

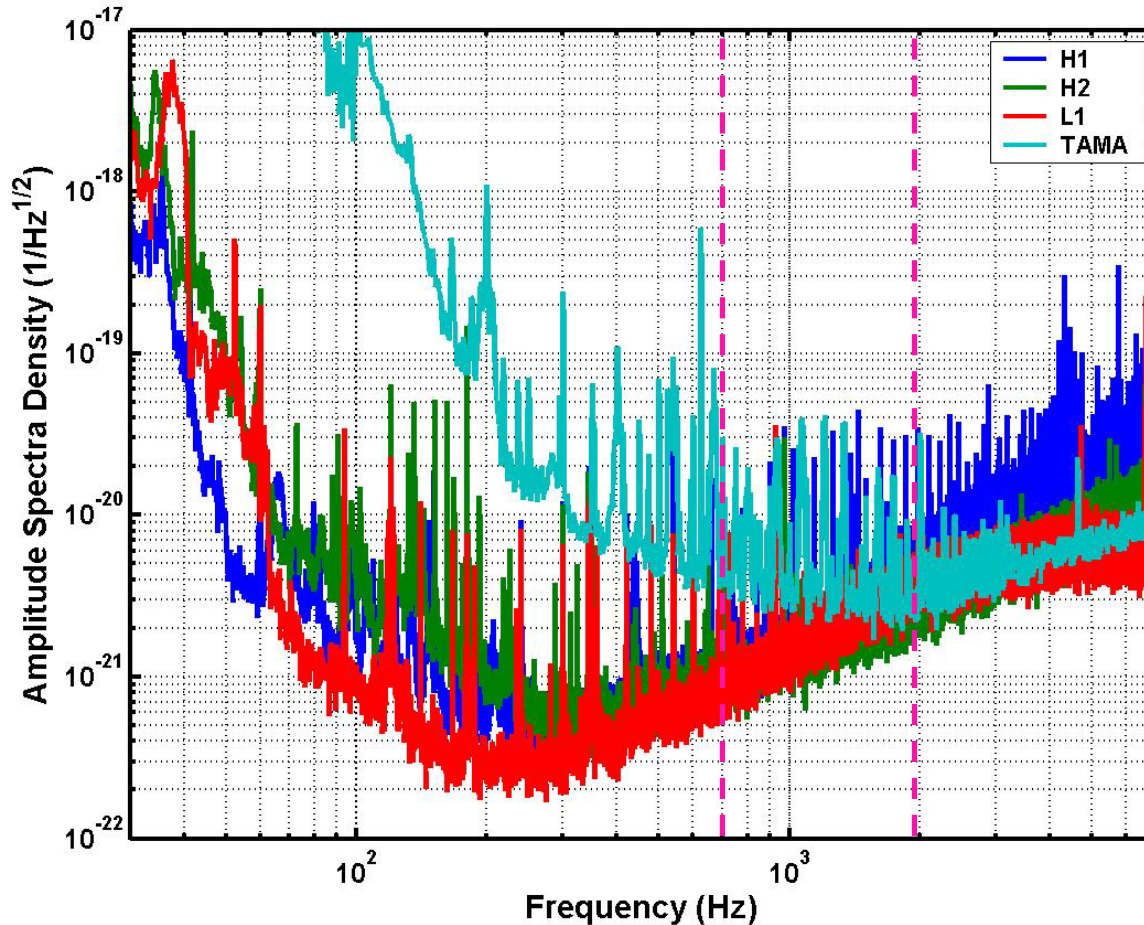
# Status of the LIGO-TAMA Joint Bursts Search

Patrick Sutton

LIGO Laboratory, Caltech, for the  
LIGO-TAMA Joint Working Group

- Background
- Science with LIGO-TAMA
- Analysis Overview
- Upper Limits
- Remaining Tasks and Schedule

- The LIGO S2 un-triggered search for generic GWBs at high frequencies has been carried out jointly with TAMA.
  - » LIGO-only search: 100-1100Hz
  - » LIGO-TAMA search: 700-2000Hz
- Philosophy of LIGO-TAMA: Use machinery and techniques common with the LIGO-only low-frequency search
  - » ETGs, coincidence, r-statistic, time lags, MDC simulation frames, low false rate ( $\ll 1/S^2$ ), rate vs strength upper limits, ...
- Seek optimal ways to combine LIGO and TAMA for best science.



LIGO and TAMA look with best sensitivity at different frequencies:

- Tune for signals near minimum of envelope, [700-2000]Hz.

Duty cycles:

H1	74%
H2	58%
L1	37%
T1	81%

- Take advantage of excellent T1 duty cycle.

- TAMA-LIGO 4X search has several interesting features:
  - » 4X coincidence allows for searches with very low false rates (<1/yr).
  - » Extra time lags allow much more accurate background estimates
    - LIGO 2-site network = 47 lags in (-115s,+115s)
    - LIGO-TAMA 3-site network =  $47^2 = 2209$  lags in (-115s,+115s).
  - » *Not yet explored*: Extra non-aligned site with long baseline: exploit for sky direction? polarization information?
- But 3X searches also valuable:
  - » Can use TAMA as substitute for a missing LIGO detector.
  - » Eg: H1-H2-T1 coincidence allows us to use the large amount of H1-H2 data that would otherwise be lost because of poor L1 duty cycle (Brady, Cadonati, Katsavounidis; also done in joint inspiral search).

H1-H2-L1-T1	250hr
H1-H2- $\bar{\text{L}}$ 1-T1	325hr
H1-H2-L1- $\bar{\text{T}}$ 1	62hr
total LIGO-TAMA	637hr
total LIGO-only 3X	312hr

$\bar{\text{L}}$ 1  $\equiv$  L1 not operating

$\bar{\text{T}}$ 1  $\equiv$  T1 not operating

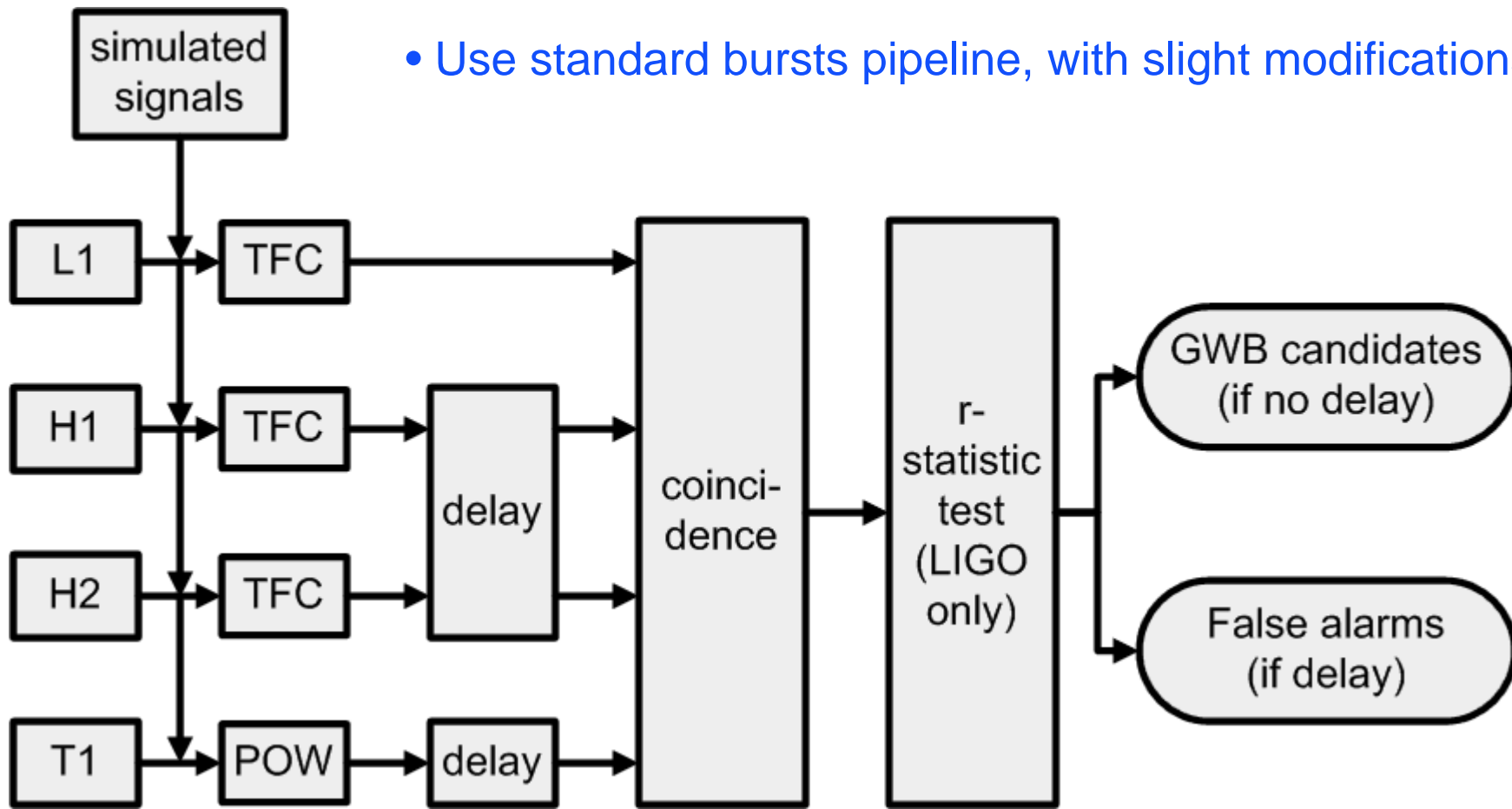
- TAMA *doubles* total usable data set
  - » Better chance of “getting lucky” in a search
  - » Cut rate upper limits in half
  - » Cost: some loss in efficiency (minor effect)

$$\longrightarrow R(h) = \frac{N}{\epsilon(h)T}$$

- Response: Analyze all H1-H2-(L1 or T1) data
  - » H1-L1-T1, H2-L1-T1: small amount of data, much higher false rate. Ignore.

- 3 independent data sets:
  - » Must derive single upper limit from 3 independent experiments.
  
- No access to TAMA data (only triggers exchanged):
  - » Compare LIGO-TFClusters triggers to TAMA-Power triggers
    - TFClusters triggers have peak time, duration, SNR, central frequency, and bandwidth.
    - TAMA triggers have peak time, duration, SNR, but no central frequency or bandwidth.
    - Coincidence is based on time overlap only. No SNR or frequency comparison.
  - » No r-statistic test with TAMA

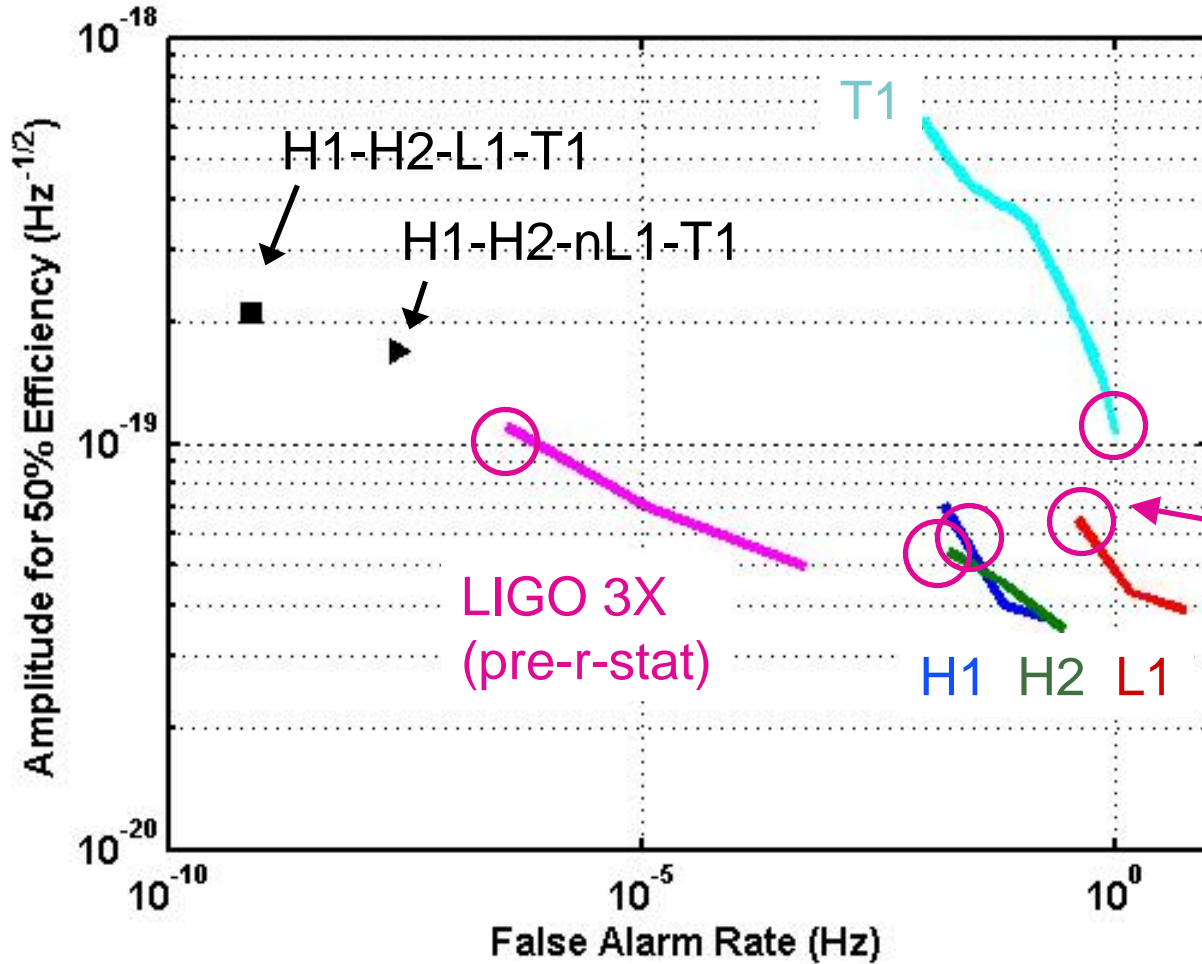
- Use standard bursts pipeline, with slight modifications





- One set of MDC frames has been exchanged: SG13
  - » sine-Gaussians
  - »  $Q = 8.9$
  - »  $f_0 = \{700, 849, 1053, 1304, 1615, 2000\}$ Hz
  - » isotropic sky distribution
  - » random linear polarization
  - » total ~16800 injections, distributed over LIGO 3X times (H1-H2-L1-T1 and H1-H2-L1-nT1)

- Use single tuning for all three data sets.
- Tune for best efficiency at each false rate.
  - » single-IFO: use to fix TFClusters parameters
  - » multi-IFO: select bpp/SNR to match efficiencies
- Select multi-ETG rate & r-statistic threshold for  $\ll 1$  event from background.
  - » H1-H2-L1-T1 and H1-H2-nL1-T1 have ~same observation time, so efficiencies are averaged
  - » H1-H2-nL1-T1 dominates false rate.



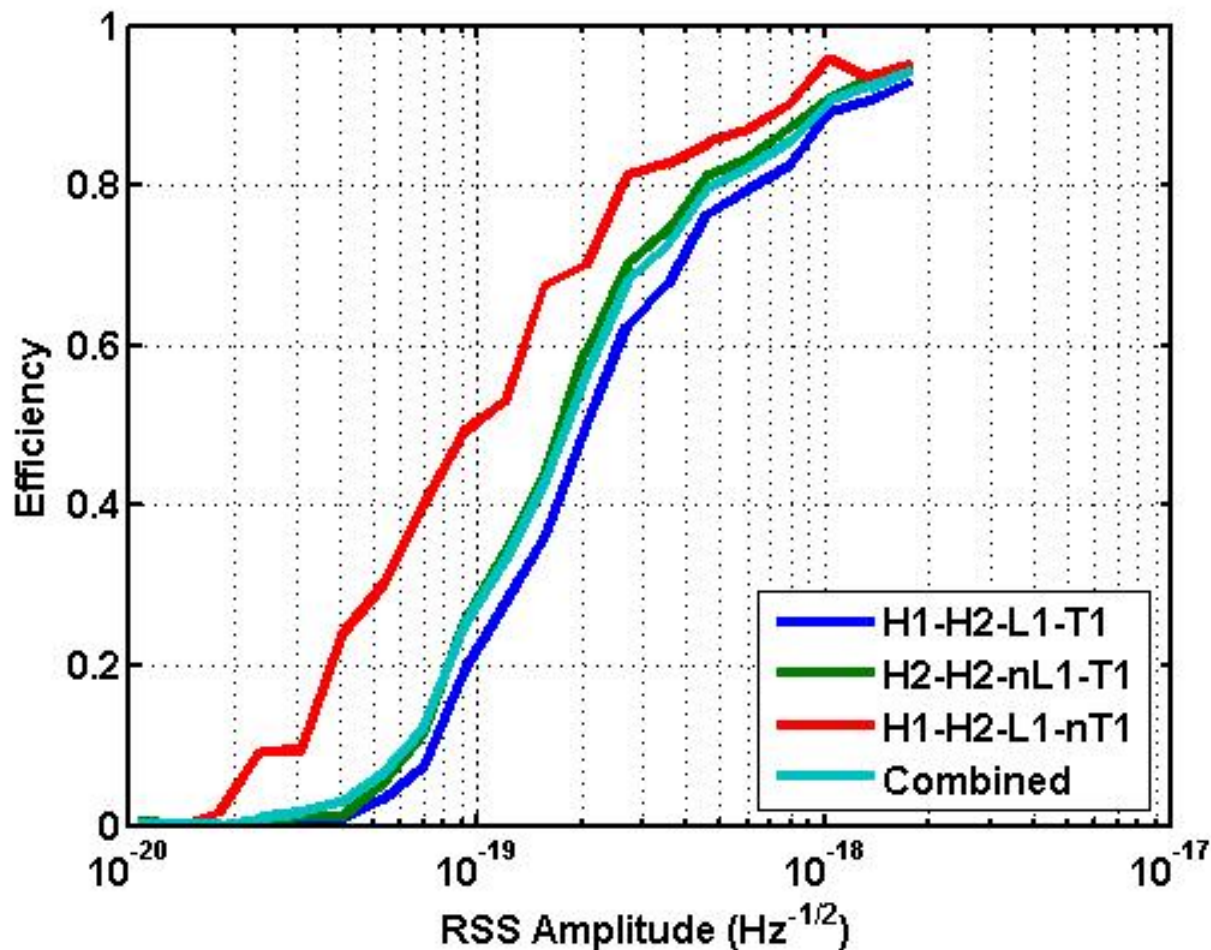
From SG13 simulations

Effective coincidence windows:

20ms (LIGO-LIGO)  
43ms (LIGO-TAMA)

Chosen single-IFO operating points

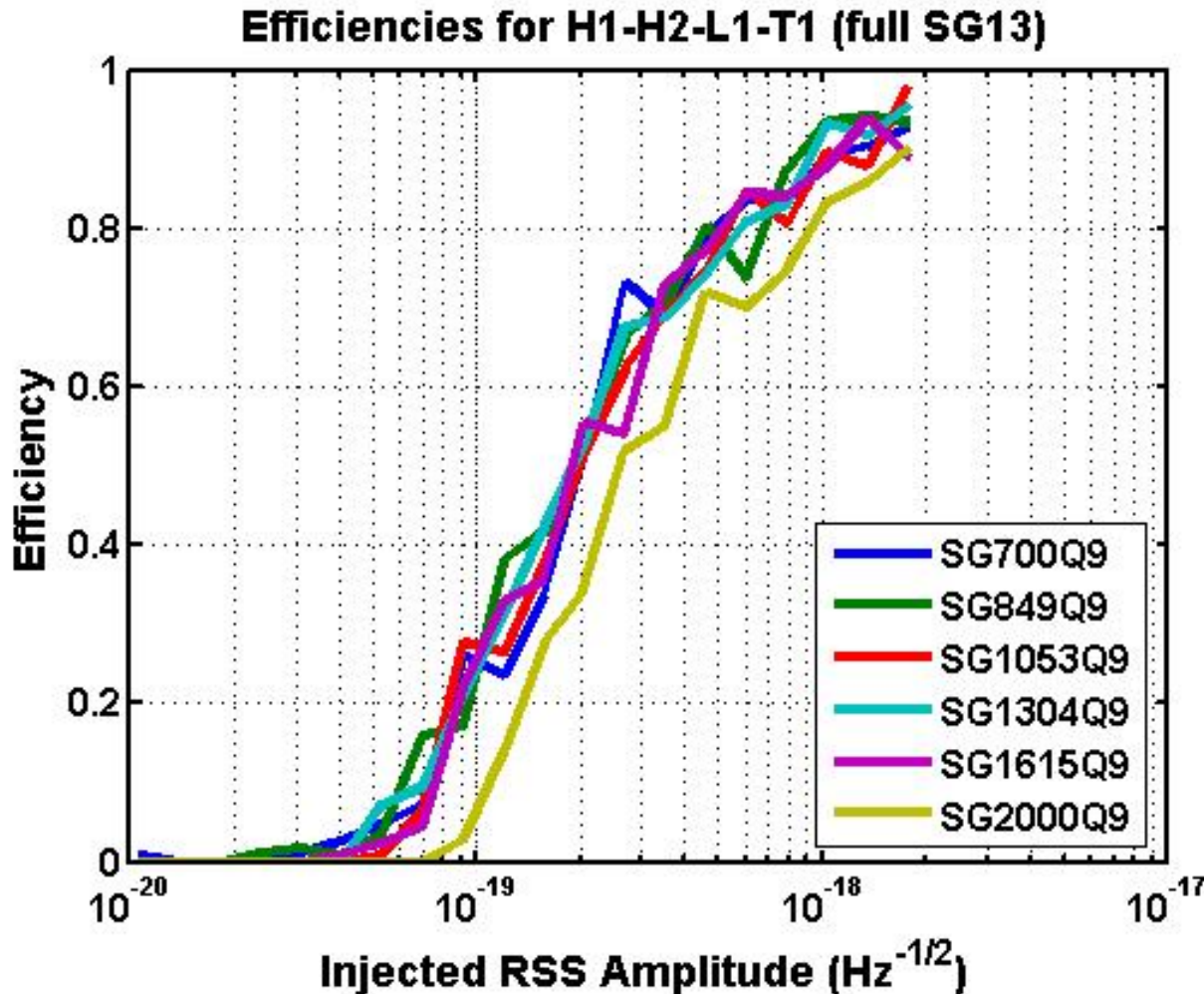
- Full data set box has been opened and final upper limits have been calculated.
  - » No surviving coincidences (after r-statistic) for any of the network combinations.
  - » Rate upper limit of 0.13/day.
  - »  $h_{\text{rss}}^{50\%} = 2 \times 10^{-19} \text{Hz}^{-1/2}$  averaged over networks, analysis band.



SG13 simulations

(Q=8.9 SG over [700,2000]Hz, with sky & polarization averaging)

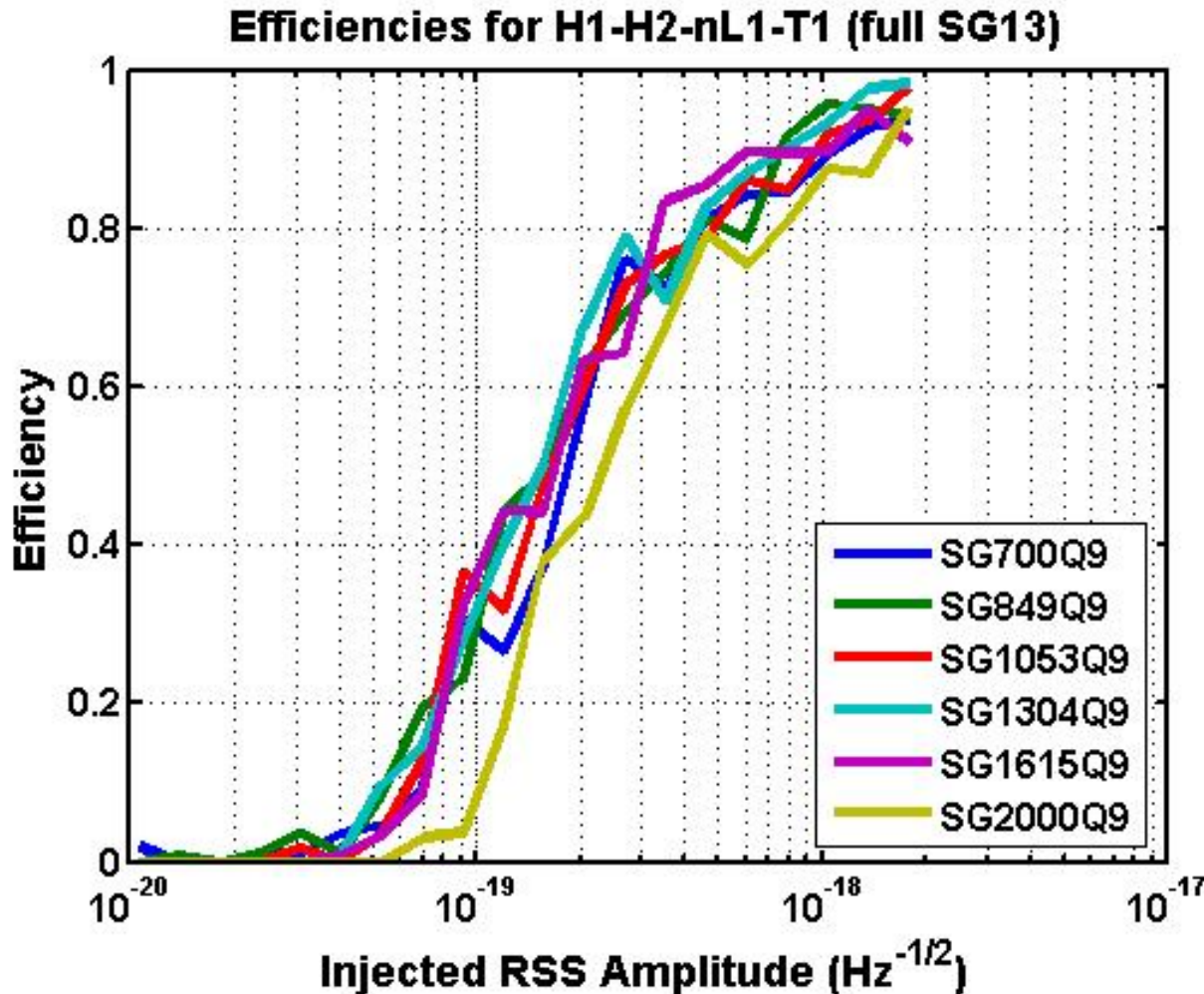
Different network combinations have similar efficiency (factor ~2 in 50% point).



4X detection

SG13 simulations  
separately by central  
frequency

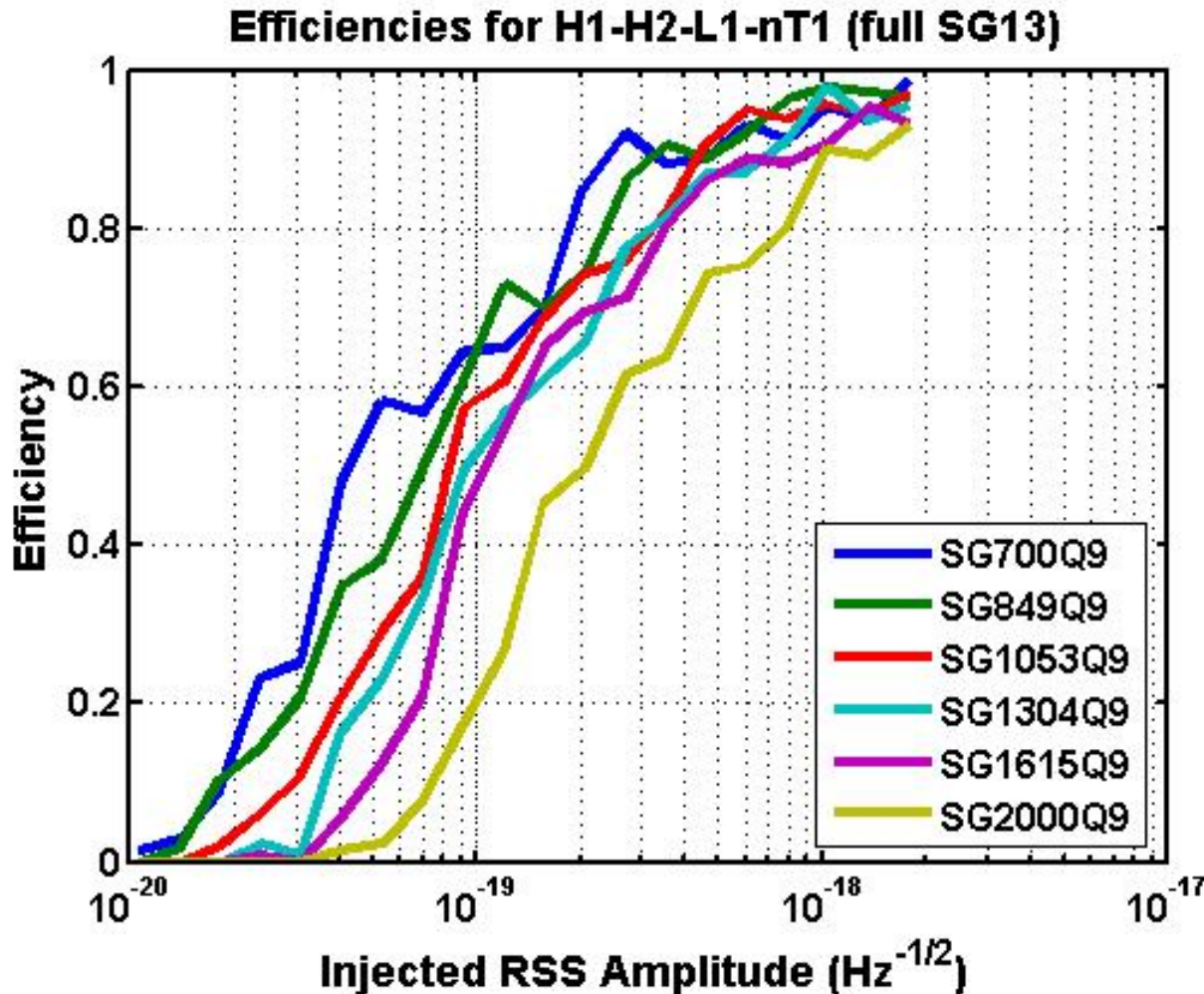
All about the same.



3X detection (no L1)

SG13 simulations  
separately by central  
frequency

All about the same.



3X detection (no T1)

SG13 simulations  
separately by central  
frequency

Better at lower  
frequencies – TAMA  
limits sensitivity  
there.



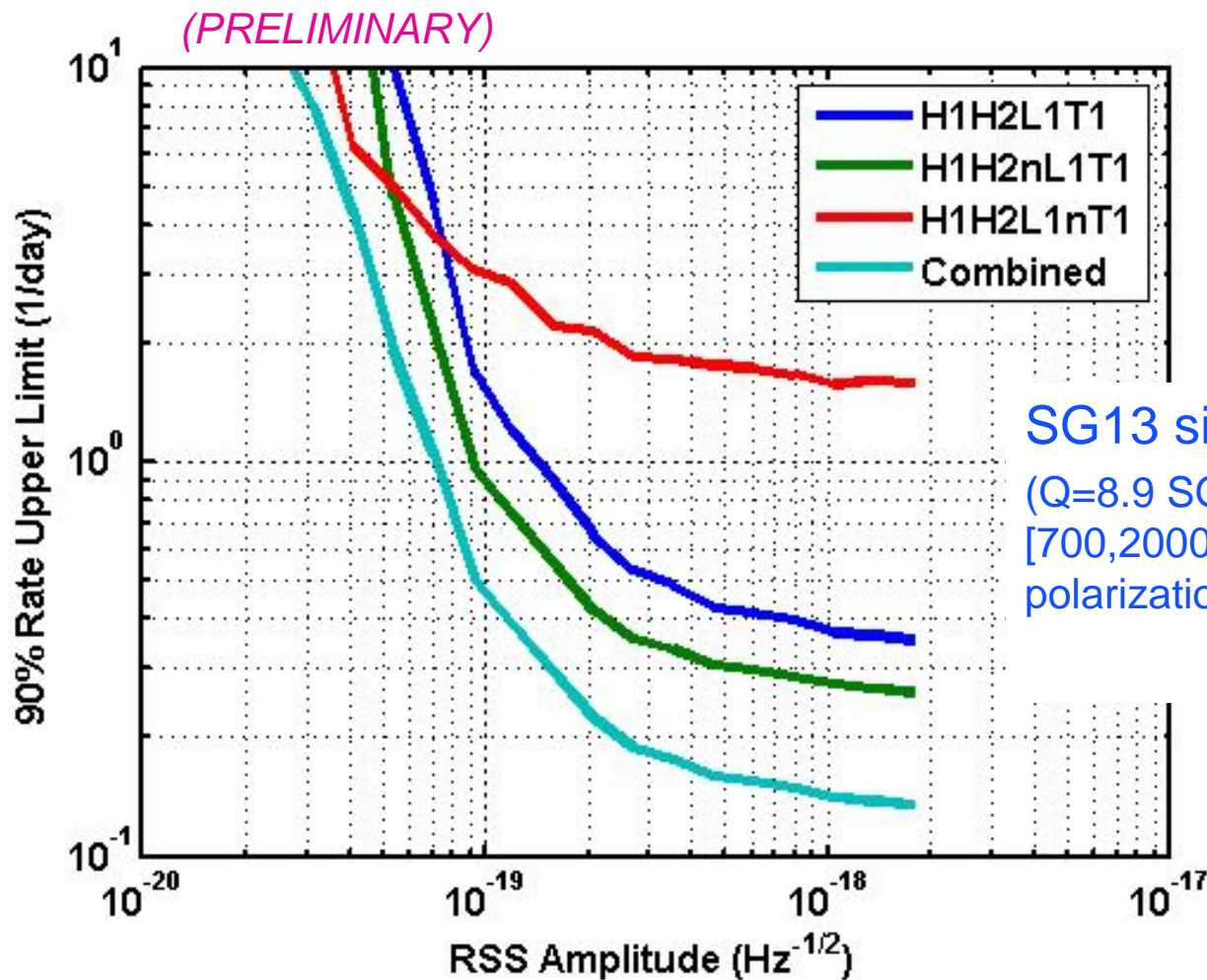
*Full data set, including N before/after the R-Statistic:*

Network	T (Ms)	N	$R_{\text{bck}}$ (nHz)	$N_{\text{bck}}$	$R_{90\%}$ (1/day)	$h_{50\%}$ ( $\text{Hz}^{-1/2}$ )
H1-H2-L1-T1	0.64*	0/0	<0.75	<5e-4	0.33	$2.1 \times 10^{-19}$
H1-H2- $\bar{\Gamma}$ L1-T1	0.84*	3/0	<27	<0.023	0.25	$1.7 \times 10^{-19}$
H1-H2-L1- $\bar{\Gamma}$ T1	0.14	0/0	<165	<0.023	2.41	$0.97 \times 10^{-19}$
Combined**	1.6	3/0	n/a	<0.046	0.13	$1.8 \times 10^{-19}$

\*TAMA livetimes to be finalized.

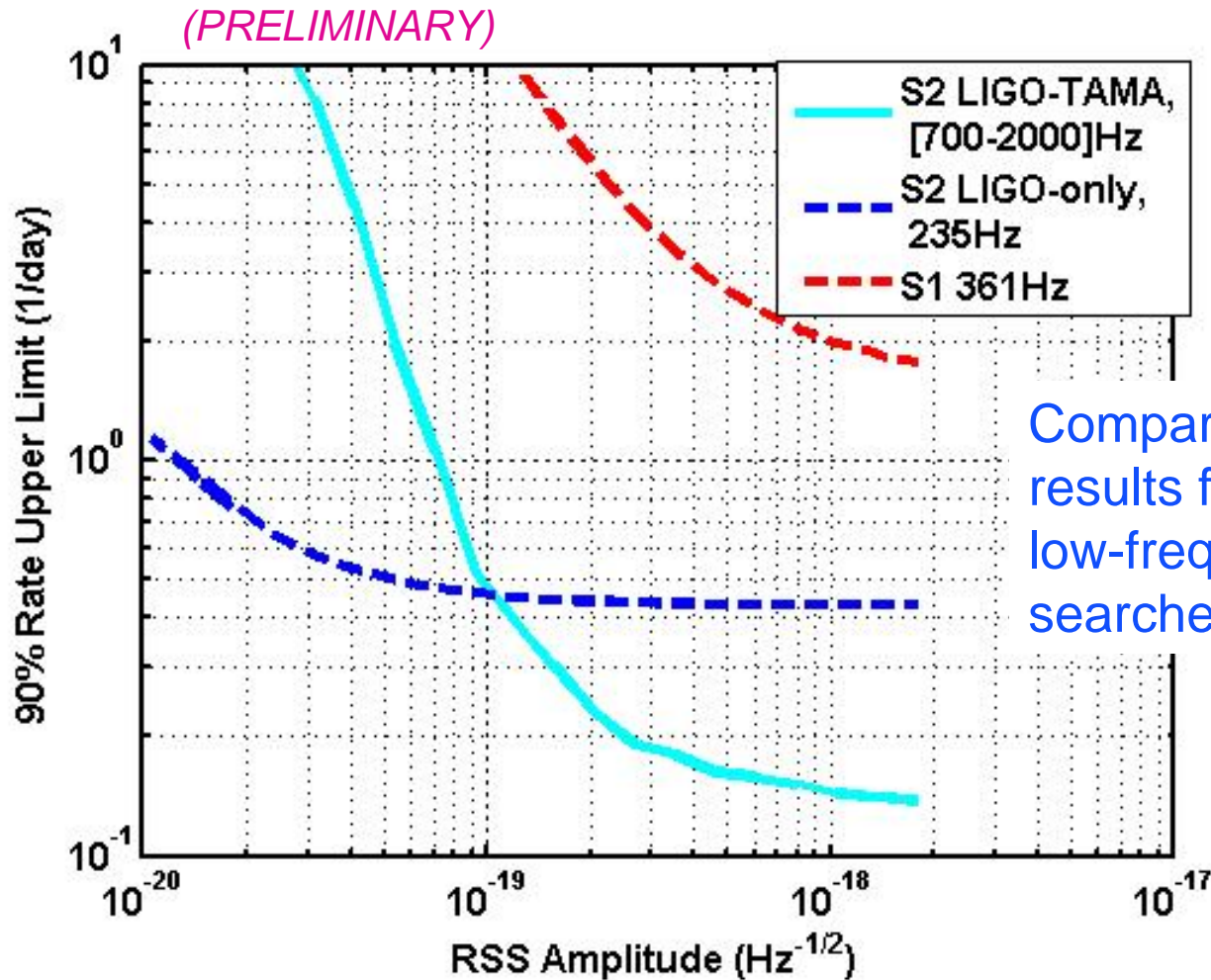
\*\*Treating all 3 data sets as one experiment (all have  $N_{\text{bck}} \sim 0$ ).

# R vs h Upper Limits



SG13 simulations  
 (Q=8.9 SG over  
 [700,2000]Hz, with sky &  
 polarization averaging)

# R vs h Upper Limits



Compare to best results from S1, S2 low-frequency searches.

- TAMA-LIGO joint search for GWBs in S2 is in late stages.
  - » High-frequency search complementary to LIGO-only search at low frequencies.
  - » Two main parts:
    - 4X: very low false rate
    - 3X: lots of additional observation time
  - » No GWB candidates survived pipeline.
    - Rate upper limit of **0.13/day**.
    - $h_{\text{rss}}^{50\%} = 2 \times 10^{-19} \text{Hz}^{-1/2}$  averaged over networks, analysis band.
- Remaining issues:
  - » livetime to be finalized (account for TAMA veto deadtime)
  - » review (esp TFClusters)
- Paper draft in preparation.
  - » Hope to present at GWDAW.
- S3?
  - » Exploring value of joint S3 search with LIGO, TAMA, GEO representatives.