequency: 500 Hz,
the direction of the test mass:
north-south

Data Quality



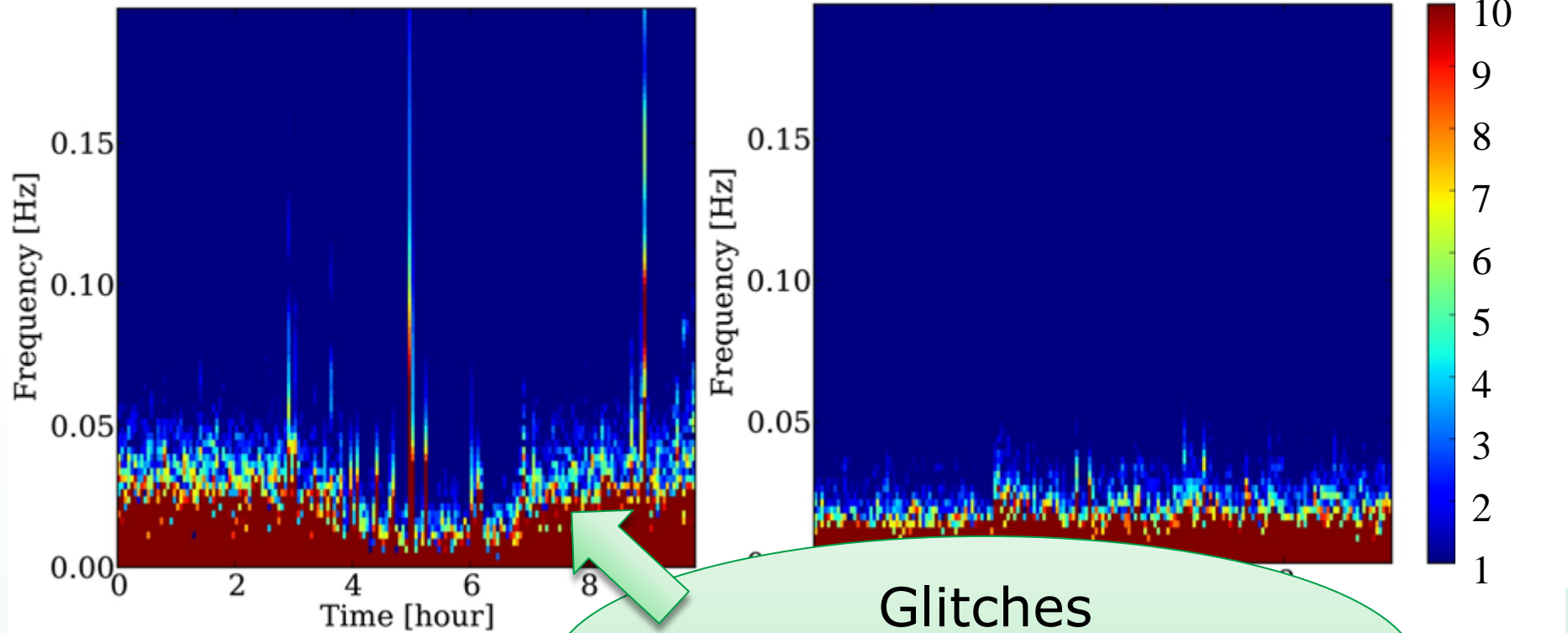
Data Quality

Spectrogram

Tokyo

Kyoto

$\times 10^{-6}$



Cross Correlation Analysis

Concept

difficult to predict the waveform of a stochastic GW background



**Search the coherent signal
on the data of the two detectors.**

$$\text{Correlation Value} = \int_{f_{\min}}^{f_{\max}} df \tilde{s}_1^*(f) \tilde{Q}(f) \tilde{s}_2(f)$$

$\tilde{s}_i(f)$: The signal of i -th detector
 $\tilde{Q}(f)$: The optimal filter
(Weighting function)

Cross Correlation Analysis

Data selection

- Divide the time series data into 200sec long segments
- Delete 10% of the segments whose noise level is worst

Calculate the cross correlation value

- Choose the analyzed frequency band as 0.035 – 0.84 Hz

Detection test

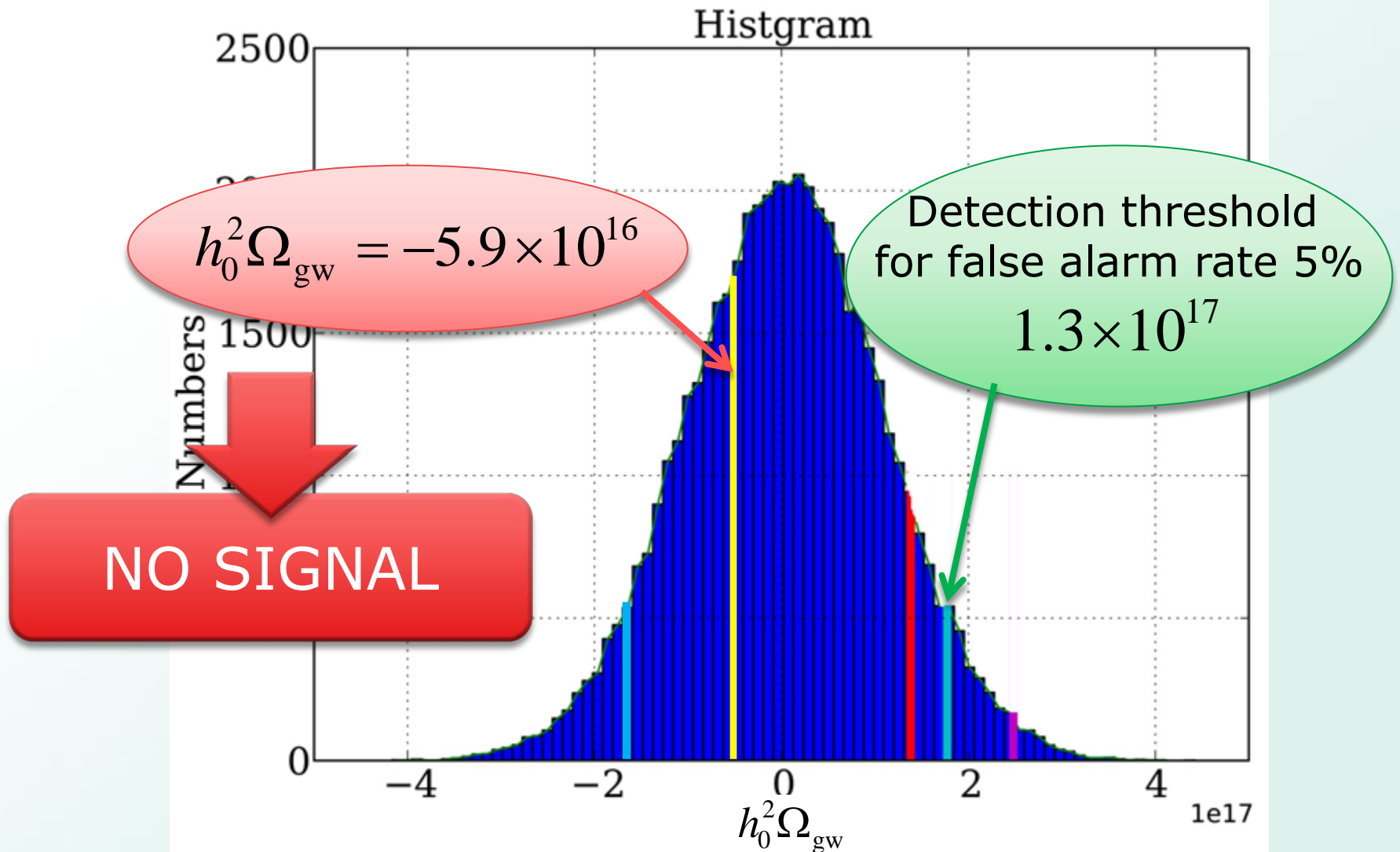
If not detected

Set the upper limit

- Inject mock signals into the real data and calculate the detection efficiency

Result

Histogram and Cross correlation value

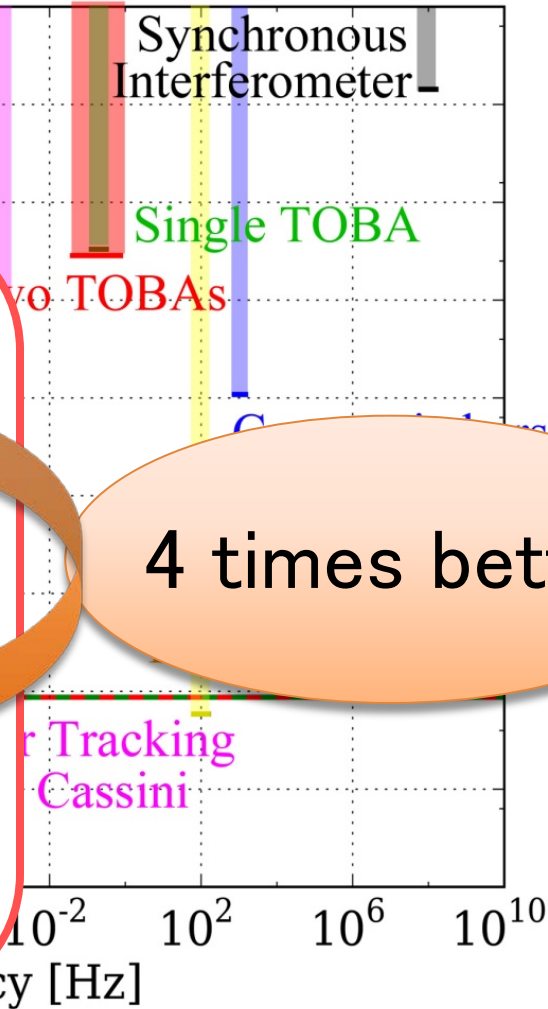
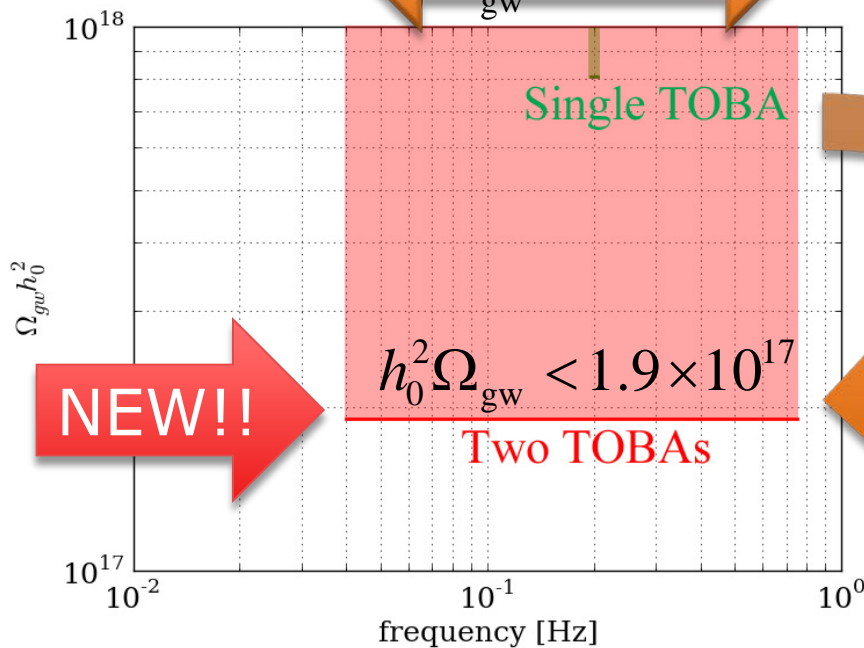


Result – upper limit

95 % confidence upper limit

$$h_0^2 \Omega_{\text{gw}} < 1.9 \times 10^{17}$$

Extend explored frequency band



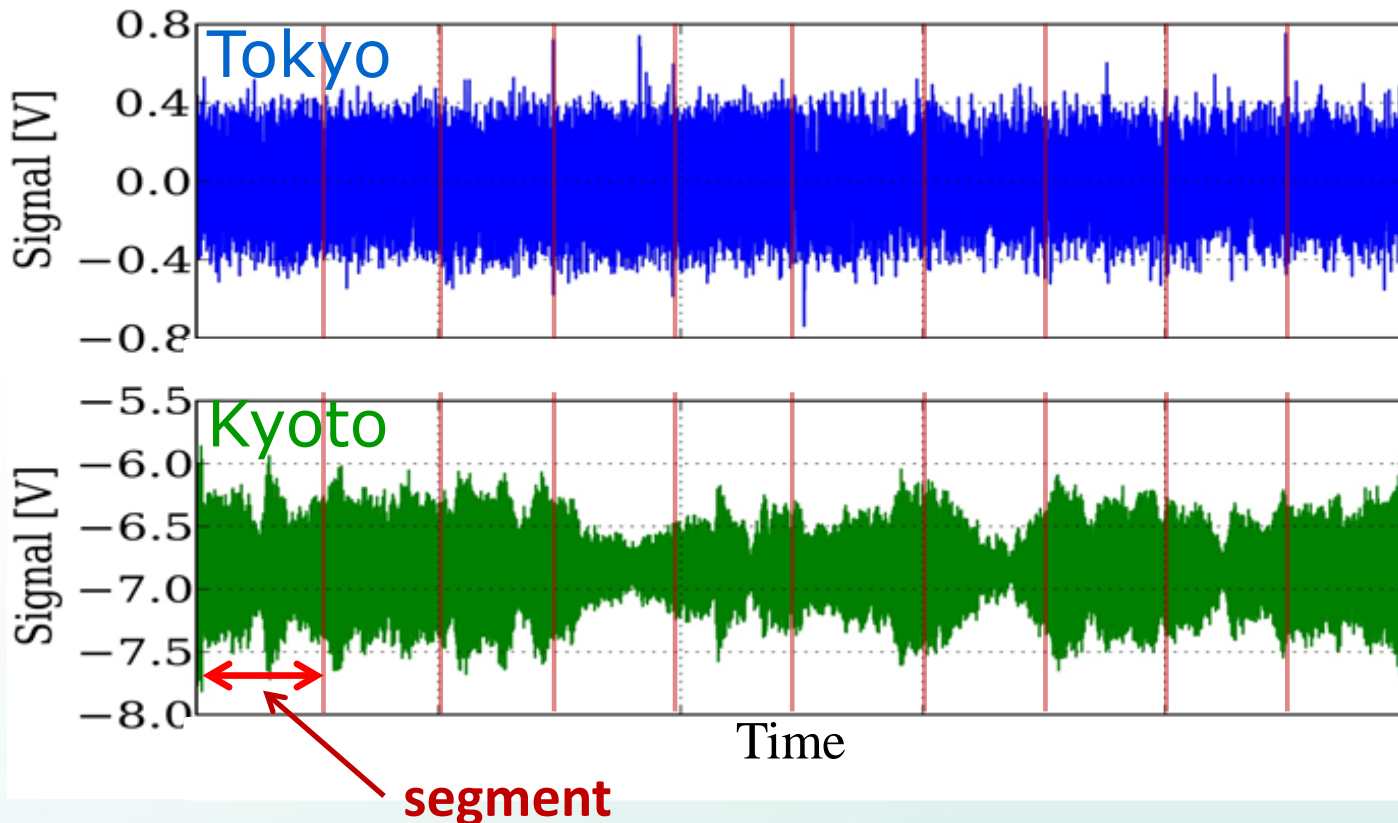
4 times better

Summary

- Established the pipeline of the cross correlation analysis with TOBAs
- The stochastic GW background signal is not detected.
- Update the upper limit on a stochastic GW background at 0.035~0.840 Hz: $h_0^2 \Omega_{\text{gw}} < 1.9 \times 10^{17}$

Backup slides

Data Selection



- Divide time series data into several segments
- Remove the segments in which RMS is big
- Calculate cross correlation with the survived segments

Cross Correlation Value

$$Y \approx \int_{f_{\min}}^{f_{\max}} df s_1^*(f) s_2(f) Q(f)$$

Optimal Filter : a filter which maximizes the signal-to-noise ratio

$$Q(f) = N \frac{\gamma(f)}{f^3 P_1(f) P_2(f)}$$

$P_i(f)$: PSD of i-th detector

Overlap reduction function

: a function which represents the difference of response to the GWs between two detectors

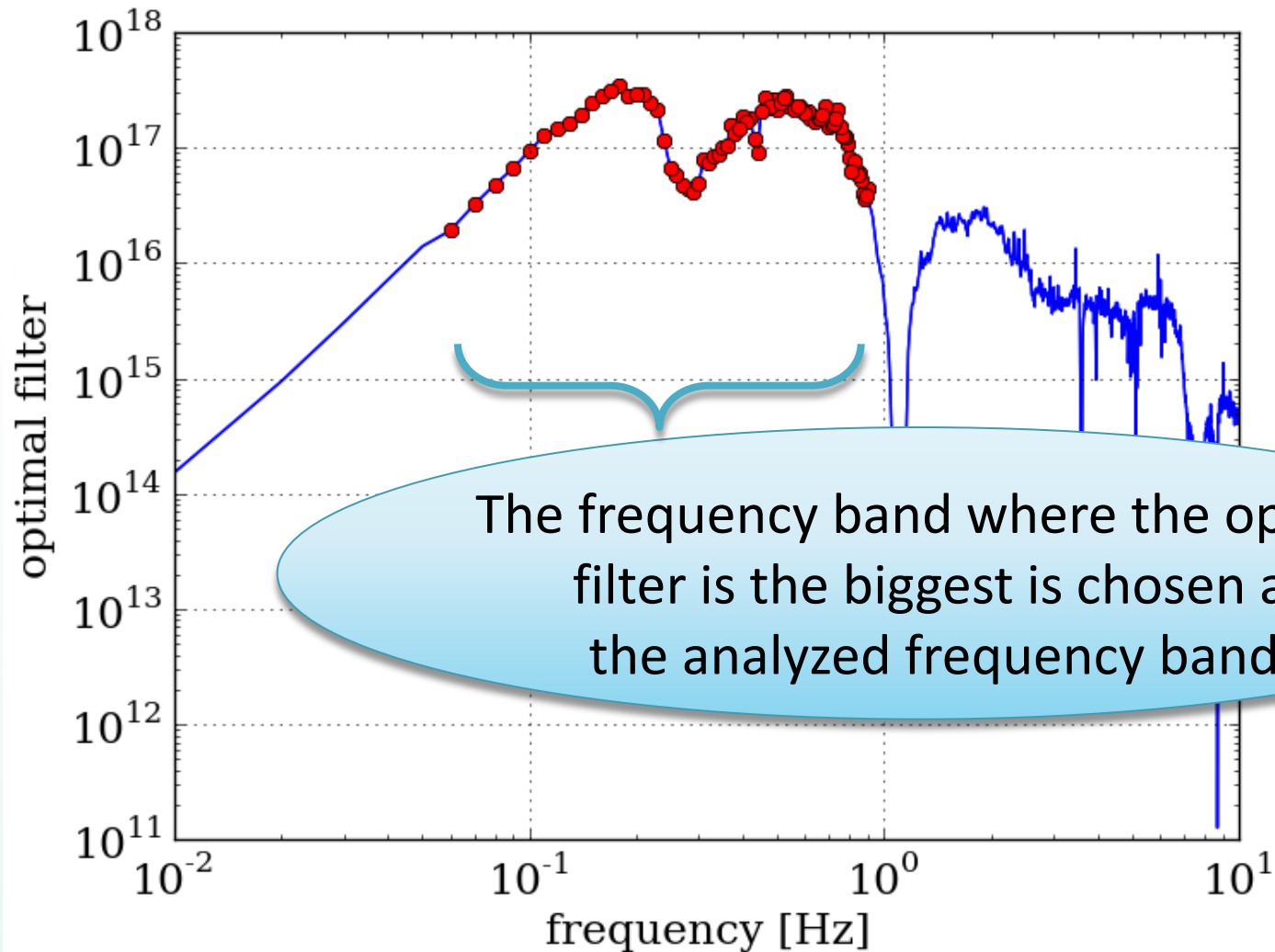
← In the case of TOBAs,
same as the interferometer's one

In this case, $\gamma(f) \approx 1$

Optimal filter

Optimal filter =

big when the sensitivity to a stochastic GW background is good

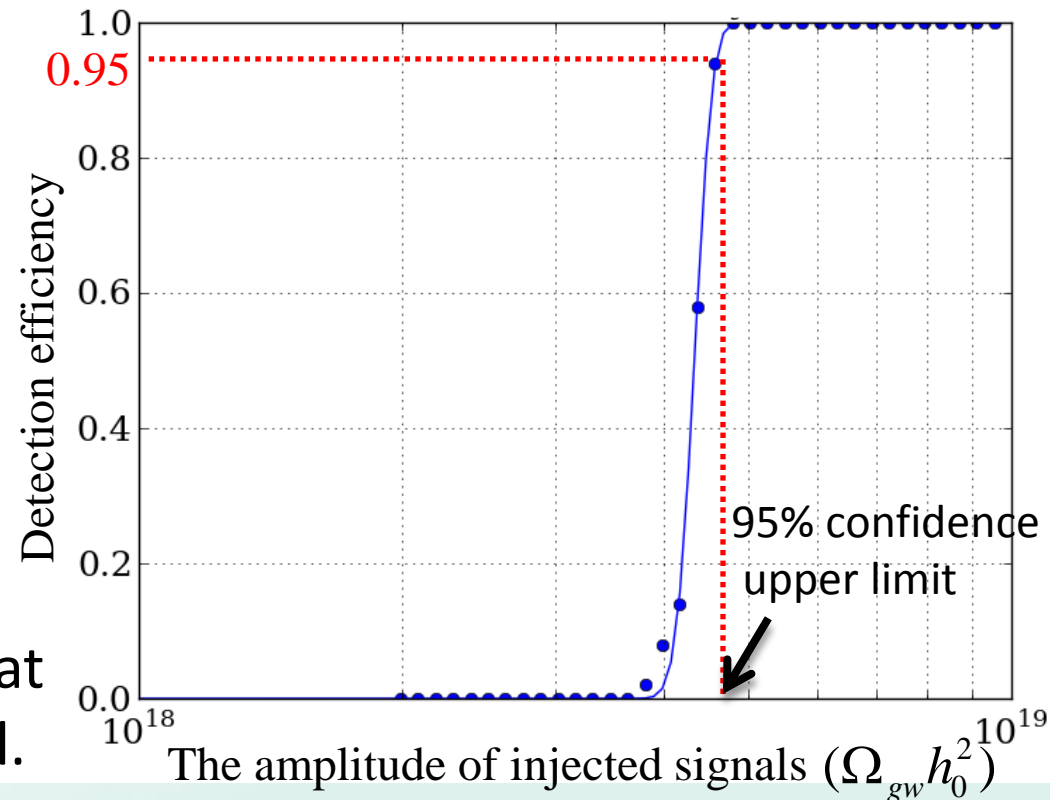


Upper Limit

How big a stochastic GW background can we detect if it would come to this data set?

1. Make a mock signal of a stochastic GW background
2. Inject the signal into the observational data
3. Perform same analysis as explained above

Repeat 1-3 to compute the rate at which we detect the mock signal.



■ Detection efficiency

Parameter Tuning

There are some parameters whose optimal values are depend on the data quality

- The length of the segments → **200 sec**
- The amount of the segments removed by data selection → **10 %**
- The bandwidth of the analyzed frequency band → **0.8 Hz**

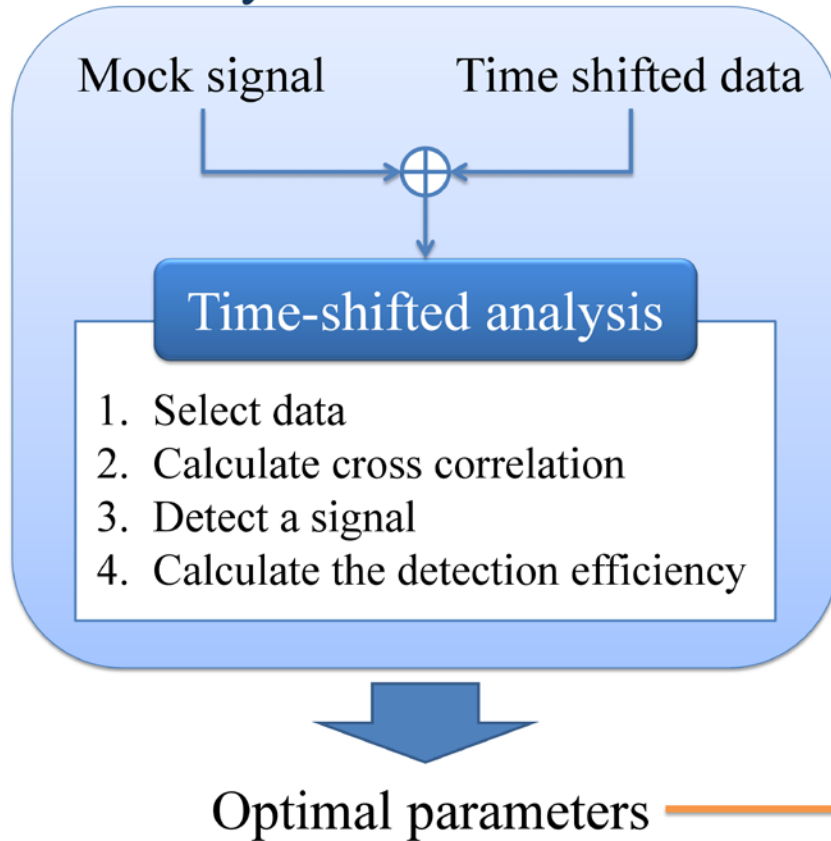
Determined

by the time shifted data.

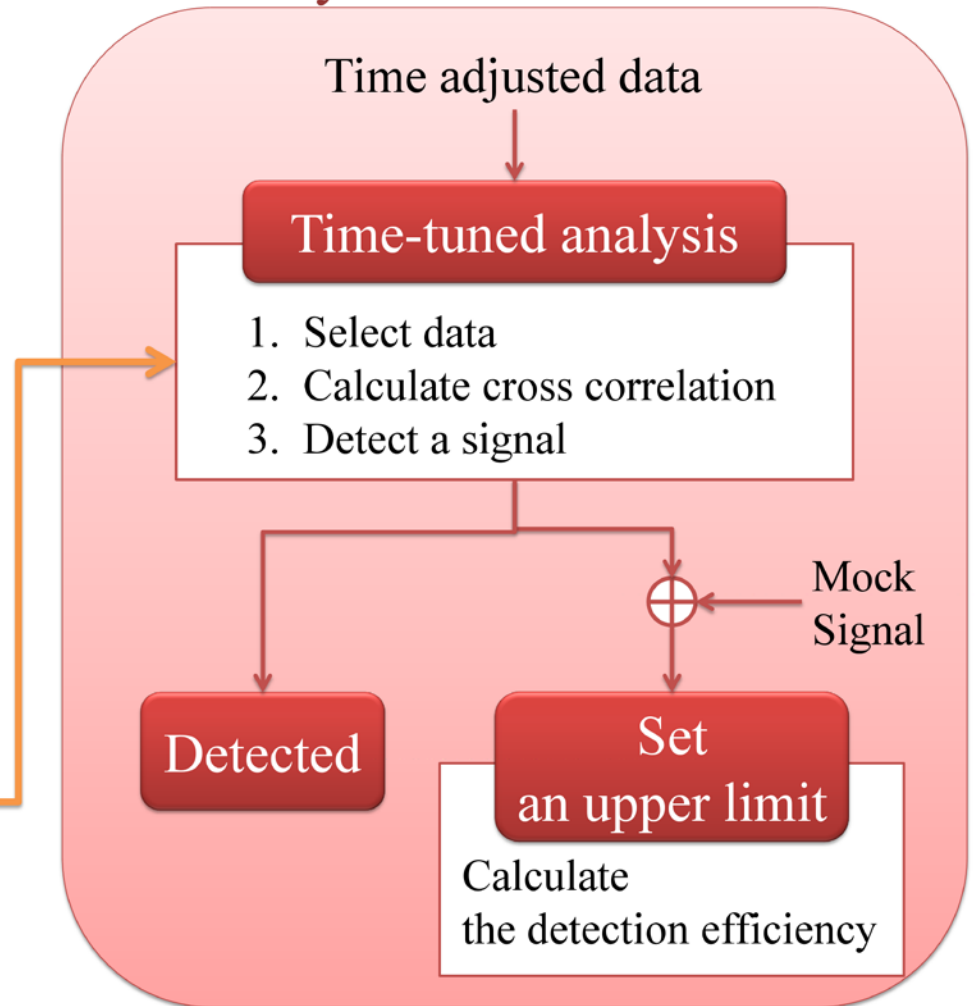
The values which make the upper limit calculated with time shifted data best is used.

Summary of Analysis

Per-Analysis

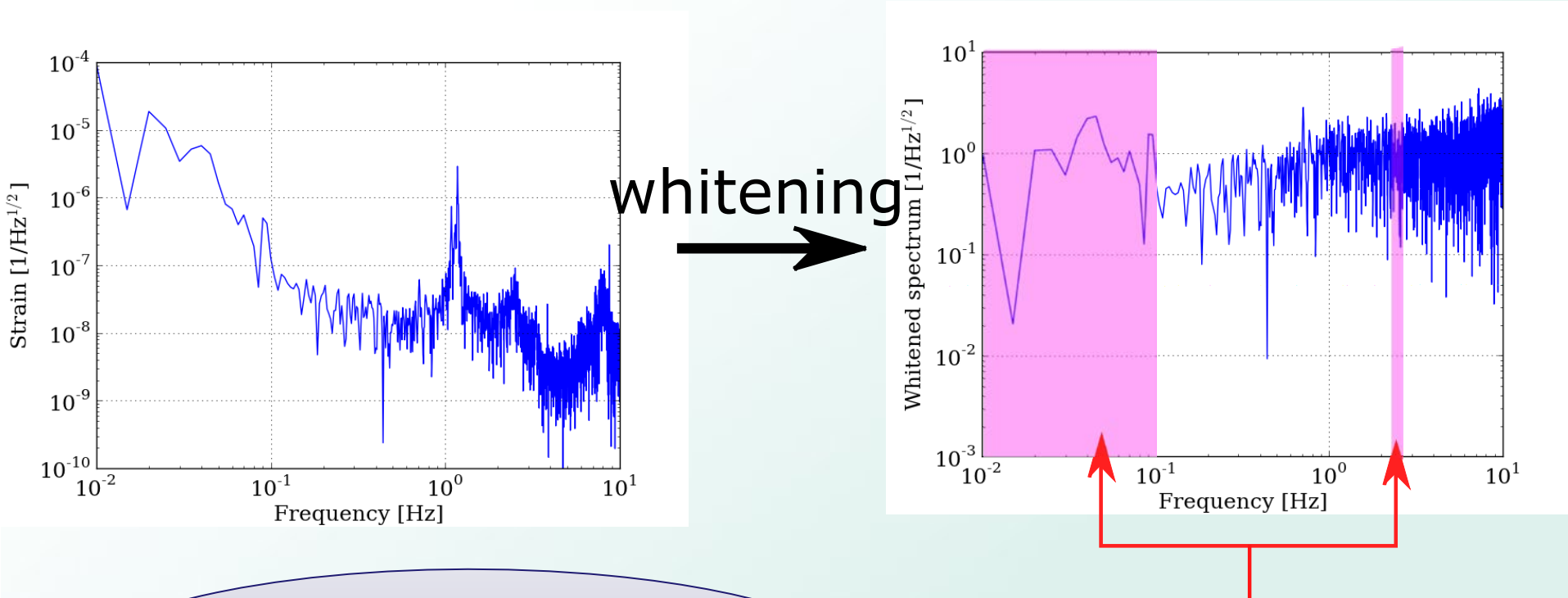


Main Analysis



Data selection

The indicator of the noise level = Whitened RMS



Avoid making the result intentionally better

The analyzed frequency band is excluded from RMS calculation

Cross correlation analysis

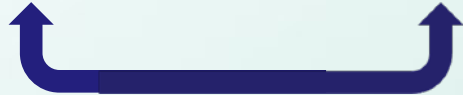
$$s(t) = h(t) + n(t)$$

signal

GW
signal

noise

Cross correlation value

$$= \int_{-T/2}^{T/2} s_1(t)s_2(t)Q(t)dt$$
$$= \int_{-T/2}^{T/2} (h_1(t) + n_1(t))(h_2(t) + n_2(t))Q(t)dt$$


Noise is reduced

$$= \int_{-T/2}^{T/2} h_1(t)h_2(t)Q(t)dt$$

Fourier transformation

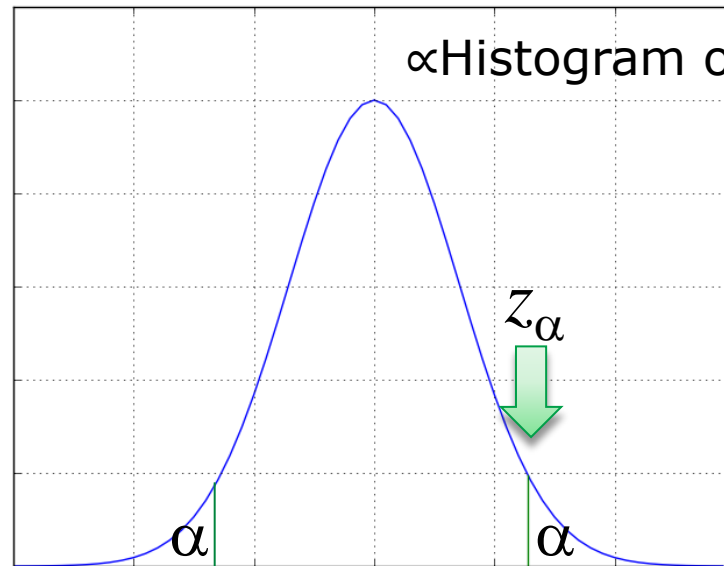
Cross correlation value

$$= \int_{f_{\min}}^{f_{\max}} df \tilde{s}_1^*(f) \tilde{Q}(f) \tilde{s}_2(f)$$

Detection Criteria

By the Neuman-Pearson criterion,

Probability distribution of $\langle Y \rangle / T_{seg}$



\propto Histogram of $\langle Y \rangle / T_{seg}$ calculated with time shifted data

$\langle Y \rangle / T_{seg}$

α : false alarm rate

Note: we do not know the sign of the two signal.



$$\langle Y \rangle / T_{seg} \geq |z_\alpha|$$

Signal is present

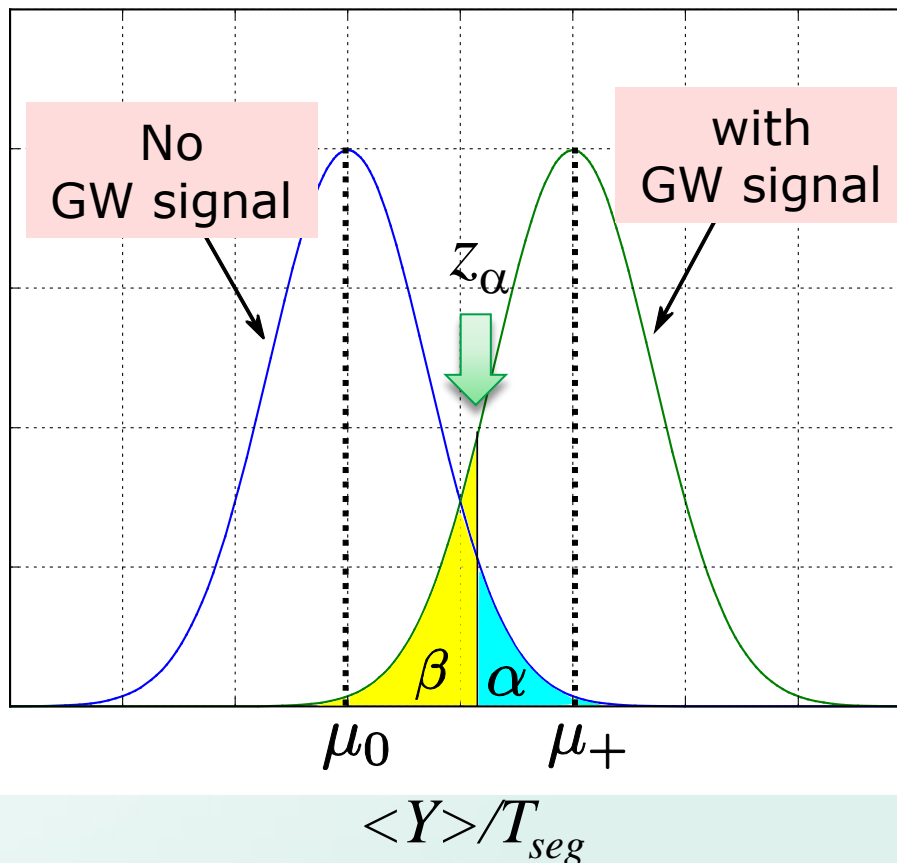
$$\langle Y \rangle / T_{seg} < |z_\alpha|$$

Signal is absent

Signal detection

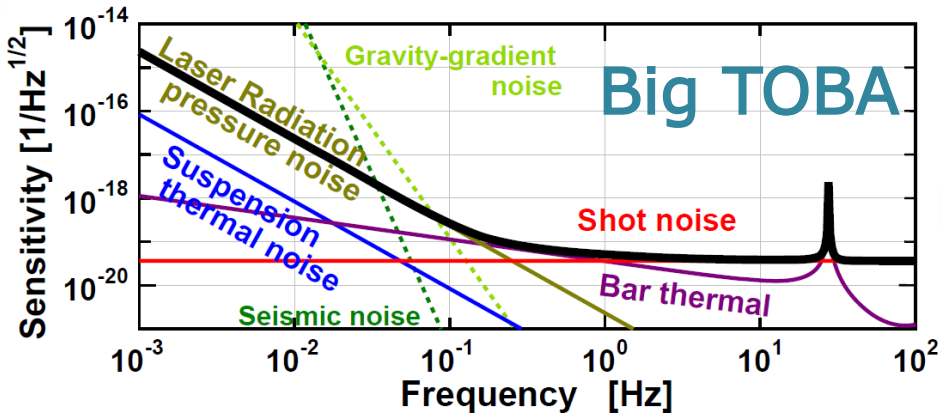
How to decide z_α :

- Calculate $\langle Y \rangle / T_{\text{seg}}$ with diversely time shifted data
- Histogram of calculated $\langle Y \rangle / T_{\text{seg}}$
= Histogram of $\langle Y \rangle / T_{\text{seg}}$ when the signal is absent.

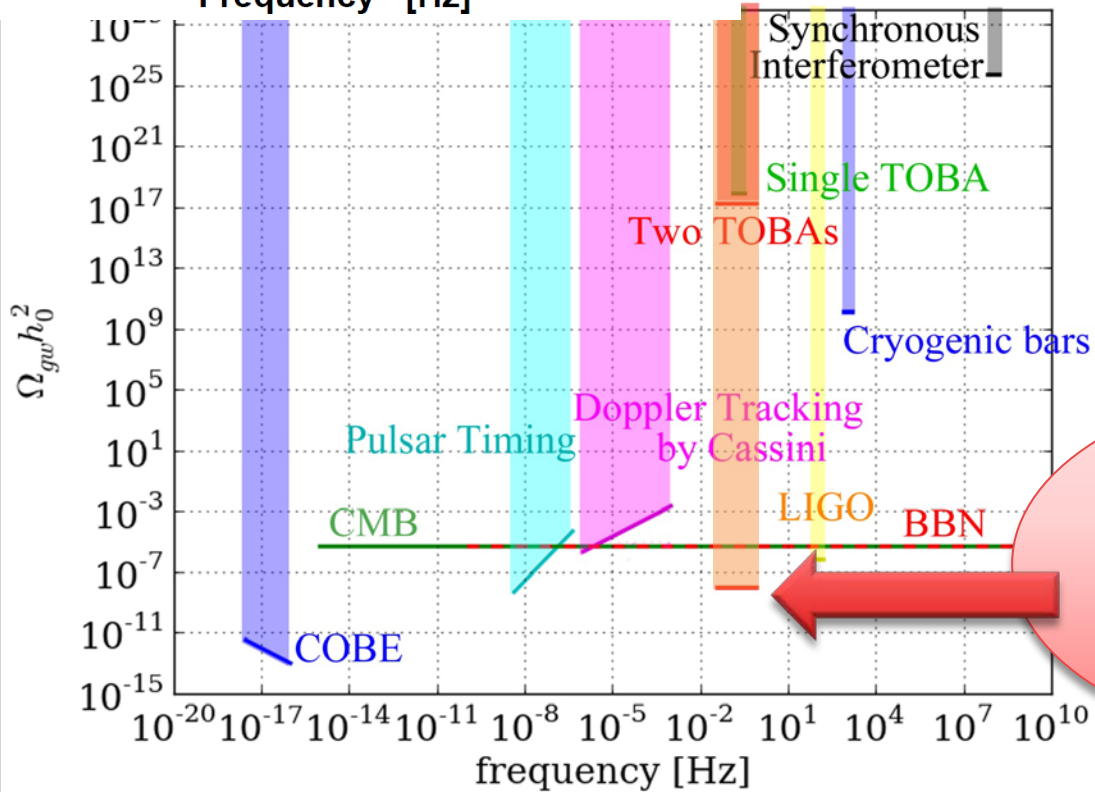


α : false alarm rate
 β : false dismissal rate

Future plan



This cross correlation analysis



1 year observation

$h_0^2 \Omega_{gw} \sim 10^{-8}$