



# Waveburst sensitivity study and event property reconstruction on S2 playground data and MDC frames

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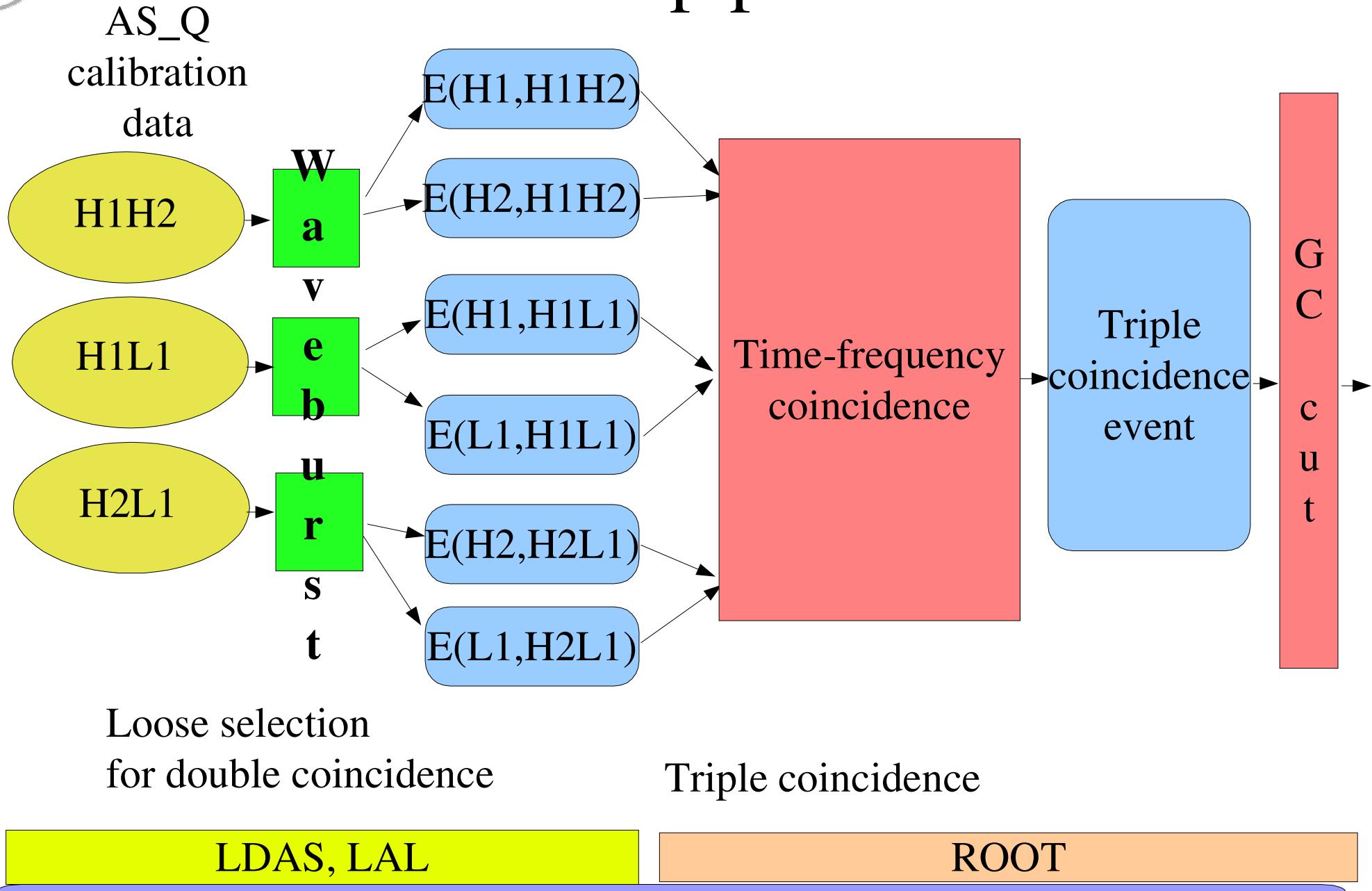
LSC meeting, March 2004

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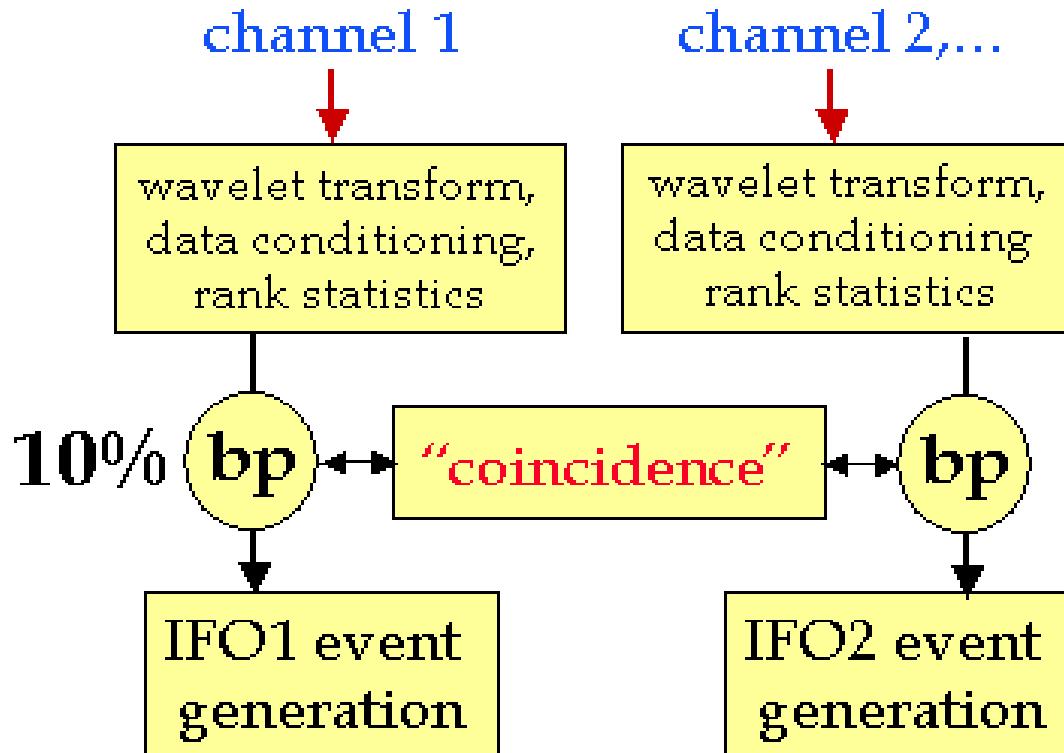
# Outline

- Waveburst data pipeline and algorithm
- Triple coincidence
- Combined confidence threshold
  - Definition of geometric confidence
  - Detection efficiency vs geometric confidence threshold
- Efficiency for sine-gaussian, gaussian, bh-bh merger MDC frames
- Central time, frequency and strain reconstruction
- Future plans

# Waveburst pipeline



# Waveburst ETG



## Settings:

time stride	- 120 sec
F resolution	- 64 Hz
T resolution	- 1/128 sec
band	<4096 Hz

NO data conditioning before wavelet transformation  
tested: no effect from LPF

The output of a single waveburst job are two independently reconstructed clusters that overlap in time and frequency.

# Time window

- (X,Y) average time

$$T_{XY} = \frac{T_{XY}^X + T_{XY}^Y}{2}$$

- Time coincidence:

$$|T_{L_1H_1} - T_{H_1H_2}| \leq T_w$$

$$|T_{L_1H_1} - T_{H_2L_1}| \leq T_w$$

$$|T_{H_2L_1} - T_{H_1H_2}| \leq T_w$$

- Frequency coincidence:



L1,H1



H1,H2



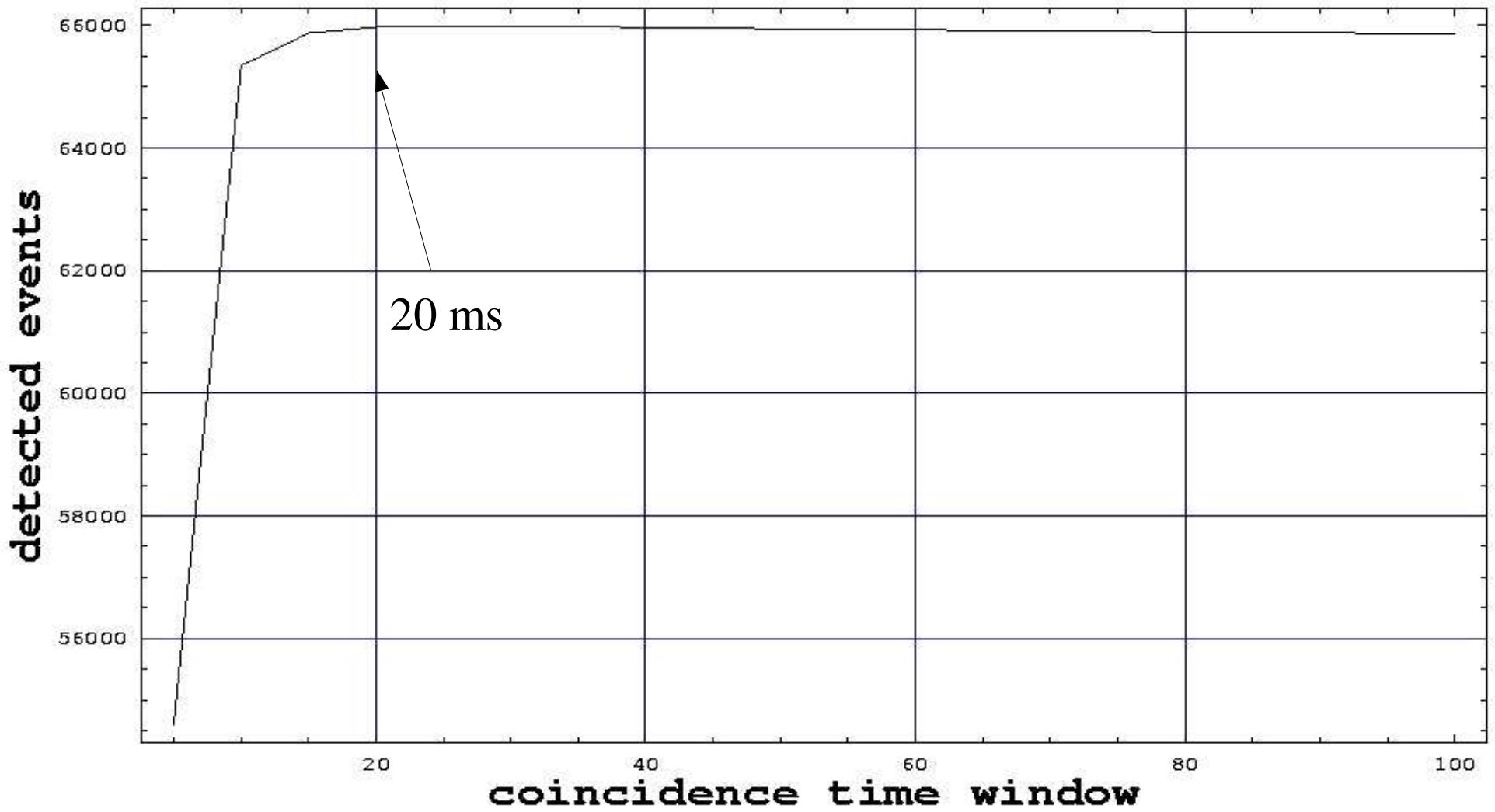
H2,L1

$T_w = 20ms$

zero tolerance overlap

# Time window

**Number of detected injections vs coincidence time window**



# Pixel and cluster significance

- After 10% of the strongest pixels is selected on each frequency layer, the strength of the remaining pixels are characterized as follows:
  - Sort pixels according to the absolute value of their wavelet amplitudes on each frequency layer, if a pixel's rank in this list is  $R$  and there are  $nP$  nonzero pixels on the frequency layer, define **non-parametric significance** of this pixel as

$$Y = -\ln \left( \frac{R}{nP} \right)$$

Non parametric  
significance of a cluster

$$Y(\text{cluster}, k) = \sum_{i=0}^k Y_i$$

# Cluster confidence

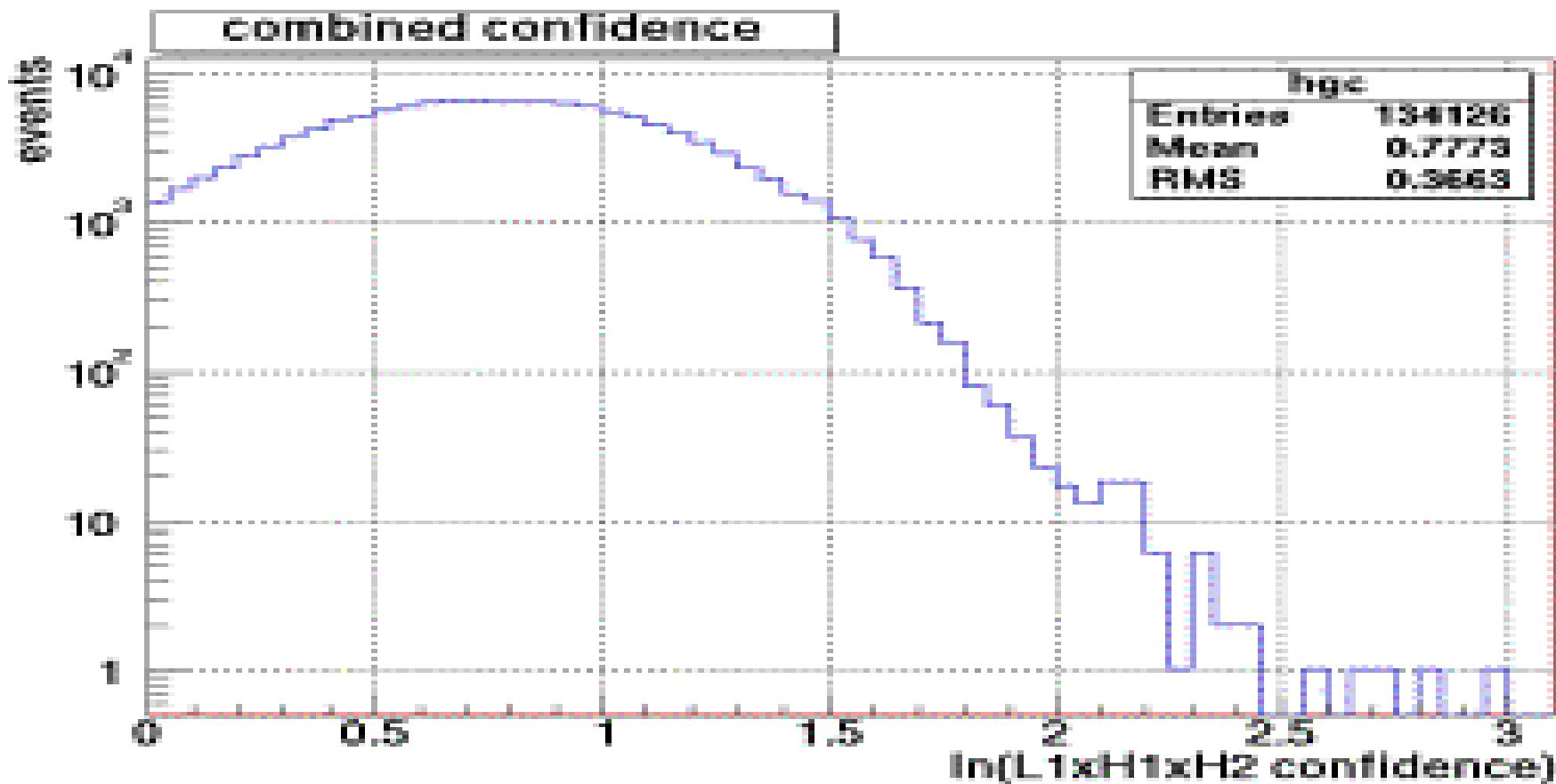
- Since pixel's significance is distributed exponentially, significance of a cluster of size k has gamma distribution.
- **Cluster confidence** is defined by its significance Y and size k and is a measure of how unlikely would it be for statistically independent pixels to form a given cluster:

$$C(Y, k) = -\ln \left( \frac{1}{\Gamma(k)} \int_Y^{\infty} x^{k-1} e^{-x} dx \right)$$

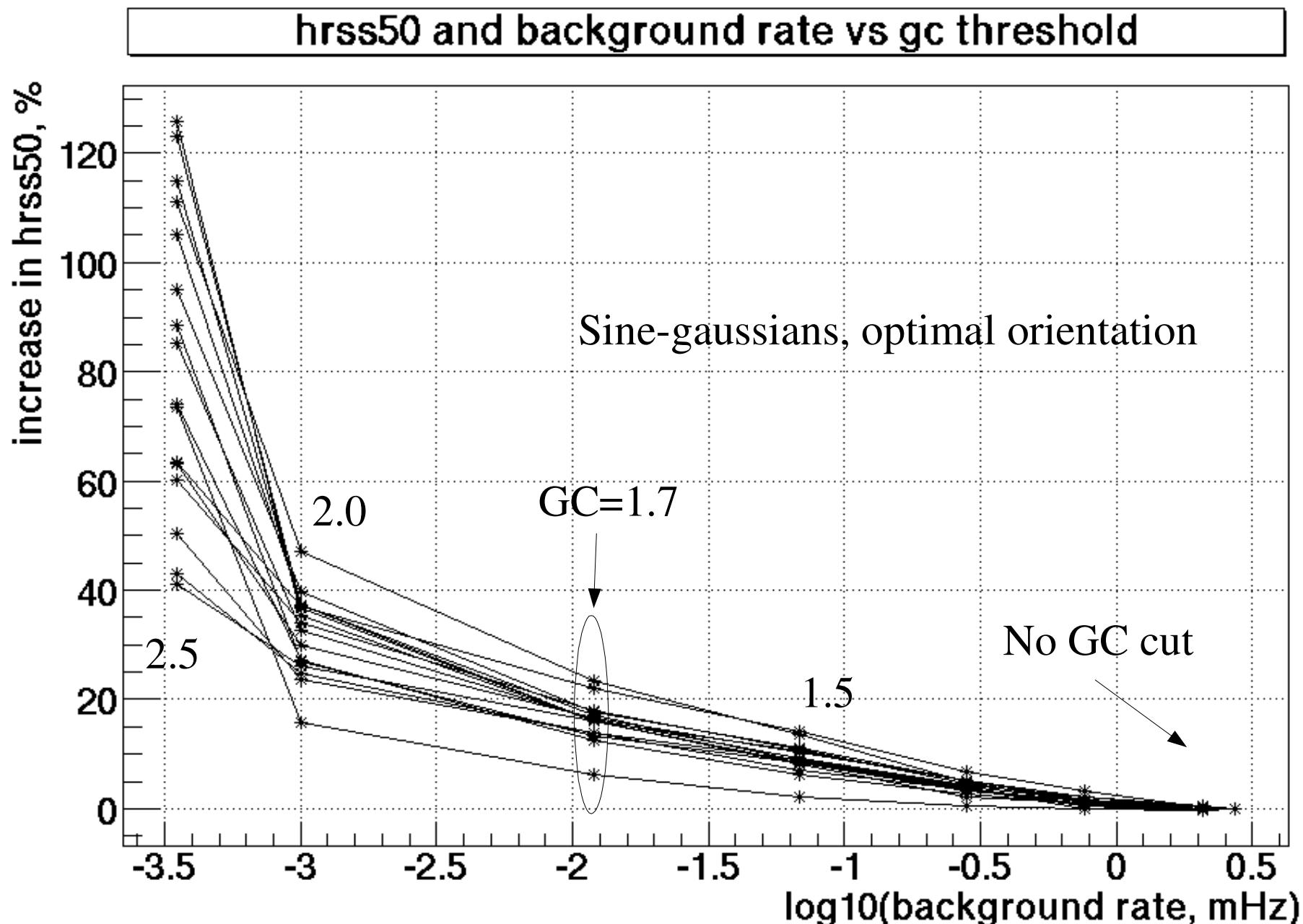
# Geometric confidence

- After time window is applied we get 6 clusters for each event. We define its combined **geometric confidence**:

$$C_G = \sqrt[6]{C_{H1L1}^{L1} C_{H1L1}^{H1} C_{H1H2}^{H1} C_{H1H2}^{H2} C_{H2L1}^{H2} C_{H2L1}^{L1}}$$

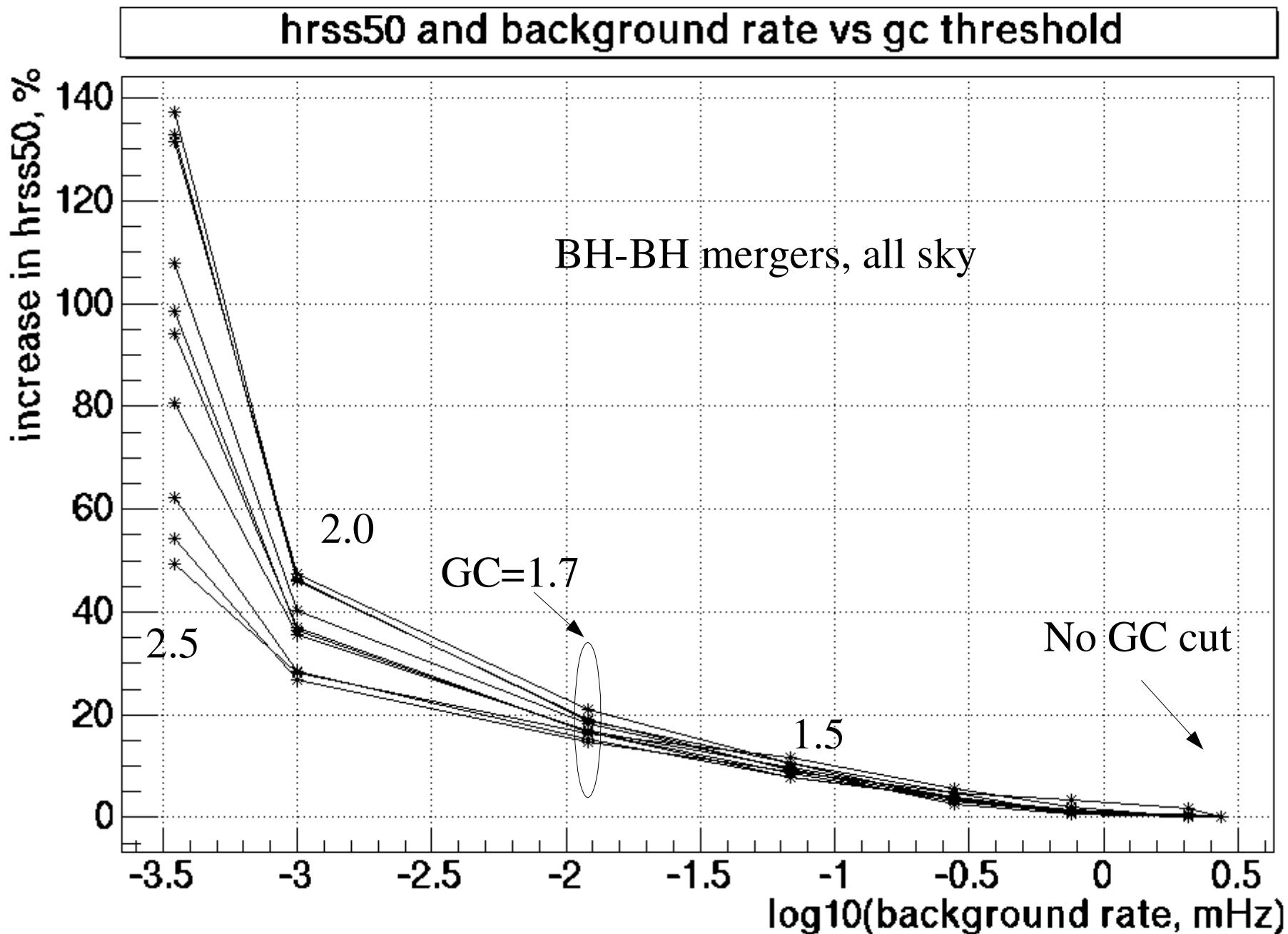


# Choosing geometric confidence threshold



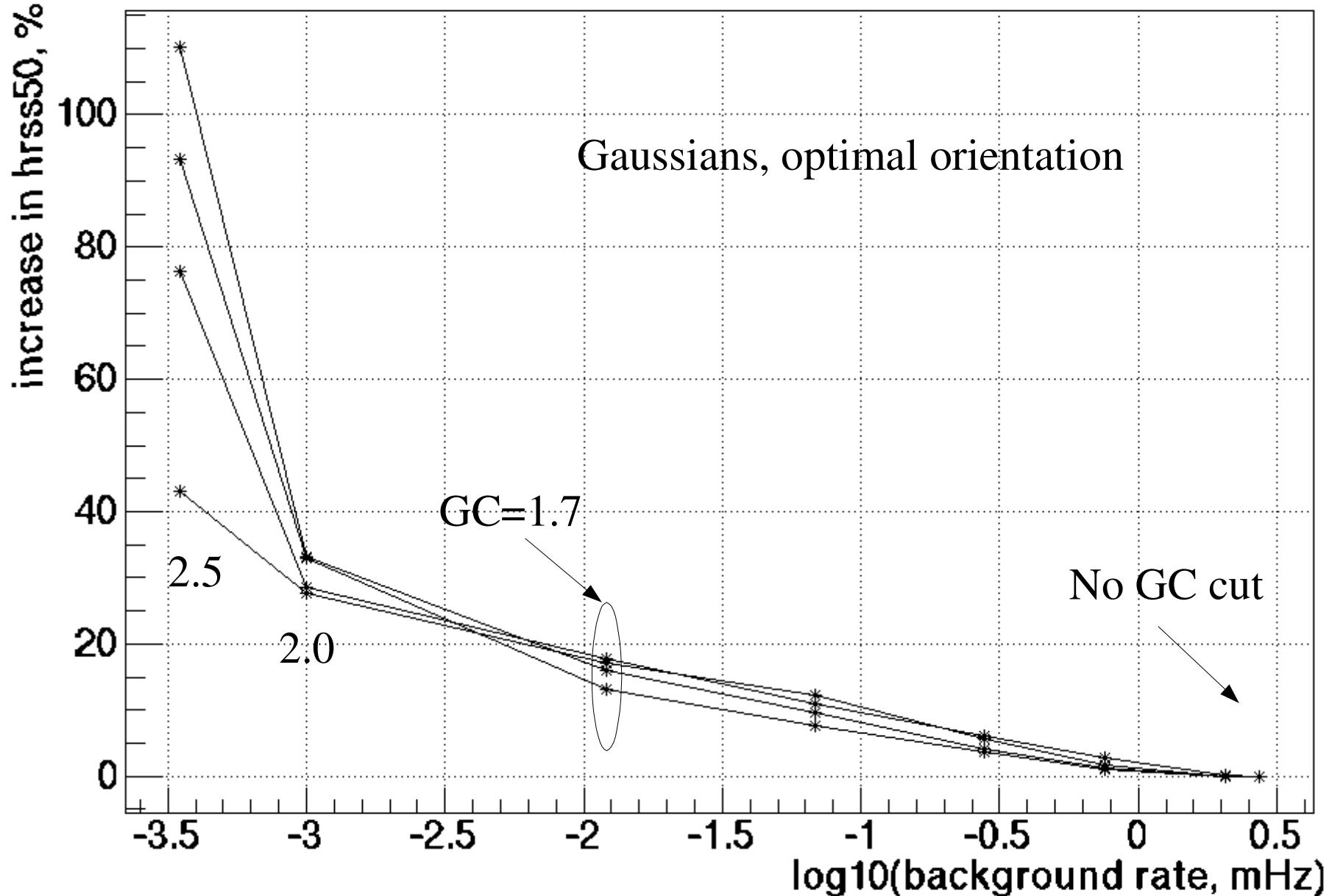


# Choosing geometric confidence threshold

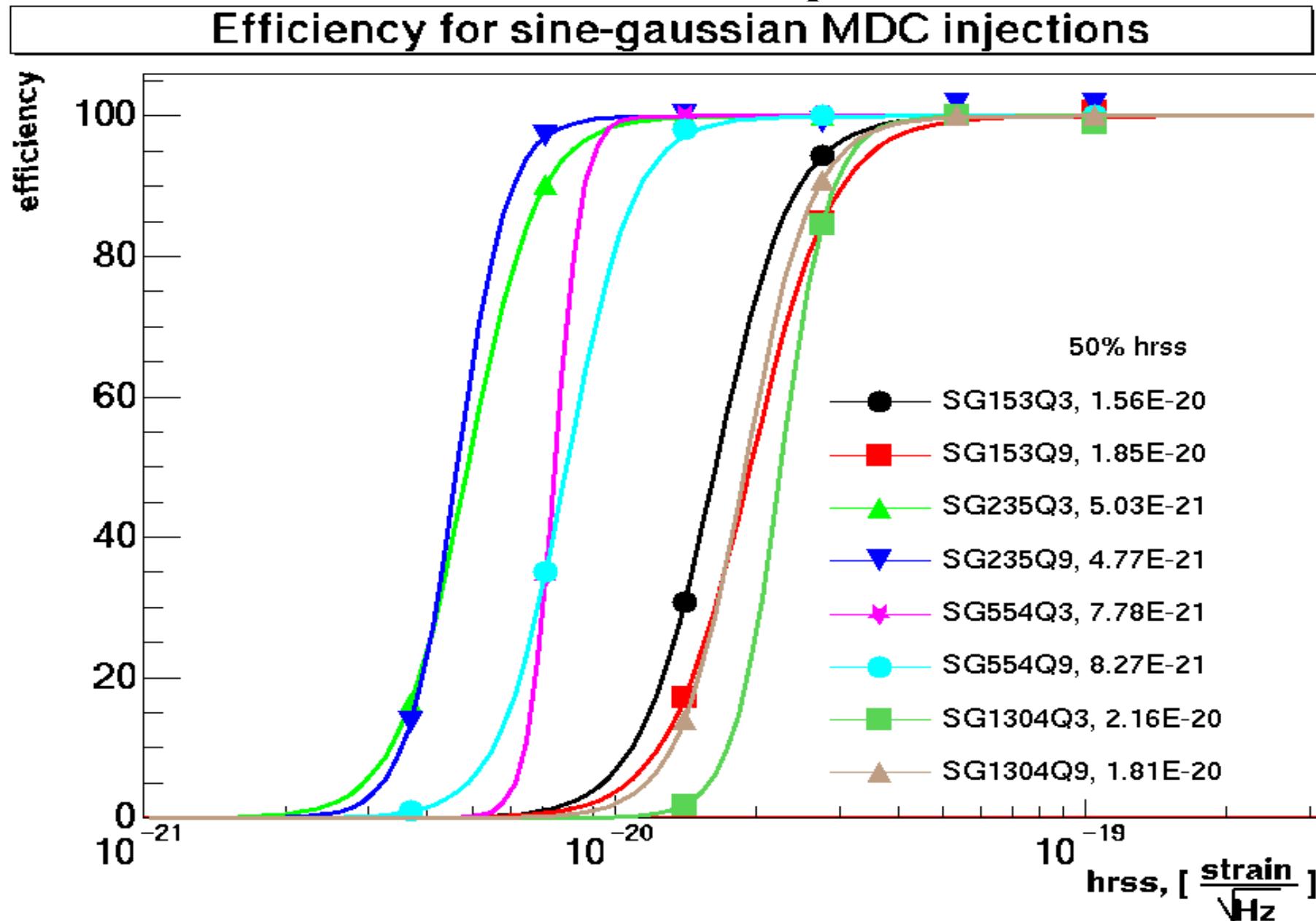


# Choosing geometric confidence threshold

hrss50 and background rate vs gc threshold

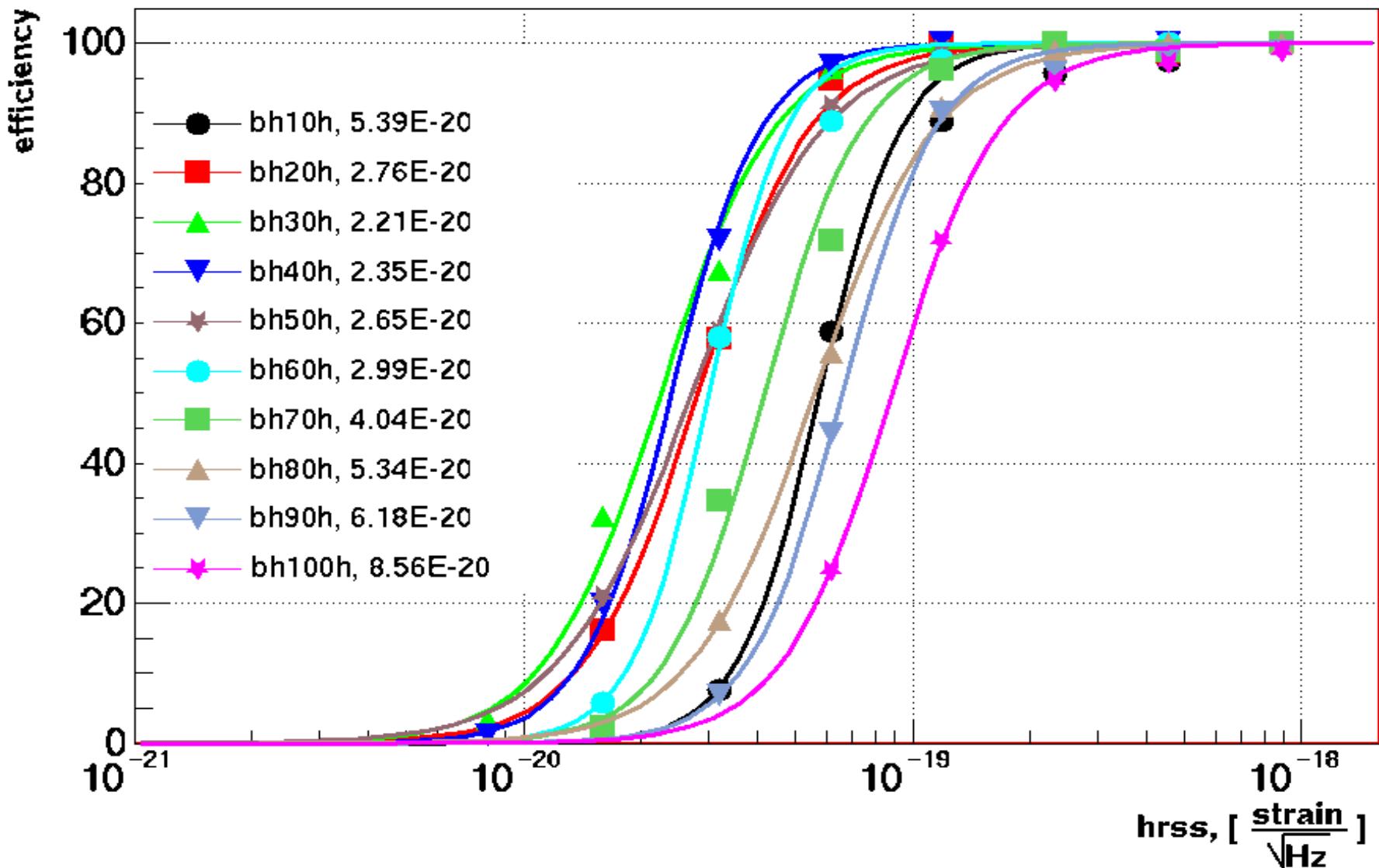


# LIGO Sensitivity: sine-gaussian, optimal orientation, one polarization



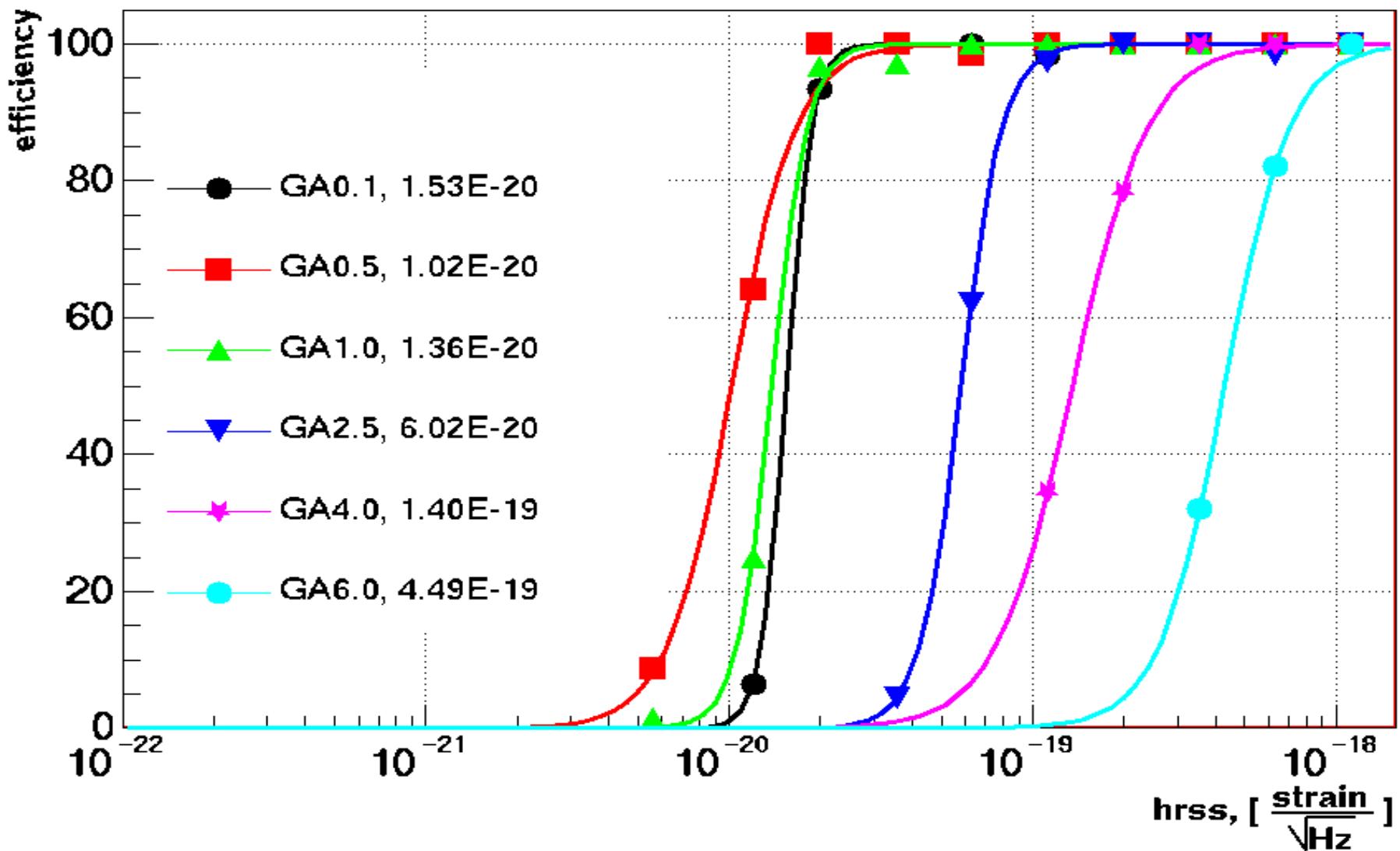
# Sensitivity: BH-BH merger, all sky, two polarizations

## Efficiency for BH–BH merger MDC injections

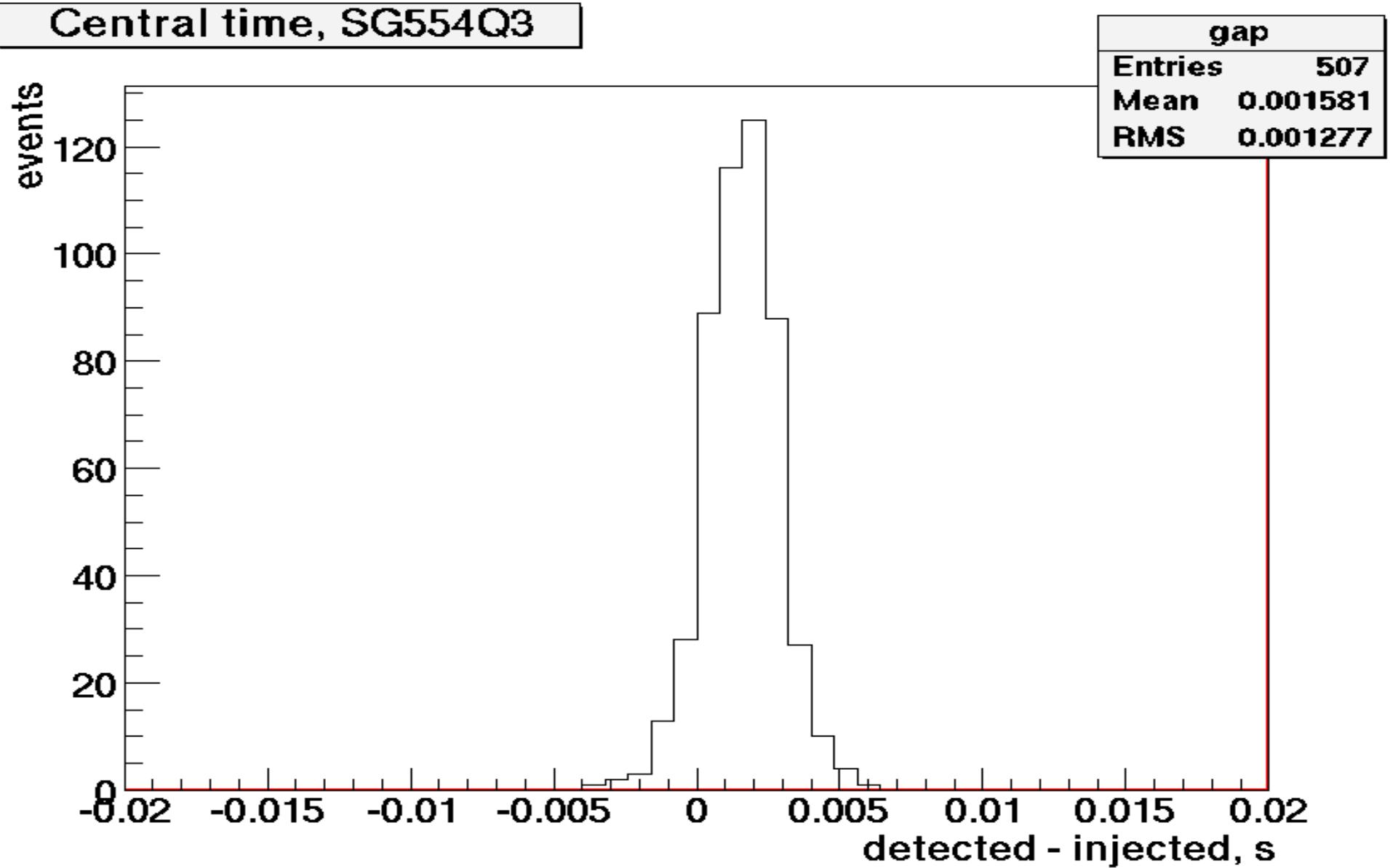


# Sensitivity: gaussian, optimal orientation, one polarization

Efficiency for gaussian MDC injections



# Central time reconstruction



# Central time reconstruction

Detected-injected central time, ms, mean(std)

## Sine-gaussian

<b>Q\F</b>	<b>153</b>	<b>235</b>	<b>554</b>	<b>1304</b>
<b>3</b>	0.01(2.3)	3.1(1.4)	1.6(1.3)	-0.3(1.6)
<b>9</b>	-0.7(2.4)	3.4(1.3)	2.6(1.6)	-0.5(1.5)

## BH-BH merger

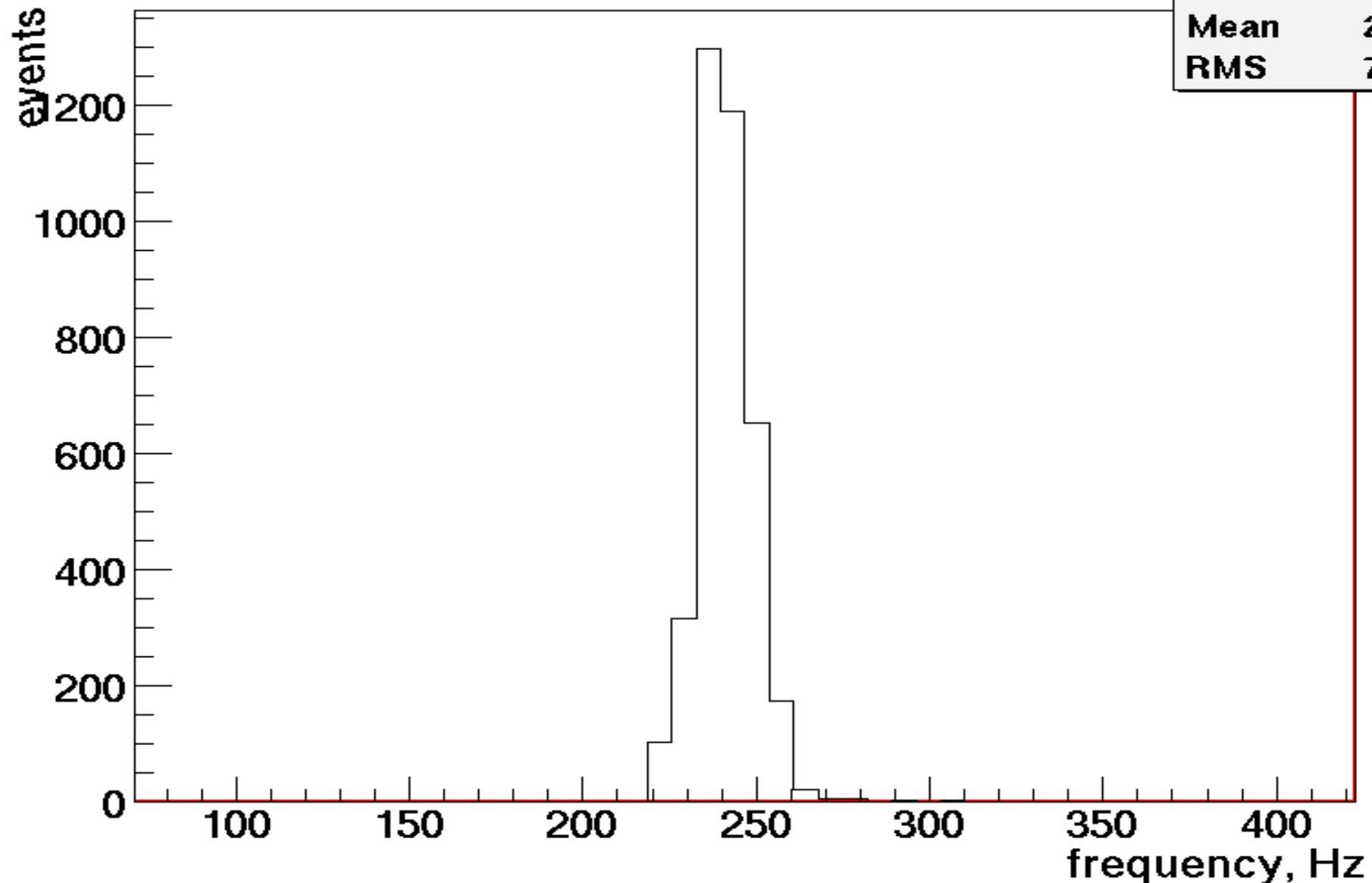
<b>M</b>	<b>10</b>	<b>20</b>	<b>30</b>	<b>40</b>	<b>50</b>	<b>60</b>	<b>70</b>	<b>80</b>	<b>90</b>	<b>100</b>
	0.2(0.8)	0.1(0.9)	1.0(1.4)	1.0(1.6)	2.7(1.5)	4.1(1.5)	4.9(1.6)	4.2(2.4)	3.6(2.4)	3.6(2.3)

## Gaussian

<b>Tau, ms</b>	<b>0.1</b>	<b>0.5</b>	<b>1</b>	<b>2.5</b>	<b>4</b>	<b>6</b>
	1.1(1.1)	1.5(1.3)	2.0(1.3)	2.3(1.9)	3.6(2.3)	4.1(3.3)

# Central frequency reconstruction

Central frequency, SG235Q9





# Central frequency reconstruction

Detected central frequency, Hz, mean(std)

## Sine-gaussians

Q\F	153	235	554	1304
3	166(8)	246(16)	538(22)	1245(43)
9	159(3)	241(8)	541(9)	1297(17)

## BH-BH mergers

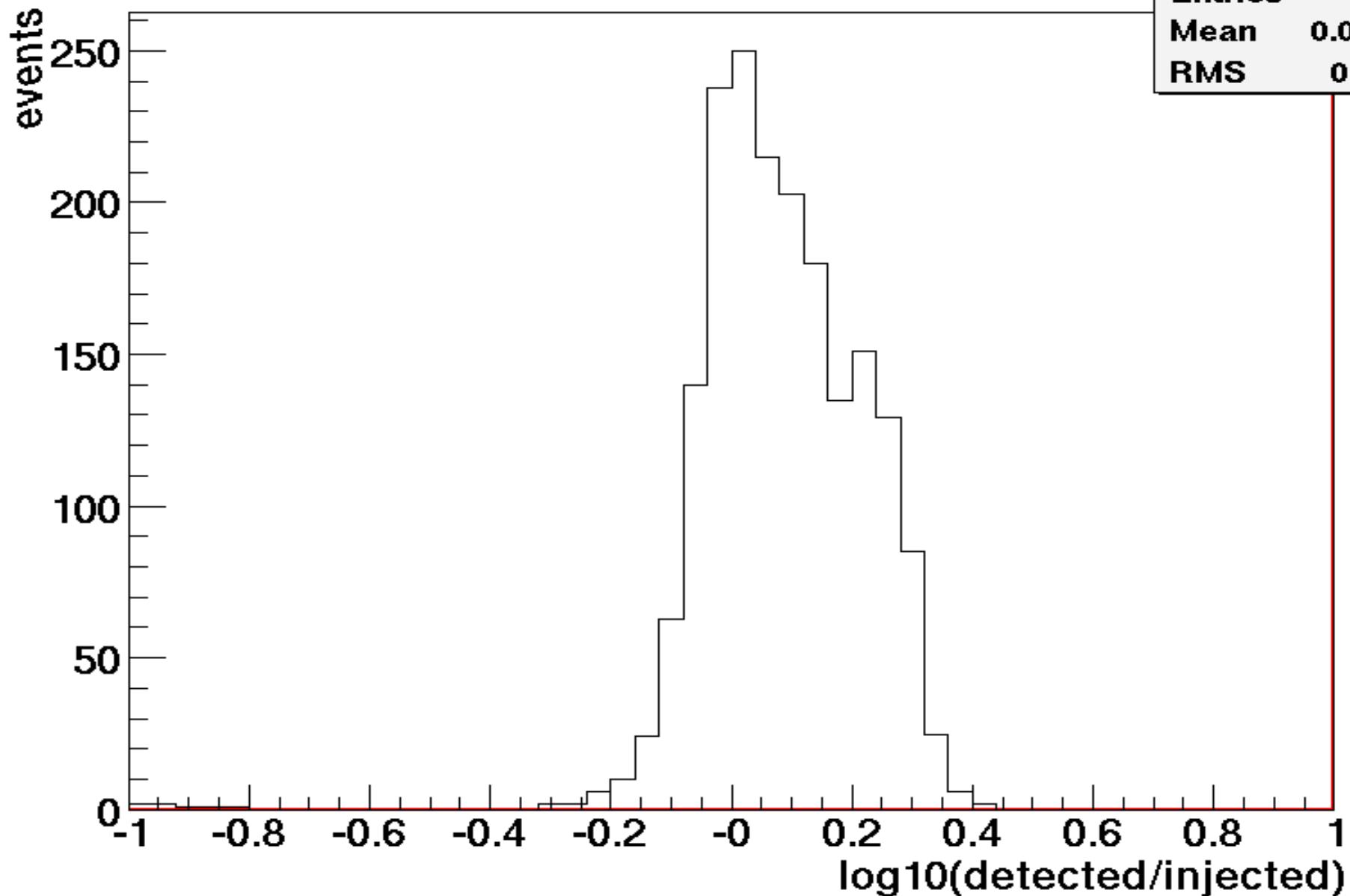
M	10	20	30	40	50	60	70	80	90	100
Detected	856(199)	517(62)	383(37)	310(26)	267(23)	237(22)	220(18)	195(20)	176(24)	163(23)
Injected, x	1132(247)	569(123)	381(82)	286(61)	229(49)	192(41)	165(35)	144(30)	128(27)	134(24)
Injected,+	1330(237)	667(119)	446(79)	334(59)	268(48)	223(40)	192(34)	168(30)	149(26)	115(24)

## Gaussians

Tau, ms	0.1	0.5	1	2.5	4	6
Detected	505(197)	318(40)	250(25)	135(24)	93(8)	103(2)
Injected	899(678)	180(136)	90(68)	36(27)	23(17)	16(11)

# Strain reconstruction

Strain, SG235Q9



# Strain reconstruction

$\log_{10}(\text{detected strain}/\text{injected strain})$ , mean(std)

## Sine-gaussians

Q\F	153	235	554	1304
3	-0.1(0.2)	0.1(0.1)	0.03(0.09)	-0.1(0.1)
9	-0.02(0.19)	0.08(0.1)	0.03(0.08)	-0.07(0.08)

## BH-BH mergers

M	10	20	30	40	50	60	70	80	90	100
	-0.5(0.2)	-0.4(0.2)	-0.3(0.2)	-0.3(0.2)	-0.3(0.2)	-0.3(0.2)	-0.4(0.2)	-0.5(0.2)	-0.5(0.2)	-0.6(0.2)

## Gaussians

Tau, ms	0.1	0.5	1	2.5	4	6
	-0.12(0.12)	0.006(0.1)	-0.11(0.12)	-0.57(0.18)	-0.66(0.16)	-1.13(0.16)

# Future plans

- S2 analysis:
  - Complete simulations on the whole S2;
  - Prepare set of tests for waveburst software validation;
  - “Low threshold” analysis.
- S3 analysis.
- Add multiresolution wavelet transform to waveburst to optimize it for black hole merger search.
- Make a grid version of waveburst.
- Explore combination of DMT and Condor as a possible future framework for waveburst.
- Work out waveform reconstruction of the detected events.

# References

1. T040040-00-Z. WaveBurst. S. Klimenko, I. Yakushin,  
G. Mitselmakher
2. Burst MDC page  
<http://www.ligo-la.caltech.edu/~igor/MDC>
3. WaveBurst S2 simulation page  
<http://www.ligo-la.caltech.edu/~igor/SIMULATION>
4. WaveBurst page  
<http://www.phys.ufl.edu/LIGO/bursts/S2>