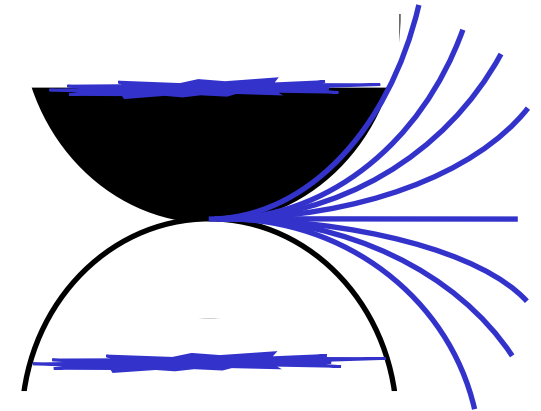
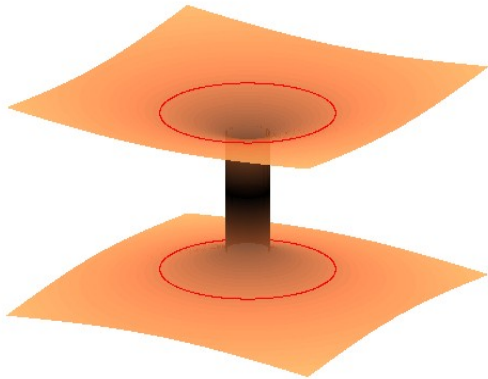
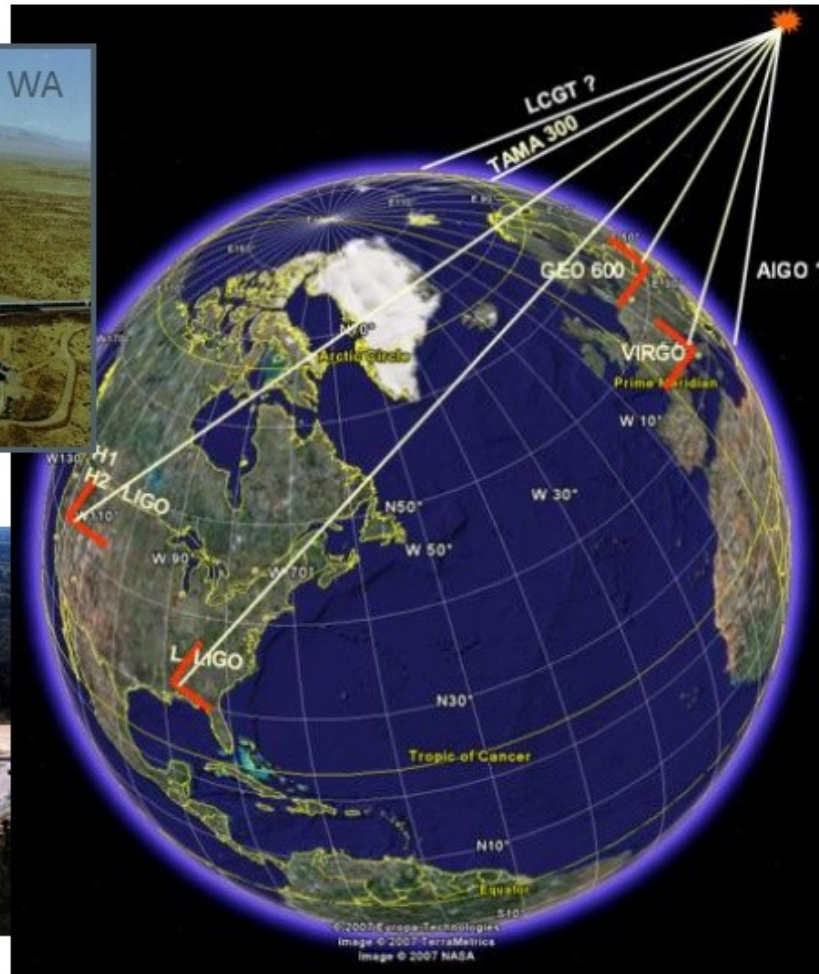
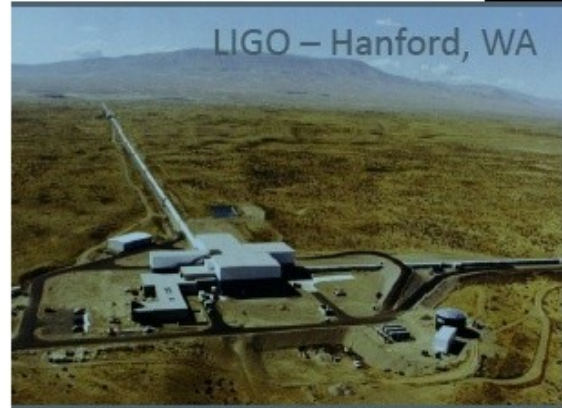


Black Holes And LIGO



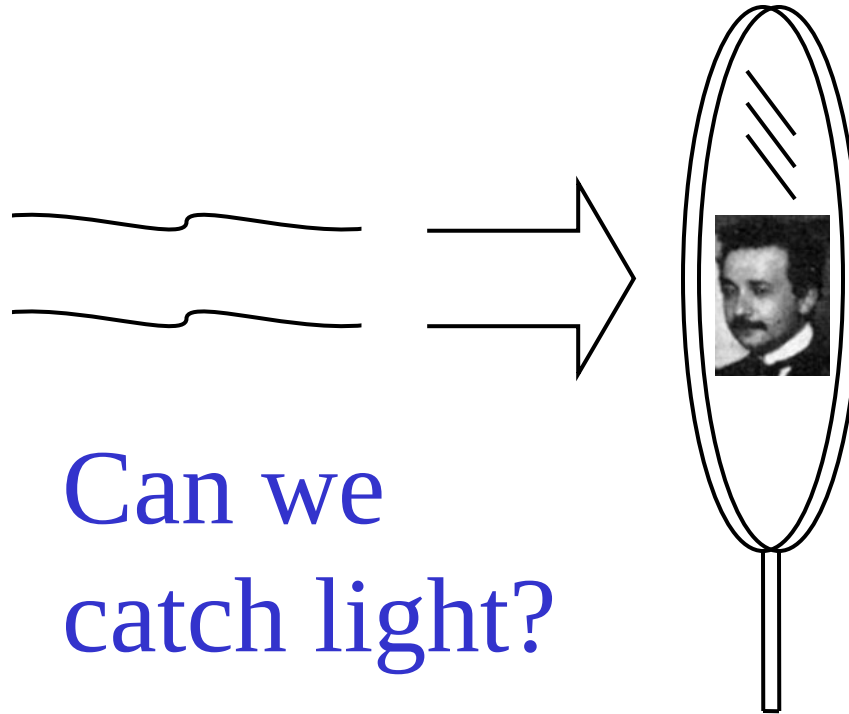
The Laser Interferometer Gravitational-wave Observatory: a Caltech/MIT collaboration supported by the National Science Foundation

Gregory Mendell
LIGO Hanford Observatory



600+ Scientist and Engineers

Einstein Wondered:



Can we
catch light?

Mirror

Photo: Albert Einstein at the first Solvay Conference, 1911; Public Domain

LIGO

Niels Bohr



Albert Einstein



Time Dilation



$$\Delta x = v \Delta t$$



$c \Delta T$

$c \Delta t$

$$\Delta T = \Delta t \sqrt{1 - v^2/c^2}$$

Δ = change in

T = time measured by motorcycle riders

t = time measured by observer at "rest"

v = speed of motorcycles

c = speed of light

Start

Warning: thought experiment only; do not try this at home.

Motorcycle: http://en.wikipedia.org/wiki/Motorcycle_racing

LIGO

The Pythagorean Theorem Of Spacetime

$$c^2\Delta T^2 + v^2\Delta t^2 = c^2\Delta t^2$$

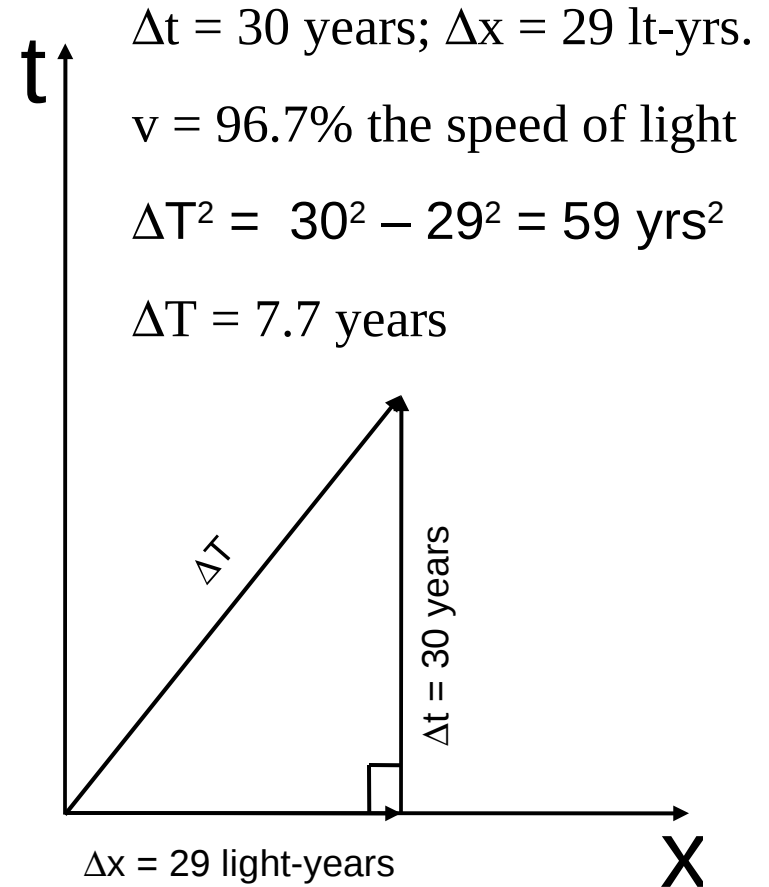
$$c^2\Delta T^2 = c^2\Delta t^2 - v^2\Delta t^2$$

$$c^2\Delta T^2 = c^2\Delta t^2 - \Delta x^2$$

$$c = 1 \text{ light-year/year}$$

$$\Delta T^2 = \Delta t^2 - \Delta x^2$$

Pythagorean Thm. of Spacetime



Spacetime

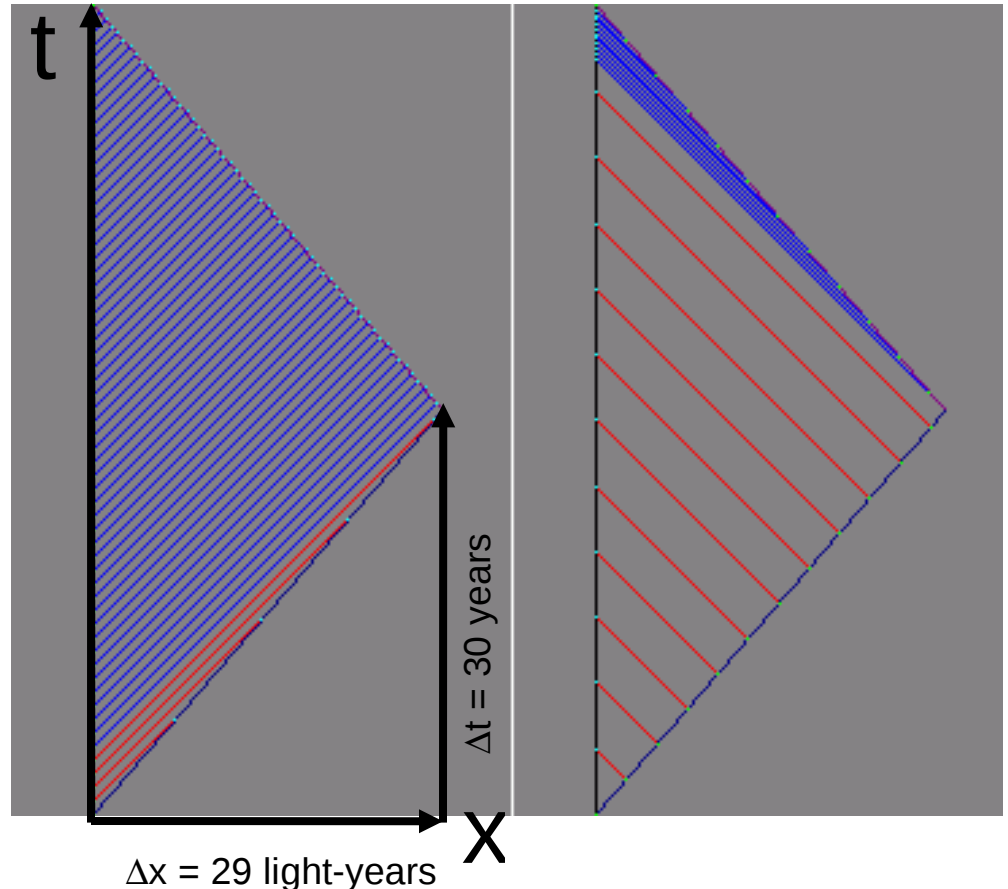
Spacetime Diagram

The Twin Paradox

- Imagine twins, Betty and Bob, separated 1 year after birth. Baby Betty & Bob: ☺ ☺

- Betty takes a rocket travelling at 96.67% the speed of light and travels 29 lt-yrs from Earth and back.

- When Betty returns she is sweet 16, and Bob is 61 years old!!!



$$\Delta T = 30 \text{ yrs} \sqrt{1 - (.9667)^2} = 7.7 \text{ yrs}$$

LIGO

Einstein's Happiest Thought: Gravity Disappears When You Free Fall



http://en.wikipedia.org/wiki/Leaning_Tower_of_Pisa



Photo: NASA

Einstein had this thought in 1907. This led to the idea that gravity is the curvature of spacetime. Here I paraphrase a thought experiment I first heard from Kip Thorne. Suppose two friends jump parallel to each other off the Leaning Tower of Pisa. For the friends, gravity has disappeared, and they believe they are in empty space. Strangely though, they find their parallel paths converging at the center of the Earth. That can happen in empty space only if that space is not flat but curved. Einstein thought about the geometry of rotating objects, and other things, and after 8 more years produced General Relativity, which is a theory of gravity and spacetime. He had help from a mathematician, Marcel Grossmann.

Warning: thought experiment only; do not try this at home.

Pythagorean Theorem and Einstein's General Theory of Relativity

$\Delta \rightarrow d =$ infinitesimal
change

$$dT^2 = g_{tt} dt^2 + g_{xx} dx^2$$

$$dT^2 = g_{\mu\nu} dx^\mu dx^\nu$$

In GR the components of a
4x4 symmetric matrix
called the metric tensor
define the curvature of
spacetime.

$$G_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R$$

$$R_{\mu\nu} = R^\alpha_{\mu\alpha\nu}; R = g^{\mu\nu} R_{\mu\nu}$$

$$R^\alpha_{\mu\beta\nu} = \partial_\beta \Gamma^\alpha_{\mu\nu} - \partial_\nu \Gamma^\alpha_{\mu\beta} + \Gamma^\alpha_{\beta\gamma} \Gamma^\gamma_{\mu\nu} - \Gamma^\alpha_{\gamma\nu} \Gamma^\gamma_{\mu\beta}$$

$$\Gamma^\alpha_{\mu\nu} = \frac{1}{2} g^{\alpha\beta} (\partial_\nu g_{\mu\beta} + \partial_\mu g_{\beta\nu} - \partial_\beta g_{\mu\nu})$$

Einstein's Field Equations

$$\frac{dx^\alpha}{dT} = U^\alpha; \quad U_\alpha = g_{\alpha\beta} U^\beta \quad U = 4\text{-Vel.}; T = \text{Proper Time}$$

$$\frac{dU_\alpha}{dT} = \frac{1}{2} \partial_\alpha g_{\beta\gamma} U^\beta U^\gamma \quad \text{Geodesic Equation}$$

Schwarzschild Black Hole

$$c^2 dT^2 = \left(1 - \frac{2GM}{rc^2}\right) c^2 dt^2 - \frac{1}{\left(1 - \frac{2GM}{rc^2}\right)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$

$$c^2 dT^2 = \left(1 - \frac{v_{esc}^2}{c^2}\right) c^2 dt^2 - \frac{1}{\left(1 - \frac{v_{esc}^2}{c^2}\right)} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\phi^2$$



Karl Schwarzschild

$$v_{esc} = \sqrt{\frac{2GM}{r}}$$

• Escape Velocity

$$R_s = \frac{2GM}{c^2}$$

• Schwarzschild Radius

<u>Object</u>	<u>Schwarzschild Radius</u>
You	1 thousand, million, million, millionth the thickness of a human hair
Earth	1 cm (size of marble)
Sun	3 km (2 miles)
Galaxy	~ trillion miles

Gravitational Time Dilation



$$\Delta T = \sqrt{1 - \frac{2GM}{rc^2}} \Delta t$$

Gravity
slows time
down!

Photo:http://en.wikipedia.org/wiki/Leaning_Tower_of_Pisa

Clock_Photos:http://en.wikipedia.org/wiki/Cuckoo_clock

LIGO

Gravity Slows Time

- Due to the orbital speed, clocks on the satellite lose 7 microseconds per day
- Due to the weaker gravitational field, clocks on the satellite gain 45 microseconds per day
- Satellite clocks gain a net of 38 microsecond per day
- Distance error = $c \cdot 38$ microseconds; $c = 186,000$ miles per second.
- Without calibrating clocks to account for Relativity, GPS distance would be off by 7 miles after one day!

See Scientific American, Sept. 1994



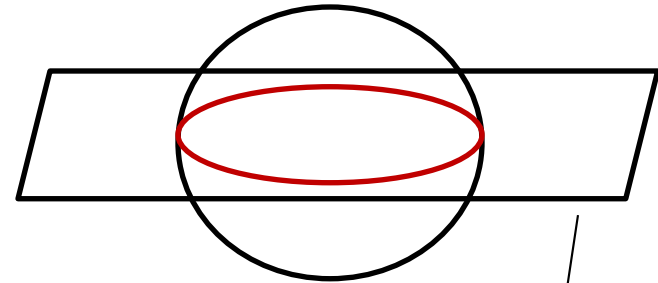
Illustration: NASA

Clock_Photos: http://en.wikipedia.org/wiki/Cuckoo_clock

Embedding Diagram

Schwarzschild for $t = 0, \theta = \pi / 2$:

$$ds^2 = \frac{1}{\left(1 - \frac{2GM}{rc^2}\right)} dr^2 + r^2 d\phi^2$$



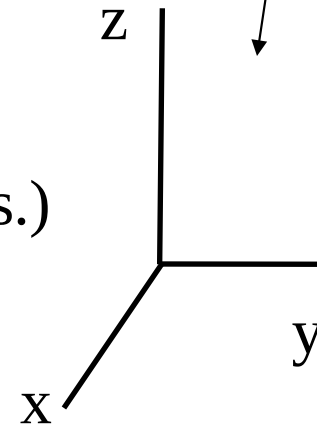
Flat space cylindrical coordinates:

$$ds^2 = dz^2 + dr^2 + r^2 d\phi^2$$

$z = f(r)$ (Surface of revolution about z-axis.)

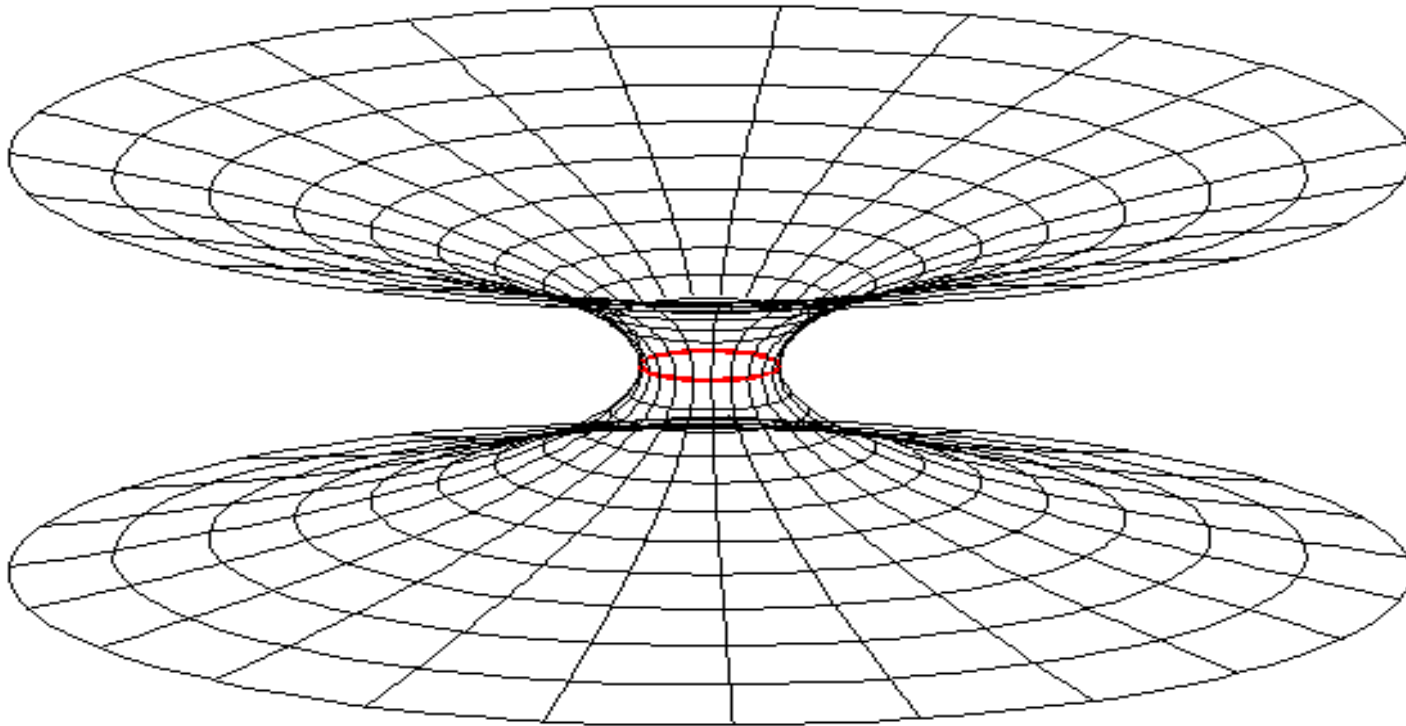
$$dz = f'(r) dr$$

$$ds^2 = [f'(r)^2 + 1] dr^2 + r^2 d\phi^2$$



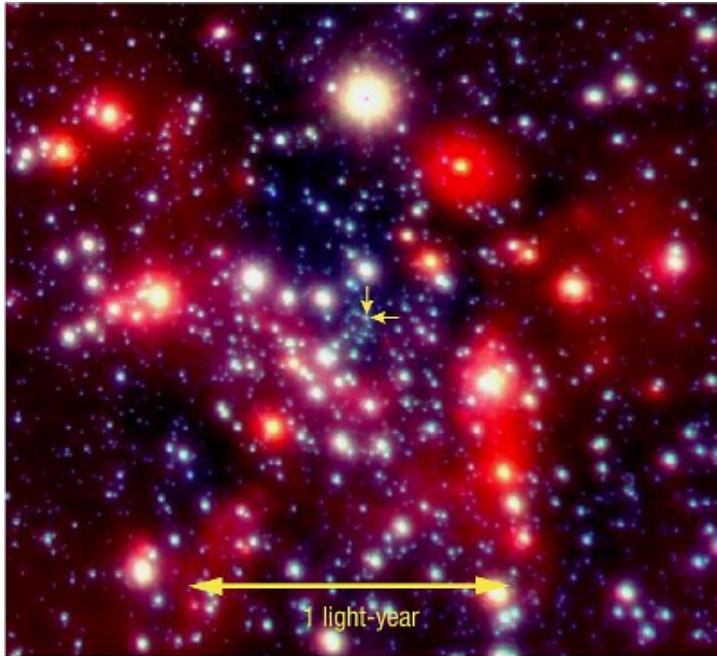
Einstein-Rosen Bridge

Our Universe



Another Universe?

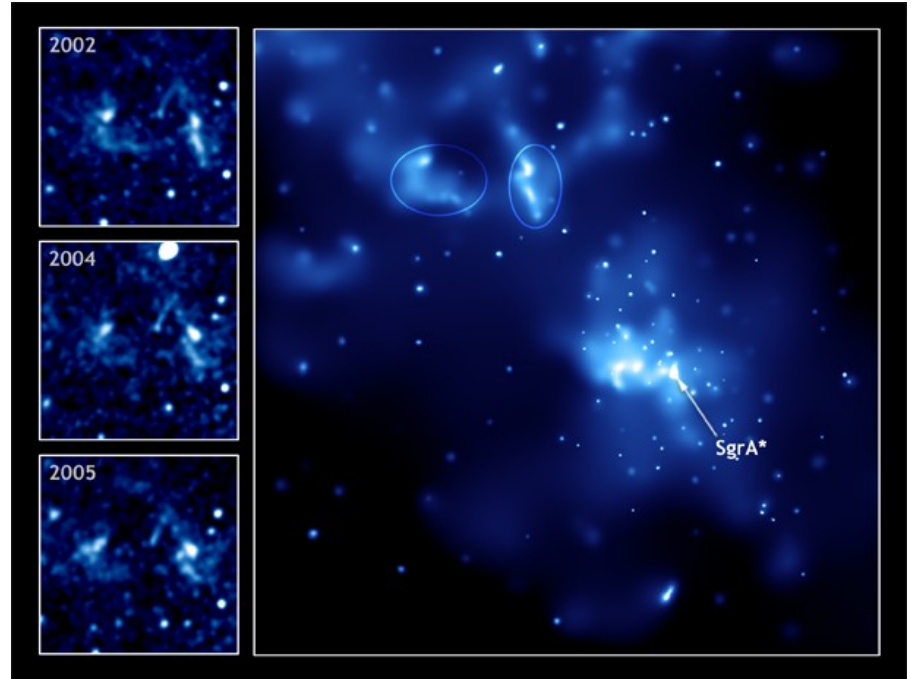
The Center Of The Milky Way



The Centre of the Milky Way
(VLT YEPUN + NACO)

ESO PR Photo 23a/02 (9 October 2002)

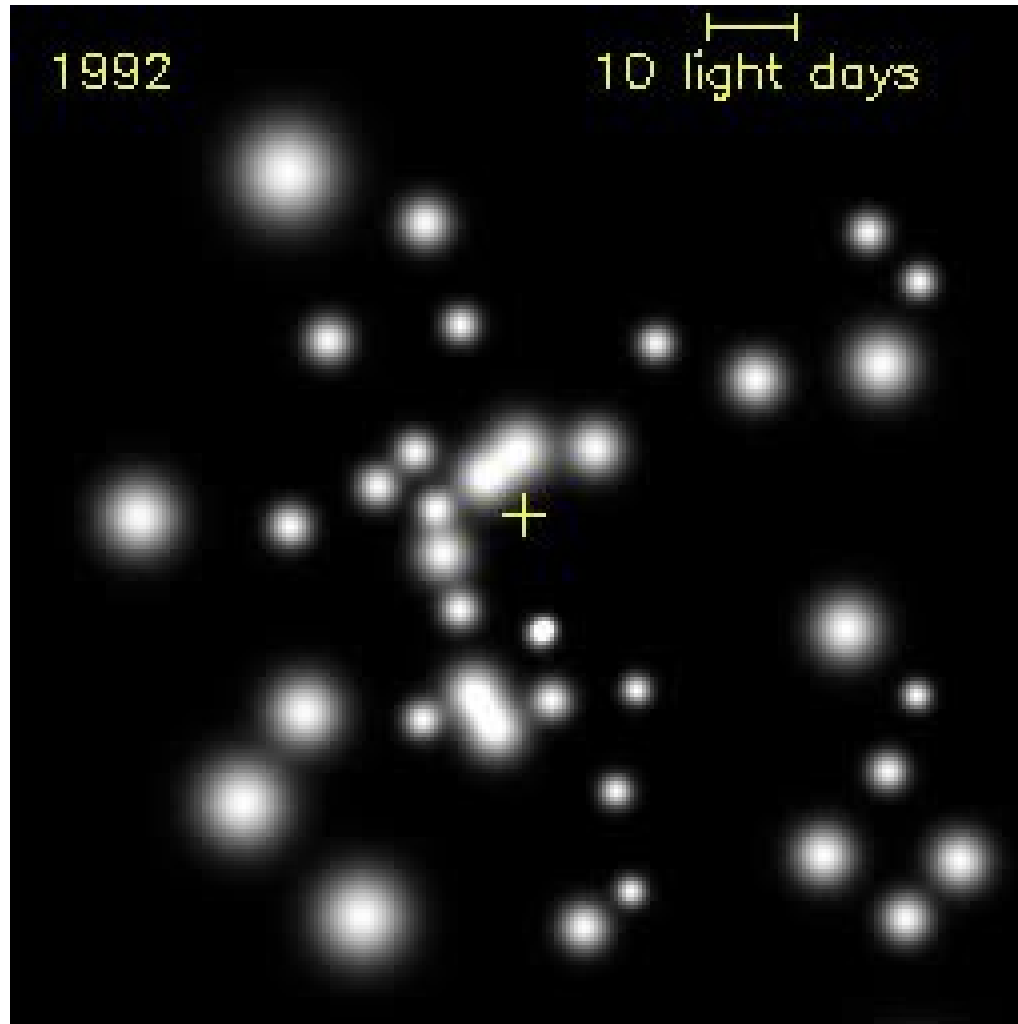
© European Southern Observatory



Credit: NASA/Chandra X-Ray Observatory

LIGO

Zooming in on the galactic center...



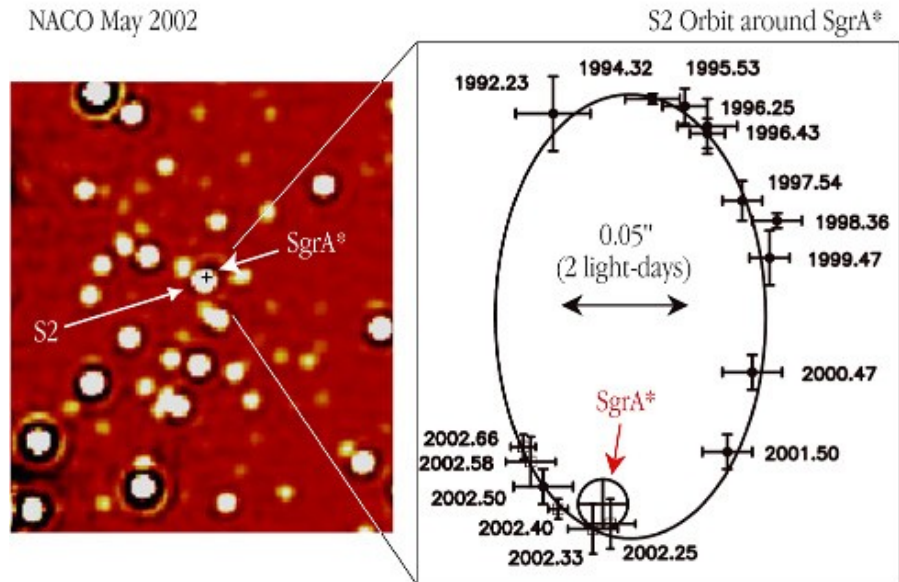
Credit: ESO PR Video Clip 02/02; ESO/European Organization for Astronomical Research in the Southern Hemisphere; Press Release 2002

Black Hole Detection

$$M = \frac{4\pi^2 a^3}{GT^2}$$

$$M = \frac{(900 \text{ AU})^3}{(15 \text{ yrs})^2}$$

= 3 million Solar Masses



The Motion of a Star around the Central Black Hole in the Milky Way

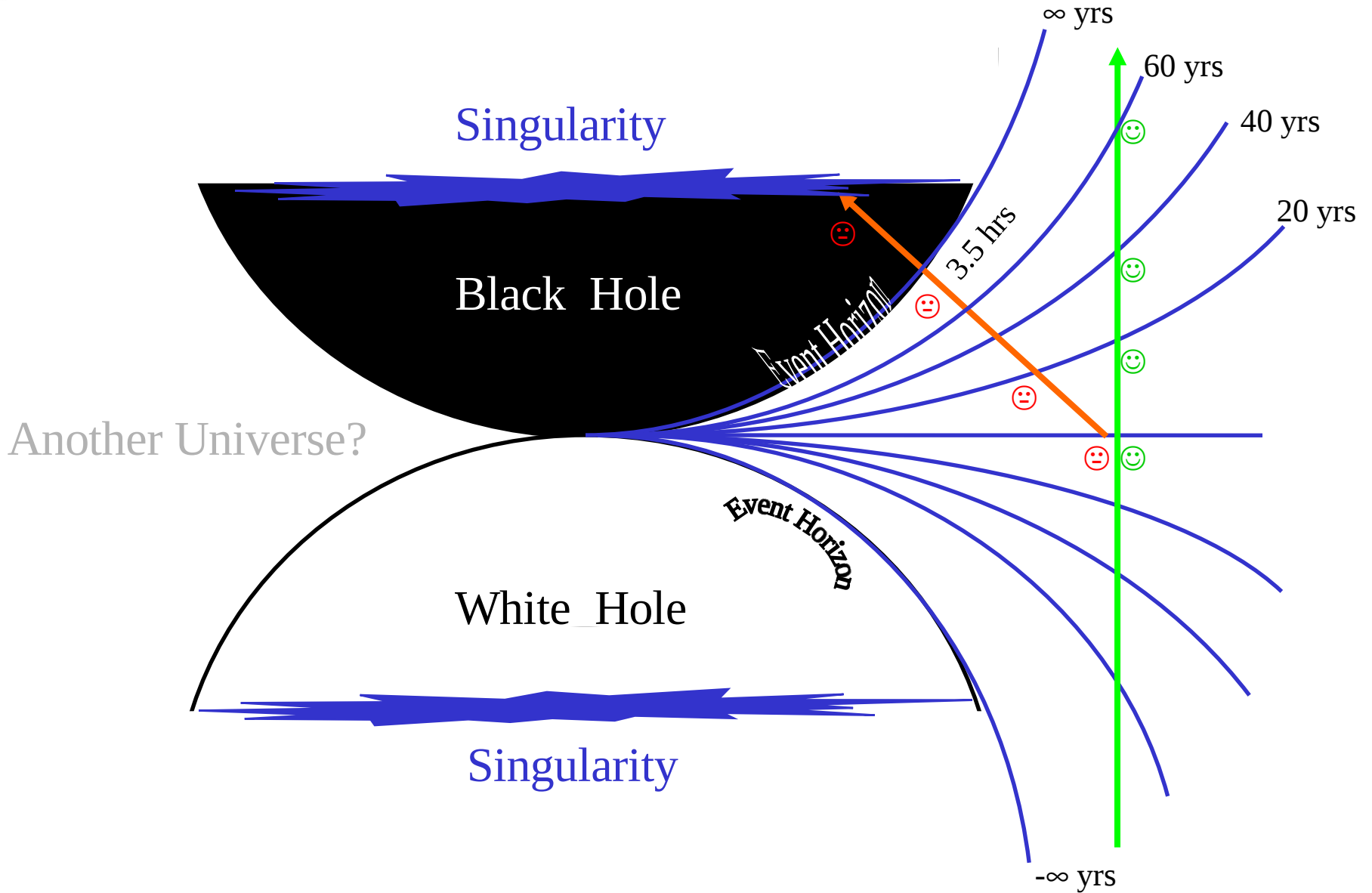
ESO PR Photo 23c02 (9 October 2002)

© European Southern Observatory



Conclusion: there is a Black Hole at the center of our Galaxy that has a mass 3 millions times (or more precisely 3.95 million times) that of the Sun. S2 orbits this Black Hole at a distance of 12000 Schwarzschild Radii.

Falling Into A Black Hole



Embedding Diagram Inside The Black Hole

Schwarzschild for $r = R$, $\theta = \pi / 2$:

$$ds^2 = c^2[2GM/(Rc^2)-1]dt^2 + R^2d\phi^2.$$

Flat space cylindrical coordinates:

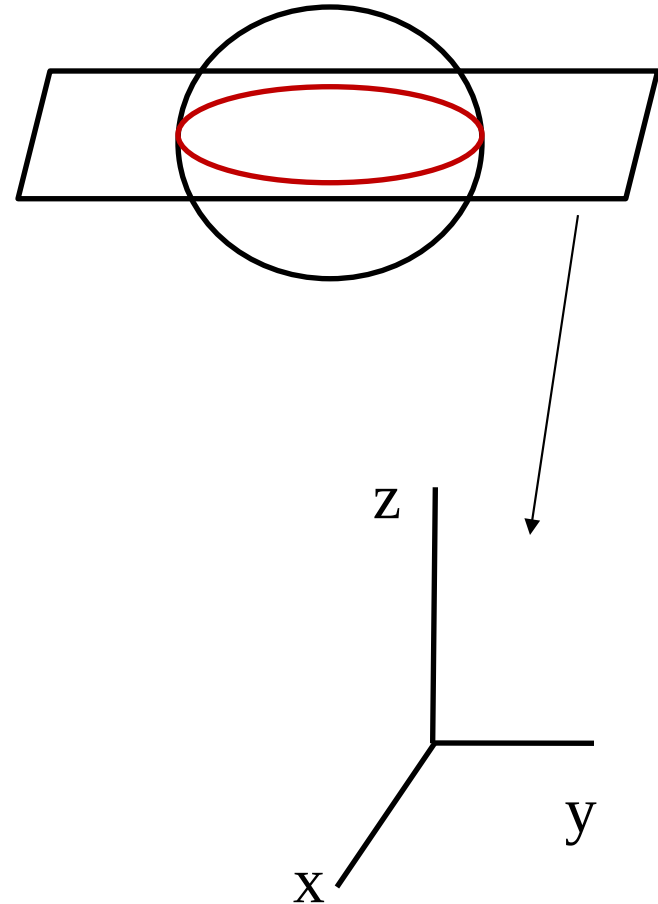
$$ds^2 = dz^2 + dr^2 + r^2d\phi^2.$$

Comparing, it looks like in the flat space $r = R = \text{constant}$, so

$$ds^2 = dz^2 + R^2d\phi^2.$$

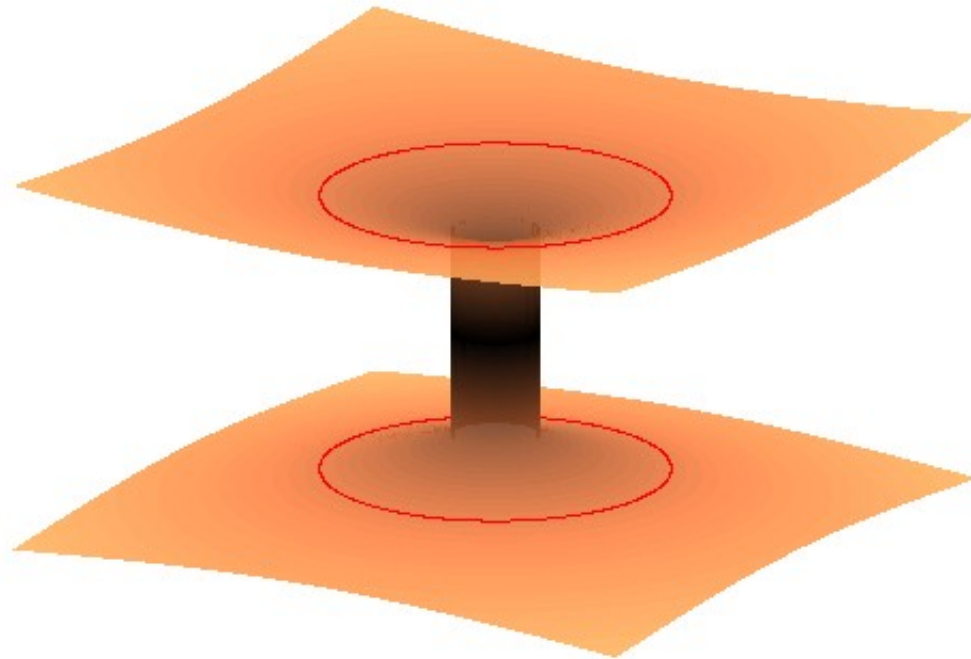
We need to match up:

$$dz^2 = c^2[2GM/(Rc^2)-1]dt^2.$$



Schwarzschild Worm Hole

Our Universe

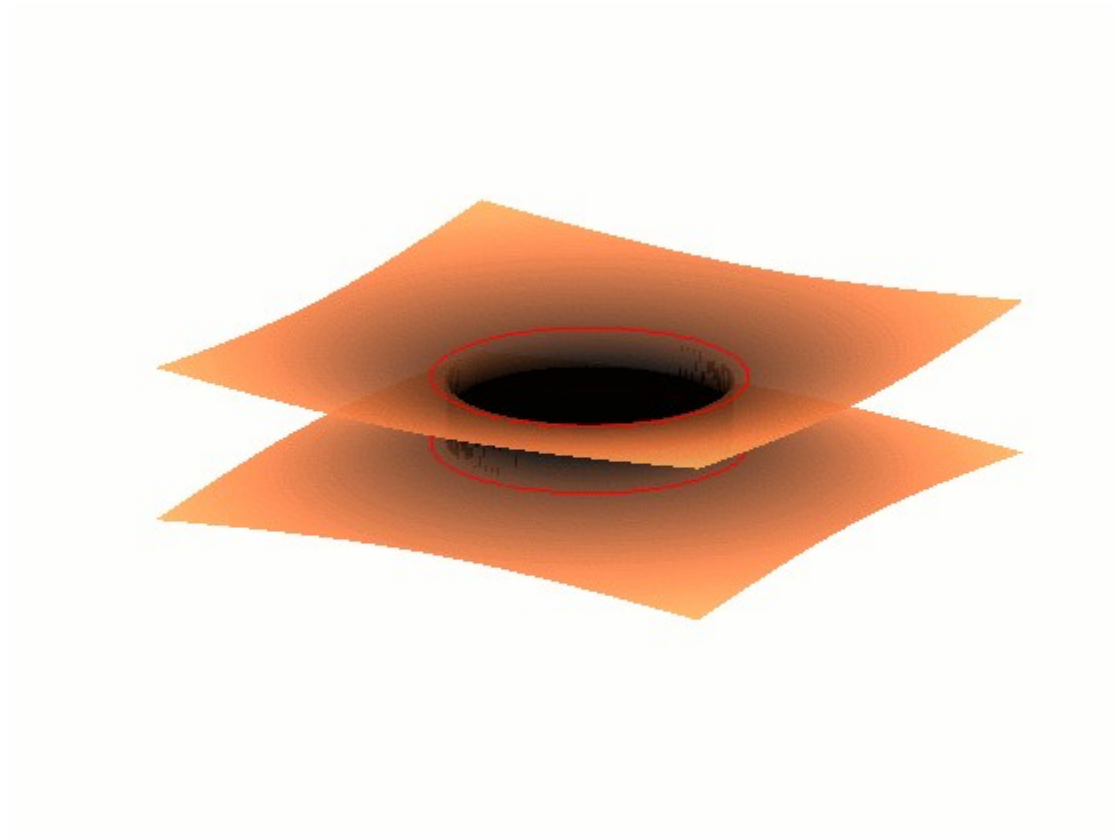


Another Universe

LIGO

Embedding With Interior Dynamics

Our Universe

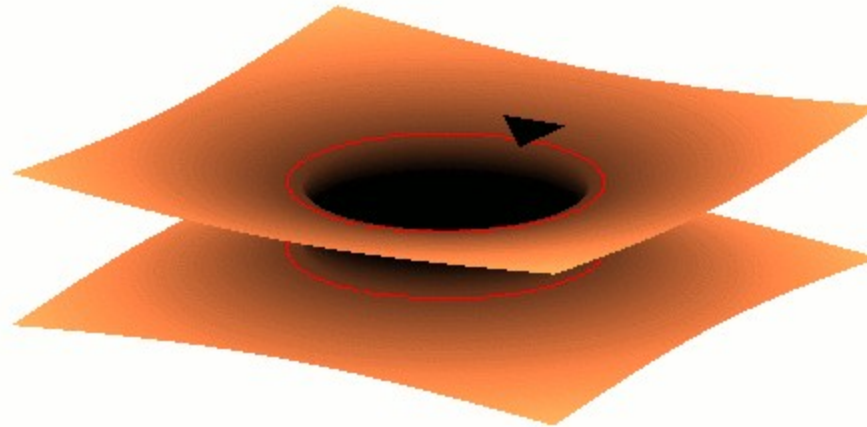


Another Universe

Nontraversable Wormhole

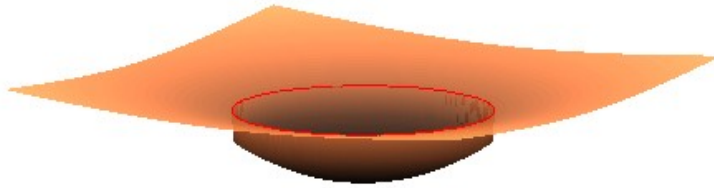
Our Universe

TIME INSIDE = 0.0 sec. ROCKET TIME = 0.0 sec. TIME OUTSIDE = 100.3 sec.



Another Universe

Stellar Collapse To Form A Black Hole

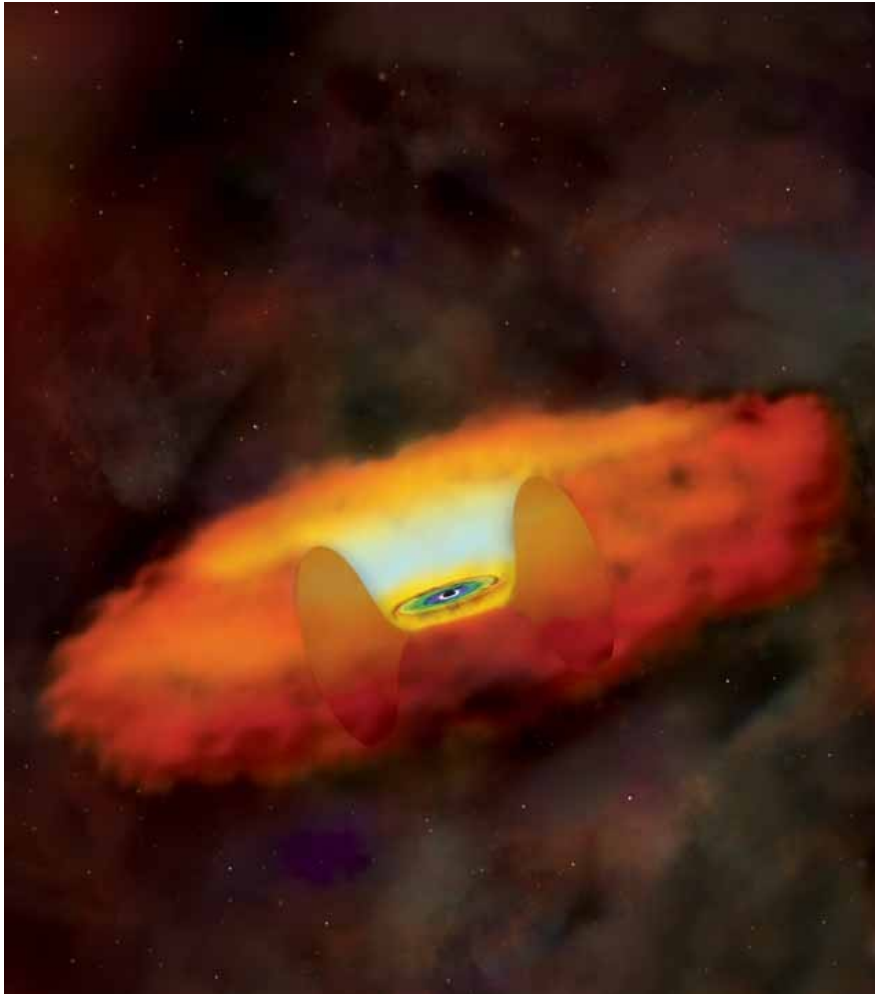


When pressure can no longer support a star's gravity its mass falls through its horizon.



And it collapses to a Singularity.

LIGO Black Holes & Accretion Disks



<http://researchnews.osu.edu/archive/fuzzballpic.htm>
(Illustration: CXC/M.Weiss)

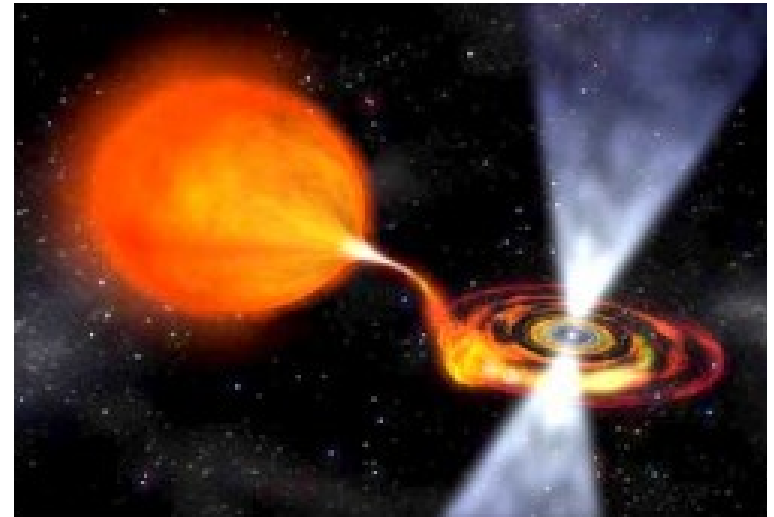
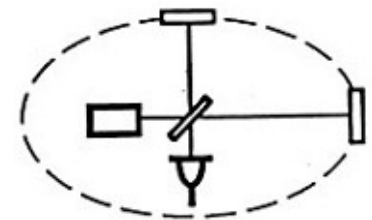
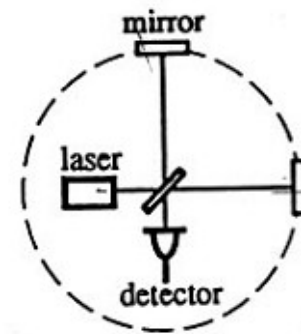
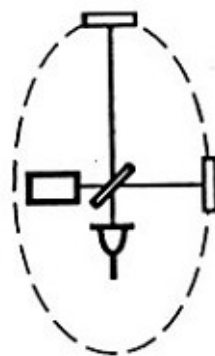
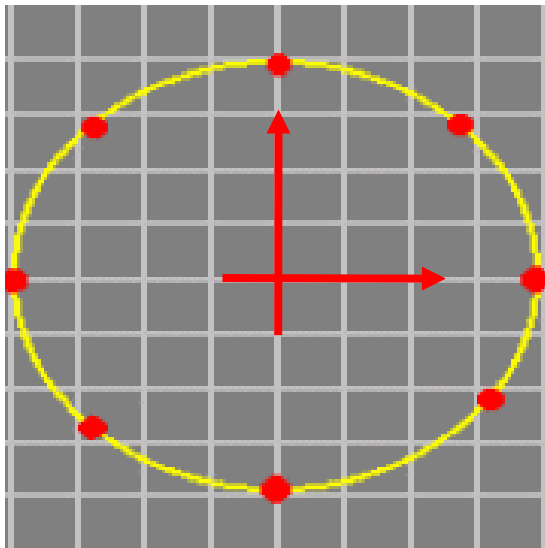
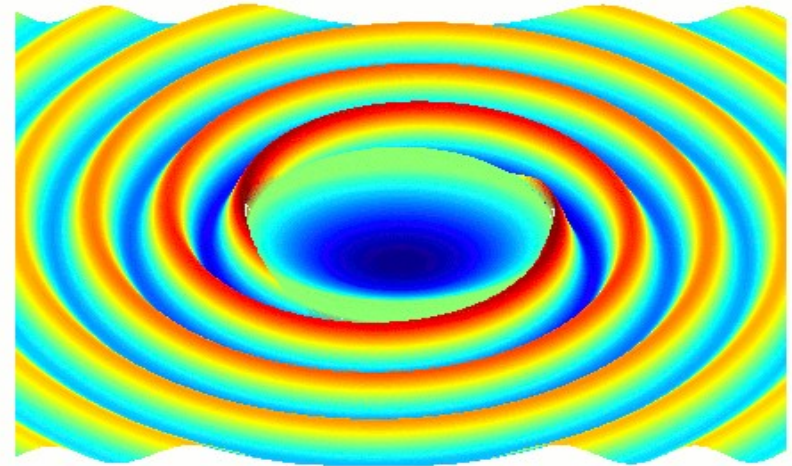


image by Dana Berry/NASA; NASA News Release
posted July 2, 2003 on Spaceflight Now.

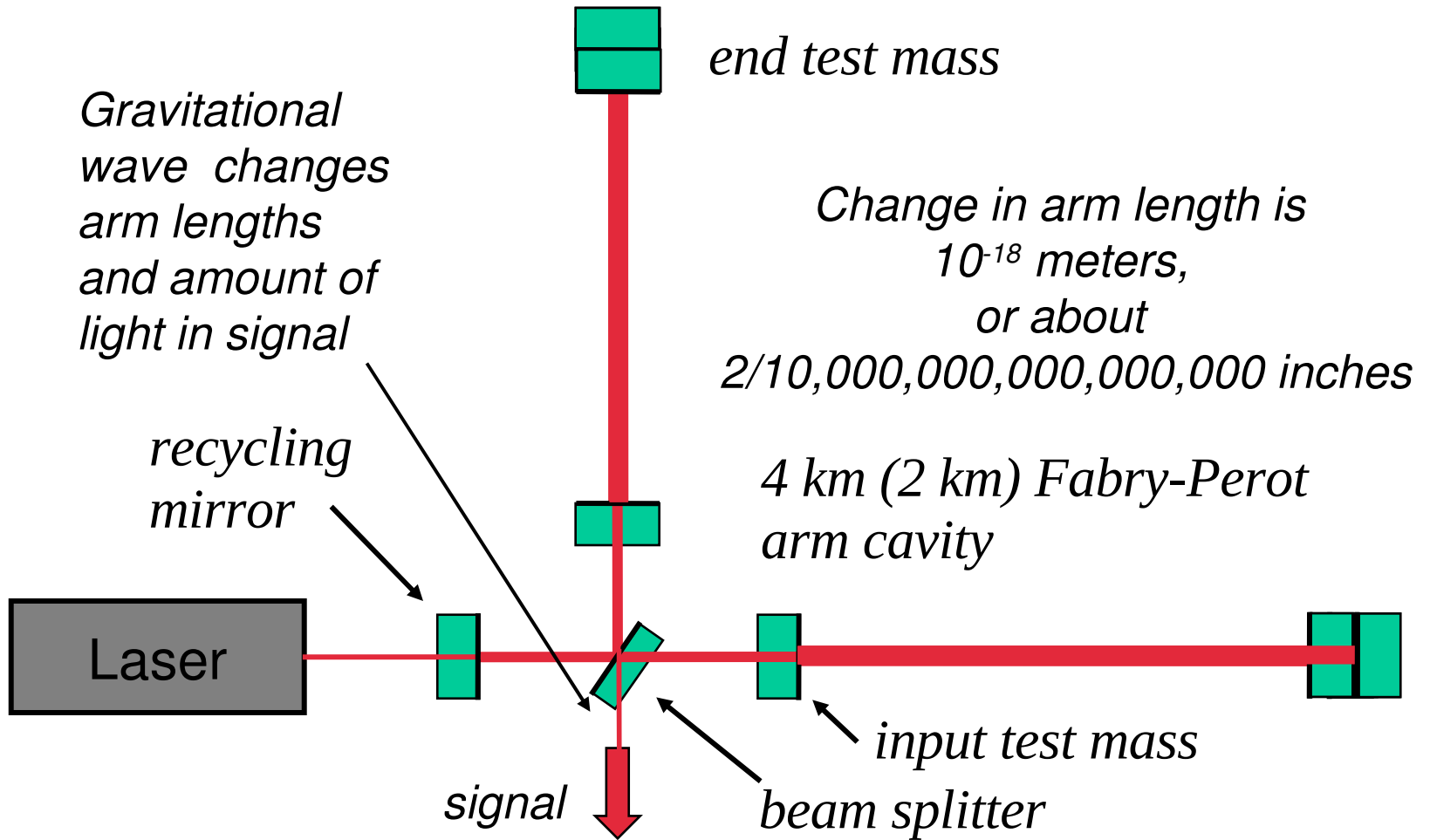
Gravitational Waves

Gravitational waves are ripples in spacetime when it is stirred up by rapidly changing motions of large concentrations of matter or energy. **The waves are extremely weak by the times they reach Earth.**

Illustration of Gravitational Waves:



Sensing the Effect of a Gravitational Wave



LIGO is in some ways like a space mission
flying a few feet off the ground



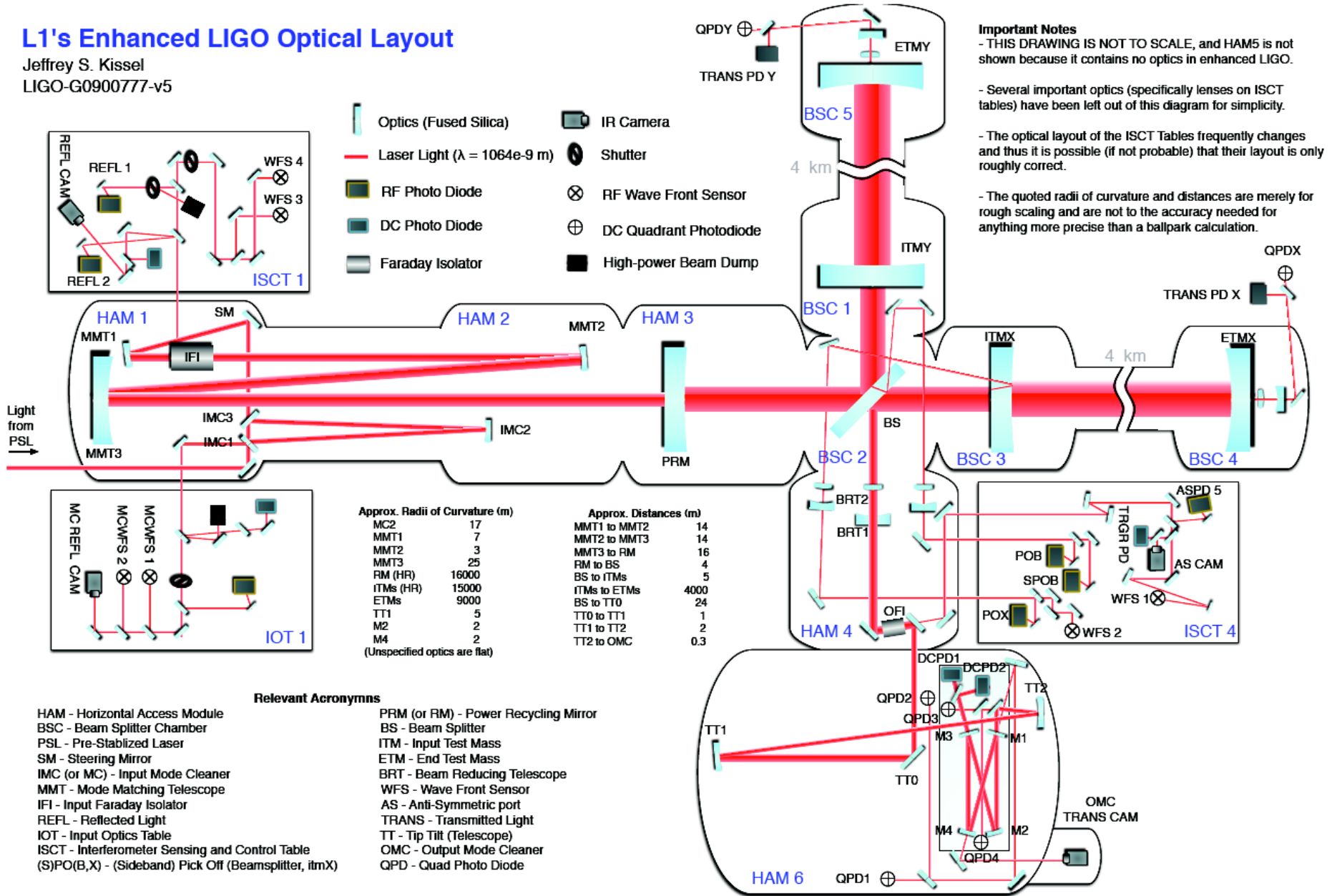
One of world's largest ultra high
vacuum systems.

~ 10,000 m³

10⁻⁹ torr

L1's Enhanced LIGO Optical Layout

Jeffrey S. Kissel
LIGO-G0900777-v5



Binary Black Hole Coalescence

Show movies from: **Simulating Extreme Spacetimes – SXS - Caltech
– Cornell Project. <http://www.black-holes.org/explore2.html>**

Credit: Introduction to LIGO & Gravitational Waves:
<http://www.ligo.org/science/GW-Inspiral.php>



Credit: Scott Hughs, MIT group:





- During their 2009-2010 science runs, the LIGO Scientific Collaboration and the Virgo Collaboration did an end-to-end test with a **blind hardware injection of a fake signal into the detectors.**

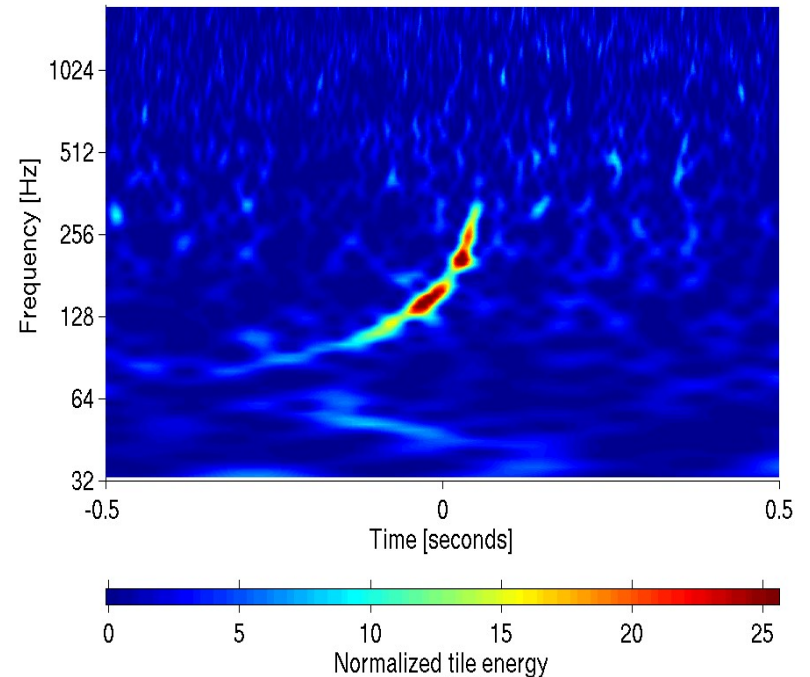
- A signal was observed by several methods on Sept. 16, 2010. Subsequent analysis suggested it was a binary coalescence involving at least one black hole, with a $1/(7000 \text{ yr})$ false alarm rate.

- The Blind Injection Envelope was opened on March 14, 2011 revealing the Sept. 16, 2010 event was the fake injection of a neutron star – black hole coalescence signal.

- See: <http://www.ligo.org/news/blind-injection.php>;
<http://www.ligo.org/science/GW100916/>

What will a detection look and sound like?

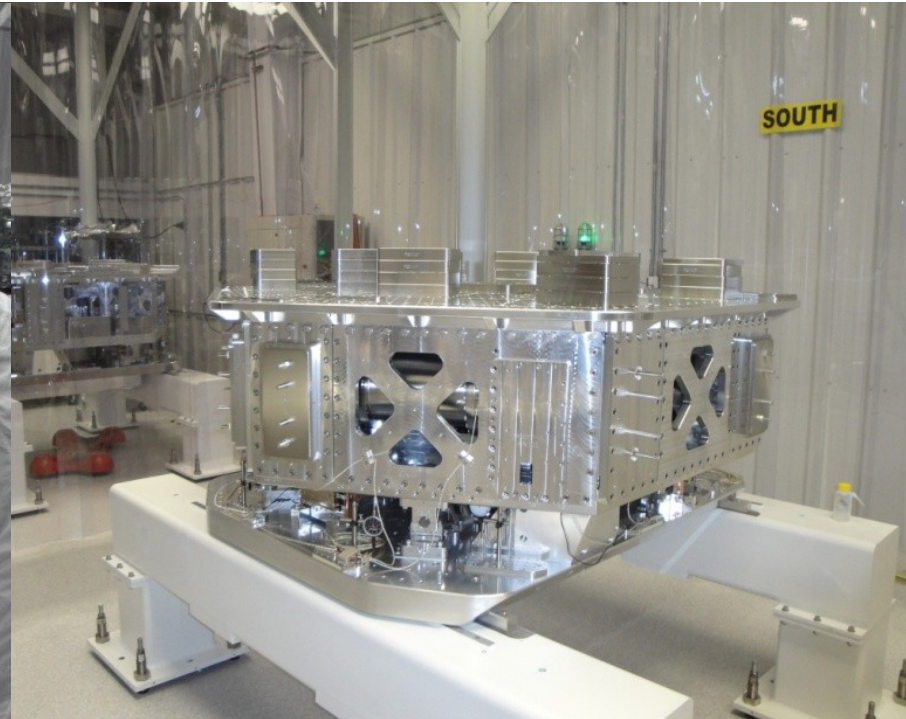
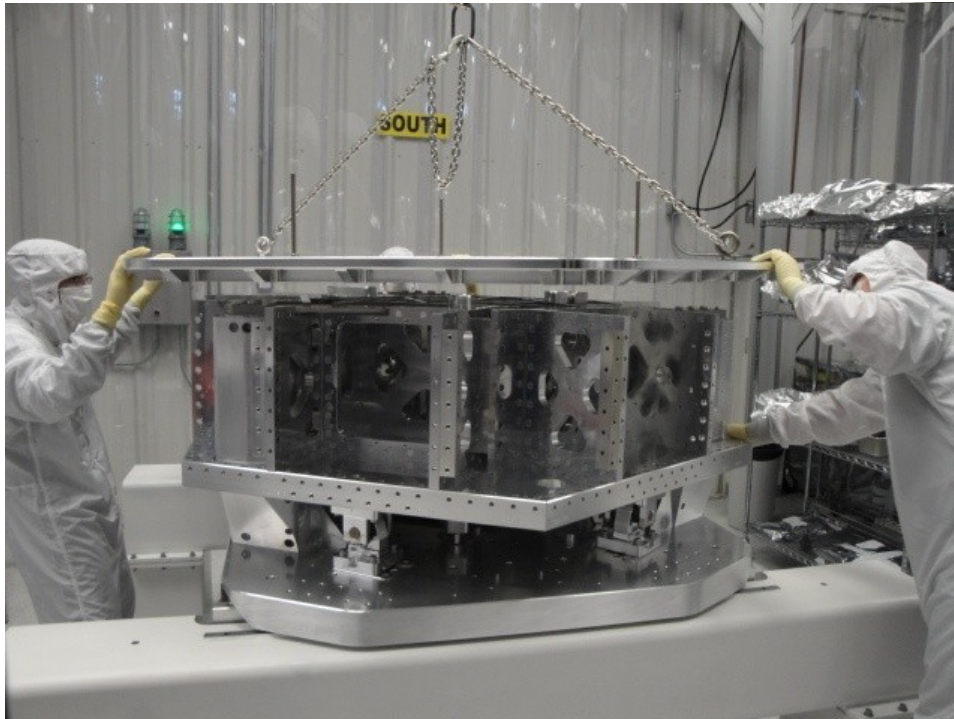
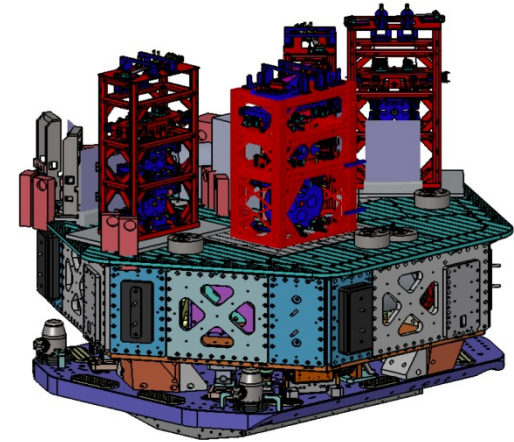
H1:LDAS-STRAIN at 968654557.955 with Q of 22.6



Data and Sound File from LIGO Hanford Observatory with **blind hardware injection of a fake signal.**

Advanced LIGO Seismic Isolation

- Assembly of the Horizontal Access Module stacks is in full swing at both observatories.
- Active feedback control will be used.
- One assembly already was used in the Enhanced LIGO configuration.



Suspension Systems

Initial Single vs. Advanced Quad Pendulum

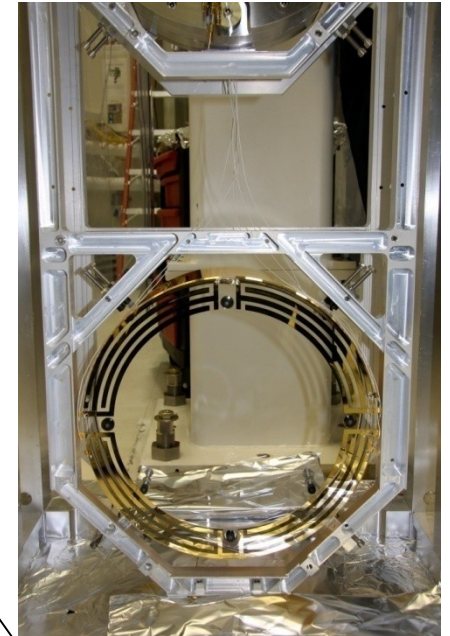
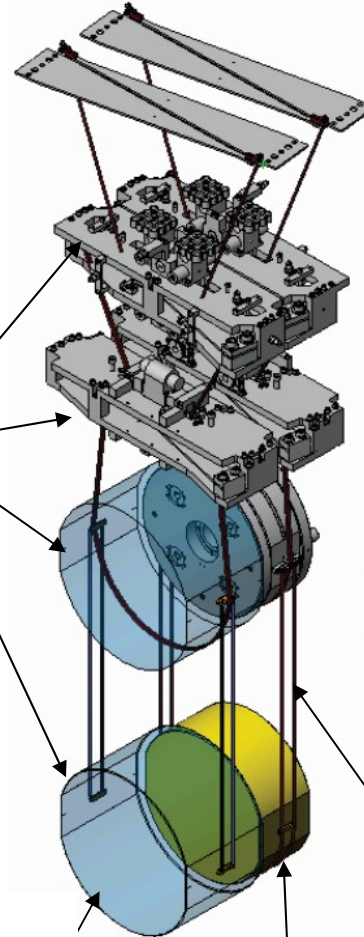
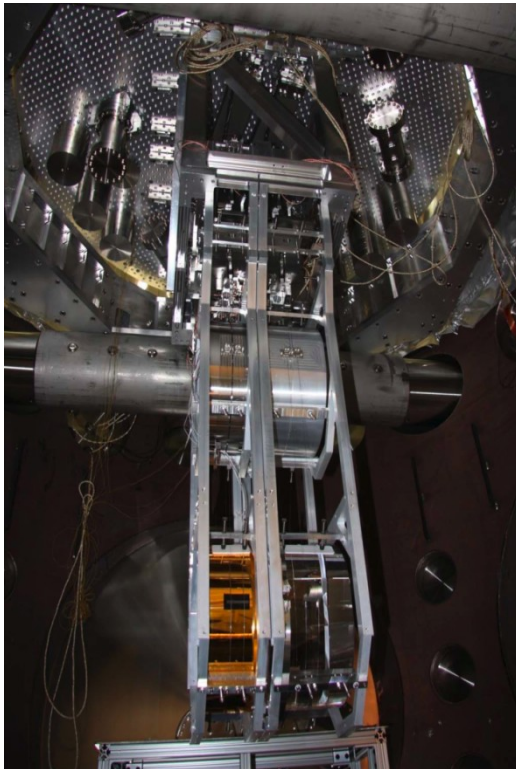
Electro Static Drive (ESD) on last stage:
Reduces noise from electromagnets

four stages

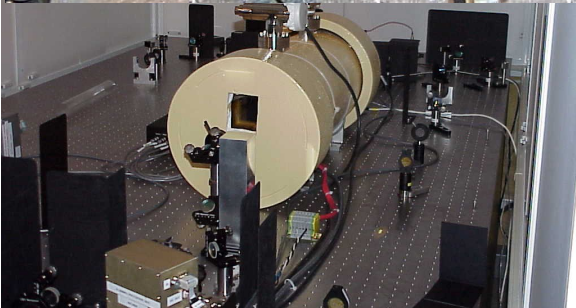
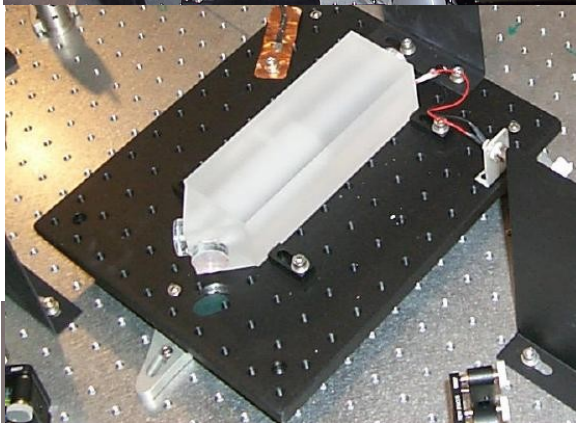
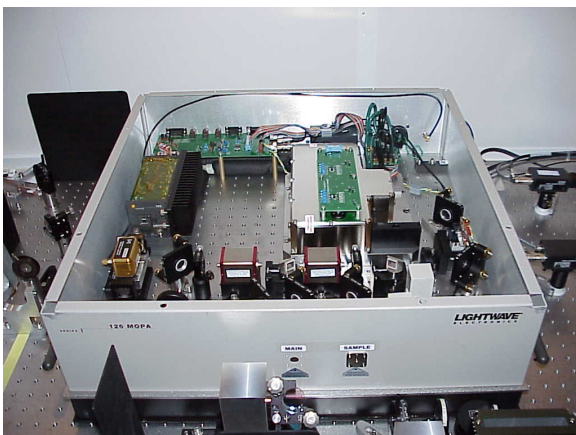
40kg silica masses

parallel reaction chain for control

silica fibers



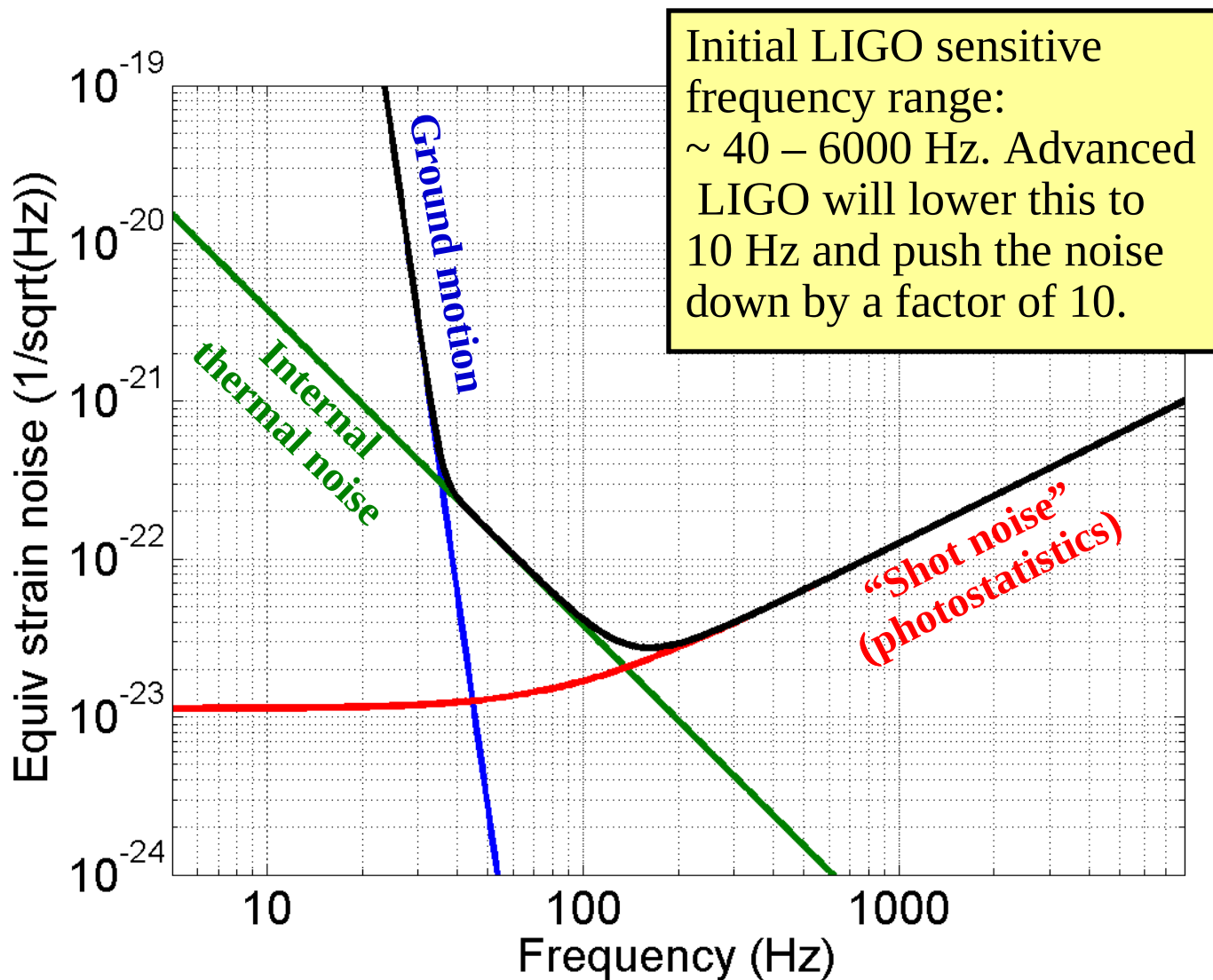
LIGO Nd:YAG Lasers: Initial LIGO 10W; Enhanced LIGO 35W; Advanced LIGO 150W



- Nd: YAG
Neodymium-doped yttrium aluminum garnet.
- 1064 nanometers
= infrared
- Stable to 1 part
per million at 100
Hz.



Limiting Sources of Noise





The End

LIGO & Gravitational Waves

Gravitational waves carry information about the spacetime around black holes & other sources.

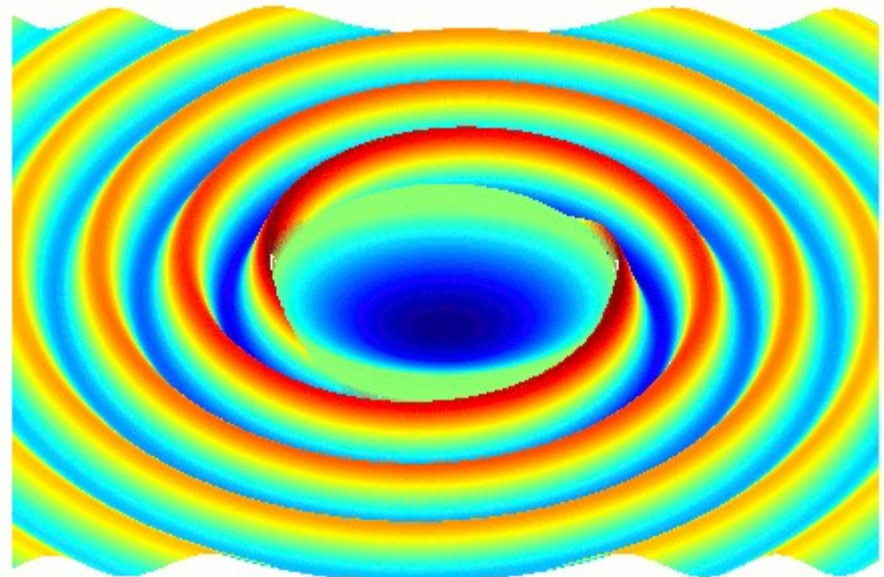
$$h_{\mu\nu}^{TT} = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & h_+ & h_\times & 0 \\ 0 & h_\times & -h_+ & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} e^{2\pi i f(t-z/c)}$$

$$dT^2 = g_{\mu\nu} dx^\mu dx^\nu$$

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$\left(\nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) \bar{h}^{\mu\nu} = 0$$

$$h_{\hat{\theta}\hat{\theta}}^{TT}(\theta = \pi/2) \propto \frac{1}{r} \cos[2\pi f(t - r/c) + 2\phi]$$



Detector Response

$$g_{\mu\nu} dx^\mu dx^\nu = 0 \quad (\text{Light Travels On Null Geodesics})$$

$$c^2 dt^2 - \begin{pmatrix} dx & 0 & 0 \end{pmatrix} \begin{pmatrix} 1+h_{xx} & h_{xy} & h_{xz} \\ h_{yx} & 1+h_{yy} & h_{yz} \\ h_{zx} & h_{zy} & 1+h_{zz} \end{pmatrix} \begin{pmatrix} dx \\ 0 \\ 0 \end{pmatrix} = 0$$

$$c^2 dt^2 = (1+h_{xx}) dx^2$$

$$c \int_0^{\Delta t} dt = \int_0^L \sqrt{1+h_{xx}} dx \cong \int_0^L \left(1 + \frac{1}{2} h_{xx} \right) dx$$

$$c\Delta t = L_x = L + \frac{L}{2} h_{xx}$$

$$\frac{\Delta L}{L} = \frac{1}{2} (h_{xx} - h_{yy}) = F_+(\theta, \phi, \psi) h_+(t) + F_\times(\theta, \phi, \psi) h_\times(t)$$

Black Holes After 1960

- Kruskal-Szekeres
Coordinates, 1960
- Wormholes, Wheeler and
Fuller, 1962
- Black Holes, popularized by
Wheeler, 1968
- Penrose Process,1969
- Black Hole Evaporation,
Hawking, 1974
- Time Machines, Morris and
Thorne, 1988
- BH Information Theory?

Black Hole History

- Dark Stars, John Michell 1784 (Also Pierre-Simon Laplace, 1796)
- General Relativity, Einstein, 1915
- Spherically Symmetric Solution, Karl Schwarzschild, 1916
- Einstein-Rosen Bridge, 1935

$$E=mc^2$$

$$c^2\Delta t^2 = c^2\Delta T^2 + v^2\Delta t^2$$

$$m^2c^4\Delta t^2 = m^2c^4\Delta T^2 + m^2c^2v^2\Delta t^2$$

$$m^2c^4\Delta t^2/\Delta T^2 = m^2c^4 + m^2c^2v^2\Delta t^2/\Delta T^2$$

$$[mc^2/(1-v^2/c^2)^{1/2}]^2 = [mc^2]^2 + [mv/(1-v^2/c^2)^{1/2}]^2c^2$$

$$[mc^2 + 1/2mv^2]^2 = E^2 = [mc^2]^2 + p^2c^2$$

$$\text{For } v = 0: E = mc^2$$

Newtonian Momentum

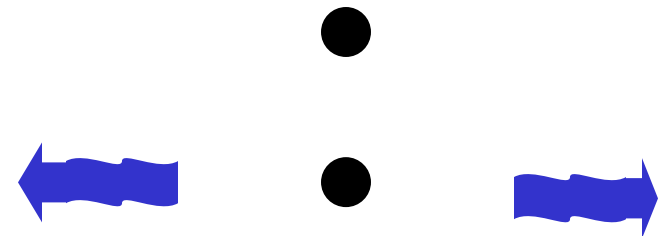
Approximate to order v^2/c^2 == Newtonian Kinetic Energy

E=mc²: What Einstein Said

“Does The Inertia Of A Body Depend Upon Its Energy-Content?” by A. Einstein, in “The Principle Of Relativity” translated by W. Perrett & G. B. Jeffery (Dover: 1952) from A. Einstein, Annalen der Physik, 17, 1905.

Consider a particle with energy U .

After it emits pulses of light with energy $0.5 E$ in opposite directions, the particles energy is H . Note that v does not change.



In a moving frame its energy is $U + K1$.



After it emits pulses of light with energy $0.5 \gamma E (1 \pm v/c)$ in opposite directions the particle's energy is $H + K2$. Note the relativistic blue/red shift factor is used.



By conservation of energy $U = H + E$ and $U + K1 = H + K2 + \gamma E$. Thus, $(\gamma-1)E = K1 - K2 = \Delta K$ so $\frac{1}{2} (v^2/c^2)E = \frac{1}{2} (\Delta m)v^2$ to lowest order. The particle lost mass Δm . For Δm max. equal to m : $E = mc^2$.

Einstein Wondered

- Einstein is famous for his thought experiments.
- In 1895, around age 16, he wondered, can we catch light?
- If yes, your image in a mirror would disappear. You would know your speed independent of any outside frame of reference. This would violate Galilean relativity.
- Einstein decides we cannot catch light; nothing can go faster than light.

The Speed of Light

$$c = 186,000 \text{ miles/s} = 670,000,000 \text{ miles/hr}$$

	v	Δt	ΔT
Car	60 mph	1 day	1 day - .35 nanoseconds
Plane	600 mph	1 day	1 day - 35 nanoseconds
Shuttle	17,000 mph	1 day	1 day - 28 microseconds
Voyager	38,000 mph	1 day	1 day - 140 microseconds
Andromeda	300,000 mph	1 day	1 day - 8.7 milliseconds
Electrons	99% c	1 day	3.4 hours



Photo: Stanford Linear Accelerator Center (SLAC); Public Domain

The faster you go the slower time goes!

Nothing can go faster than light!

Eddington Finkelstein Coordinates

If we introduce the following form of the Eddington Finkelstein time coordinate, t' ,

$$t = t' - (2GM/c^2) \ln |rc^2/(2GM) - 1|$$

outside the horizon, and

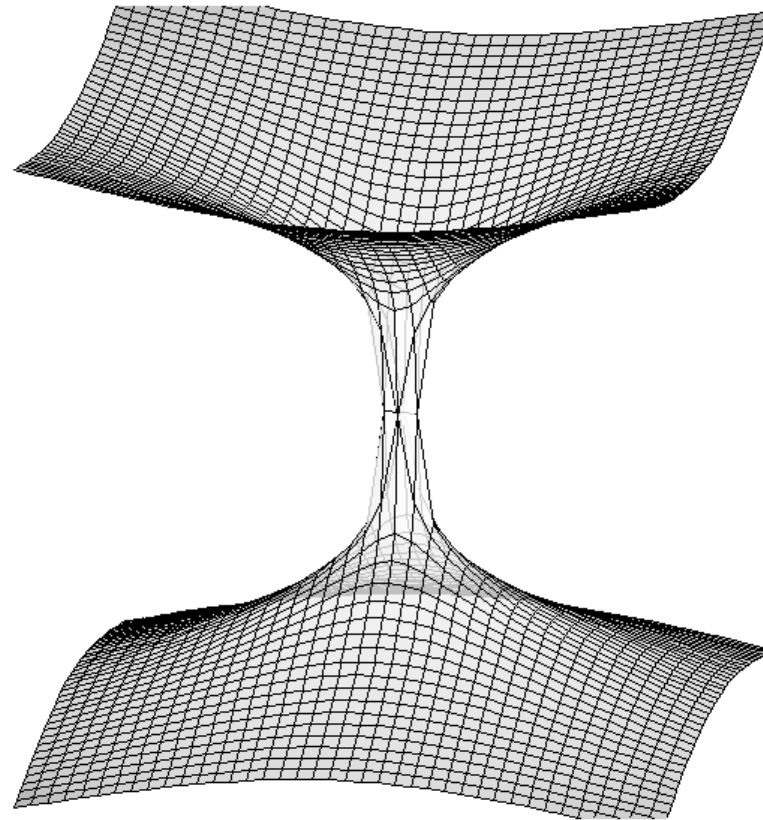
$$t = t' - (2GM/c^2) \ln |1 - rc^2/(2GM)|$$

inside the horizon, then inside or outside, we get

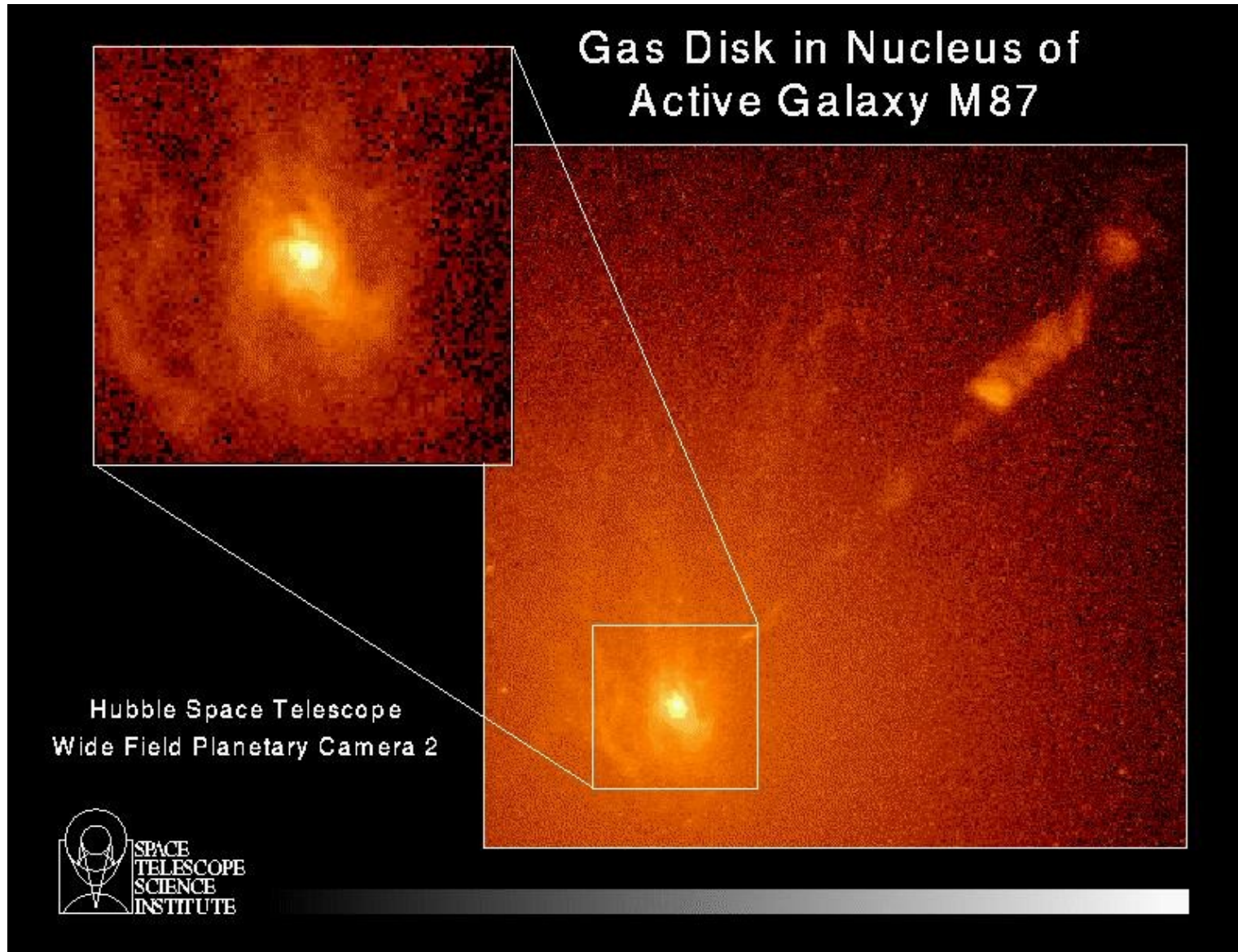
$$ds^2 = -c^2[1 - 2GM/(rc^2)]dt'^2 + 4GM/(rc^2)dt'dr + [1 + 2GM/(rc^2)]dr^2 + r^2d\theta^2 + r^2\sin^2\theta d\phi^2.$$

Note that there is no coordinate singularity at the horizon.

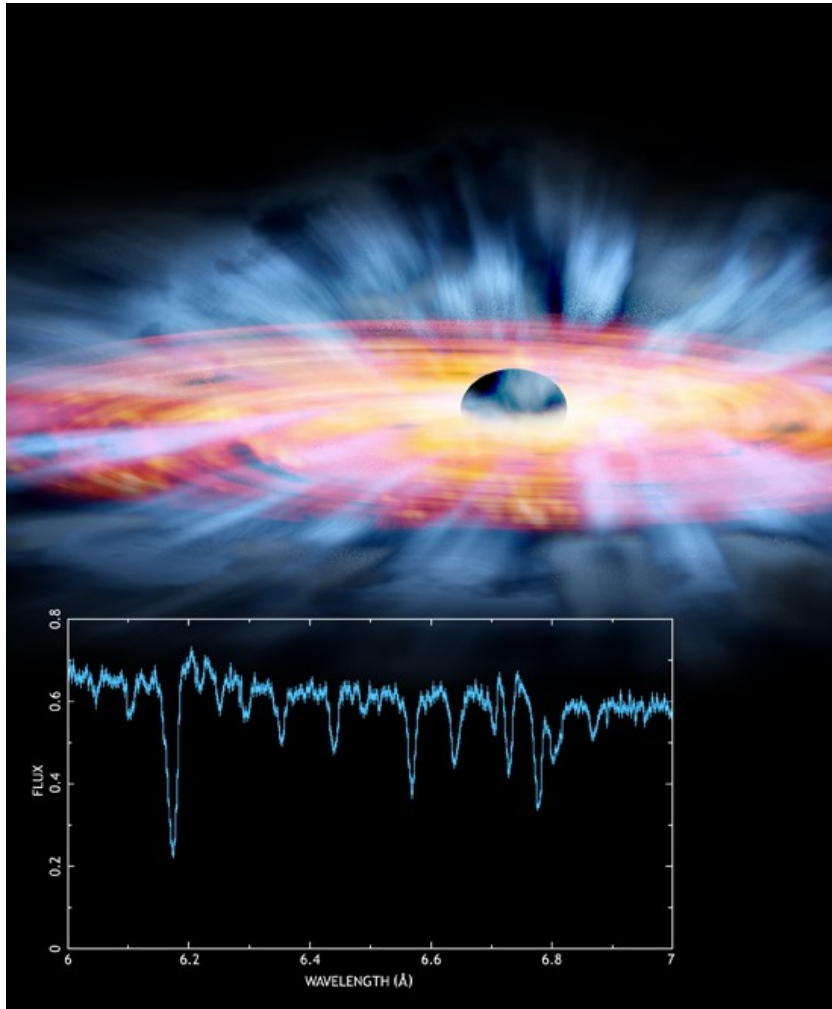
Schwarzschild Worm Hole



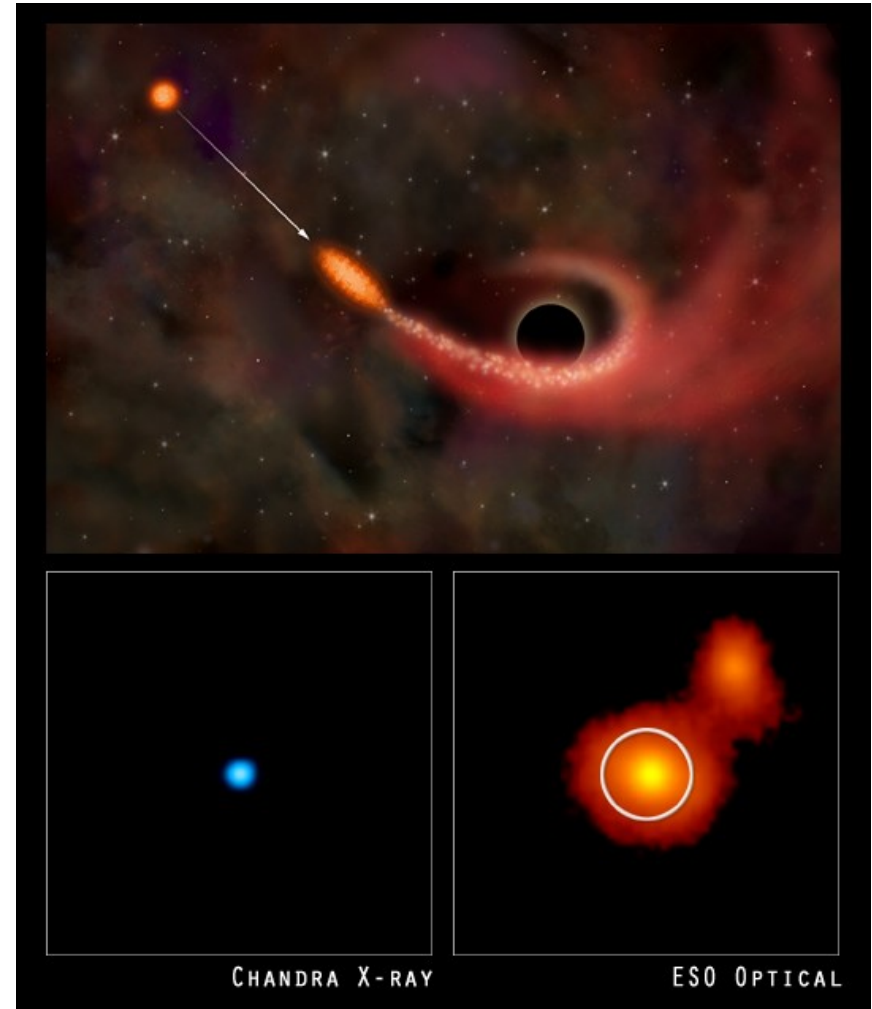
Black Hole Detection



Black Hole Detection



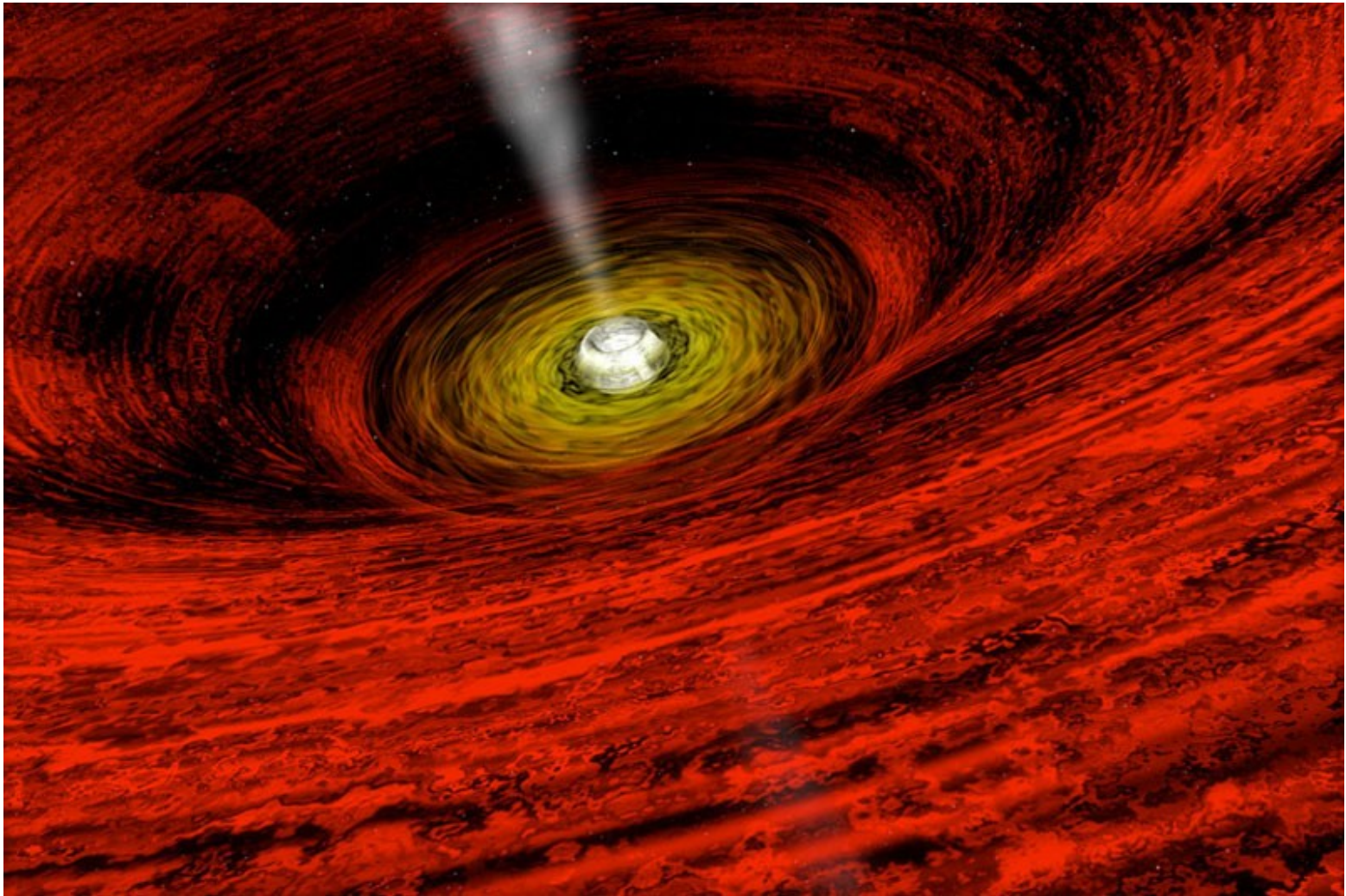
<http://chandra.harvard.edu/photo/2006/j1655/>; Credit:
Illustration: NASA/CXC/M.Weiss; X-ray Spectrum:
NASA/CXC/U.Michigan/J.Miller et al.



<http://chandra.harvard.edu/photo/2004/rxj1242/index.html>;
Credit: Illustration: NASA/CXC/M.Weiss; X-ray:
NASA/CXC/MPE/S.Komossa et al.; Optical:
ESO/MPE/S.Komossa

LIGO

Black Holes Detection



<http://antwrp.gsfc.nasa.gov/apod/ap060528.html>; GRO J1655-40: Evidence for a Spinning Black Hole; Drawing Credit: A. Hobart, CXC