

LIGO Caltech – Virgo LAL Orsay meeting

August 28th, 2006

Status of the collaboration

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Caltech - August 28th, 2006

LIGO-G060491-00-R

Status of the e2e simulation

e2e simulation:

- **40mInLockState** package → **almost done!**
- **40mInLockSeism** package → **under investigation**
- **AdVirgoFP** package → **done! Can be improved**
- **AdVirgo** package → **under investigation**

Status of the 40m activities

40m commissioning:

- note on mirror velocity reconstruction → **almost finished**
- note on the Gouy telescope
for the MC WFSs → **In progress**
- note on the calibration of the GW
signal channel → **In progress**

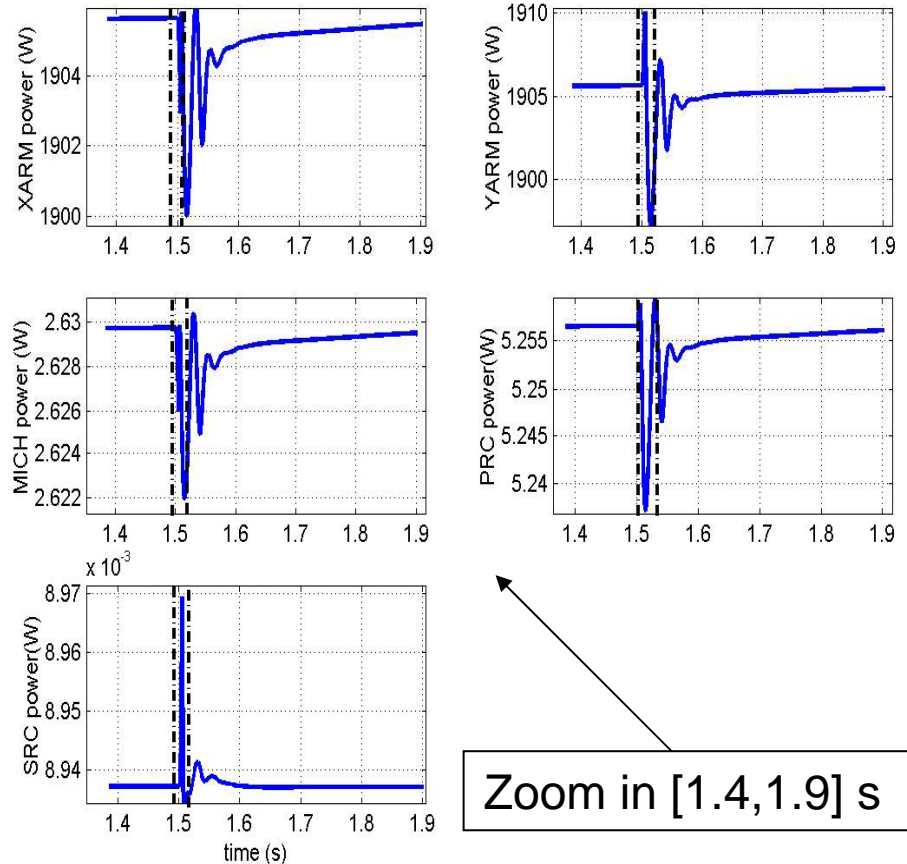
40minLockState package

Elimination of the offset on the CARM
(POX and POY error signals combination)
thanks to the new demod angles determined
dithering each end mirror independently.

Lock stable for 7s

Control system active @ 20ms
Radiation Pressure Force @ 0.1s
Seismic noise injected @ 1s
Open DARM, CARM loops @ 1.5s
Close DARM, CARM loops @ 1.505s

Cavities power



Caltech -

Zoom in [1.4, 1.9] s

40mInLockState package (2)

Controlled d.o.f. :

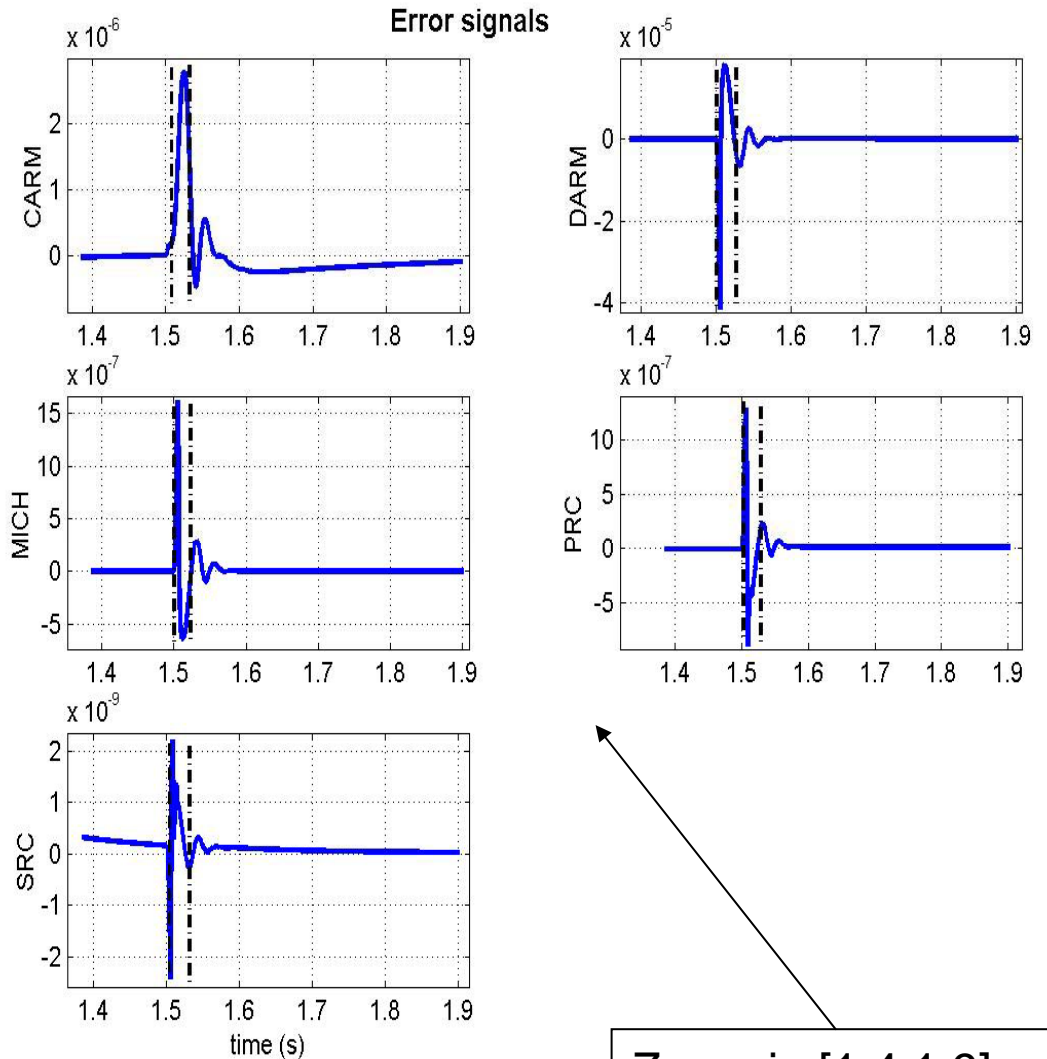
DARM : AP166Q

CARM : POX33I+POY33I

MICH : AP33I

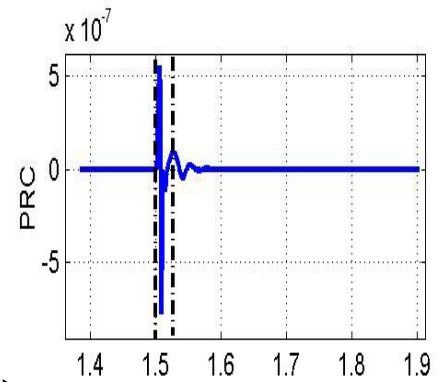
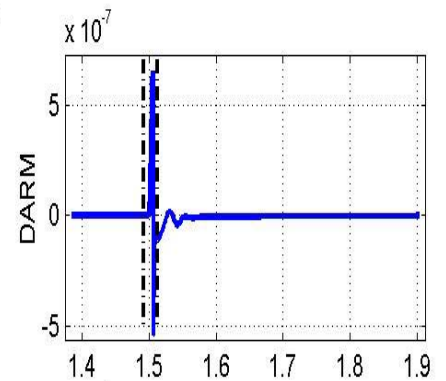
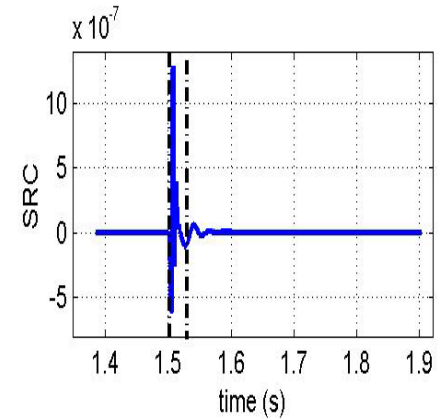
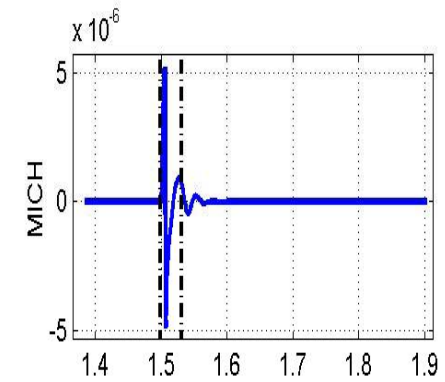
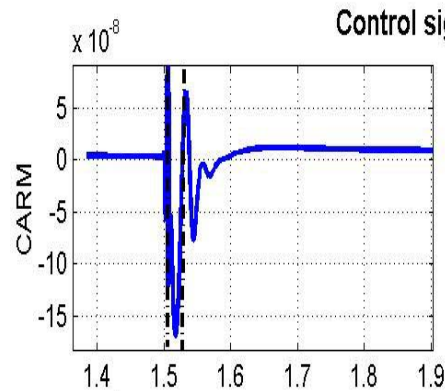
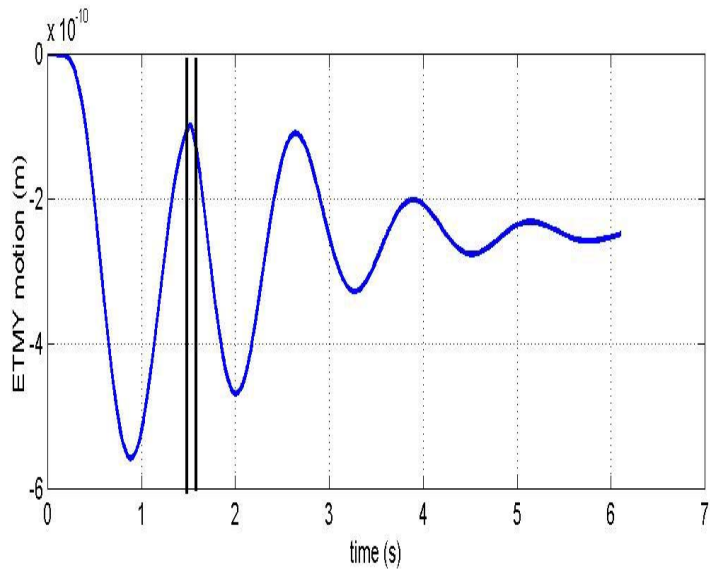
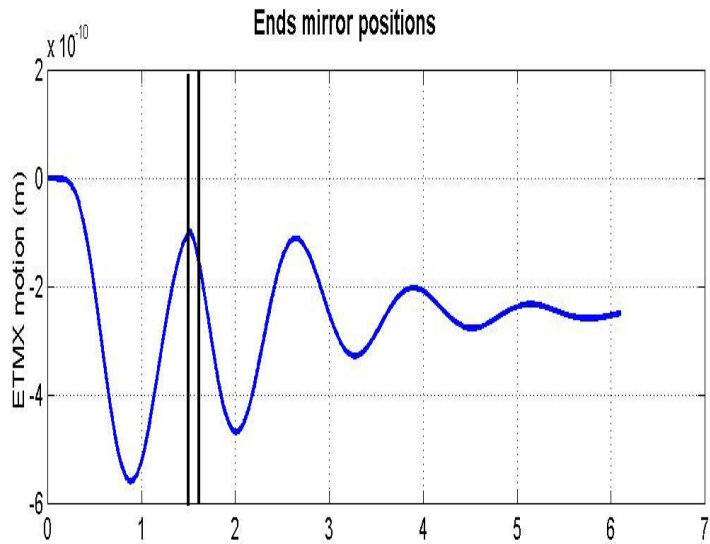
PRC : SP133I

SRC : POX199I



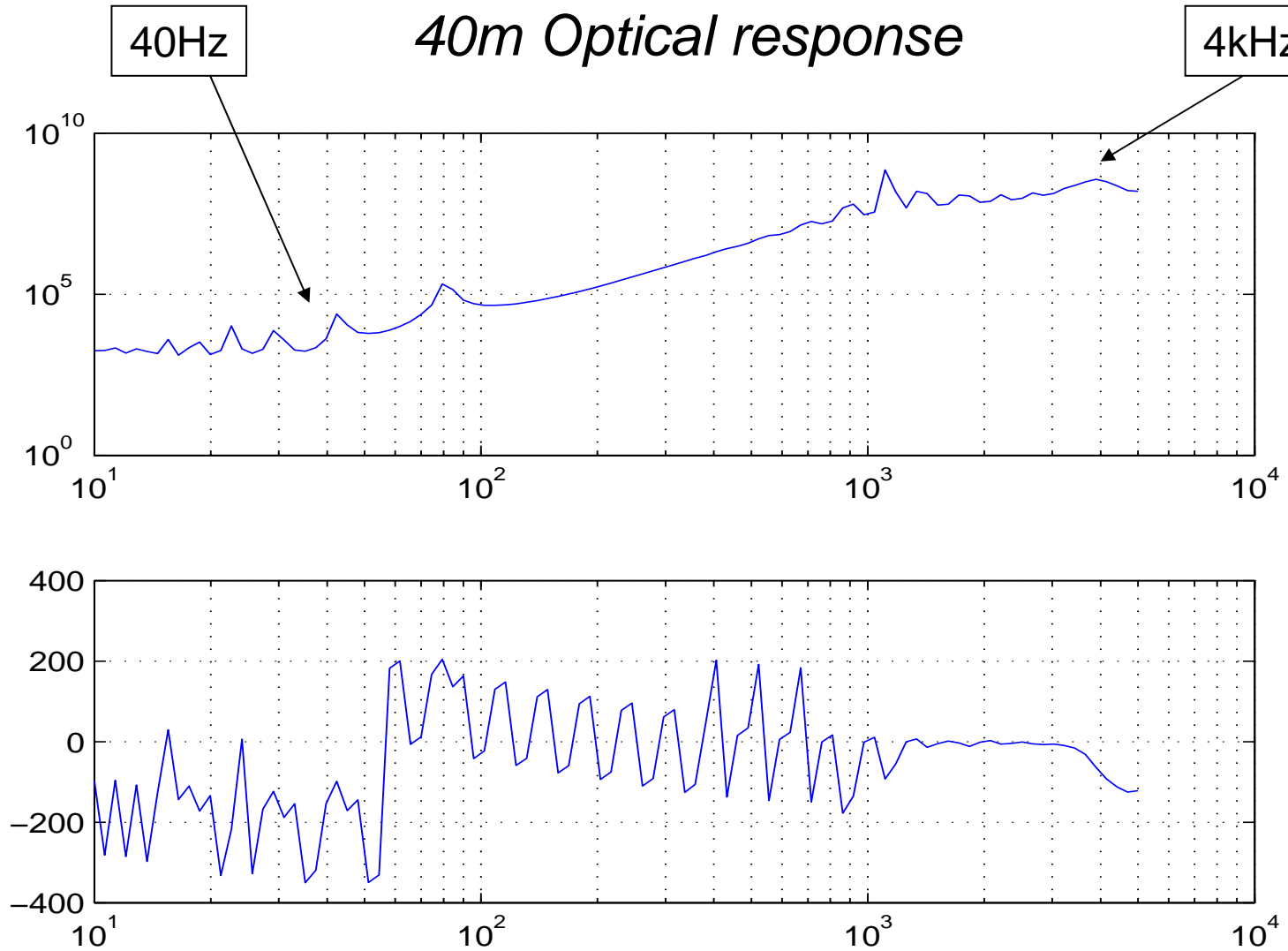
Zoom in [1.4, 1.9] s

40mInLockState package (3)



Zoom in [1.4, 1.9] s

40mInLockState package (4)



AdVirgoFP package

Power = 20W

Demod freq = 6.26 MHz

Arm length = 3km

$T_{itm} = T_{etm} = 0.005$

$L_{itm} = L_{etm} = 50\text{ppm}$

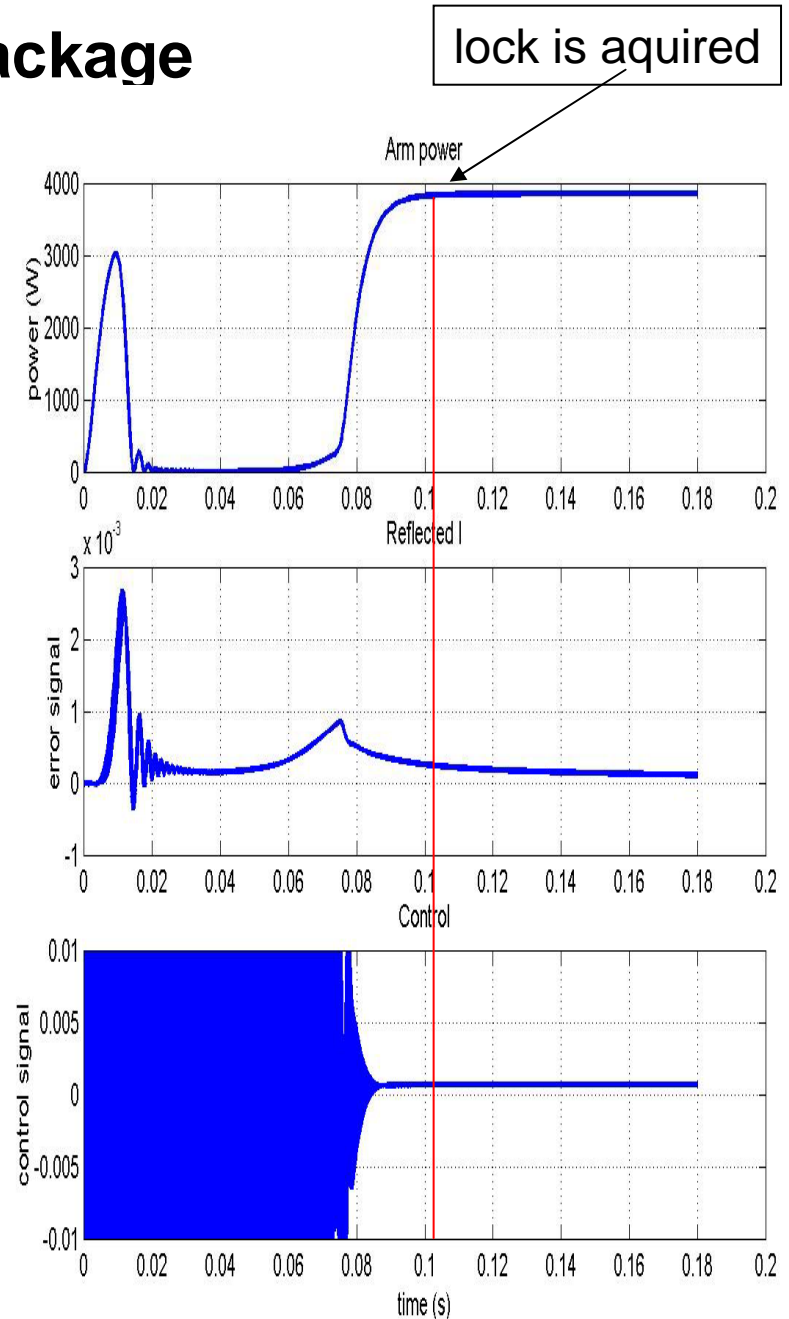
Radiation pressure on ITM and ETM

Cavity controlled with REFL_I

Lock acquired at 0.1s

It is not possible to lock with a laser power of more than 20 W

Strategies to increase the power after acquiring the lock at low power are under investigation.



AdVirgoFP package (2)

Control system:

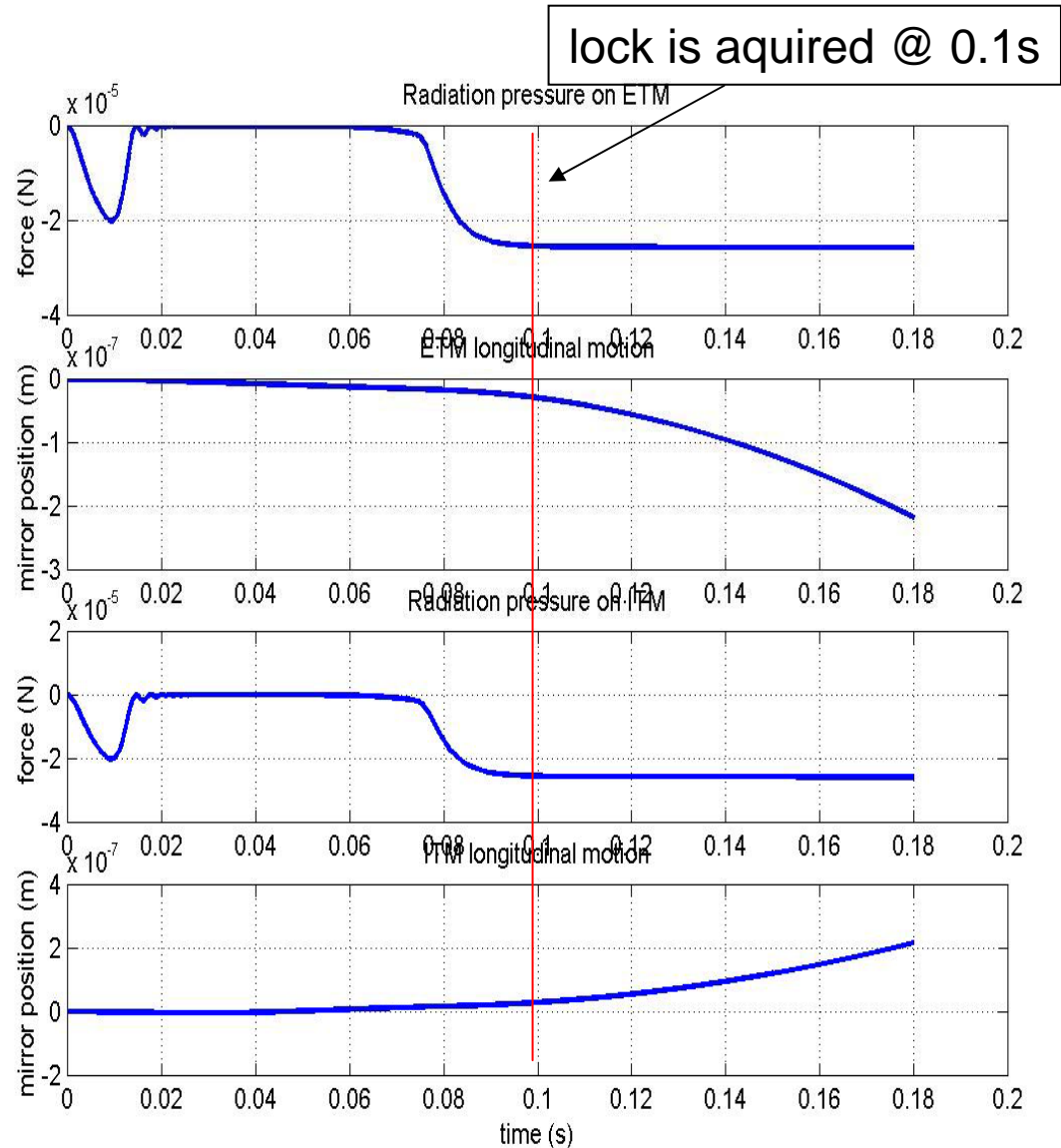
The cavity can be controlled also using the TRANS_I error signal when it is not too small

Suspension:

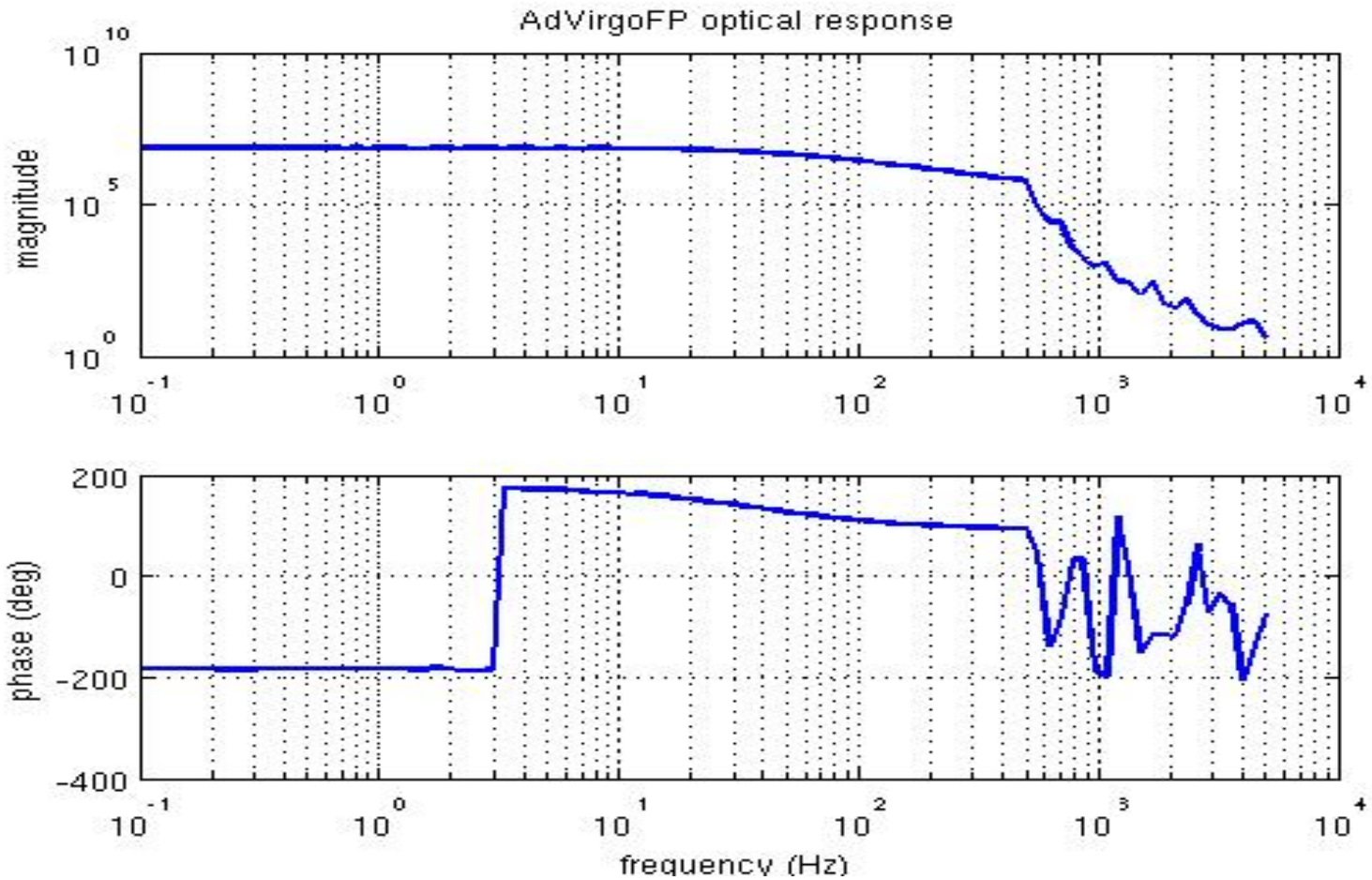
simple pendulum up to now

Seismic Noise:

10^{-6} m/root(Hz)



AdVirgoFP package (3)

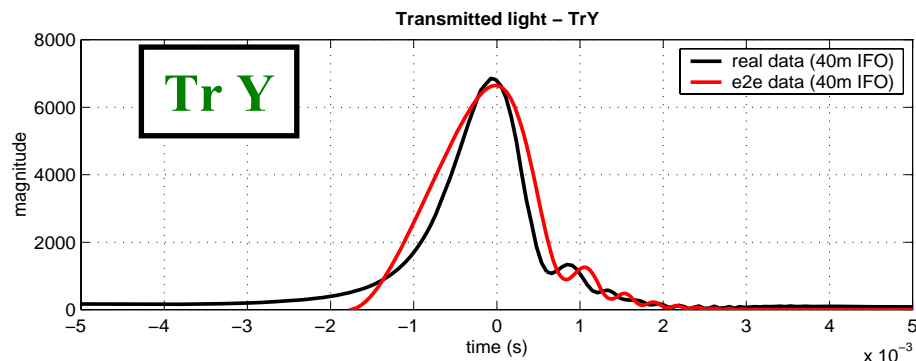
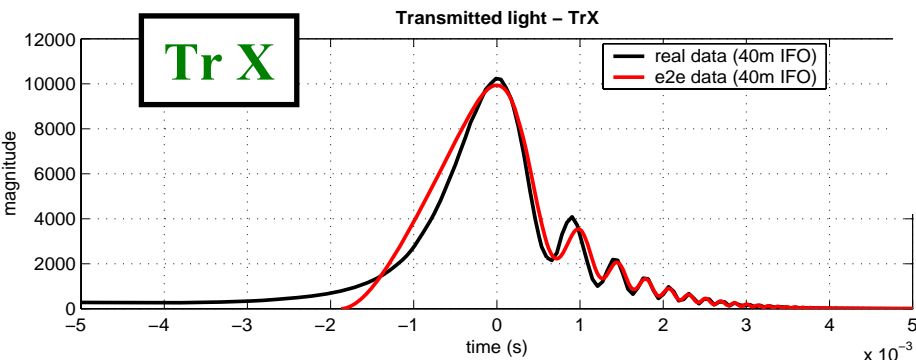
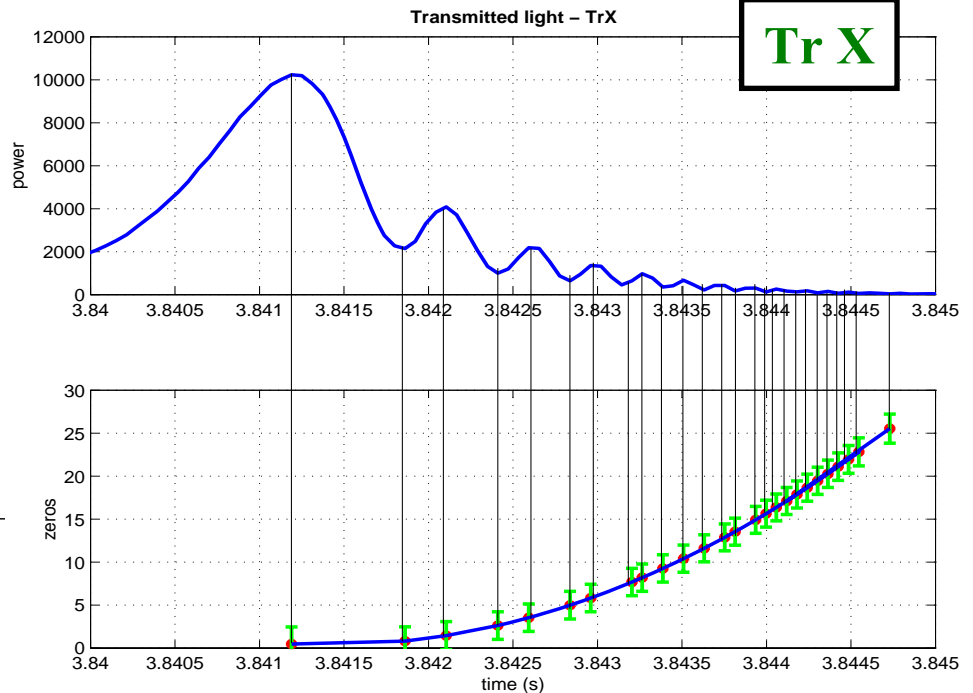


Mirror velocities reconstruction

Real data have been used to estimate relative mirror velocity for both the arms:

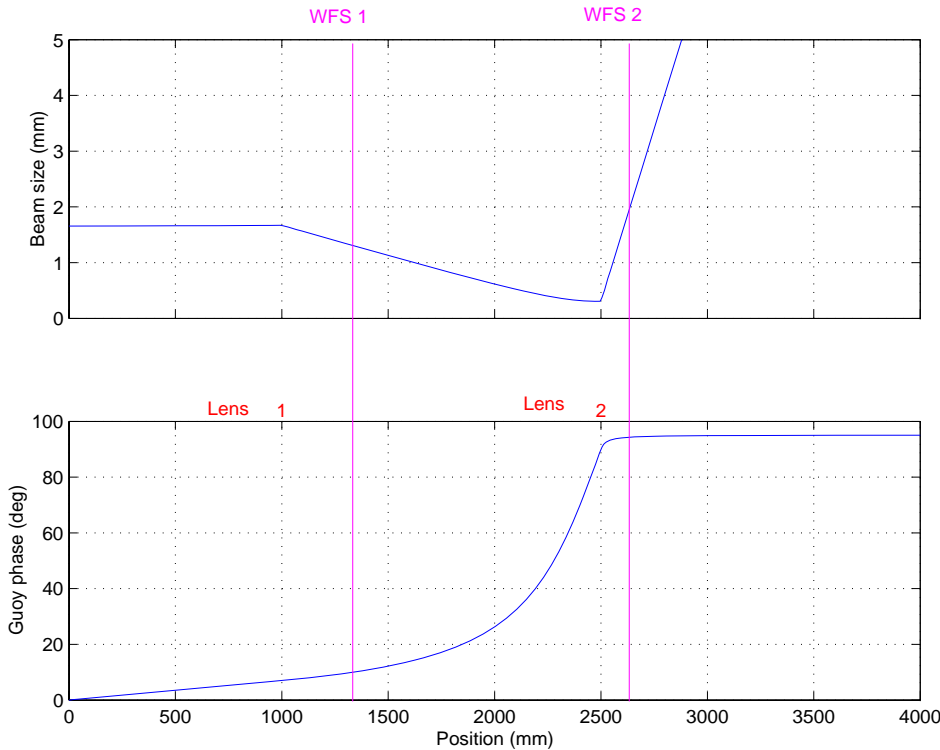
$$V_{xarm} = (0.35 \pm 0.13) \mu\text{m/s}$$

$$V_{yarm} = (0.26 \pm 0.13) \mu\text{m/s}$$



Comparison between real data (black) and e2e simulated data (red) of the transmitted light for both the arms: the mirror velocities used in e2e simulation are the values obtained fitting the real data

Design of a new Gouy telescope for the MC WFS



Beam code to determine the lenses and the WFS positions

WFS output matrix after diagonalization

CIT 40m C1100_WFS_BASIS Fri Aug 25 12:57:44 200

Pitch				Yaw			
WFS1	WFS2	Gain		WFS1	WFS2	Gain	
197.700	58.407			-33.451	-46.333		
MC1	1.000	-1.000	-1.000 ▶ -135.289	MC1	-0.073	0.652	-1.000 ▶ 20.937
MC2	0.158	-0.740	-1.000 ▶ 14.458	MC2	-1.000	0.349	-1.000 ▶ -26.217
MC3	-0.910	0.700	-1.000 ▶ 137.655	MC3	0.627	-1.000	-1.000 ▶ -18.071
PZT1	0.000	0.000	1.000 ▶ 0.000	PZT1	0.000	0.000	1.000 ▶ 0.000
PZT2	0.000	0.000	1.000 ▶ 0.000	PZT2	0.000	0.000	1.000 ▶ 0.000

What's next?

- Future of the collaboration on the Signal Recycling
- AdVirgo with or without SR option (white paper)
- Notes
- Questions, Comments, Discussion