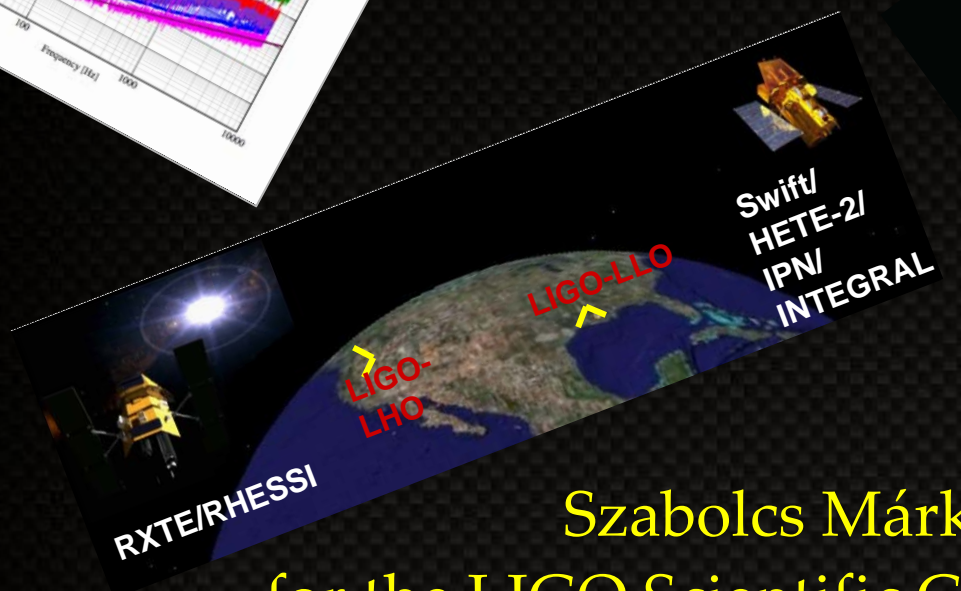
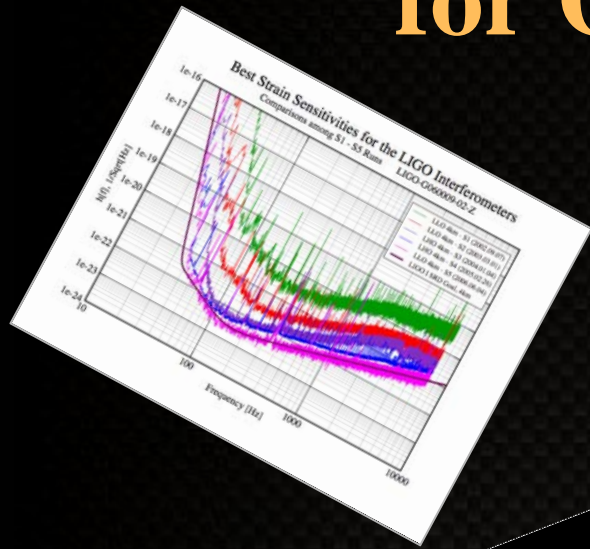


# Astrophysically Triggered Searches for Gravitational Waves



Szabolcs Márka  
for the LIGO Scientific Collaboration

AAS 211, Austin, TX, 2008

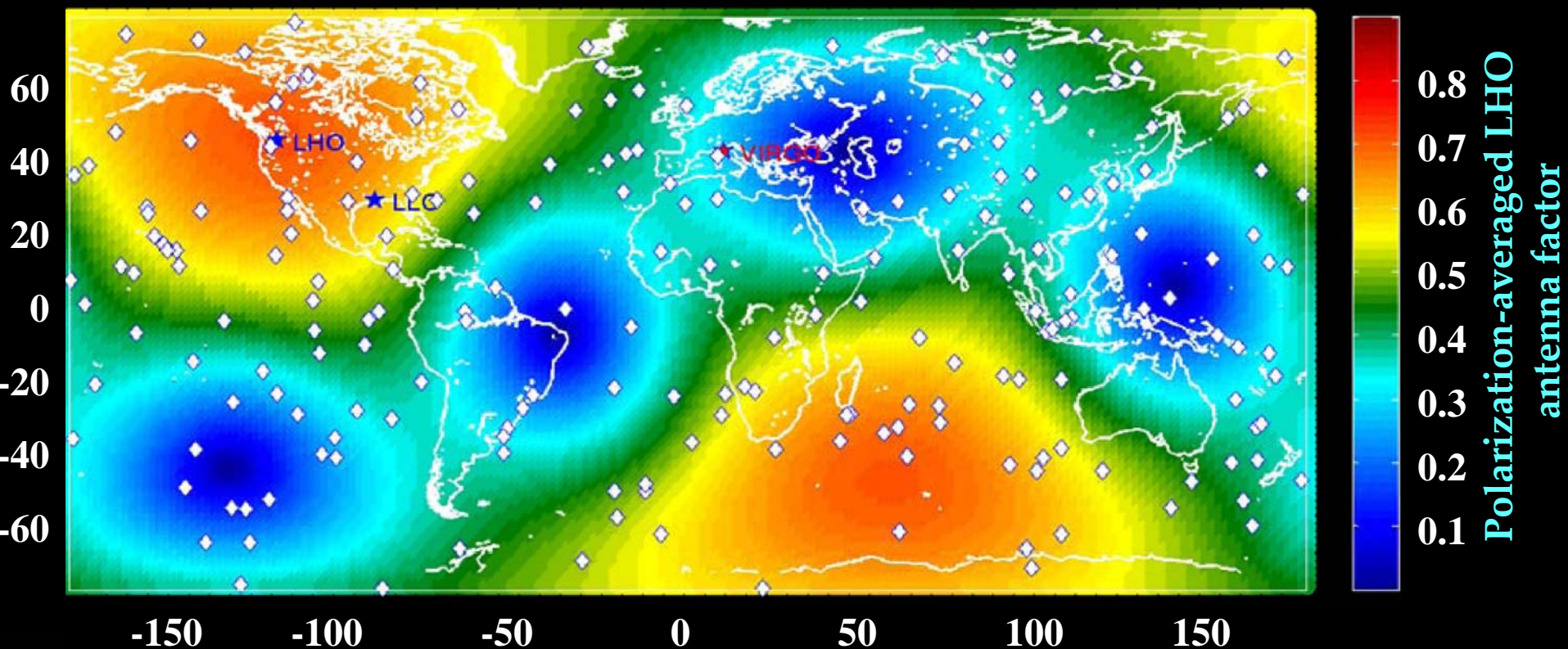
- Gamma-ray transients (GRBs, SGRs)
- Optical transients
- Neutrino events
- ...

- Correlation in time
- Correlation in direction
- Information on the source properties
- ...



- ✓ Confident detection of GWs (eventually).
- ✓ Better background rejection  $\Rightarrow$  Higher sensitivity to GW signals.
- ✓ More information about the source/engine.
- ⚠ ✓ Even upper limits can have interesting implications. ⚠

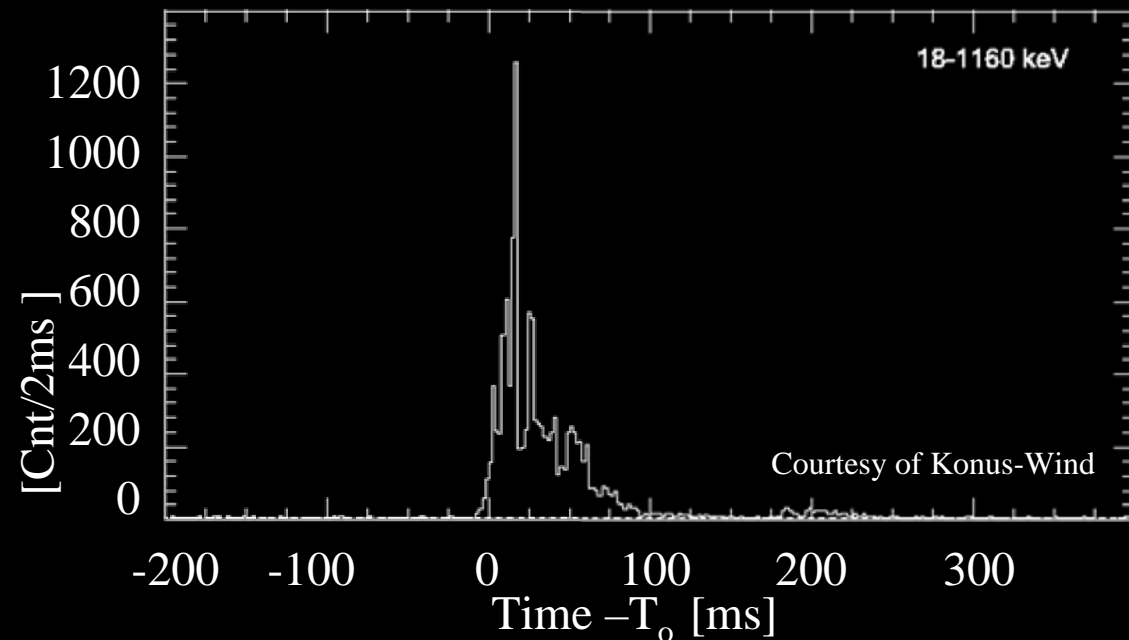




– **GRB triggers** (mostly from Swift, IPN, INTEGRAL, HETE-2)

- ~70% with double-IFO coincidence LIGO data ⚠
- ~40% with triple-IFO coincidence LIGO data
- ~25% with measured redshift
- ~15% short-duration GRBs

Detected by Konus-Wind, INTEGRAL, Swift, MESSENGER



Described as an

*“intense short hard GRB”* (GCN 6088)

**Duration ~0.15 seconds,**  
followed by a weaker, softer pulse  
with duration ~0.08 seconds

**R.A. = 11.089 deg,**  
**Dec = 42.308 deg**

**Antenna responses of LIGO Hanford:**

$$F_{RMS} = \sqrt{F_+^2 + F_\times^2} / \sqrt{2} = 0.304$$

$$h(t) = F_+(\theta, \phi, \psi)h_+(t) + F_\times(\theta, \phi, \psi)h_\times(t)$$



# GRB 070201 – Sky Location

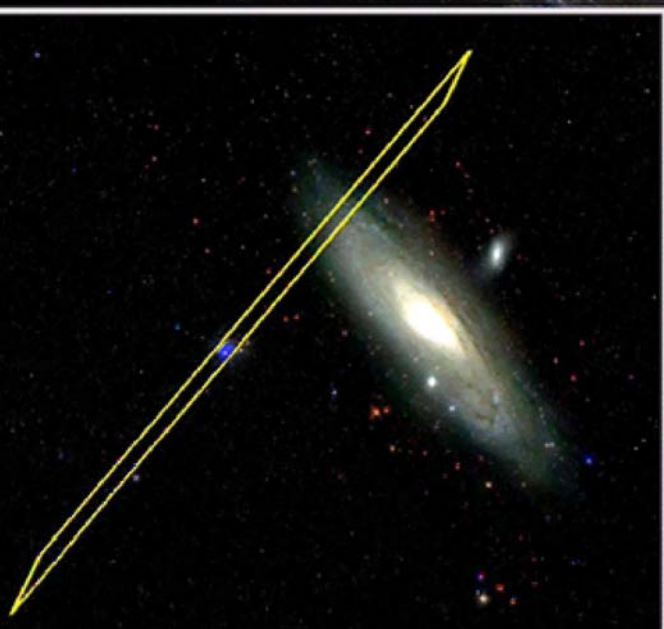
**R.A. = 11.089 deg,  
Dec = 42.308 deg**

**$D_{M31} \approx 770$  kpc**

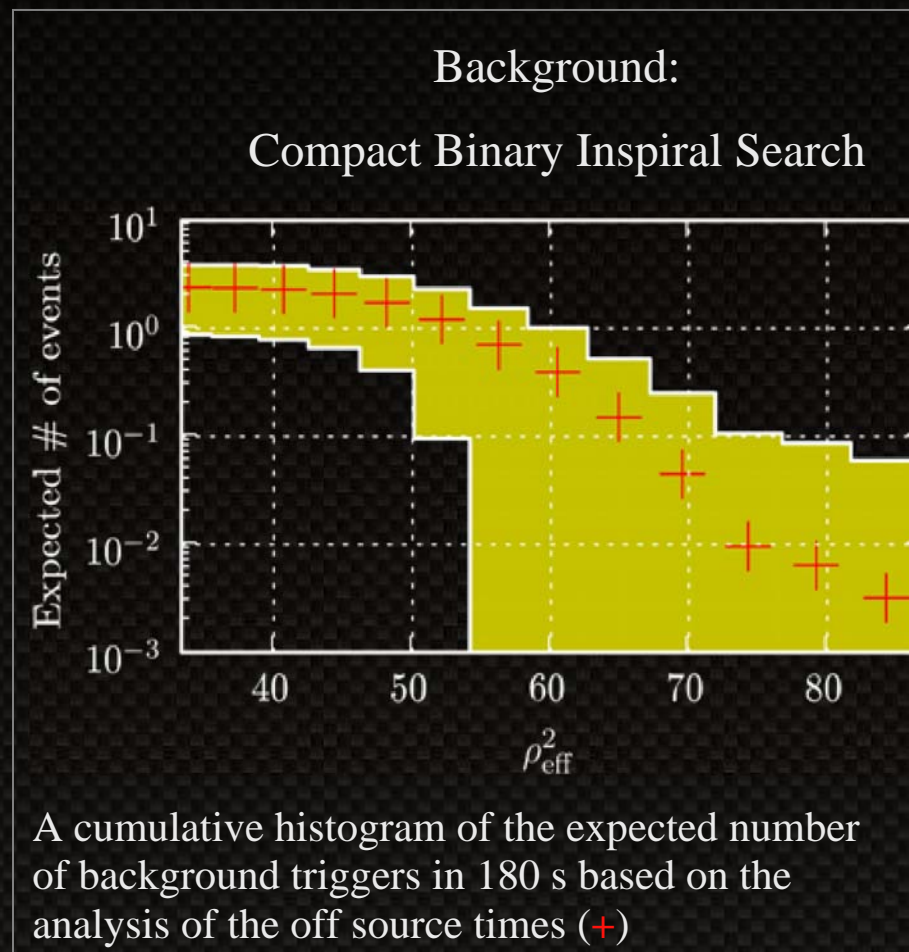
**Possible progenitors for short GRBs:**

- **NS/NS or NS/BH mergers**  
Emits strong gravitational waves
- **SGR**  
May emit GW but weaker

**$E_{\text{iso}} \sim 10^{45}$  ergs**  
**if at M31 distance**  
(more similar to SGR than GRB energy)



- **In case of a detection:**
  - Confirmation of a progenitor (e.g. coalescing binary system)
  - GW observation could determine the distance to the GRB
- **In case of non-detection:**
  - Exclude progenitor/model in mass-distance region
  - Assumed M31 distance to hypothetical GRB  $\Rightarrow$   
exclude binary progenitor
  - Bound the GW energy emitted by a source in M31
  - ...

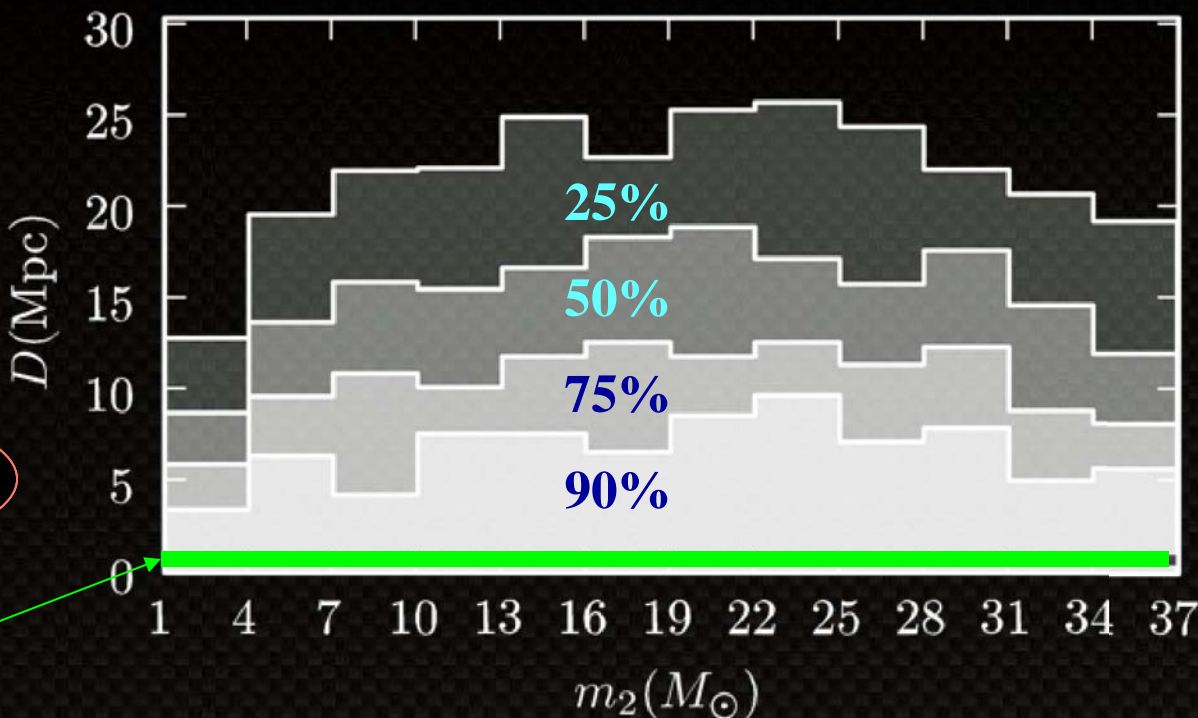




Exercise matched filtering techniques for inspiral waveform search

*No plausible gravitational waves identified*

$D_{M31}$



**Exclude** compact binary progenitor with masses

$1 M_{\odot} < m_1 < 3 M_{\odot}$  and  $1 M_{\odot} < m_2 < 40 M_{\odot}$  with  $D < 3.5$  Mpc away at 90% CL

**Exclude** any compact binary progenitor in our simulation space

at the distance of M31 at  $> 99\%$  confidence level

Excess Power type search

No plausible gravitational waves identified

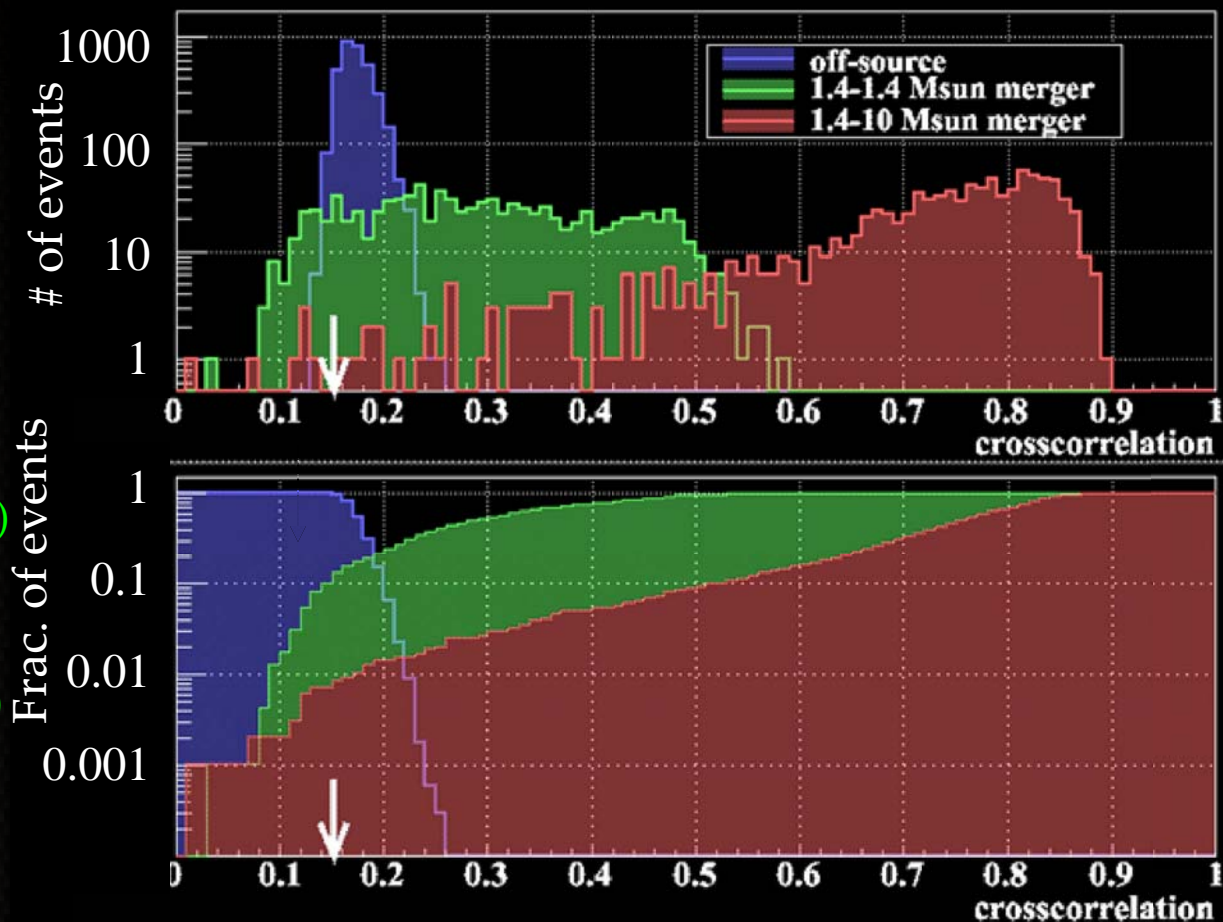
Injected simulated waveforms

- NS-NS inspirals ( $1.4-1.4 M_{\odot}$ )
- and
- NS-BH inspirals ( $1.4-10 M_{\odot}$ )

at nominal M31 distance

Efficiency  $> 0.878$ ,  $1.4 - 1.4 M_{\odot}$

Efficiency  $> 0.989$ ,  $1.4 - 10 M_{\odot}$



These results give an independent way to reject hypothesis of a compact binary progenitor in M31



# Soft Gamma-ray Repeater in M31 ?



SGR: highly magnetized neutron star;  
 can have giant flares (rare)  
 (arXiv:0712.1502)

STARQUAKE ON A MAGNETAR releases a vast amount of magnetic energy—equivalent to the seismic energy of a magnitude 21 earthquake—and unleashes a fireball of plasma. The fireball gets trapped by the magnetar's field.

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Scientific American, February 2003

- Giant flare from an SGR:
  - a hypothesized explanation for 070201 burst
- Energy release in gamma rays consistent with SGR model
  - Measured gamma-ray fluence =  $2 \times 10^{-5}$  ergs/cm<sup>2</sup> (Konus-Wind)

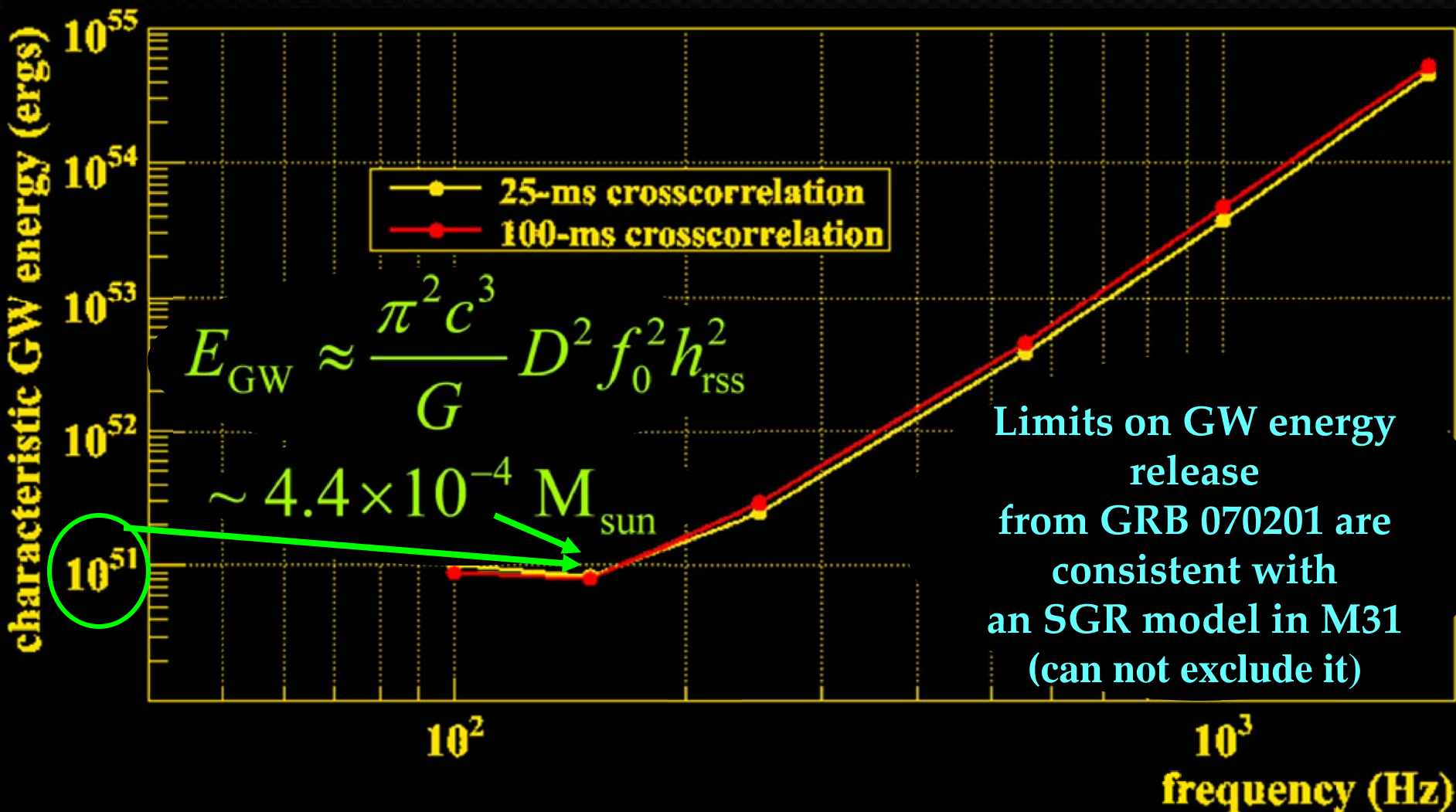
Corresponding gamma-ray energy, assuming isotropic emission, with source at  $D = 770$  kpc (M31):

$$E_{\gamma, iso} = \phi \times 4\pi D^2 \approx 10^{45} \text{ ergs}$$

- SGR models predict energy release in GW to be no more than  $\sim 10^{46}$  ergs

# Model independent burst search result

- Corresponding gamma-ray energy, assuming isotropic emission, with source at  $D = 770$  kpc (M31):
 
$$E_{\gamma, \text{iso}} = \phi \times 4\pi D^2 \approx 10^{45} \text{ ergs}$$
- SGR models predict energy release in GW to be no more than  $\sim 10^{46}$  ergs





*“Implications for the Origin of GRB 070201 from LIGO Observations”*  
(arXiv:0711.1163) paper accepted for publication in ApJ

- No plausible gravitational waves were identified
- Excluded compact binary progenitor in M31
- Corresponding limits on isotropic energy emission in GW do not exclude an SGR model in M31

Analysis is ongoing to search for gravitational waves associated with

- The sample of 213 GRB triggers contemporaneous with LIGO S5 run
- Other external triggers...

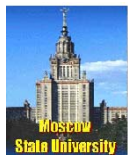
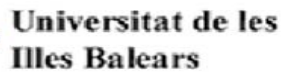
...and the future is bright...

# LIGO

# LIGO Scientific Collaboration

# LSC

- Australian Consortium for Interferometric Gravitational Astronomy
- The Univ. of Adelaide
- Andrews University
- The Australian National Univ.
- The University of Birmingham
- California Inst. of Technology
- Cardiff University
- Carleton College
- Charles Sturt Univ.
- Columbia University
- Embry Riddle Aeronautical Univ.
- Eötvös Loránd University
- University of Florida
- German/British Collaboration for the Detection of Gravitational Waves
- University of Glasgow
- Goddard Space Flight Center
- Leibniz Universität Hannover
- Hobart & William Smith Colleges
- Inst. of Applied Physics of the Russian Academy of Sciences
- Polish Academy of Sciences
- India Inter-University Centre for Astronomy and Astrophysics
- Louisiana State University
- Louisiana Tech University
- Loyola University New Orleans
- University of Maryland
- Max Planck Institute for Gravitational Physics



- University of Michigan
- University of Minnesota
- The University of Mississippi
- Massachusetts Inst. of Technology
- Monash University
- Montana State University
- Moscow State University
- National Astronomical Observatory of Japan
- Northwestern University
- University of Oregon
- Pennsylvania State University
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- Rutherford Appleton Lab
- University of Rochester
- San Jose State University
- Univ. of Sannio at Benevento, and Univ. of Salerno
- University of Sheffield
- University of Southampton
- Southeastern Louisiana Univ.
- Southern Univ. and A&M College
- Stanford University
- University of Strathclyde
- Syracuse University
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- Univ. of Texas at Brownsville
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- University of Western Australia
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