



Beam Profile Optimization for Thermal Noise Reduction in Advanced IFOs: Lower Bounds, Margins of Progress and Degrees of Freedom

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Outlook



- Thermal Noise Components vs. Beam Profile
 a general formula;
- Absolute Lower Bounds Variational Solutions hard-clipped beams (unphysical);
- Finite Spatial Bandwidth vs. Diffraction Losses: degrees of freedom and effective dimension;
- A More Realistic Bound (via rLSP Expansion);
- Conclusions

Thermal Noise PSD A General Formula



[G. Lovelace, ArXiv:gr-qc/0610041 (2007) R. O'Shaughnessy, CQG 23 (2006) 7627]

rco

$$S = C \int_0^\infty \kappa^{q+1} \left\{ \mathcal{H} \left[|\Phi|^2 \right] (\kappa) \right\}^2 d\kappa, \quad q = \begin{cases} 0\\ 1\\ -1 \end{cases}$$

Coating (Brownian & Thermoelastic) Substrate Brownian (SiO₂)

Substrate Thermoelastic (Al_2O_3)

$$\mathcal{H}[F](\xi) \equiv \int_{0}^{\infty} F(\zeta) J_{0}(\xi\zeta) \zeta d\zeta \qquad \text{(Hankel transform)}$$
$$\Phi(r) \Big|^{2} \equiv \text{ beam intensity distribution at mirror}$$

beam intensity distribution at mirror

- axisymmetric field distribution Assumptions: - infinite (thick) test-mass - low frequency limit

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[A.E. Siegman, Lasers, Univ. Sci. Books, Mill Valley, US, 1998]

$$\gamma \Phi(r) = \int_0^a K(r, r') \Phi(r') r' dr' \quad \text{(integral eq., eigenvalue problem)}$$
$$K(r, r') = \frac{\imath k}{L} J_0\left(\frac{krr'}{L}\right) \exp\left\{\imath k \left[-L + h(r) + h(r') - \frac{(r^2 + r'^2)}{2L}\right]\right\}$$

 $h(r) \equiv$ mirror profile (departure from flatness)

 $a \equiv \text{mirror radius}$

 $L \equiv$ cavity length; $k = 2\pi / \lambda \equiv$ wavenumber

Mapping between : a mirror profile h(r)a set of eigenstates $\Omega[h] = \{\gamma_n, \Phi_n\}$



Diffraction Loss Constraint



Light spillover (diffraction) beyond mirror should be limited:

$$\mathcal{L}[\Phi] \equiv \int_{a}^{\infty} |\Phi(r)|^2 \, r dr \, \leq \mathcal{L}_T$$

It's always possible to make (will be assumed throughout)

$$\int_{0}^{\infty} \left| \Phi(r) \right|^{2} r dr = 1$$

so as to rewrite the diffraction loss constraint as

$$1 - |\gamma|^2 \le \mathcal{L}_T$$

(selects diffraction-loss admissibile eigenstates)



Minimum Noise Mirror (Beam) Profile



Formal Mirror Optimization Procedure

- Assume suitable (e.g., C^{∞}) functional class Λ for h(r);
- Denote as $\Omega_{
 m c}[h]$ the subset of the eigenstate set $\Omega[h]$: $1-|\gamma|^2 \leq \mathcal{L}_T$
- Find $h^* \in \Lambda$ such that : $\min_{\substack{\phi \in \Omega_c[h^*]}} S[\phi] \leq \min_{\substack{\phi \in \Omega_c[h]}} S[\phi], \quad \forall h \in \Lambda : h \neq h^*,$





- For most h(r), the field integral equation can only be attacked numerically → need to parameterize sought function h(r) in terms of a finite number of unknowns
 ↓ ("best" (minimum size) representation ? size of problem ?
- Numerical solution may be *hard to obtain* due to (paramterization dependent) problem's ill-posedness and/or non-convexity (*robust* optimization algos required).
- "Exact" solution could be *technologically unfeasible* .



Available Results & Research Trends



Reference solution: Gaussian Beams (GB);

Mesa-Beams (MB) (Mexican Hat (MH) mirrors) [E. D'Ambrosio, PRD67 (2003) 102004, etc.];



Higher Order Gauss-Laguerre Modes (HOGL) (keep std. mirrors; larger a/w; excitation issues) [B. Mours et al., CQG 23 (2006) 5777]



Hyperboloidal-Beams and related representations (mitigate tilt instability affecting nearly flat MH mirror cavities) [M. Bondarescu and K. Thorne, PRD74 (2006) 082003; V.Galdi et al., PRD73 (2006) 127101]

• Infinite-radius mirror eigenstates used throughout in computing diffraction losses (mirror clipping approximation);

Scalings

LIGO





Absolute (Lower) Noise PSD Bounds



- Cope with diffraction loss constraint by forcing $\phi(\overline{r})$ to vanish outside [0,1] (no-diffraction, compact support beams).
- Don't care about field (eigenvalue) equation. Just seek for an *intensity* profile $f(\overline{r}) = |\phi(\overline{r})|^2 \ge 0$ for which PSD is minimum.
- Translates into simple (constrained) variational calculus problems, with unique *exact solutions* [Castaldi et al., 2007]

$$f(\overline{r}) = (q+2) \left[1 - \overline{r}^2 \right]^{q/2}, \quad 1 \le q \le 1, \ q \in \mathbb{Z}, \ 0 \le \overline{r} \le 1$$

yielding:

$$\bar{S}_1^{(min)} = 2^{q+1} \Gamma\left(\frac{q}{2} + 1\right) \Gamma\left(\frac{q}{2} + 2\right)$$



becomes:

$$\int_{0}^{1} \overline{r} d\overline{r} \left(\int_{0}^{\infty} d\overline{\kappa} \overline{\kappa}^{q+1} J_{0}(\overline{\kappa}\overline{r}) \int_{0}^{1} \overline{r}' d\overline{r}' f(\overline{r}') J_{0}(\overline{\kappa}\overline{r}') - \mu \right) \xi(\overline{r}) = 0, \quad 0 \le \overline{r} \le 1$$
becomes:

$$\int_{0}^{1} d\overline{\kappa} \overline{\kappa}^{q/2} J_{q/2+1}(\overline{\kappa}) J_{0}(\overline{\kappa}\overline{r}) = 2^{q/2} \Gamma\left(\frac{q}{2}+1\right), \quad 0 \le \overline{r} \le 1, \quad -1 \le q \le 1$$
becomes:

$$\int_{0}^{\infty} d\overline{\kappa} \overline{\kappa}^{q/2} J_{q/2+1}(\overline{\kappa}) J_{0}(\overline{\kappa}\overline{r}) = 2^{q/2} \Gamma\left(\frac{q}{2}+1\right), \quad 0 \le \overline{r} \le 1, \quad -1 \le q \le 1$$
becomes:

$$\int_{0}^{\infty} d\overline{\kappa} \overline{\kappa}^{q+1} J_{0}(\overline{\kappa}\overline{r}) \int_{0}^{1} \overline{r}' d\overline{r}' (1-\overline{r}'^{2})^{q/2} J_{0}(\overline{\kappa}\overline{r}'), \quad 0 \le \overline{r} \le 1$$
becomes:

$$\int_{0}^{\infty} d\overline{\kappa} \overline{\kappa}^{q+1} J_{0}(\overline{\kappa}\overline{r}) \int_{0}^{1} \overline{r}' d\overline{r}' (1-\overline{r}'^{2})^{q/2} J_{0}(\overline{\kappa}\overline{r}') = 2^{q} \Gamma^{-2}(q/2+1)$$
whence:

$$f = \mu 2^{-q} (1-\overline{r}'^{2})^{q/2} \Gamma^{-2}(q/2+1), \quad \|f\| = 1 \Leftrightarrow \mu = (q+2) 2^{q} \Gamma(q/2+1)$$

$$qed$$

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Absolute (Lower) Noise PSD Bounds, contd.





Remarks/Caveats

Optimal field-intensity profile for coating noises flat as expected; for substrate Noises, *not exactly flat,* and not obvious.

Obtained (scaled) field-intensity profiles yield absolute but likely *loose* lower bounds for the noise PSDs.

The no-diffraction field assumption made is indeed violated by *any* solution of the field equation.

How close can we go to these bounds using physically *admissible* fields ?

IGO Spatial Band-Limitedness of Cavity Field



From the obvious properties:

$$\mathcal{H}_1[f(\bar{r})] = \mathcal{H}[\Pi(\bar{r})f(\bar{r})], \ \Pi(\bar{r}) = \begin{cases} 1, \ 0 \le \bar{r} \le 1\\ 0, \ \text{elsewhere} \end{cases}, \ \mathcal{H}[\mathcal{H}[f]] = f \end{cases}$$

by applying $\,\mathcal{H}\,$ operator to both sides of field (eigenvalue) equation we obtain

$$\mathcal{H}\left[\phi\exp\left(iV\right)\right]\left(\pi N_{D}\bar{r}\right) = i\frac{\pi N_{D}}{\bar{\gamma}}\Pi\left(\bar{r}\right)\exp\left[-iV\left(\bar{r}\right)\right]\phi\left(\bar{r}\right)$$

The Hankel transform (wavenumber spectrum) of $\exp[\iota V(\overline{r})]\phi(\overline{r})$ has *compact support*, vanishing outside $[0, \pi N_D]$. Accordingly $\exp[\iota V(\overline{r})]\phi(\overline{r})$, and hence ϕ , *cannot* vanish identically for $\overline{r} > 1$.



The (Radial) Slepian Landau Pollack Basis



• The (real valued) eigenstates of

$$\bar{\eta}\varphi(\bar{r}) = i\pi N_D \mathcal{H}_1\left[\varphi\right] \left(\pi N_D \bar{r}\right)$$

(modal fields of a confocal-spherical finite-mirror FP cavity) play *a special role* (Slepian-Landau-Pollack radial wavefunctions).

- Among *all* L^2 bases, they allow to approximate *any* exact solution of the field equations (corresponding to an arbitrary mirror profile), using the minimum number N_{ε} of terms for *any* prescribed L^2 error ε (*minimum-redundant basis*).
- Technically, N_{ε} is referred to as the number of *degrees of freedom* of our cavity fields at the *resolution level* ε .

[D. Slepian et al., Bell System Tech. Journal, 40 (1961) 43 and 65; ibid. 41 (1962) 1295]



Peculiar Properties of rSLP Eigenstates





Peculiar Properties of rSLP Eigenstates, contd.



infinite-mirror (Gauss-Laguerre) modes also shown dashed



SLP Diffraction Loss Constraint



 N_T

- Let sought field be expanded in terms of rSLP modes : $\phi(\bar{r}) = \sum_{n=1}^{1} b_n \varphi_n(\bar{r})$
- Diffraction loss constraint rephrases into (in view of double-orthogonality)

$$\mathcal{L}[\phi] = \int_{1}^{\infty} \bar{r} d\bar{r} |\phi(\bar{r})|^{2} = \sum_{n=1}^{N_{T}} (1 - |\eta_{n}|^{2}) |b_{n}|^{2} \leq \frac{\max}{n=1,2,\dots,N_{T}} (1 - |\eta_{n}|^{2}) \sum_{n=1}^{N_{T}} |b_{n}|^{2} \leq (1 - |\eta_{N_{T}}|^{2})$$

last inequality follows from : i) the fact that $\{|\eta_n|\}$ is monotonic - decreasing; ii) Parseval theorem; iii) the fact that $||\phi|| = 1$.

• The diffraction loss constraint dictates the *effective dimension* $N_T \sim N_D$ of our optimization problem (number of unknown coefficients in the rSLP modal expansion of the cavity field)



rLSP Approximants of Variational Solutions



- As a natural next step, we construct L^2 approximants of the (unphysical) fields obtained from minimal-noise variational-solutions, by suitable linear combinations of the lowest N_D rLSP-eigenstates.
- At variance of the compact-spatial-support fields deduced from the variational solutions, these fields will satisfy *both* the diffraction-loss constraint *and* the compact-spectral-support condition.
- However, there is NO guarantee that such fields may be decently approximated by the lowest (or any other pure) eigenstate corresponding to *some* mirror profile.

LIGO rLSP Approximants of Variational Solutions, contd.



LIGO How Far did We Reach ?



	\overline{S}_{\min}	$\overline{S}_{\scriptscriptstyle SLP}$ / $\overline{S}_{\scriptscriptstyle \min}$	$\overline{S}_{M\!B}$ / \overline{S}_{\min}	\overline{S}_{GB} / \overline{S}_{\min}
Substrate (Br)	1.5708	1.145	2.044	2.97
Coating (Br+TE)	2	1.225	3.227	6.92
Substrate (TE)	4.712	1.381	4.455	13.66

 $a = 16cm (N_D = 14);$ $\mathcal{L}_T = 1ppm;$ $w_{MB} = (N_D)^{-1/2}$ (minimum spreading)

...sensible possible improvement, e.g. by a factor 2.65 for the coating noise!



Conclusions



- Absolute noise lower bounds, corresponding to compactspatial support intensity profiles have been identified, together with the intensity profiles themselves, via a variational approach;
- The effective dimension of the optimization problem has been related to the diffraction-loss constraint and found to be of the order of $N_D = 2a^2/\lambda L$;
- A field solution (superposition of N_D rLSP-modes) coping w. *both* the diffraction-loss bound, *and* the compact-spectralsupport property of eigenstates has been shown to get *fairly close* to the absolute noise bounds, for $N_D = 2a^2/\lambda L$ *sufficiently large* (until the infinitely thick mirror approximation breaks down);





- While, there is *NO guarantee* that such fields may be the eigenstates of *some* mirror profile, the gap in terms of noise levels between the *best* currently available solutions (MB, HOGL) and the above lower bounds is *pretty large* (compared, in particular, to the expected infinitely thick mirror assumptions related inaccuracy).
- This *suggests* that there's margin for further *substantial* noise reduction through mirror/beam optimization, and that the related conceptual/computational research effort is *worth*.



What is a *Good* "Optimized" Mirror ?



- "Optimized" Mirror (and superimposed coating) must be technologically feasible (slope constraints, tolerances, etc.);
- "Optimized" field (eigenstate) must be easy to launch (...ideally, a *dominant* mode...)
- "Optimized" field must be robust w.r.t. cavity (& coupling) drifts/tolerances;
- "Optimized" field must yield noise levels as close as possible to lower bounds stemming from (competing) diffraction loss constraint and compact support property of eigenstates.

LIGO A Possible Constructive Approach Being Explored



- Parameterize mirror profile *consistently* (to prevent ill-conditioning) to effective dimension $N_T \sim N_D = 2a^2/L\lambda$ of optimization problem;
- Derive lowest order eigenstate(s) using an *efficient* (fast and reasonably accurate) algorithm, e.g., Nystrom [J. Comp. Phys. 146 (1998) 627], or perhaps Donsker-Kac [J. Res. NBS 44 (1954) 551; V. Galdi et al., Electromagnetics, 18 (1998) 367];
- Use *robust* (e.g., *genetic*) *optimization* engine to tweak unknown mirror parameters (e.g., polynomial coefficients) to bring noise to a minimum, while coping with diffraction-loss *and suitable technological* constraints.