



# **Stack-a-flare** SGR burst search review readiness

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Photo: Joe Becker

LIGO-G080591-00-Z



# Project Science Goals

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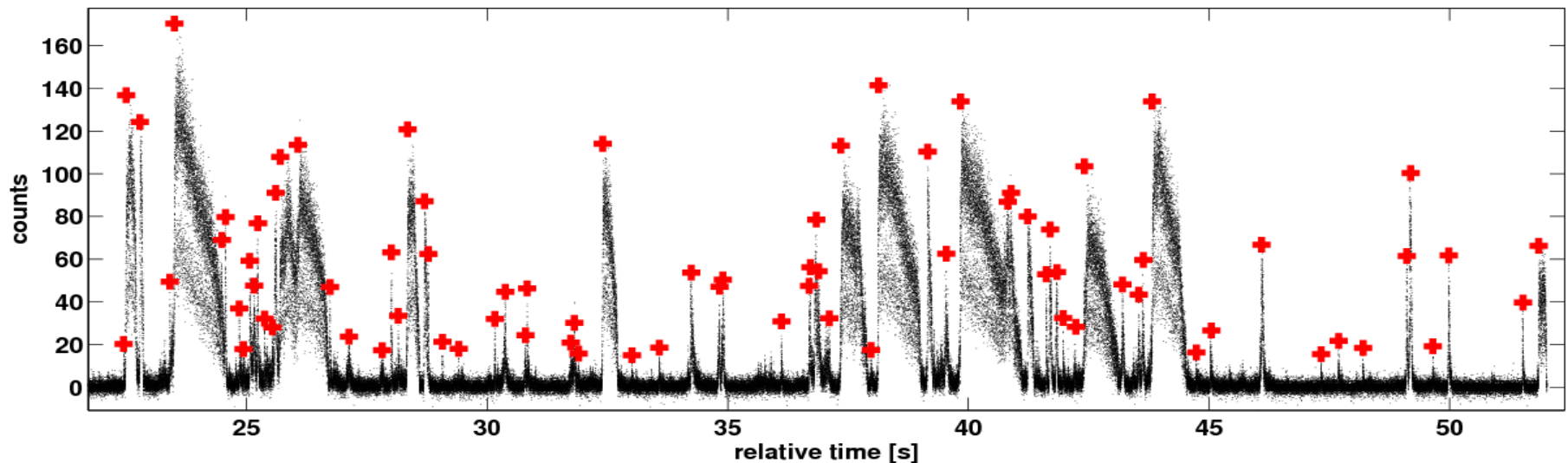
Astrophysical target: SGR 1900+14 storm (March 29 2006)

Use assumption of SGR burst similarity to improve SGR GW search

Detection statement

Improved upper limits

Probe new astrophysics with different stacking scenarios





## Other Project Goals

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Keep it simple

**Stack-a-flare:** Flare pipeline does heavy lifting

Review will be significantly quicker than for S5y1 SGR search

Publication target: ApJL

We expect significant sensitivity gains over S5y1

We are probing new astrophysics via stacking model scenarios



# Search Parameter Space

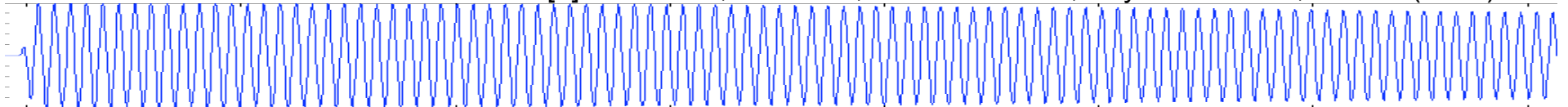
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**Presentation of the search will be the same as for S5y1 SGR**

## Ringdowns 1– 3 kHz

Simulation frequencies: 1090, 1590, 2090, 2590 Hz  
3 kHz upper bound: strange quark stars  
1.5 kHz lower bound: lightweight star with stiff EOS [1]  
Simulation tau: 200 ms  
predicted range is 140-380 ms

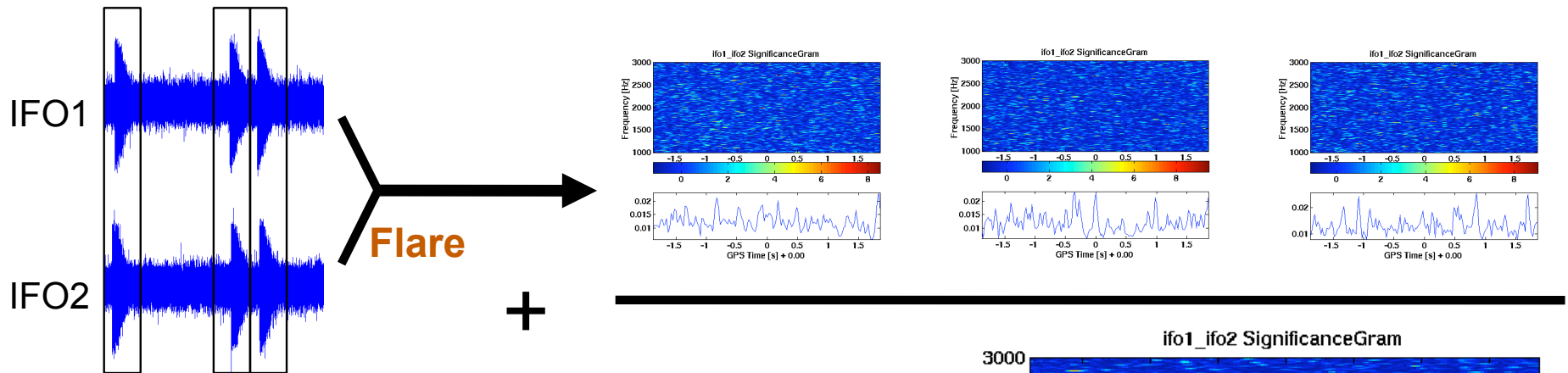
[1] O. Benhar, V. Ferrari, and L. Gualtieri, Phys. Rev. D 70,124015 (2004)



## WNB below 1 kHz

WNB injections to estimate upper limits; 11 ms and 100 ms durations  
Band-limited to detector's sensitive regions:  
100 – 200 Hz (small band)  
100 – 1000 Hz (large band)

# P-Stack Method



1. Apply Flare N times at EM burst times
2. Add up resulting **P**ower matrices

We have chosen parameters  
We have obtained closed box results

Pros:

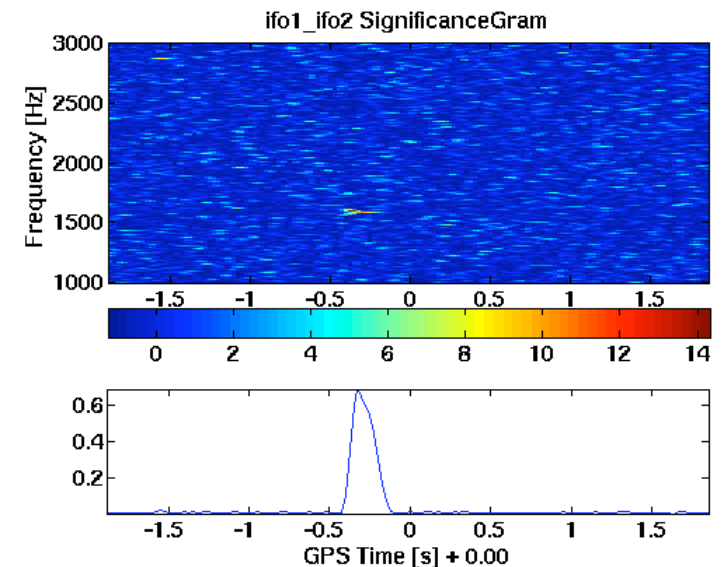
Suited for stochastic (WNB) SGR search

Less timing precision needed

$N^{1/4}$  amplitude sensitivity dependence in WN

Cons:

Less sensitive than coherent method (T-Stack)





# SGR 1900+14 storm

Proposed stacking scenarios (preliminary)

Flat (N=11) – most energetic bursts

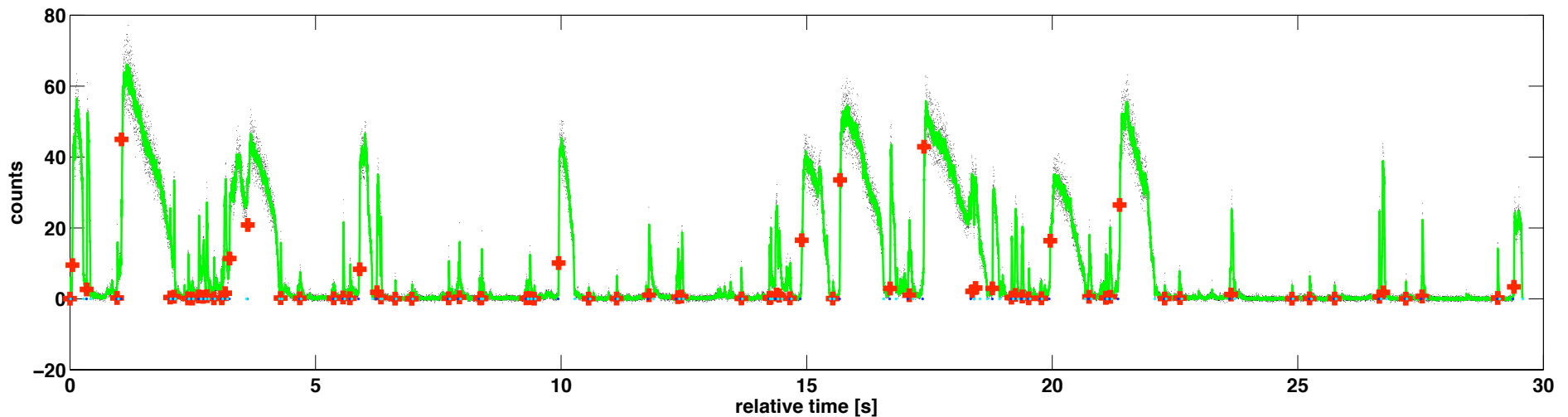
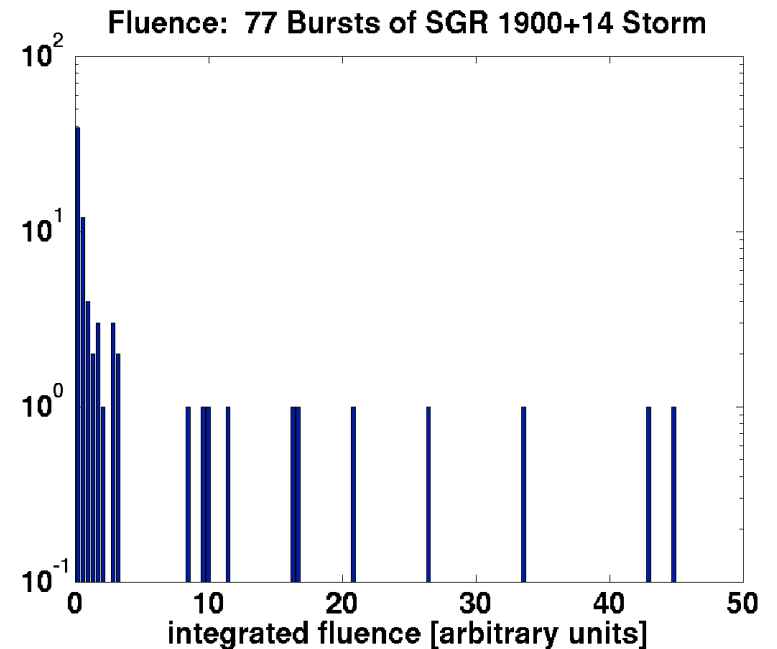
Integrated fluence weighting (N=77)

Timing error can be simulated

5 ms preliminary timing error

Swift BAT light curve, 100 us bins

G. L. Israel et al., ArXiv e-prints 805 (2008), 0805.3919





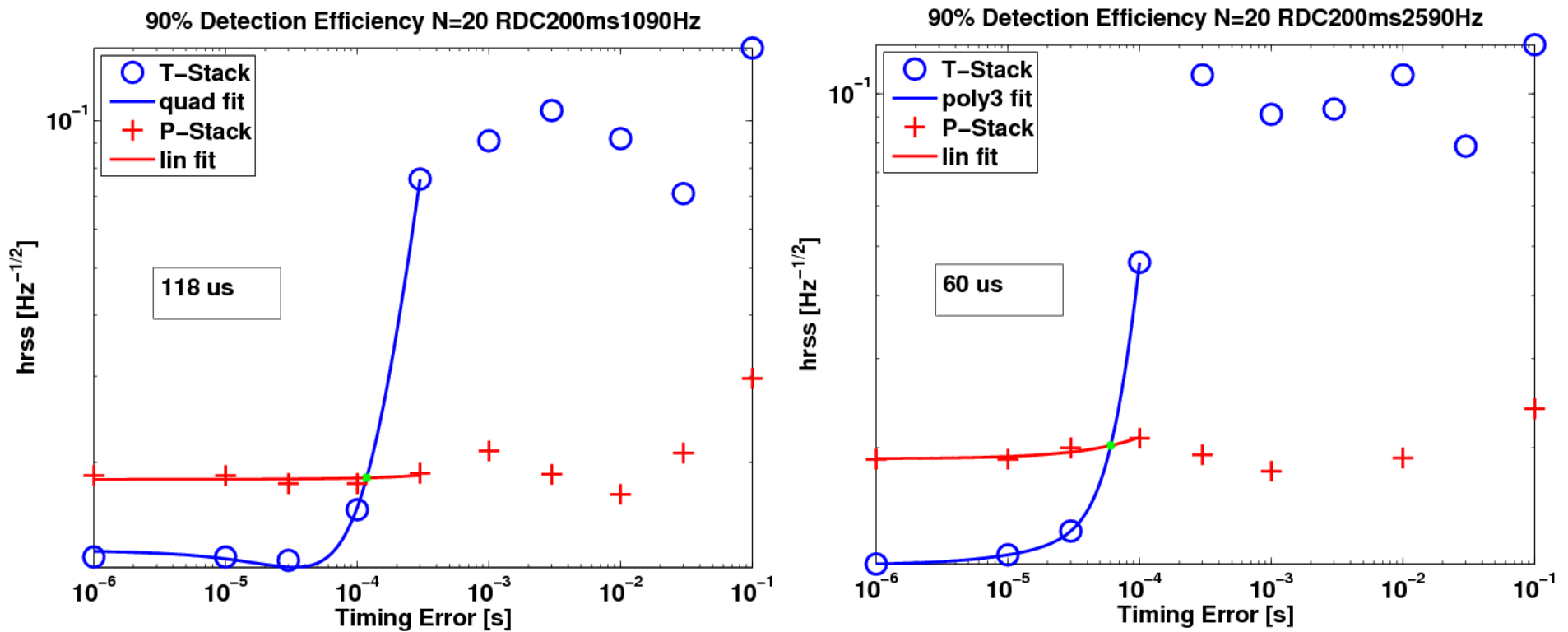
# Stack-of- $\sigma_{\Delta T}$

Timing Error is  $\sigma$  of distribution used in Monte Carlo

P-Stack (red) is insensitive to time shifts (so long as signal TF pixels overlap)

T-Stack is sensitive. Crossover points are given

hrss50 plots look very similar





# Storm search strategy

On-source choice  $\pm 2$  s **as before**

BAT timing precision:  $< 100$  usec (after propagation to geocenter)

GW emission delay: probably less than 100 ms (S5y1 SGR paper)

Empirical test:  $\pm 1$  s hrss90 limits are 2.4% lower than  $\pm 2$  s (24 trials)

$[-1000, 1000]$  second background region **as before**

estimate  $\mu(f)$ ,  $\sigma(f)$  used by Flare pipeline

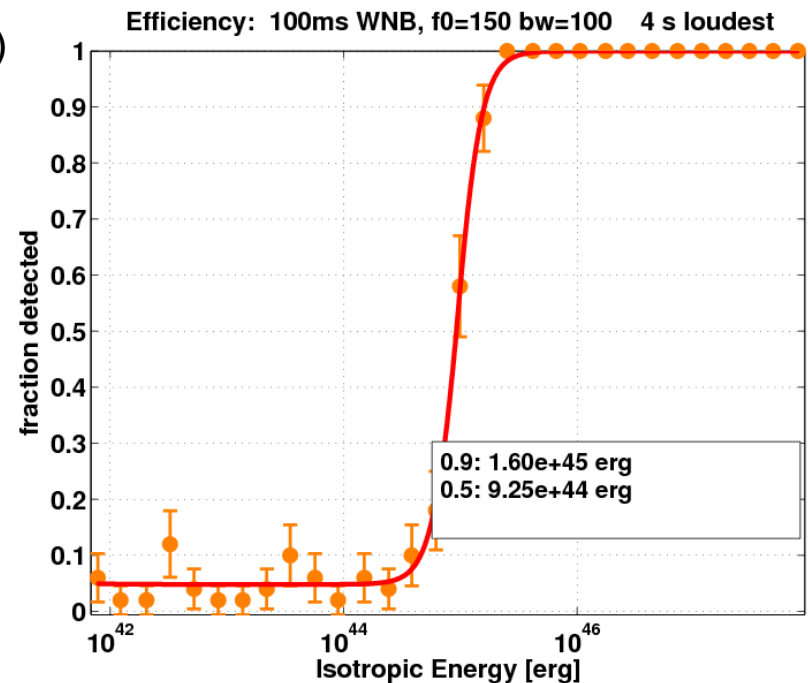
estimate local false alarm rate (FAR)

background data treated identically (i.e. stacked)

same bg region; choose time slide (new to stack)

Follow up events with significant FAR **as before**

Loudest event upper limits **as before**







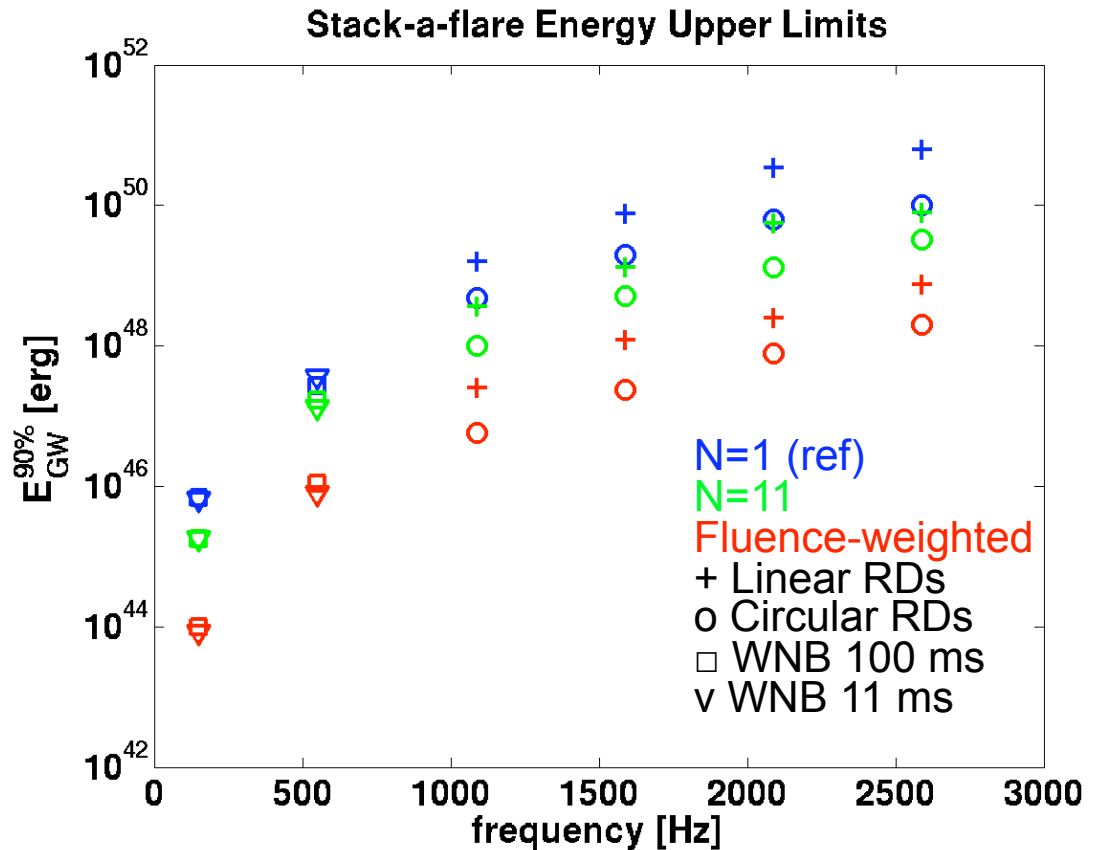
# Closed Box Results Energy Limits

SGR 1900+14 simulation  
400 s after open box trigger  
4 s on-source  
50 trials per result  
5 ms simulated time error

Flat N=11  
 $11^{1/2} = 3.3$  (expected, over N=1)  
Mean = 4.4 (observed)

Fluence-weighted  
Simulation mean  $E_{GW}$  is presented

Isotropic GW energy at 10 kpc  
90% detection efficiency



# PRELIMINARY

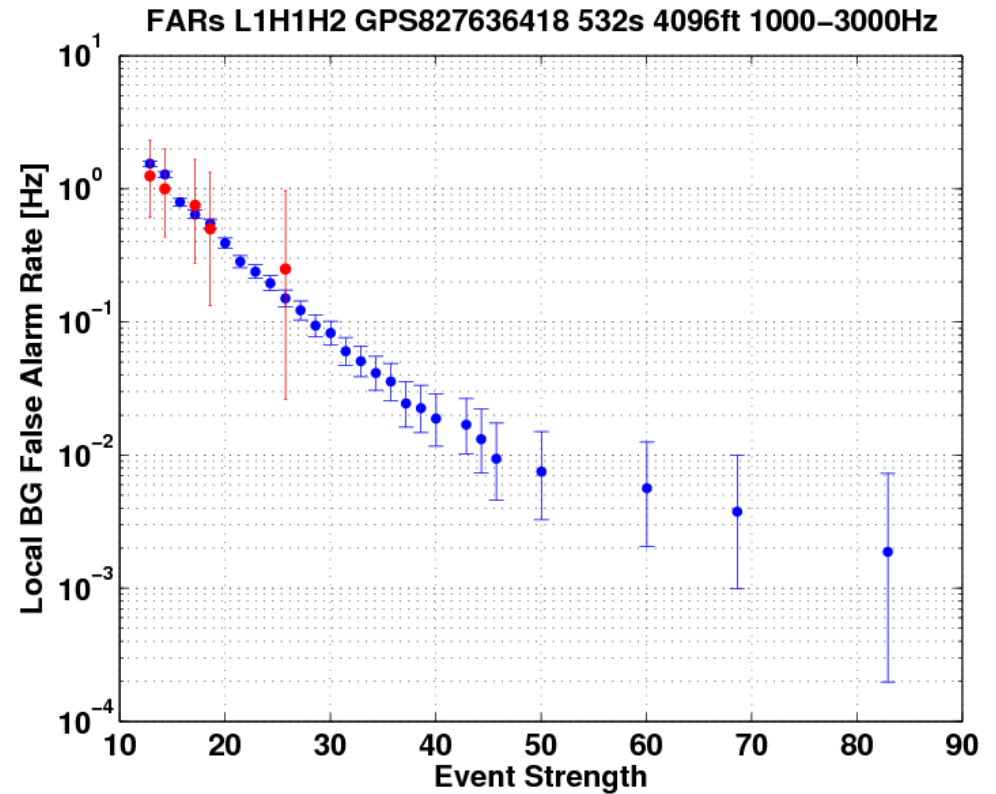


# Closed Box Results Detection Statement

## PRELIMINARY

Same procedure as before  
Live time is stacked live time  
T is time from stack center

Example here:  
Flat N=11 1000-3000 Hz on-source events



no.	Z	FAR(local)	T	DT	F	DF	N
1	2.61e+01	1.43e-01	1.453	0.100	1044.0	0.0	5
2	1.86e+01	5.43e-01	-1.218	0.200	1549.0	4.0	4
3	1.76e+01	6.11e-01	-1.324	0.050	2280.0	0.0	3
4	1.45e+01	1.23e+00	0.811	0.025	1469.3	4.0	3
5	1.30e+01	1.52e+00	1.103	0.150	2713.3	12.0	3



# Documentation

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Project summary page:

<https://www.lsc-group.phys.uwm.edu/cgi-bin/bag-enote.pl?nb=burs5trig&action=view&page=33>

Includes links to config files

Includes instructions for installation and running

Method is described in a chapter of my thesis

Methods paper draft has been started

Results paper not started

Code is in DASWG CVS:

<http://www.gravity.phys.uwm.edu/cgi-bin/cvs/viewcvs.cgi/matapps/src/searches/burst/Stac?cvsroot=lscsoft>

This talk has addressed the burst group's review readiness requirements



## EXTRA SLIDES

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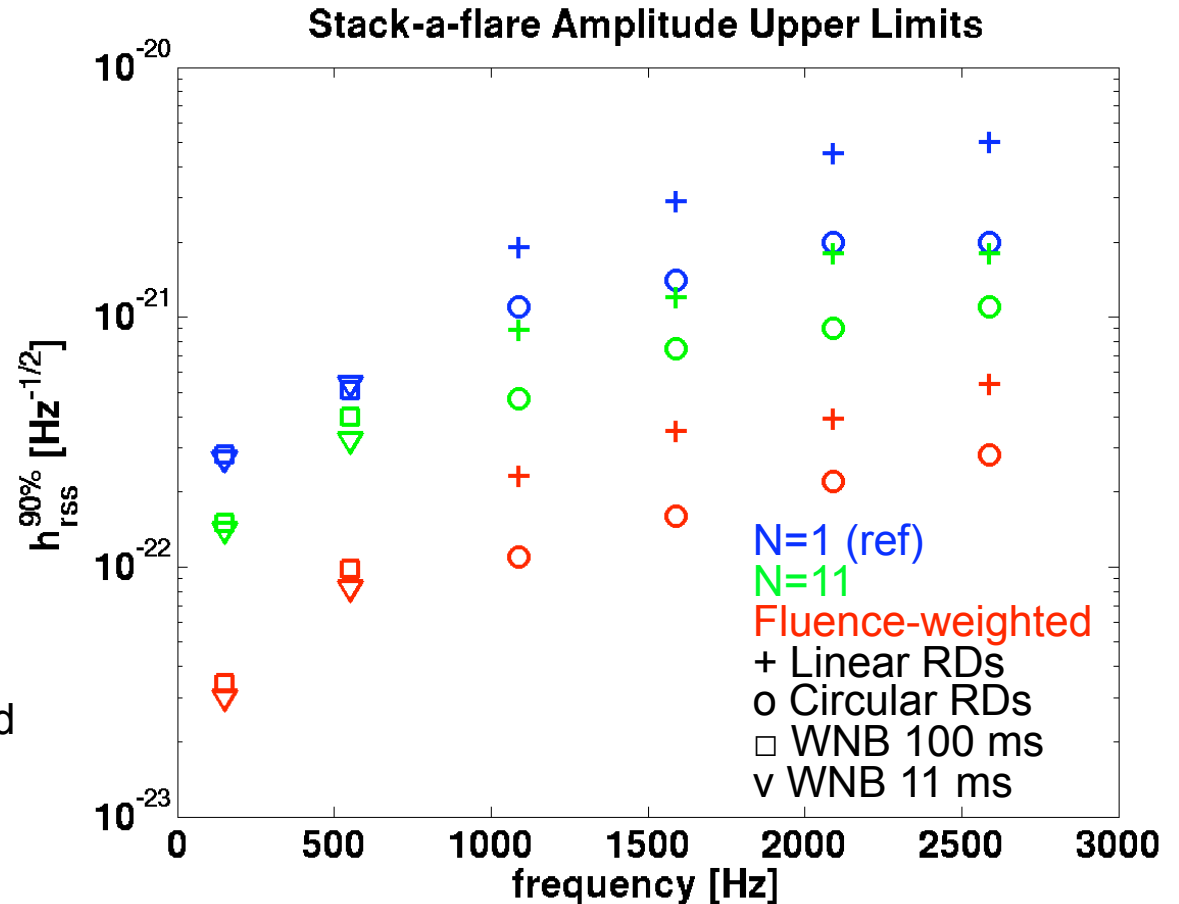


# Closed Box Results Amplitude Limits

SGR 1900+14 simulation  
400 s after open box trigger  
4 s on-source  
50 trials per result  
5 ms simulated time error

Flat N=11  
 $11^{1/4} = 1.8$  (expected, over N=1)  
Mean = 2.1 (observed)

Fluence-weighted  
Simulation mean hrss is presented



# PRELIMINARY



# Stacking Isolated Bursts

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## P-Stack timing:

- Propagation of light crossing times to detectors
- Light curve analysis for precise burst start times

Weight each burst according to antenna pattern before stacking

## Combining bursts

- All bursts from all SGR sources?
- All bursts from particular SGR sources?
- Brightest N bursts from a given source?
- Weight each burst according to fluence?

Theorist input required

# Future of Stack-a-flare



**P-Stack** is worth developing  
storm unmodeled search, isolated bursts

**T-Stack** issues

Requires timing precision of ~50 us or better to be worth it  
Cannot do WNB

Concentrate first on the storm

Relative timing between bursts **much** easier to get precisely

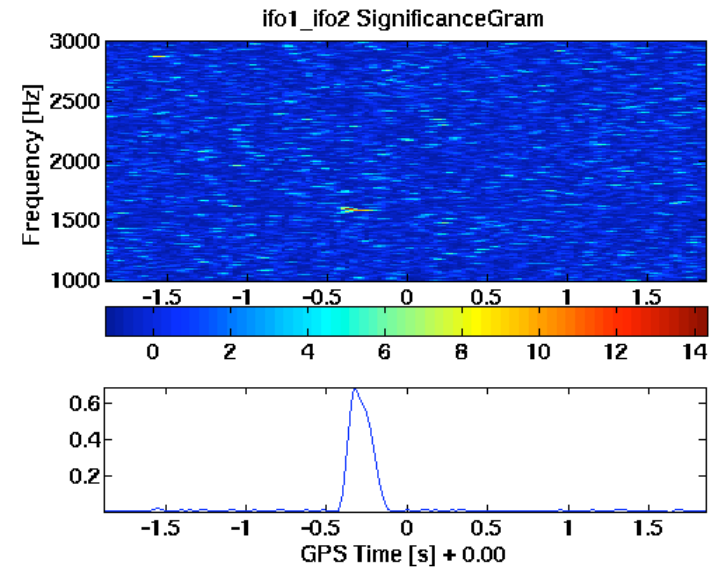
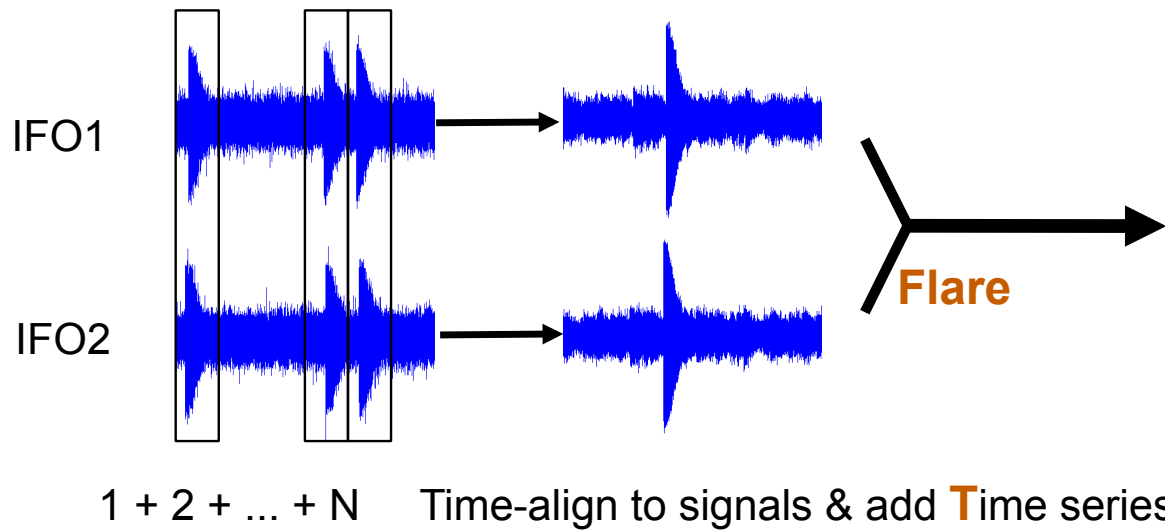
**T-Stack** *may* be possible; theorist input needed

Get **P-Stack** S5 closed box results

Time-to-publish should be small, thanks to Flare LSC review

Experience from storm work may help approach isolated bursts

# T-Stack Method



Pros:

Greater potential sensitivity. Stacking *amplitude*  
 expect  $\sim N^{1/2}$  amplitude sensitivity dependence in WN

Cons:

Either precise timing, or expensive time shift combinatorics, is needed  
 Sensitive to relative sign between detectors

**T-Stack will NOT be used for initial search**





# Prototype Characterization

Benchmark: individual burst simulations & vanilla Flare

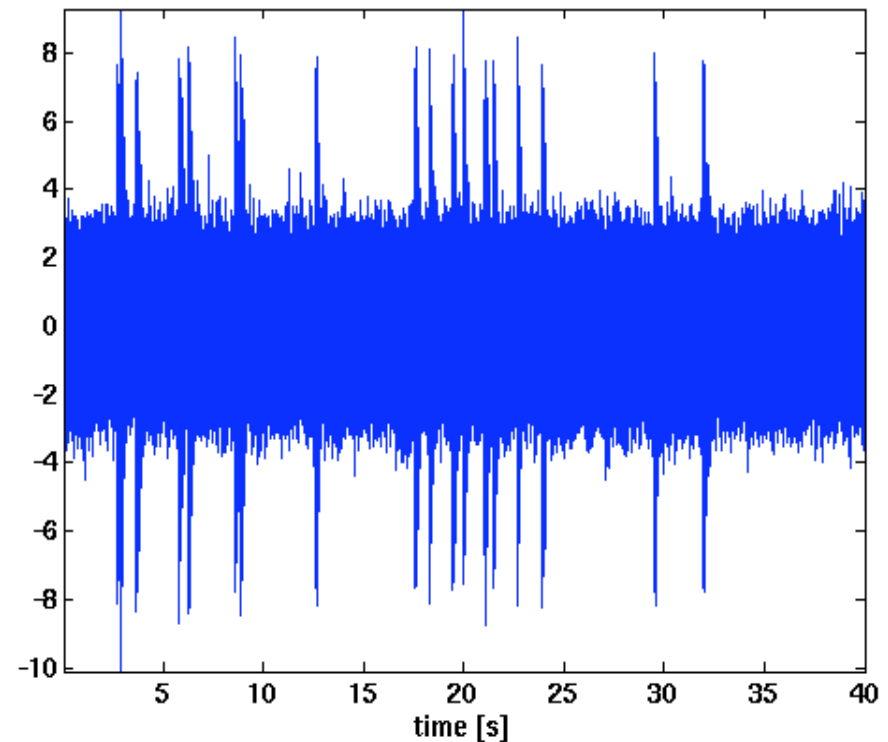
Characterize:

sensitivity dependence on **N**

sensitivity dependence on  $\sigma_{\Delta T}$

Shift component injections by small times  $\Delta T$   
simulates timing imprecision  
subsample shifting

Two simulation designs tell consistent story  
Rough simulated SGR 1900+14 storm  
18 identical ringdowns into  $\sigma=1$  WN  
**N** evenly spaced injections (RDs or WNBs)

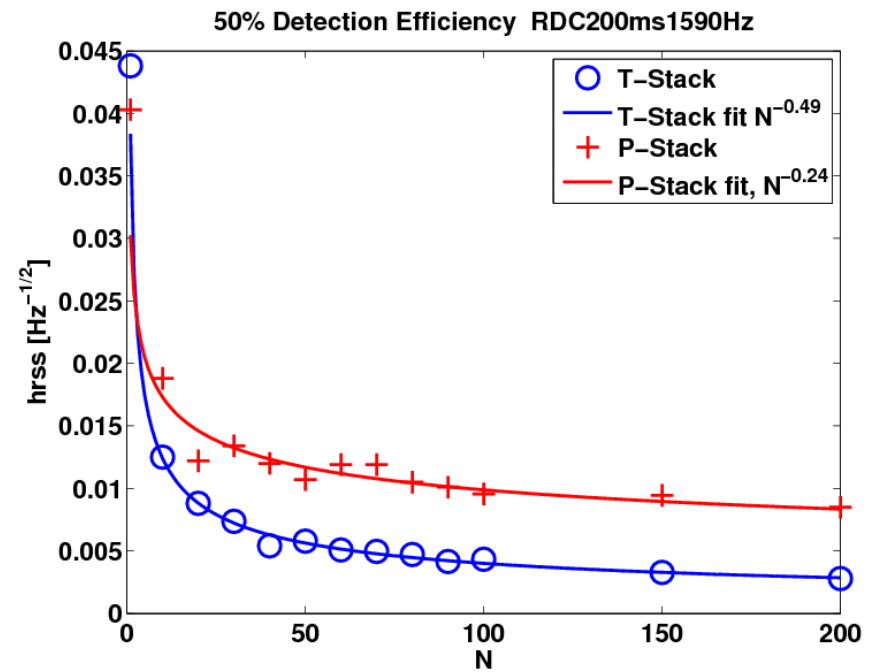
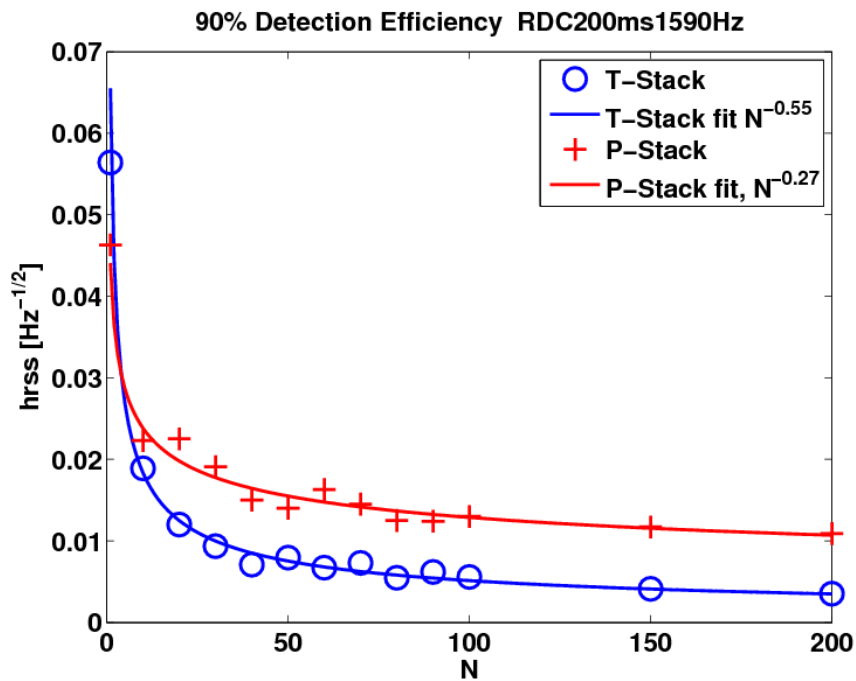




# RD Stack-of-N

P-Stack goes as  $N^{1/4}$

T-Stack goes as  $N^{1/2}$





# WNB Stack-of-N

P-Stack goes as  $N^{1/4}$

T-Stack is flat

