



LISA

Laser Interferometer Space Antenna

Fundamental Limit in Frequency Stabilization of Lasers

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Abstract



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Frequency stabilization of lasers with rigid cavity

- Widely used in many fields
 - Including LISA
- Stability achieved with rigid cavity
 - Unbeatable and unidentified noise

Thermal fluctuation (Brownian noise) of rigid cavity

- We point out this as a fundamental limitation.
 - Rigorous evaluation with experiments and calculations
 - Agreement with world-highest level stabilization results

New insights for precision measurement communities



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1. Frequency stability of laser



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Why frequency stabilization?

– Wide range of application

- Optical frequency standards
- High-resolution spectroscopies
- Fundamental physics tests
 - Ex.) Michelson-Morley type experiment: basis of Special/General Relativity
- Interferometric measurements
 - -Ex.) Gravitational wave detection using laser interferometer: LISA, LIGO...

– Used as wavelength reference

- Laser frequency assumed to be fixed to one frequency
- In reality, it fluctuates! : Needs to be stabilized.
 - LISA requirement: 30Hz/rtHz



Methods of frequency stabilization

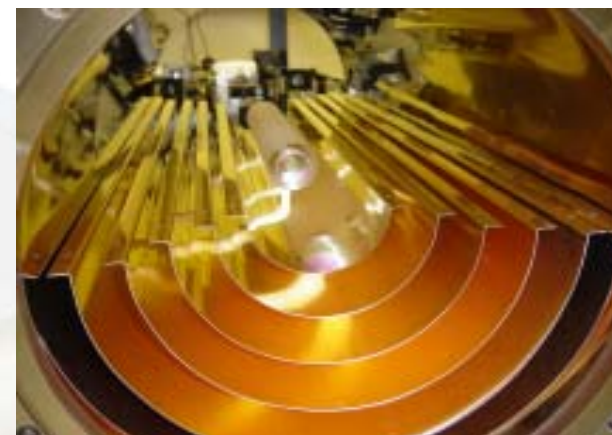
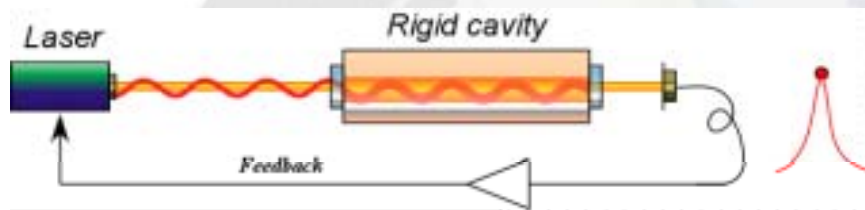


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Two major frequency references

– Rigid cavity (Basic design for LISA) <<<

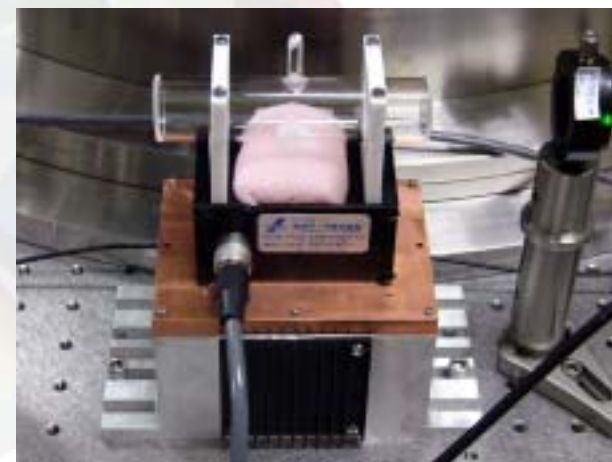
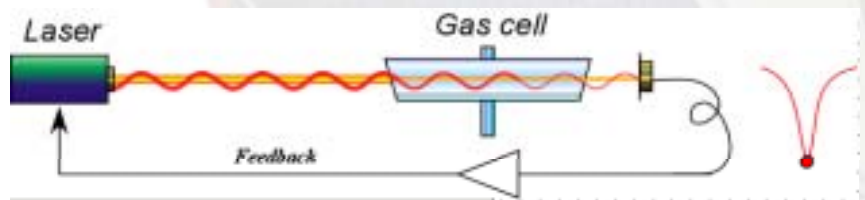
- Made of low CTE material
- Thermal shield to minimize length variation
- Laser controlled to be stored within the cavity



ULE rigid cavity with thermal shields

– Atomic absorption line (Basic design for TPF)

- Cooling to avoid Doppler broadening
- Use of hyperfine structure
- Laser controlled to be absorbed by one line



Iodine cell with cooling system

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Stability of rigid cavity



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Possible noise sources

– Non-fundamental noise source

- Length change due to temperature variation
- Length change due to vibration (seismic noise)
- Length change due to mirror heating --- coupled to laser intensity noise
- Coupling from RF amplitude noise
- Pointing noise coupled to misalignment
- Circuit noise
- Etc...

– Fundamental noise source

- **Thermal noise as a result of statistical physics** <<<

- Hasn't been evaluated from 1970's!!

● Calculation of thermal noise spectrum $G(f)$

- Based on **FDT** (Fluctuation-Dissipation Theorem)

$$G(f) = -\frac{4k_B T}{\omega} \text{Im}[H(\omega)]$$

$H(f)$: transfer function

- Useful form

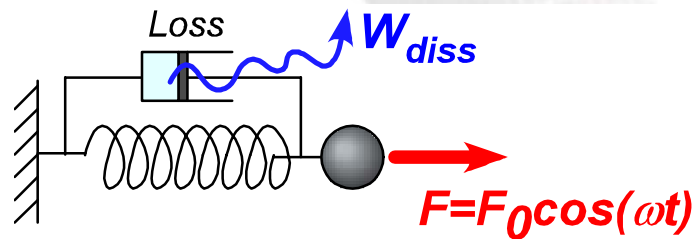
$$G(f) = \frac{4k_B T}{\pi^2 f^2} \frac{W_{diss}(f)}{F^2}$$

W_{diss} : dissipated energy under cyclic force

F : force amplitude

ε : strain energy under cyclic force

$$W_{diss}(f) = \int \varepsilon(\vec{r}) dV / Q$$



Stored strain energy

Calculation

Quality factor <<<

Experiment

- Basic information in estimating thermal noise level
 - **Mechanical quality factor (Q)**
 - Measured by “ring-down” method
 - Vibration decay measured by Michelson interferometer
 - Sample supported by thin wires in vacuum to reduce external loss
 - **Samples**
 - Low CTE materials
 - Imitating rigid cavity



ULE sample



Measurement system



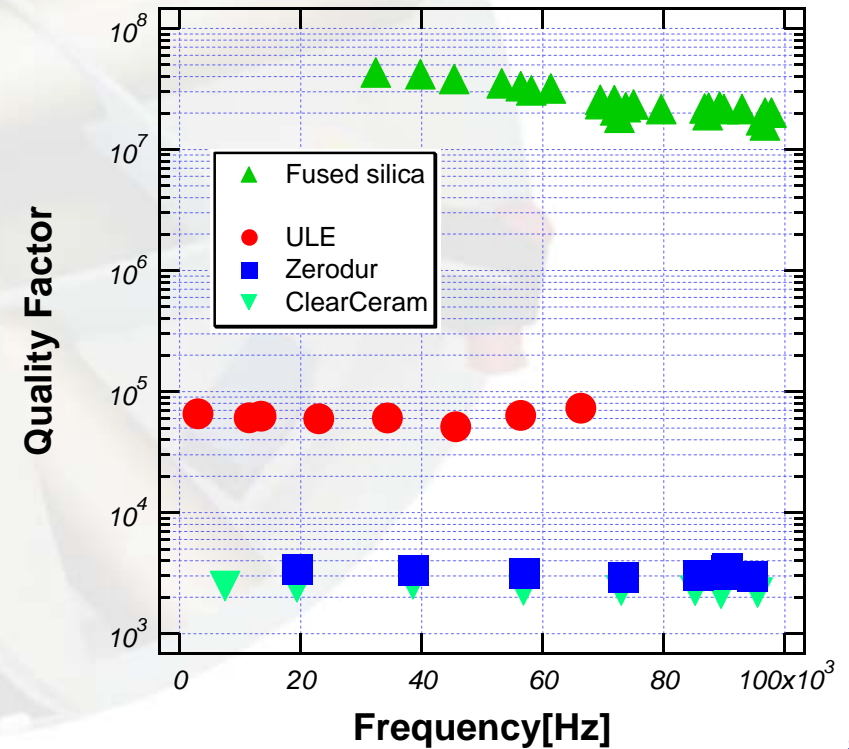
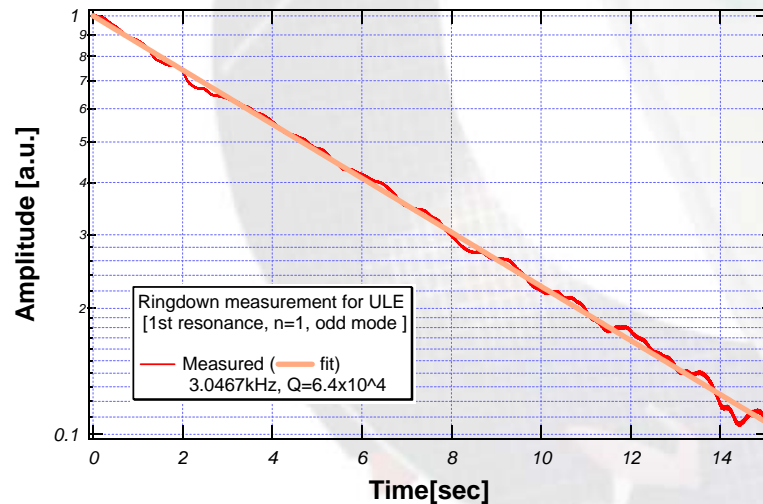
Experimental result



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Fairly low quality factor measured

- ULE : $Q \sim 61000$, Zerodure: $Q \sim 3100$, etc.
 - Usually high Q materials used when thermal noise matters
 - Ex. Ground based gravitational-wave detector: mirror made of silica: $Q > 10^7$
- Constant Q measured



3. Calculation



- Calculation of dissipated energy
 - Calculation of strain energy under cyclic force
 - Done by solving Equation Of Motion (EQM) of the system
 - Numerical approach adopted
 - Finite Element Method (FEM)
 - Procedure
 - 1) Prepare rigid cavity mechanical model
 - 2) Apply cyclic force to the observing (beam-illuminating) points
 - 3) Calculate strain energy within the system based on EQM

$$W_{diss}(f) = \int \varepsilon(\vec{r}) dV / Q$$

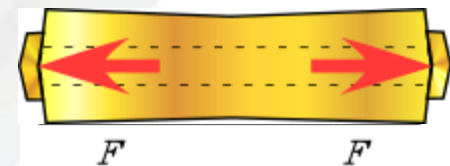


Stored strain energy

Calculation

Quality factor

Experiment



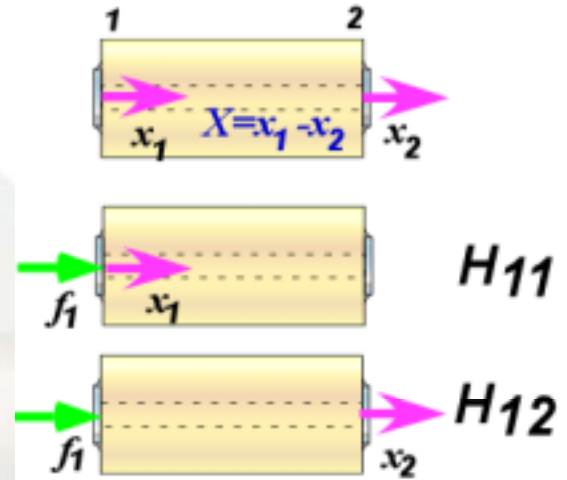
Calculation of correlated thermal noise

- Individual calculation but with correlation term

- $G_X = G_{11} + G_{12} - 2G_{12}$

- Apply two forces simultaneously to measured points (areas)

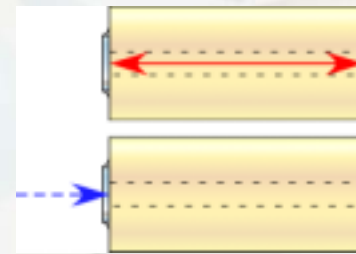
- These two give us equivalent results



Solving equation of motion

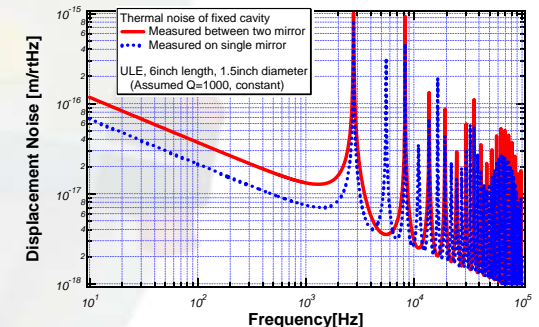
- Numerical approach adopted

- FEM to get W_{diss}
 - Can be applied for any:
 - Frequency, shape, loss distribution/frequency dependence, weighing...
 - (This method itself should be published somewhere.)



- In the following

- Thermal noise in rigid cavity solved
 - Problem includes:
 - Finite sized mass, material combination (loss distribution), Gaussian beam, coating loss, correlated (generalized coordinate) etc...



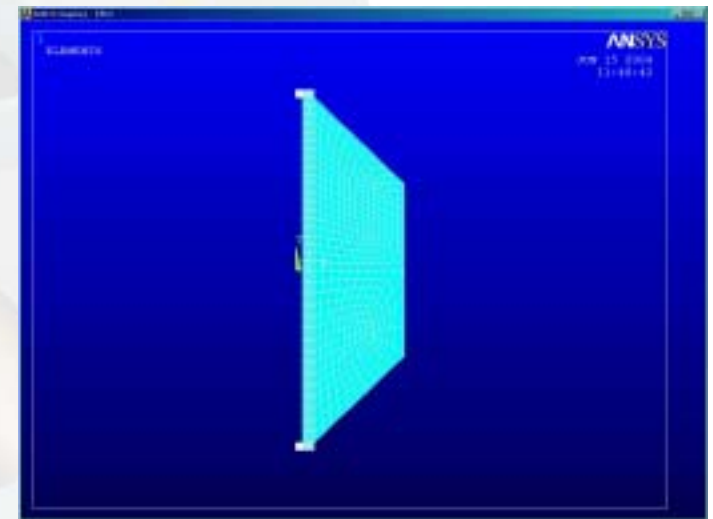
World highest stability achieved by NIST/VIRGO

– Calculation assumptions

- Spacer
 - Material: ULE (Q=60000)
 - Length: 15cm (tapered), diameter: 24cm
- Mirror (optical contacted)
 - Material: ULE
 - Diameter: 1inch, thickness: 5mm
 - Beam radius: 240um on both mirrors
- Coating
 - Thickness: 2um, $\phi(1/Q)=4 \times 10^{-4}$

– FEM model

- ANSYS
- Semi-3D model (2-D axisymmetric)



Half of the cross section model of the cavity

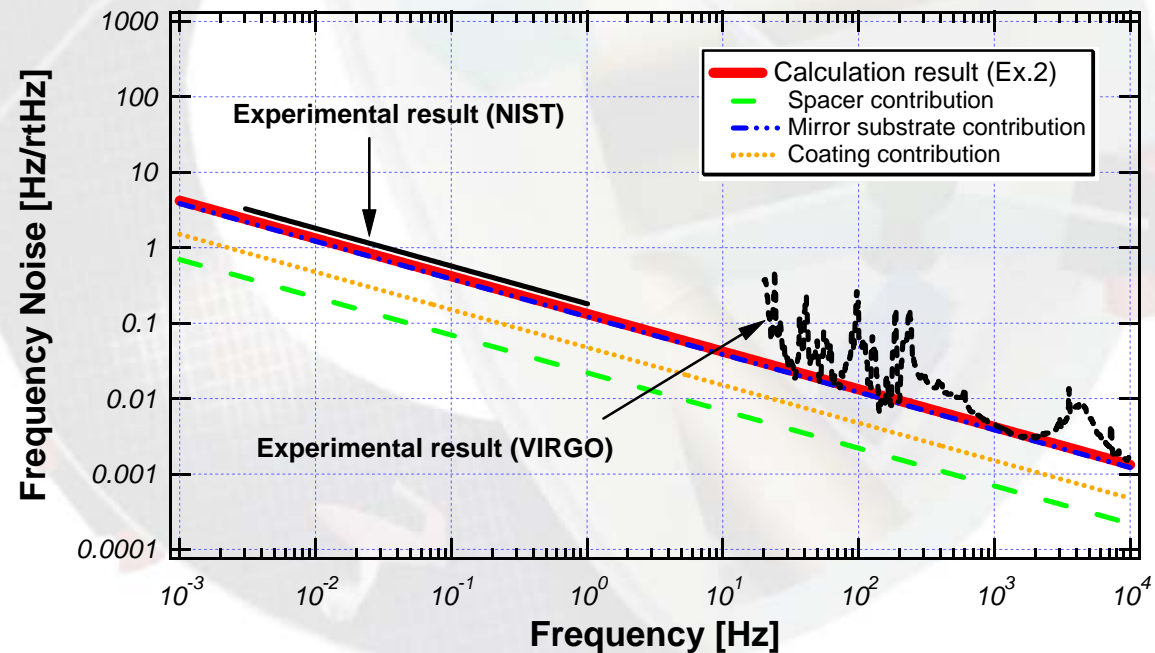


Comparison with experiment



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- Agreed pretty well with the measurement
 - ~1Hz/rtHz@0.01Hz level (563nm wavelength)
 - We cannot neglect thermal noise (Brownian motion) anymore!
 - Use of low loss mirrors, larger beam diameter, cooling etc
 - Expected to renew world highest frequency stability



- Frequency stabilization of laser
 - Wide-range of demands and applications in physics and engineering
 - Use of rigid cavity
 - Basic design for LISA
 - Any length fluctuation of cavity limits frequency stability
- Fundamental limit in frequency stabilization with rigid cavity
 - Thermal noise as a result of statistical physics
 - We evaluated the noise level with the FDT.
 - Experiment: Q measurement of cavity materials
 - Calculation: Numerical analysis of strain energy
 - Importance of thermal noise pointed out
 - Agreement with world-highest level stabilization results
 - See PRL **93** (2004) 250602 for details.