



### A RTP crystal electro-optic modulator for next generation gravitational wave detector

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Part of the technical concerns for Advanced LIGO

- Potential 'thermal lensing' headaches due to the application of 180W laser
- Stringent requirement on the laser amplitude and frequency noise

Demanding requirements on EOMs:

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- Ultra-low optical absorption on the 1064nm laser
- Negligible residual amplitude and phase noise associated with the phase modulation



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# RTP crystal EOM

A novel RTP (RbTiOPO4) crystal EOM is developed for Advanced LIGO

- Optical absorption coefficient < 1000 ppm/cm (upper limit according to our measurement)
- Maximum power loss of 00-mode ~ 1.4%

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**RFAM** 





LIGO



Measured RFAM produced by RTP EOMs under high power laser heating  $\Delta I_{\Omega} / I_{DC} \sim 10e-5$ 

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- EOM imposes the additional noise on the transmitted light
  - Creates intensity noise of both the carrier and the sidebands
  - Introduces sidebands phase noise

(The phase fluctuation between the carrier and the sidebands causes laser frequency noise via the feed back loop)



# LIGO Measurement of the intrinsic EOM



- Phase lock two lasers with a frequency offset
- Measure the beat signals Carrier-Carrier (C-C), Carrier-Sidebands (C-S)



# Amplitude noise & phase noise

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- Noise in two beat signals (C-C, C-S) in common mode
  - Beam jitter between two laser beams
  - Laser intensity noise
- Noise in two beat signals in differential mode (amplitude & phase noise)
  - produced by the EOM

Common-mode noise

rejection analysis

Amplitude

 $fft \left\{ \begin{array}{ll} amplitude & of \quad C-S \\ \hline amplitude & of \quad C-C \end{array} \right\}$ 

Phase

 $fft \{ \phi_{C-S} - \phi_{C-C} \}$ 



# Data analysis





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# Future improvement

Modulation index

$$m = \frac{\pi L}{\lambda} r_{33} n_z^3 \frac{V_z}{d}$$

Phase retardation

 $\Delta \phi = \frac{2\pi}{\lambda} (n_z - n_y) L$ Temperature stabilization is needed

$$L = L_0 + \kappa \Delta T \qquad n = n_0 + \frac{dn}{dT} \Delta T$$

X Z

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Use wedge crystal to eliminate the cavity effect









- Solving the noise problems associated with EOMs is a challenging task !
- Current estimation of the requirement on the amplitude and the phase noise:
  - Laser intensity noise ~  $2 \times 10^{-9} / \sqrt{\text{Hz}}$
  - Sideband phase noise ~  $10^{-11} / \sqrt{Hz}$  at 100 Hz

Based on the calculation in (K. Somiya, Y. Chen, S. Kawamura, and N. Mio, PRD, 2006)