



Combining the Two Hanford-Livingston Stochastic Background Limits in the Presence of Local Environmental Correlations Between the Hanford Interferometers

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Motivation

- The S1 stochastic analysis exposed environmental correlations between H1 and H2
 - » Precluded use of this measurement for setting an upper limit on the stochastic background
 - » Made combining the H1-L1 and H2-L1 results potentially tricky due to the known H1-H2 correlations
 - H1-L1 and H2-L1 measurements made when the other interferometer was not operating may be added assuming no correlations between the measurements
 - see original Allen&Romano article, also T. Regimbau note in stochastic e-log 02 Nov 2003:
<http://www.lidas-sw.ligo.caltech.edu/ilog/pub/ilog.cgi?group=stochastic>
 - 2X measurements made during periods of 3X coincident operations in general cannot be combined in this way -- subject of this talk
 - see <http://www.ligo.caltech.edu/docs/T/T030250-01.pdf>
 - » Formulation discussed during recent face-to-face meeting at MIT

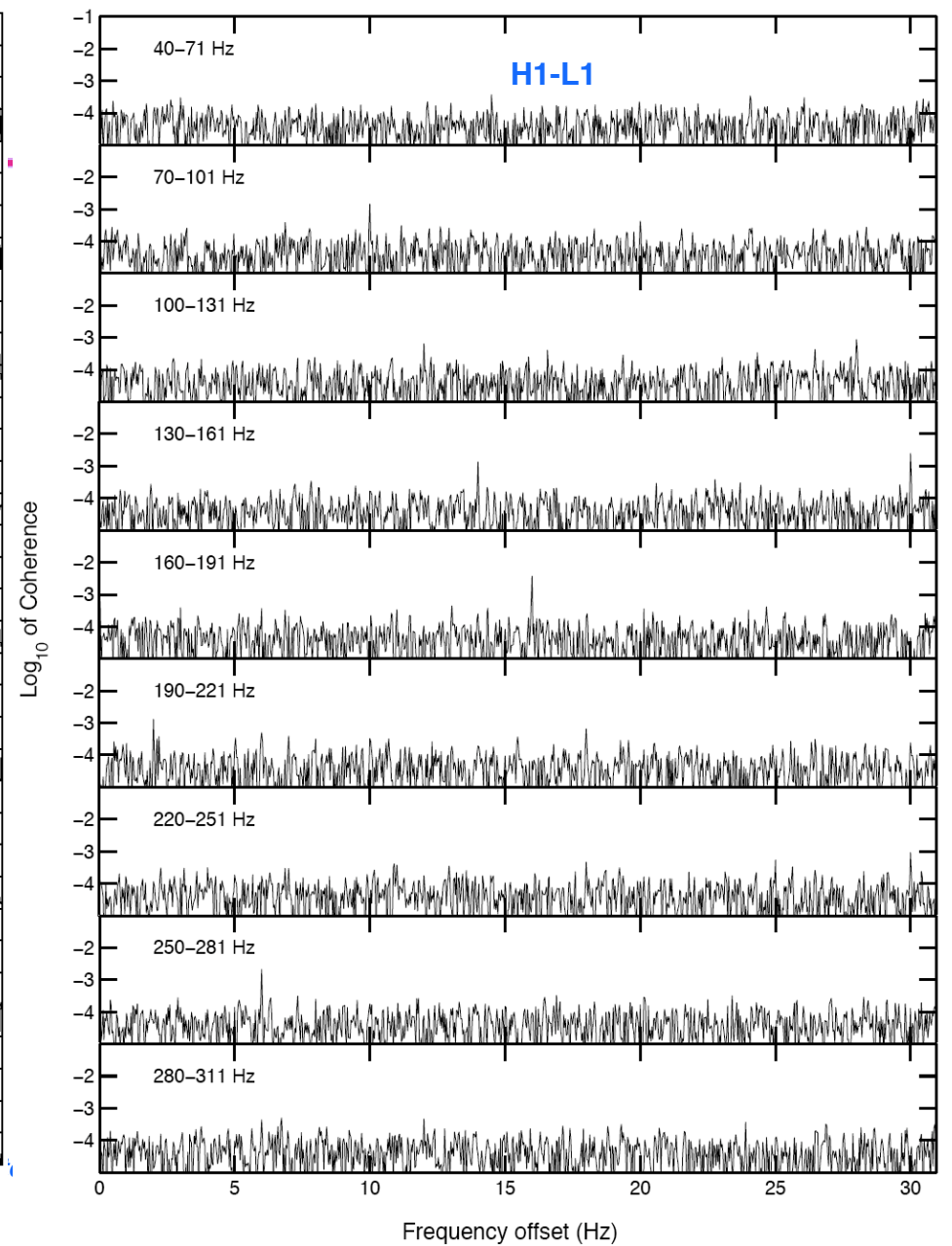
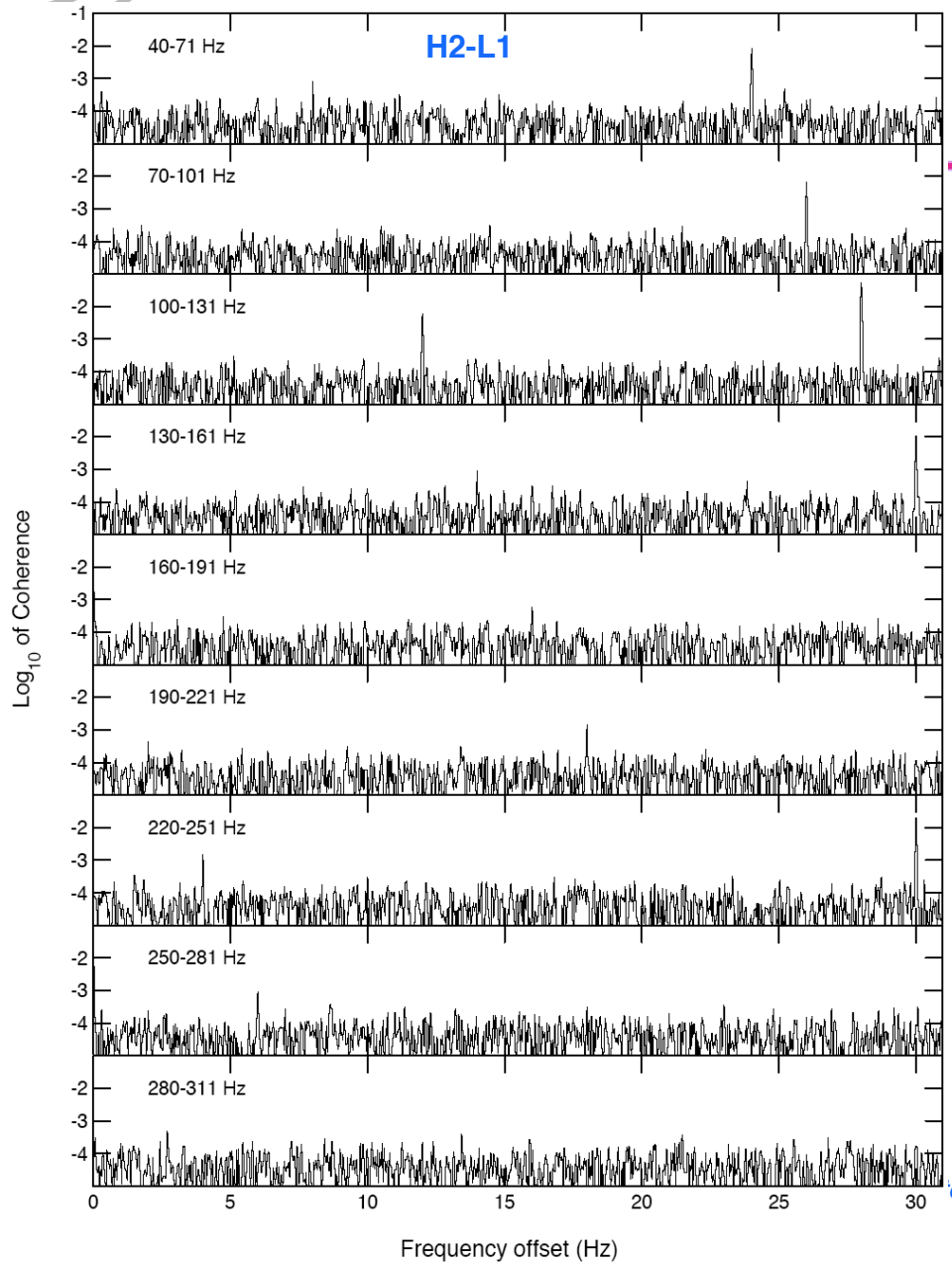


Optimally using the H1-H2-L1 data for stochastic background measurements

- Idea:
 - » First form a Hanford *pseudodetector* data set that takes into account local instrumental and environmental correlations
 - » Then use the *pseudodetector* data in the transcontinental cross-correlation measurement
 - » Naturally combines three interferometer datastreams to produce a *single* H-L estimate
- Assumes:
 - » No sources of broadband correlations between LIGO sites
 - Supported by S1, S2 long-term coherence measurements^{*}
 - * Except for very narrow lines related to GPS timing and DAQS
 - Local H1-H2 coherence is dominated by environment, instrumental noise
 - Supported by character, magnitude of the H1H2 coherence measurements during S1, S2
 - Turns out that so long as H1 and H2 calibrations are *accurate* linearly melding H1 + h2 does not affect GW component

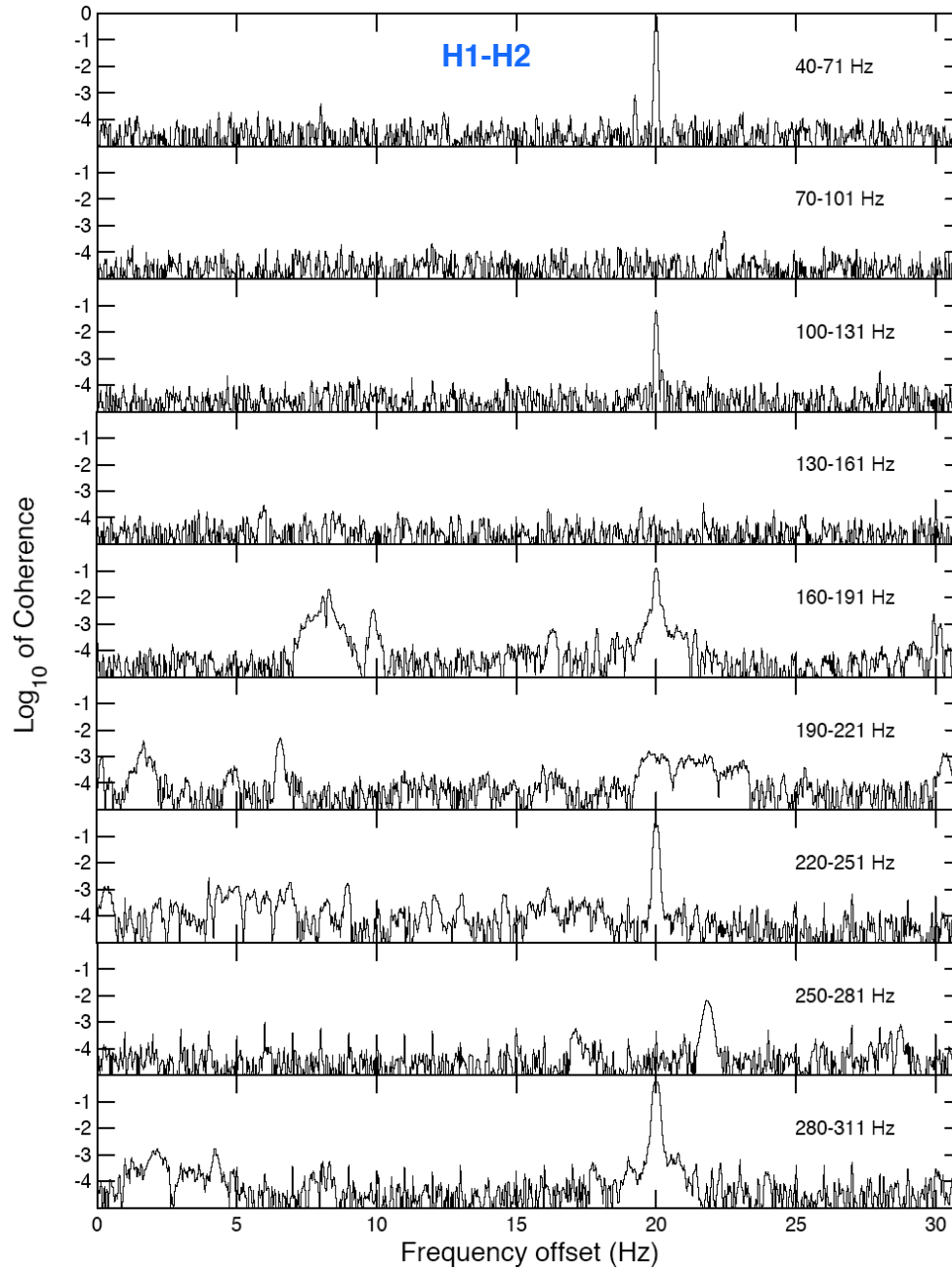


Run-averaged Coherences - S1





Run-averaged Coherences - S1

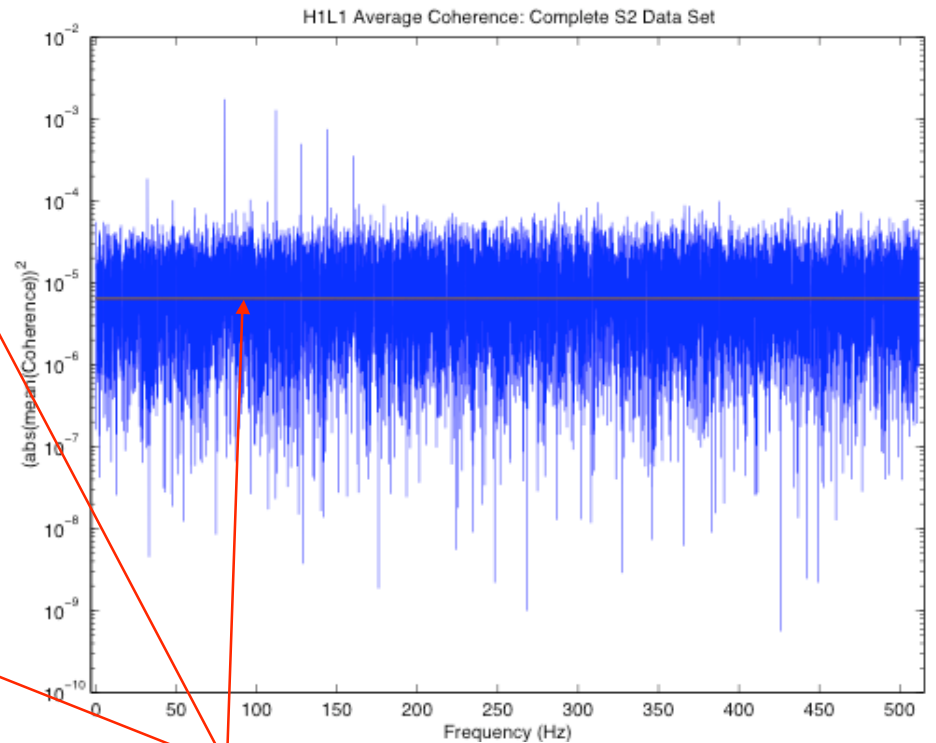
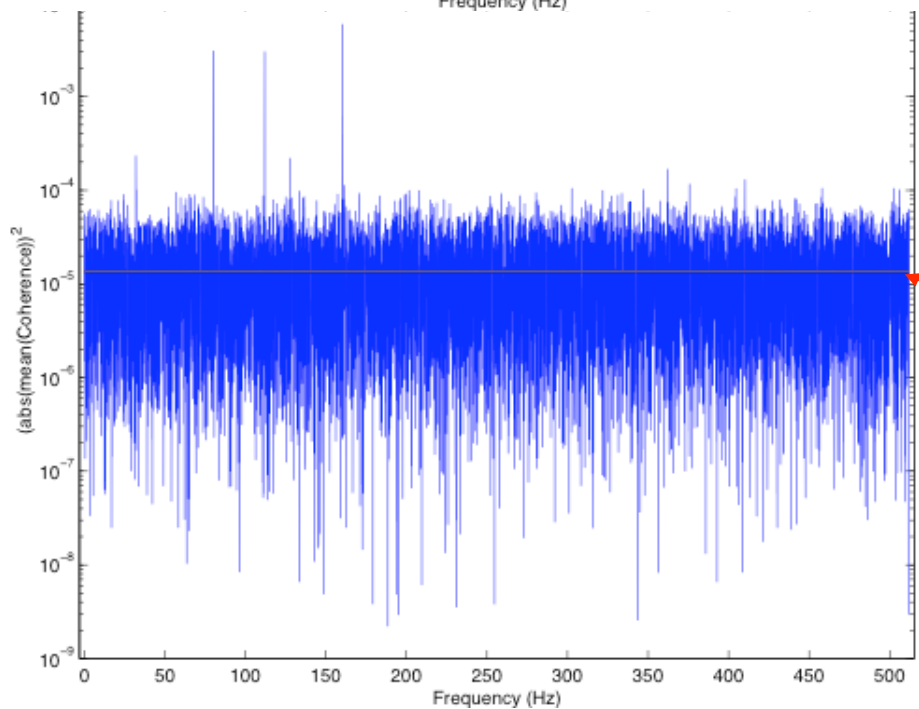
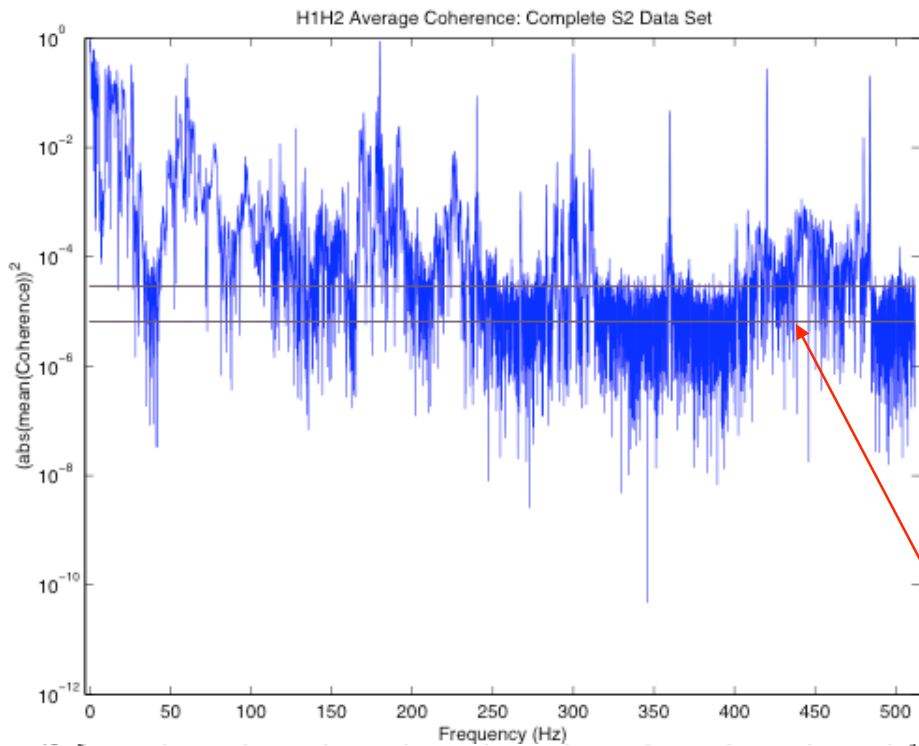


LIGO-G030553-0:

LSC Meeting 2003.

Run-averaged Coherences - S2

S1



Theoretical levels for no correlation after integration time of S2



Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford

$$\begin{aligned} s_{H1}(t) &= h(t) + n_{H1}(t) \\ s_{H2}(t) &= h(t) + n_{H2}(t) \end{aligned} \quad \Longrightarrow \quad \begin{aligned} \tilde{s}_{H1}(f) &= \tilde{h}(f) + \tilde{n}_{H1}(f) \\ \tilde{s}_{H2}(f) &= \tilde{h}(f) + \tilde{n}_{H2}(f) \end{aligned}$$

$$\langle \tilde{n}_i(f) \rangle = \langle \tilde{h}(f) \rangle = 0$$

$$\langle \tilde{n}_i^*(f) \tilde{h}(f) \rangle = 0$$

$$\begin{aligned} \langle \tilde{n}_i^*(f) \tilde{n}_j(f') \rangle &= P_{ij}(f) \delta(f - f')^* \\ &= \rho_{ij}(f) \sqrt{P_i(f) P_j(f)} \times \\ &\quad \delta(f - f') \end{aligned}$$

$$P_{ii}(f) := P_i(f)$$

$$\langle \tilde{h}^*(f) \tilde{h}(f') \rangle = P_\Omega(f) \delta(f - f')$$

$$P_\Omega(f) \ll P_i(f)$$

$$\Gamma_{ij}(f) := |\rho_{ij}(f)|^2$$

Not required for final conclusion to be correct



LIGO Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (2)

- Form linear combination of two interferometer signals:

$$\tilde{s}'(f) = \tilde{\alpha}(f)\tilde{s}_1(f) + \tilde{\beta}(f)\tilde{s}_2(f)$$

- Require s' to be an *unbiased* estimate of h :

$$\begin{aligned}\langle \tilde{h}^*(f) \tilde{s}'(f') \rangle &= P_{\Omega}(f)\delta(f - f') \\ &\rightarrow \tilde{\alpha}(f) + \tilde{\beta}(f) = 1\end{aligned}$$

- Require s' to have *minimum variance*:

$$\begin{aligned}Var(s') &:= P_{s'}(f) \\ \langle \tilde{s}'^*(f) \tilde{s}'(f') \rangle &= P_{s'}(f)\delta(f - f')\end{aligned}$$



LIGO Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (3)

$$P_{s'}(f) = |\tilde{\alpha}(f)|^2 P_1(f) + |1 - \tilde{\alpha}(f)|^2 P_2(f) + \left(\tilde{\alpha}^*(f)(1 - \tilde{\alpha}(f))\rho_{12}(f) + \tilde{\alpha}(f)(1 - \tilde{\alpha}^*(f))\rho_{12}^*(f) \right) \times \sqrt{P_1(f)P_2(f)}$$

- Minimize $P_{s'}$:

$$\begin{pmatrix} \frac{\partial P_{s'}(f)}{\partial \tilde{\alpha}(f)} \\ \frac{\partial P_{s'}(f)}{\partial \tilde{\alpha}^*(f)} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$$



Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (4)

Result for $\tilde{\alpha}$:

$$\tilde{\alpha}(f) = \frac{P_2(f) - \rho_{12}(f)\sqrt{P_1(f)P_2(f)}}{P_1(f) + P_2(f) - (\rho_{12}(f) + \rho_{12}^*(f))\sqrt{P_1(f)P_2(f)}}$$

and for $P_{s'}$:

$$P_{s'}(f) = \frac{P_1(f)P_2(f)(1 - \Gamma(f))}{P_1(f) + P_2(f) - (\rho_{12}(f) + \rho_{12}^*(f))\sqrt{P_1(f)P_2(f)}}$$



LIGO Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (6)

Limits:

I. If $\rho_{12}(f) \rightarrow 0$: Then $\tilde{\alpha}(f)$ becomes,

$$\tilde{\alpha}(f) \rightarrow \frac{P_2(f)}{P_1(f) + P_2(f)}$$

IIa. If $P_1(f) \rightarrow P_2(f)$: Then $\tilde{\alpha}(f)$ is independent of $P(f)$,

$$\tilde{\alpha}(f) \rightarrow \frac{1 - \rho_{12}(f)}{2 - (\rho_{12}(f) + \rho_{12}^*(f))}$$

IIb. If $\rho_{12}(f) \rightarrow 1$, then $\rho_{12}(f) = \rho_{12}^*(f) = \sqrt{\Gamma(f)}$. If also $P_1(f) \rightarrow P_2(f)$, then $P_{s'}(f) \rightarrow P_1(f)$:

$$P_{s'}(f) = \lim_{\Gamma(f) \rightarrow 1} \frac{P_1(f)}{2} \frac{1 - \Gamma(f)}{1 - \sqrt{\Gamma(f)}} = P_1(f)$$



LIGO Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (7)

Limits :

III. For H1 and H2 the limiting design performance will have $P_2(f) = 4P_1(f)$ due to the 1 : 2 arm length ratio,

$$\tilde{\alpha}(f) \rightarrow \frac{2(2 - \rho_{12}(f))}{5 - 2(\rho_{12} + \rho_{12}^*)}$$

If the noise were either completely correlated ($\rho_{12} \rightarrow 1, \tilde{\alpha}(f) \rightarrow 2$) or anti-correlated ($\rho_{12} \rightarrow -1, \tilde{\alpha}(f) \rightarrow \frac{2}{3}$), then it would be possible to exactly cancel the noise in the signals s_i .



Optimal estimate of pseudostrain in the presence of instrumental correlations at Hanford (8)

The correlation kernel for L1- H becomes (assuming $\Gamma_{GW}(f) = \text{const.}$):

$$Y(f) \approx \frac{\gamma(|f|)s_{L_1}^*(f) \left((P_{H_2}(f) - \rho_{H_1H_2}(f)\sqrt{P_{H_1}(f)P_{H_2}(f)})\tilde{s}_{H_1}(f) + (P_{H_1}(f) - \rho_{H_1H_2}^*(f)\sqrt{P_{H_1}(f)P_{H_2}(f)})\tilde{s}_{H_2}(f) \right)}{|f|^3 P_{L_1}(|f|)P_{H_1}(f)P_{H_2}(f)(1 - \Gamma_{H_1H_2}(f))}$$

Implementation issues/details:

- Noise properties of a 4x product of spectra
- Need to modify the correlation code to take in 3 interferometer channels, condition, etc.
- $\Gamma(f)$ and $\Gamma^*(f)$ should be calculated over the entire run -- read them in as frequency series, similar to R(f) data.
- Apply this ONLY to 3X data stretches



Results from S2 representative spectra

