

LIGO Advanced System Test Interferometer

PAC meeting 16 Nov 98
David Shoemaker

Initial steps

- System proposed at PAC MIT June 97
- Plan for vacuum system adopted by LIGO in Fall 97
- Space approved by MIT in Fall 97
- order placed Fall 97 with PSI (LIGO Vacuum Equipment vendor)

Since last PAC meeting

- envelope completed in May 98
- MIT Lab, high bay completed; Lab moved 15 July 98
- vacuum equipment in high bay 5 Nov 98
- seismic isolation equipment ordered; some delivered

Update on research and infrastructure

- activities at MIT, in LIGO Lab, in Ligo Science Collaboration
- LASTI: where it stands, where it is going

SCANNED

LSC Suspension/Isolation Working Group

2004: First opportunity for significant changes

- first science run finishing in ~2003
- allows a ~5 year cycle of research, development, engineering, test

Low-Frequency Performance goals dictated by thermal noise

- target materials (quartz, sapphire) loss characterized
- practical techniques needed for assembly
- isolation goal: let optimistic level for thermal noise dominate
- only moderate improvements in isolation required

Leads to near-term plan for research:

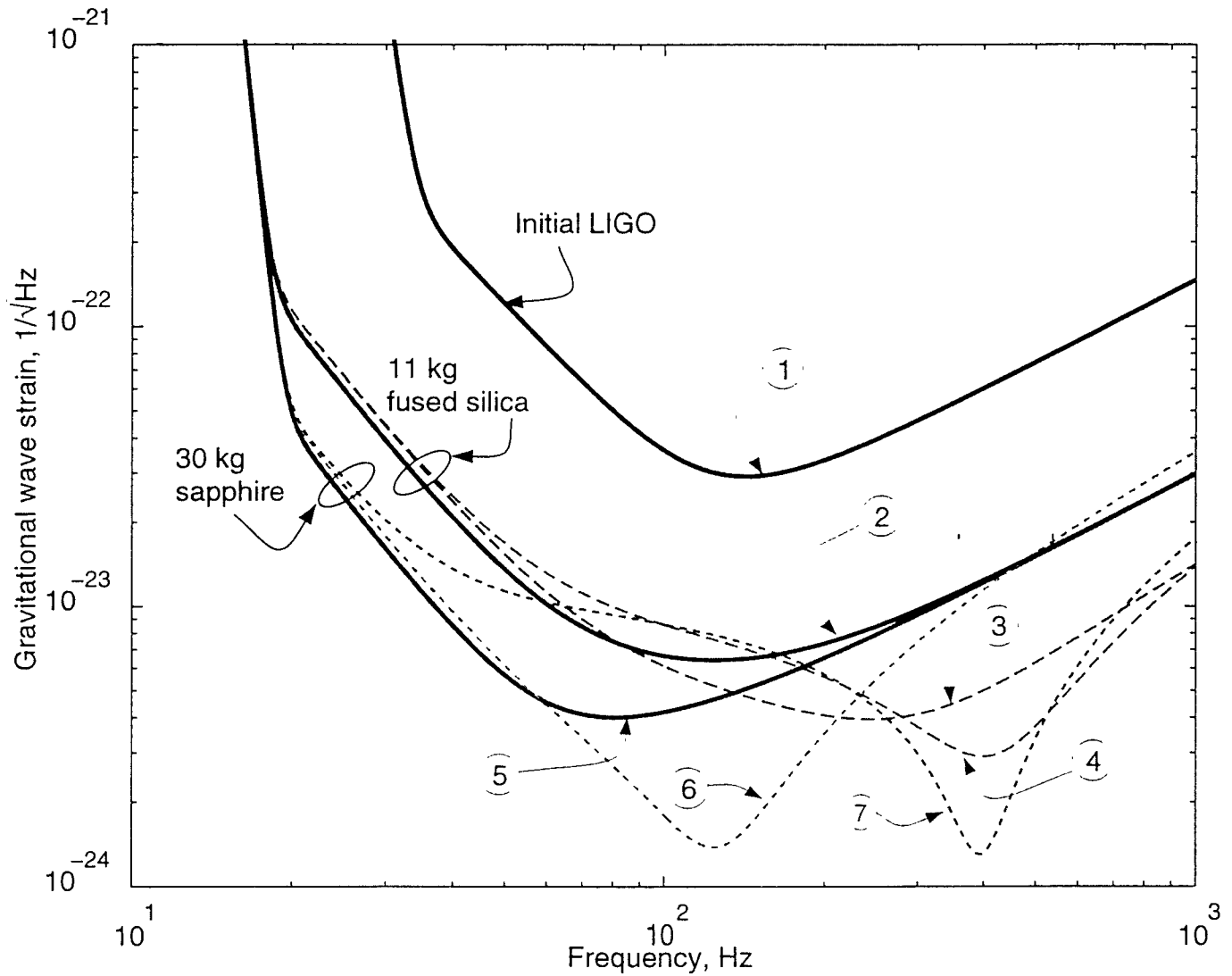
- model and plan the system: requirements, trades
- leave LIGO I passive isolation in place, augmented by...
- multiple-pendulum suspension, and
- modest active isolation system

Long-term goal: gravity gradient limited performance

- probably rather limited by pendulum thermal noise

interest in LIGO and community for aggressive application of sensing/servo techniques, but all approaches to be considered

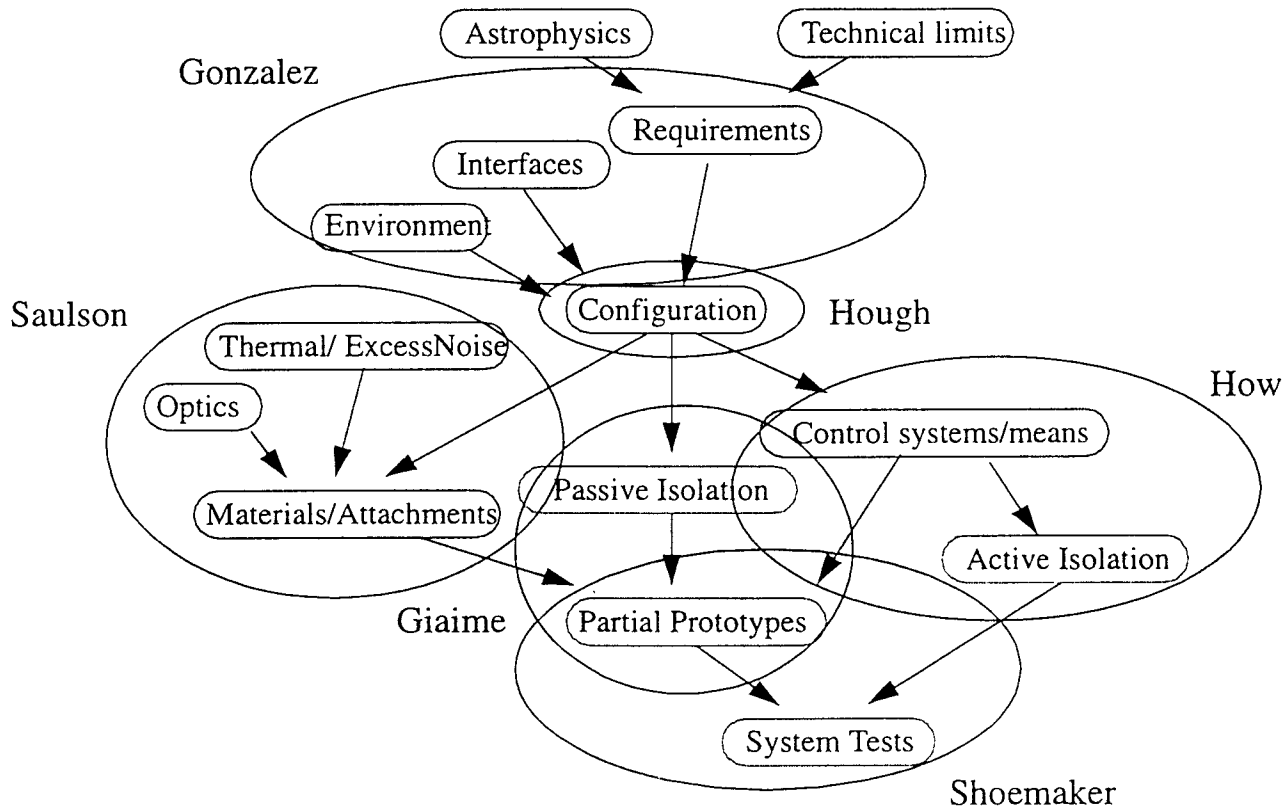
Performance goals



Considerable progress in defining strawman and models

- realistic parameters for losses
- design concepts for control hierarchy, mechanical assembly
- some staging of implementation possible (test mass material)

Roadmap to Stochastic forces research



LIGO Science Collaboration Suspension Working Group

- important force in planning for LASTI
- steering group meeting monthly, larger physical meetings ~2-3 months
- coordinated effort to develop near, far-term solutions
- here, discuss ways that the Full-Scale Test Interferometer contributes

Active Seismic Isolation

Plan to re-use the bulk of LIGO I passive isolation system

- with double pendulum, allows thermal noise to dominate in GW band
- considerable cost/complexity savings if incorporated in new design
- will be well-understood and characterized system
- changes (e.g., damping of springs) if indicated by experience

Active pre-isolation system

- Objective: to reduce required dynamic range in suspension; goal: factor 30 reduction, 0.1-10 Hz
- target design is (principally) external to vacuum
 - › may include sensors in vacuum to control stack dynamics
- MIT contributing commercial system for test to help make MIT LIGO campus site workable

Servo-control aspects

- fully digital system
- payload sensors, internal sensors, ground sensors
- commercial system comes with expert control development help

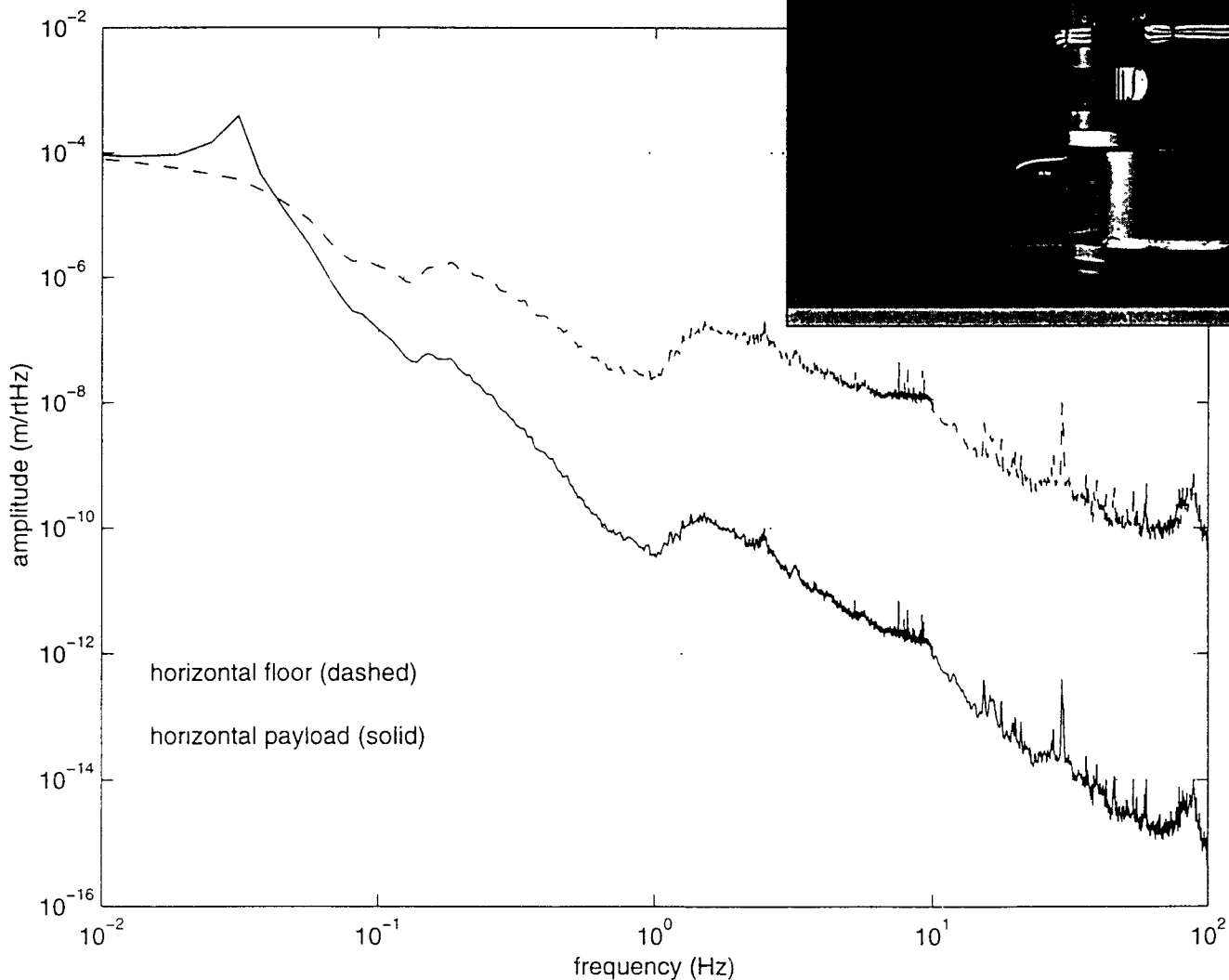
Goal: element of LIGO II isolation system

- off-load actuators on test mass (to zero!)

Active Isolation

Status

- working relationship established with Integrated Dynamics Engineering (IDE)
- in design, with some fundamental choices still unresolved
- MIT staff/faculty working with IDE to reduce noise in air-bearings
- possible change in design to horizontal-only air bearing
- December: fix the basic design choices



Passive Isolation

LIGO II objective: thermal/quantum noise to dominate

- put ‘seismic wall’ intercept $>10x$ above sensitivity limit
- combination of LIGO I stack, multiple pendulum, pre-isolator sufficient for horizontal isolation

Additional vertical isolation required

- earth’s curvature gives vertical-horizontal coupling
 - › best case: 3×10^{-4} ; difficult to measure, and likely greater
 - › LIGO I passive isolation $\sim 1000x$ worse in vertical than horizontal
- simple pendulum has $\sim 10\text{-}15$ Hz ‘bounce mode’; no real isolation
- additional isolation must be close to test mass to filter passive stack internally generated thermal noise
- development work in GEO and LIGO/Caltech

LASTI activities

- modify test-mass chamber isolation/suspension interface
- raise optics table without modification of stack
 - › spacers between support tubes and support table
 - › spacers between support table and first stack element
- tests in LASTI of target system for vertical isolation (once integrated with suspension)

Suspension Prototypes

Transfer of GEO double pendulum to MIT Lab

- fused-quartz system with masses comparable to LIGO II
- requires vertical compliant member
- requires instrumentation for testing

Separate test setup

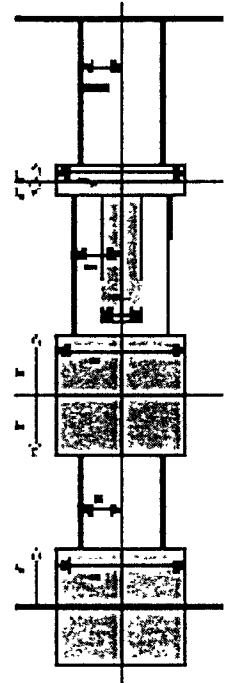
- in Phase Noise Interferometer tank
- allows early start
- allows 'dirty' tests and instrumentation

Measurement of transfer functions, cross coupling

- looking at particularities of fused quartz connections
- difficulties/improvements in static alignment strategy

Charge monitoring and control

- need to track static charge and time varying charge
- electrometer with oscillating probe near surface
- will probably need means to discharge in situ, non-destructively
- ultraviolet probably not permissible solution



System tests

Once design well developed, need end-to-end testing

- tests of modifications of interfaces to LIGO Seismic Isolation
- acquisition with coupled cavities; alignment control
- noise performance

Complete LASTI to enable this research

- isolation systems in Test Mass Chamber, HAMs
- active isolation customized for suspension, isolation designs
- fabrication of suspension systems

Tests as required to qualify design

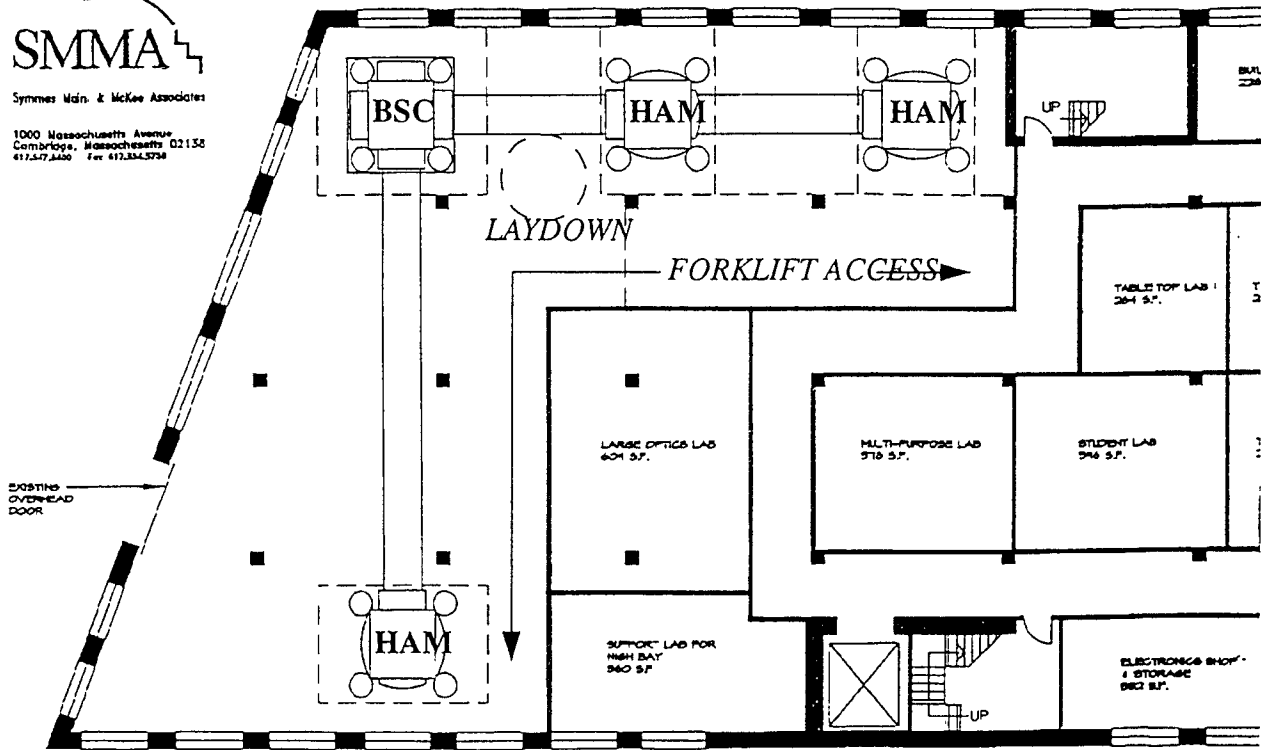
- transfer functions
- control issues: length [distribution of authority], alignment
- thermal noise (upper limits---small laser beam diameters)
- internally-generated noise
- installation problems/procedures
- reliability

Any optics/configurations issues

- as lab, LSC see appropriate

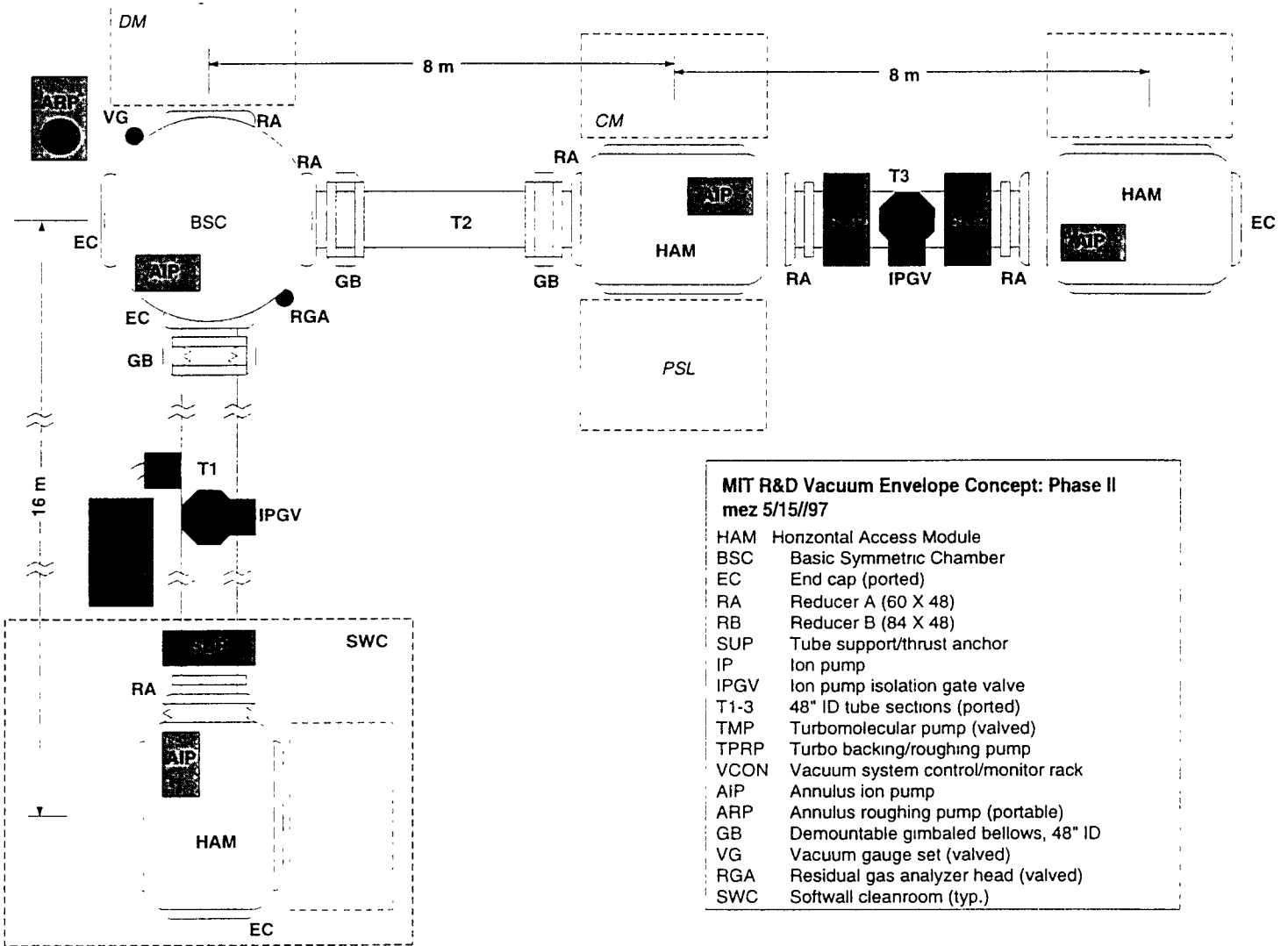
Test Interferometer

SMMA
Symmes Main & McKee Associates
1000 Massachusetts Avenue
Cambridge, Massachusetts 02138
617.547.8460 Fax: 617.554.3738



- Shown in place in new MIT Lab high-bay
 - › overhead clearance ~4' less than LIGO sites
 - › would require segmenting Test Mass Chamber cover for internals of maximum height; PSI sees no problem
 - › no problem with LIGO I/II seismic installation
- overhead crane for manipulating covers, tops
- control room, supporting labs on same floor
- staff, visitors' offices on two floors above

Vacuum system layout



LASTI Seismic Isolation elements

Objectives:

- give LIGO-like environment for tests (LIGO II, possibly LIGO I)
- keep costs for re-engineering/startup minimal (add on to LIGO orders)
- not interfere with LIGO fab/install (put at end of queue)

BSC (test mass chamber):

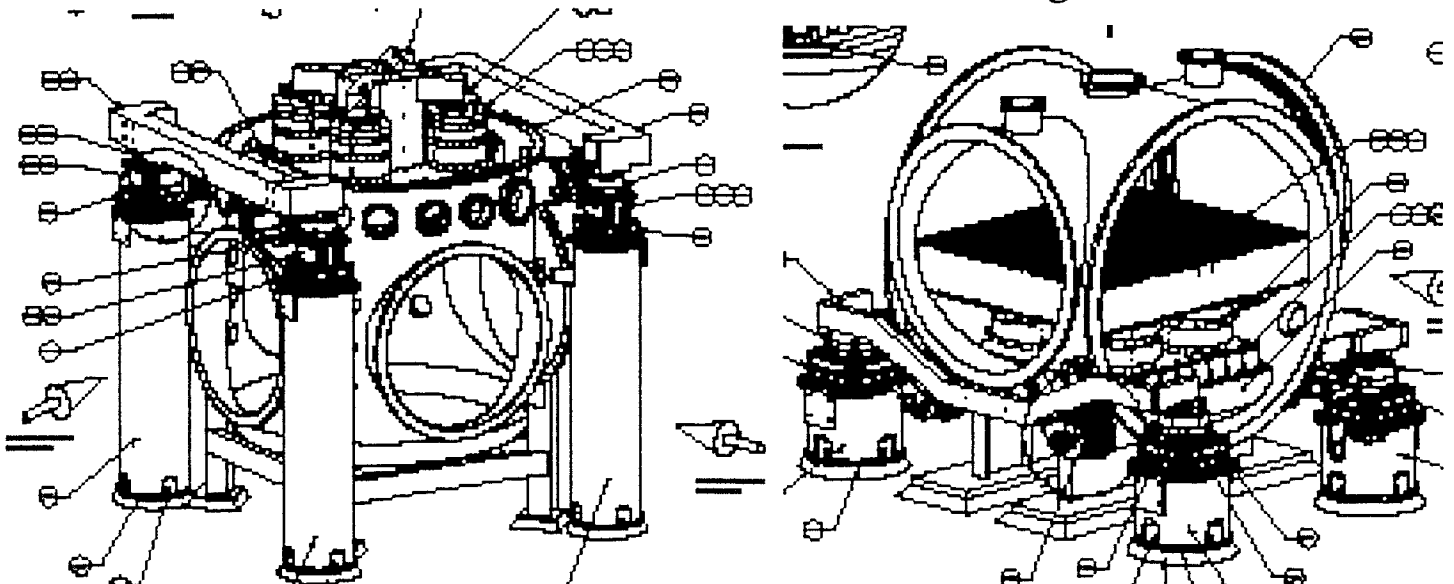
- complete support and isolation system, except
- no 'fine actuator' or 'coarse actuator' (replaced by IDE system)
- additional spacers to create headroom for multiple suspension

HAM #1:

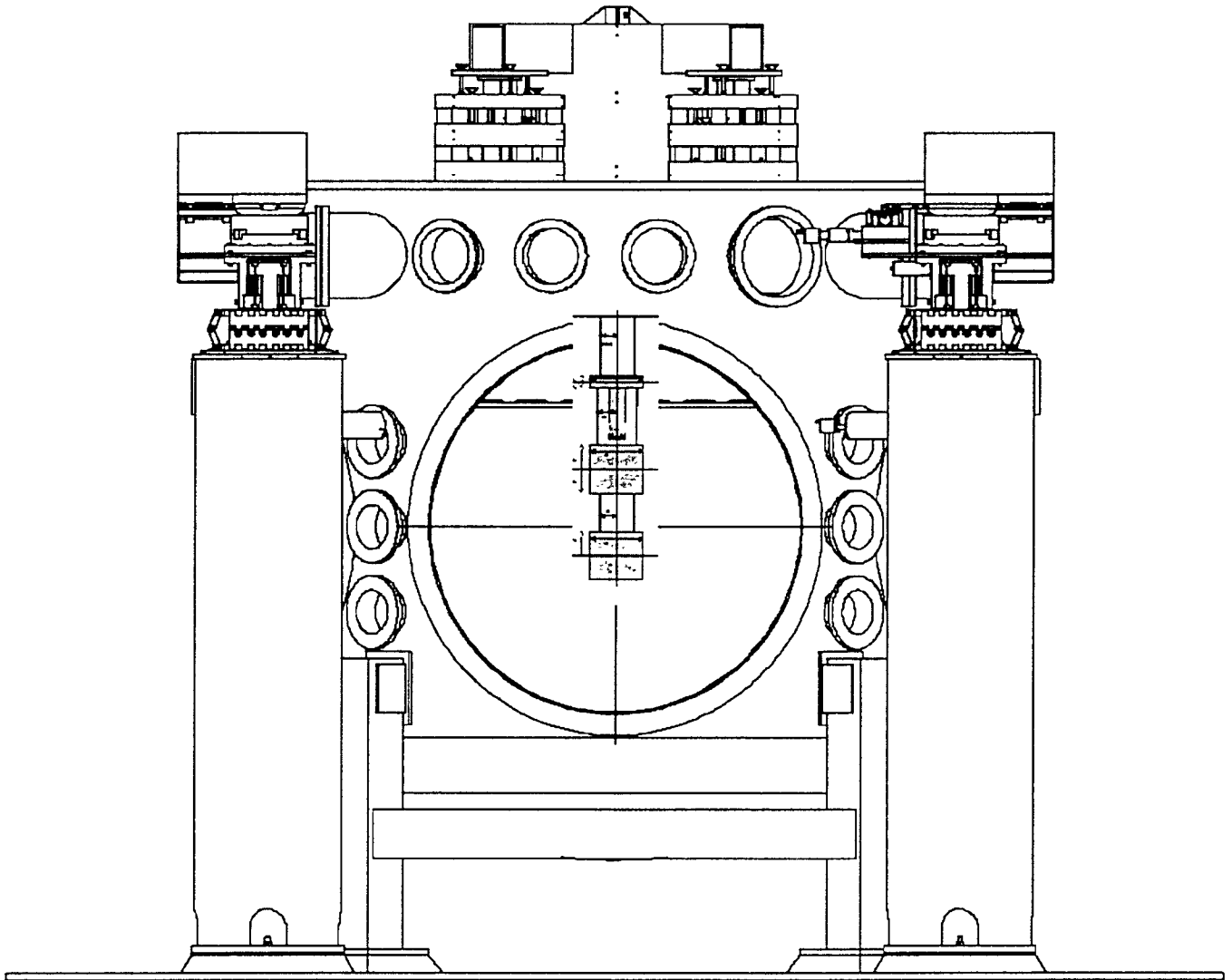
- complete support and isolation system

HAM #2, 3:

- complete support system and optics table; no masses, springs
- could retrofit if indicated, or substitute alternative designs



View of BSC, GEO Suspension



Scheduling/staffing

LASTI schedule tied to LSC Suspension development plans:

LSC Suspension Group	Epoch	LASTI
Top-level Requirements review	4th qtr '98	start envelope installation
Conceptual design; Design summit	1st qtr '99	complete envelope installation; start pier installation
Internal requirements, component research, configuration trades; Preliminary Design Review	2nd qtr '99	start seismic isolation installation; start prototype fab
	4th qtr '99	pre-isolation functional; start partial prototype tests
Component Prototypes; Test Review	3rd qtr '00	add isolation/infrastructure as needed for system tests
Fabrication of Engineering Prototypes	1st qtr '01	install engineering prototypes
System tests; Final Design Review	3rd qtr '02	...start to plan for LIGO III

Staff

- Joe Giaime (50%), Andrea Stolte (grad), post-doc (search), dhs
- Mike Zucker, tech help on vacuum envelope design/install/shakedown

State of the Art

