

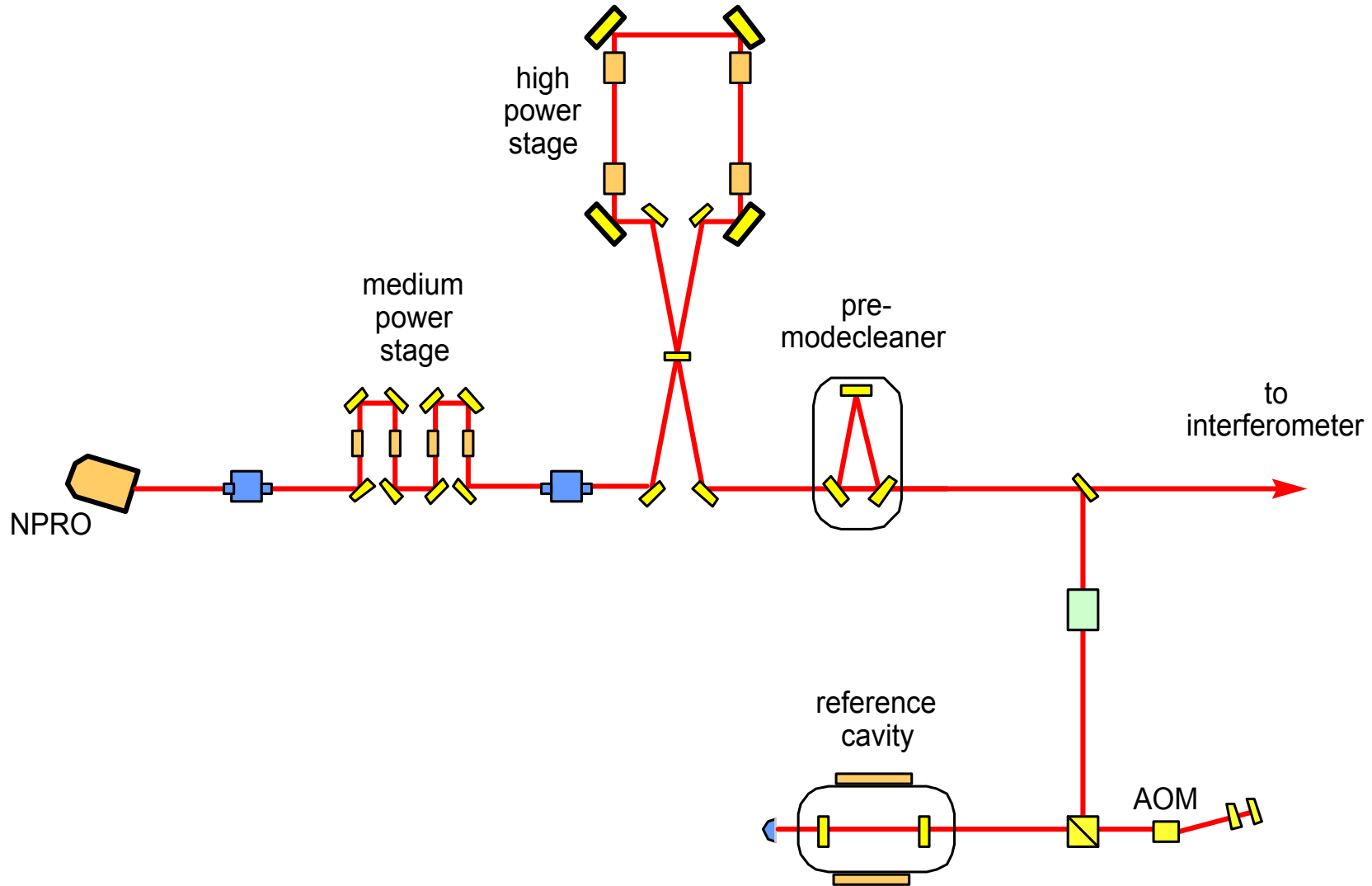


Status of the Advanced LIGO PSL development

Benno Willke
for the PSL team

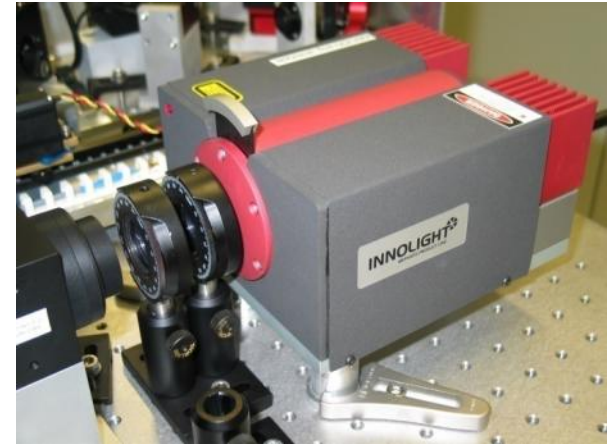
LSC Virgo meeting, Caltech, March 2008

LIGO G080143-00-Z



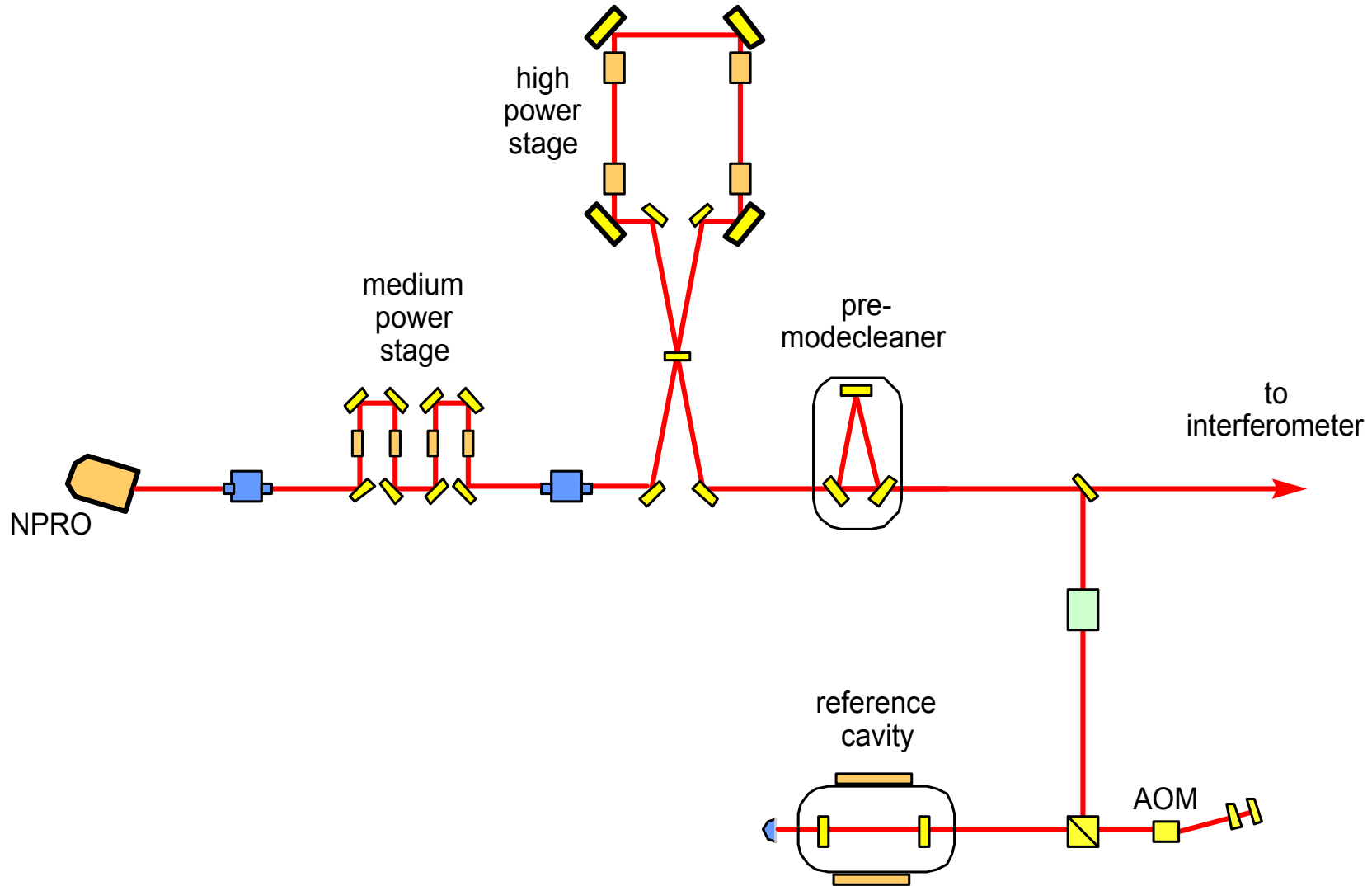


- all 8 Master lasers (2W NPROs) are delivered
- they have a special interface and specially selected LDs
- characterization program
 - power, slope, power in p-pol
 - RIN:
 - noise spectrum 1Hz – 100kHz,
 - time series (60min) rms
 - frequency noise
 - spectrum 1Hz – 100kHz
 - upper limit for drift
 - PZT and slow actuator calibration
 - beam quality
 - higher order mode content
 - beam pointing
- LDs power is de-rated by 30% to increase their lifetime and give headroom for power adjust



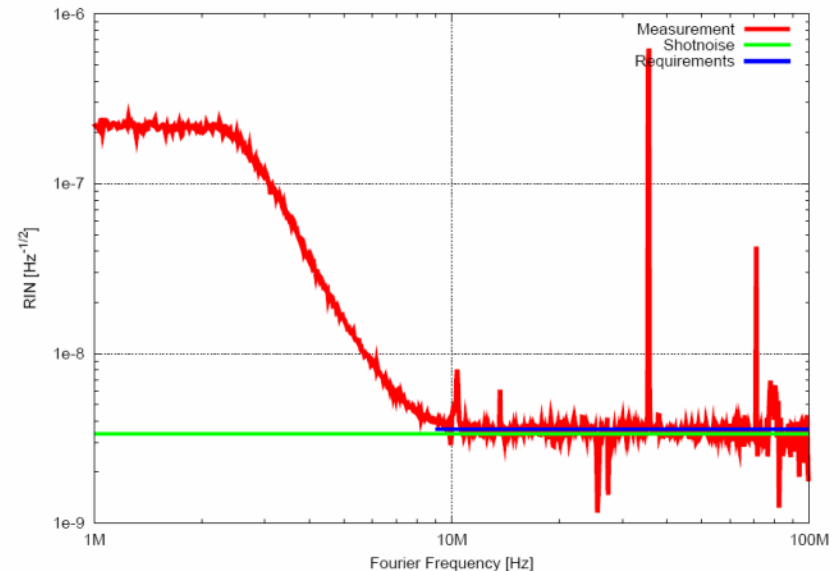
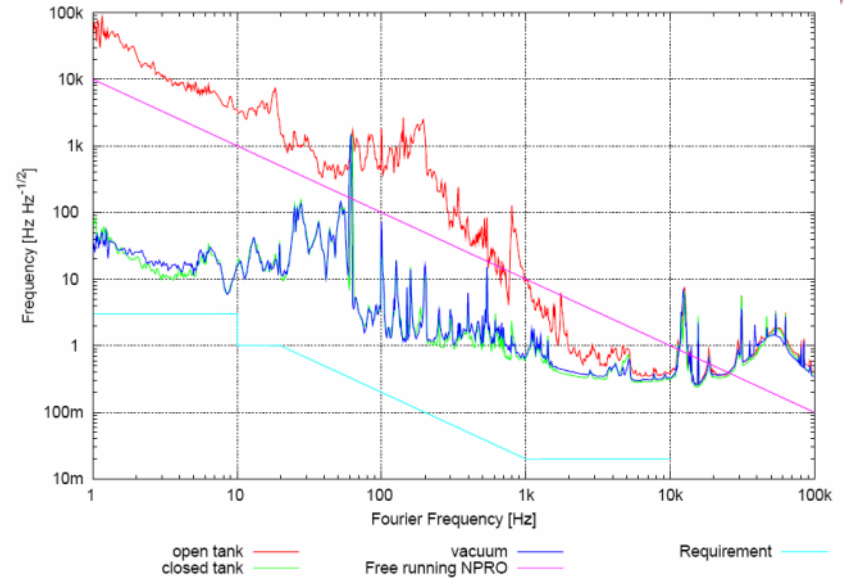
LD1 Power Monitor
GND
LD2 Power Monitor
GND
CTC1 Error
GND
n.c.
GND
n.c.
GND
LDTTC1 Error
GND
LDTTC2 Error
GND
Guard LDTC1
GND
Guard LDTC2
GND
n.c.
GND
n.c.
GND
NE Monitor
GND
Interlock

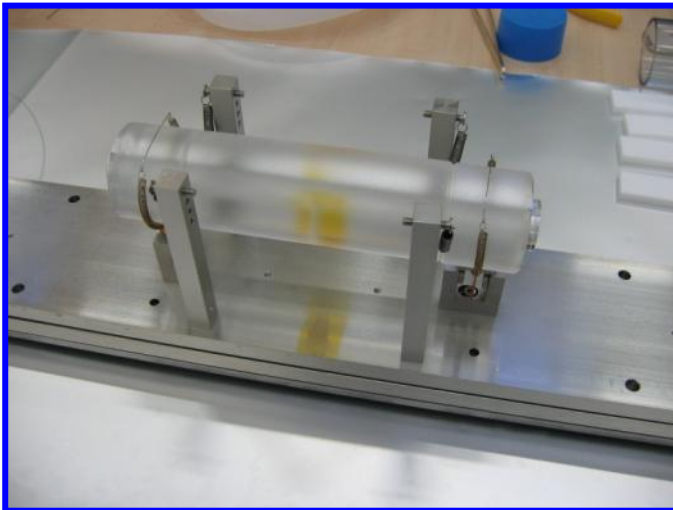
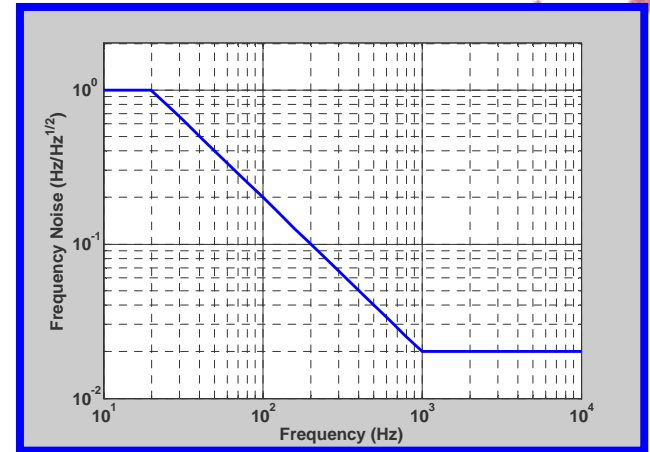
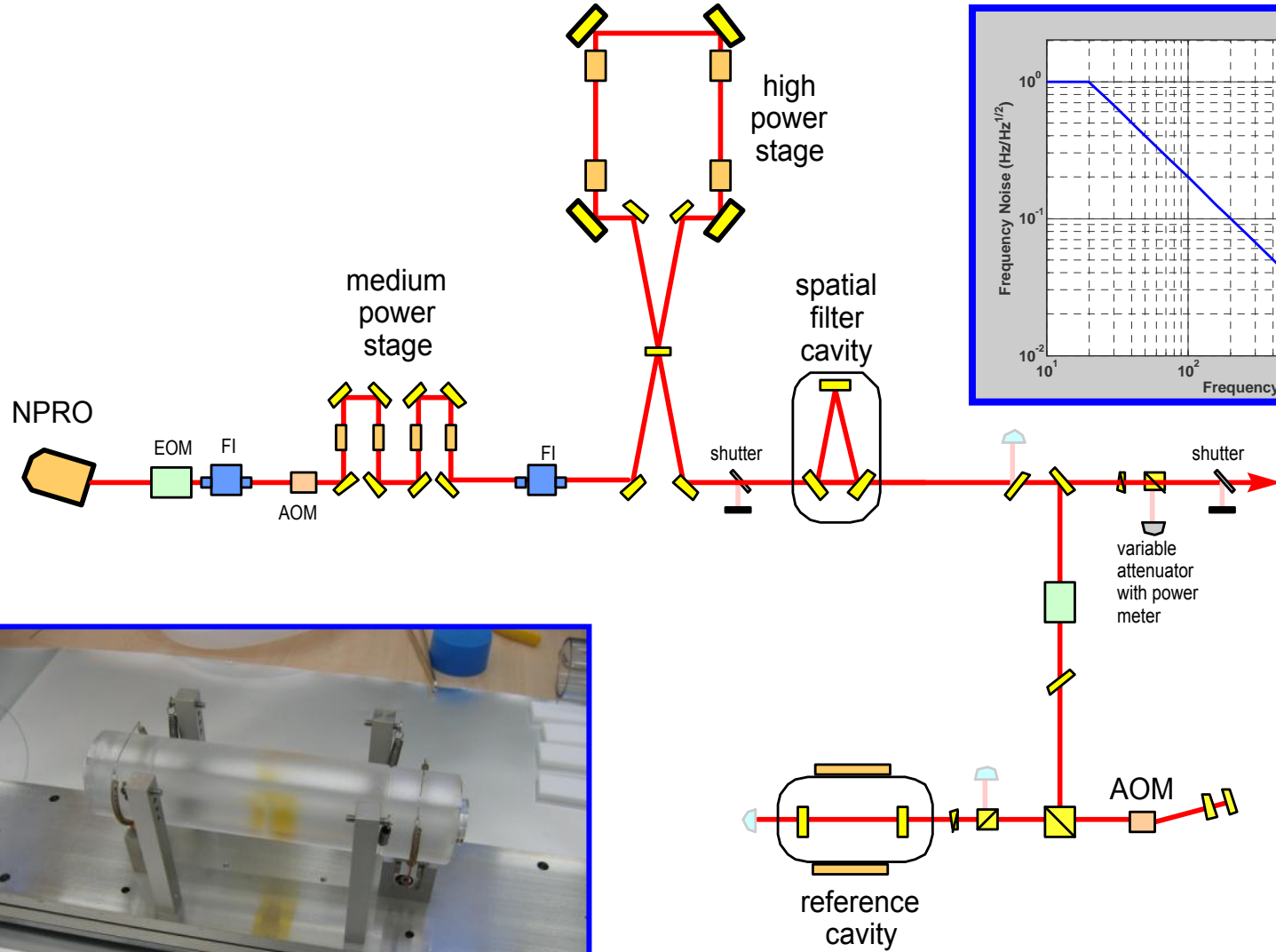
Set Temperature LD1
GND
Act. Temperature LD1
GND
Set Temperature LD2
GND
Act. Temperature LD2
GND
Set Temperature Crystal
GND
Act. Temperature Crystal
GND
Set Current
GND
Act. Current
GND
Status Laser ON
GND
n.c.
GND
Remote SW
GND
Remote SW ON
GND
Remote SW OFF

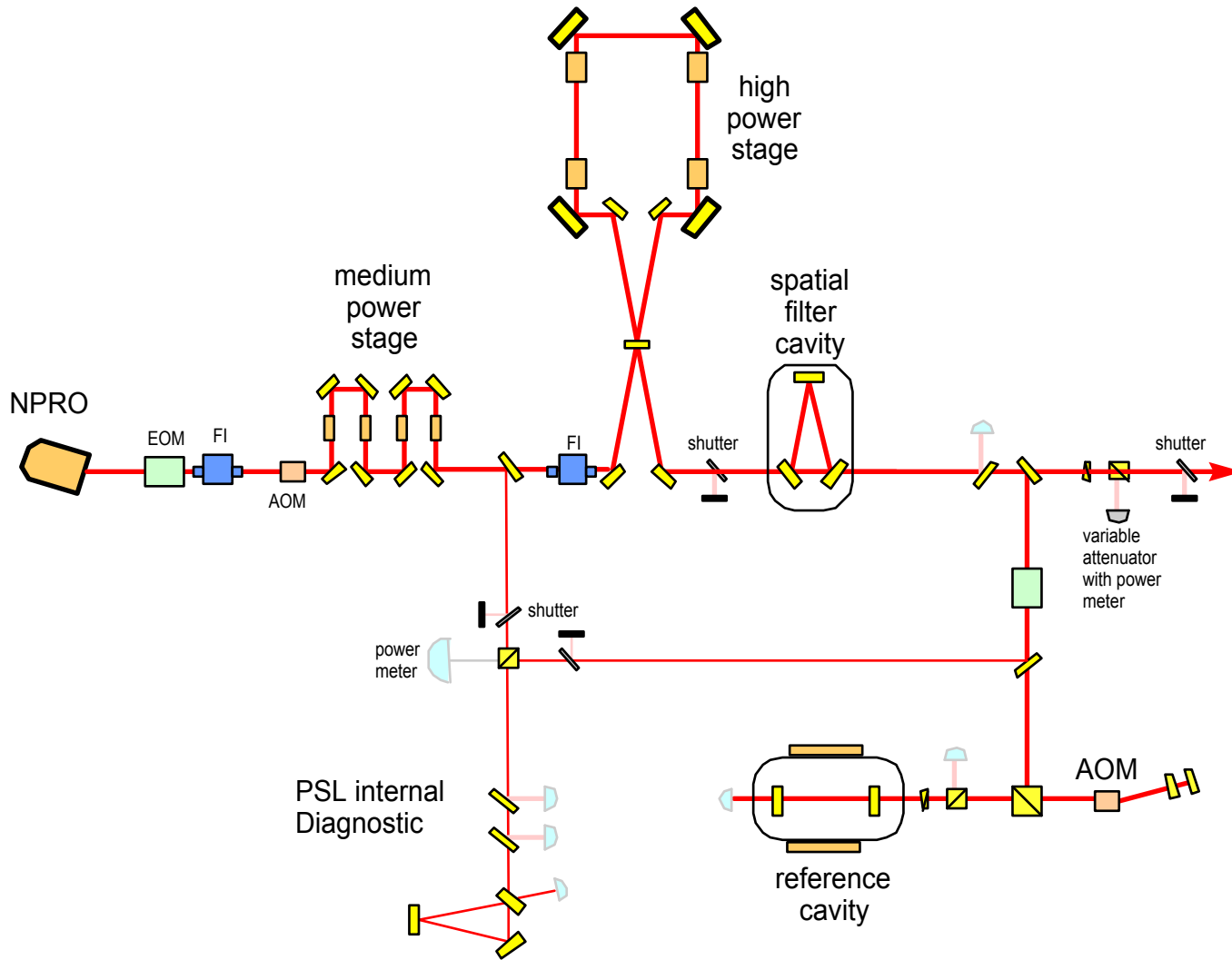




- first prototype fabricated at AEI
 - Finesse 46 / 383
 - since Jan 2008 operated in s-pol with 35W injected (60% higher circulating power than in AdvLIGO)
 - PZT
 - FSR \leftrightarrow 140V
 - first resonance @ 10kHz
 - aluminium spacer to allow for long range thermal actuator
- in sealed housing, probably no vacuum required
- rf filtering
 - 4dB @9MHz, sufficient to meet AdvLIGO requirements

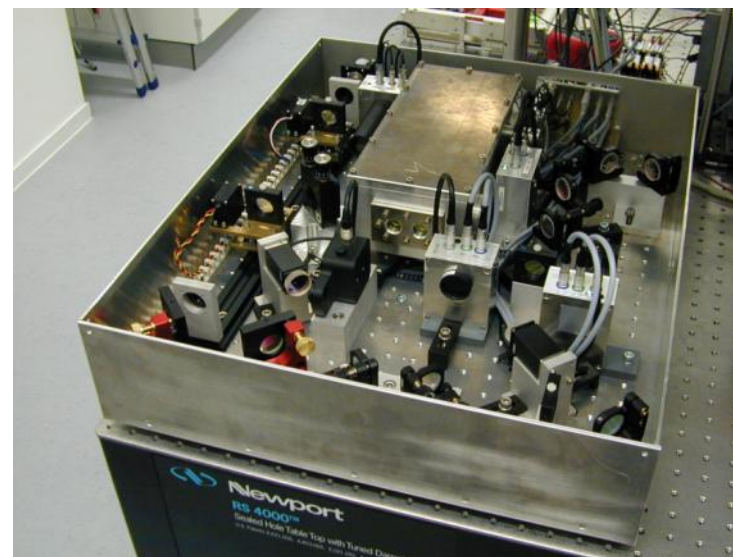
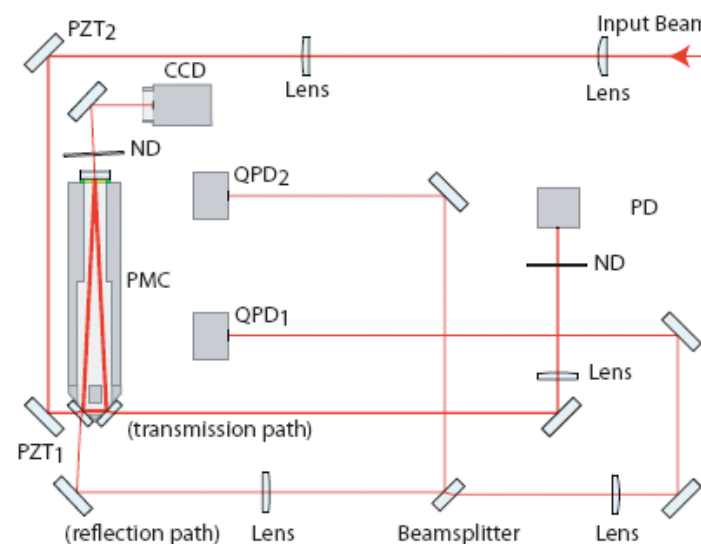


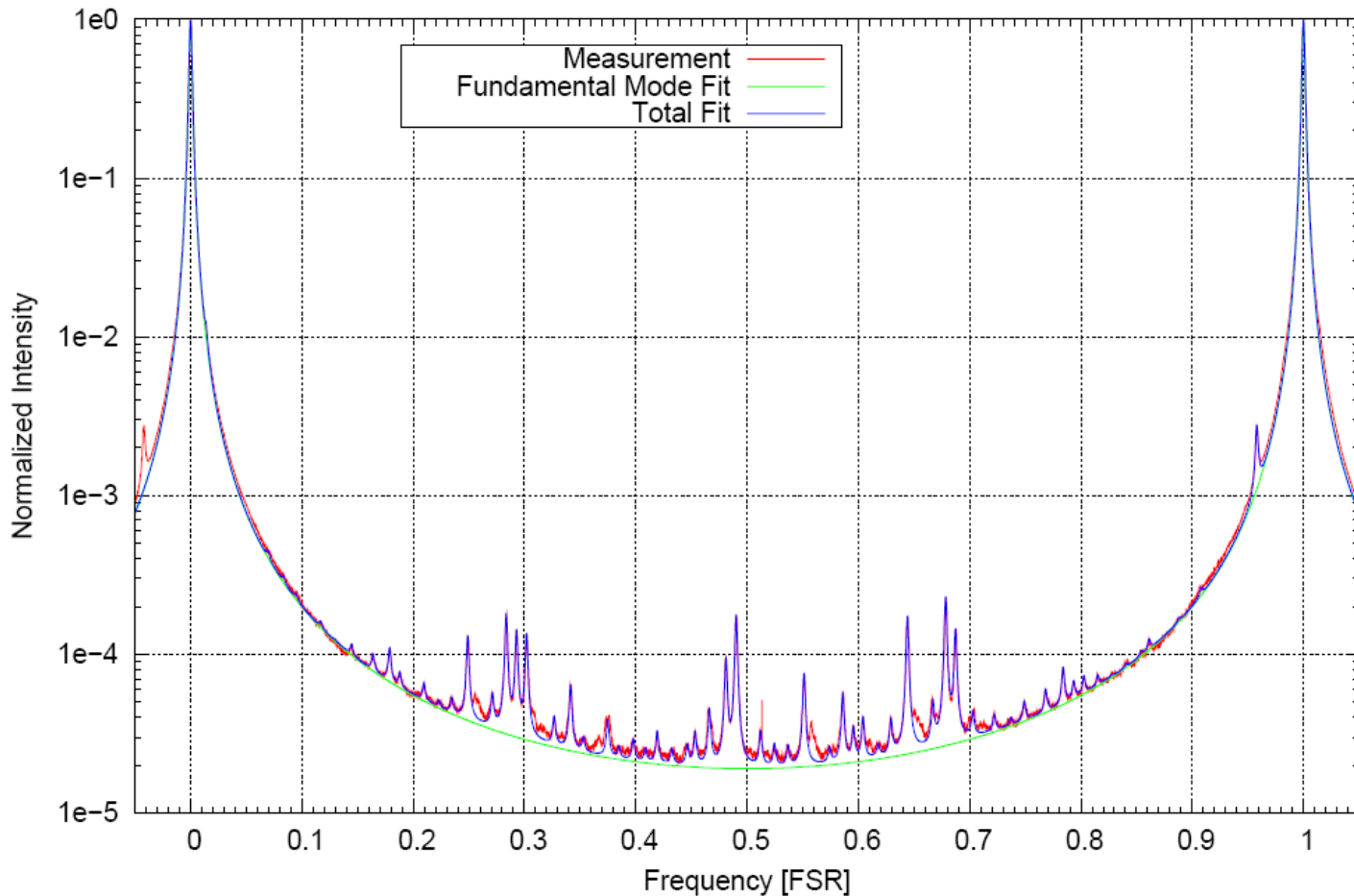




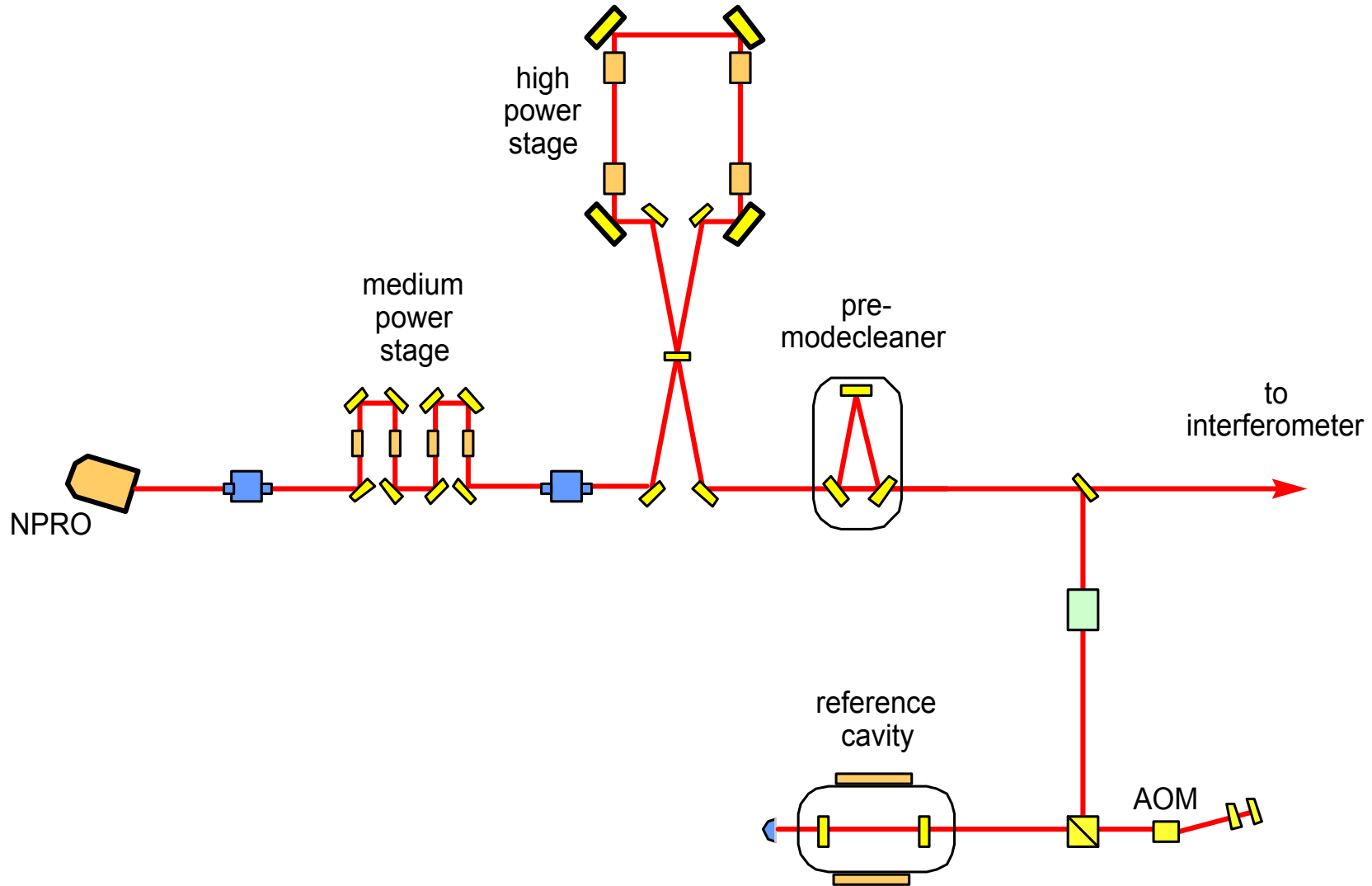


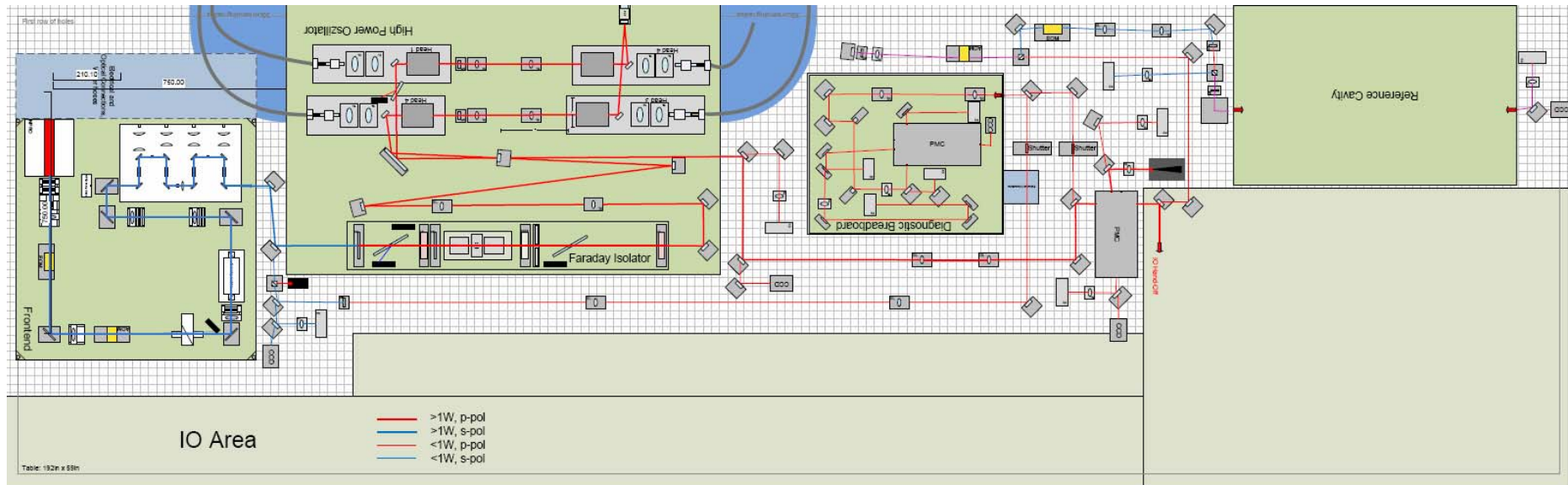
- diagnostic tool to measure
 - power noise (low f and rf)
 - frequency noise
 - higher order mode content
 - beam pointing (differential wavefront sensing)
- fully automated
 - automatic length and alignment control
 - can switch from lock to scan mode
 - performs complete beam analysis without human interaction (at night during long term test)
 - allows fast turn around between laser optimization and characterization



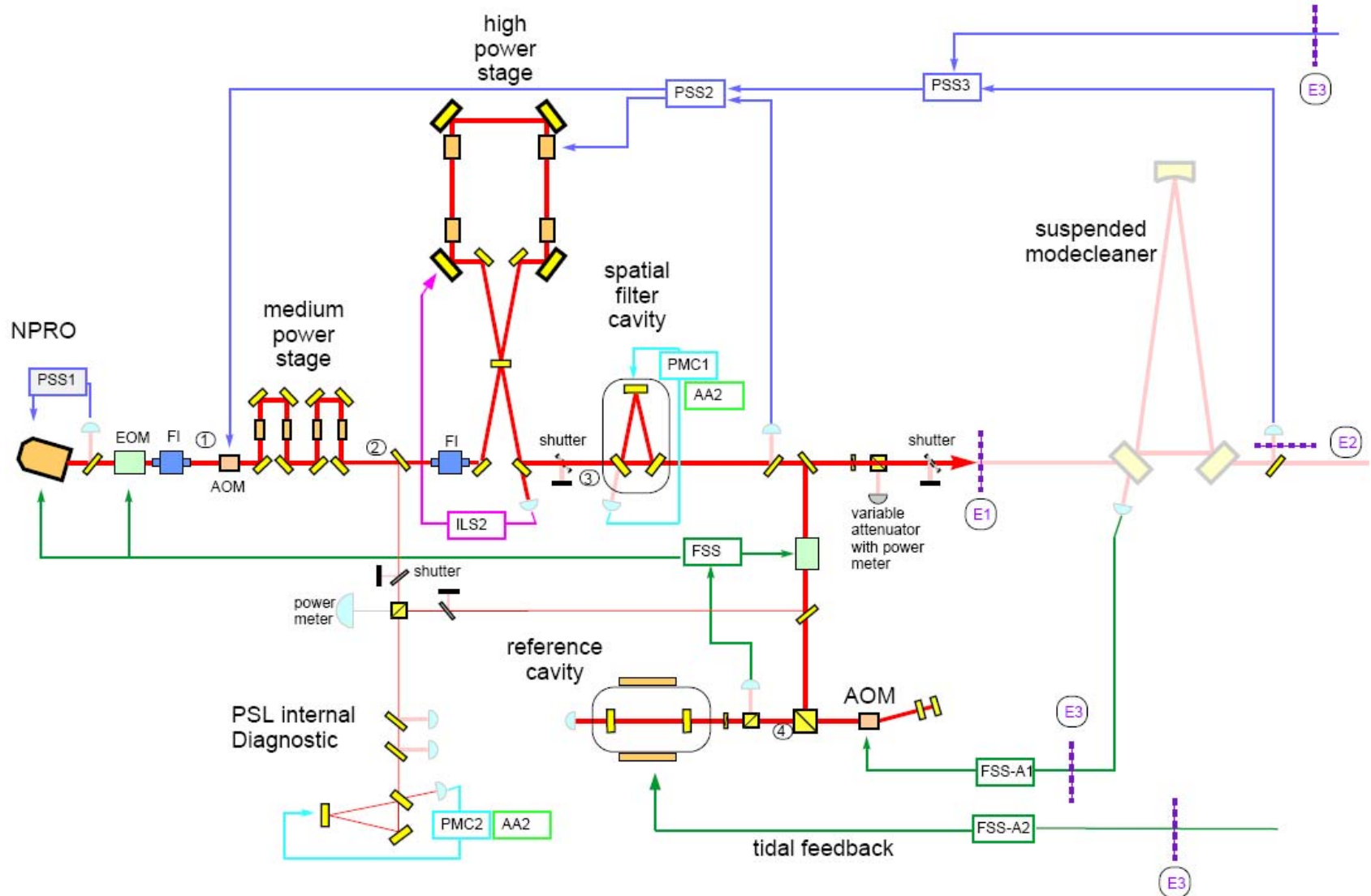


Finesse: 366 ± 5
 higher order mode power: $0.56\% \pm 0.3\%$





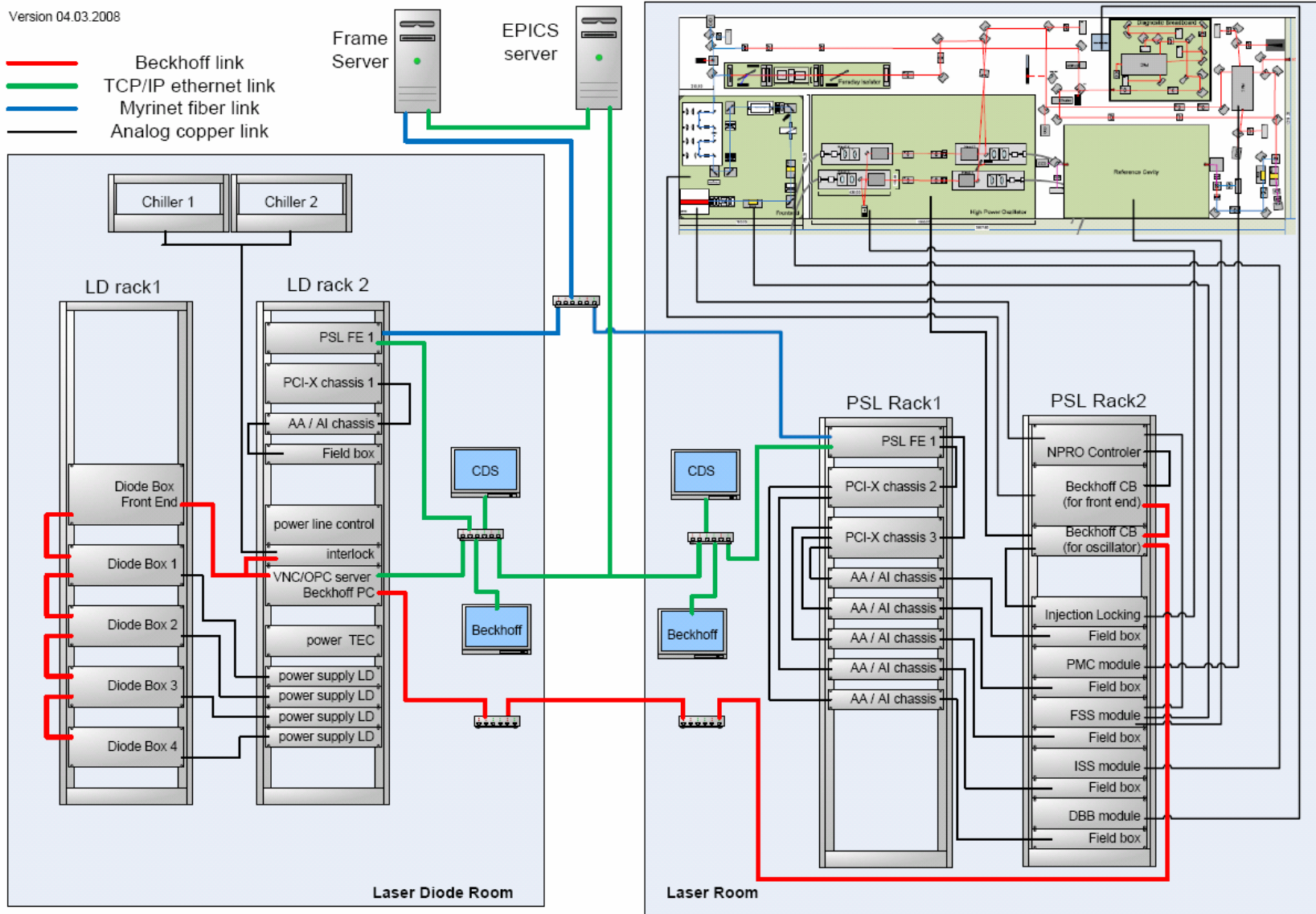
- real estate on the PSL table is currently negotiated between IO and PSL
- layout sketch above will change
- PSL table more crowded than in initial LIGO
- table size probably just big enough
- can DBB be used for IO diagnostic purposes ?
- cleanliness, easy access, small number of components in 180W laser path

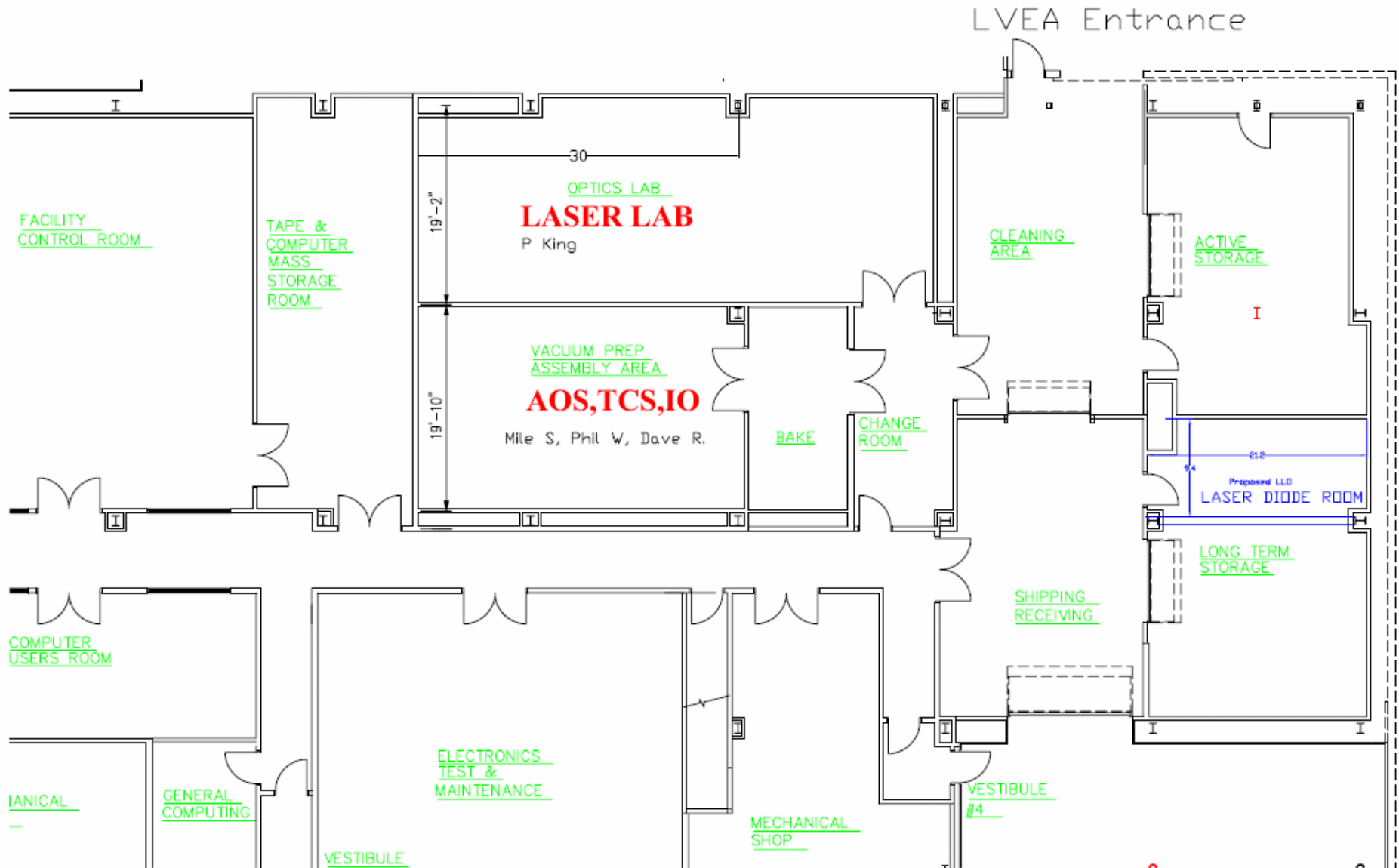


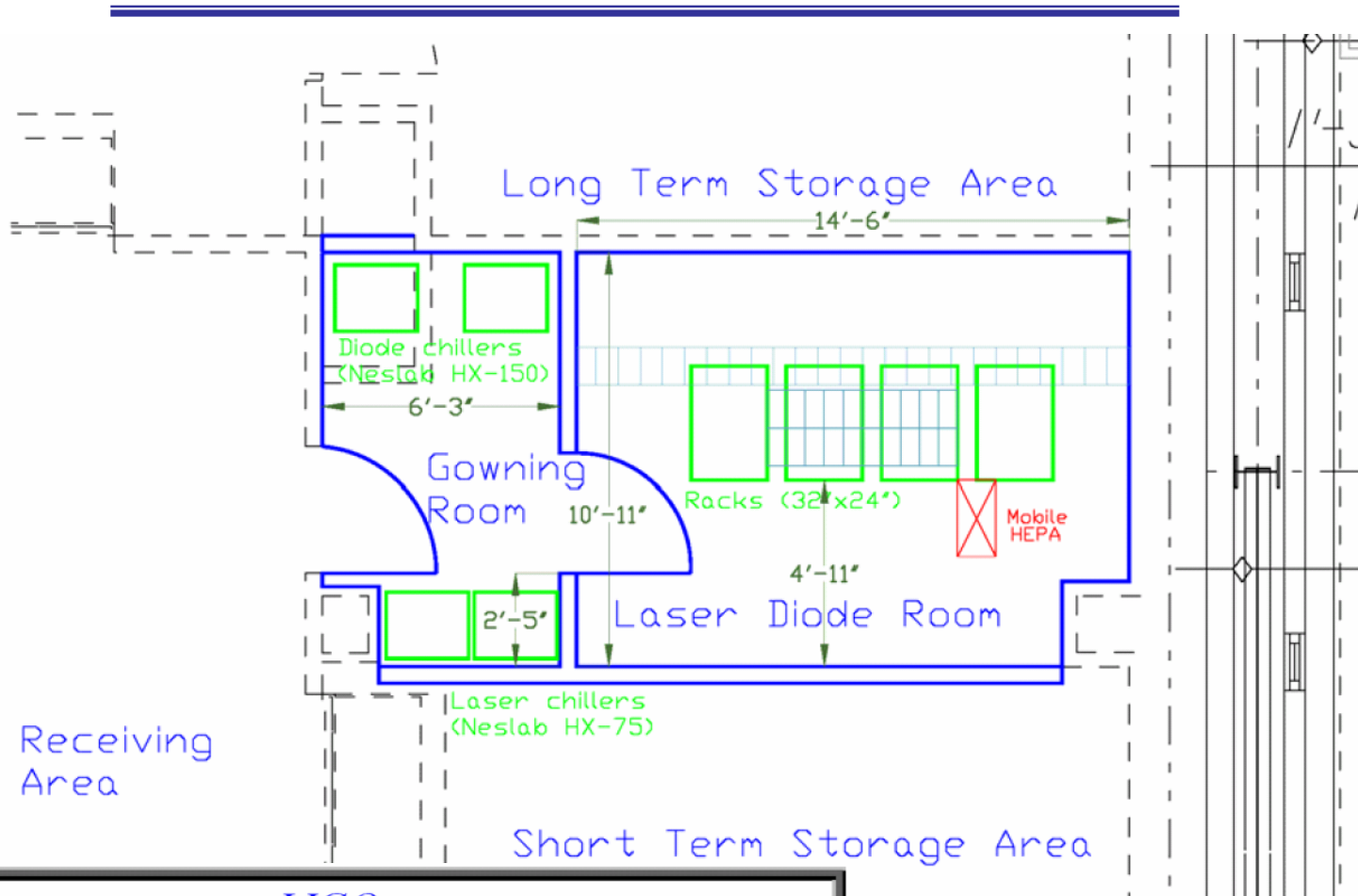


Version 04.03.2008

- Beckhoff link
- TCP/IP ethernet link
- Myrinet fiber link
- Analog copper link







LIGO-T070195-04-W

LIGO

Aug. 17, 2007

Revised Sept. 5, 2007

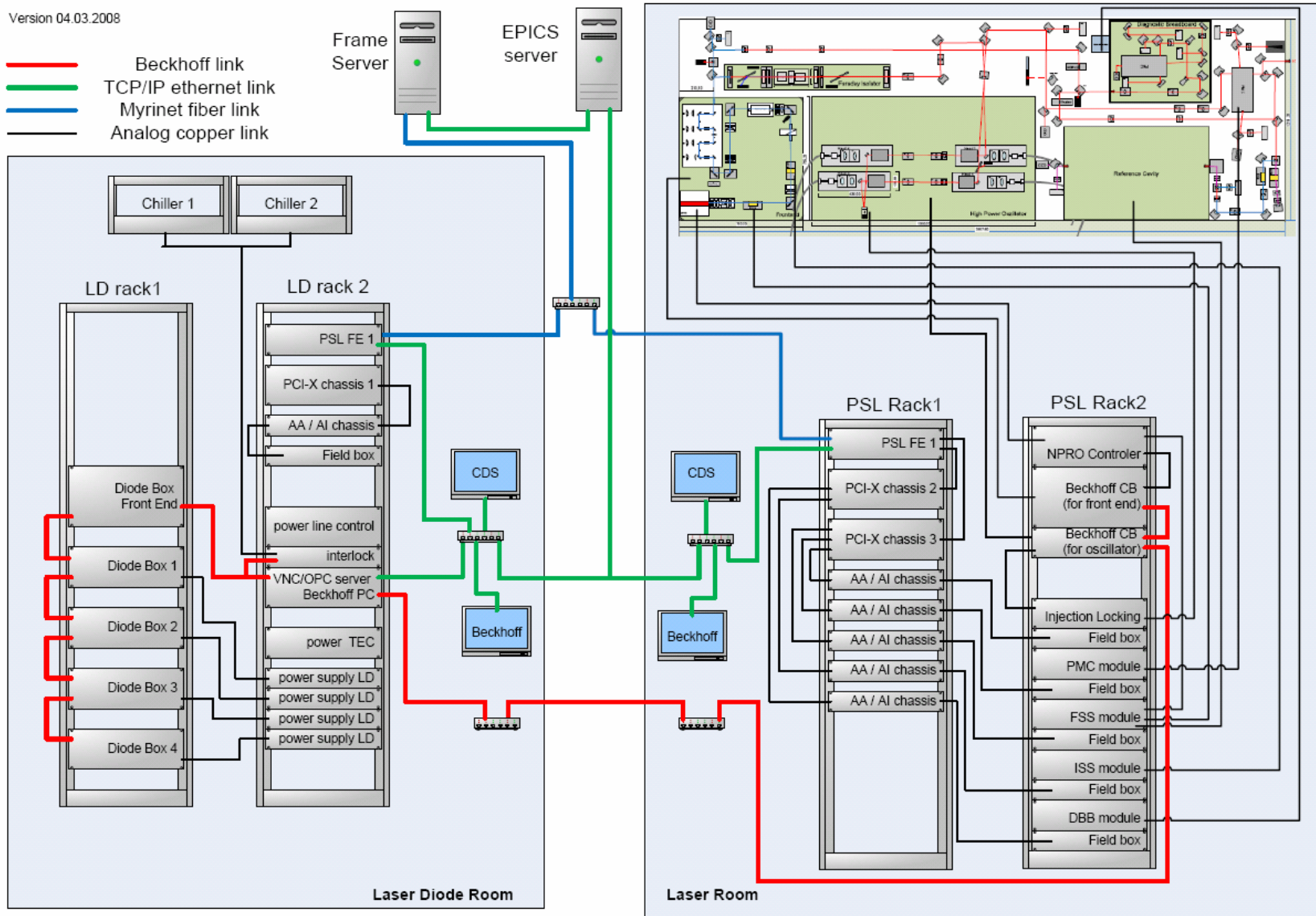
Preliminary Design – AdvLIGO PSL Laser Diode Room

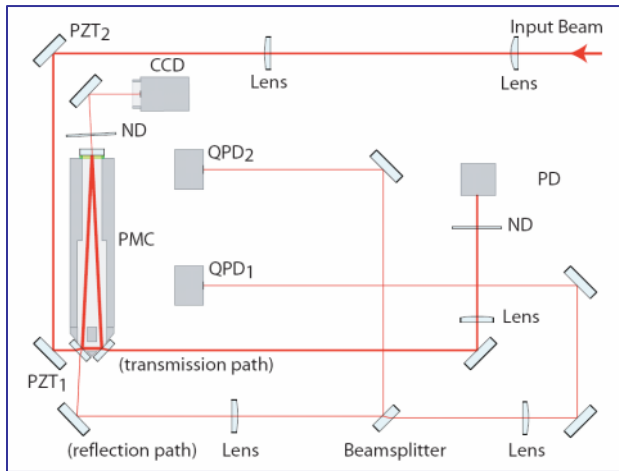
A. Effler, P. King, R. Savage, R. Wooley



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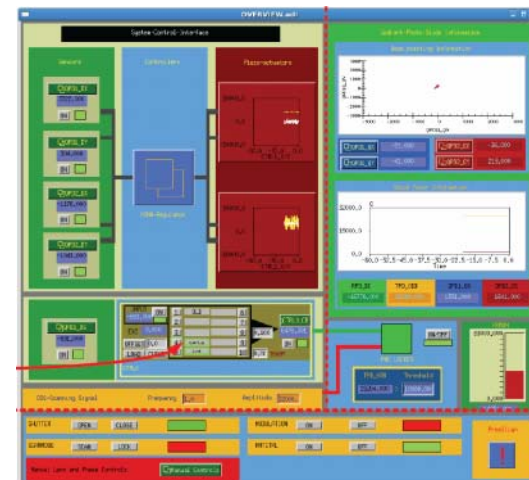
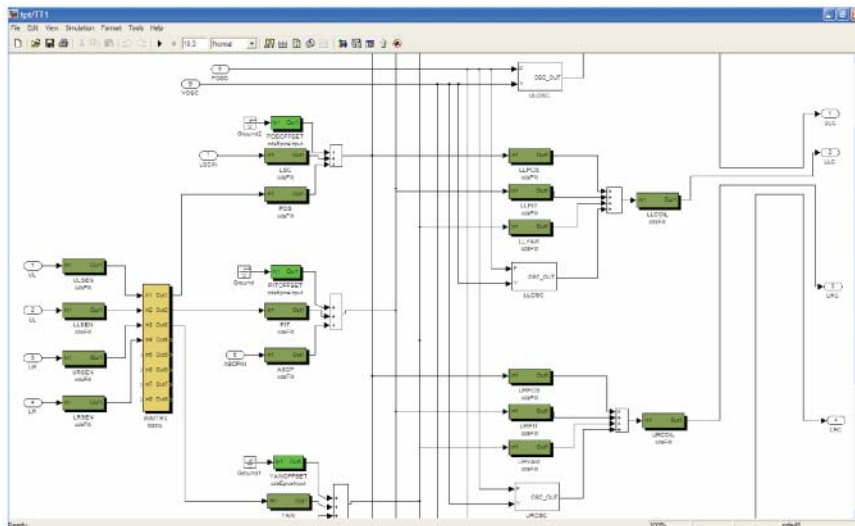
- RTLinux/Epics system was installed at AEI
- length and alignment control of ring cavity in DBB via digital control
- now used to control FSS
- next step: automate lock acquisition for ref. cavity and include PMc

STUDIENARBEIT

Erweiterung eines CAD-Systems um die Funktionen der Modellbildung und des Reglerentwurfs

von

Anatoli Fedynitch





- finalize functional prototype of 180W laser and move to engineering prototype
- measure free running noise and actuator transferfunctions to design Pstab loop
- continue long term performance measurements and integration test at AEI
 - LIGO provides additional DAC/ADC hardware
 - implement PMC and ISS loop
 - install automation and measure cross couplings between loops
 - find phase delay in FSS and test different sensing points
- PSL preliminary design review in summer