



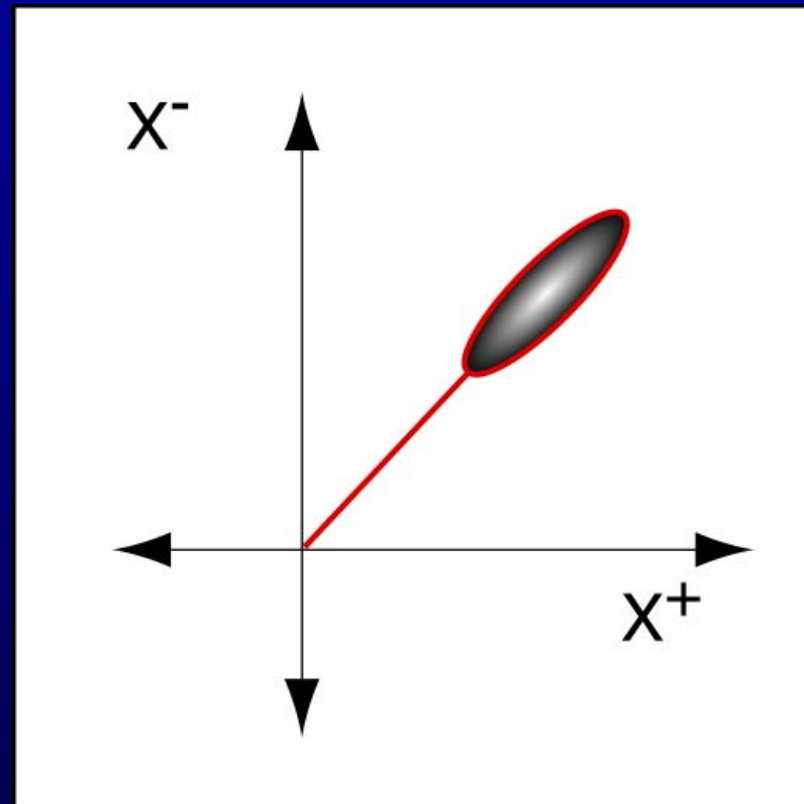
# Recent Developments toward Sub-Quantum-Noise-Limited Gravitational-wave Interferometers

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Aspen  
January 2005

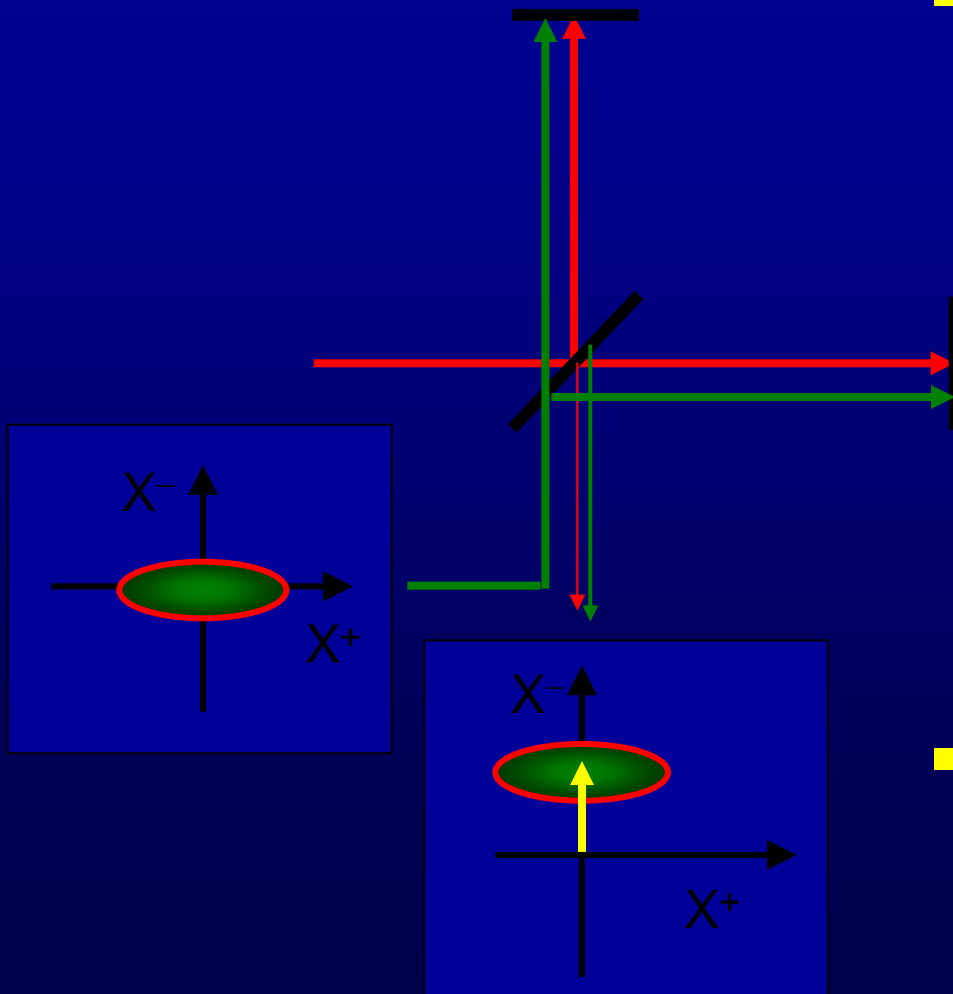
LIGO-G050044-00-R

## Some quantum states of light

- Analogous to the phasor diagram
- Stick  $\rightarrow$  dc term
- Ball  $\rightarrow$  fluctuations
- Common states
  - Coherent state
  - Vacuum state
  - Amplitude squeezed state
  - Phase squeezed state



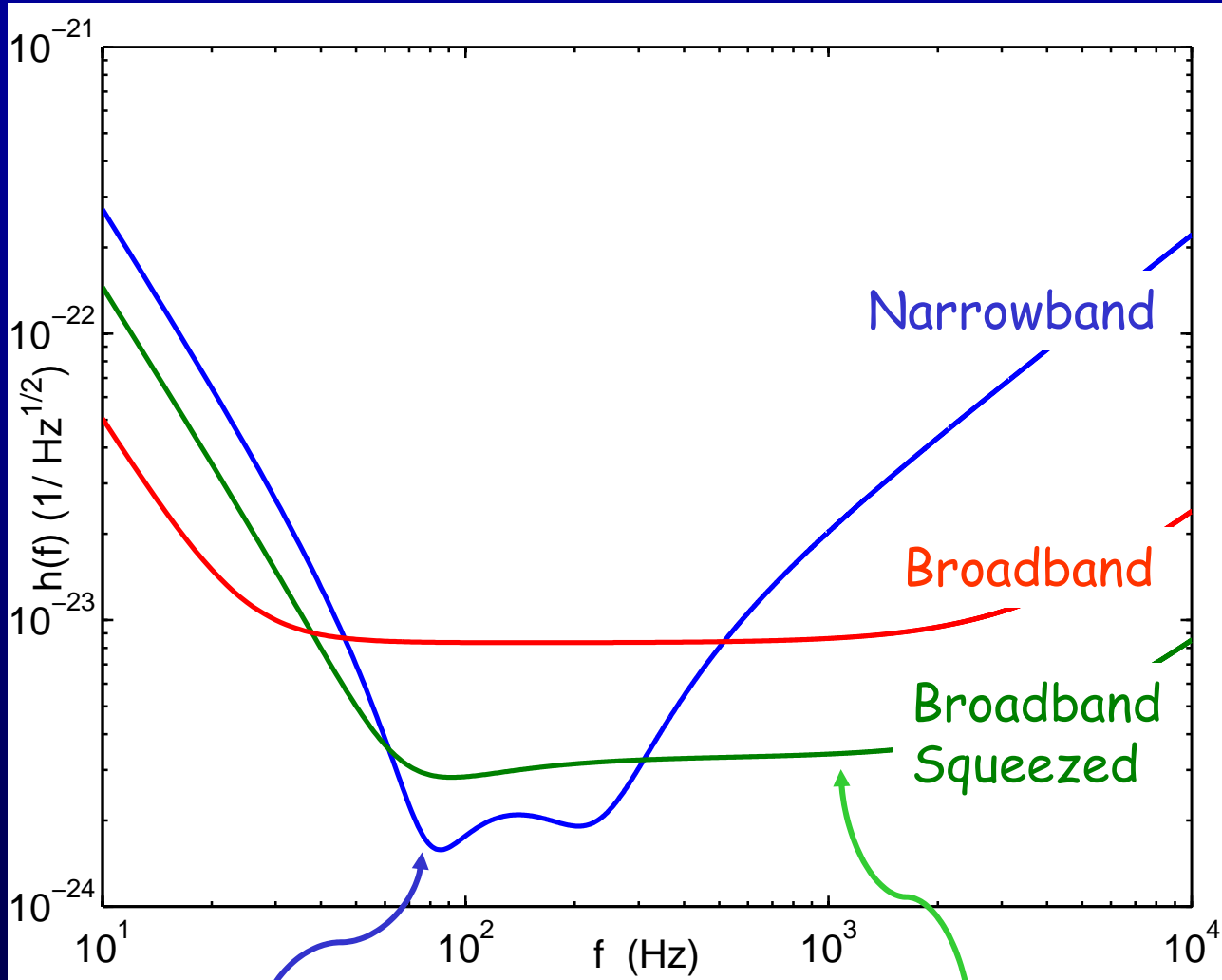
## Squeezed input vacuum state in Michelson Interferometer



- GW signal in the phase quadrature
  - Not true for all interferometer configurations
  - Detuned signal recycled interferometer  $\rightarrow$  GW signal in both quadratures
- Orient squeezed state to reduce noise in phase quadrature

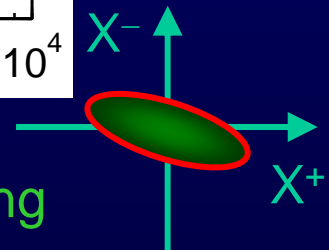


# Sub-quantum-limited interferometer



Quantum correlations

Input squeezing



# Squeezed vacuum

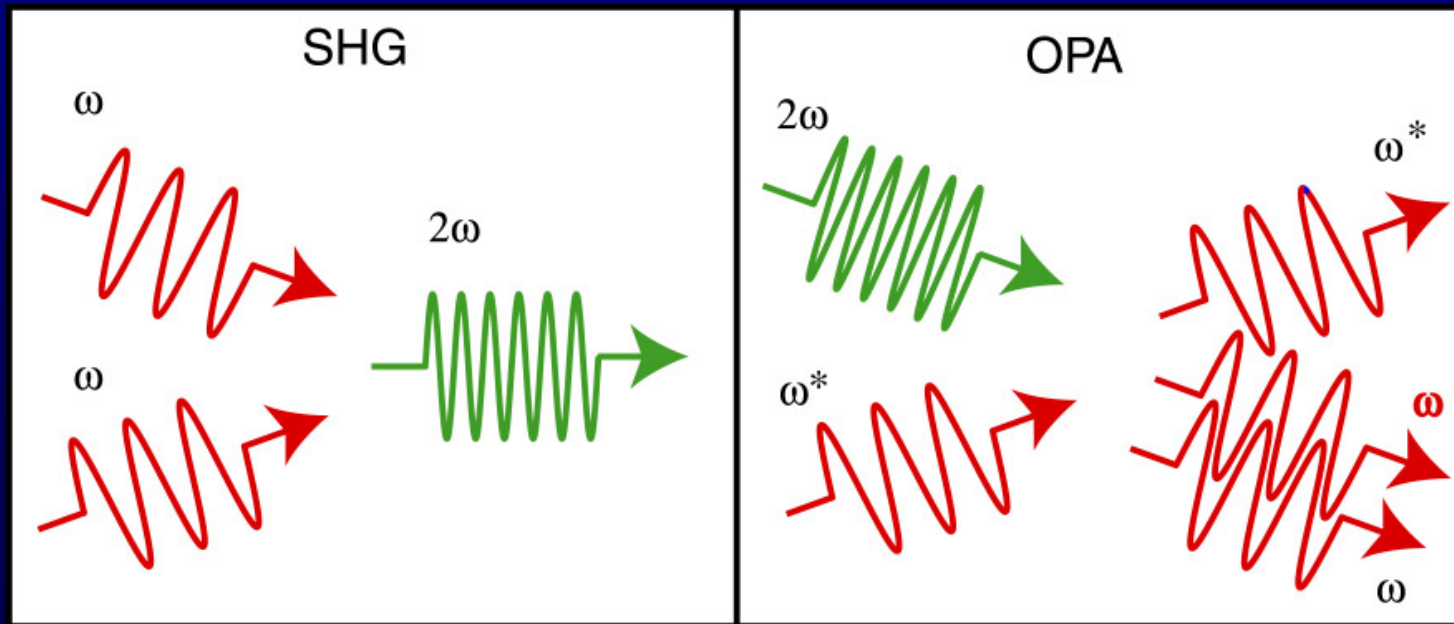
- Requirements
  - Squeezing at low frequencies (within GW band)
  - Frequency-dependent squeeze angle
  - Increased levels of squeezing
- Generation methods
  - Non-linear optical media ( $\chi^{(2)}$  and  $\chi^{(3)}$  non-linearities) ← crystal-based squeezing
  - Radiation pressure effects in interferometers ← ponderomotive squeezing
- Challenges
  - Frequency-dependence → filter cavities
    - Amplitude filters
    - Squeeze angle rotation filters
  - Low-loss optical systems



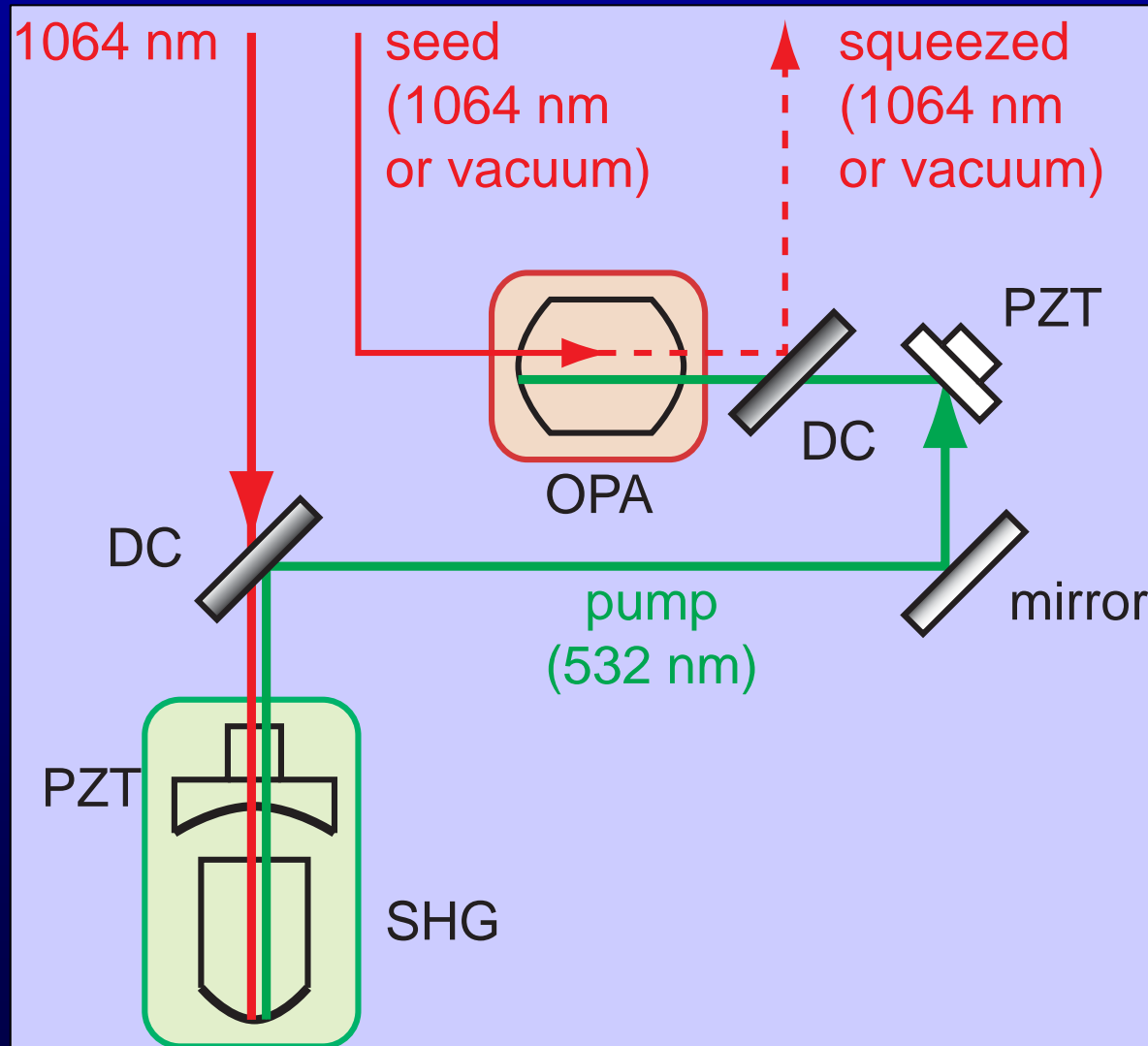
# Squeezing using nonlinear optical media

# Non-linear crystals

- Optical Parametric Amplification (OPA)
- Three (or four) wave mixing
  - Pump (532nm)
  - Seed (1064nm)



# Optical Parametric Oscillator

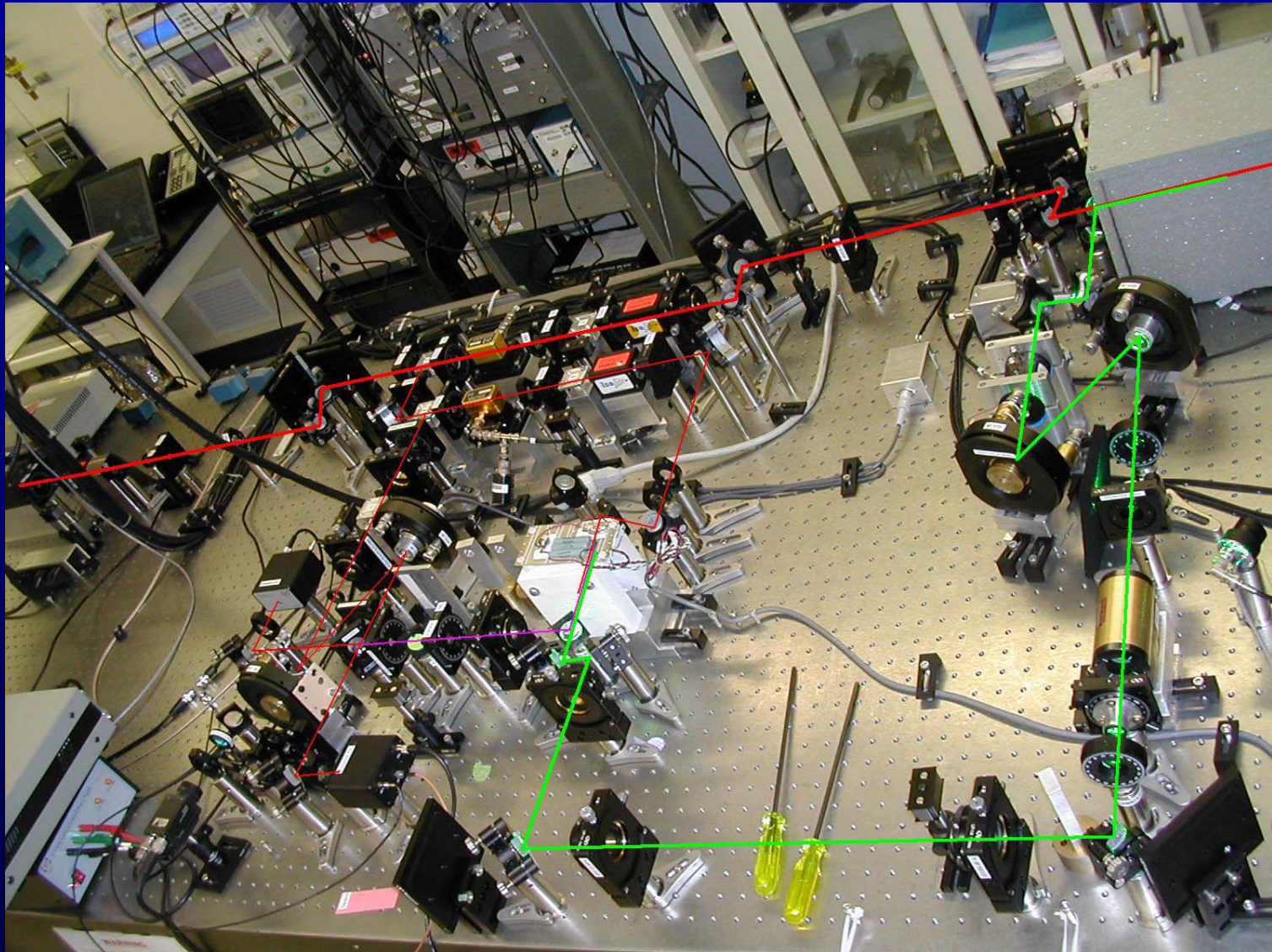




## What's new since last year?

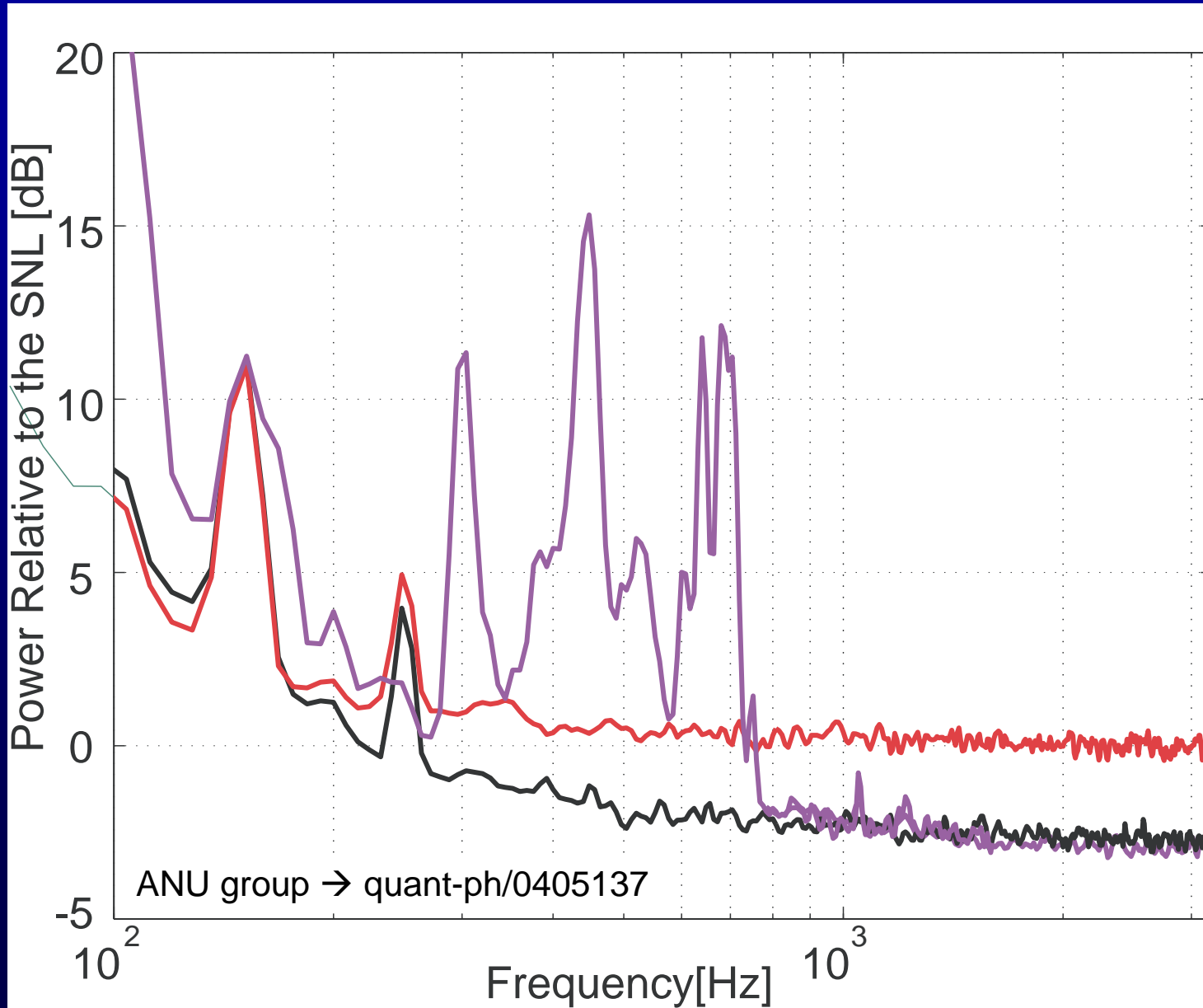
- Squeezing at audio frequencies (ANU, Caltech)
- Next-generation crystals in use (Hannover)
- Testing filter cavities (Hannover, MIT)
- Testing noise couplings (ANU, MIT)
- Detailed calculations of noise budget (ANU, MIT)
  - Photo-thermal noise not a problem
  - Pump noise coupling being considered

# Typical Experimental Setup





# Low frequency squeezing at ANU



## What's next

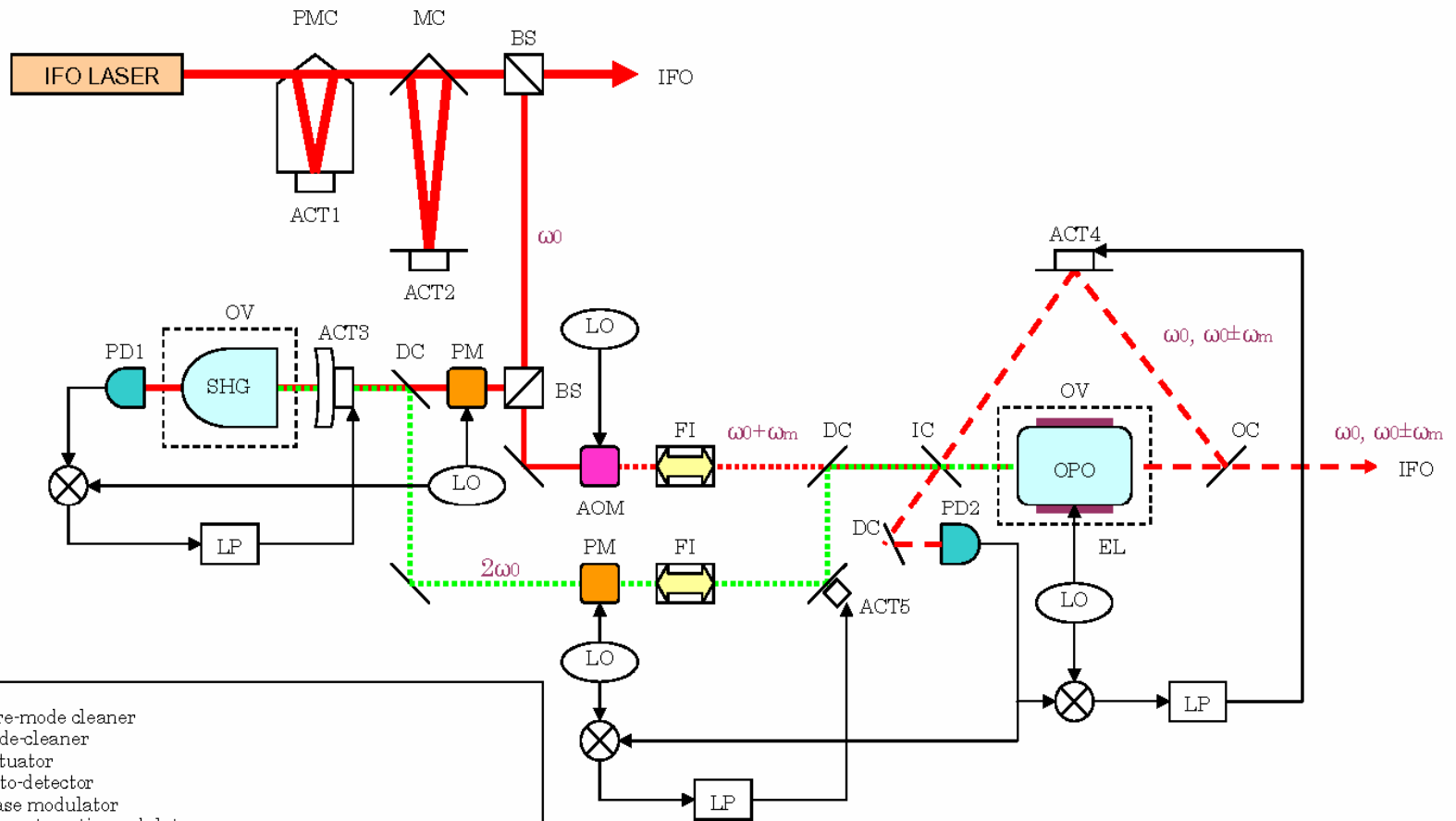
- Ultimate goal

PERFORM A SUSPENDED  
INTERFEROMETER TEST

- Issues to work out

- Coupling into interferometer dark port through output mode cleaner etc
- Error signals for optimum quadrature

# Injected Squeezing into Interferometer



PMC: pre-mode cleaner  
 MC: mode-cleaner  
 ACT: actuator  
 PD: photo-detector  
 PM: phase modulator  
 AOM: acousto-optic modulator  
 LP: low-pass filter  
 DC: dichroic mirror  
 BS: beam-splitter  
 EL: electrode  
 FI: Faraday isolator  
 IFO: interferometer  
 LO: local oscillator  
 OV: oven  
 SHG: second-harmonic generator  
 OPO: optical parametric oscillator  
 IC: input coupler (HR at fundamental, AR at second-harmonic)  
 OC: output coupler (HR at both fundamental and second-harmonic)

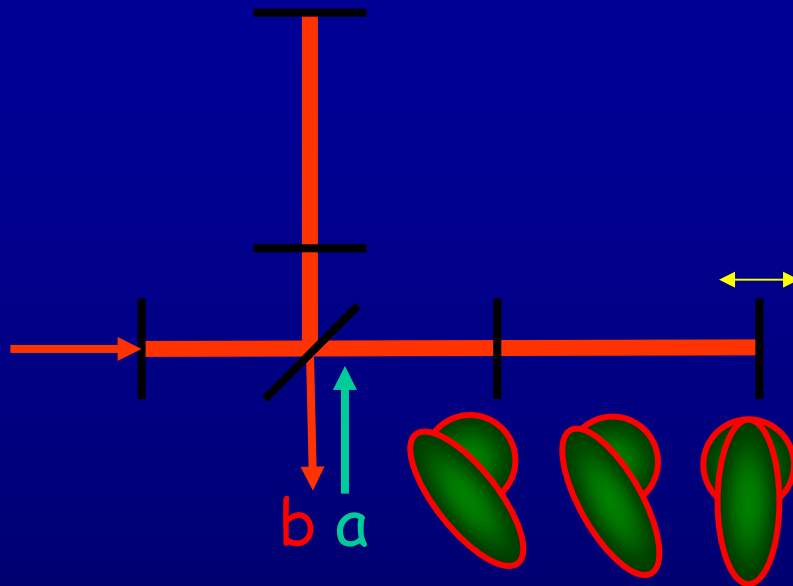
One Example  
 Cavity Length = 30 cm  
 FSR = 500 MHz  
 Cavity Linewidth = 13 MHz

Main Laser: —————  
 Second-Harmonic Laser: — · — · — · —  
 Frequency-Shifted Sub-Carrier: · · · · ·  
 Squeezed Vacuum + FSSC (Signal) + Idler: - - - - -

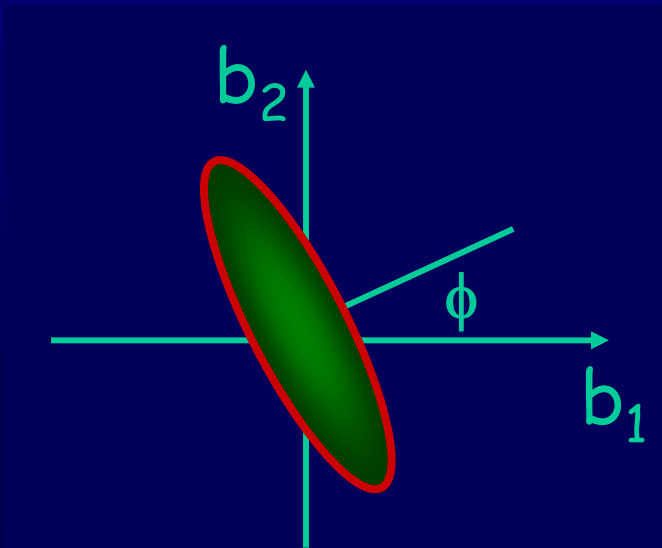


Squeezing using  
back-action effects

# Back Action Produces Squeezing



- Vacuum state enters anti-symmetric port
- Amplitude fluctuations of input state drive mirror position
- Mirror motion imposes those amplitude fluctuations onto phase of output field



Squeezing produced by back-action force of fluctuating radiation pressure on mirrors

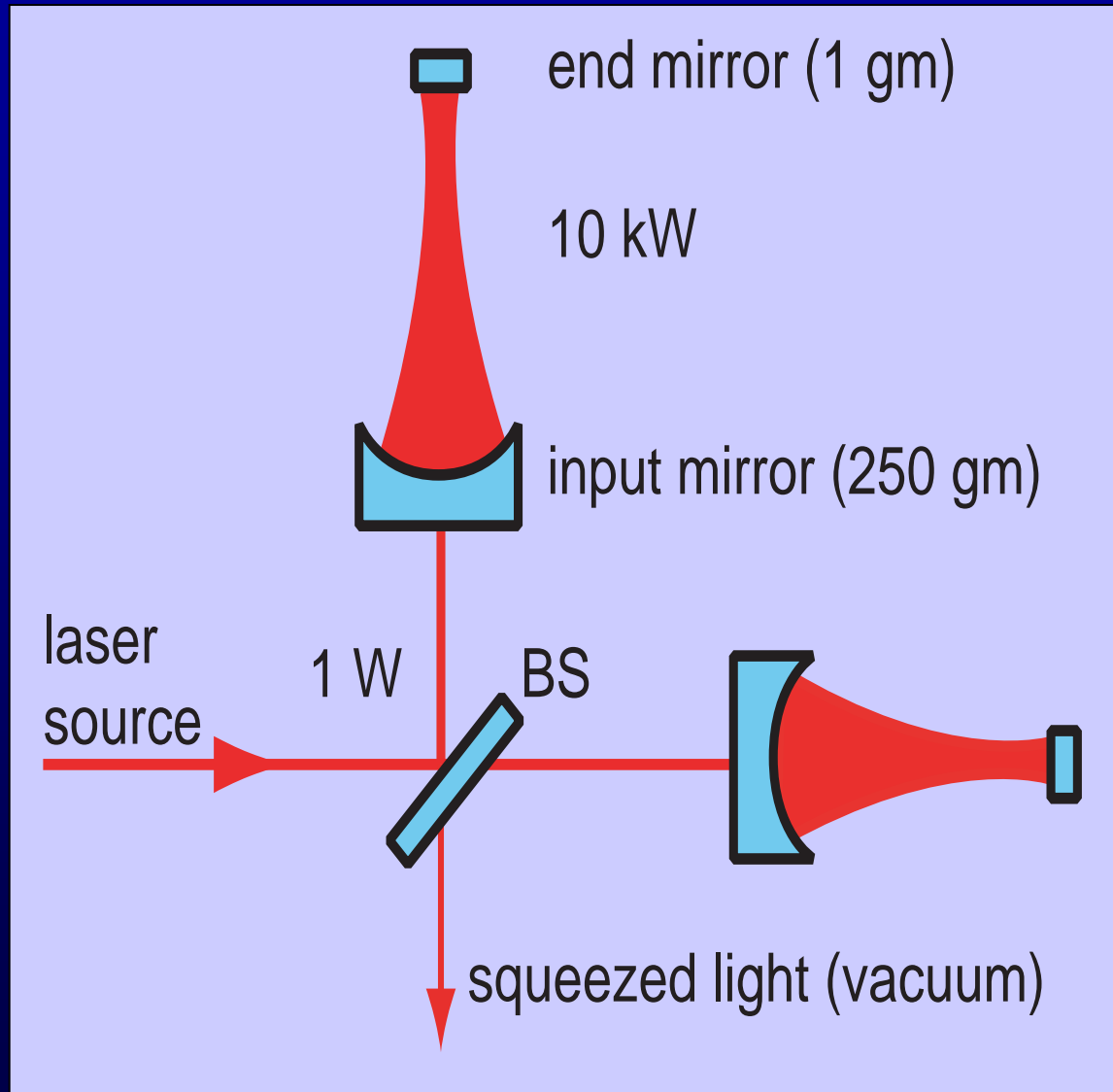
## The principle

- A “tabletop” interferometer to generate squeezed light as an alternative to nonlinear optical media
- Use radiation pressure as the squeezing mechanism
- Relies on intrinsic quantum physics of optical field–mechanical oscillator correlations
- Squeezing produced even when the sensitivity is far worse than the SQL
  - Due to noise suppression a la optical springs





# The Ponderomotive Interferometer



## Key ingredients

- High circulating laser power
  - 10 kW
- High-finesse cavities
  - 15000
- Light, low-noise mechanical oscillator mirror
  - 1 gm with 1 Hz resonant frequency
- Optical spring
  - Detuned arm cavities

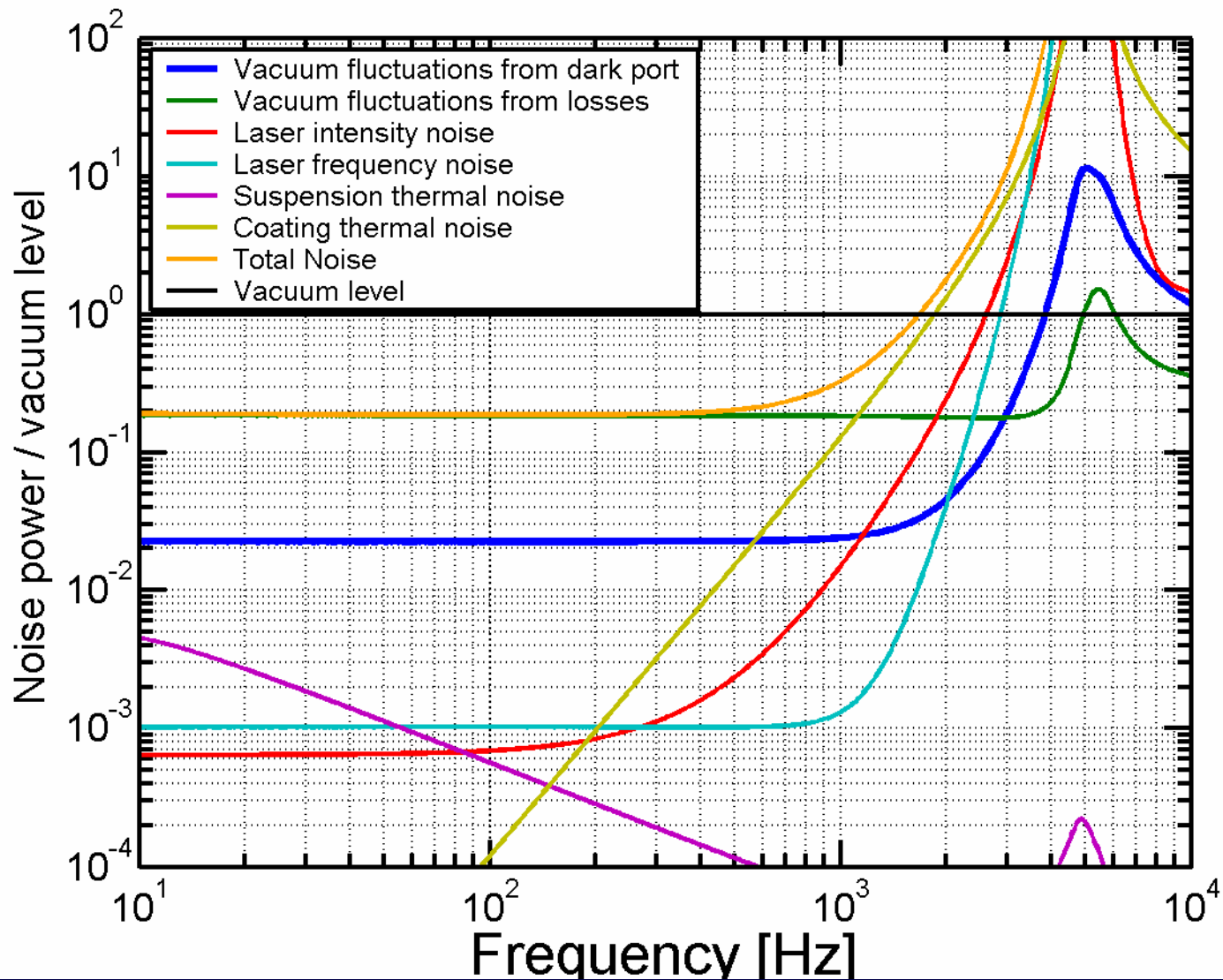


# Assumed experimental parameters

Parameter	Symbol	Value	Units	Parameter	Symbol	Value	Units
Light wavelength	$\lambda_0$	1064	nm	Input mirror trans.	$T_{ITM}$	$4 \times 10^{-4}$	-
Input mirror mass	$M_{ITM}$	0.25	kg	End mirror mass	$M_{ETM}$	1	g
Arm cavity finesse	$\mathcal{F}$	$1.6 \times 10^4$	-	Loss per bounce	-	$5 \times 10^{-6}$	-
Input power	$I_0$	1	W	Arm cavity detuning	$\delta$	$10^{-5}$	$\lambda_0$
BS refl. imbalance	$\Delta_{BS}$	0.01	-	Mich. phase imbalance	$\Delta\alpha_M$		
Mich. loss imbalance	$\Delta\epsilon_M$			Input mirror mismatch	$\Delta_T$	$5 \times 10^{-6}$	-
Detuning mismatch	$\Delta_\delta$	$10^{-7}$	$\lambda_0$	Arm cavity loss mismatch	$\Delta_\epsilon$	$2 \times 10^{-6}$	-
Susp. resonant freq.	$\Omega_0$	1.5	Hz	Susp. mech. loss angle	$\phi$	$10^{-6}$	-
Laser intensity noise	-	$10^{-8}$	$\text{Hz}^{-1/2}$	Laser frequency noise	-	$10^{-4}$	$\text{Hz}/\sqrt{\text{Hz}}$



# Noise budget



## Work so far

- Detailed simulation of noise couplings
  - Uses first fully quantum mechanical simulation code for a GW interferometer
- Location and infrastructure
  - LASTI laser, vacuum envelop and seismic isolation
- Cavity geometrical parameters
- Monolithic fused silica suspensions for mini-mirror

## What's next

- Design completion
  - Suspension
  - Control system
- High finesse cavity tests
  - Fixed mini-mirror – optical tests
  - Suspended mini-mirror – includes mirror dynamics and radiation-pressure coupling
- Complete interferometer

## Why is this interesting/important?

- First ever demonstration of ponderomotive squeezing
- Probes quantum mechanics of optical field-mechanical oscillator coupling at 1 g mass scales
- Test of low noise optical spring
  - Suppression of thermal noise
- Simulations and techniques useful for AdLIGO and other GW interferometers
  - Quantum optical simulation package
  - Michelson detuning
- Role of feedback control in these quantum systems

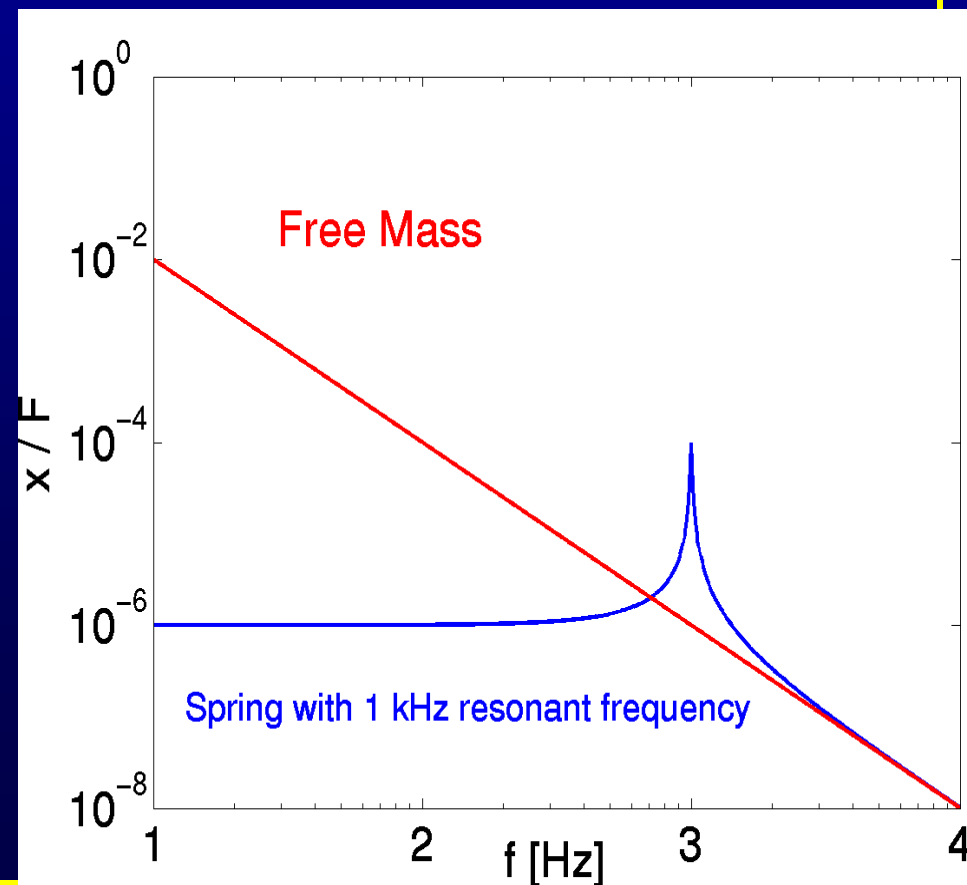


The End



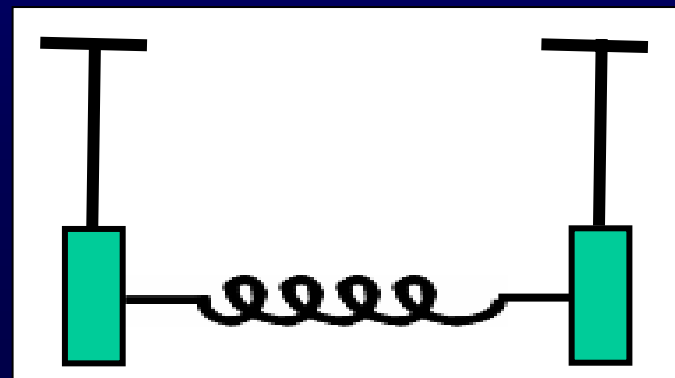
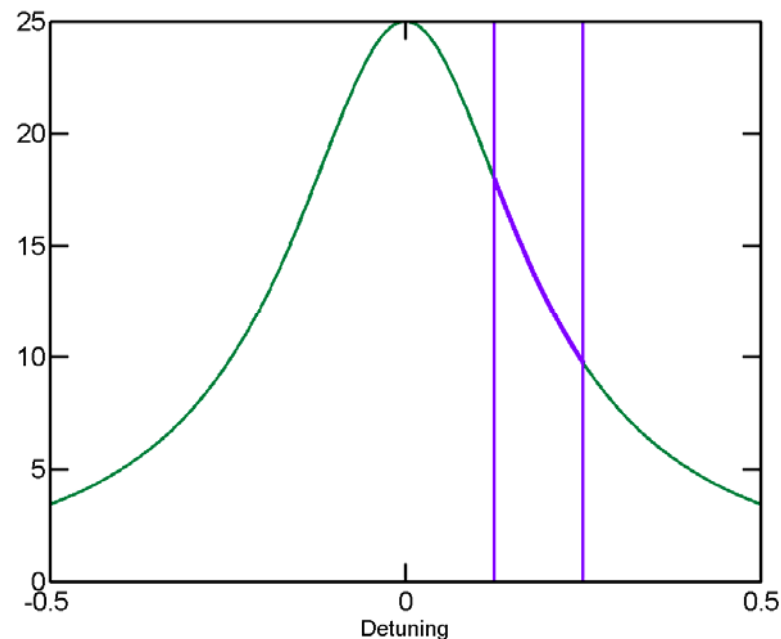
# Optical Springs

- Modify test mass dynamics
- Suppress displacement noise (compared to free mass case)
- Why not use a mechanical spring?
  - Displacements due to thermal noise introduced by the high frequency (mechanical) spring will wash out the effects of squeezing
- Connect low-frequency mechanical oscillator to (nearly) noiseless optical spring
- An optical spring with a high resonant frequency will not change the thermal force spectrum of the mechanical pendulum
  - Use a low resonant frequency mechanical pendulum to minimize thermal noise
  - Use an optical spring to produce a flat response out to higher frequencies



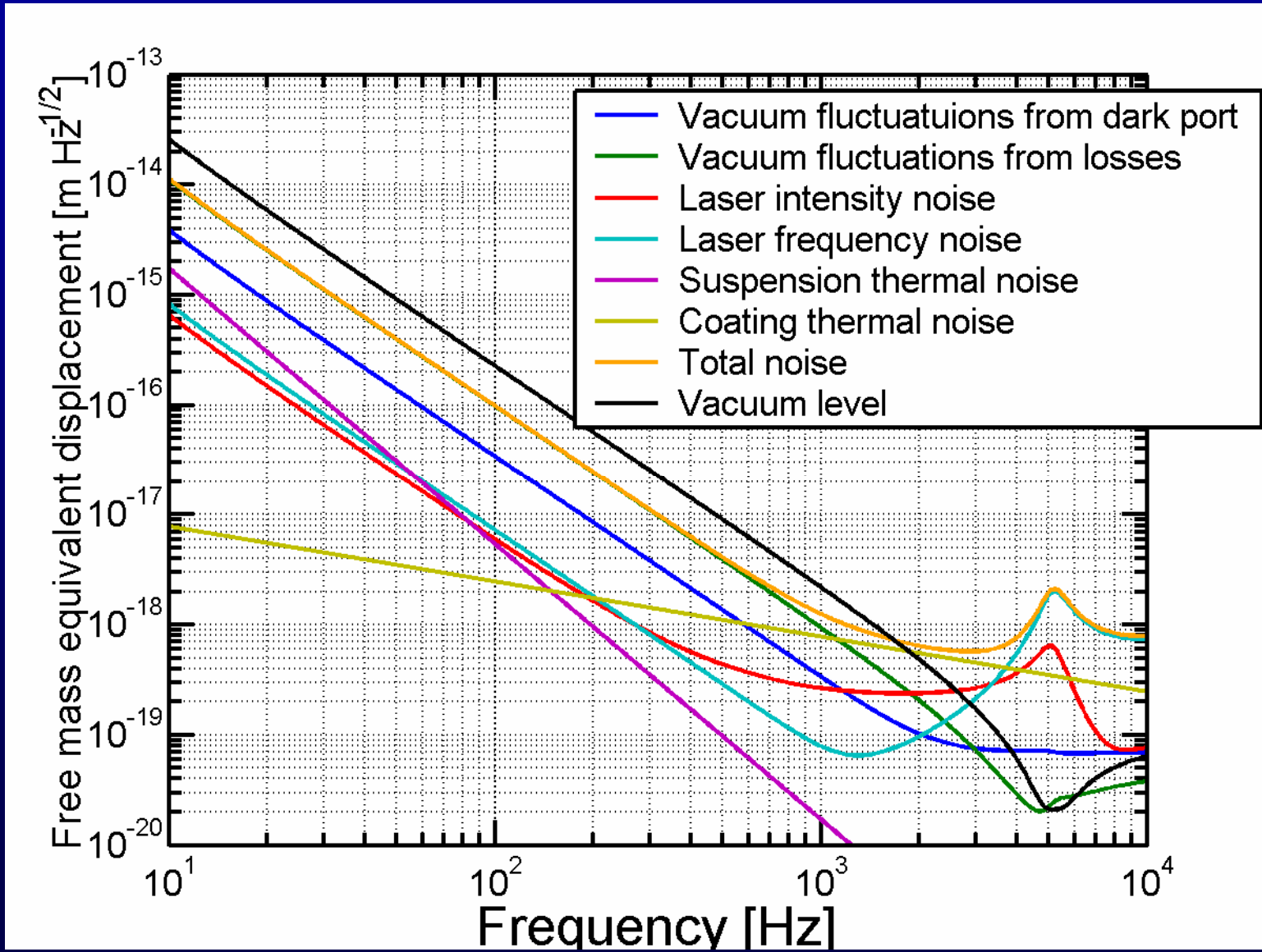
# Detuned cavity for optical spring

- Positive detuning
  - Detuning increases
  - Cavity becomes longer
  - Power in cavity decreases
  - Radiation-pressure force decreases
  - Mirror 'restored' to original position
  - Cavity becomes shorter
  - Power in cavity increases
  - Mirror still 'restored' to original position





# Noise budget - Equivalent displacement





# Squeezed Vacuum

