



Reducing False Alarms in Searches for Gravitational Waves from Coalescing Binary Systems

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M.S. Thesis Defense

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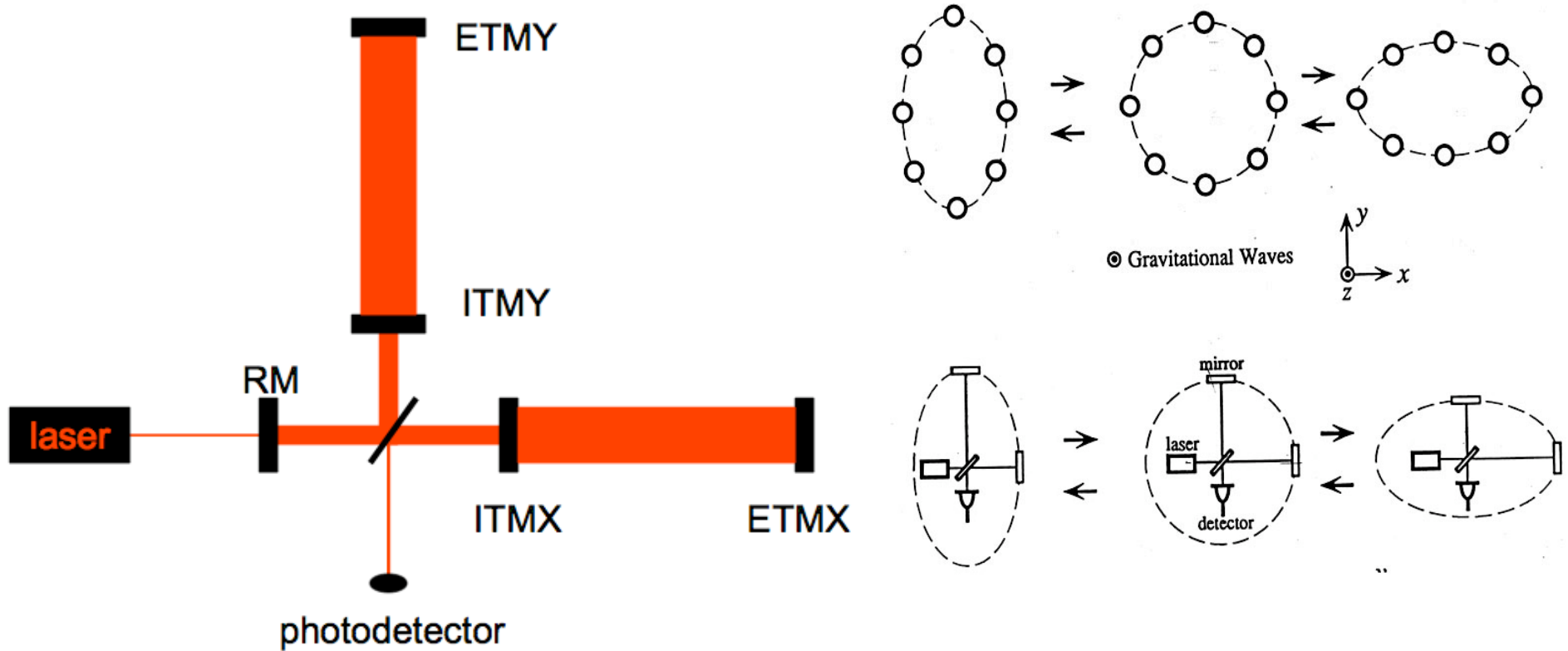




Outline

- LIGO
- Sources of Gravitational Waves
- Data Analysis for Coalescing Binary Systems
- Methods to Reduce False Alarms - χ^2 Veto, r^2 Test
- LIGO S3 Primordial Black Hole Search
- Conclusions

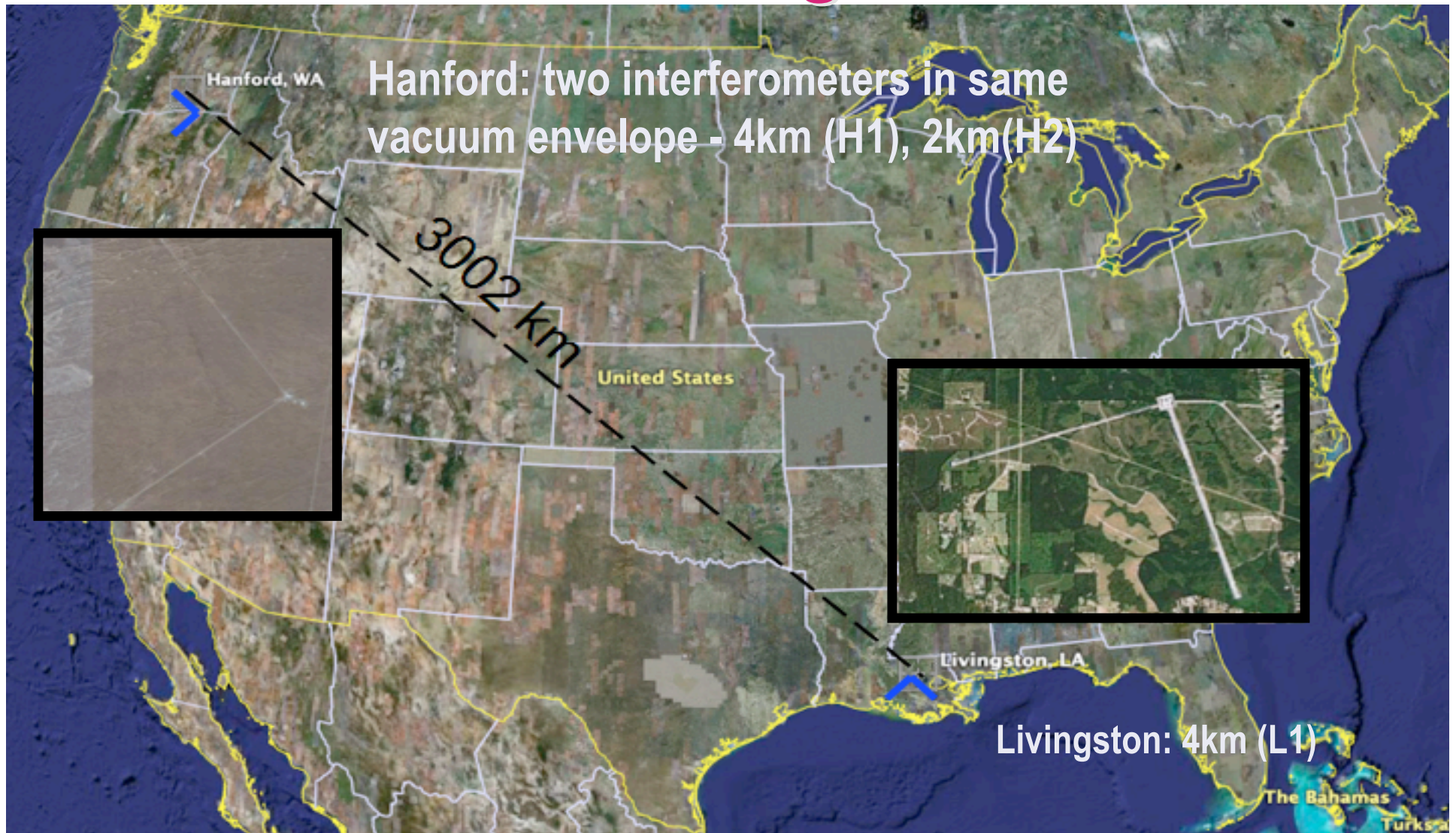
Laser Interferometer Gravitational Wave Observatory (LIGO)



$$s(t) = [\Delta L_x - \Delta L_y]/L$$

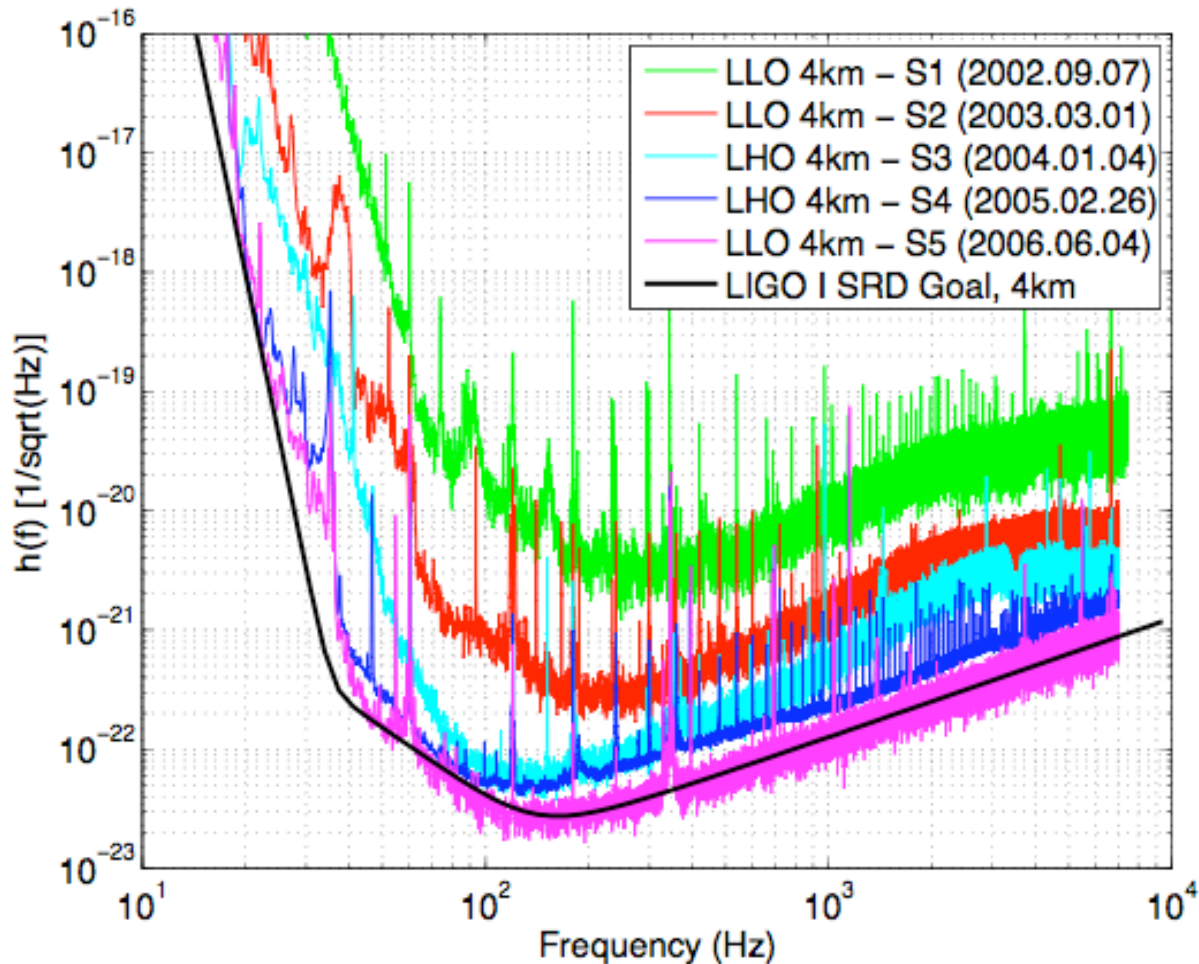


LIGO Sites lsc.org





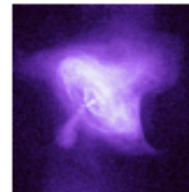
Improved Sensitivity, LIGO Science Runs



- **S1** Aug '02 - Sept '02
- **S2** Feb '03 - April '03
- **S3** Oct '03 - Jan '04
- **S4** Feb '05 - March '05
- **S5** Nov '05 - current

Sources of Gravitational Waves

- *periodic signals*: pulsars

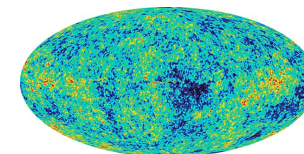


Crab pulsar (NASA, Chandra Observatory)

- *burst signals*: supernovae, gamma ray bursts

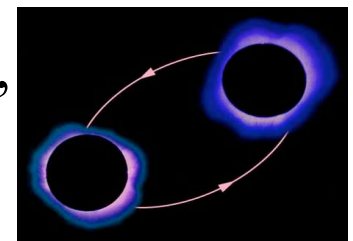


- *stochastic background*: early universe, unresolved sources



NASA, WMAP

- *compact binary coalescing (CBC) systems*: neutron stars, black holes





Coalescing Binary Systems

- General Relativity predicts the decay of the binary orbit due to the emission of gravitational radiation.
- Waveforms can be well approximated by 2nd order post-Newtonian expansion.

Non -spinning waveforms parameterized by:

- » masses: m_1, m_2 , total mass M , or reduced mass μ
- » orbital phase ϕ & orientation ι
- » position in the sky: θ, ϕ

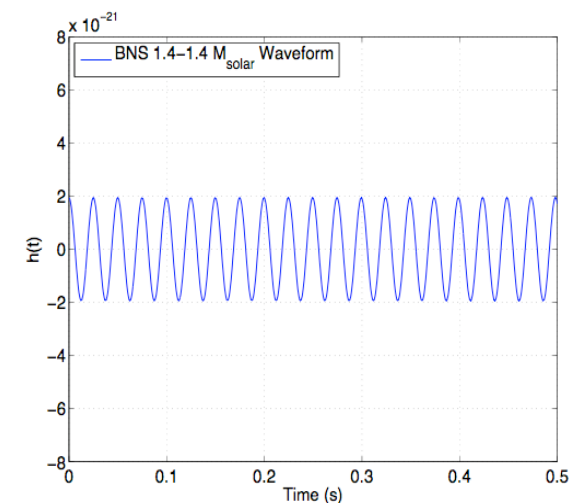
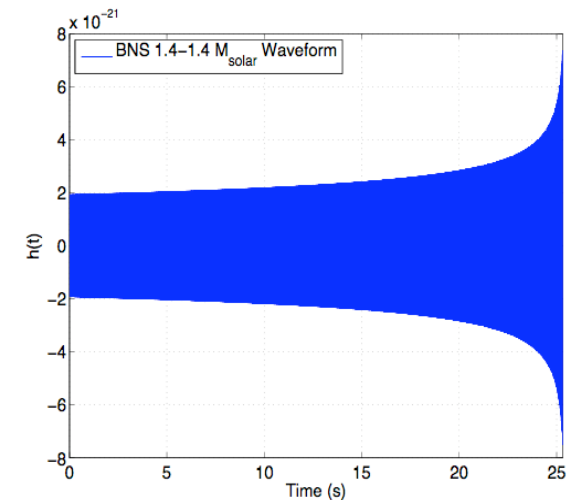
$$F_+ = -\frac{1}{2}(1 + \cos^2 \theta) \cos 2\phi$$

$$h(t) = F_+ h_+(t) + F_\times h_\times(t)$$

$$F_\times = \cos \theta \sin 2\phi$$

$$h_+(t) = \frac{1 + \cos^2 \iota}{2} \left(\frac{GM}{c^2 D} \right) \left(\frac{t_c - t}{5GM/c^3} \right)^{1/4} \times \cos \iota [2\phi_c - 2\phi(t - t_c; \mathcal{M}, \mu)]$$

$$h_\times(t) = \cos \iota \left(\frac{GM}{c^2 D} \right) \left(\frac{t_c - t}{5GM/c^3} \right)^{1/4} \times \sin \iota [2\phi_c - 2\phi(t - t_c; \mathcal{M}, \mu)]$$



Coalescing Binary Systems II

- Effective Distance:

$$D_{\text{eff}} = \frac{D}{\sqrt{F_+^2(1 + \cos^2 \iota)^2/4 + F_\times^2(\cos \iota)^2}}$$

- Chirp Mass: $\mathcal{M} = \frac{(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}}$

- Symmetric Mass ratio η :

$$\eta = \frac{m_1 m_2}{(m_1 + m_2)^2} = \frac{\mu}{\mathcal{M}}$$

Matched Filtering

- Data stream searched using matched filtering between data & template waveform: $s(t) = n(t) + h(t)$

- Matched Filter: $z(t) = 4 \int_0^\infty \frac{\tilde{h}^*(f)\tilde{s}(f)}{S_n(f)} e^{2\pi i f t} df$

Effective Distance:

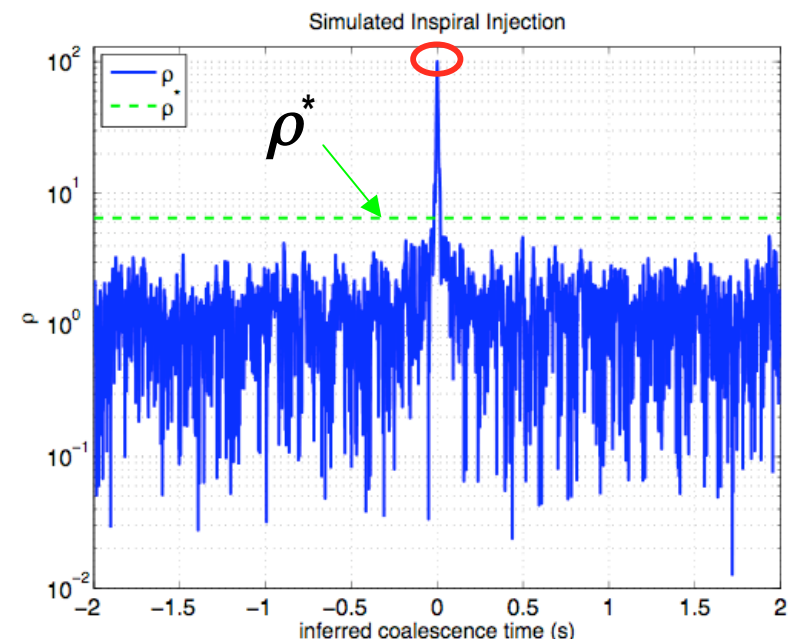
$$D_{\text{eff}} = (1\text{Mpc}) \frac{\sigma}{\rho}$$

- Characteristic Amplitude:

$$\sigma^2 = 4 \int_0^\infty \frac{|\tilde{h}(f)|^2}{S(f)} df$$

- Signal to Noise Ratio (SNR): $\rho(t) = \frac{|z(t)|}{\sigma}$

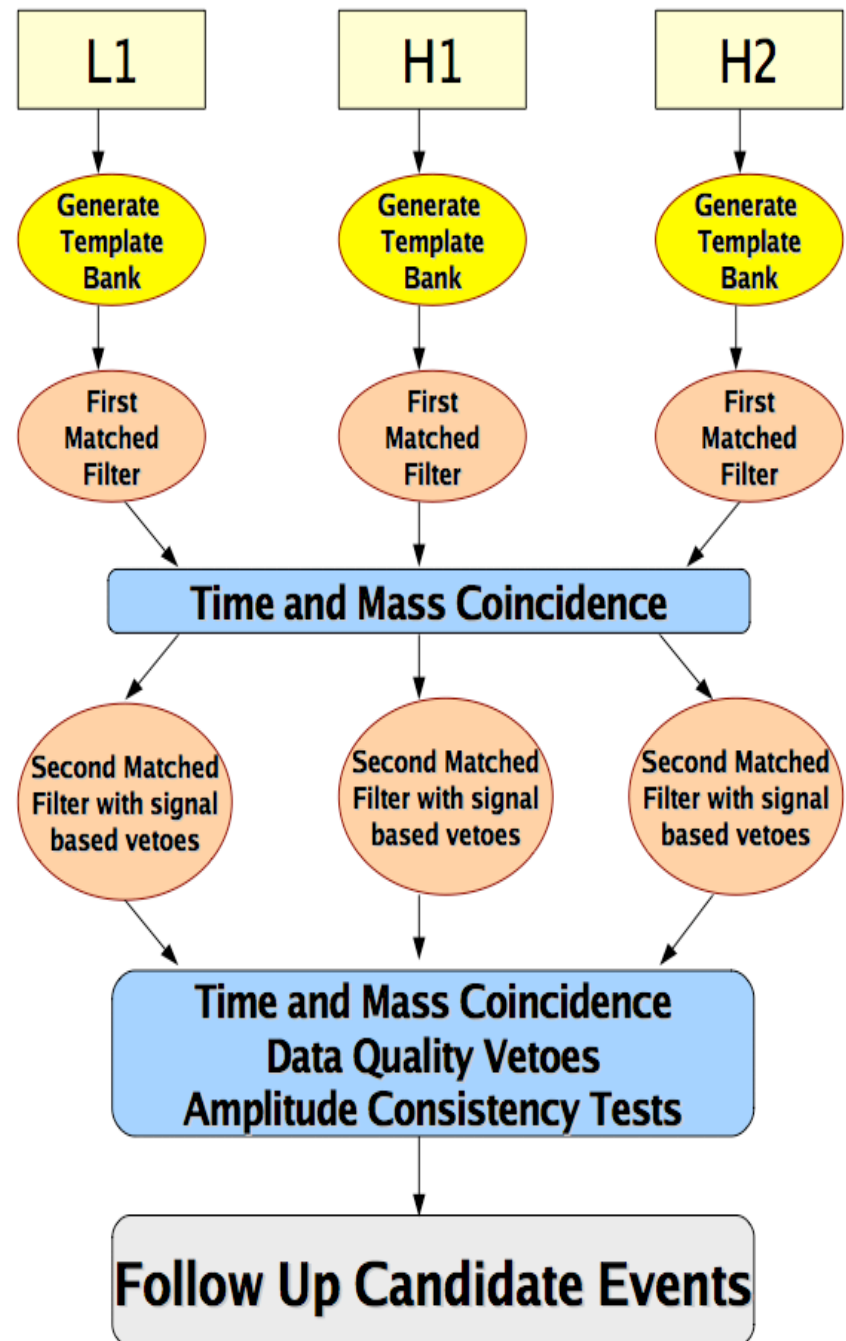
- Inspiral Trigger: $\rho = \frac{|z_{\text{max}}|}{\sigma}$ & $\rho > \rho^*$





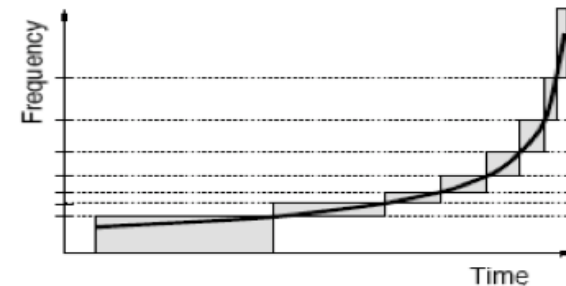
Inspiral Coincidence Pipeline

- Data Collection
- Template Bank Generation:
 $h(f)$
- Matched Filter 1: $z(t)$, $\rho(t)$,
ask $\rho > \rho^*$
- 1st Coincidence (time, η , M_c)
- Matched Filter 2:
 χ^2 , r^2 Test
- 2nd Coincidence (time, η , M_c)
- Follow Up



Methods to Reduce False Alarms I: χ^2 Veto

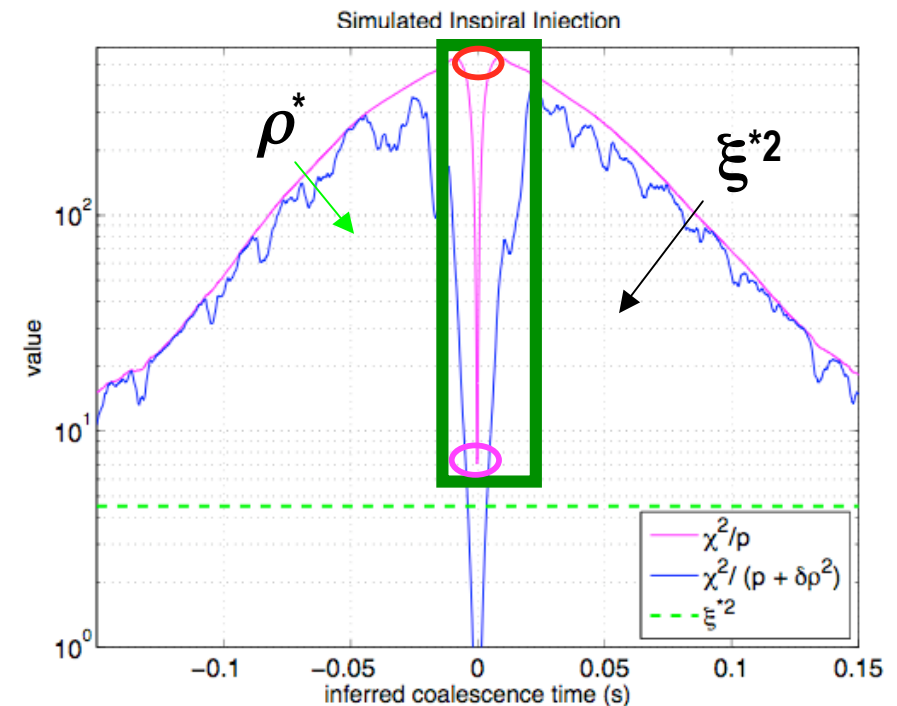
$$\chi^2(t) = p \sum_{i=1}^p |\rho_i(t) - \rho(t)/p|^2$$



$$\xi^2 = \frac{\chi^2}{p + \delta\rho^2}, \quad r^2 = \frac{\chi^2}{p}$$

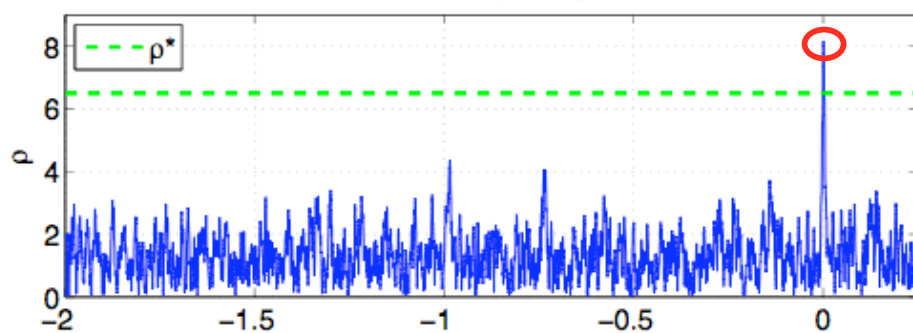
Inspiral Trigger:

$$\rho > \rho^* \quad \& \quad \frac{\chi^2}{p + \delta\rho^2} < \xi^{*2}$$

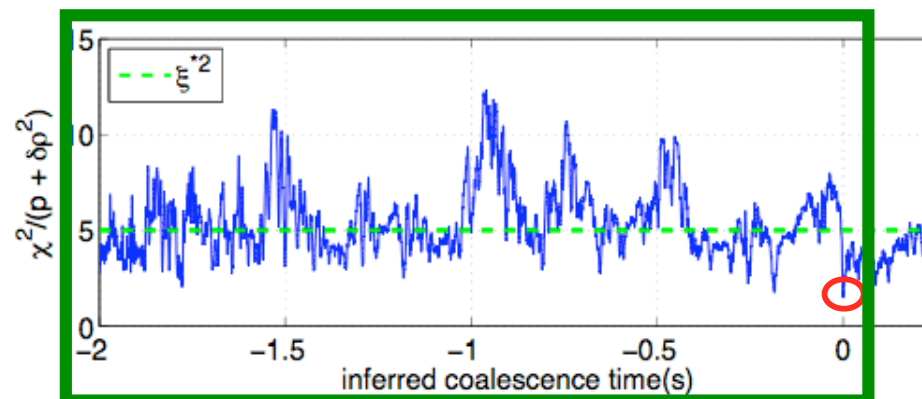
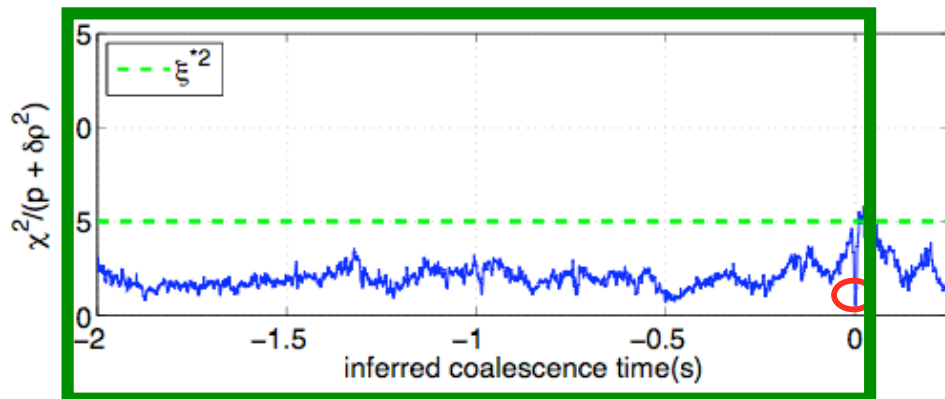
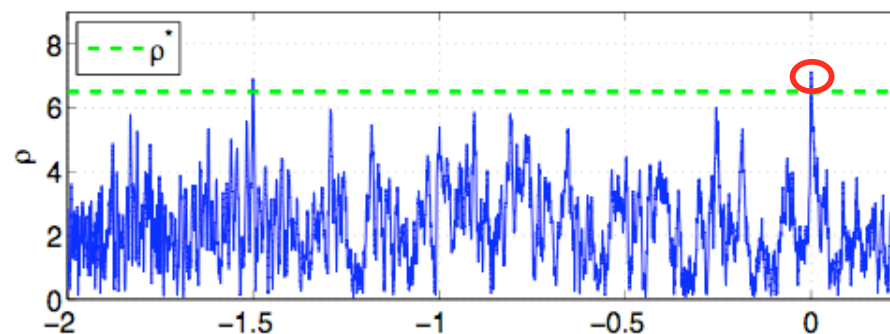


Simulated Waveform vs False Alarm

Simulated Inspiral Injection



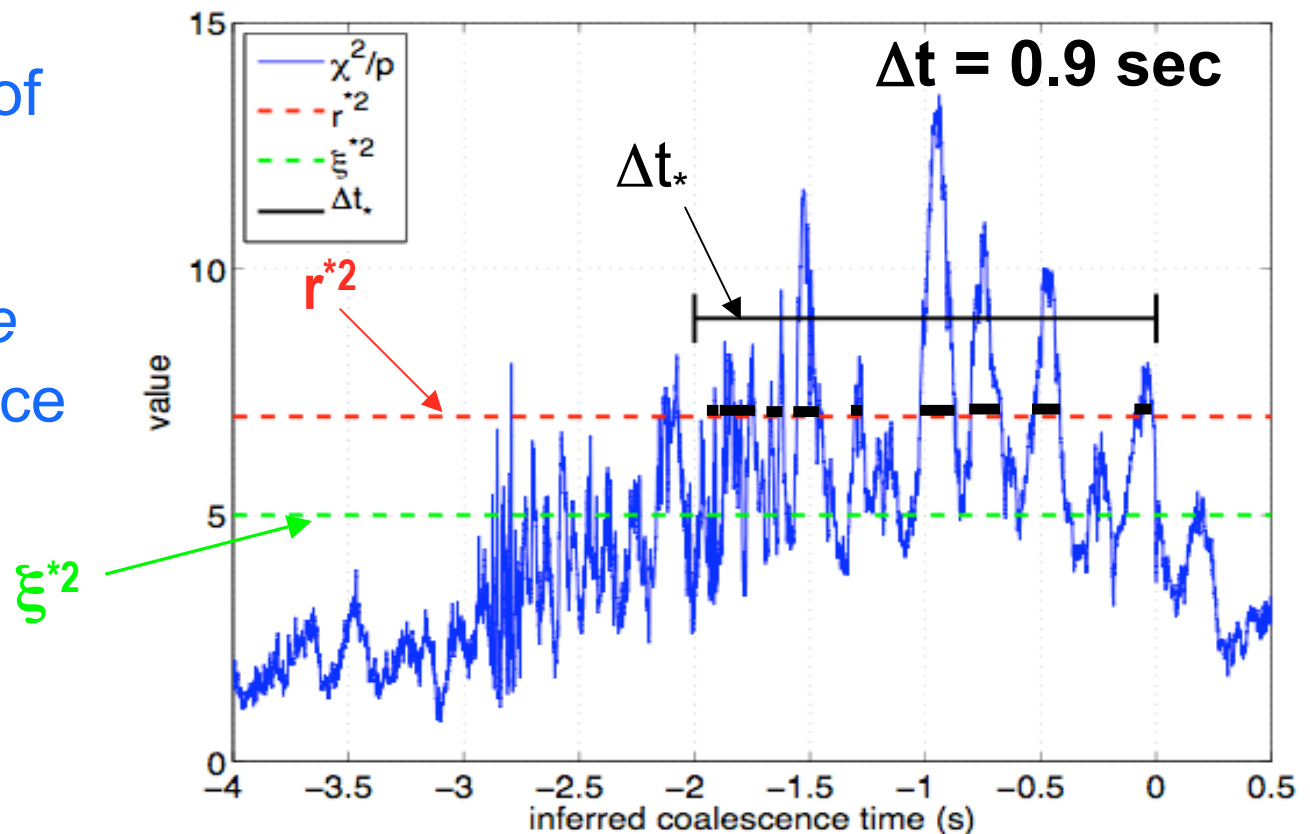
False Alarm



Methods to Reduce False Alarms II: r^2 Test

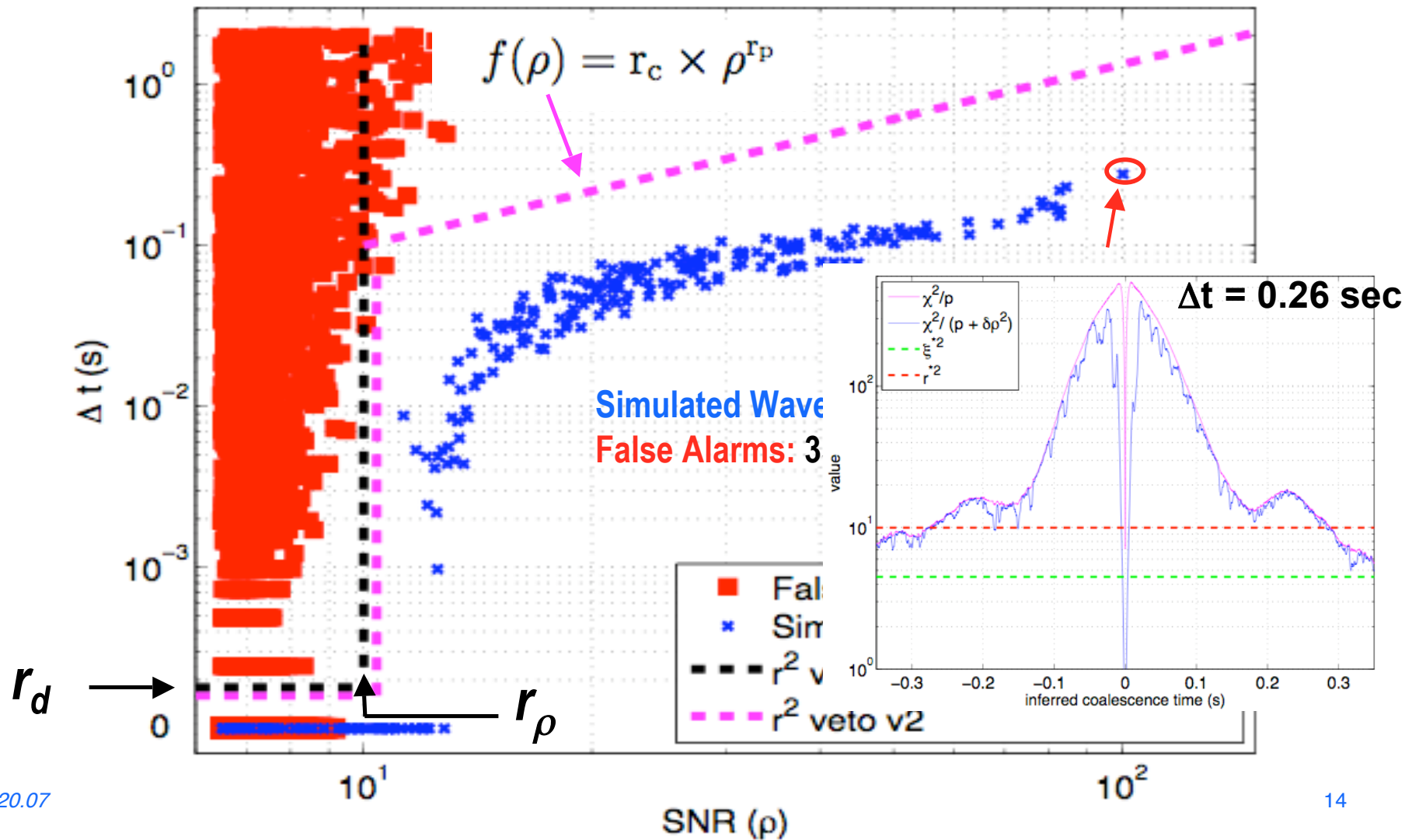
- Use the r^2 time series (χ^2/p) as a method to search for excess noise.
- Impose a higher r^2 threshold (r^{*2}) than the search employs.

» Count the number of time samples (Δt) above r^{*2} in a time interval (Δt_*) before inferred coalescence



Methods to Reduce False Alarms II

r^2 Test Result: S4 BNS Search





r^2 Test Results: Searches performed by CBC Group in the LSC

r^2 Test Results for LIGO CBC Searches

	r^2 veto version	Falsely Dismissed Injections (%)	Vetoed False Alarms (%)
S3BNS	1	0.001	43.0
S3PBH	1	0.0	26.5
S4BNS	1	0.0	35.0
S4PBH	1	0.0	35.0
S5BNS (epoch 1)	2	0.001	26.9
S5BBH (epoch 1)	2	0.12	19.1

- r^2 test is included in current LIGO searches:
S5 Low Mass ($M < 35M_{\text{SOL}}$) 1 year

S3 Primordial Black Hole (PBH) Search I

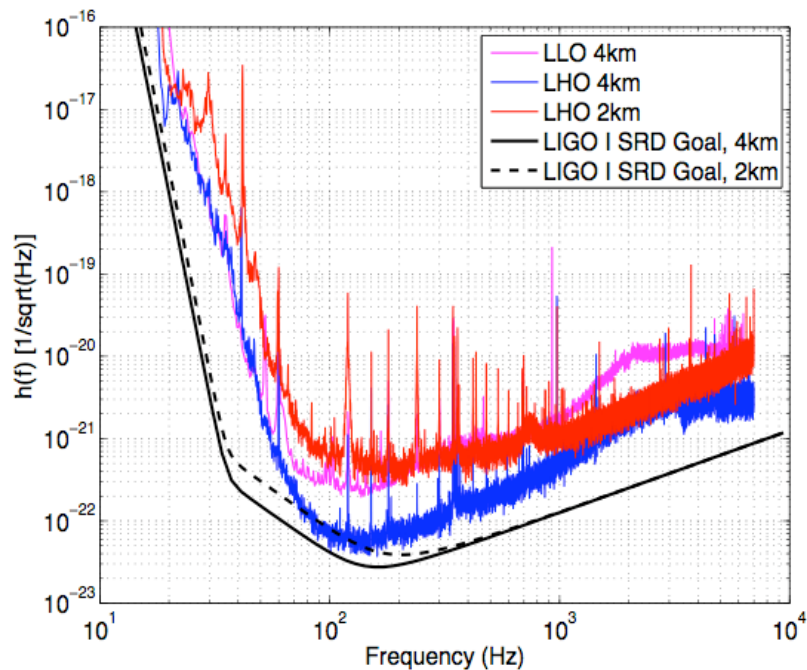
- Black hole composed with mass $< 1.0 M_{\text{SOL}}$ is believed to be a primordial black hole (PBH).
- They are compact objects may have formed in the early, highly compressed stages of the universe immediately following the big bang.
- Speculated to be part of the galactic halo or constituent of dark matter (small fraction).
- Binary system composed of two PBH's will emit gravitational waves that may be detectable by LIGO.
- A PBH binary composed of $2 \times 0.35 M_{\text{SOL}}$ objects would have a coalescence frequency of 2023Hz and would spend about 22 seconds in LIGO's sensitive band.



S3 PBH Search II

- Target Sources:

$m_{\min}(M_{\odot})$	$m_{\max}(M_{\odot})$	$f_L(\text{Hz})$	N_b	$D_{\max}(\text{s})$
0.35	1.0	100	4500	22.1



S3: October 3 - January 09: 2004

S3	
H1-H2-L1 times	184 (167) hrs
H1-H2 times	604 (548) hrs
Total times	788 (715) hrs

* Numbers in () represent time left after data used for tuning search.

S3 PBH Search III - Tuning

- Coincidence Windows

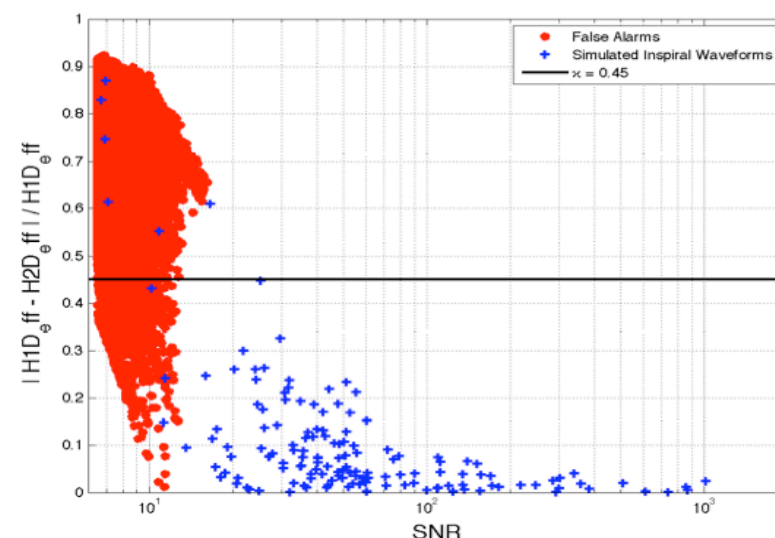
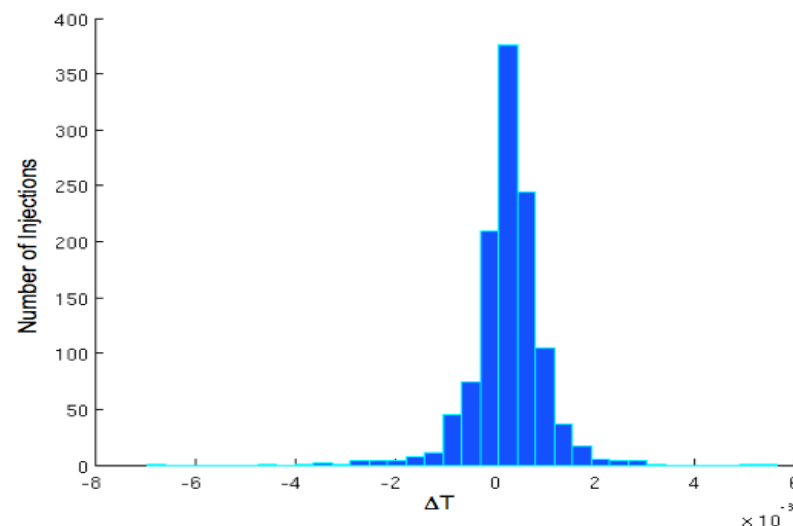
$$\eta = \frac{m_1 m_2}{(m_1 + m_2)^2} \quad \mathcal{M} = \frac{(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}}$$

- H1H2 Effective Distance Cut

$$\frac{|H1D_{\text{eff}} - H2D_{\text{eff}}|}{H1D_{\text{eff}}} < \kappa \quad , \quad \kappa = 0.45$$

- Parameters Selected:

ΔT (milliseconds)	$\Delta \mathcal{M}_c (M_\odot)$	$\Delta \eta$
4×2	0.002×2	0.06



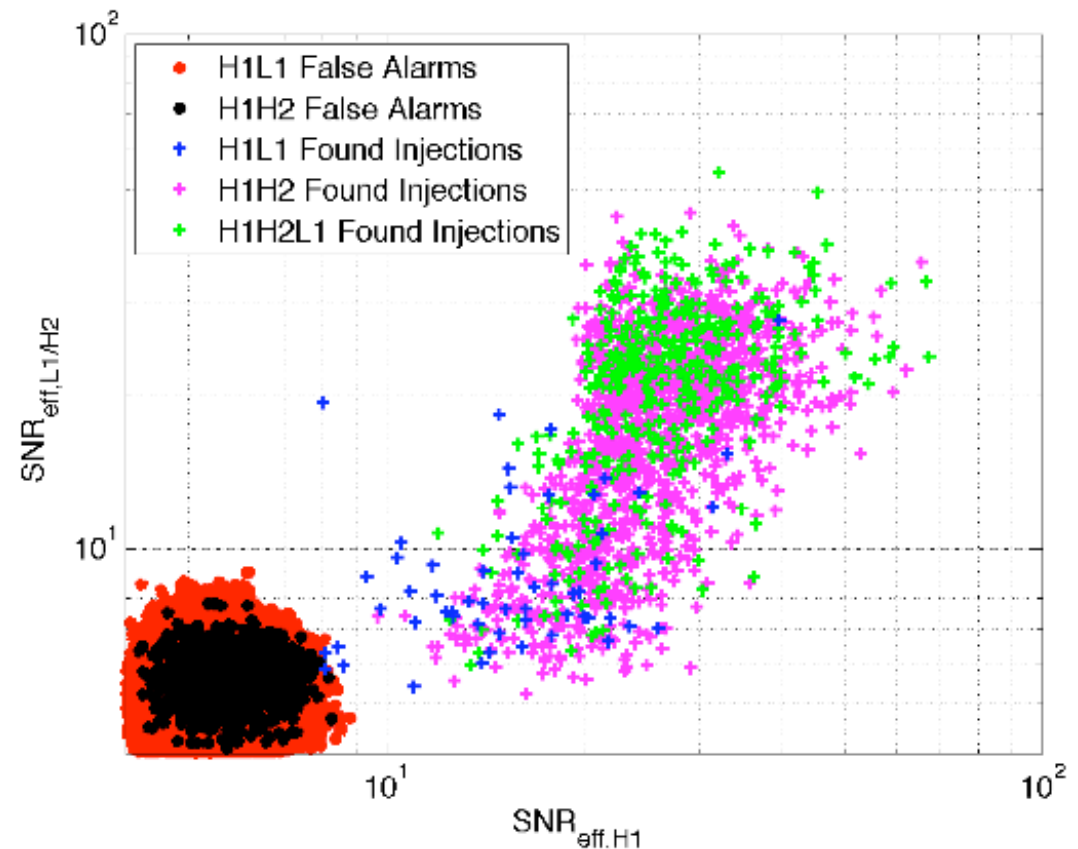
S3 PBH Search IV - Tuning

- Effective SNR:

$$\rho_{\text{eff}}^2 = \frac{\rho^2}{\sqrt{\left(\frac{\chi^2}{2p-2}\right) \left(1 + \frac{\rho^2}{250}\right)}}$$

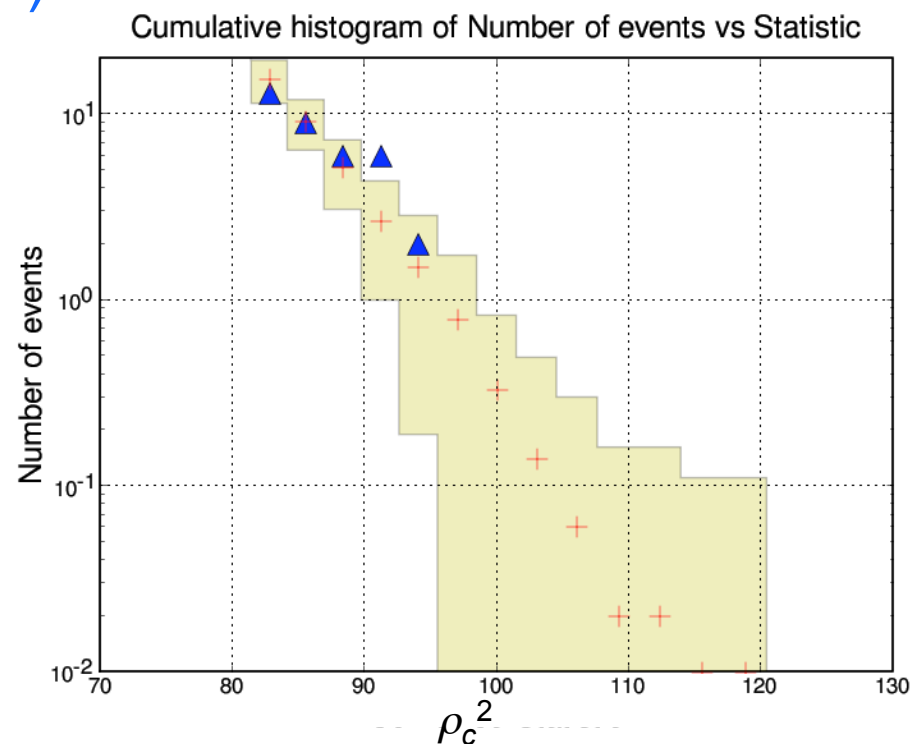
- Combined SNR:

$$(\rho_c)_{\text{PBH}}^2 = \sum_i^N \rho_{\text{eff},i}^2$$



S3 PBH Search V: Result

- No triple coincident foreground candidate events or background events were found.
- Number of double coincidences found (▲) consistent with measured background (+).





Conclusions

- r^2 test greatly reduces rate of false alarms in LIGO CBC searches.
- The r^2 test be incorporated into future LIGO searches: S5 Low Mass ($M < 35M_{\text{SOL}}$) 1 year.
- A search for primordial black hole binary systems ($M < 1M_{\text{SOL}}$) in LIGO's S3 run was performed with results consistent with measured background.



Thank You!

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LSC: Prof. Peter Shawhan, Prof. Alan Weinstein, Prof. Laura Cadonati, Dr. John Whelan,
Prof. Vicky Kalogera, Dr. Bill Kells

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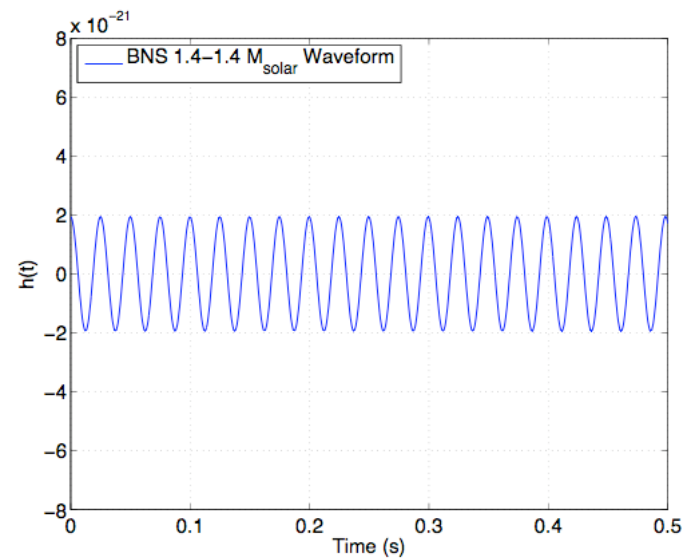
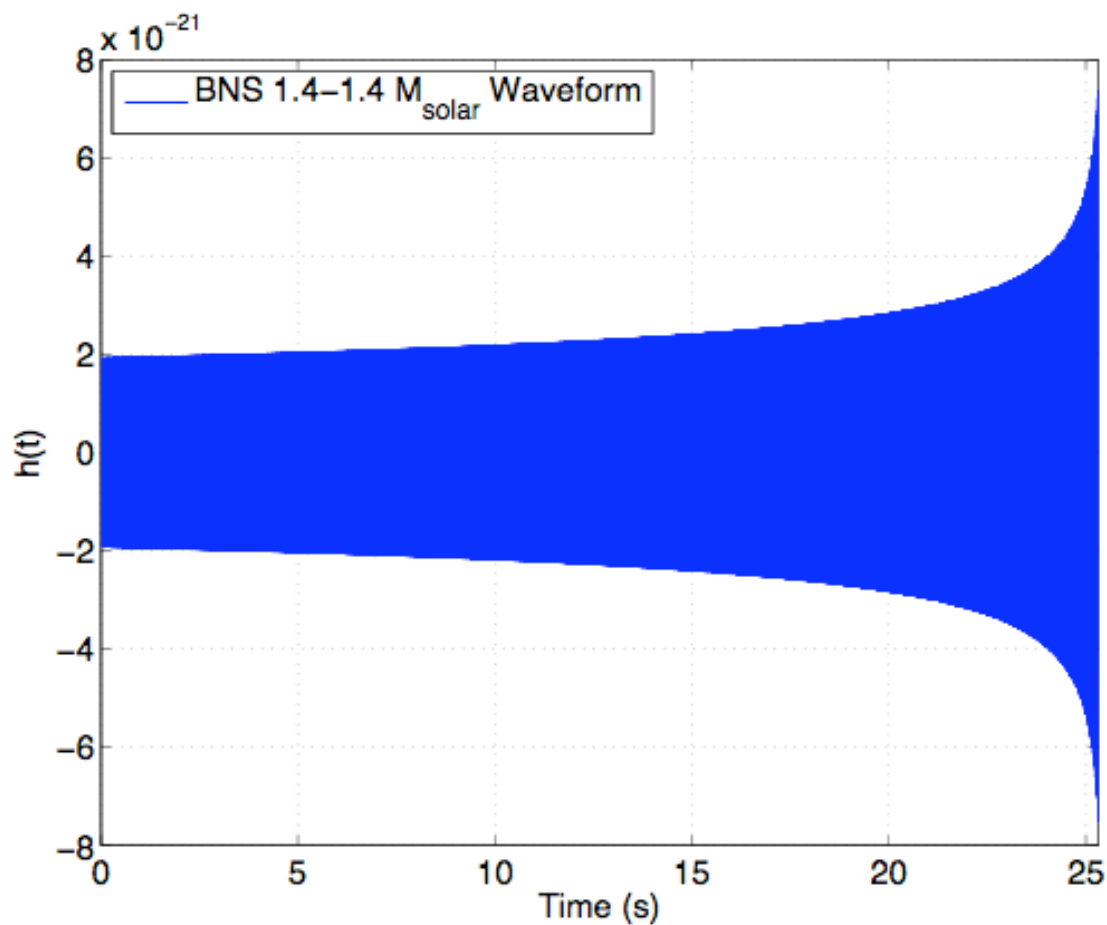
*This research has been supported by National Science Foundation grants:
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Extra Slides

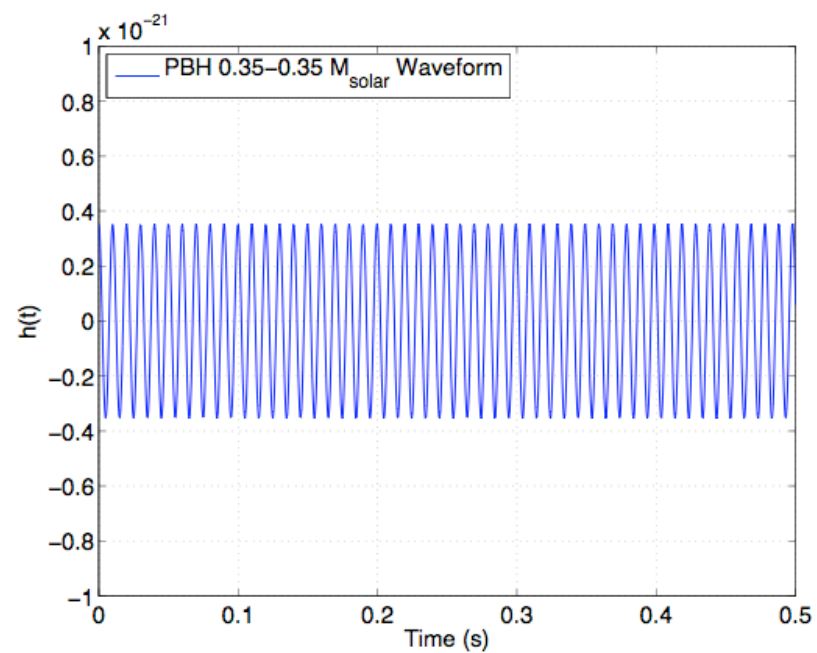
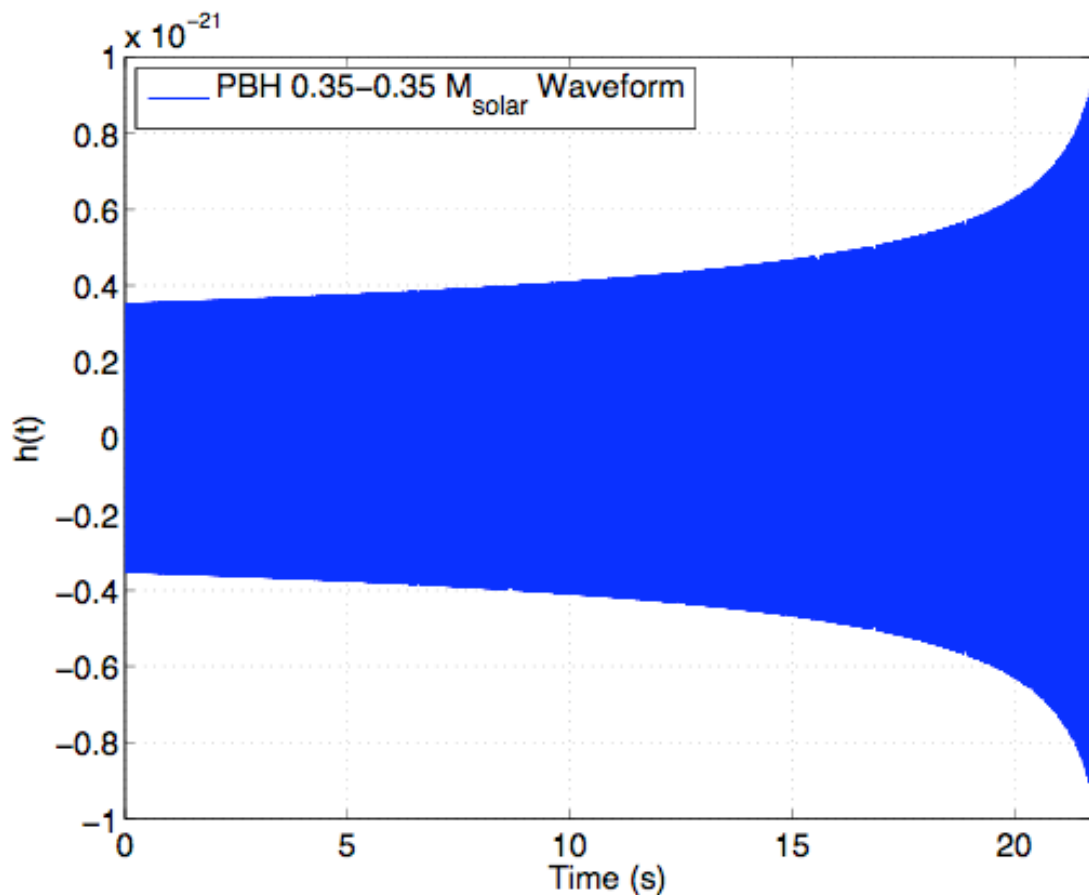


BNS Waveform



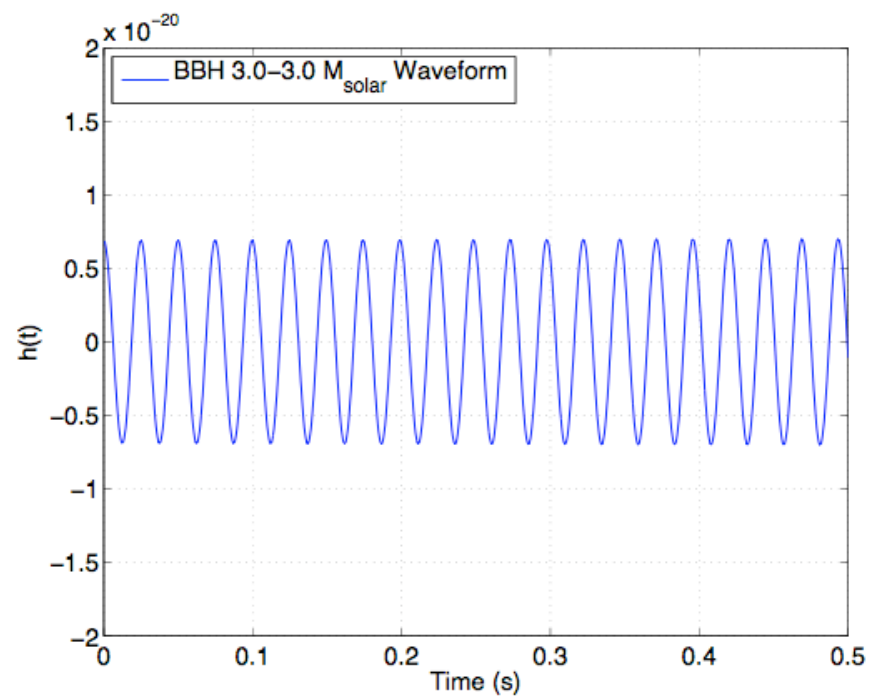
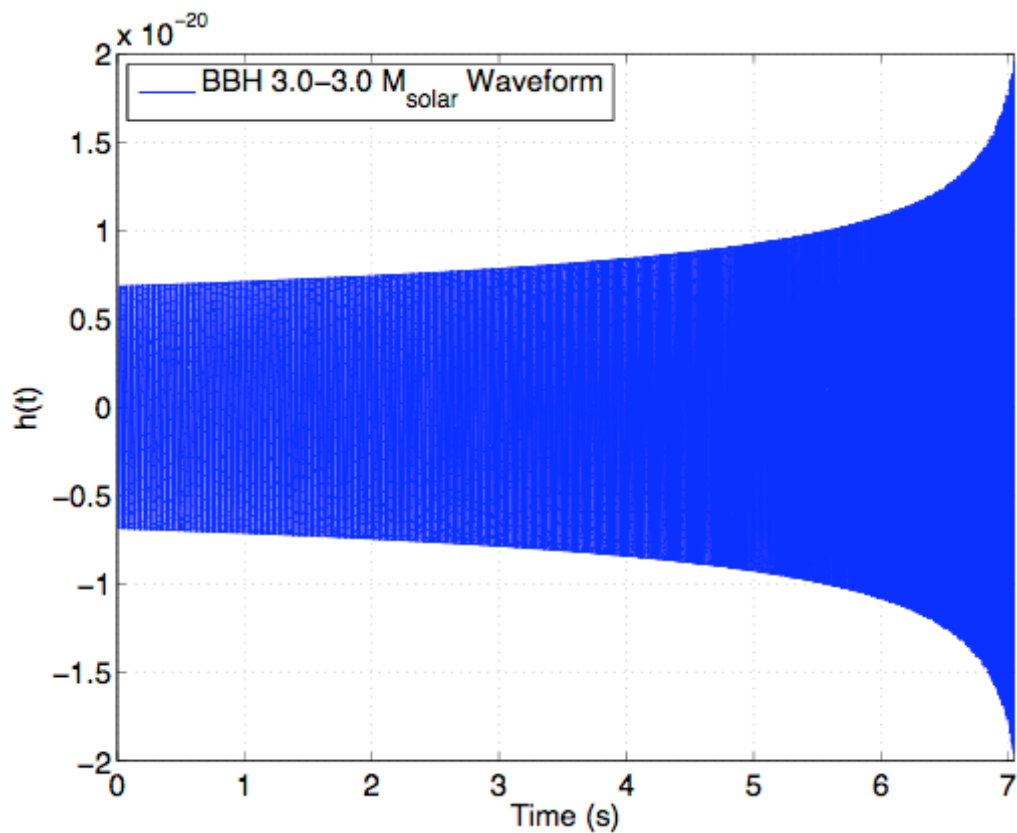


PBH Waveform

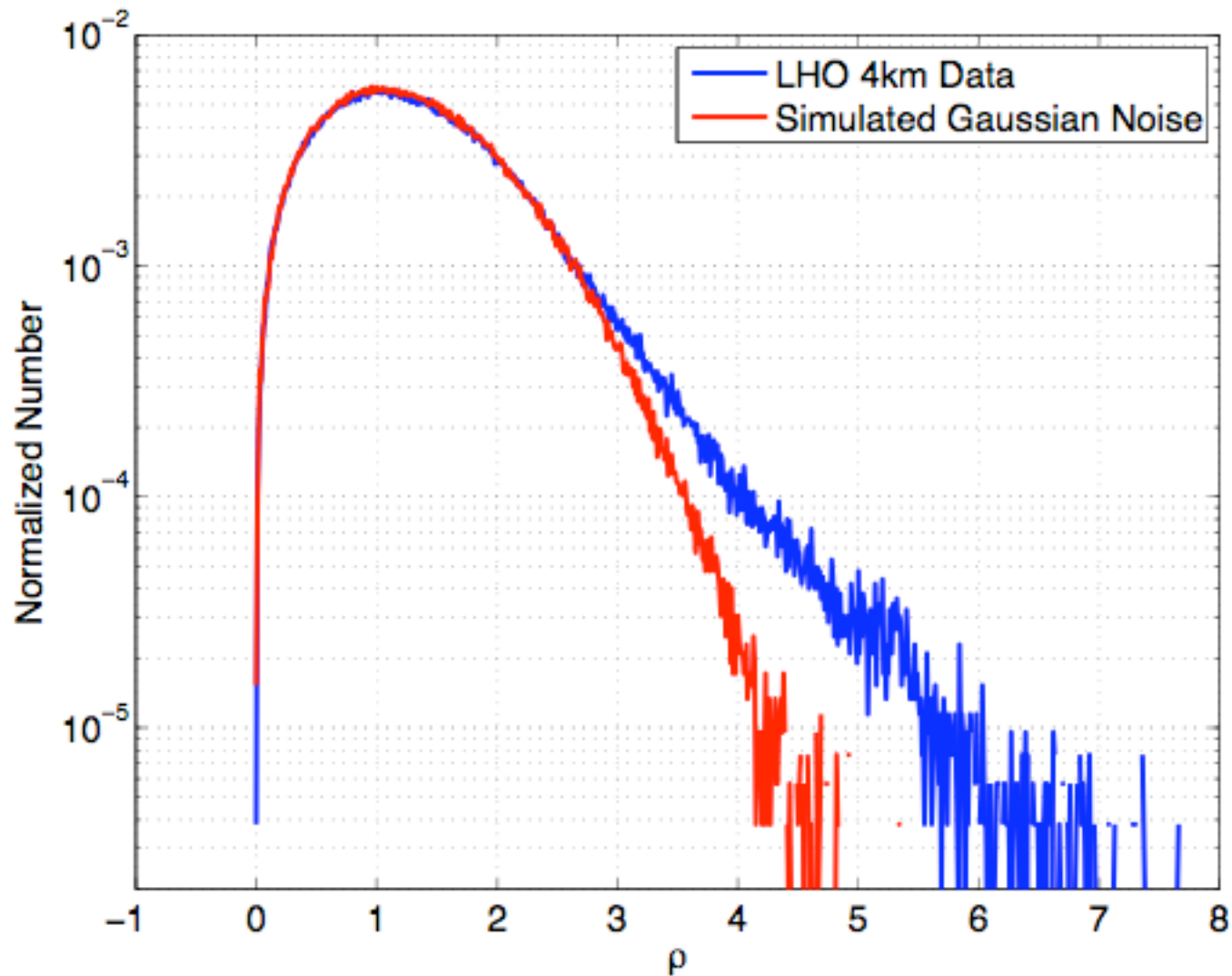




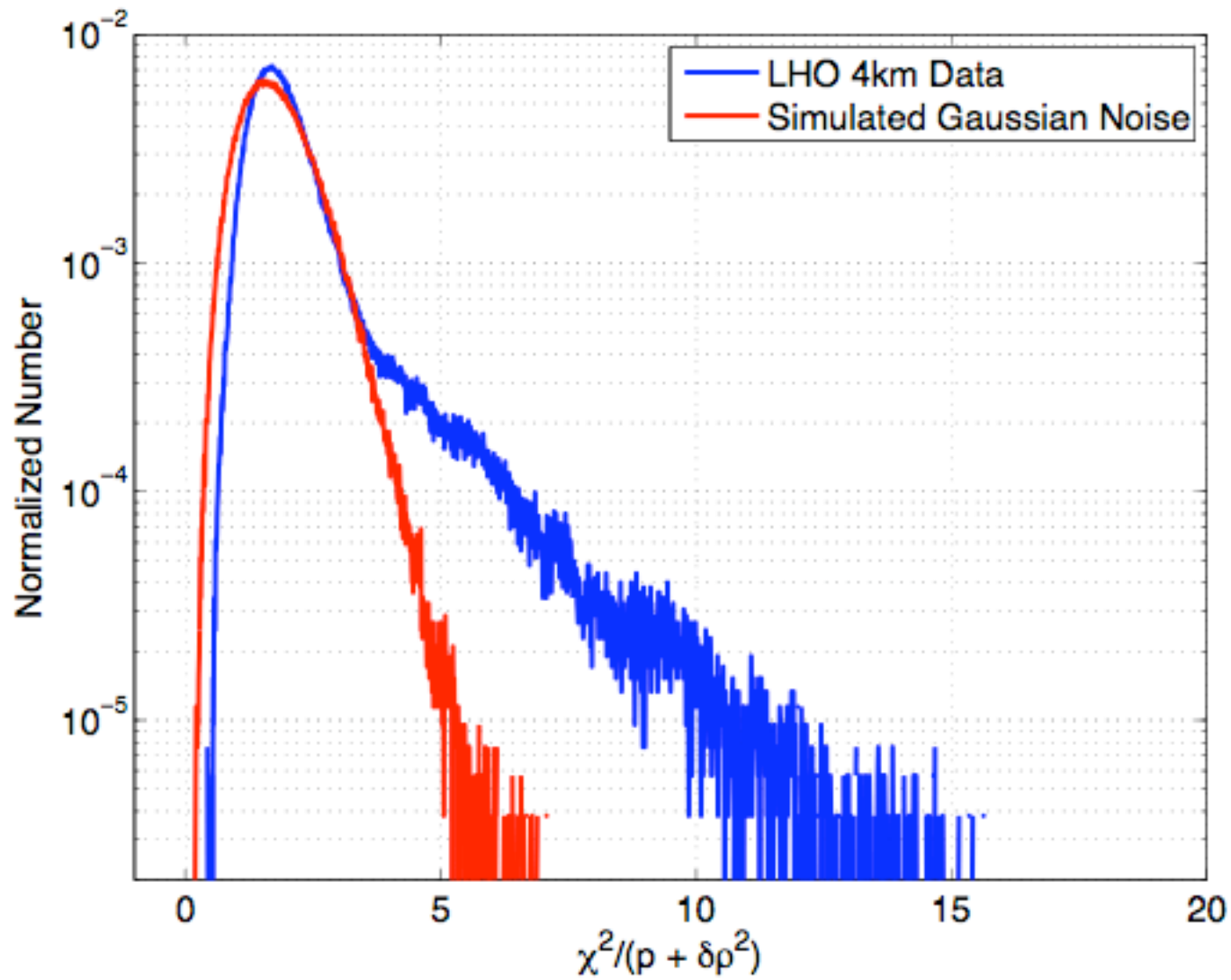
BBH Waveform



SNR

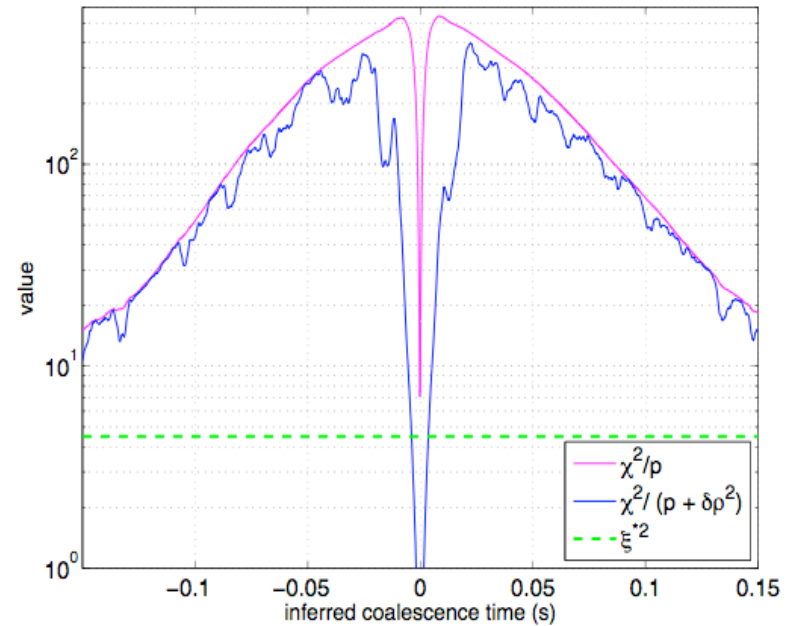
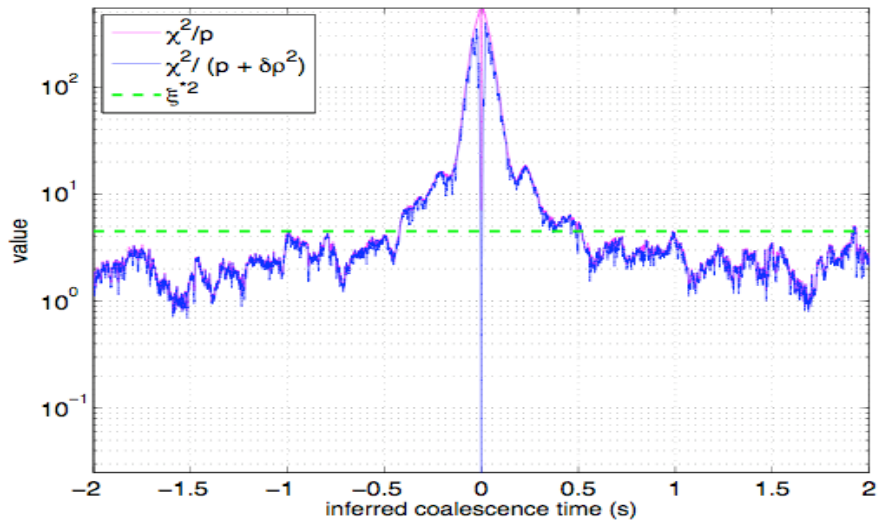
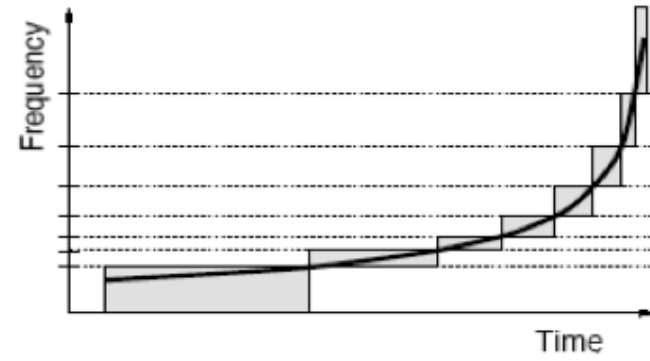
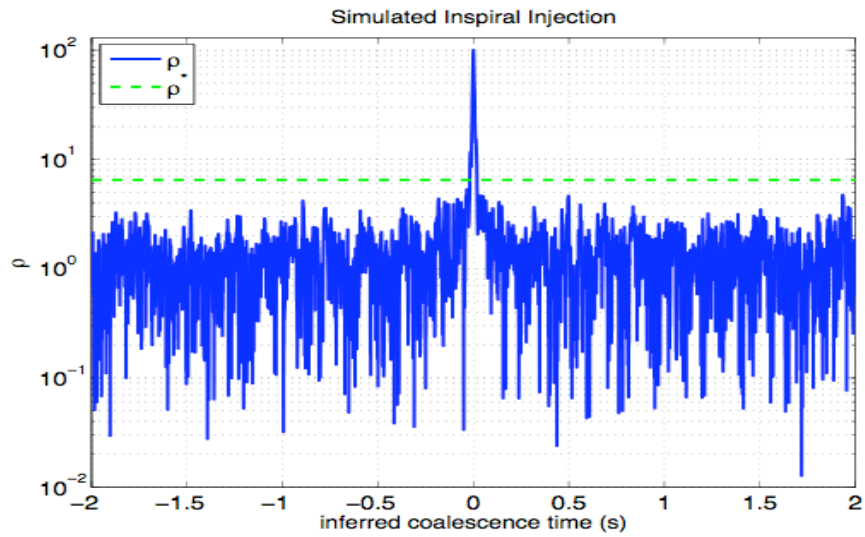


χ^2

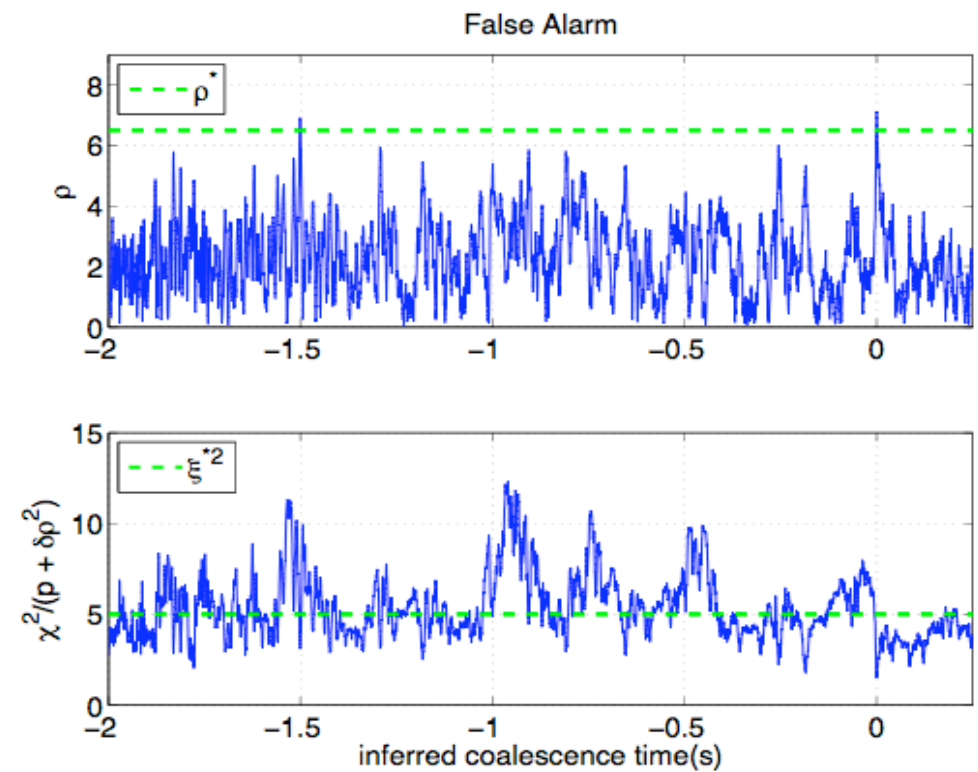
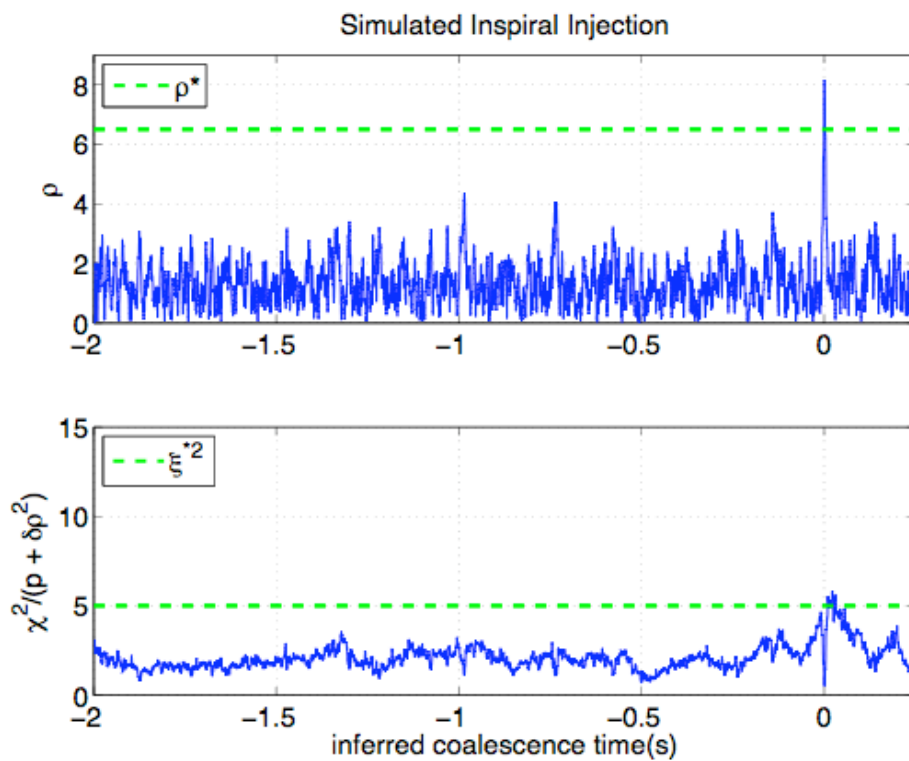




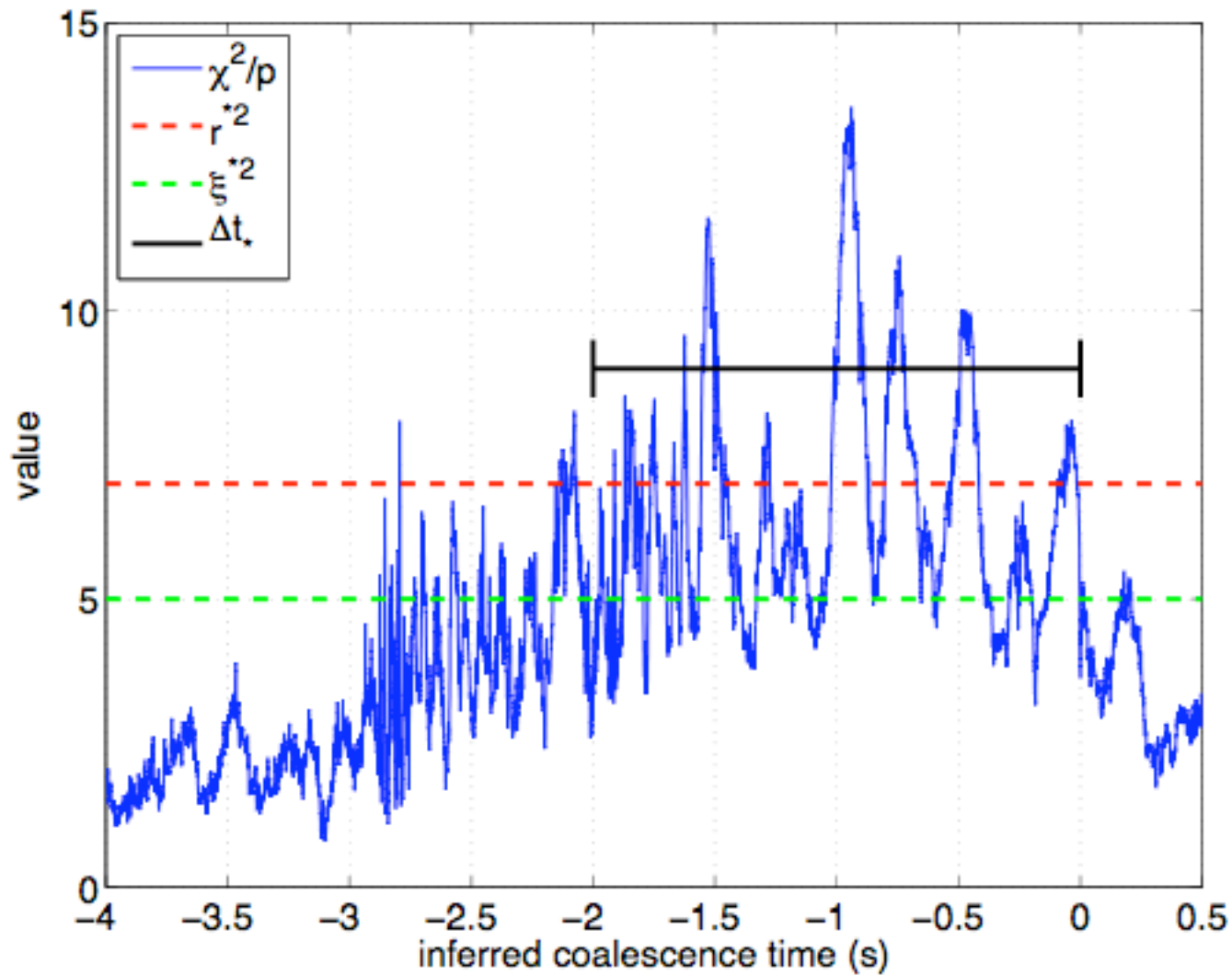
Injection



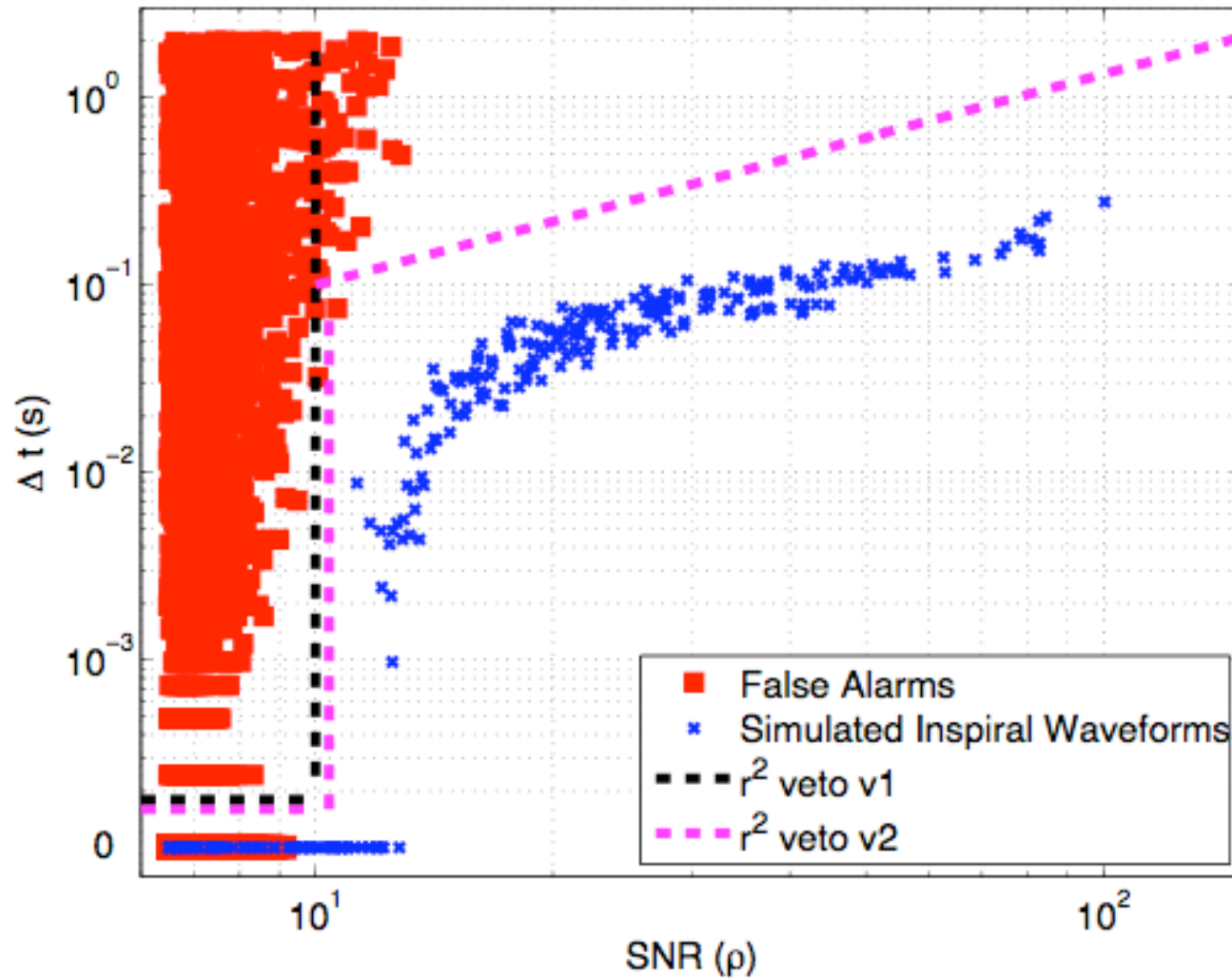
Injection vs Trigger



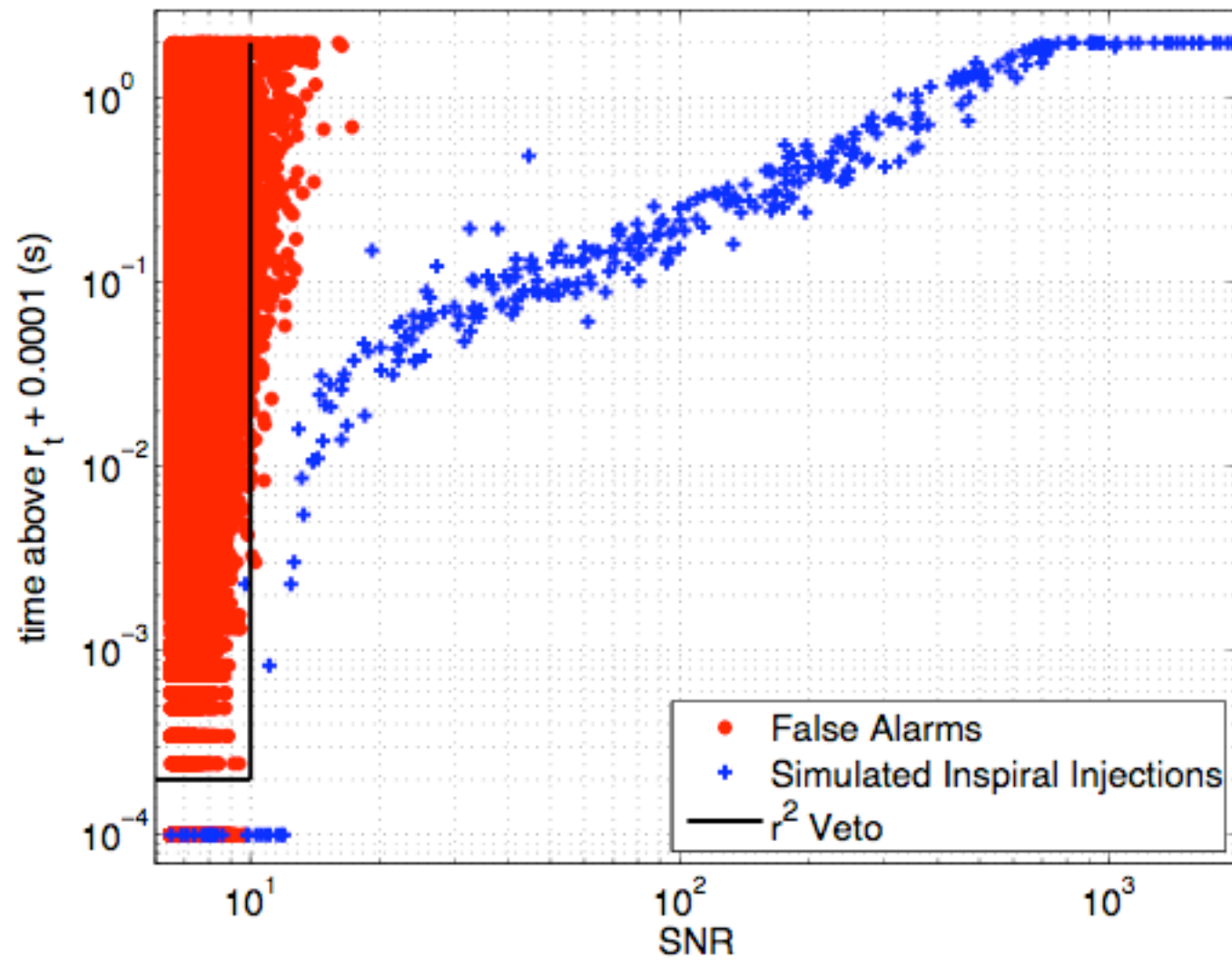
r^2 test



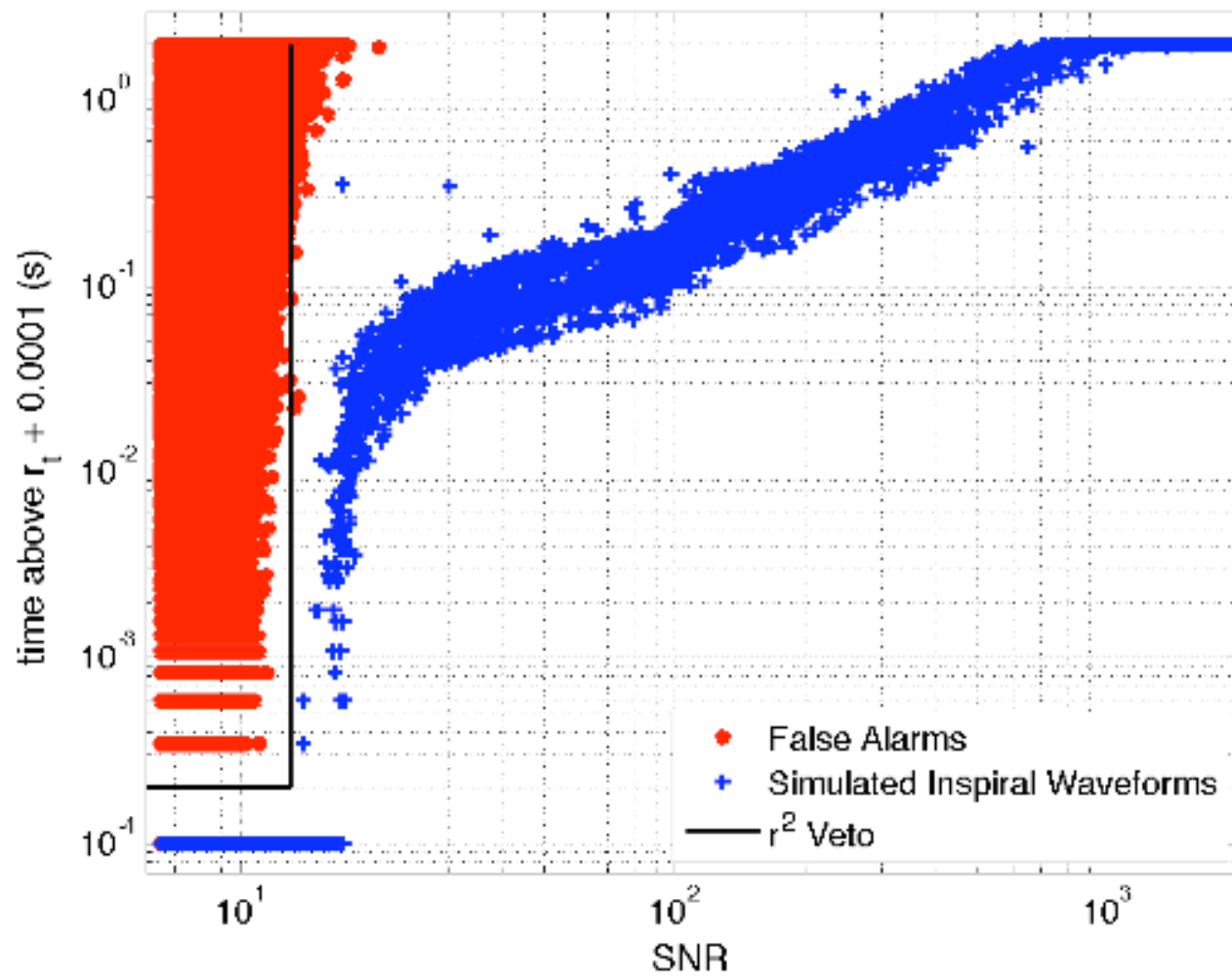
r^2 example



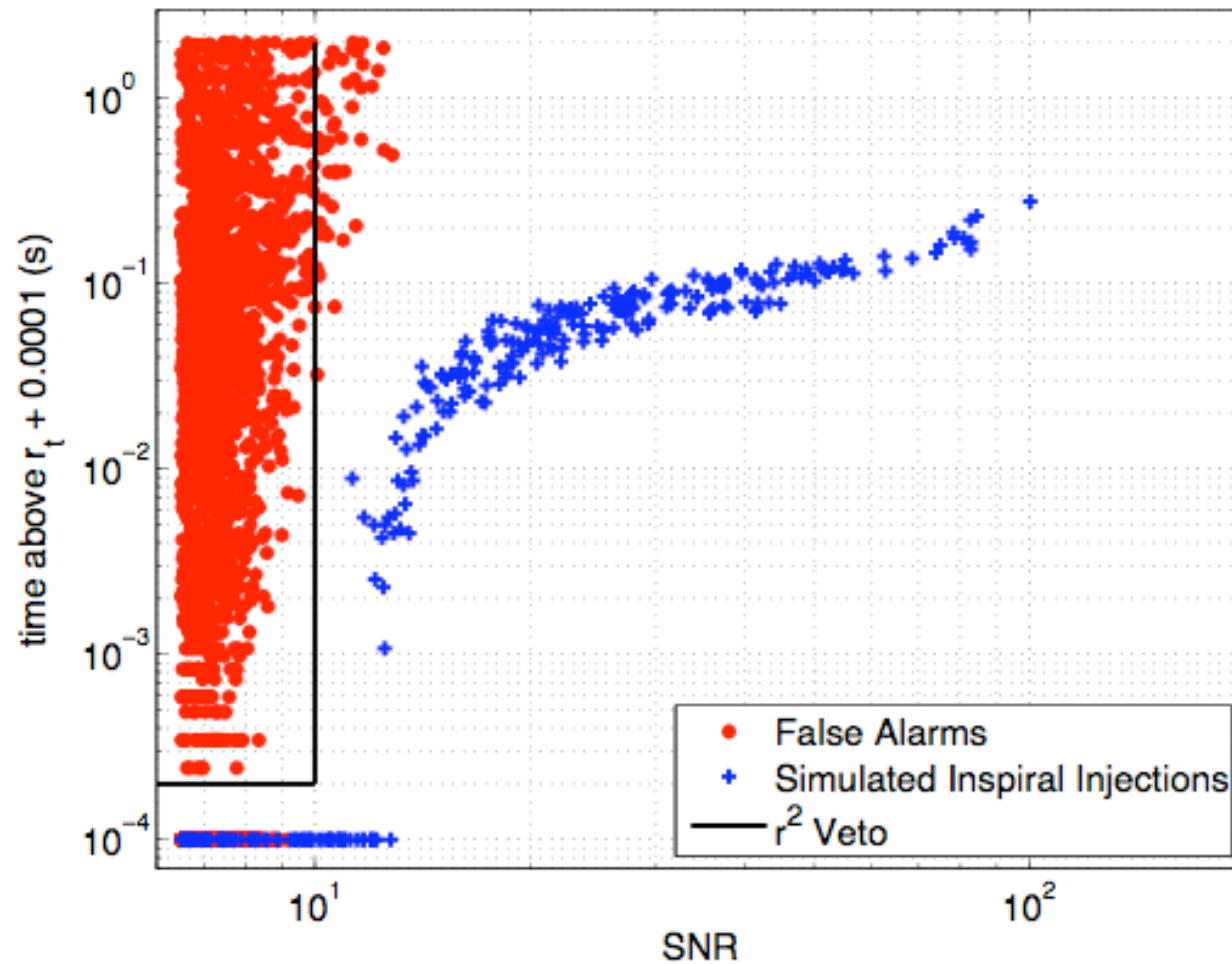
r^2 - S3 BNS



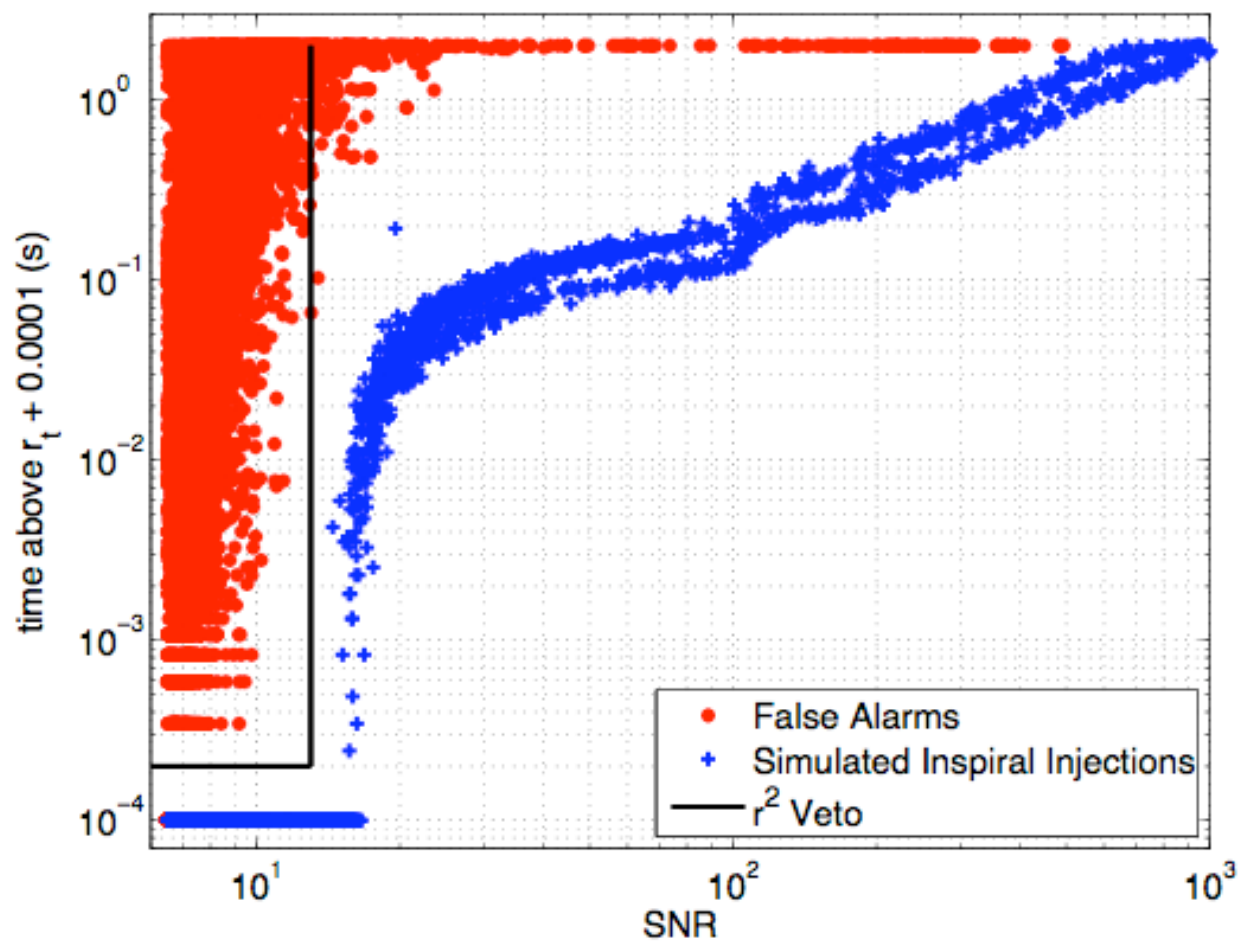
r^2 - S3 PBH



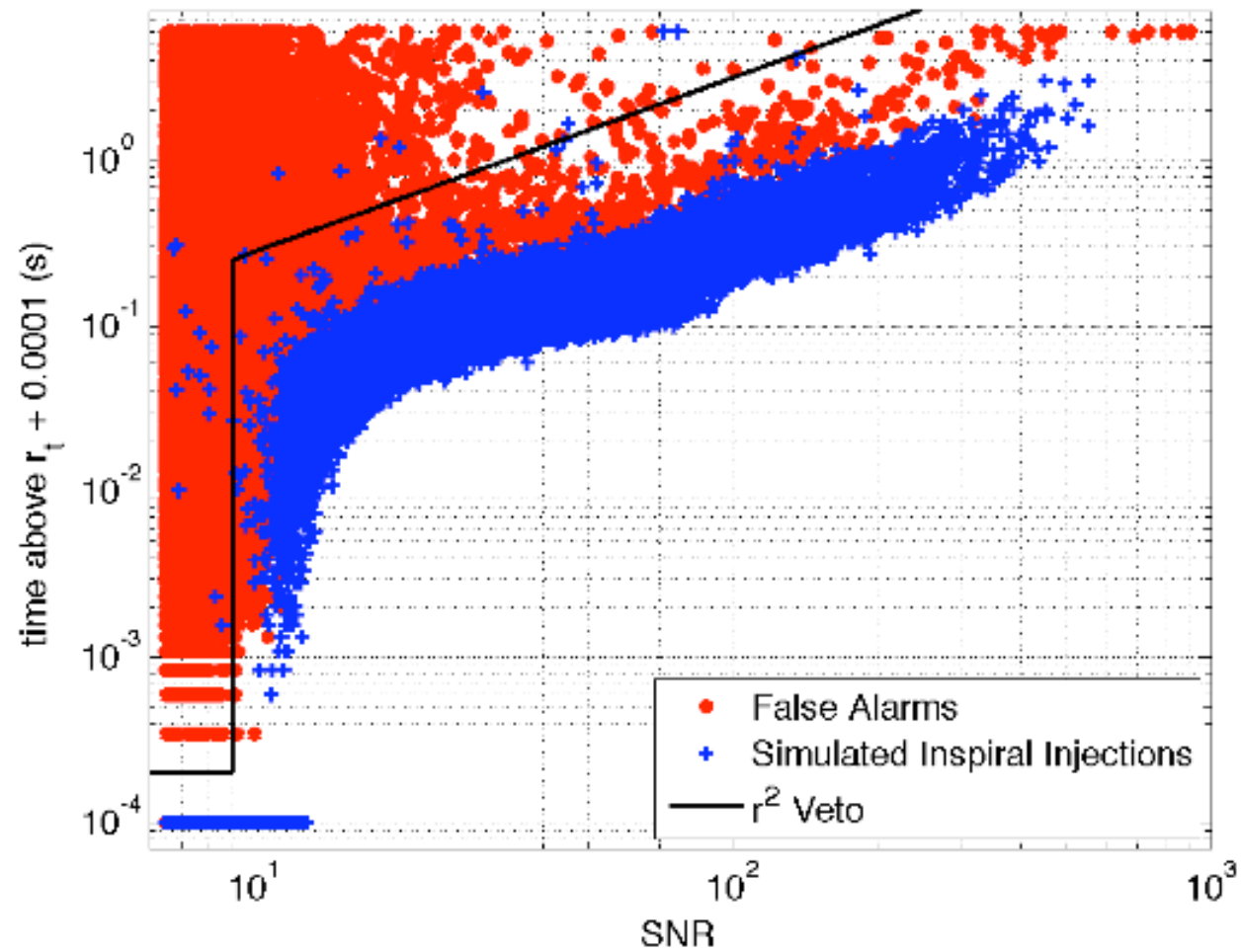
r^2 - S4 BNS



r^2 - S4 PBH

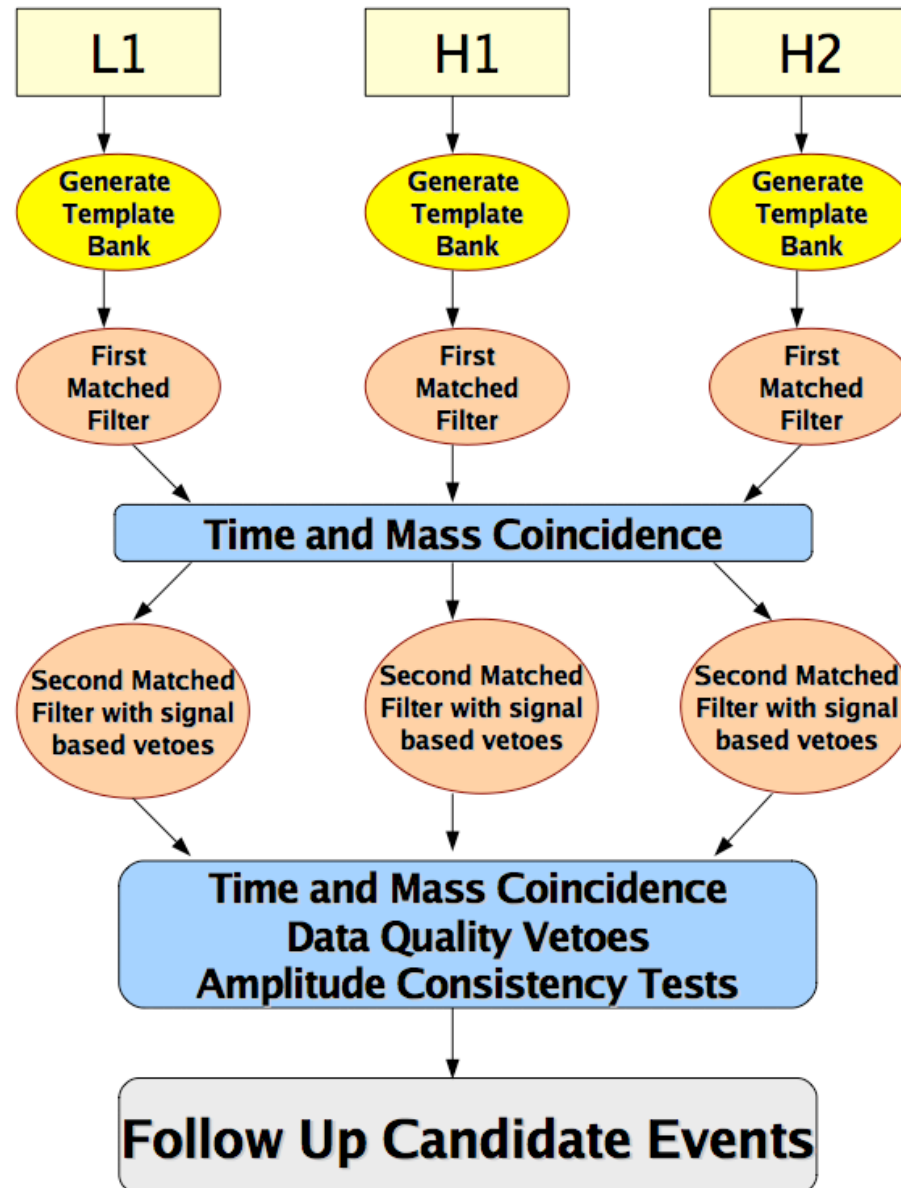


r^2 - S5 BBH epoch 1

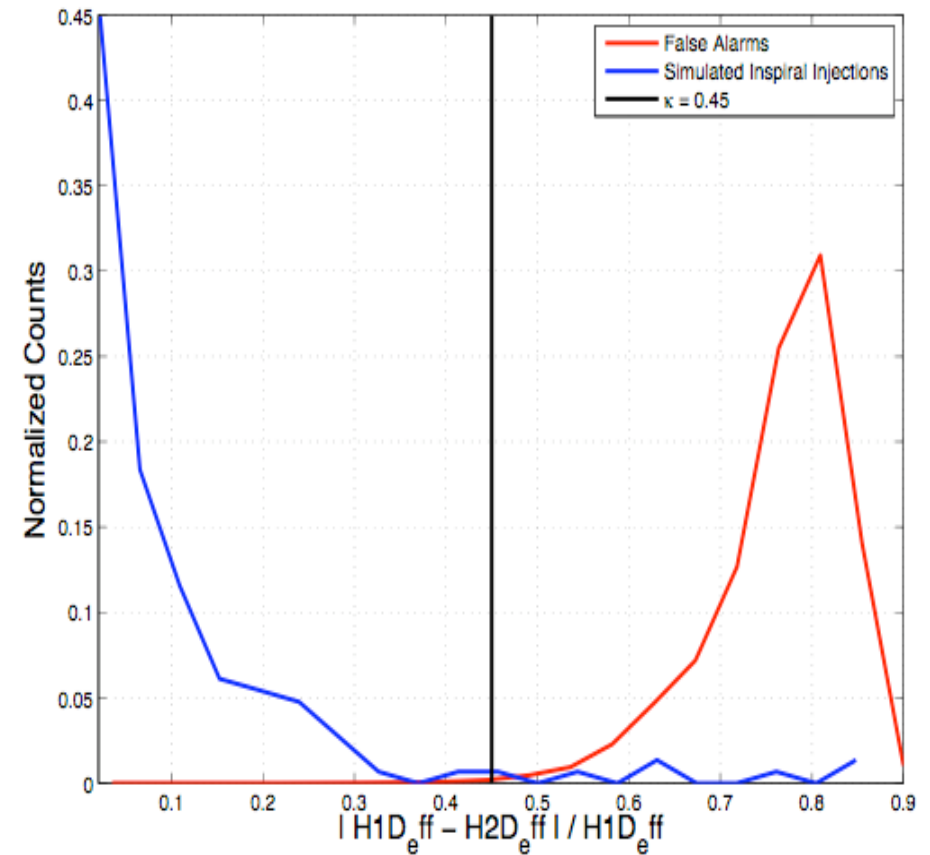
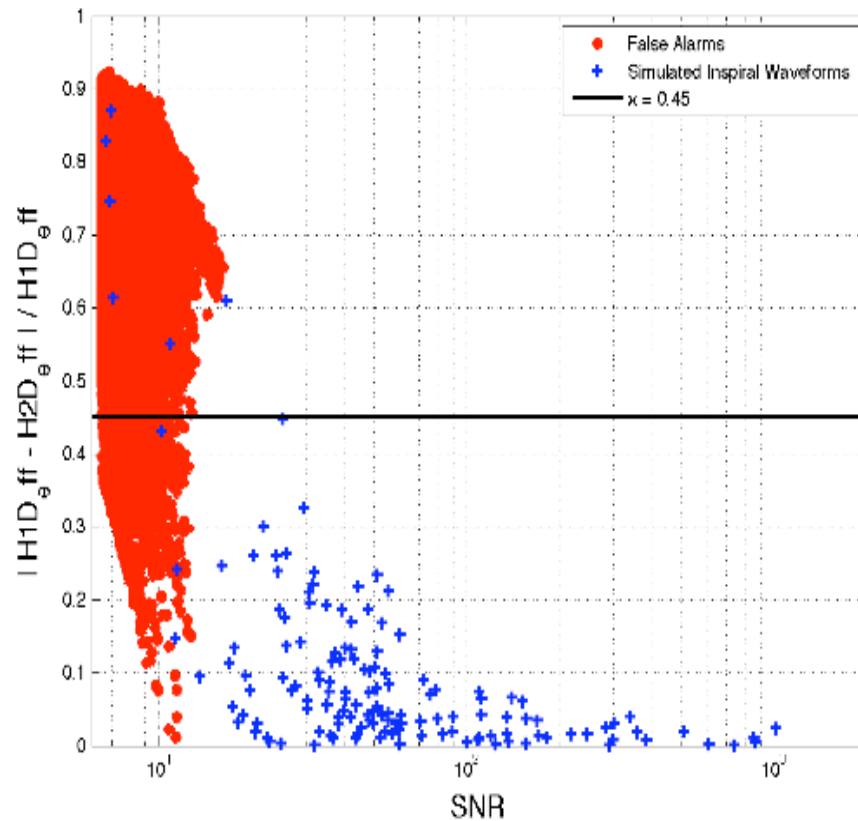




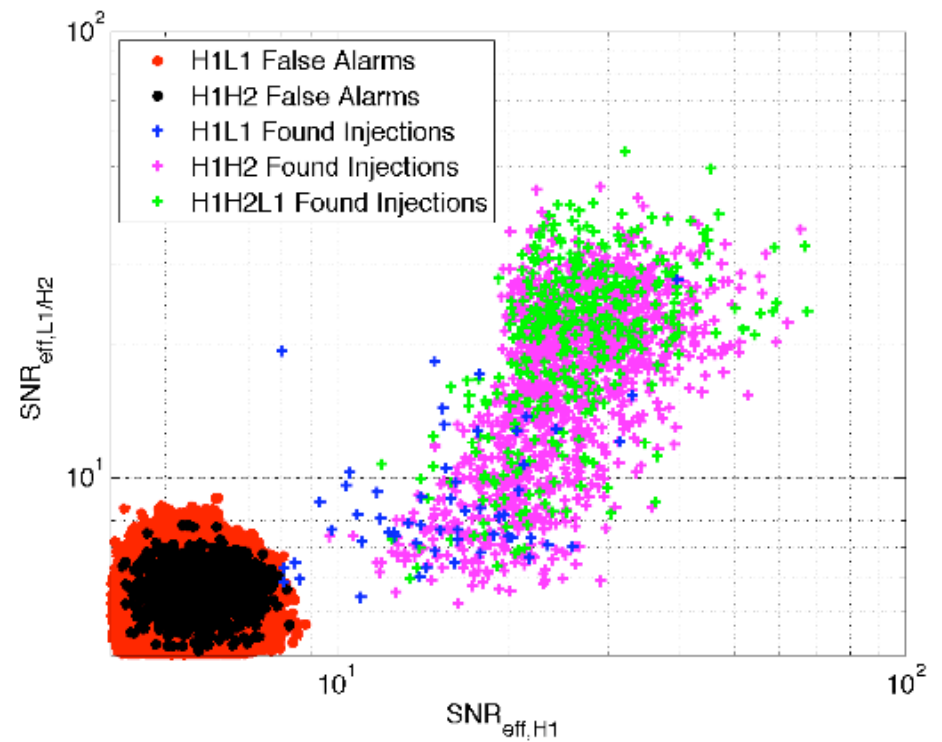
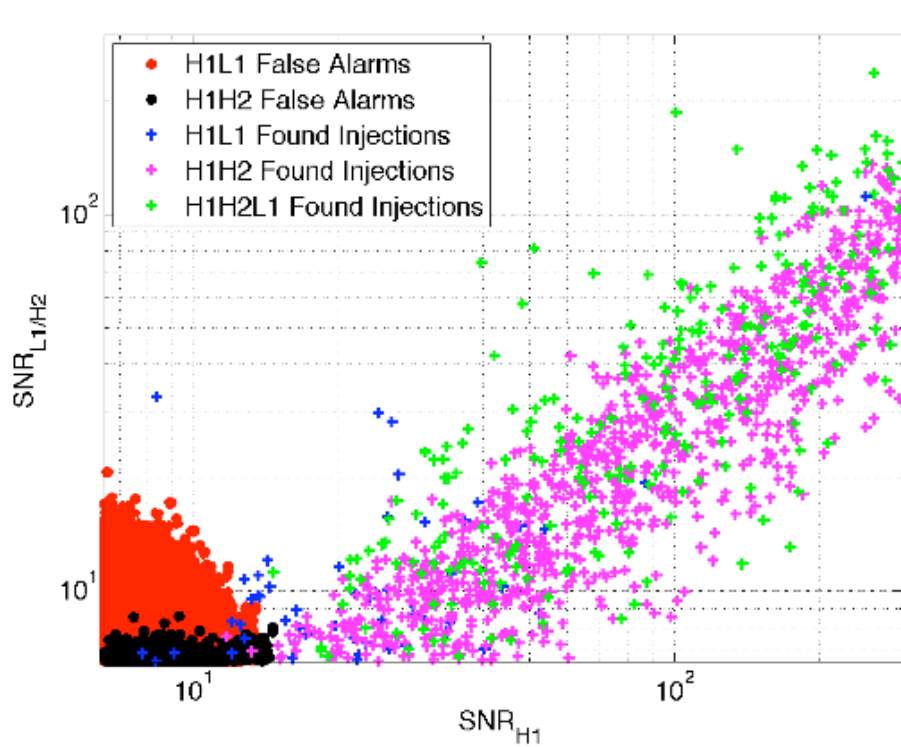
Pipeline



S3 PBH - Effective Distance Cut



Effective SNR - S3 PBH





Tuning - S3PBH

