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# The LIGO Project

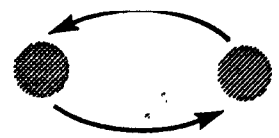
**Barry Barish**

**March 20, 1996**

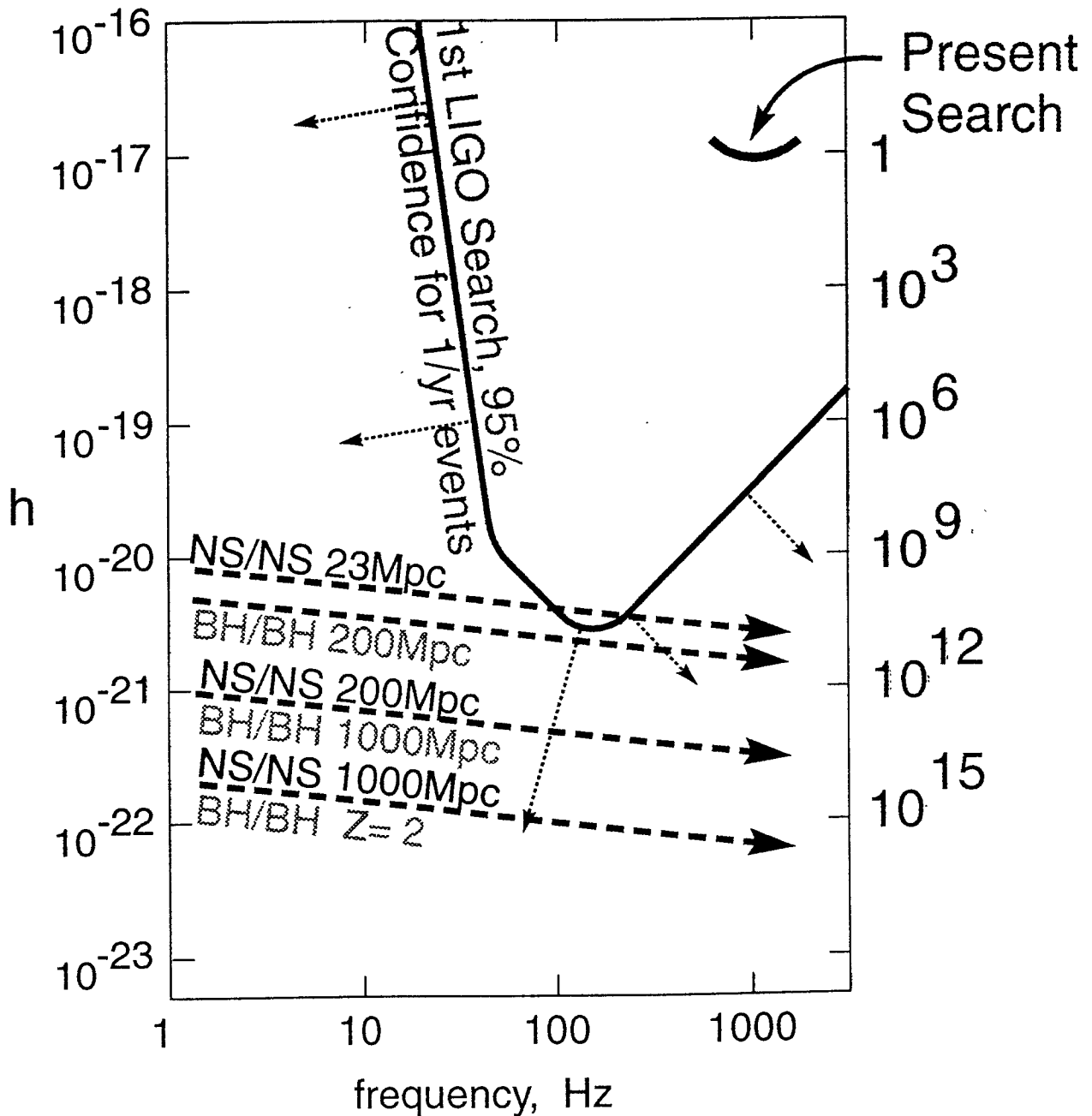
*MIT NSF Review*



# NEUTRON STAR BINARIES



[“Near-Guaranteed” source]

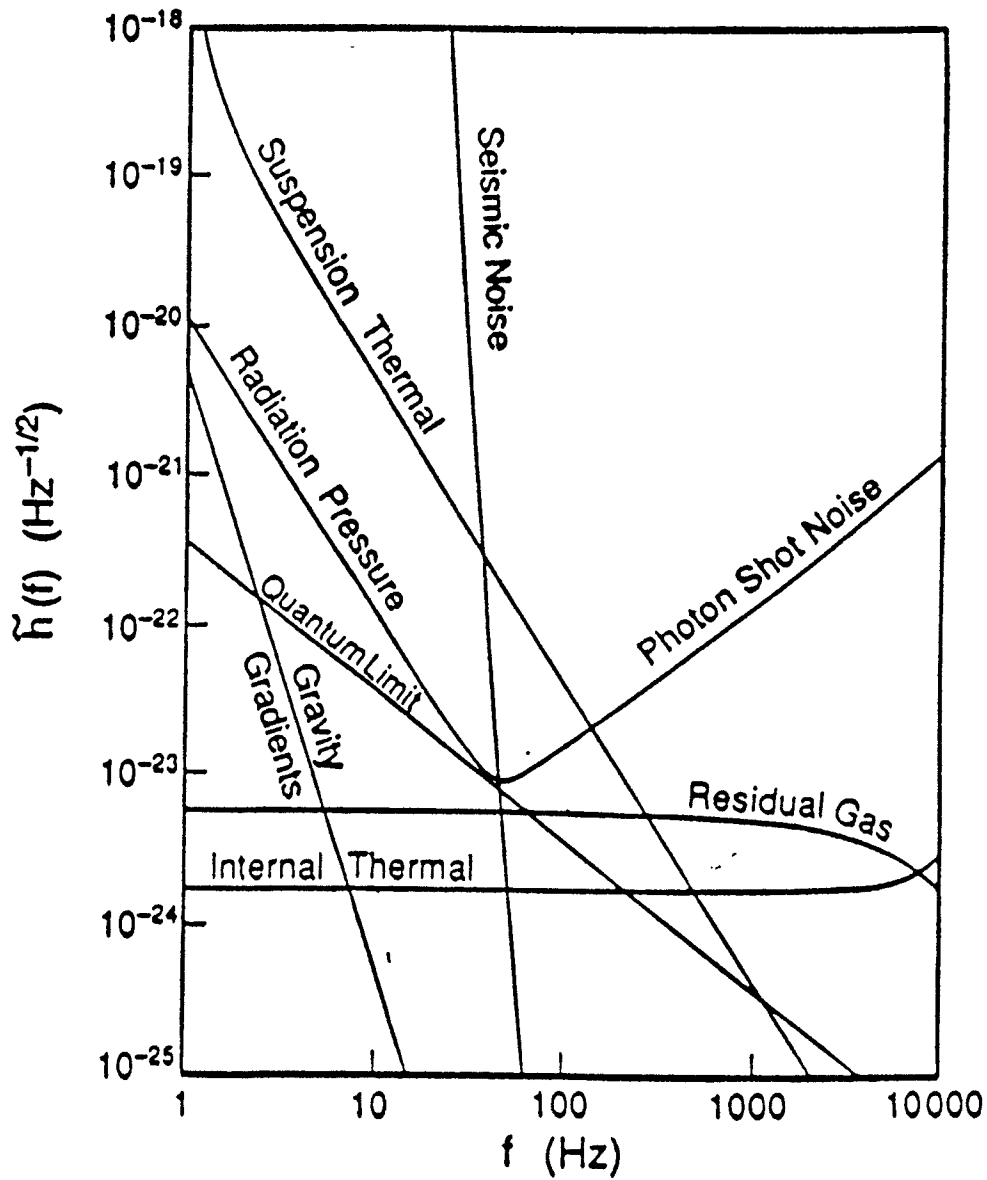


■ 15 minutes & 10,000 orbits in LIGO band

■ Rich information in waveforms:  
masses, spins, distance, direction,  
nuclear equation of state

# Noise Budget For First LIGO Detectors

- 5 Watt Laser
- Mirror Losses 50 ppm
- Recycling Factor of 30
- 10 kg Test Masses
- Suspension  $Q=10^7$



# LIGO Project

## *Technical*

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- Major Facilities
  - » Beam Tube
  - » Vacuum Systems
  - » Civil Construction
  
- Detector
  - » Detection Strategy
  - » Interferometers
  
- R&D
  - » Noise Sources and Sensitivity
  - » Demonstration Experiments
  
- Status and Plans

# LIGO Facilities

## *Civil Construction*

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### ● Characteristics

- » Structures, Foundation, Roads, etc...
  - Large and Clean Laboratory Bldg.
  - Beam Enclosures
  - Office/Lab Space
- » Requirements
  - Seismic Stability, Noise Sources, etc...
  - Cleanliness

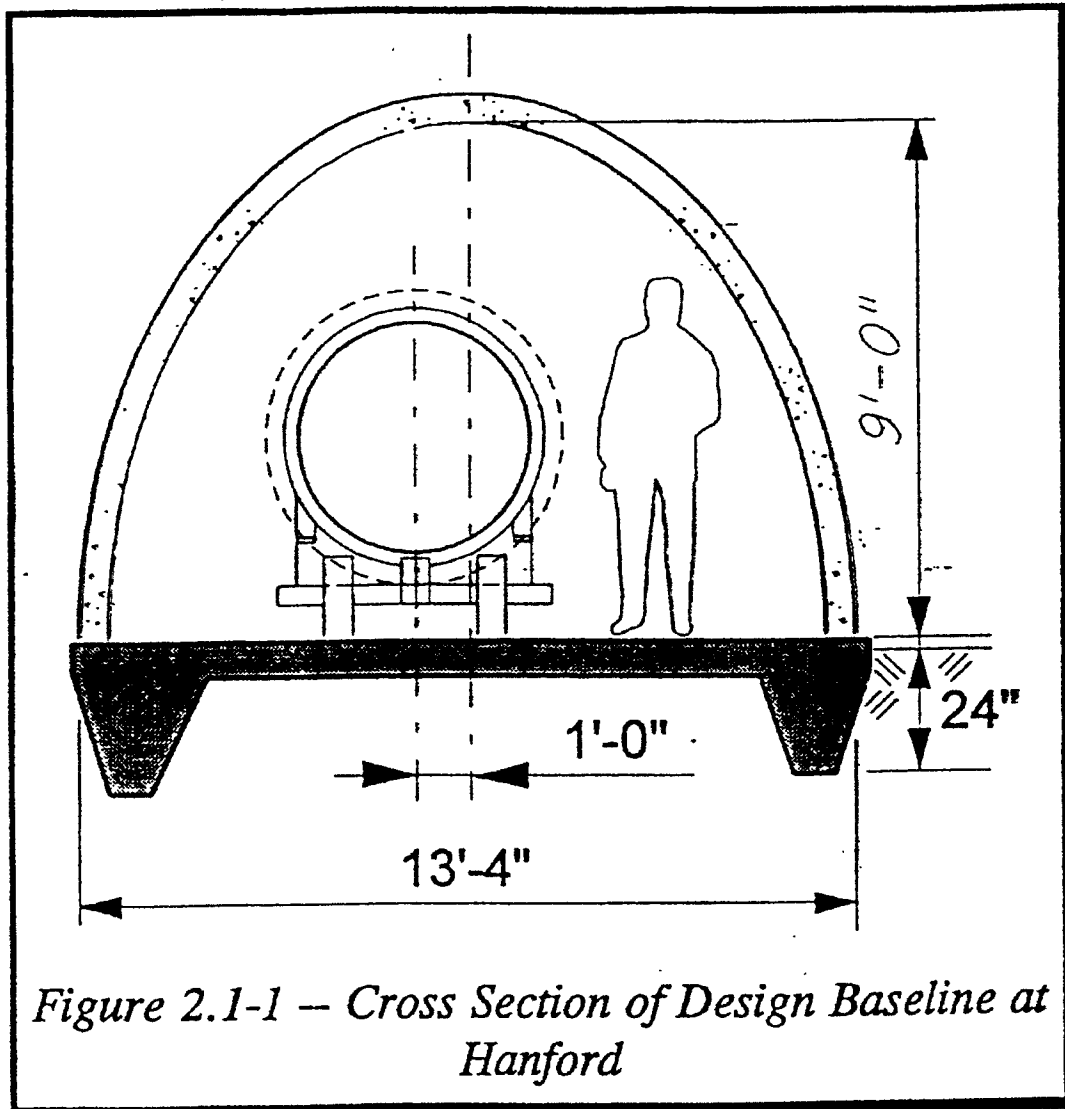
### ● Status and Plans

- » Both Sites Acquired
  - Washington cleared; graded; pouring slab and constructing enclosure
  - Louisiana cleared, grubbed; compacting, building berm underway; pipelines mitigated
- » Facilities Design
  - H.M. Parsons - slab, enclosures complete; FDR buildings April 26
  - trade Studies; value engineering
- » Construction Management
  - LIGO/Parsons -- in place in Washington

# LIGO Facilities

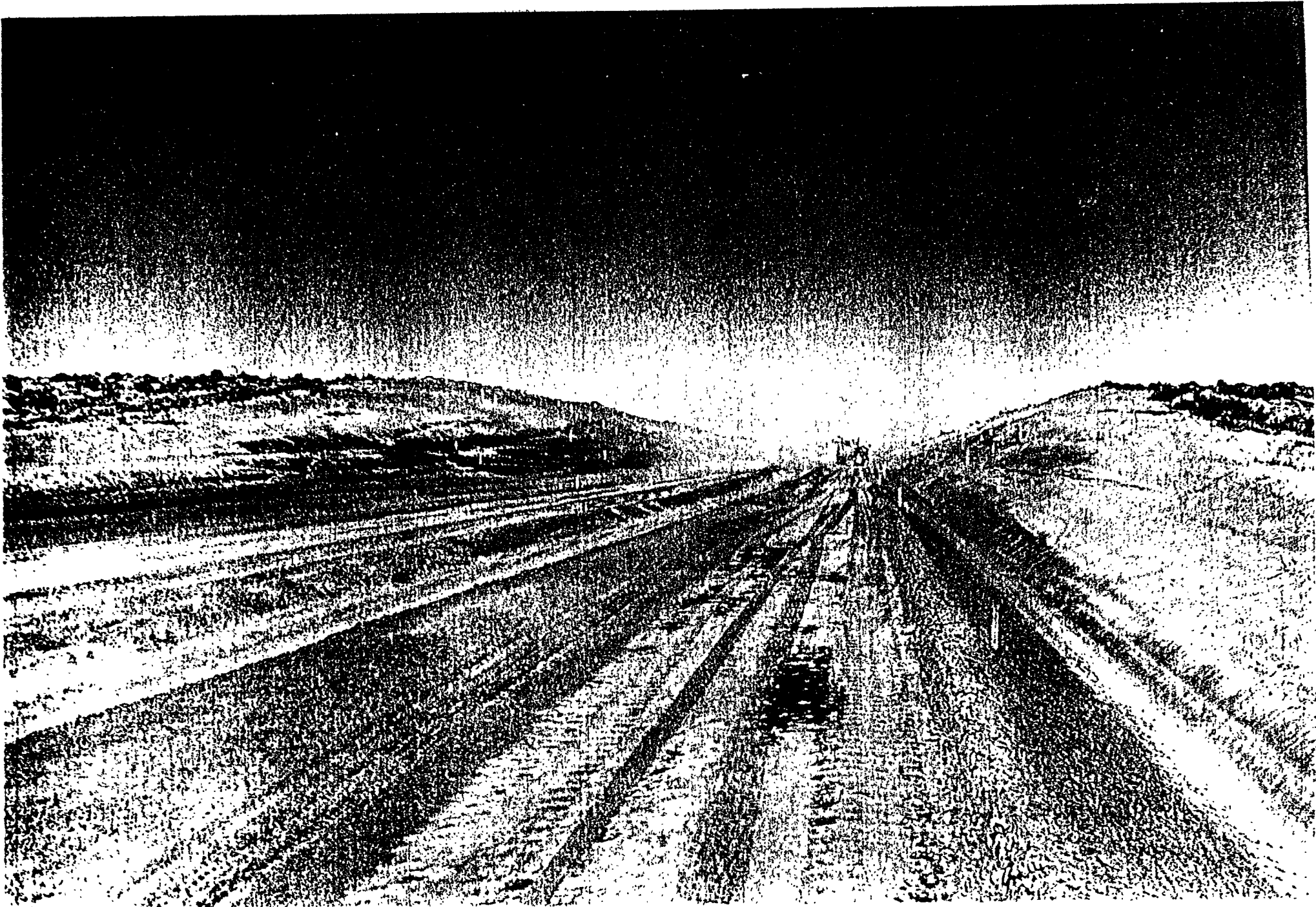
## *Beam Tube Enclosure*

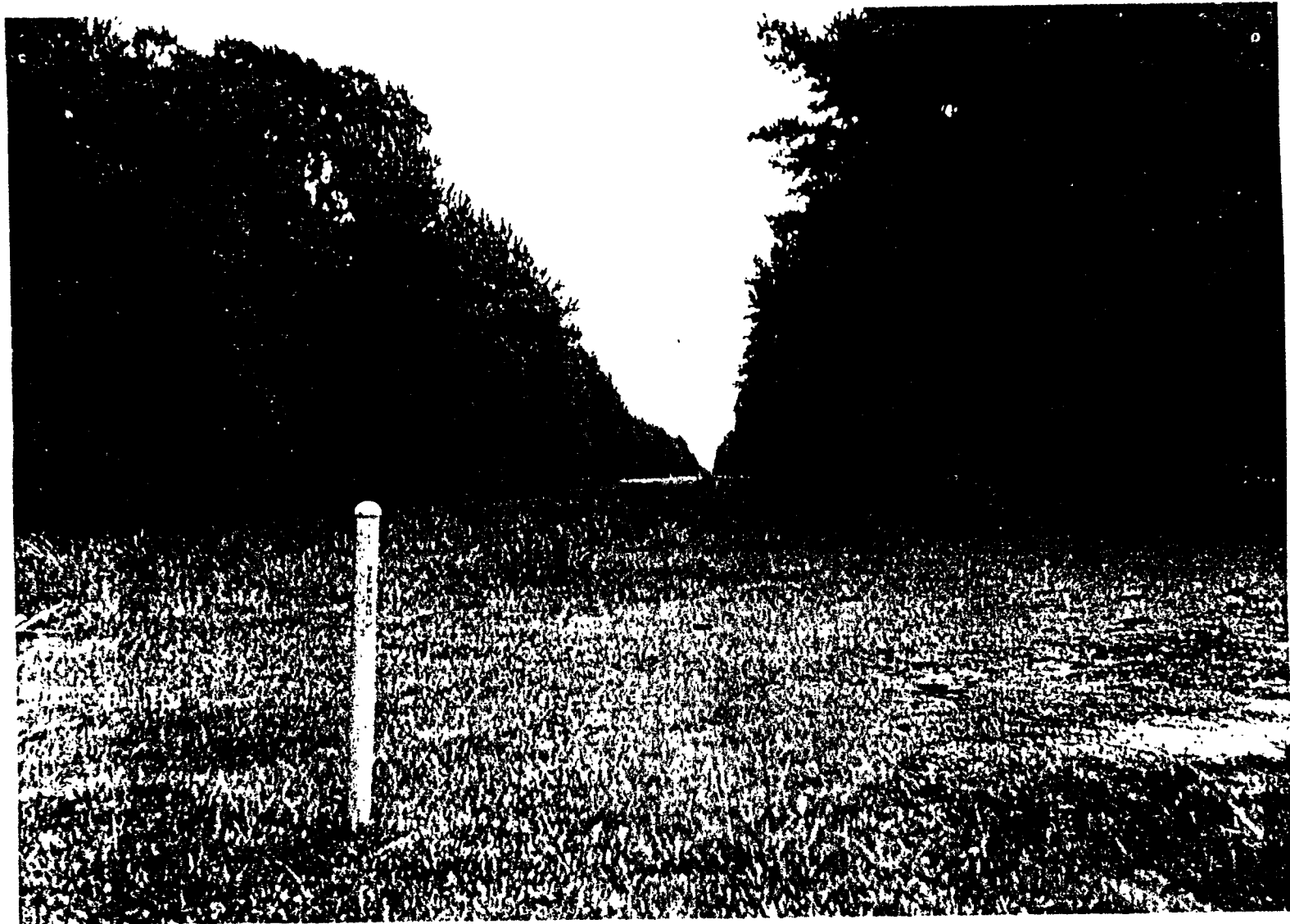
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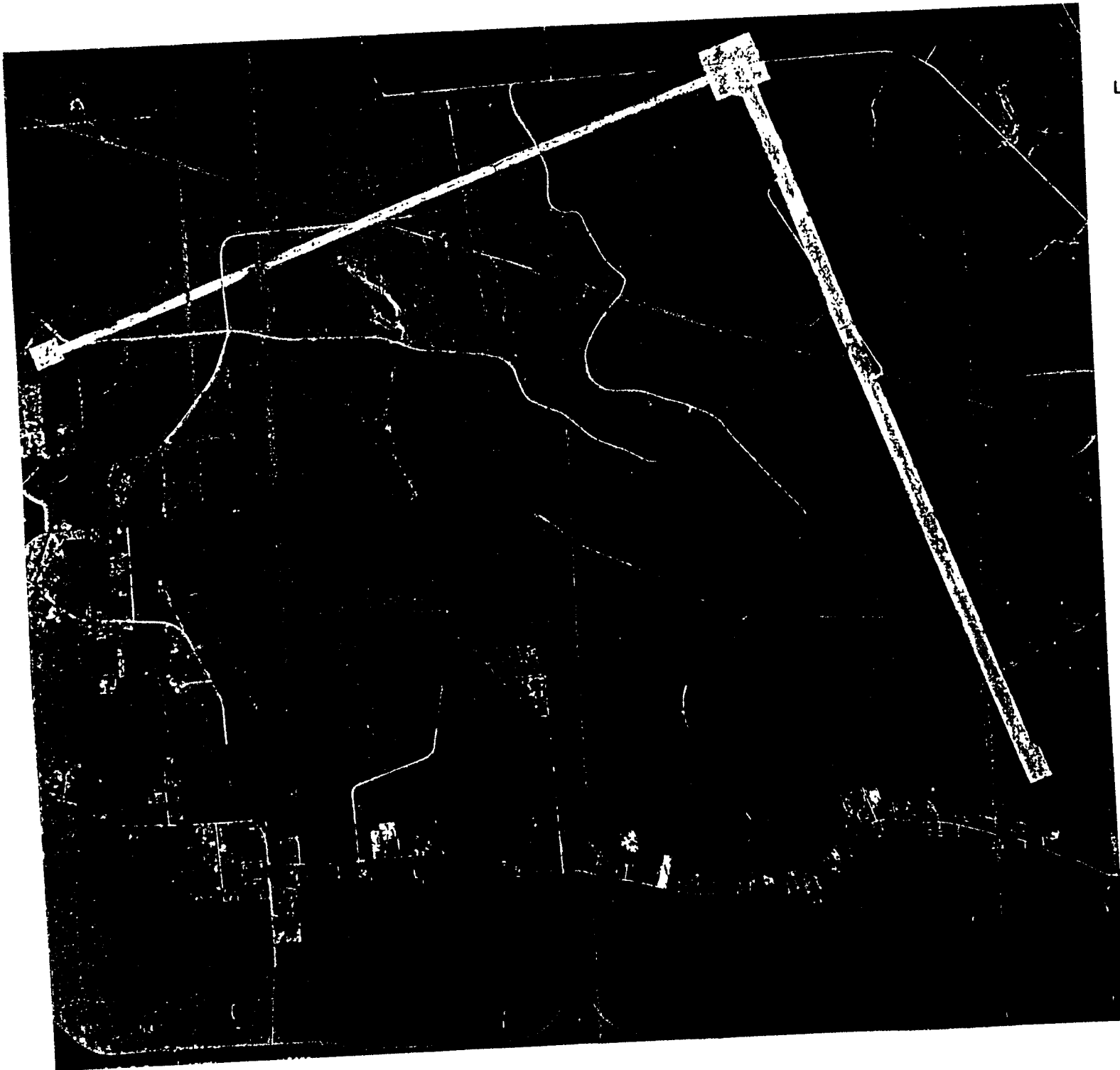
*Figure 2.1-1 – Cross Section of Design Baseline at Hanford*

63 8111









**LIGO**

**LIVINGSTON PARISH**

**LOUISIANA**

1A

**AERIAL PHOTO BY:**  
**GULF COAST AERIAL MAPPING**  
**FLOWN: AUGUST 26, 1995**  
**ALTITUDE: 12,000 FEET**

# LIGO Facilities

## *Vacuum Equipment*

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### ● Characteristics

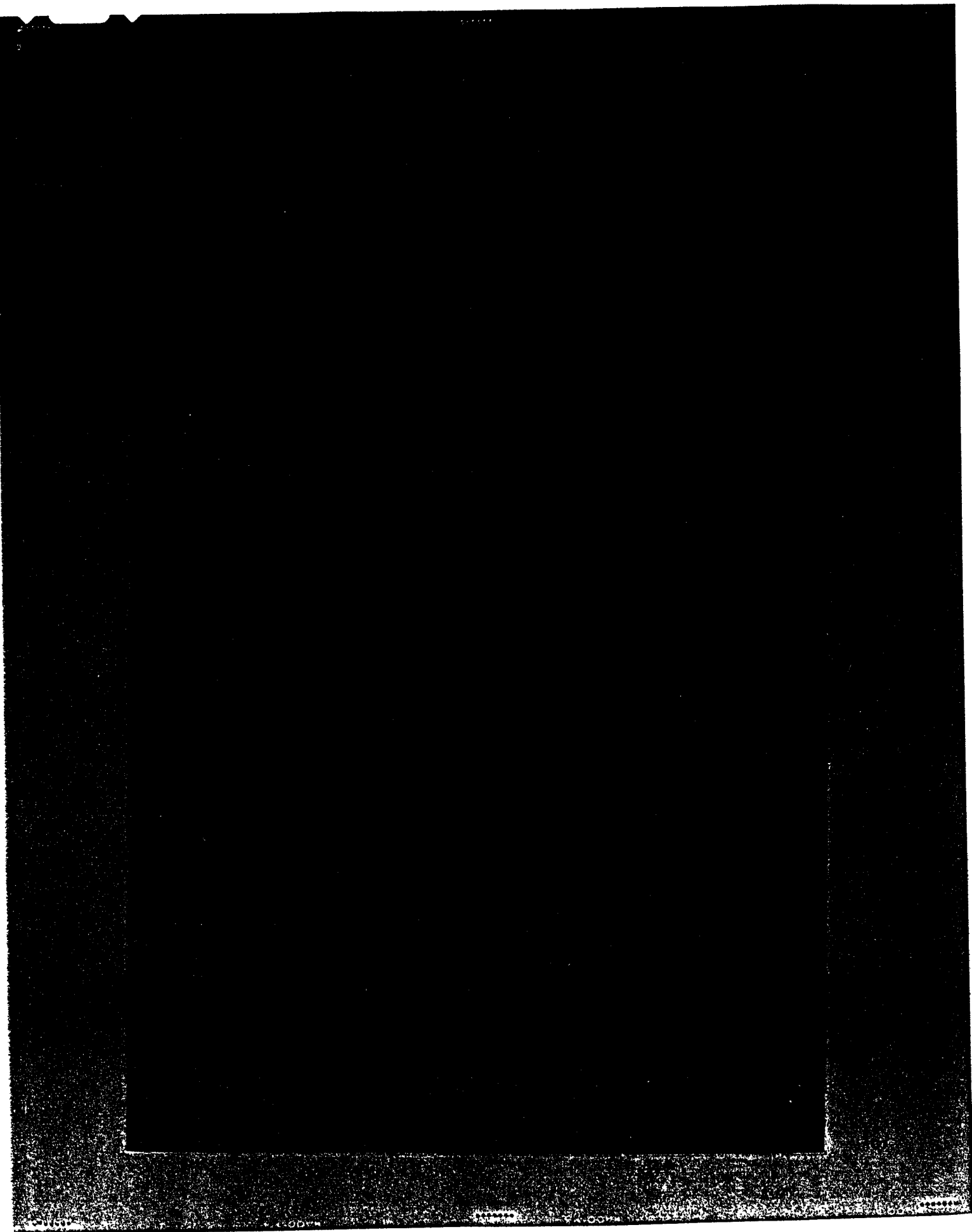
- » mostly standard vacuum equipment
  - 1st stage roughing atm -> 0.1 torr
  - 2nd stage roughing 0.1 torr ->  $10^{-6}$  torr
  - steady state - ion/getter pumps
- » large gate valves (4 ft diam)
  - access and flexibility
- » controls and monitoring

### ● Status

- » Science requirements and review 6/94
- » RFP issued for design contract only
- » Two competitive contracts awarded (CB&I, PSI)
- » Final design and manufacturing
  - down select (6/95) to PSI
  - CDR approved 10/95
  - FDR May 96; some prototype/acquisitions now

1950

1950



# LIGO Facilities

## *Beam Tube*

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- Characteristics

- » Arm Lengths - 4km
- » Tube Diameter - 4 ft
- » Initial Detector
  - $10^{-6}$  torr Hydrogen;  $10^{-7}$  torr Water
- » Advanced Detectors
  - $10^{-9}$  torr Hydrogen;  $10^{-10}$  torr Water
- » Quality Control
  - (materials, welding, cleaning, etc..)

- Status and Plans

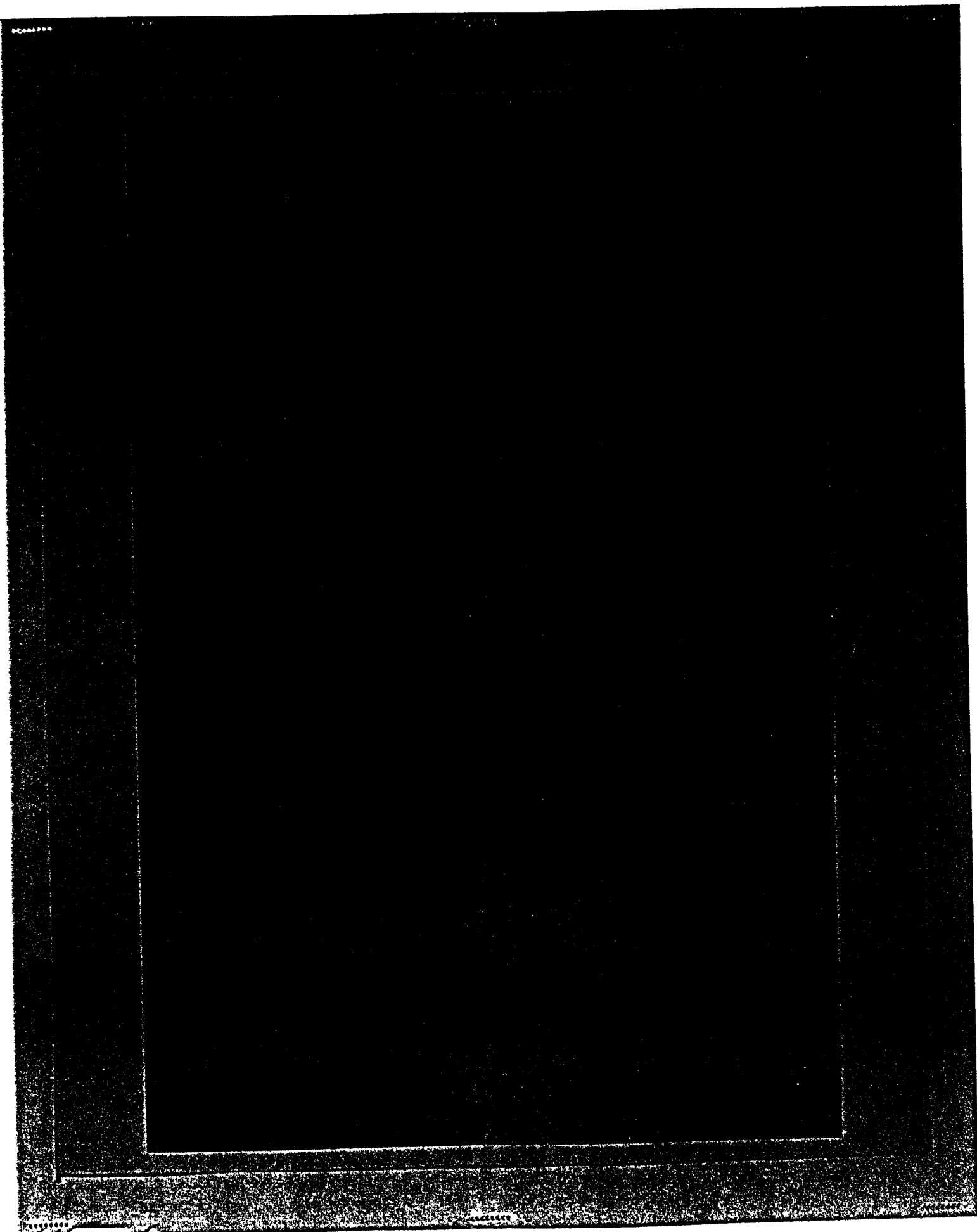
- » Design Contract was with CBI
  - Final Design Report Accepted (6/94)
- » Qualification Test
  - 130 ft Section - success (4/95)
- » Contract Options

# LIGO Facilities

## *Beam Tube*

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- Characteristics:
  - » length = 4 km (4 arms)
  - » diameter = 4 ft
  - » volume = 20, 000m<sup>3</sup>
  
- Design Contract with CBI
  - » Design report accepted
    - thin wall stainless spiral weld structure
    - 65 foot sections with bellow
    - quality control in material selection, welding, cleaning, etc..
  - » Qualification test -
    - 130 ft section assembled, cleaned, baked and tested -- achieved design
  
- Construction contract
  - » negotiated option with CB&I ( prepared to compete if necessary)
  - » contract placed December 95
  - » design review - March 27



# LIGO Detectors

## *R&D Program*

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- Sensitivity

- » main features of 40 m spectrum understood
- » monolithic test masses improve sensitivity

- Demonstration Experiments

- » optical recombination demonstrated on 40 m
- » acquisition locking with LIGO controls
- » MIT phase noise experiments

- Pre- [detector design freeze][<1998]

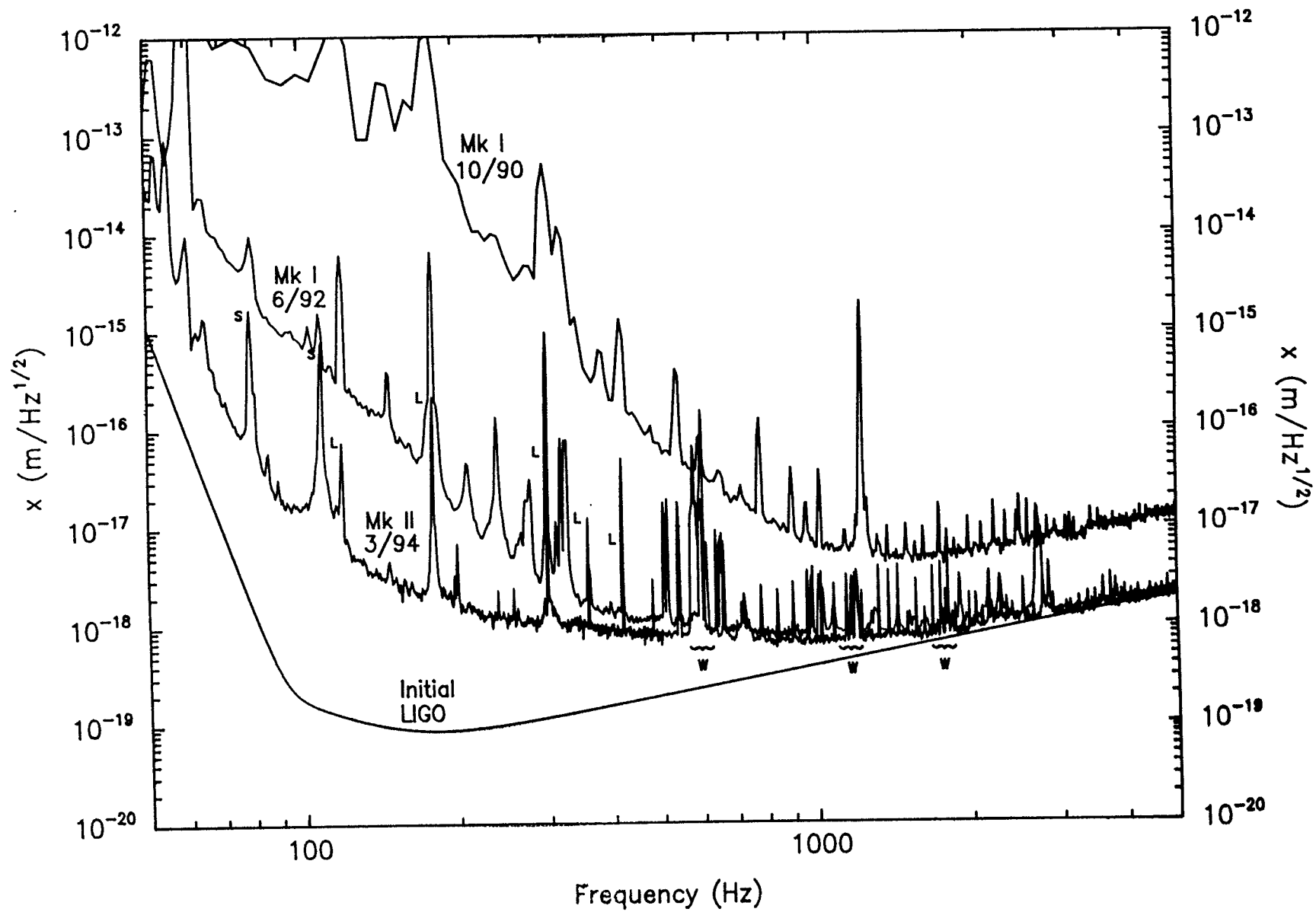
- » Program testing directed at tasks that could effect design over the next two years

- Post- [detector design freeze][>1998]

- » Program directed at improved sensitivity;
- » understand performance of LIGO initial interferometer
- » gain experience running an interferometer facility



Displacement Sensitivity of Caltech 40 m Interferometer



# LIGO Detectors

## *Detector Implementation*

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- Detector baseline established for Costing (9/94).
  - » This baseline is 'consistent' with sensitivity goals.
- Detector requirements established from sensitivity goals
- Switch to Nd:YAG lasers for LIGO
  - » decision in Sept '95
  - » YAG chosen for reliability and most direct path toward high power (improved sensitivity)
  - » Now integrating into LIGO baseline
  - » official change control action when plan is developed and cost/schedule impact understood

# LIGO Detectors

## *Integration*

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- Science Goals for initial and improved detectors
  
- Establish Systems Requirements
  - » First draft document
  
- Modeling
  - » environment - AVS
  - » end to end model - underway
  - » lock acquisition, optics, etc.
  - » help final design and understand performance
  
- Interfaces
  - » facilities, detector

# **LIGO Management**

## *Approach*

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### **Project Management Plan**

- Sets out project organization, management, project controls, etc

### **NSF Budget Approval**

- Construction + R&D Budget - \$292.1M
  - » Construction Completed in 1999
- Commissioning/Operations - \$68.7M
  - » Commissioning Commences in 1997
  - » 3-fold Sensitivity  $10^{-20}$  by July 2000
  - » First Search for Gravitational Waves
  - » Sensitivity  $10^{-21}$  by end of 2001
- Established as our Working Baselines

### **Project Organization**

- Parallels WBS - Product Oriented
- Integrated Detector and R&D Efforts

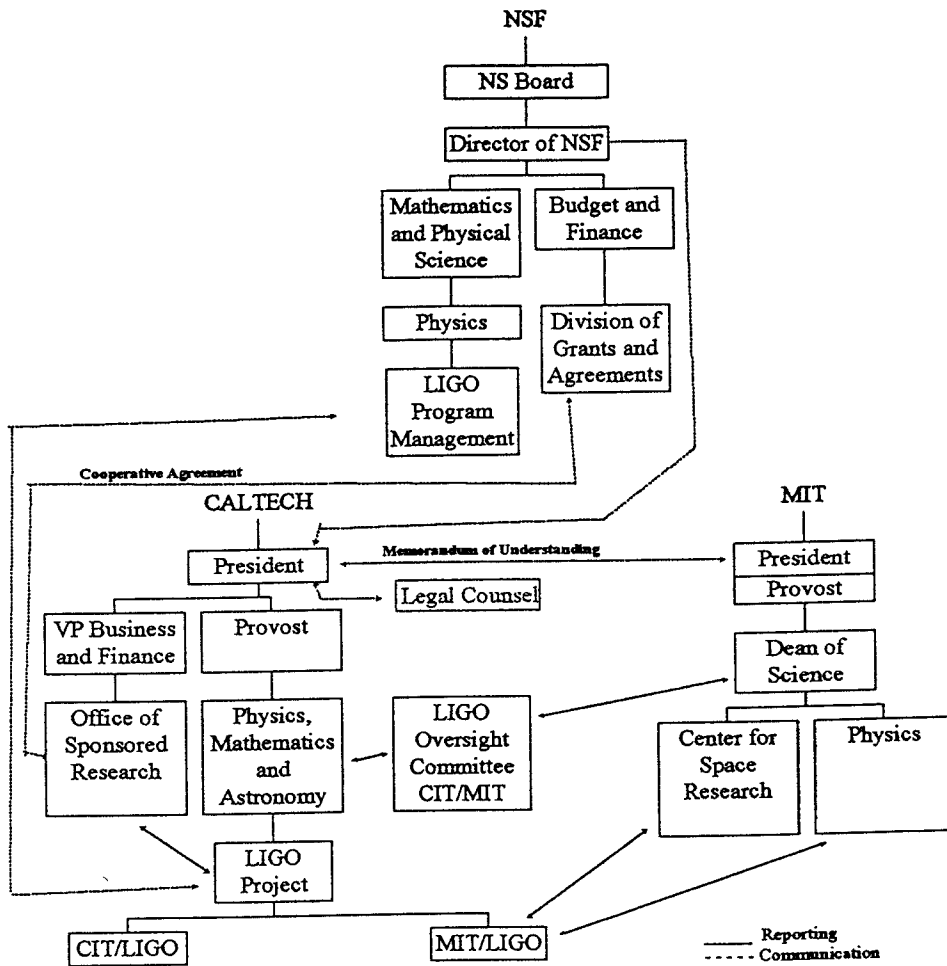


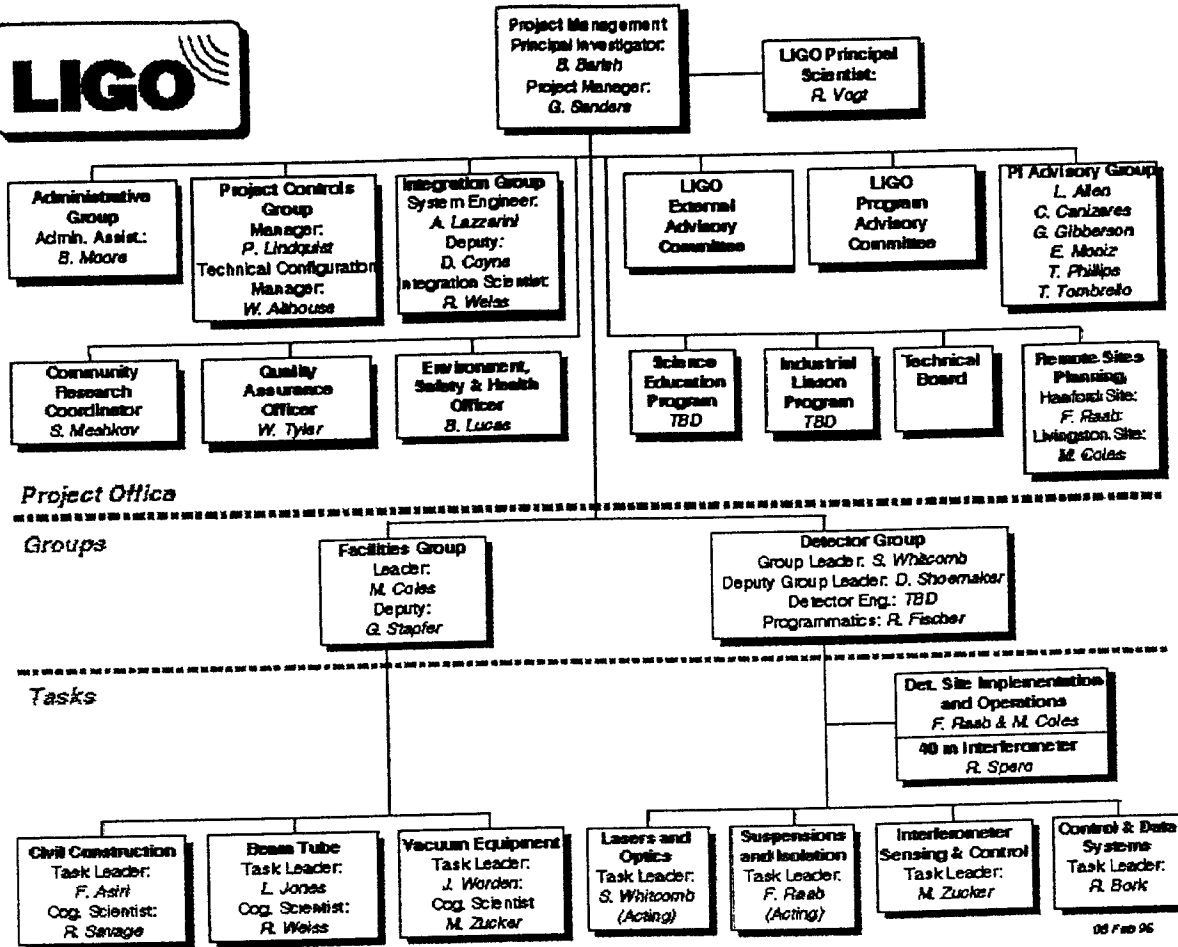
Figure 2-1. Organizational hierarchy.

# LIGO

## *Approved Funding*

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Fiscal Year	Construction	R&D	Operations	Total
1992	\$15.9M	\$3.2M		\$19.1M
1993	\$20.0M	\$4.0M		\$24.0M
1994	0	\$4.0M		\$4.0M
1995	\$85.0M	\$4.0M		\$89.0M
1996	\$55.0M	\$4.0M		\$59.0M
1997	\$55.0M	\$0.8M	\$0.3M	\$56.1M
1998	\$41.2M		\$7.3M	\$48.5M
1999			\$20.9M	\$20.9M
2000			\$21.1M	\$21.1M
2001			\$19.1M	\$19.1M
<b>Total</b>	<b>\$272.1M</b>	<b>\$20.0M</b>	<b>\$68.7M</b>	<b>\$360.8M</b>



Here is a link to the LIGO roster.

Here is a link to the LIGO meetings calendar for the next six months.

# LIGO

## *Staffing as of June '95*

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<b>Task</b>	<b>Caltech</b>	<b>MIT</b>	<b>Total</b>
Proj Mgt	4	0	4
Admin	6	2	8
P Control	8	0	8
Syst Int	6	1	7
Facilities	7	0	7
Detector	17	2	19
R&D	10	7	17
Grad Stud	4	5	9
<b>TOTAL</b>	<b>62</b>	<b>17</b>	<b>79</b>



# LIGO

## *Commissioning/Operations*

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- Management for Hanford:
  - » F. Raab (Head)
  - » O. Matherny (Facilities Manager)
  - » J. Worden (Vacuum Engineer)
  
- Management for Livingston
  - » M. Coles (Head)
  - » G. Stapfer (Facilities Manager)
  - » A. Sibley (Vacuum Engineer)
  
- MIT / Caltech
  - » support operations and running
  - » advanced detector development
  - » data analysis and physics

# LIGO

## *The Community*

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- LIGO Users

- » Aspen Workshops (Jan 95,96)
- » Formed charter, membership of LRC (172 members)
- » Elected Executive Board; Chairman (S.Finn)
- » Meeting @ Spring APS

- Collaborations

- » VIRGO : Workshop in April on Data Formats
  - cooperate on technology, data analysis; MOU in process
- » GEO 600:MOU - visitors program
- » TAMA: MOU - visitors program
- » MOUs :
  - K. Thorne (Caltech), Bender (JILA), Cutler (Penn State), Saulson (Syracuse); Sandeman (Australian groups), etc.

# LIGO Status and Plans

## *Conclusions*

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- The LIGO Project
  - » technical considerations for civil construction and vacuum systems in hand
  - » major contracts established and actual construction of facilities well underway
- R&D and Detector \*
  - » active and productive R&D program
  - » development of the detector design
  - » reconcile with requirements
- Detailed site planning underway
  - » staffing and organization
  - » plan for installation, commissioning and operations
  - » integration with the construction schedule
- Outside community being included, informed and organized