



HAM **Passive** Seismic Attenuation System (SAS) System Performance, Fabrication, Assembly , Installation

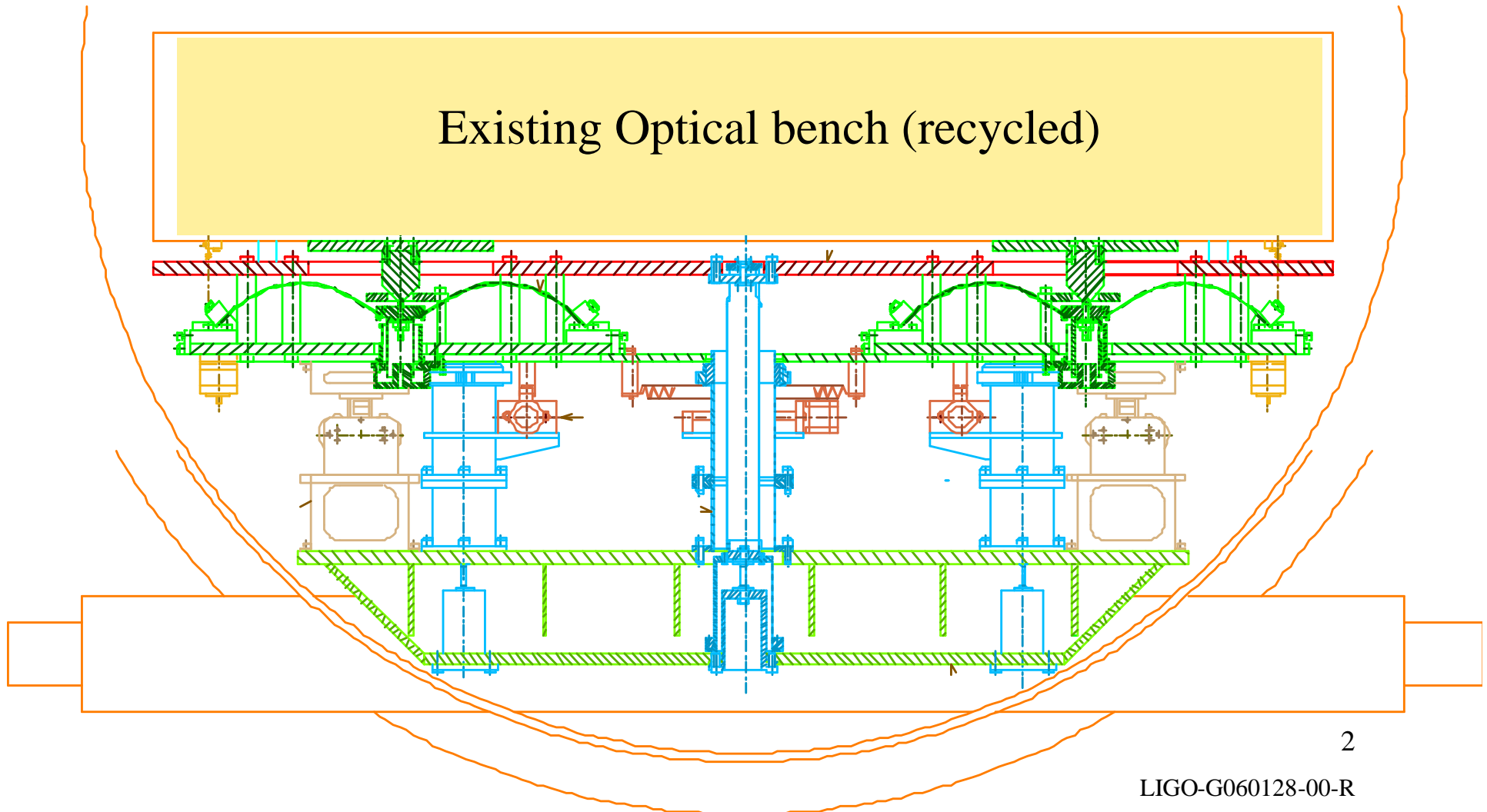
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Yumei Huang, Virginio Sannibale
LIGO Gravitational Wave Observatories
California Institute of Technology



A seismically attenuated optical bench for the HAM chambers

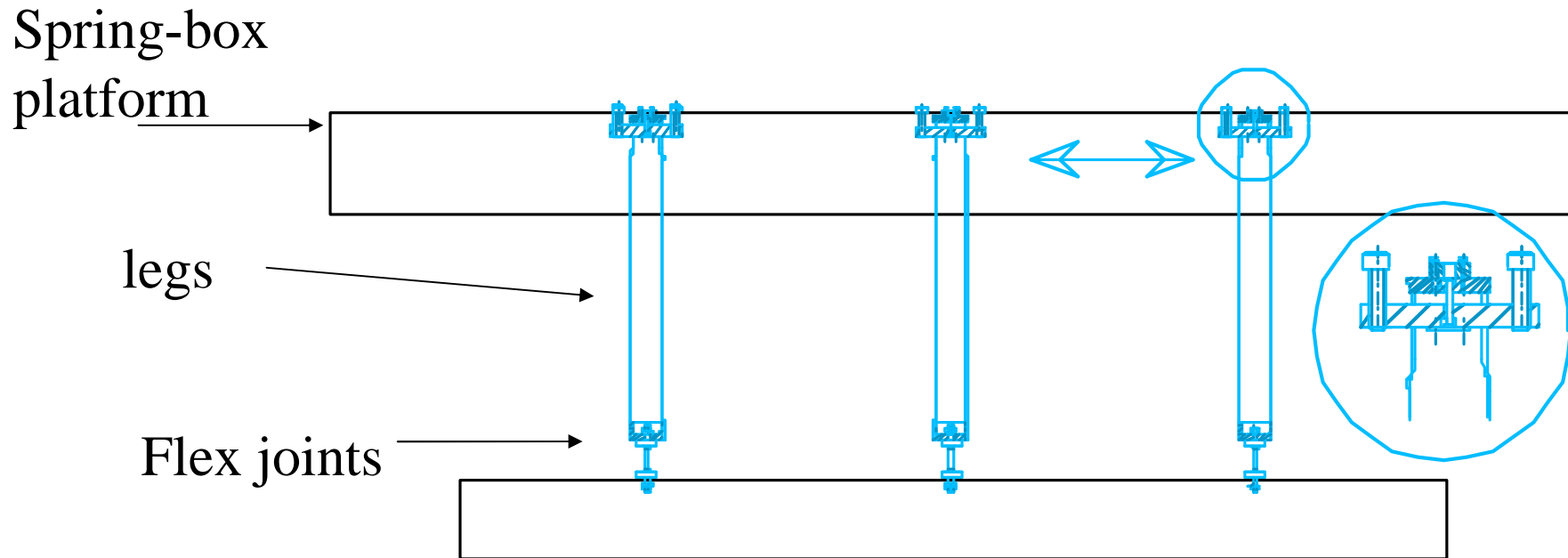


Existing Optical bench (recycled)



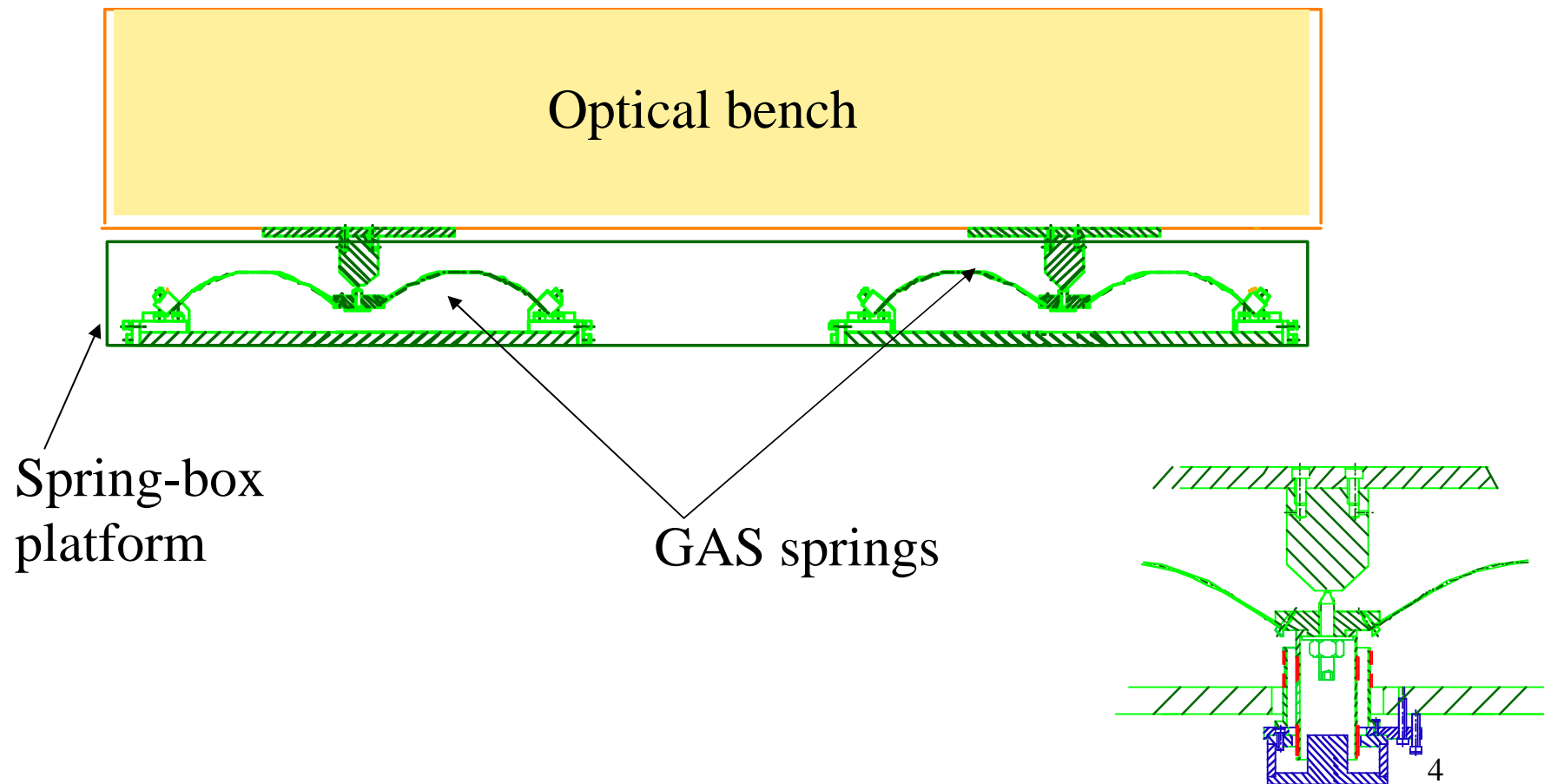


Horizontal direction, x, y, φ the **Inverted Pendula**





Attenuation in the vertical direction, the GAS springs





Performances

- Deliver more than 60 dB attenuation at > 1 Hz
- Single, passive layer attenuation to satisfy requirements and minimize complexity
- Significant attenuation at the micro seismic peak
- Internal damping for minimized control burden
- Tidal control with pointing accuracy at \sim nm level
- No standing control forces
- Earthquake protection for up to ± 12 mm shakes



Performances



- **Reliability !!!**
 1. No active components in vacuum
 - Only coils in vacuum
 - No electronics failures in vacuum !!!
 - No power dissipation under vacuum !!!
 2. No sealed gas volumes in vacuum
 - No chance of crippling virtual leaks !!!
 3. Four actuator/sensor groups for each three d.o.f.
 - functions unimpeded by one sensor/actuator failure
 4. Functionality unimpeded by power losses
 - (earthquake protection)



Difficulties



- Careful tuning and weight watching required.



Assembly philosophy

- Clean assembly and factory tuning
 - Minimize expense of LIGO manpower
 - Training fabricators to our procedures
- Shipping clean assembly
 - Develop clean installation techniques in HAMs
 - Install populated optical bench

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Illustration of vertical attenuation



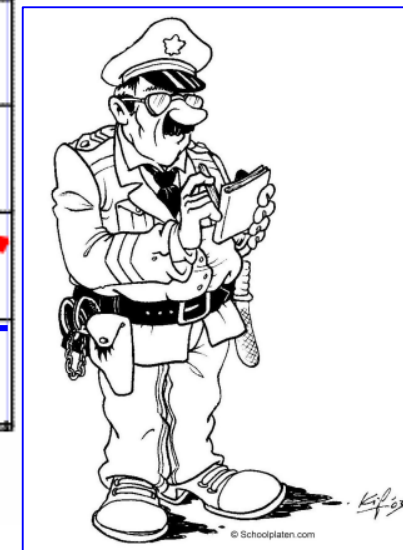
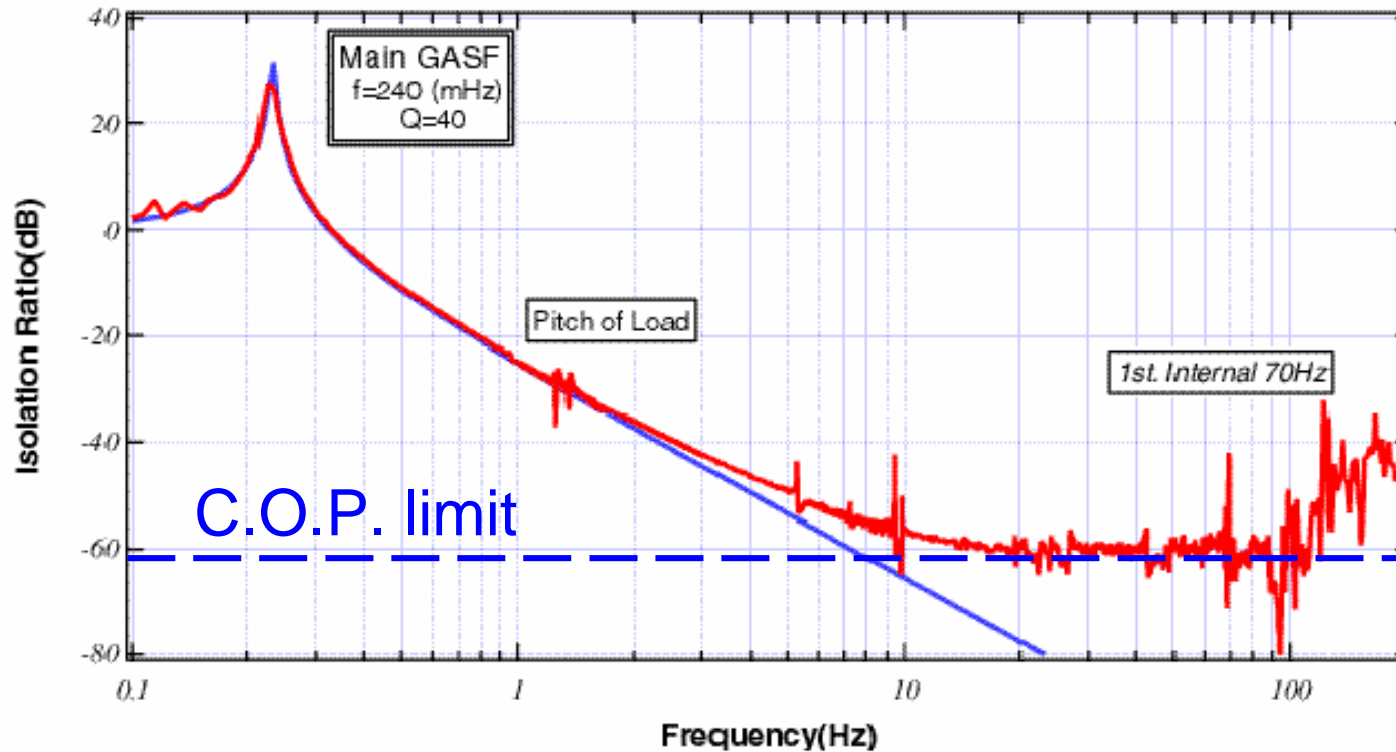
QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.



Typical GAS performance

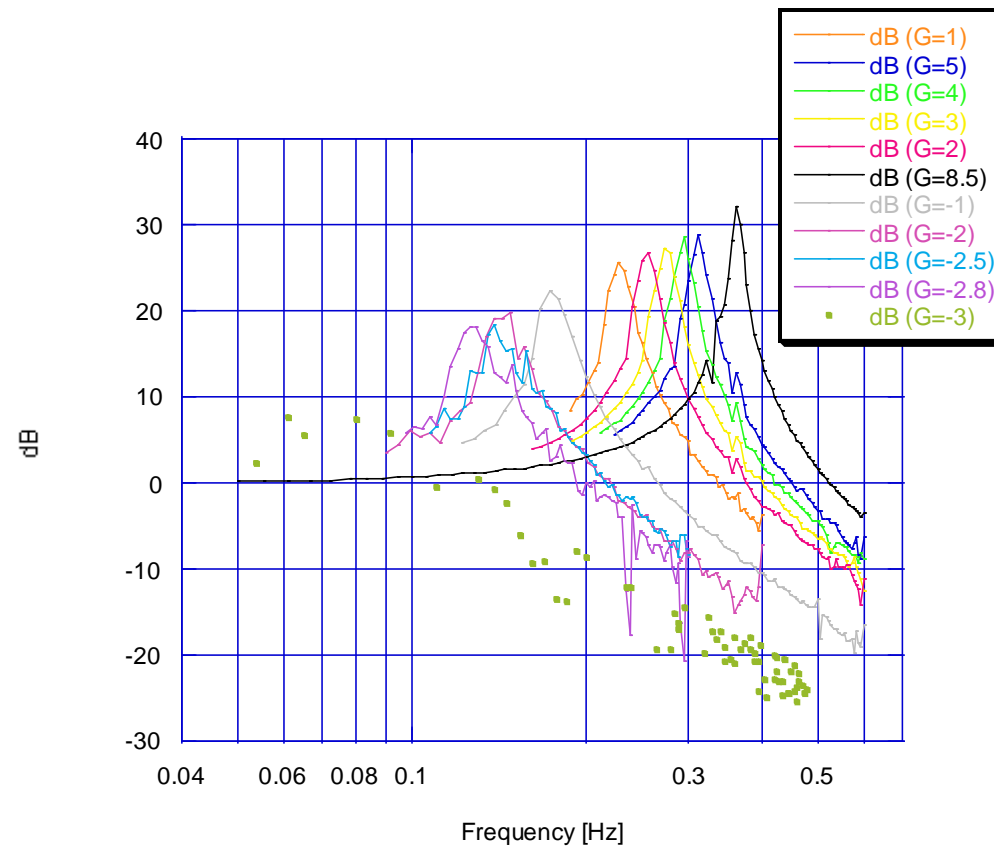
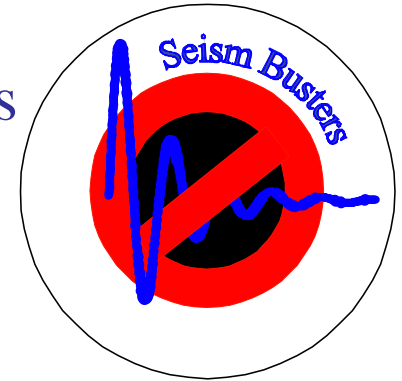


- LF tune sets the attenuation startup
 - Can tune to 30 mHz
- 60 dB c.o.p. attenuation limit
 - Can exceed with CW.





Transfer Function with Different Gain values



Lowering the system stiffness to 30 mHz

The Transfer Function shifts to lower frequencies,

The Q factor decreases



LIGO Origin of Q-factor decrease advantage



- The GAS process cancels the return forces
- Reduces the oscillation frequency ω
- Not the dissipation processed
- Energy dissipation_(per cycle) $\delta E = \text{constant}$
- Kynetic Energy $E_k \sim \omega^2$
- $Q = E_k / \delta E \Rightarrow Q \rightarrow 0 \sim 1 / \omega^2$

No damping required !

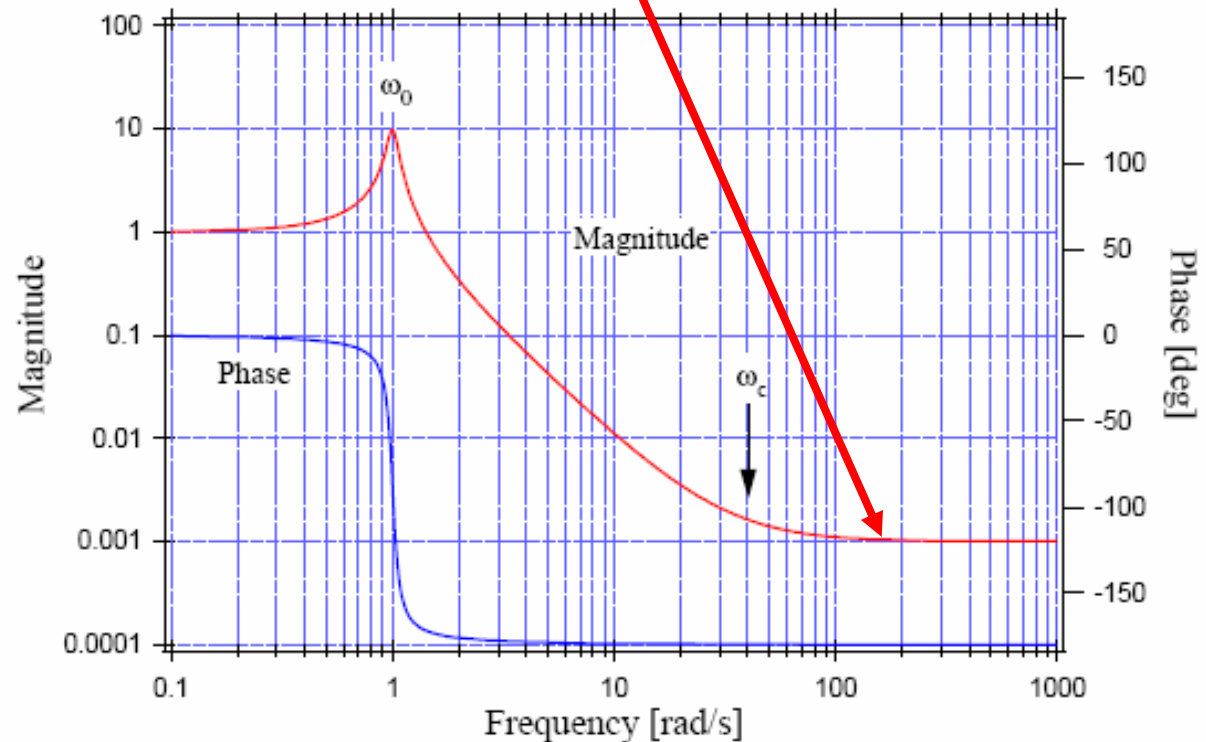


Center Of Percussion effect



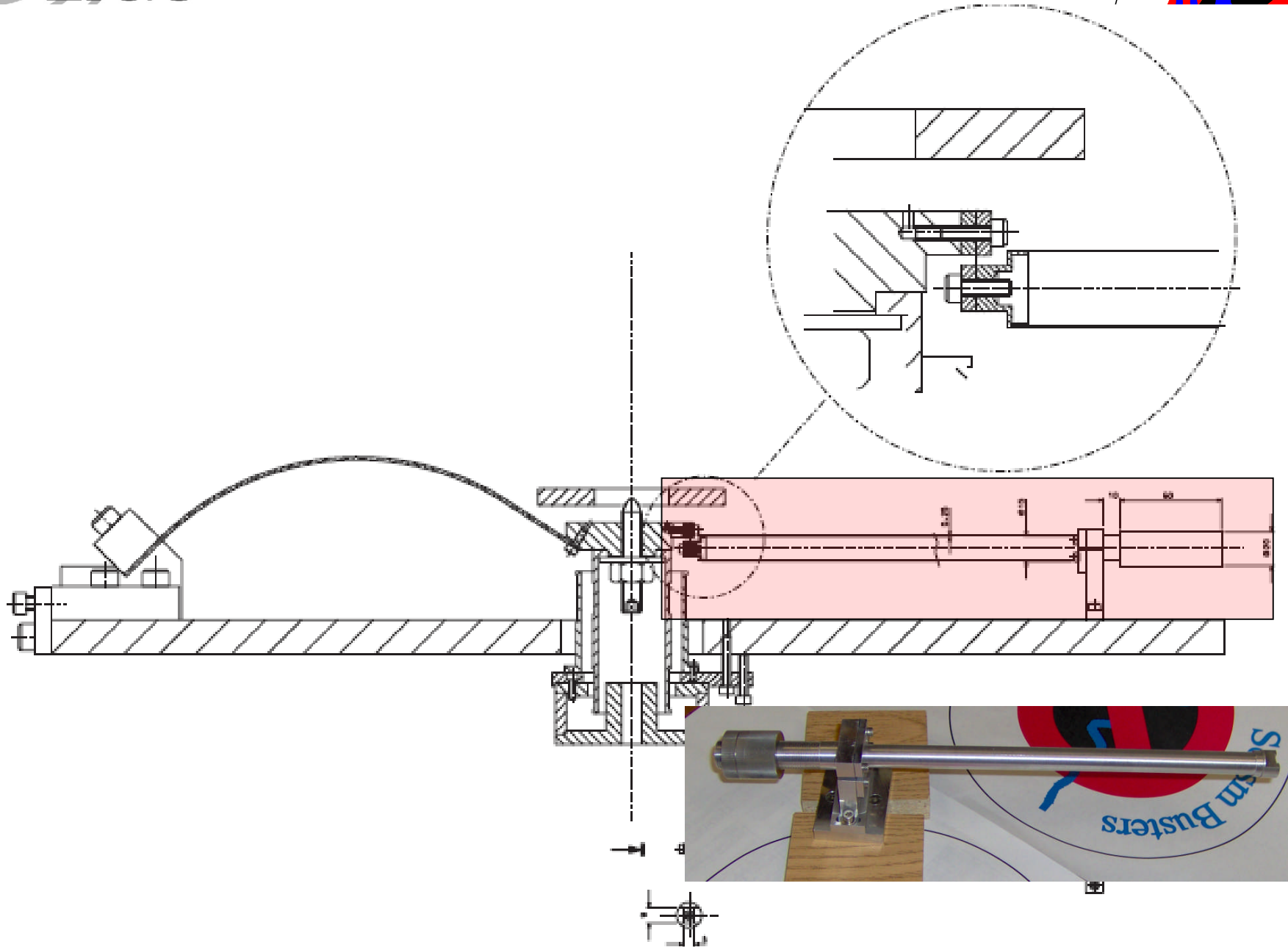
$$H_x \equiv \frac{\omega_0^2 (1+i\phi) + \frac{M}{m} \omega^2}{\omega_0^2 (1+i\phi) - \omega^2 + i \frac{\gamma}{m} \omega}$$

- Mass term limiting the attenuation performance





Magic Wand Implementation

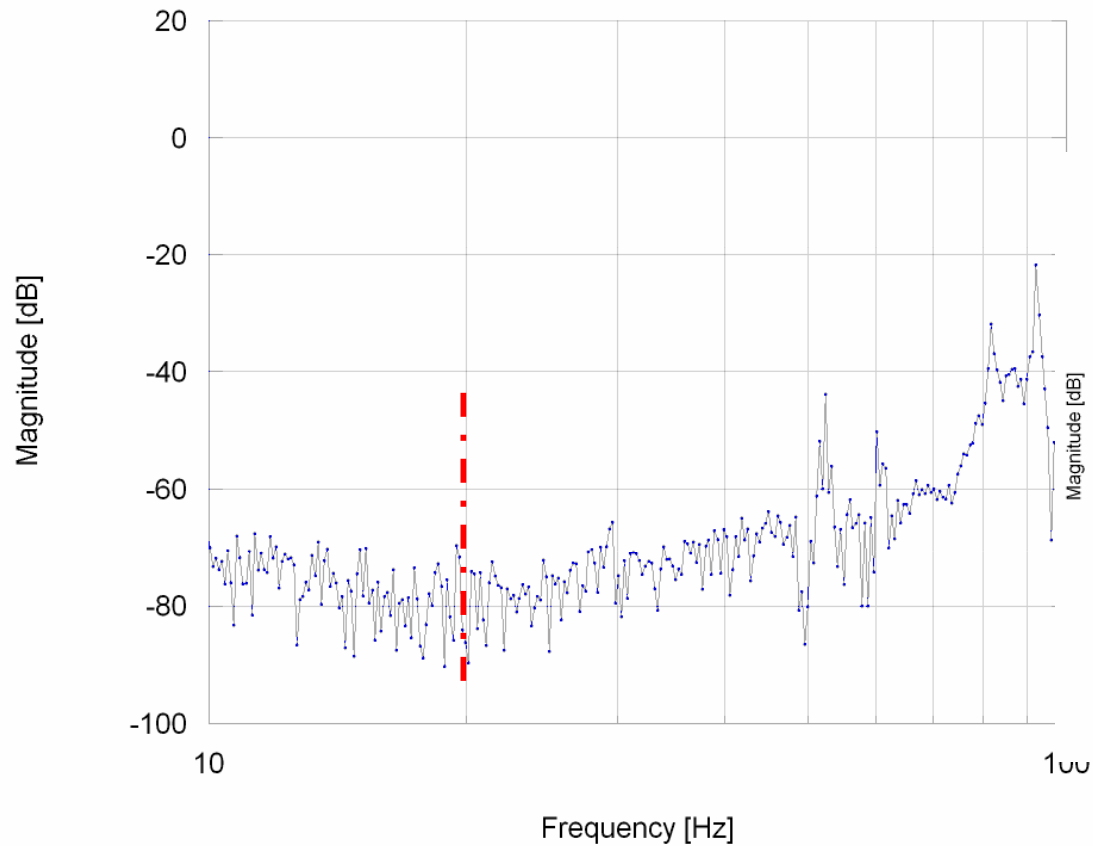




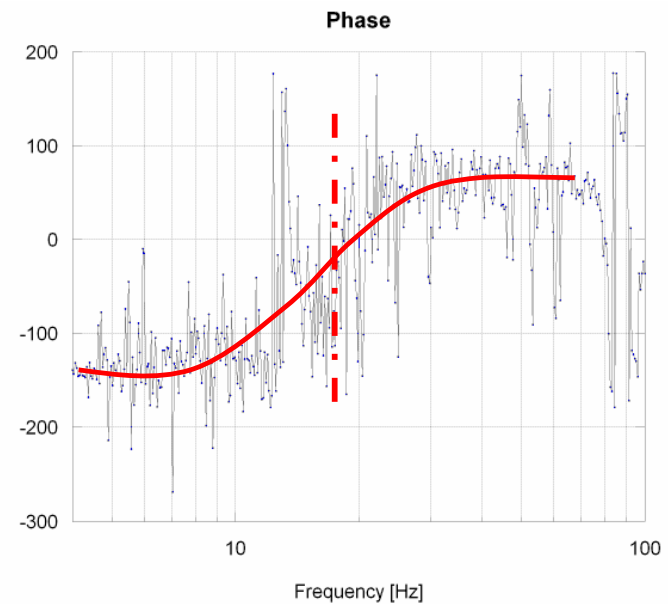
Magic wand tuning 80 dB GAS springs vertical isolation



Transfer Function - Two Booms 1 CW Inner

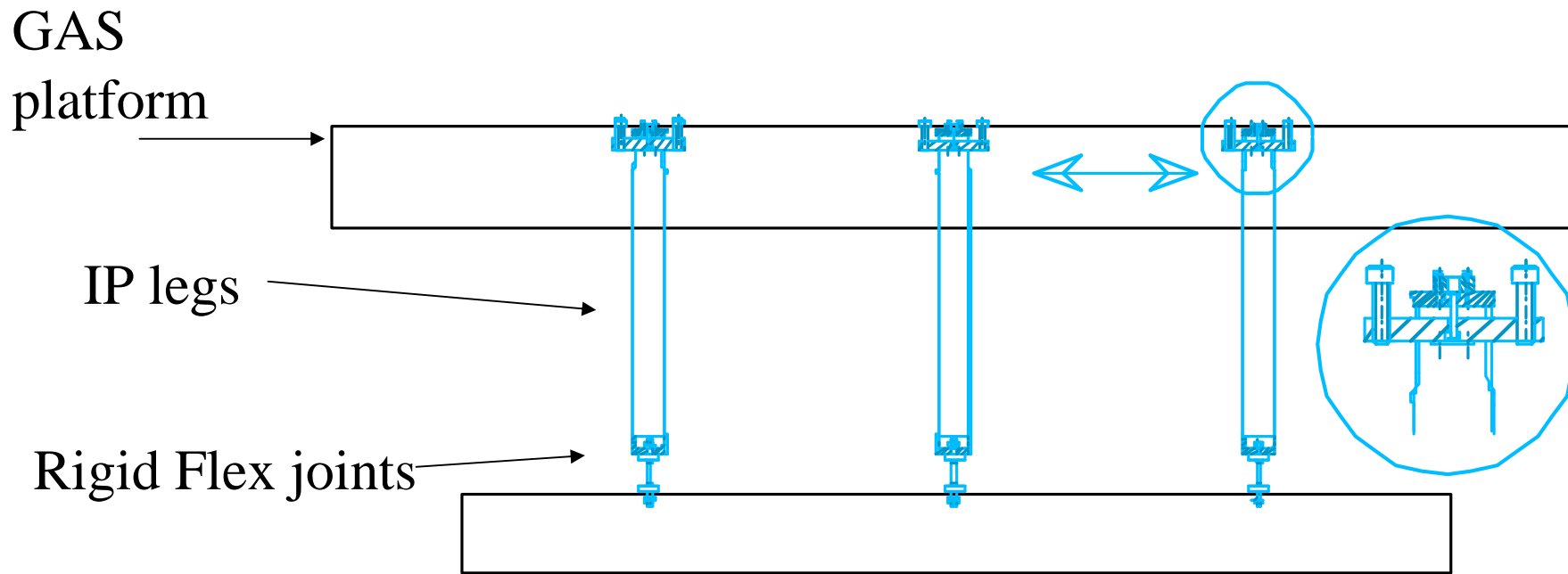


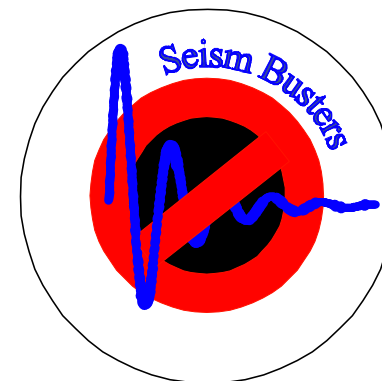
- (Still overcompensated)





Horizontal direction, x , y , ϕ the Inverted Pendula

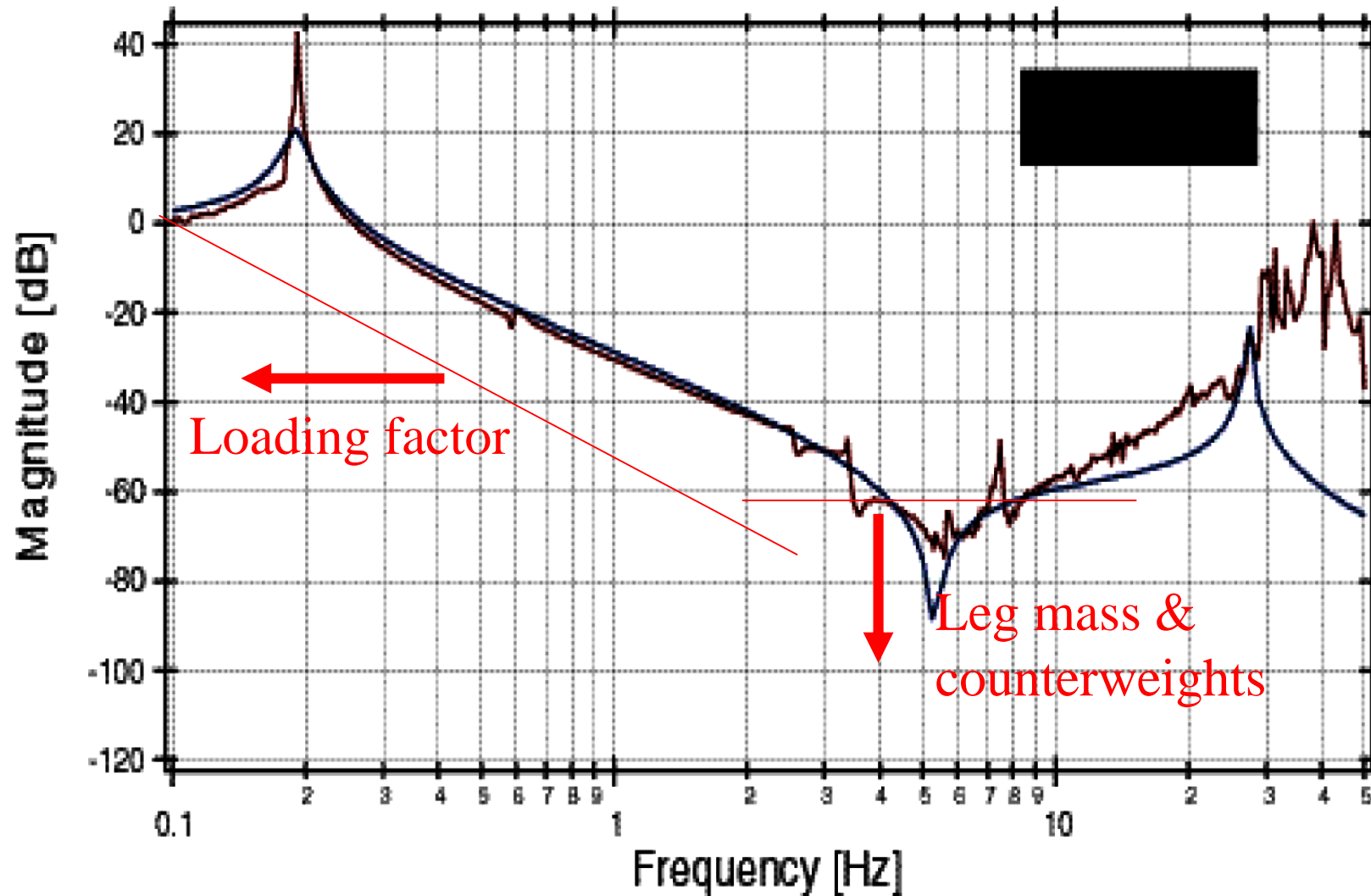




QuickTime™ and a
YUV420 codec decompressor
are needed to see this picture.



Typical IP Horizontal performance





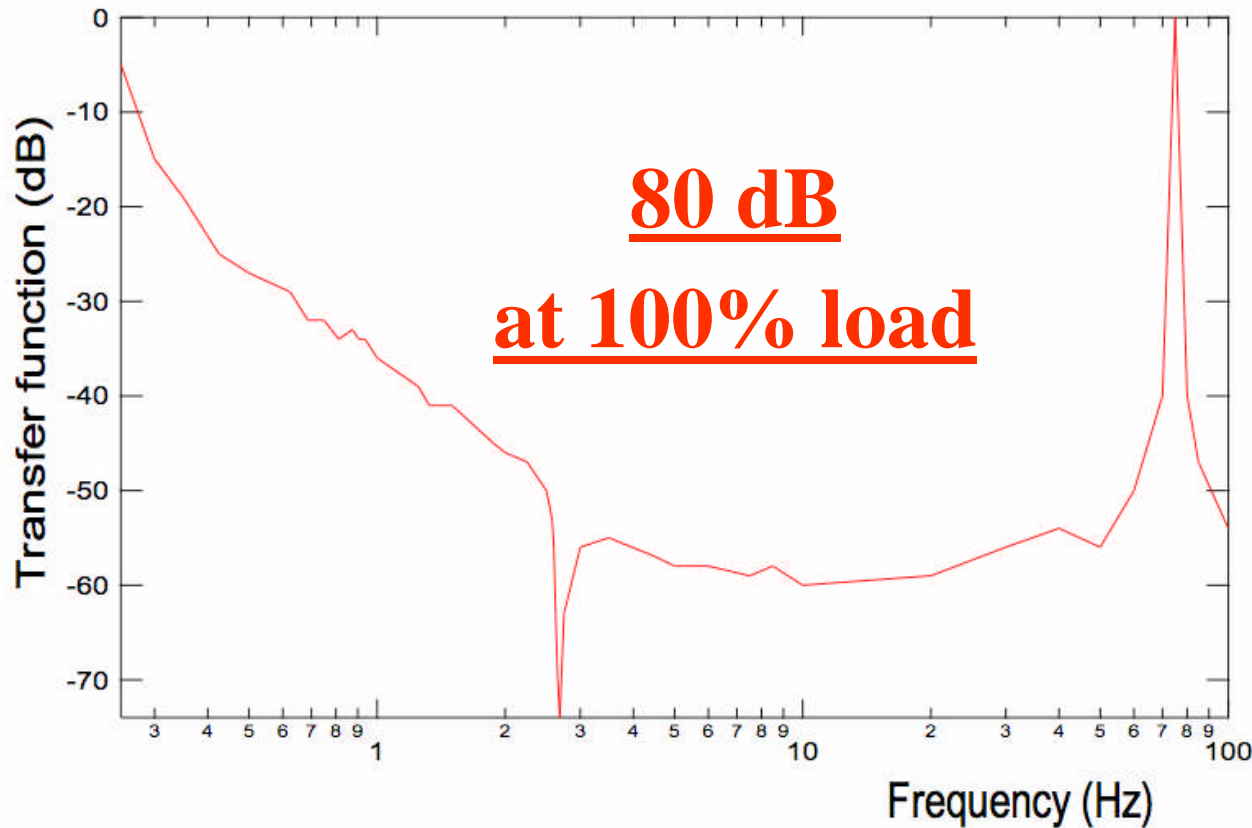
IP performance and improvements



- Minimize the mass of the legs
 - => up to 80 dB attenuation
- Add counterweights to null the COP effect
 - => beyond 80 dB attenuation



Attenuation performance
10% payload
No counterweights



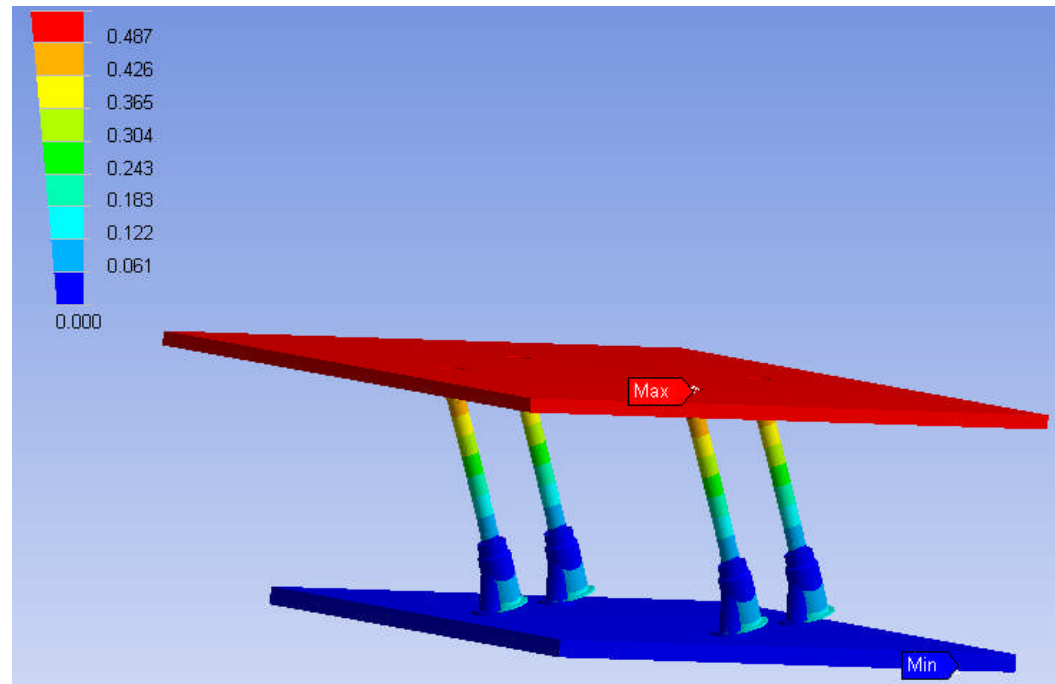
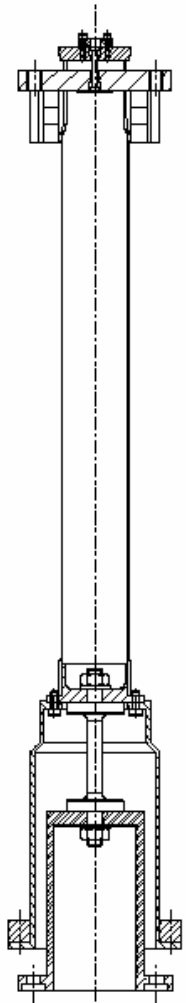


LIGO

Introducing counterweights



from 80 dB to **100 dB**
horizontal attenuation



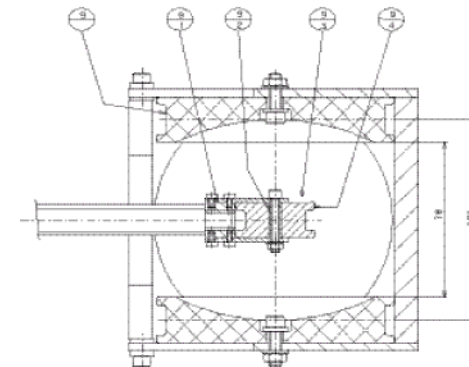
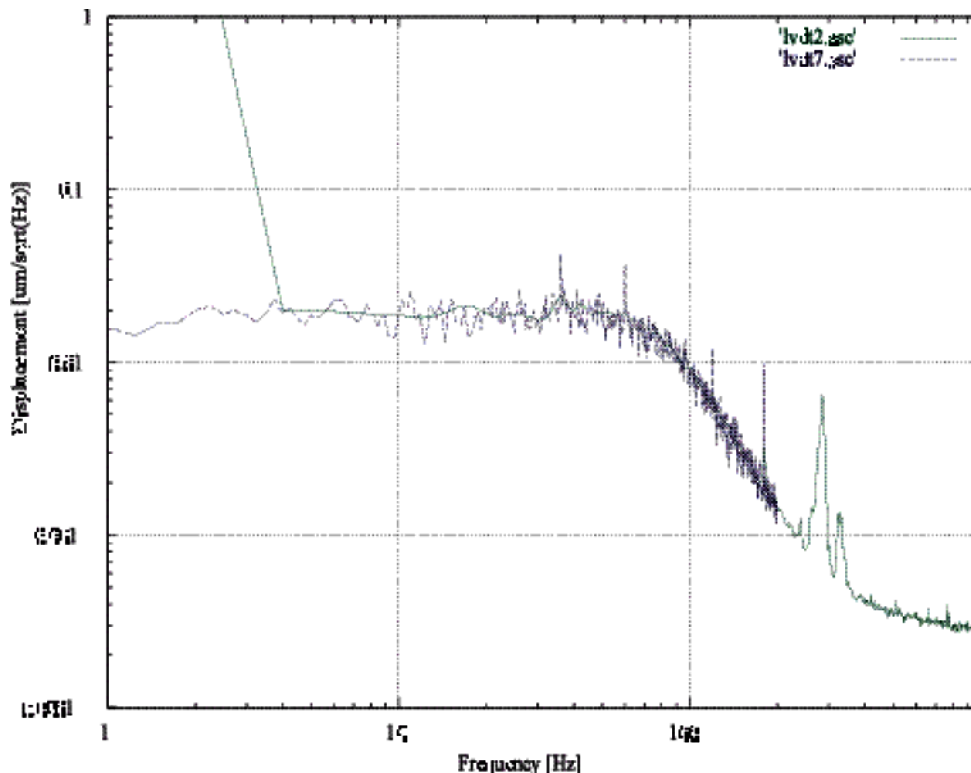
•10% CW tuning precision => x 10 perf LIGO-G060128-00-R



Static and dynamic controls MICRO POSITIONING AND POINTING



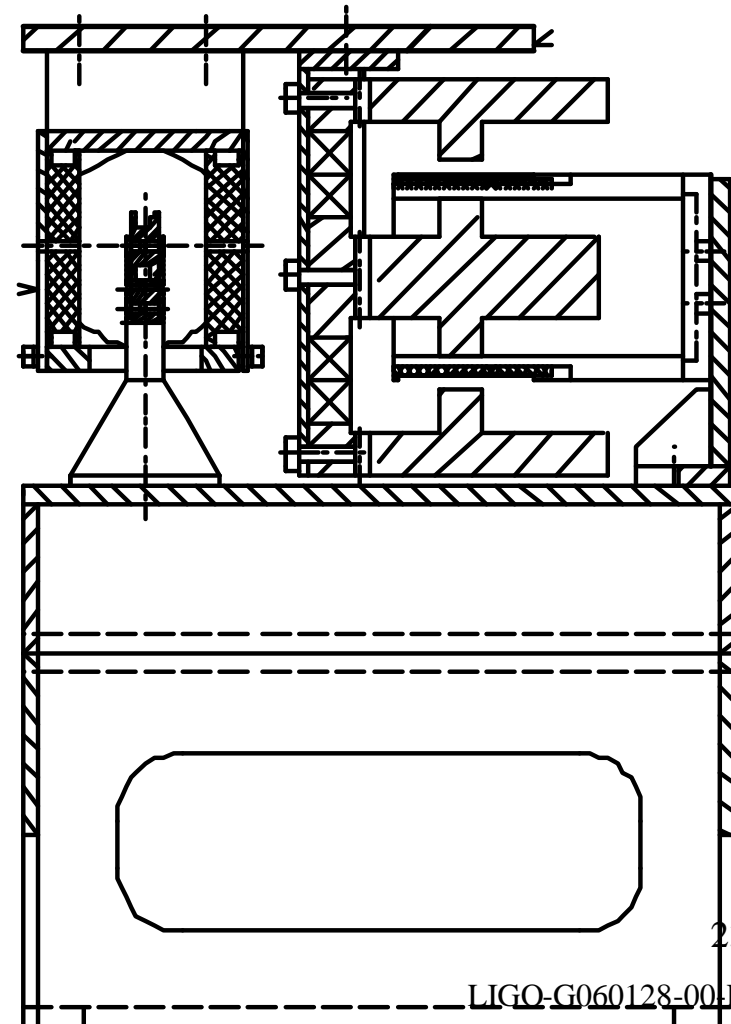
- LVDT for local nanometer positioning memory
- Voice coil actuator dynamic controls
- Position and alignment controls < 30 mHz



Sensors and coil actuators



- produced with UHV compatible materials and procedures
 - TAMA resolution ($\text{nm}/\sqrt{\text{Hz}}$)





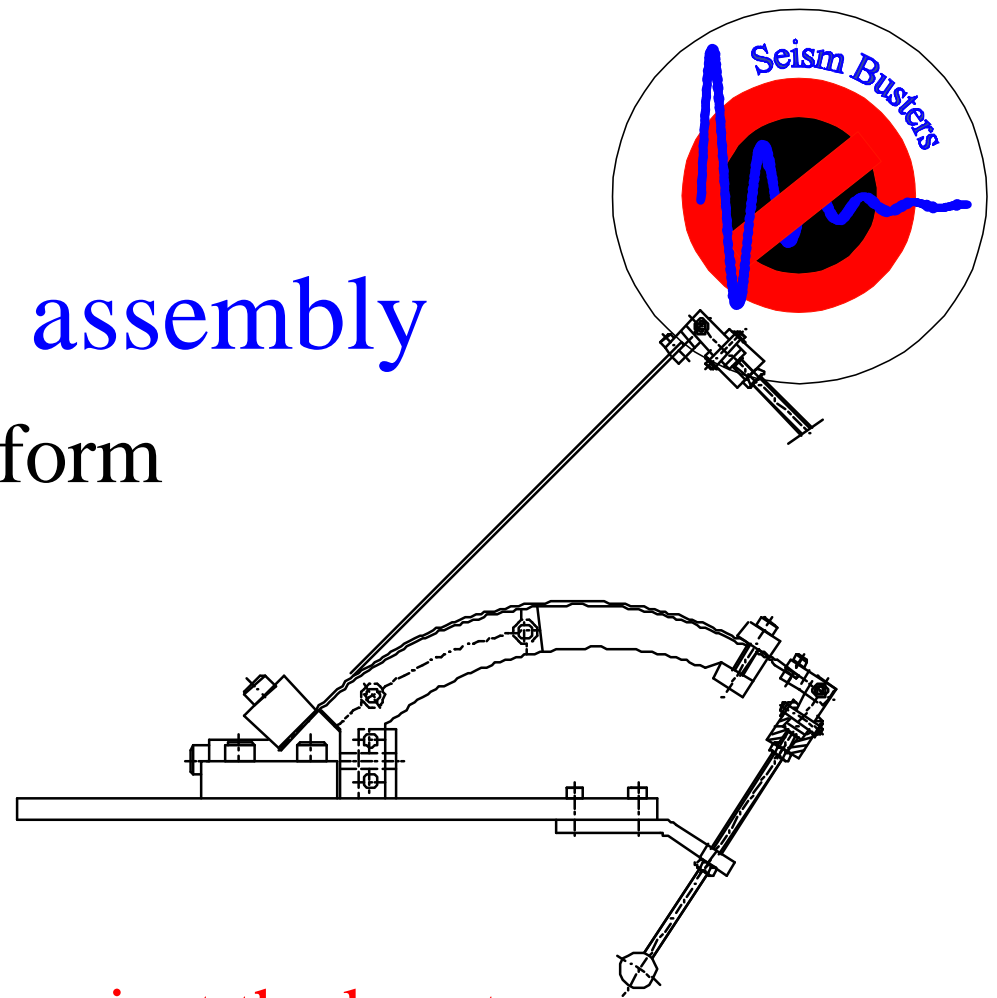
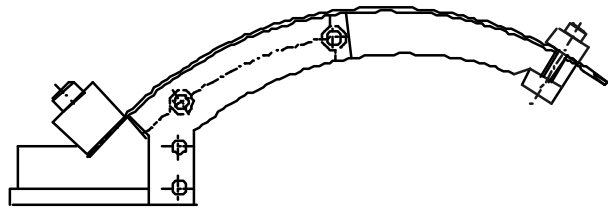
Assembling a real attenuator system



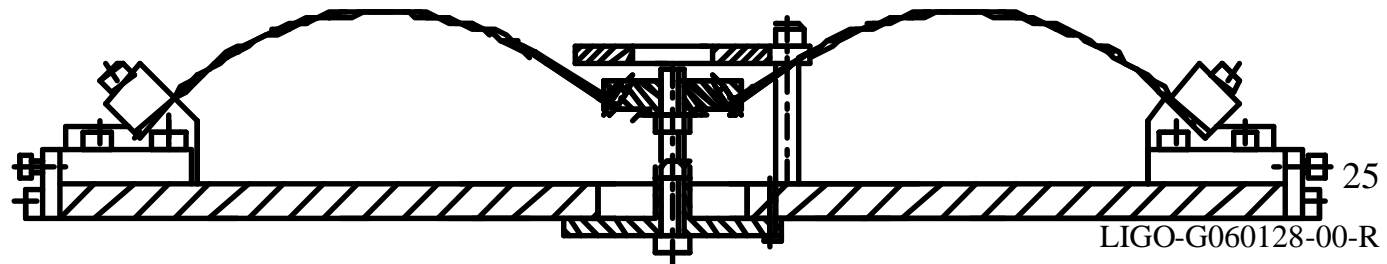
- What are we building
- How are we putting it together

HAM SAS assembly

- Pull the blade over a form
- Clamp for transport



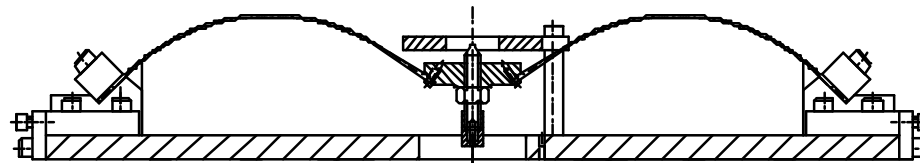
- Mount on the base and against the keystone
- Transfer the load and tune



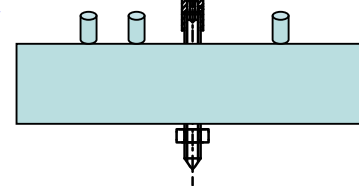


Tuning the GAS filter

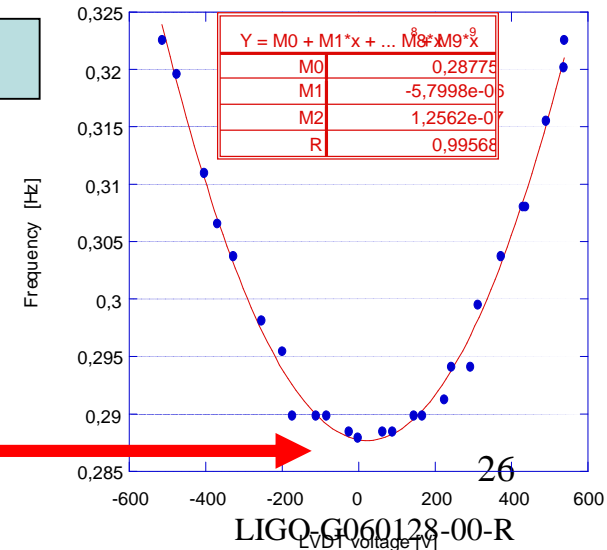
- Use screws for radial compression tuning



- Add mass to change working point

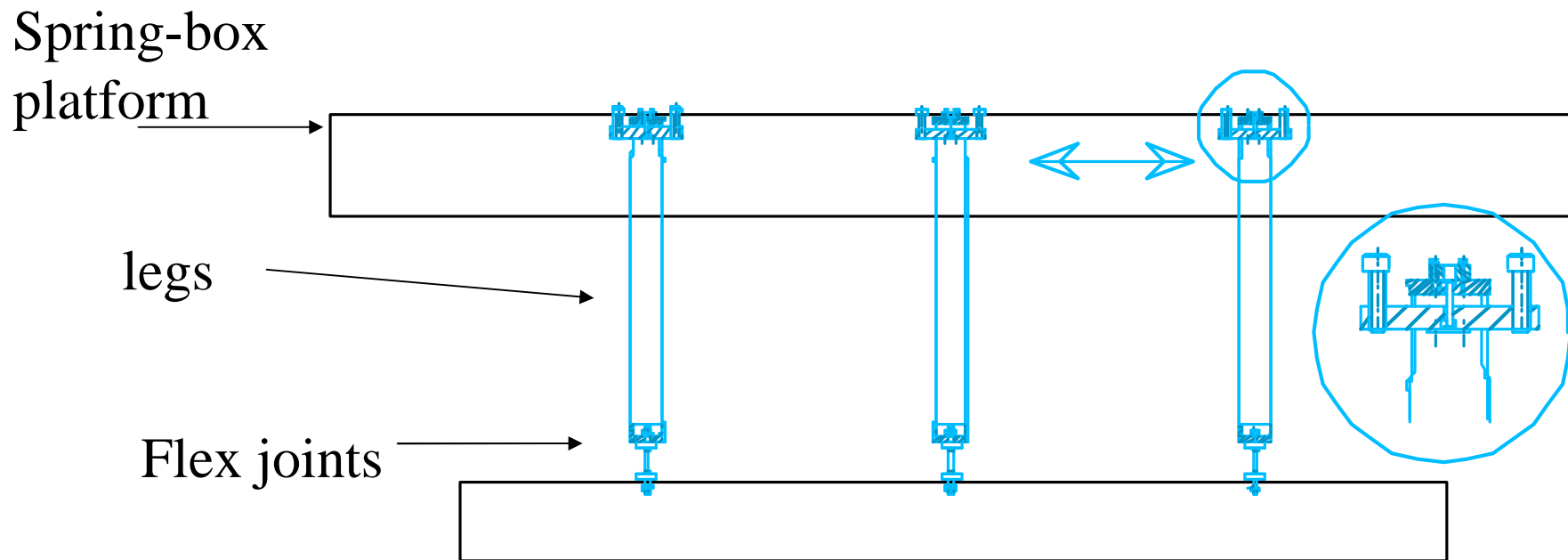


- Best mechanical working point





Horizontal direction, x, y, φ the **Inverted Pendula**

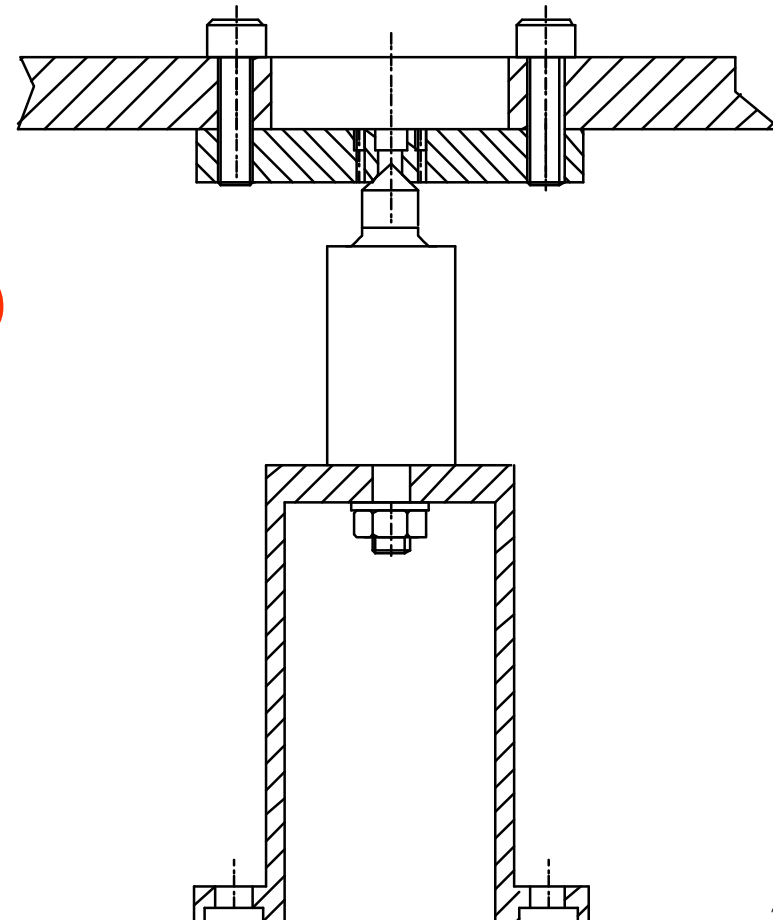




Leg alignment procedure: avoid cradle effect



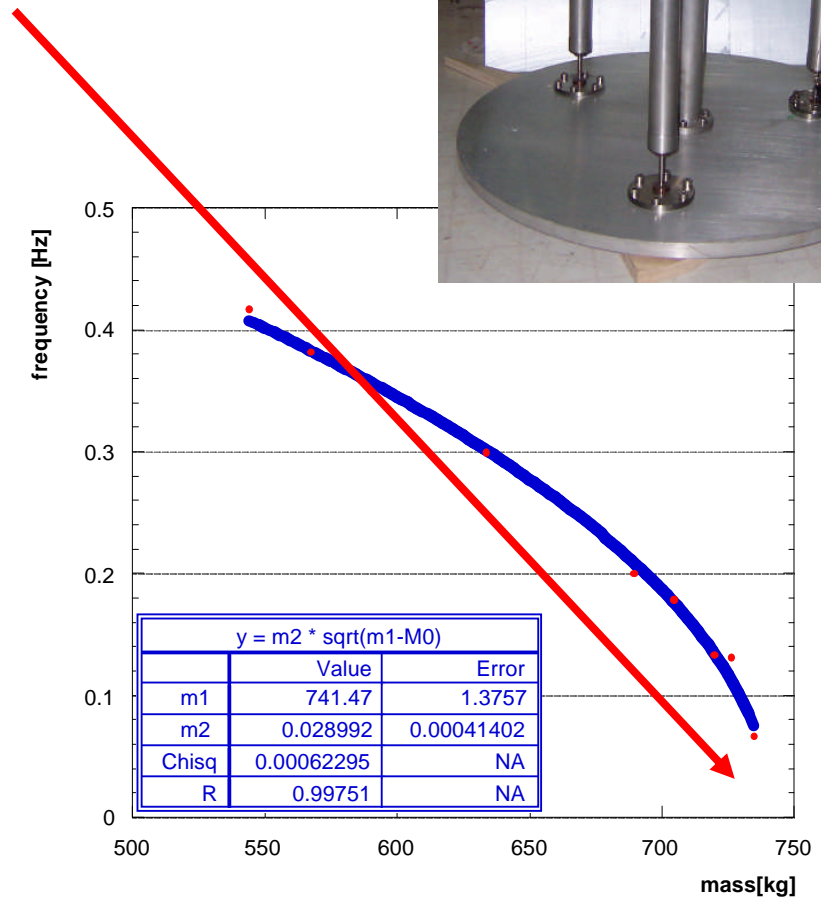
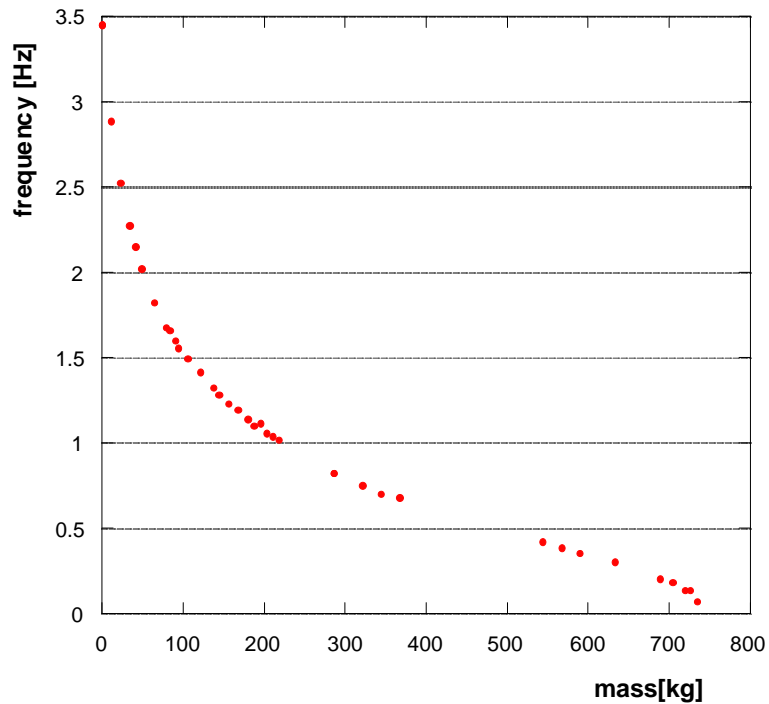
- Legs aligned
(cradle effect depressed)
to $2.5 \cdot 10^{-4}$ m/m





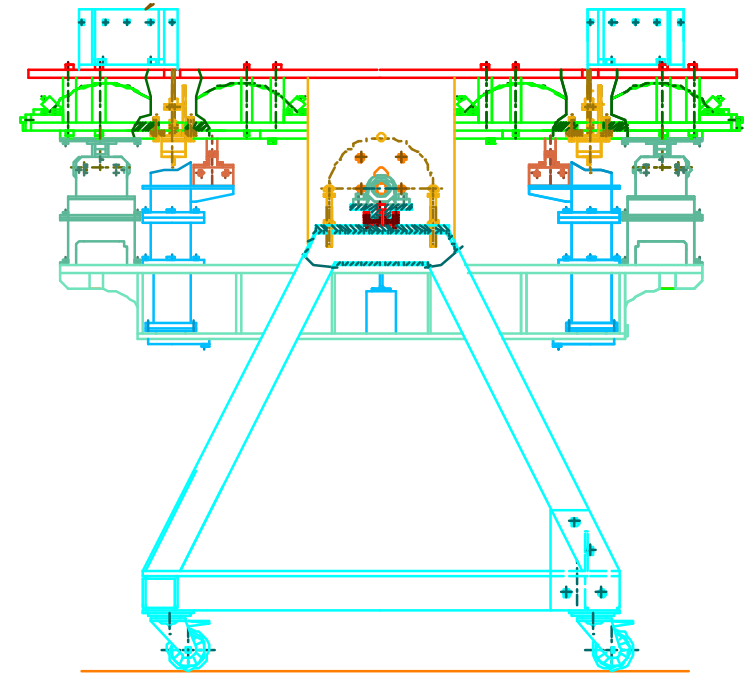
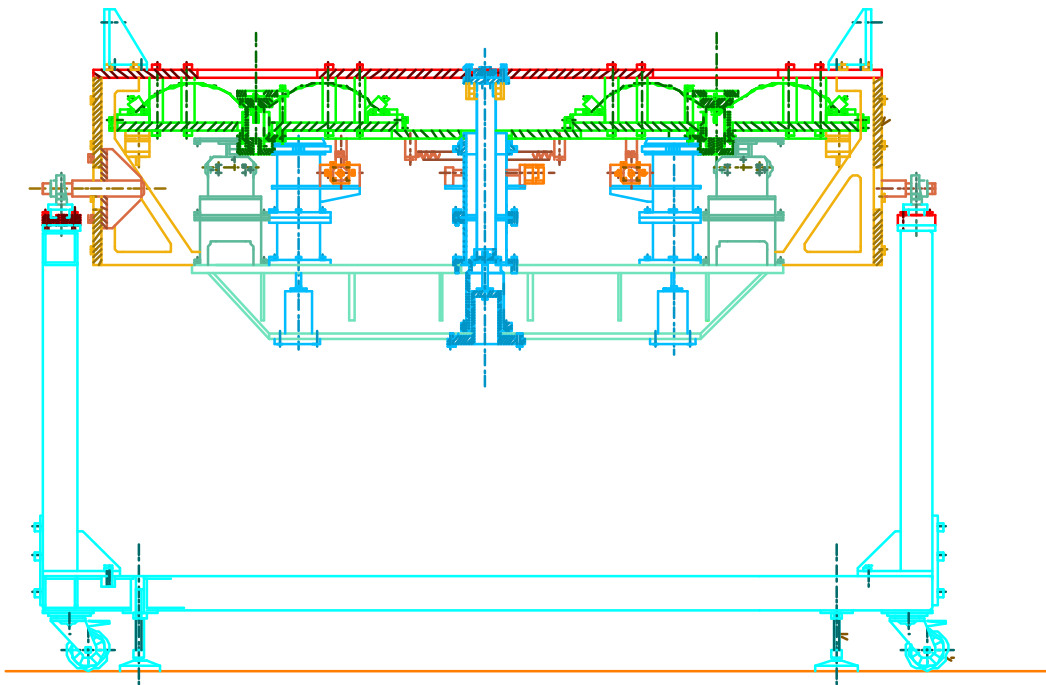
Frequency/Load tuning

<30 mHz tune on stable ground



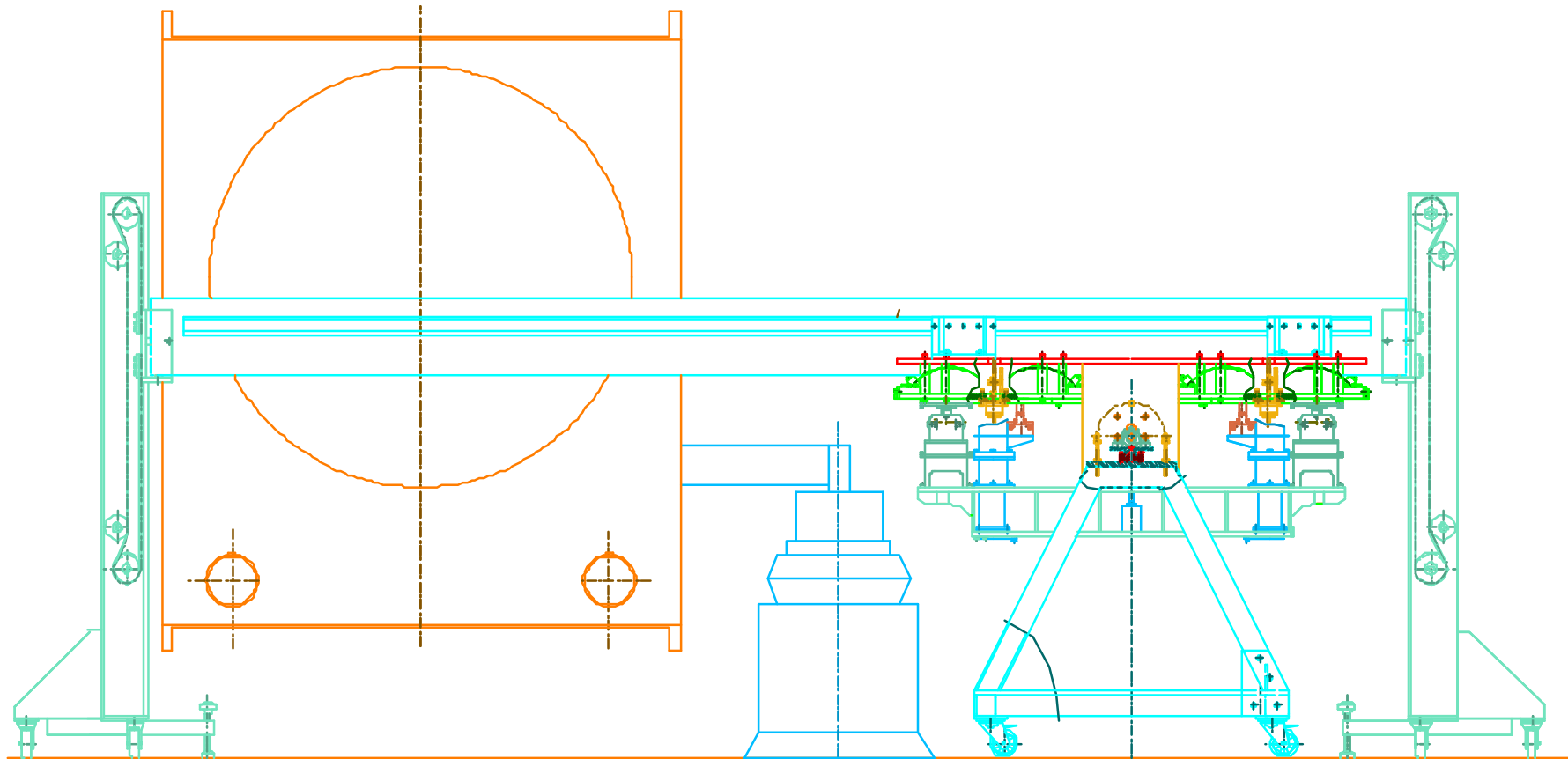


Installing SAS in the HAMs





Installing SAS in the HAMs





Installing SAS in the HAMs



- Two long rails are installed across the HAM doors extending two meter outside the chamber, resting on synchronous jacks on installation carts
- The rails are lowered to extract the optical bench from the chamber
- The optical bench slides off the HAM chamber and is lowered on a cart

- The rails descend to pick HAM-SAS from its cart
- The rails are raised to slide HAM-SAS inside the HAM
- The rails are lowered to position HAM-SAS on cross tubes

- The operation is repeated to pick-up the optical bench and lower it over HAM-SAS. The optical bench can be installed with most pre-assembled optics



HAM-SAS

primary seismic attenuation
for Advanced LIGO

