



LIGO G080082-00-R



Optimization of Coating Design for Reduced Thermal Noise

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LSC / VIRGO Joint Meeting, 17-20 March, 2008, Caltech
Workshop on Optical Coatings in Precision Measurements



SUMMARY



Coating Optimization Method:

Stacked Doublet Coatings
Barebone Optimization Code
Coating Noise Model;
How Did We Get Here -
Genetic Selection
Another Look at the Optimum

Results:

Prototype(s)
Robustness
Losses
Event Rate Boost
Usable Reflectance Windows

Perspectives:

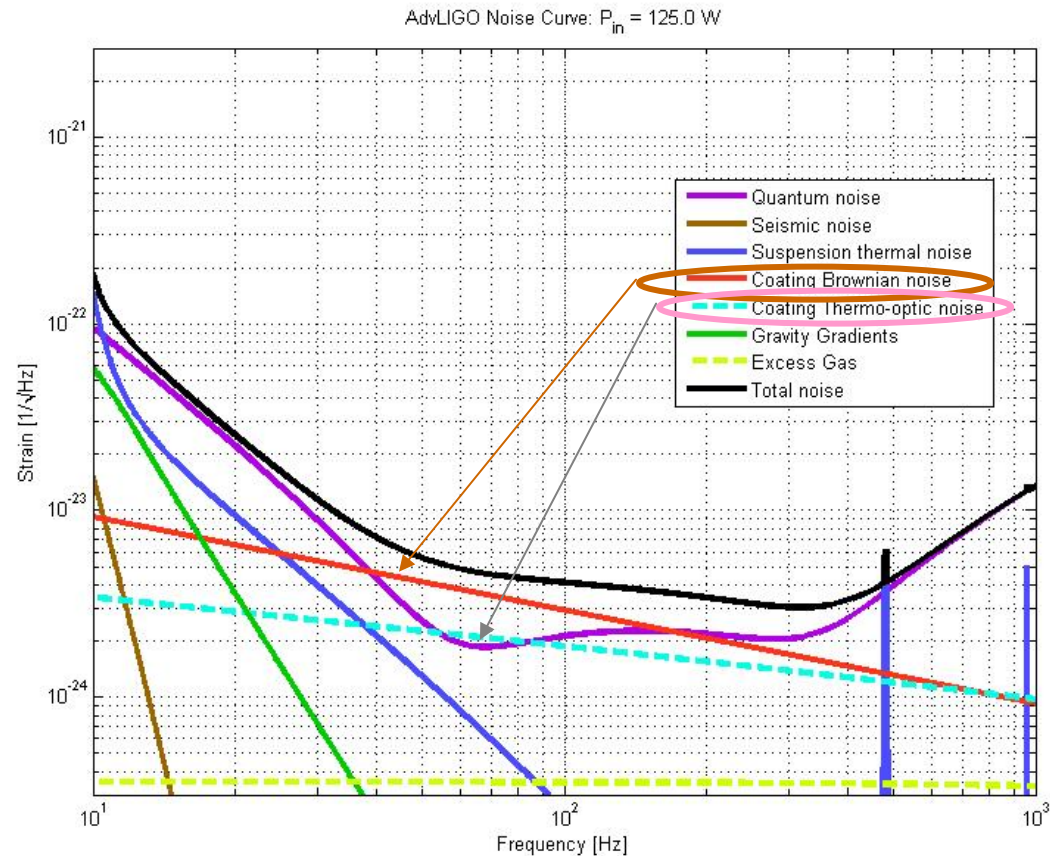
M-ary Coatings
Band-gap Engineered Coatings

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AdLIGO NOISE BUDGET (Quarter Wavelength Design)



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BINARY COATING OPTIMIZATION IN A NUTSHELL, I

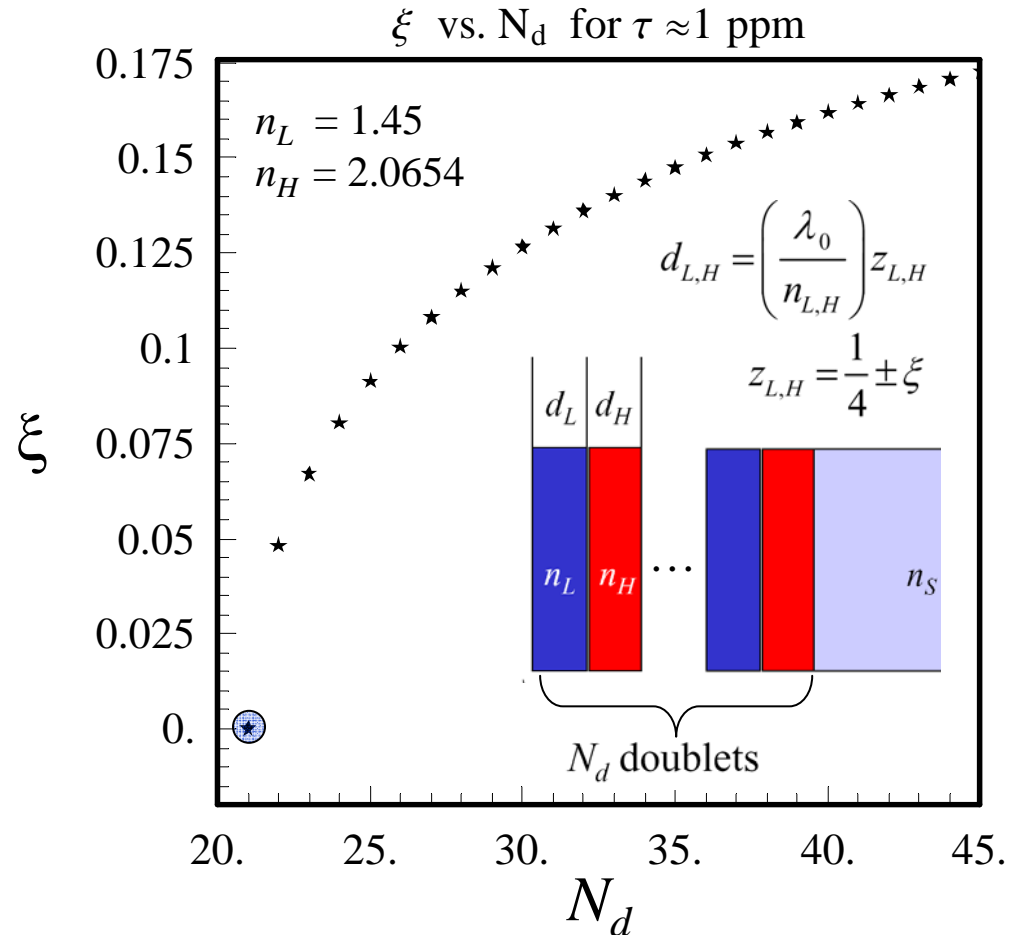


ISO-TRANSMITTIVE STACKED DOUBLET, “BRAGG” BINARY COATINGS

Reference QWL ($\xi=0$)
design closest to 1ppm
has

$$N_D = 21$$

$$\tau = 0.9727 \text{ ppm}$$



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BINARY COATING OPTIMIZATION IN A NUTSHELL, II



AMONG ISO – TRANSMITTIVE ALTERNATIVE
DESIGNS, PICK **OPTIMAL ONE**, IN TERMS OF
THERMAL NOISE

WARNING : *OPTIMAL SHOULD BE ALSO GOOD*,
IN TERMS, OF

- DESIGN ROBUSTNESS [TOLERANCE W.R.T. DEPOSITION
ERRORS & MATERIAL PARAMETERS.UNCERTAINTIES].
- OPTICAL LOSSES;
- REFLECTANCE ON OTHER USABLE WAVELENGTHS
- ANGULAR ACCEPTANCE, etc.

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Prescribe transmittance : $\tau_P = 1 - |\Gamma_p|^2$

Find smallest N_d : $\tau_{QWL}(N_d) = \tau^* \leq \tau_P$

Do $N_d = N_d + 1$,

find (z_L, z_H) :

$$\begin{cases} z_L + z_H = 1/2 \\ \tau(z_L, z_H, N_d) = \tau^* \end{cases}$$

while $PSD(z_L, z_H, N_d) \leq PSD(z_L, z_H, N_d - 1)$

End layer
tweaking

Tweak (trim) topmost L-layer for maximum reflectance

Tweak (trim) bottom H-layer to bring back reflectance to design value (trims noise further)

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COATING NOISE MODEL



Three basic components should be included:

BROWNIAN - Key references:

[G. Harry, LIGO-T040029-00-R, 2004]

THERMOELASTIC - Key references:

[V.B. Braginsky and S.A. Vyatchanin, ArXiv:cond-mat/0302617, post-print of Phys. Lett. A312 (2003) 244; contains important corrections; M.M. Fejer et al., PRD-70 (2004) 082003]

THERMOELECTRIC - Key references:

[V.B. Braginsky, et al., Phys. Lett. A 271 (2000) 303; I.M. Pinto et al., LIGO T-070159-00-Z (2007), A. Gretarsson, LIGO G-08151-00-R (2008).

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[G. Harry, LIGO-T040029-00-R, 2004]

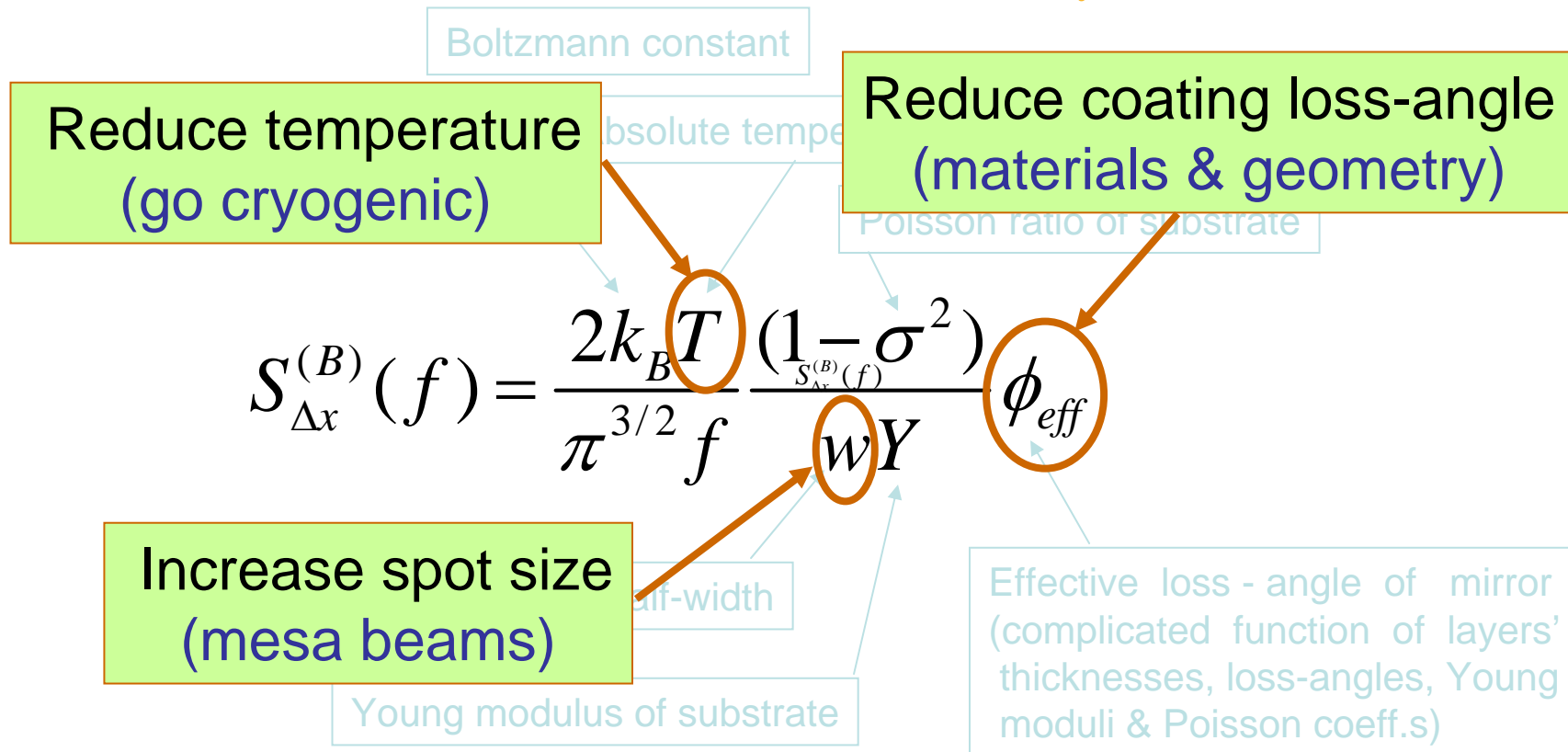
$$S_{\Delta x}^{(B)}(f) = \frac{2k_B T}{\pi^{3/2} f} \frac{(1 - \sigma^2)}{wY} \phi_{eff}$$

Diagram illustrating the components of the Brownian noise spectral density equation:

- Boltzmann constant** (k_B)
- Absolute temperature** (T)
- Poisson ratio of substrate** (σ)
- Beam half-width** (w)
- Young modulus of substrate** (Y)
- Effective loss - angle of mirror** (ϕ_{eff}) (complicated function of layers' thicknesses, loss-angles, Young moduli & Poisson coeff.s)

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[G. Harry, LIGO-T040029-00-R, 2004]



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COATING LOSS ANGLE : APPROXIMATE



...can be boiled down to simplest formula

$$(v_{L,H,S} \ll 1)$$

$$\phi_{\text{eff}} = (b_L d_L + b_H d_H)$$

$$b_{L,H} = \frac{\lambda_0}{\pi^{1/2} w} \frac{\phi_{L,H}}{n_{L,H}} \left(\frac{E_{L,H}}{E_s} + \frac{E_s}{E_{L,H}} \right)$$

$$\gamma = b_H / b_L \in [5, 10] \quad \text{plain Ta}_2\text{O}_5$$

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[G. Harry, LIGO-T040029-00-R, 2004]

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COATING NOISE: ii) THERMAL ELASTIC & THERMAL REFRACTIVE



Effective fluctuations of the test-mass (coated mirror) front - face position with respect to the mirror center of mass may occur as an effect of

- Thermal expansion of the coating layers (**thermoelastic effect**),

$$\Delta x^{(TE)} = \alpha_{eff} d_{tot} \Delta T$$

effective coating expansion coeff. coating thickness

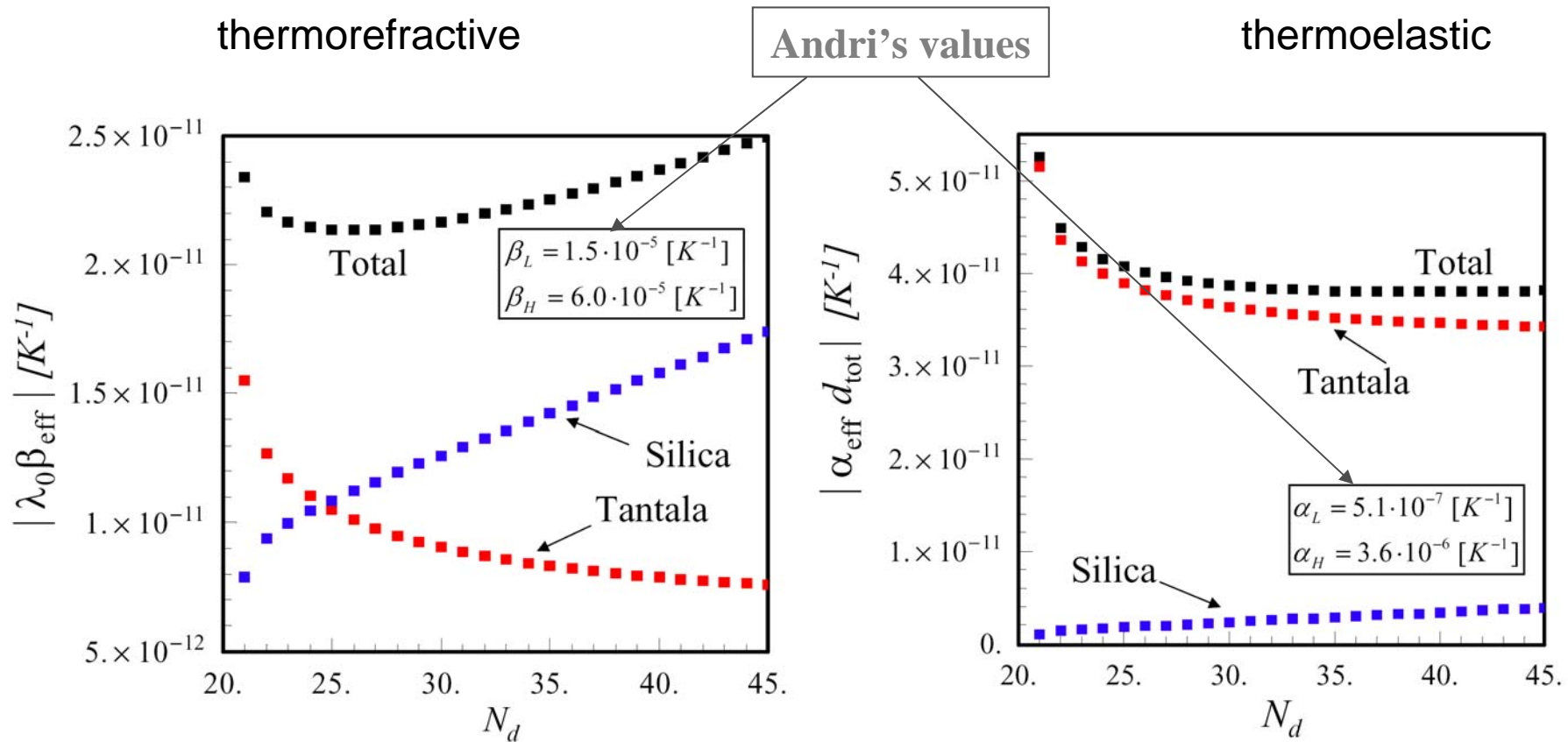
- Thermal variations of the refraction indexes $n_{H,L}$ of the coating materials (**thermorefractive effect**),

$$\Delta x^{(TR)} = \beta_{eff} \lambda_0 \Delta T$$

effective thermorefractive coefficient optical wavelength (vacuum)

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$$\begin{aligned}
 S_{\Delta x}(f) &= \mathcal{F}_{\tau \rightarrow f} \langle \Delta x(t) \Delta x(t + \tau) \rangle_t \stackrel{\text{quasi-static}}{=} \left(\frac{\Delta x}{\Delta T} \right)^2 \mathcal{F}_{\tau \rightarrow f} \langle \Delta T(t) \Delta T(t + \tau) \rangle_t \stackrel{\text{Wiener-Khinchin th.}}{=} \\
 &= \left(\frac{\Delta x}{\Delta T} \right)^2 S_{\Delta T}(f) \quad \text{PSD of } T \text{ - fluctuations in the coating} \\
 S_{\Delta T}(f) &= S_{\Delta T}^{(\Theta)}(f) + S_{\Delta T}^{(\Phi)}(f) \\
 &\quad \begin{array}{l} \text{Intrinsic fluctuations of} \\ \text{thermodynamic origin} \end{array} \quad \begin{array}{l} \text{add} \\ \text{in-coherently} \end{array} \quad \begin{array}{l} \text{Photo-thermal fluctuations} \\ \text{arising from laser shot noise} \\ \text{through optical absorption} \end{array}
 \end{aligned}$$



LIGO COATING TEMPERATURE FLUCTUATIONS



$$S_{\Delta T}^{(\Theta)}(f) = \frac{k_B T^2}{\pi^{3/2} r_0^2 \sqrt{f \kappa_s C_s \rho_s}}$$

[V. Braginsky, Phys. Lett A264 (1999) 1]

single photon energy
power abs. in coating

mass density
specific heat capacity
thermal conductivity } of substrate

$$S_{\Delta T}^{(\Phi)}(f) = \frac{P_{abs} E_\lambda}{4\pi^3 r_0^4 \kappa_s \rho_s C_s f}$$

[S. Rao, PhD Thesis, Caltech, 2003, etd-05092003-153759]

$$E_\lambda \cong 1.867 \cdot 10^{-19} J @ \lambda = 1064 nm$$

$$P_{abs} = 0.4 W \text{ for Adv LIGO}$$

(a different formula for $S_{\Delta T}^{(\Phi)}$
applies for sapphire substrates)

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Total Coating Noise PSD

$$S_{\Delta x}^{(tot)}(f) = S_{\Delta x}^{(B)}(f) + \left(\frac{\Delta x^{(TE)}}{\Delta T} + \frac{\Delta x^{(TR)}}{\Delta T} \right)^2 S_{\Delta T}(f)$$

Brownian-structural;

Thermally - driven elastic and refractive fluctuations.
may likely add incoherently or : indeed, the temperature
 in the coating does *not* vary



- on the space-scale (thickness) of the coating,
- on the time scales whereby the field in the coating builds up.



SD COATINGS: HOW DID WE GET HERE ? GENETIC SELECTION !



Educated ignorance attitude

- *no a-priori assumption on structure;*
- *easy inclusion of heterogeneous design constraints;*

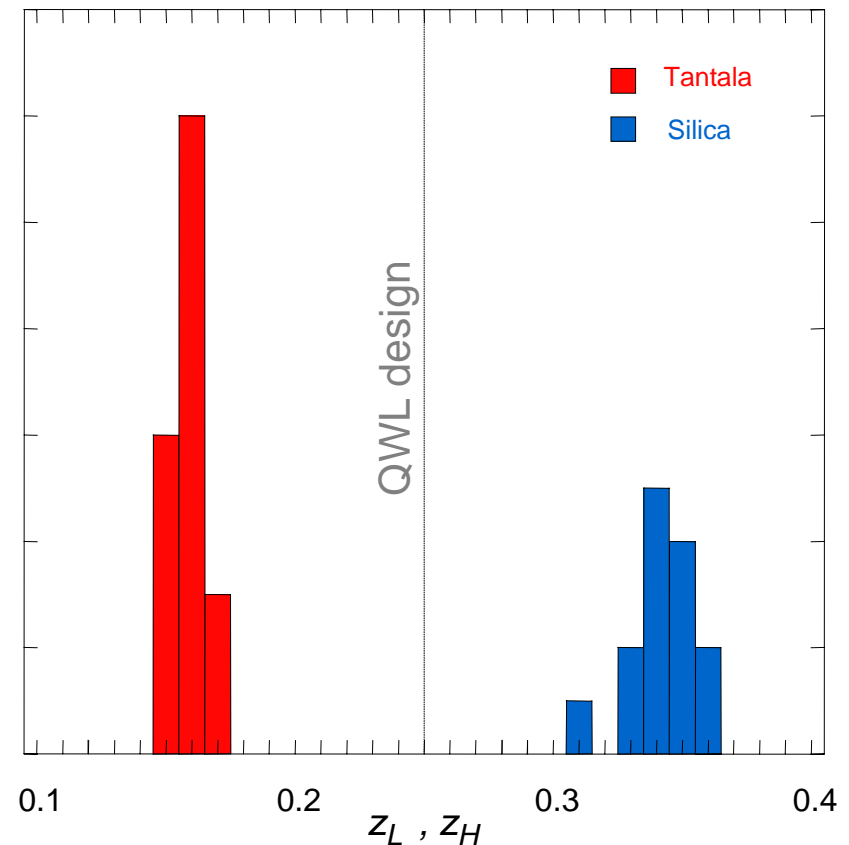
GA - engineered minimum noise *binary* coatings show trend toward non - QWL quasi - Bragg ($z_L + z_H = 0.5$) stacked-doublet (SD) configurations;

Deviations from trend are confined to fewest end - layers (first, last);

Suggests sequential design recipe :

- a) Design minimum-noise SD;
- b) Tweak terminal layers;

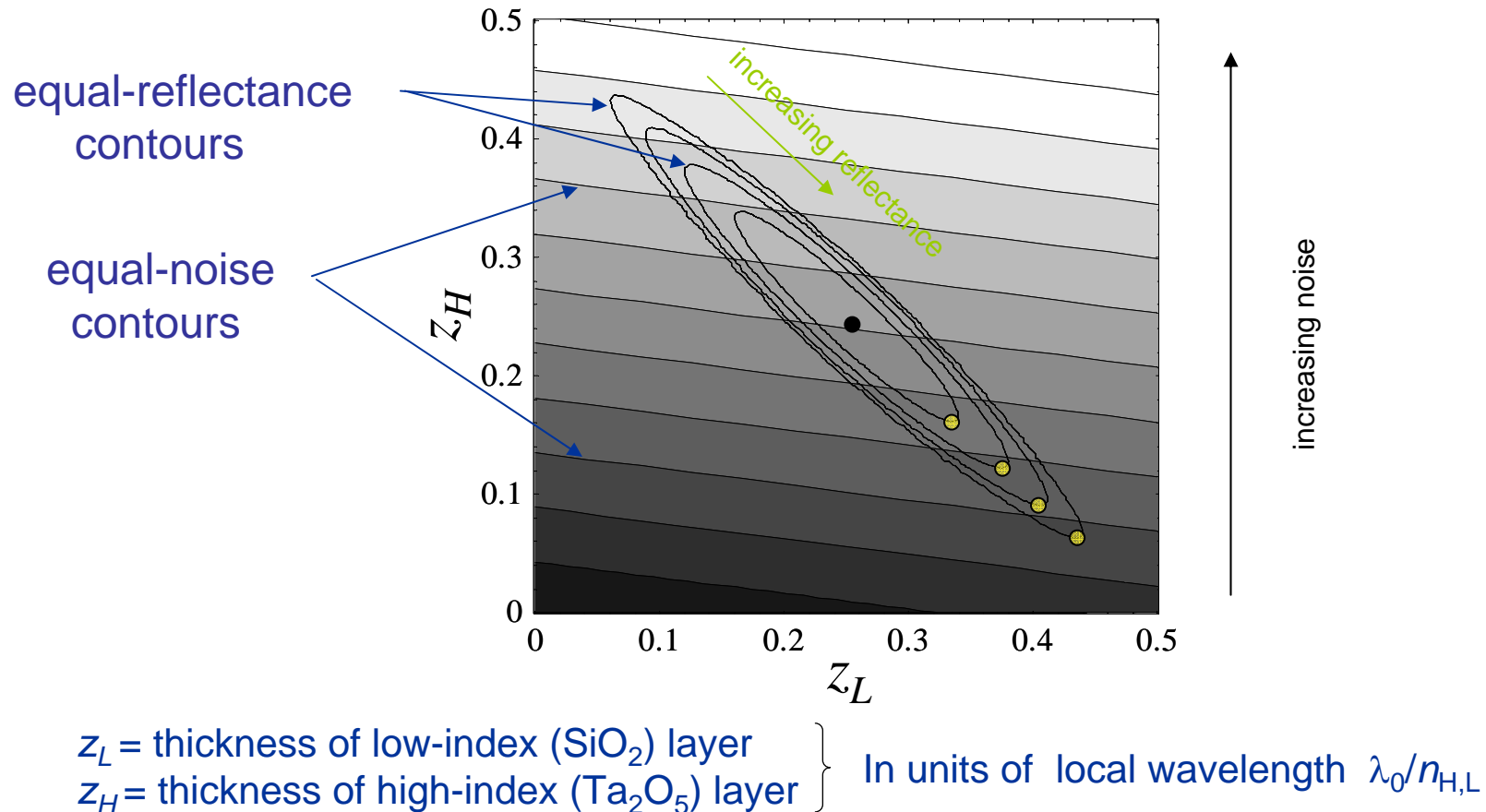
GA-optimized 20ppm transmittance prototype.
 z_L and z_H histograms after 10^4 generations.



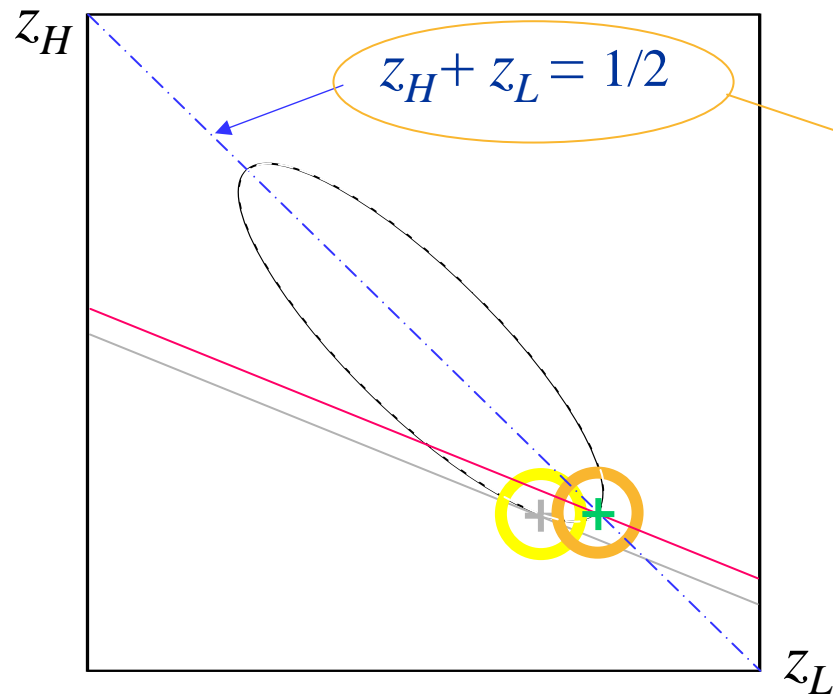
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BINARY COATINGS MINIMUM - NOISE BRAGG SD SYNTHESIS



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Bragg condition

Iso-reflectance contour lines
very thin: difference between
optimal and quasi – optimal
synthesis **negligible** (absorbed
by material uncertainties)

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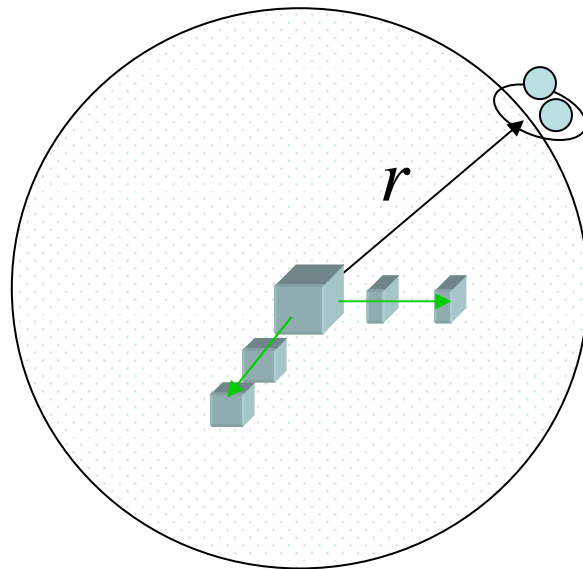
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$$h \propto r^{-1}$$

$$h_{min} \propto PSD_{floor}^{1/2}$$

$$r_{max} \propto PSD_{floor}^{-1/2}$$

$$\left. \begin{array}{l} \text{visibility volume,} \\ \text{event rate} \\ \text{(isotropic source distrib.)} \end{array} \right\} \propto PSD_{floor}^{-3/2}$$

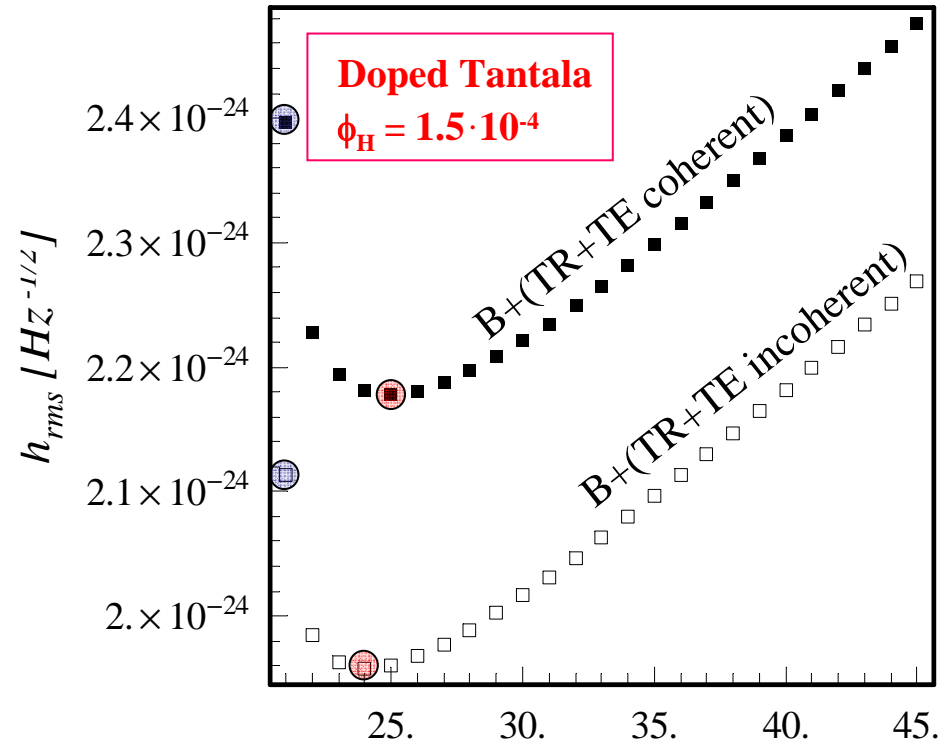
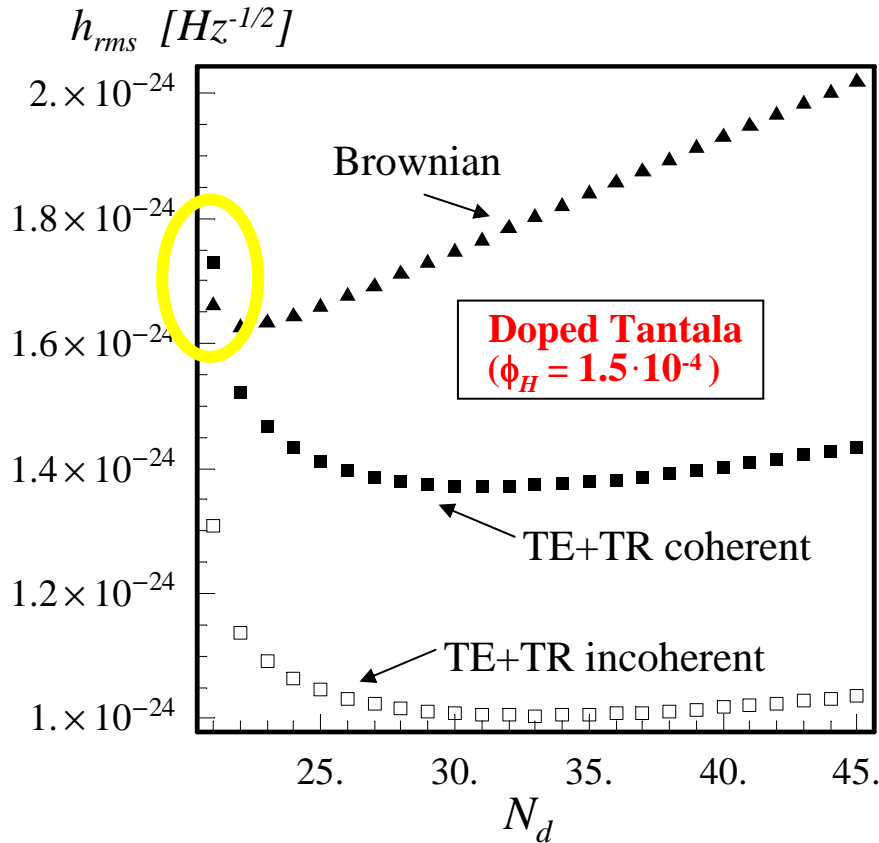
➔ A 13% reduction in PSD_{floor} boosts the event rate by 23%, etc.



1ppm OPTIMIZED COATING : DOPED TANTALA



$$\tau = 0.9727\text{ppm}, f = 100\text{Hz}$$



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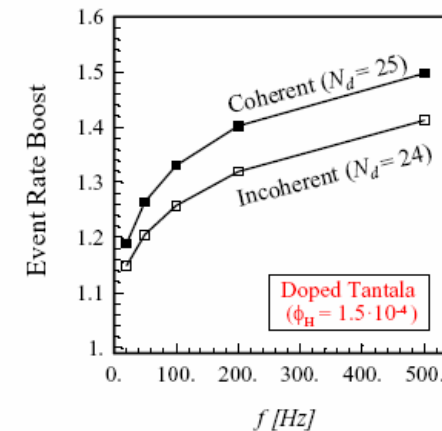
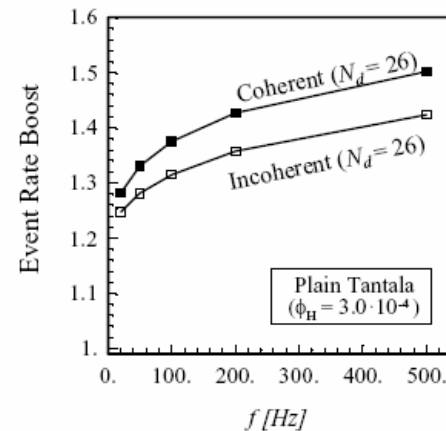


EVENT RATE BOOST (1ppm designs)



(Total noise budget included)

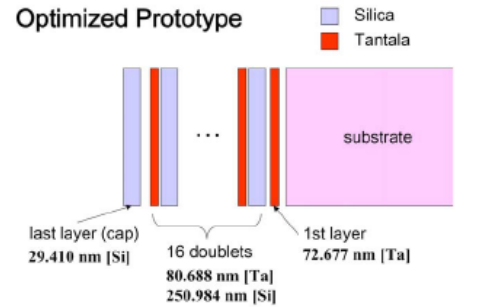
$\tau = 0.9727 \text{ ppm}$	ER boost @ 100Hz
Plain Tantala, QWL	1
Plain Tantala, OPT	1.38
Doped Tantala, QWL	1.54
Doped Tantala, OPT	2.05



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Coating	Substrate
$n_{Ta} = 2.035 - i 1.8 \cdot 10^{-7}$	$n_s = 1.46543 - i 4 \cdot 10^{-8}$
$n_{Si} = 1.46543 - i 4 \cdot 10^{-8}$	$Y_s = 727 \cdot 10^6$
$Y_{Ta} = 140.0 \cdot 10^9$	$\sigma_s = 0.167$
$Y_{Si} = 727.0 \cdot 10^9$	$\phi_s = 0.0$
$\sigma_{Ta} = 0.23$	
$\sigma_{Si} = 0.167$	
$\phi_{Ta} = 3.0 \cdot 10^{-4}$	
$\phi_{Si} = 5.0 \cdot 10^{-5}$	



COAT - 300ppm optimized mirror, final design, november 2006

$$\tau_p = 278.421 \text{ ppm (OPT \& QWL)}$$

$$\frac{PSD_{OPT}}{PSD_{QWL}} = 0.843$$

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bolle

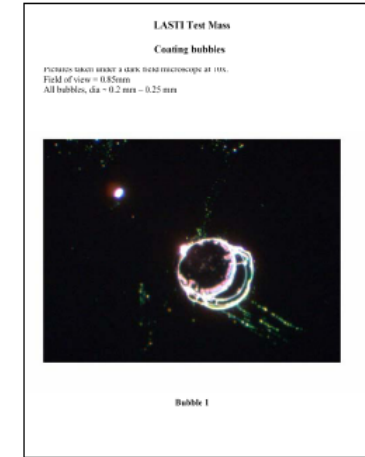
Friday, August 24, 2007 11:45:25 PM

From: desalvo@ligo.caltech.edu
To: pinto@sa.infn.it
Coating bubbles.pdf (2673.2KB)

Innocenzo, as you can see from the attached file, the LMA bubble problem is there also on a standard coating ($\lambda/4$, this time the problem appears in the first LIGO mirror (34 cm diameter, 40 Kg, and a lot expensive)). Also, you were not alone... We found 8 holes across the mirror that LMA recognizes as bubble remnants. They formed during annealing. They can be eliminated using a different annealing profile to allow for the Argon trapped in the coating during deposition to diffuse out without ripping away the coating itself. The "successful" optimized mirrors for TNI were coated just after this run, and I asked to re-inspect them for bubbles and holes. This should not generate any further delay.

The gun has moved to its new location. I'm awaiting for electrical connection of the big fans, and set of clean air filters.

Ciao,



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TNI Mirrors - July 2007

Reference	Substrate Reference	Average Scattering	Average Absorption	Average Transmission T ¹ incidence	Remarks	
C07052/2	IV 231	20 ppm Ø 70 mm	6 ppm Ø 50 mm	0.37 ppm	304 ppm	-
C07052/3	IV 230	21 ppm Ø 70 mm	9 ppm Ø 50 mm	0.40 ppm	274 ppm	-
C07053/2	IV 533	10 ppm Ø 70 mm	8 ppm Ø 50 mm	0.35 ppm	286 ppm	HR Coating on back side
C07053/3	IV 532	12 ppm Ø 70 mm	10 ppm Ø 50 mm	0.38 ppm	286 ppm	HR Coating on back side

Remarks: the concave side is marked by an arrow on the edge of the substrate and also marked on this box. The transportation boxes must be opened in clean room conditions.

Laboratoire des Matériaux Avancés
Université Clermont - CNRS - IN2P3
23, Bd. Naut. Bldg.
63022 - Vichy-Meuse Cedex, FRANCE

July 2007

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optimized coating

Wednesday, December 05, 2007 10:09:02 AM

From: c.comtet@lma.in2p3.fr

To: pinto@sa.infn.it; morgado@lma.in2p3.fr; desalvo@ligo.caltech.edu; gharry@ligo.mit.edu; black_e@ligo.caltech.edu

Dear Innocenzo

Recently we made 0 measurements on non doped tantalum and silica coating annealed at 500°C. We found similar losses for than a higher standard annealing (3.02 · 10⁻⁴ for tantalum losses and 0.5 · 10⁻⁴ for silica losses). The geometry of the cantilever used for the measurement is 110micron thick by 45mm long by 5mm large. The thickness of the coating is 500 nm

With regards

Christophe

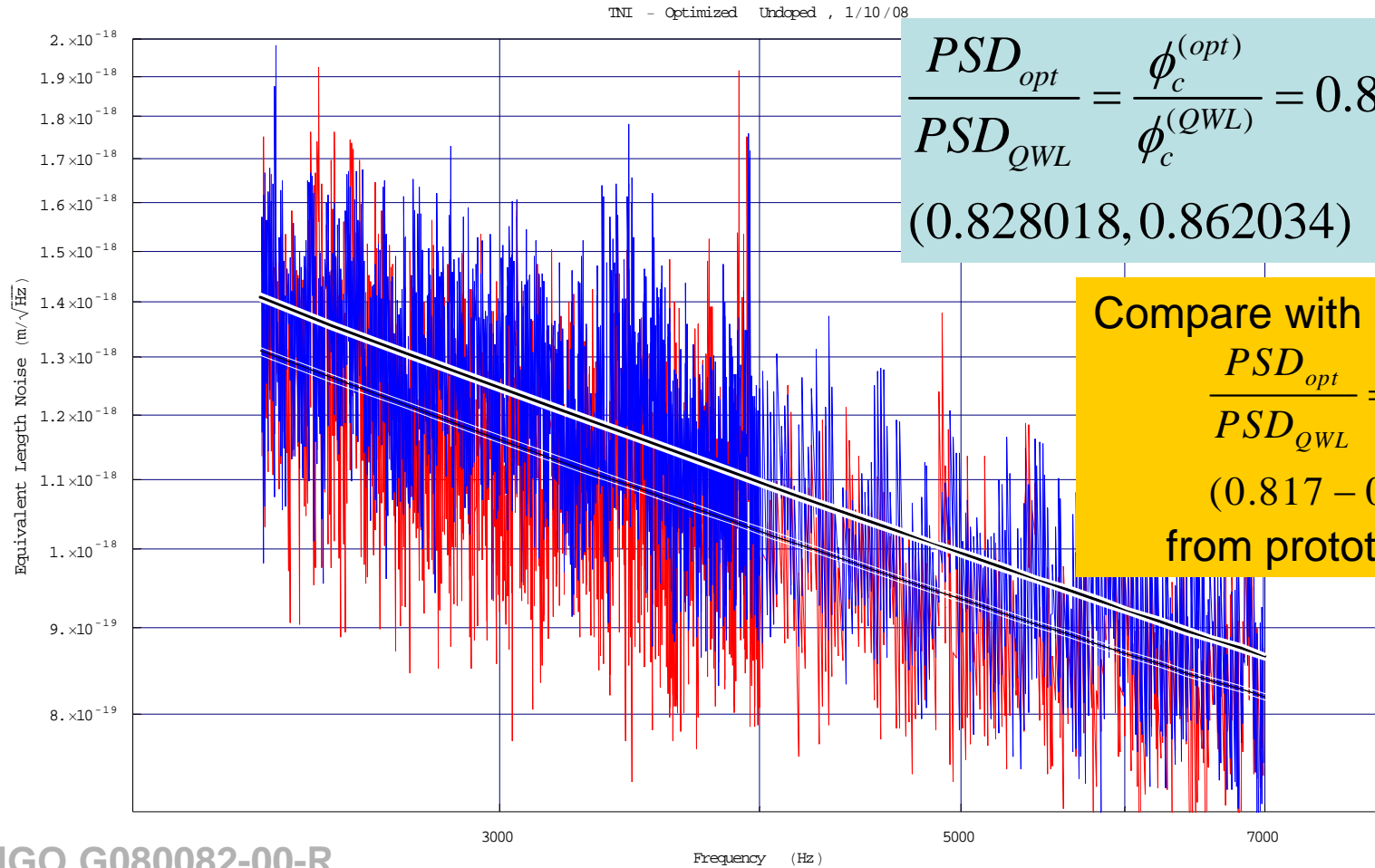
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TNI PROTOTYPES TESTED



[see Akira's talk, this meeting]



$$\frac{PSD_{opt}}{PSD_{QWL}} = \frac{\phi_c^{(opt)}}{\phi_c^{(QWL)}} = 0.844869$$

(0.828018, 0.862034)

Compare with

$$\frac{PSD_{opt}}{PSD_{QWL}} = 0.843$$

(0.817 – 0.876)
from prototype model

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(TNI prototype)

	N = 16		N = 17		N = 18	
γ	Plain { z_L, z_H } {0.330169, 0.169831}	Tweaked { z_L, z_H } {0.330169, 0.169831} { z_I, z_N } {0.0338288, 0.157079}	Plain { z_L, z_H } {0.345676, 0.154324}	Tweaked { z_L, z_H } {0.345676, 0.154324} { z_I, z_N } {0.0405062, 0.139003}	Plain { z_L, z_H } {0.357797, 0.142203}	Tweaked { z_L, z_H } {0.357797, 0.142203} { z_I, z_N } {0.0457719, 0.124762}
10	0.896	0.871	0.901 (+0.005)	0.876 (+0.005)	0.912 (+0.016)	0.886 (+0.015)
7	0.866 (+0.003)	0.847 (+0.004)	0.863	0.843	0.867 (+0.004)	0.847 (0.004)
5	0.842 (+0.012)	0.827 (+0.013)	0.833 (+0.003)	0.817 (+0.003)	0.830	0.814

Table of PSD values relative to QWL design with HWL cap (N=14), for various values of γ in

$$S_x^{(B)} = C[z_L + \gamma z_H] \quad (z_L, z_H = \text{layer thicknesses in units of local wavelength})$$

N = number of high/low index layers; optimum (minimum noise) syntheses highlighted in yellow. Numbers in brackets are $\{z_L, z_H\}$ (plain design; first line in tweaked design) and $\{z_I, z_N\}$ (second line in tweaked design).

→ The N=17 design yields the minimum degradation (in brackets) compared to optimum design, if γ is allowed to change throughout the interval [5,10].

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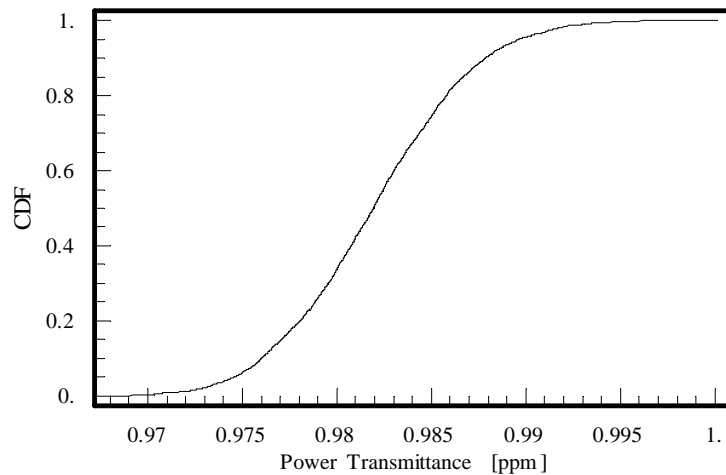


ROBUSTNESS - II

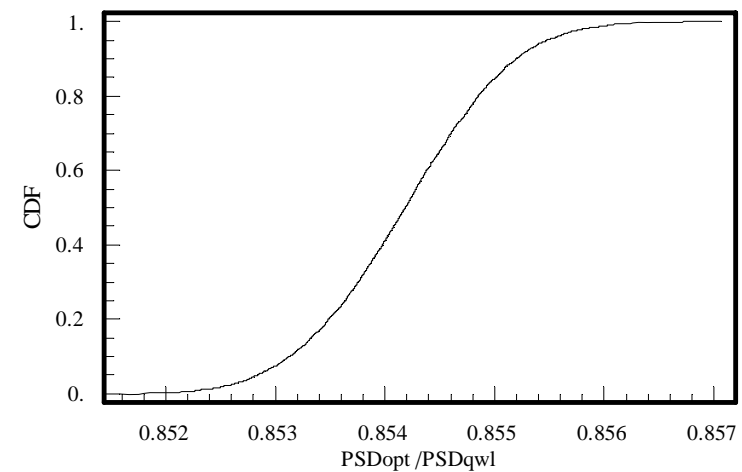


Random uniform layer-thickness errors, $|\delta\ell| \leq 1nm$, 1ppm prototype.

Power transmittance



Noise PSD ratio @ 100Hz



(10^4 realizations)

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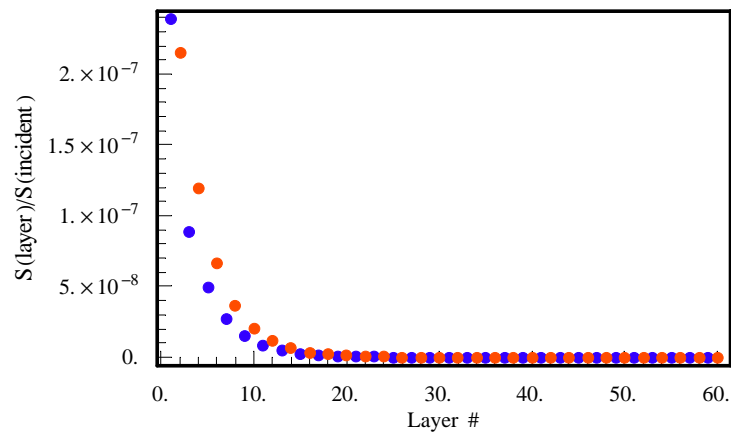


OPTICAL LOSSES

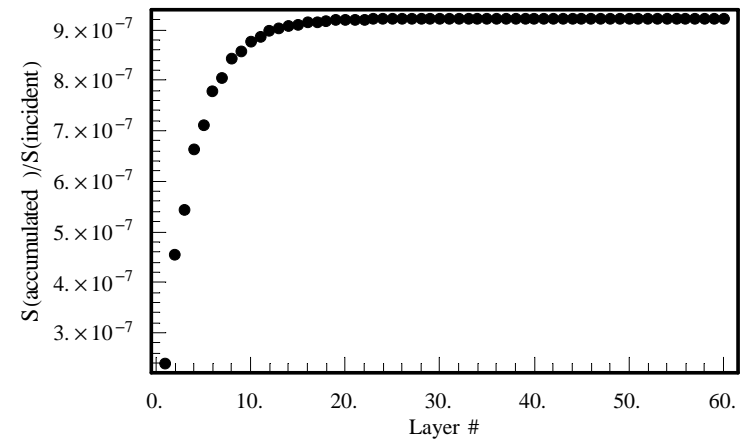


(1ppm prototype).

Optical Power Loss

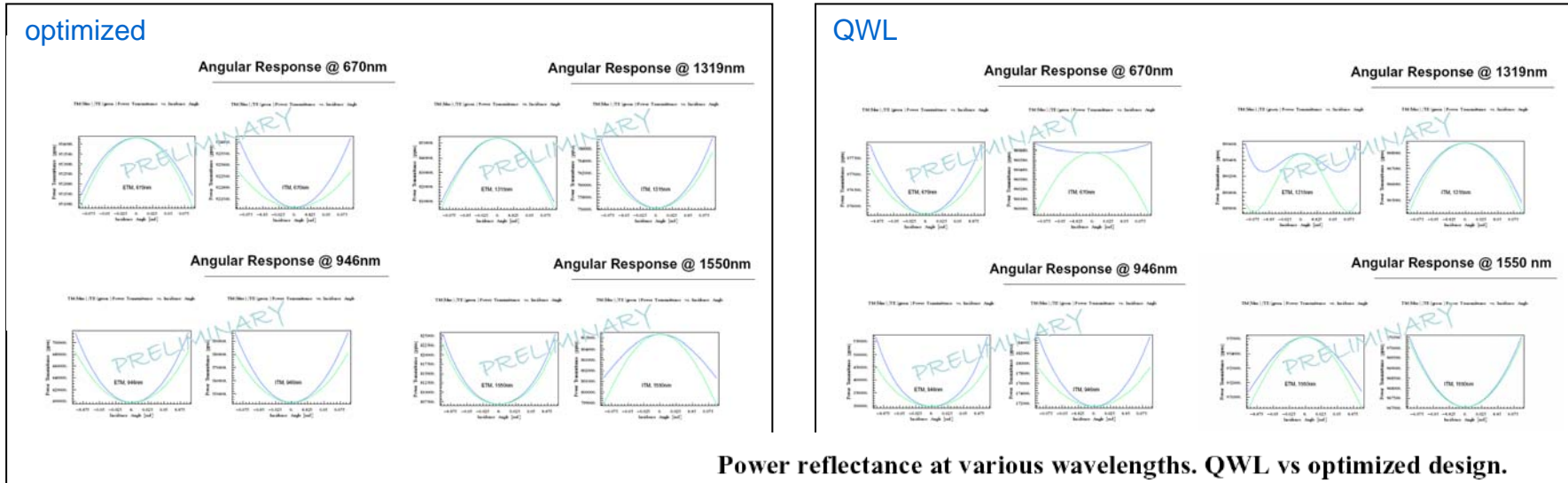


Accumulated Optical Power Loss



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Power reflectance at various wavelengths. QWL vs optimized design.

- No need to constrain noise - optimization in order to get useful power reflectance at other (laser) wavelength(s) .
- Optimized design no worse than QWL in terms of power - reflectance at other (laser) wavelengths (worse but still OK @ 946nm).
- No critical dependence of result on material parameters.

λ [nm]		670		946		1319		1550	
		QWL	OPT	QWL	OPT	QWL	OPT	QWL	OPT
ETM	$R_{TE, TM}(0 \text{ deg})$	0.024	0.046	0.56	0.40	0.11	0.15	0.045	0.19
	$R_{TE}(\pm 5 \text{ deg})$	0.023	0.049	0.60	0.31	0.11	0.19	0.050	0.18
	$R_{TM}(\pm 5 \text{ deg})$	0.022	0.048	0.61	0.28	0.10	0.19	0.048	0.17
ITM	$R_{TE, TM}(0 \text{ deg})$	0.039	0.079	0.83	0.46	0.031	0.24	0.033	0.19
	$R_{TE}(\pm 5 \text{ deg})$	0.040	0.077	0.82	0.42	0.036	0.23	0.030	0.20
	$R_{TM}(\pm 5 \text{ deg})$	0.039	0.076	0.81	0.40	0.035	0.23	0.029	0.20

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Nominal power transmittances @ 1064nm, normal incidence:
6.324 ppm (ETM), 14172 ppm (ITM)



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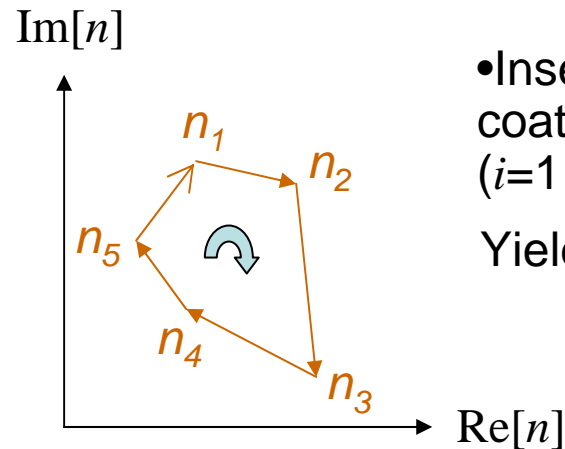
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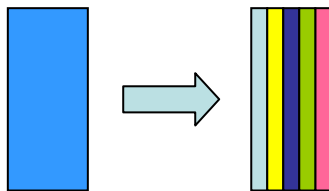


- Insertion of a (different) dielectric layer in a stacked doublet coating boosts reflectance provided $\text{Im}[(n_{i+1} - n_i)/(n_i - n_{i-1})] < 0$ ($i=1$ is top layer) [J. I. Larruquert, Opt. Comm. 206 (2002) 259]

Yields simple material selection rule (“turn clockwise”).

May result into **reduced coating noise**
Material downselection **TBD**.

- Use materials with *different* properties (contrast, losses) in topmost / bottom layers to **reduce coating thickness without increasing optical loss** [P.G. Verly, Appl.Opt. 37 (1998) 7327]



- Stack of *sub-wavelength* layers “equivalent” to *homogeneous* medium [E. Tuncer, Physical Review B 71 (2005) 012101].
 E, n, ϕ etc. obtained from appropriate *mixture* formulas.
Way to engineering *new* materials ?

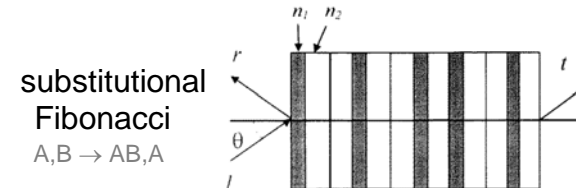
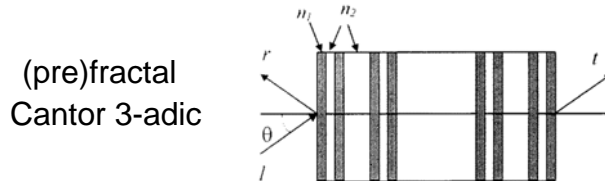
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BAND-GAP ENGINEERED COATINGS



“Photonic Band Gap Engineering” [Yablonovitch, Opt. Lett.23 (1998) 1648]

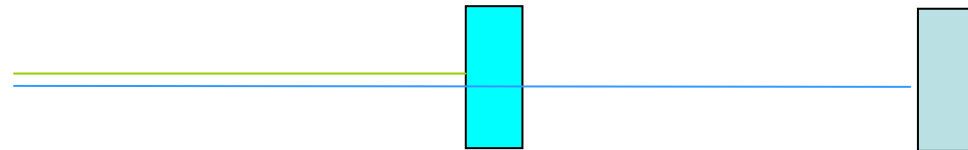


Nearly *omni-directional (and wideband) reflection* obtainable [D. Lusk and F. Placido, Thin Solid Films, 392 (2005) 226];

⇒ Expected to *mitigate misalignment instability* in (otherwise stable) spherical / confocal mesa - beam (or hyperboloidal beam) cavities w. coated mirrors

Perfect transmission bands also obtainable [R.W. Peng et al., Appl. Phys. Lett. 80 (2002) 3063]

⇒ May permit building *two (almost independent) interferometers in a single beam pipe*...



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- Coating thickness optimization *almost mandatory* to minimize coating noise when using doped Tantalum, yielding a *substantial boost* ($> 30\%$ @ 100Hz) in the expected event rate, as compared to QWL design.
- Optimal design *almost the same for both* the incoherent and the coherent thermo-optic noise formula.
- Optimal design is *robust* against thickness deposition errors, and can be made *judiciously tolerant* w.r.t. uncertainties in the relevant material parameters.
- Among all proposed coating noise reduction techniques (new materials, cryogenic mirrors, flat-top beams) thickness optimization is the *cheapest reliable option*



ONGOING & FUTURE WORK



-
- Design of optimized doped-Tantala mirrors for TNI.
 - Sensitivity study w.r.t. to uncertainties in material params.
 - Identification of most tolerant quasi-optimal design
Funding requested to INFN.
 - M-ary and sub-wavelength coatings optimization study started.
 - Band-gap engineered optimised coatings study restarted.

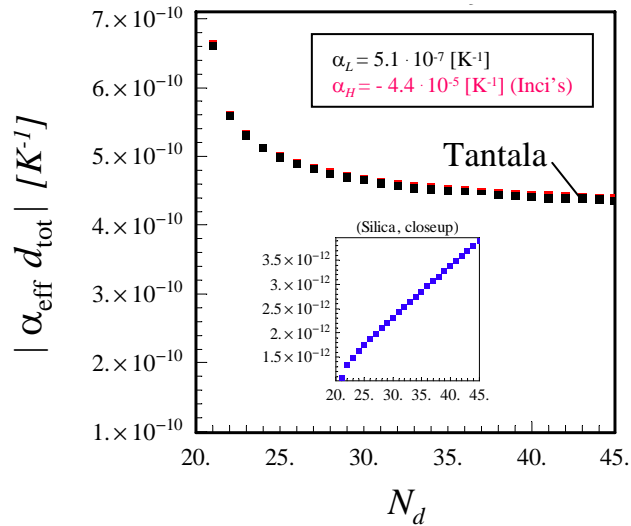
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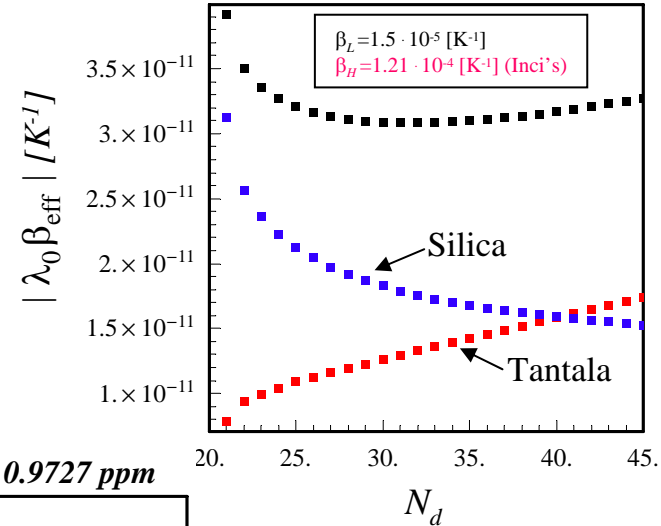


INCI'S NUMBERS

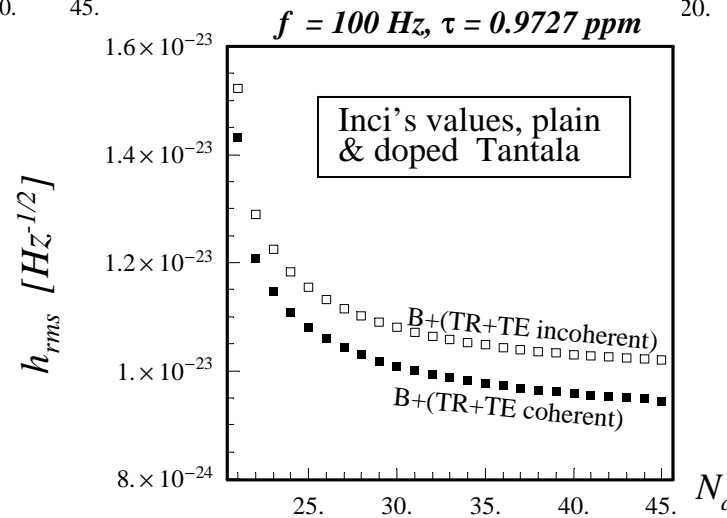
[Naci-Inci, J. Phys. D37 (2004) 3151.]



(TE Coefficient)



(TR Coefficient)



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