

Status of LIGO

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LIGO-G040206-00-Z

Outline

- Gravitational Waves
- LIGO Project
- LIGO Searches
- Searching for Gravitational–Wave Bursts
- Outlook

A New Window on the Universe



Gravitational waves provide a new and unique view on the universe. Produced by coherent motion of bulk matter; information content is complementary to that in EM waves

Expected signals & sources:

- » Transients: supernovae, inspiraling binaries, black-hole ringing
- » Stochastic background: big bang
- » Continuous monochromatic: pulsars

Possibility for the unexpected is very real!

Gravitational Waves

- Quadrupolar fluctuations in geometry of space.
- Effect of gravitational wave coming out of screen on da Vinci's Vitruvian Man:



Second polarization is rotated by 45° about direction of propagation.

Interferometric GW Detectors

• A laser is used to measure the relative lengths of two orthogonal cavities (or arms).



• LIGO: Current technology allows one to measure $\delta L/L \sim 10^{-22}$.

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LIGO

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The LIGO Project

- Flagship project of NSF.
- 3 detectors at 2 sites.

LIGO

- First operations 1999.
- Scientific data taking runs:
 - » Aug–Sept 2002
 - » Feb-Apr 2003
 - » Oct 2003 Jan 2004



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LIGO Hanford Observatory



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LIGO Sky-Averaged Sensitivity to Inspiraling Neutron Star Binaries (S3)



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LIGO Sensitivity to GW Bursts (S3)



 Black-hole binary merger & ringdown ~ few Mpc



• Supernovae ~ few kpc

GW Searches

• Four Search Groups:

- » Inspirals (eg, neutron-star binaries) Shawhan
- » Bursts (unspecified signals) Marka
- » Periodic (eg, pulsars)
- » Stochastic Background (eg, big bang)
- Formal publications from S1 in press.
- S2 analysis in final stages, S3 in progress.
- Collaborative searches with other projects:
 - » GEO600
 - » TAMA300
 - » Virgo (in negotiation)





Summary

- LIGO is up and running, and approaching initial design sensitivity.
 - » Upgrades to seismic isolation & many other systems in progress
 - » Better duty cycle & sensitivity expected
- First "S1" searches conducted
 - » No detections yet! Set upper limits on GWs.
 - » Served as testing ground for analysis techniques.
- Further analyses in progress
 - » S2, S3 data: 10-100 times more sensitive, ~4 times longer than S1.
- International partnerships are strengthening
 - » TAMA300, GEO600, Virgo

Looking Ahead

- One year of integrated data at design sensitivity by the end of 2006.
- Advanced detectors with dramatically improved sensitivity in 2007+.
 - » Advanced LIGO is under active consideration by the National Science Board
- First detections: ???