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# $h \rightarrow$ AS\_Q Filtering in the Time Domain

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# Outline

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- Goals
- Calibration review
- Mindless first attempts
- Response to problems
- Status/Future

# Motivation

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- Ability of event trigger generators to detect gravitational wave signals is tested by adding simulated signals to the AS\_Q data.
- Efficiency tests need mechanism to convert simulated signals from units of strain to AS\_Q counts.
- BlockNormal bursts ETG is a time-domain search tool; we would like to have time-domain  $h \rightarrow$  AS\_Q filters.
- Other nice properties of time domain filters:
  - » causal
  - » can use online
  - » good for long data streams

# Inverse Calibration ( $h \rightarrow AS\_Q$ )

- Inverse calibration in the frequency domain:

$$X_{AS\_Q}(f) = \frac{\alpha(t) C(f)}{1 + \alpha(t)\beta(t) G(f)} X_h(f)$$

- » Open-loop gain  $G$ , sensing function  $C$  are functions of frequency only.
- » Optical gain  $\alpha$  and DARM gain  $\beta$  are functions of time only.
- Goal: Construct digital time-domain filters  $[b,a]$  with this frequency response for any  $\alpha\beta$ .

# Calibration vs. Inverse Calibration

- Calibration function  $AS\_Q \rightarrow h$  is time-dependent sum of two fixed filters:

$$\frac{1 + \alpha\beta G}{\alpha C} = \frac{1}{\alpha C} + \frac{\beta G}{C}$$

- Inverse calibration function  $h \rightarrow AS\_Q$ :

$$\frac{\alpha C}{1 + \alpha\beta G} \neq \frac{\alpha C}{1} + \frac{C}{\beta G(f)}$$

Must mass-produce  $h \rightarrow AS\_Q$  filters (every  $\alpha\beta$  in  $0.5 < \alpha\beta < 2.5$  ).

# Simple-Minded Approach

- Let MatLab do the work:

```
>> help invfreqz
```

```
INVFREQZ Discrete filter least squares fit to  
frequency response data.
```

```
[B,A] = INVFREQZ(H,W,NB,NA) gives real numerator and  
denominator coefficients B and A of orders NB and NA  
respectively, where H is the desired complex  
frequency response of the system at frequency points  
W, and W contains the normalized frequency values  
within the interval [0, Pi] (W is in units of  
radians/sample).
```

# Knobs

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- Knobs:
  - » The orders  $n_b$ ,  $n_a$  for the desired filter.
  - » The relative weighting  $w(f)$  of various frequencies in fitting the filter to the target inverse calibration function.
  - » Free to play with transfer function outside of frequency range of interest.
- Questions:
  - » What filter orders give good fits?
  - » What frequency weights give good fits?
  - » How does the quality of the fit vary with  $\alpha\beta$ ?

# Filter Quality

- Use fractional RMS error:

$$\text{Error} \equiv \sqrt{\frac{1}{N} \sum \left| \frac{Cb_{\text{filt}} - Cb}{Cb} \right|^2}$$

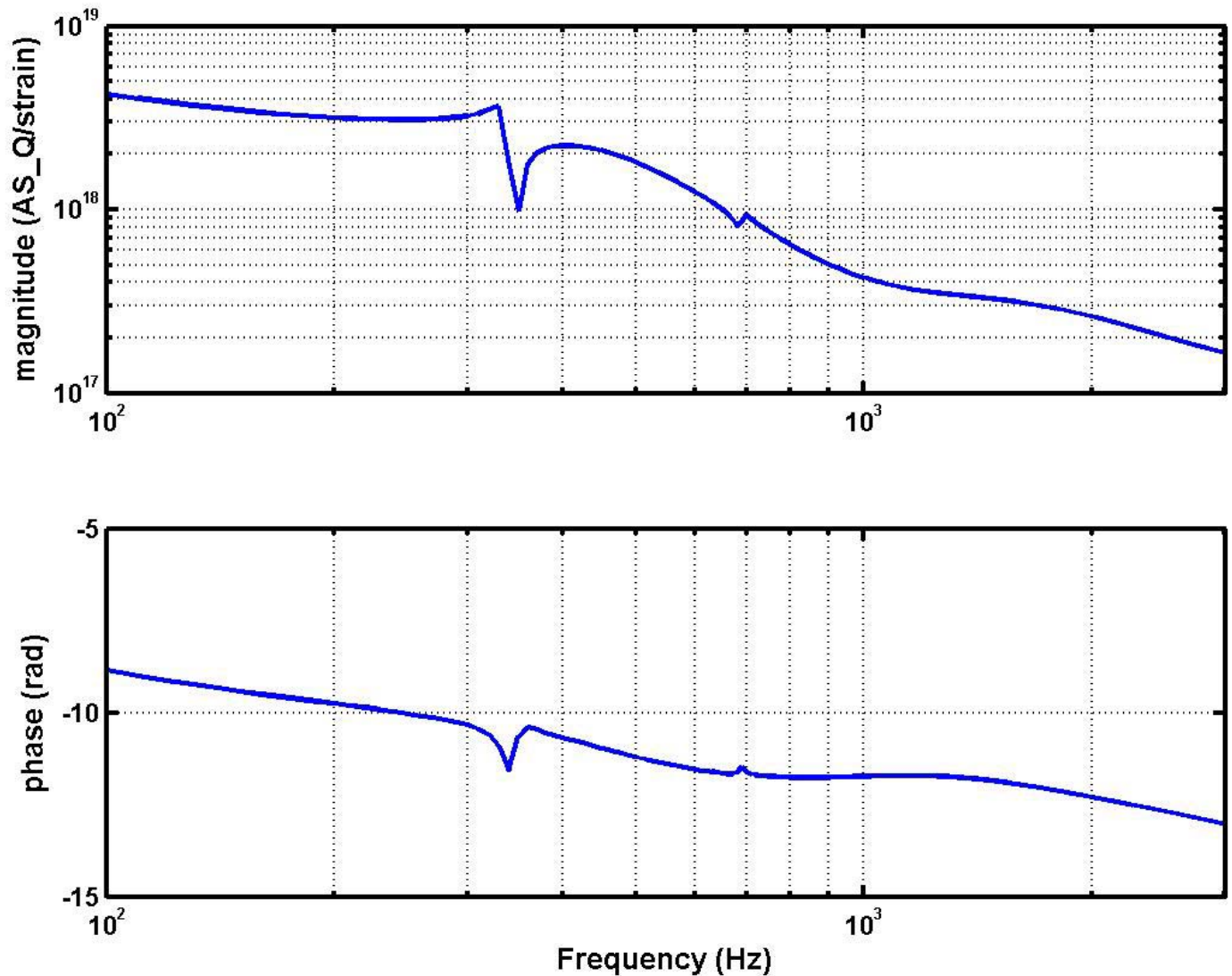
Compare desired to actual frequency response of filter at N frequencies. Ignore frequencies outside of the bands of interest ( $\sim[128,2048]$ Hz minus violin-mode harmonics).

- Desire error  $\leq$  few %.



# Example for H1:

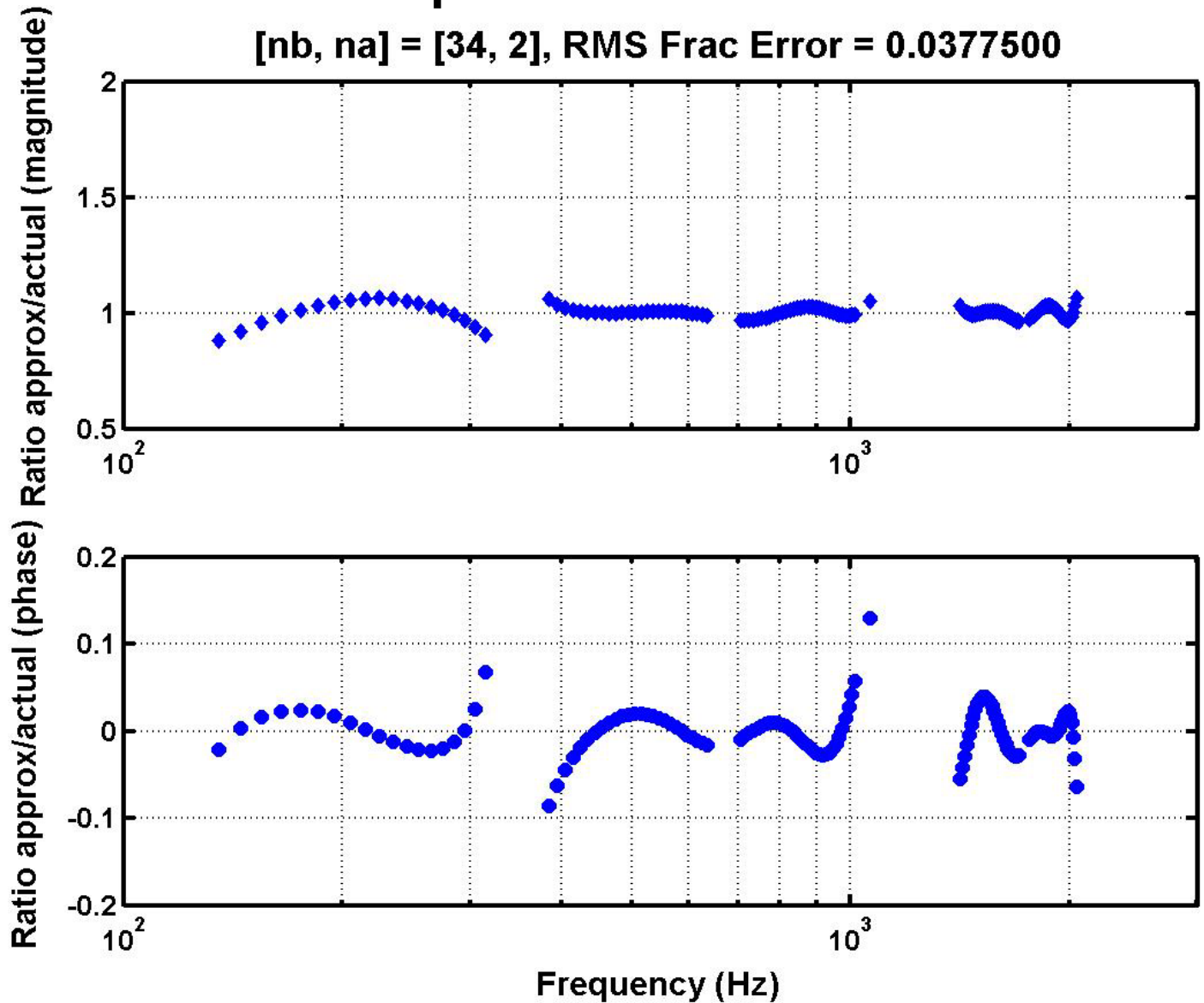
Transfer  
function  
to match:



# Example for H1:

[nb, na] = [34, 2], RMS Frac Error = 0.0377500

How  
well  
we do:



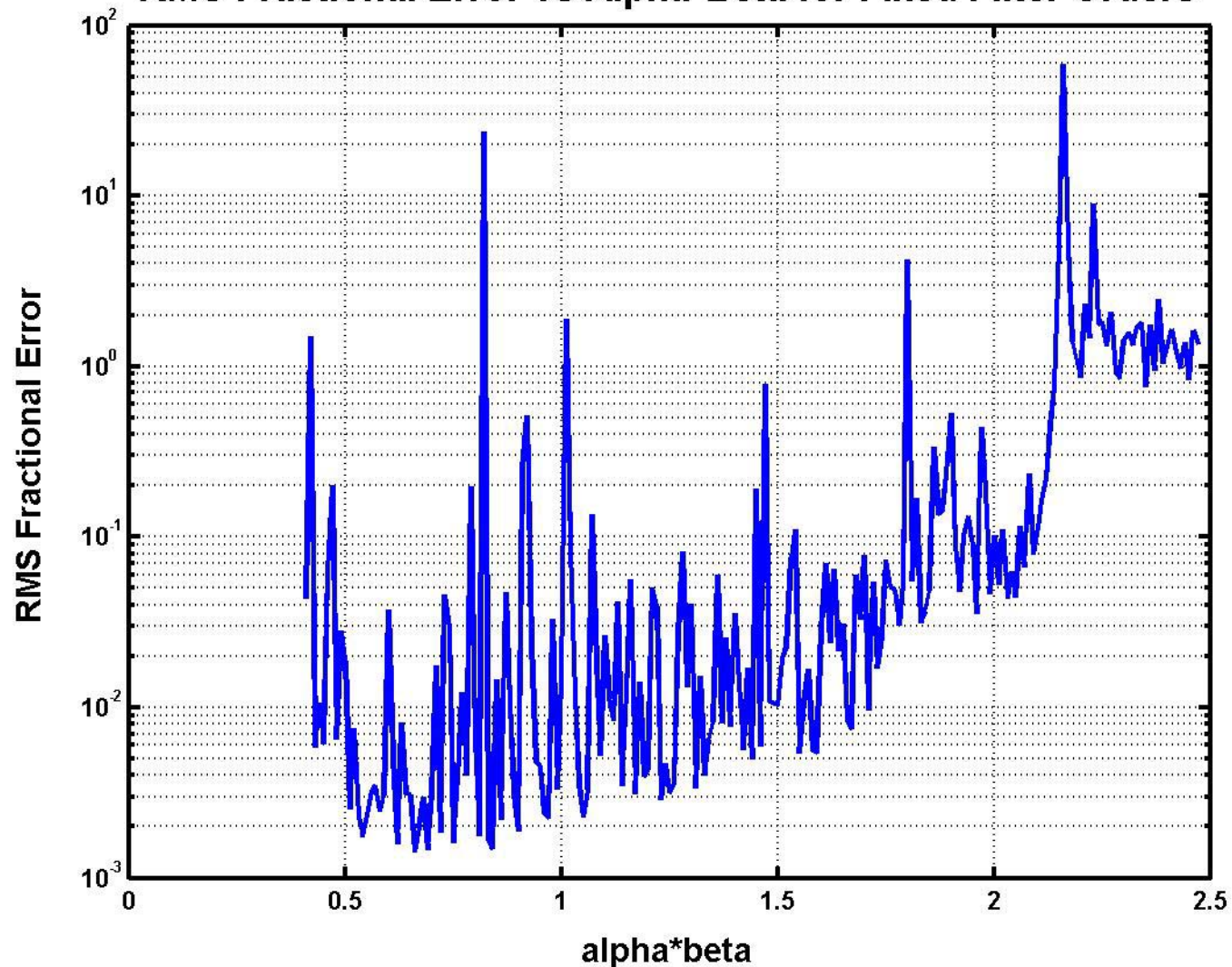
# Problem solved?

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- Can guarantee `invfreqz` filters are stable.
- *Not* guaranteed to be accurate. Vary  $\alpha\beta$ , see what happens...

# Disaster!

RMS Fractional Error vs Alpha\*Beta for Fixed Filter Orders



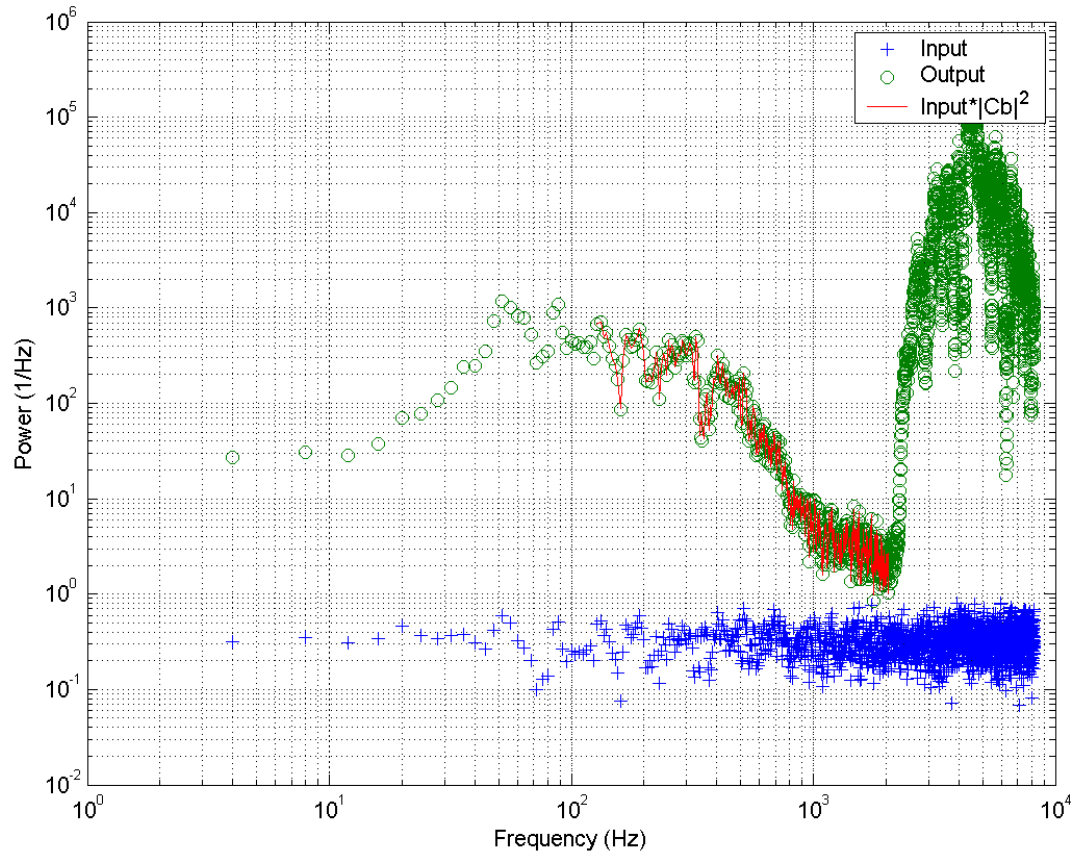
Vary  $\alpha\beta$ ,  
see what  
happens...

# Implications

- **invfreqz** is not robust: changes of a few percent in  $\alpha\beta$  can lead to filters with errors that differ by orders of magnitude.
- Problem is generic across filter orders  $n_b$ ,  $n_a$ .
- Reponse:
  - » don't generate filters on-the-fly.
  - » generate a fixed bank of filters for each IFO for closely-spaced  $\alpha\beta$ . Test each for quality before adding to the set.

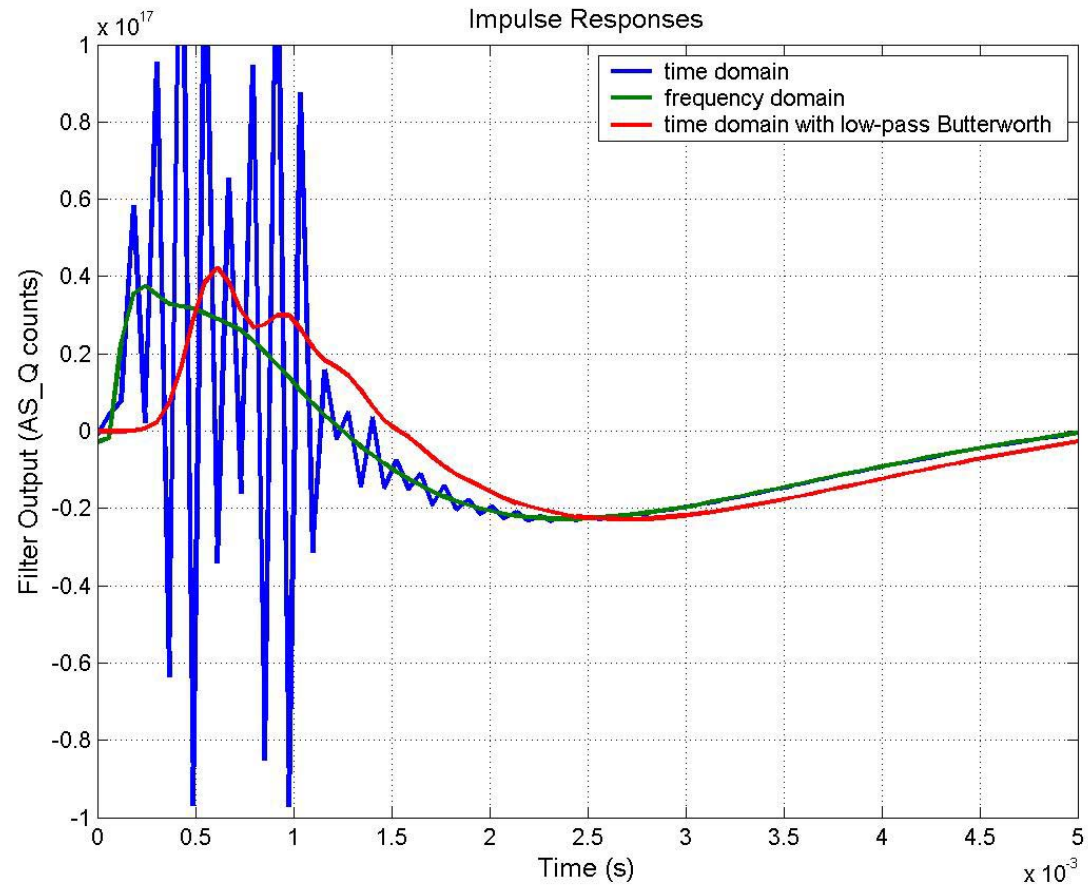
# Next problem:

- Calibration info and/or frequency band of interest usually doesn't go up to Nyquist frequency; can't ignore this region.
- In the frequency domain:



# Next problem:

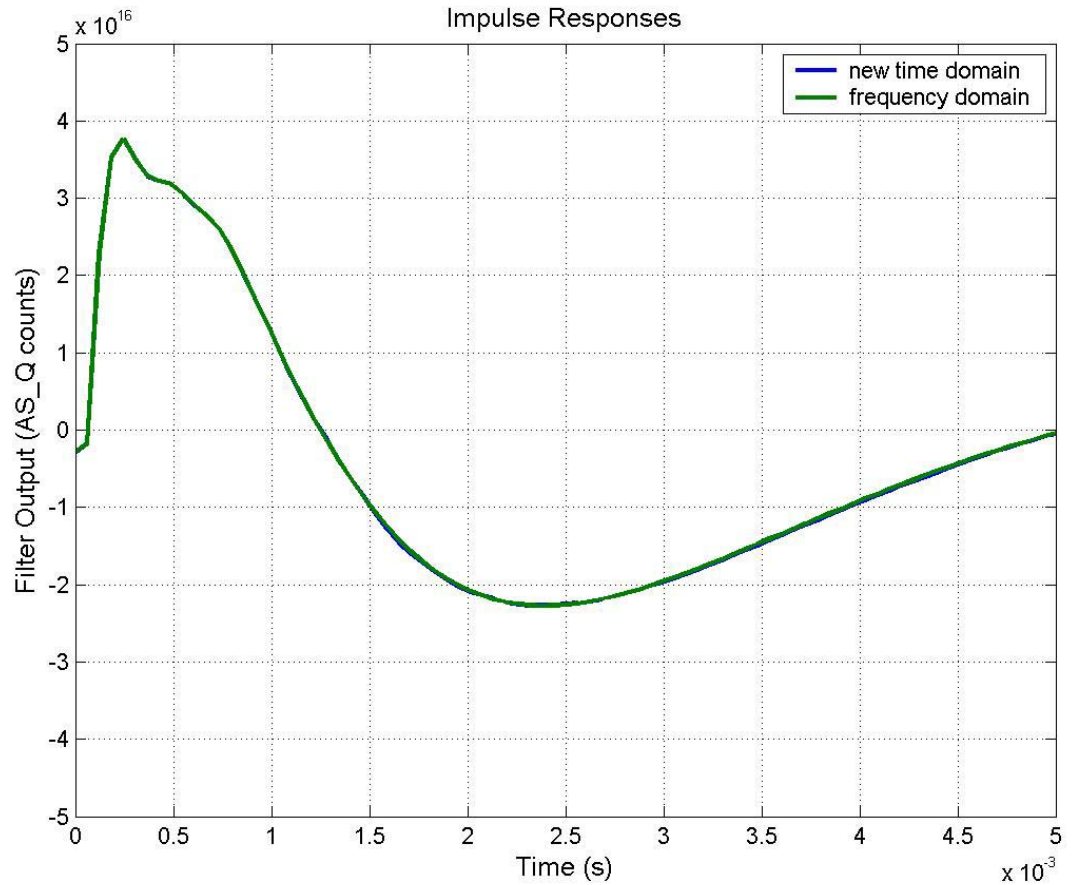
- In the time domain (impulse response):





# Solution:

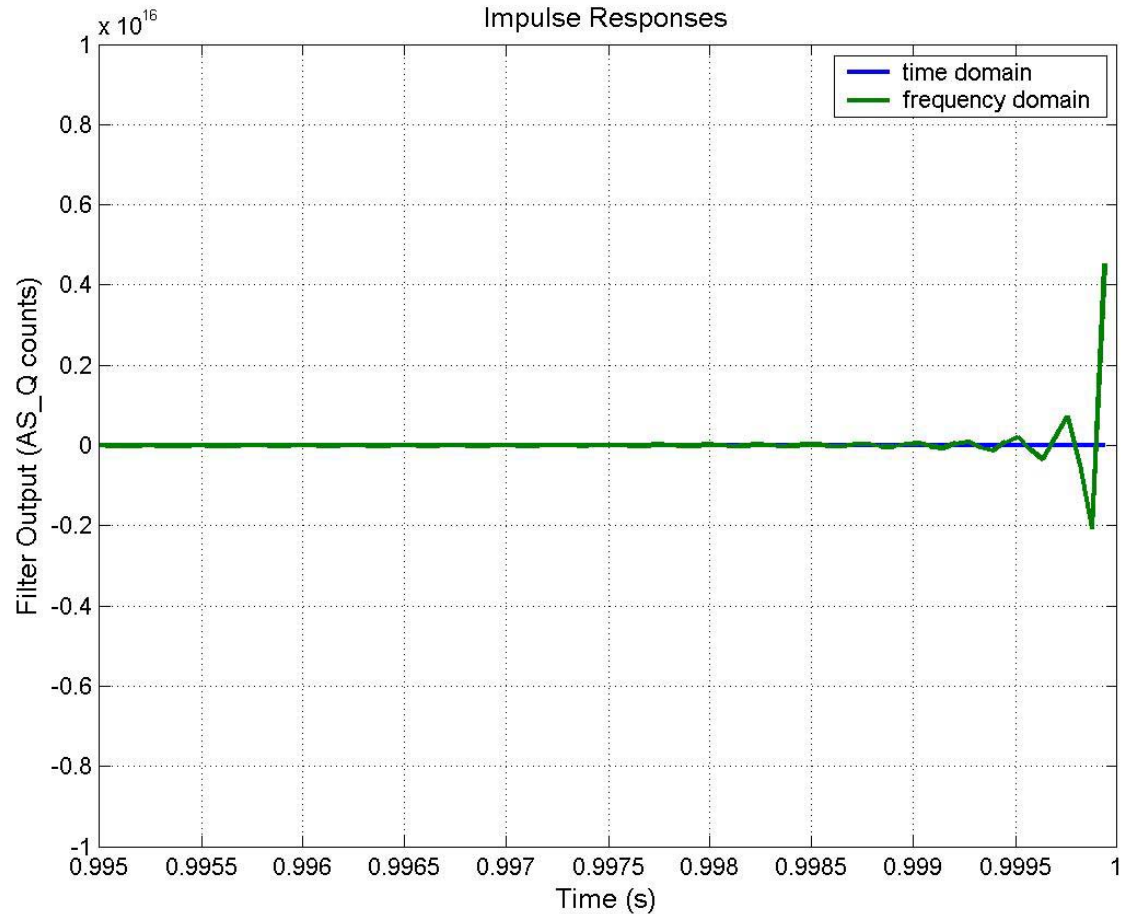
- Extrapolate inverse calibration function smoothly to zero outside frequency range of interest:





# Nice bonus of time domain:

- Filter is causal:



# Summary/Outlook

- Calibration: **good**.
- Time domain calibration: **hard**, but possible if careful.
- Next steps: Generate S2 H1/H2/L1 filter banks for closely spaced values of  $\alpha\beta$ .
- Other possibility: get smart and use knowledge of IFOs. See what Adhikari's MatLab IFO model can do for us.
- Suggestions are welcome!

Log<sub>10</sub> Maximum RMS Fractional Error vs Filter Order

