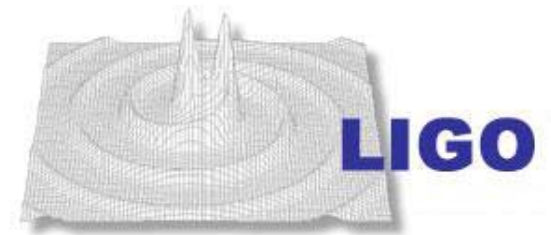


LIGO G030504-00-E

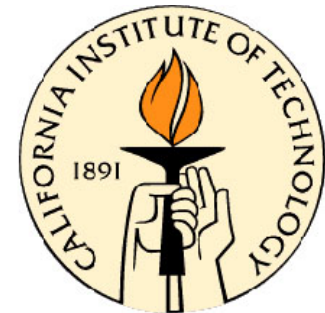


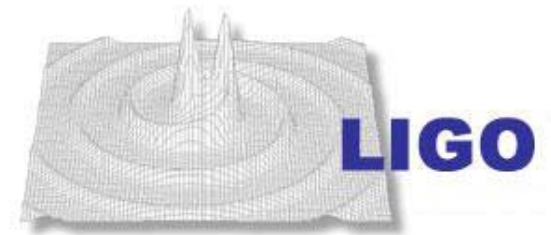
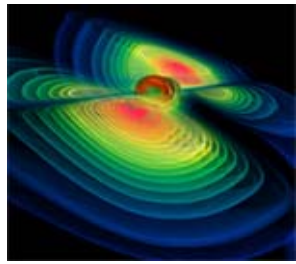
# How a Theoretical Physicist.. ..can have fun!



**Dott. Juri Agresti**

*Labor Day party  
01-Sep-03*



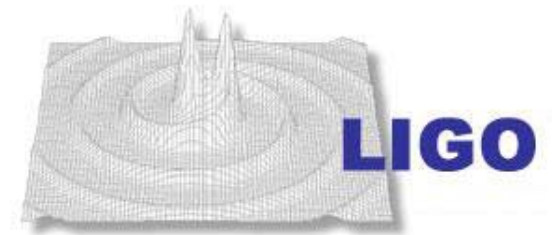
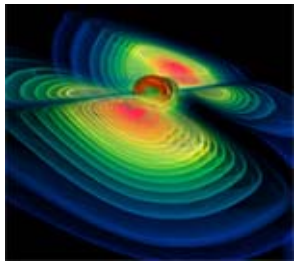


## Who am I ?

Graduated from Florence  
University with a Thesis that  
sounds like:



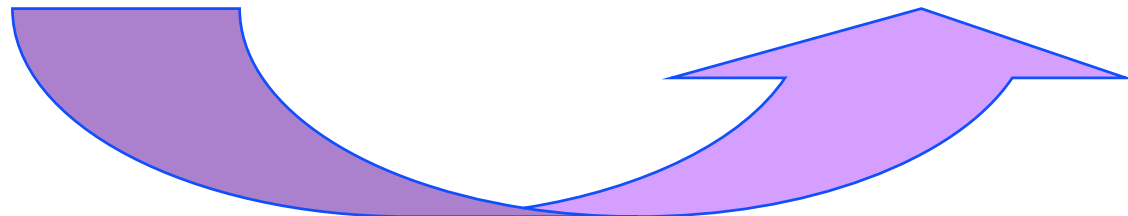
“Hamiltonian linearization of the rest-frame instant form of tetrad gravity in a completely fixed 3-orthogonal gauge: a radiation gauge for background-independent gravitational waves in a post-Minkowskian Einstein space-time.”

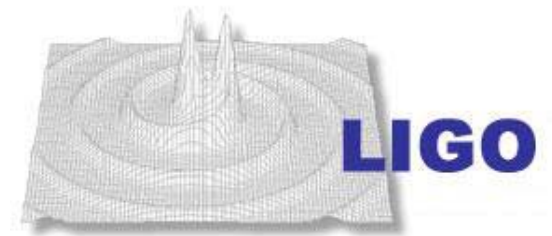
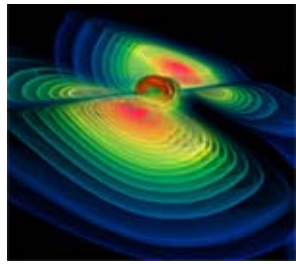


## And Now ?

Ph.D. student of Pisa University with a LIGO grant.

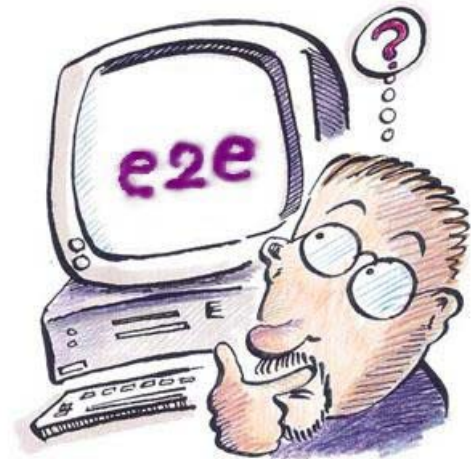
Caltech's visitor during this summer with my wife Chiara (SURF student)

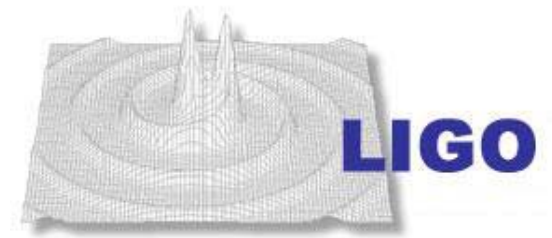
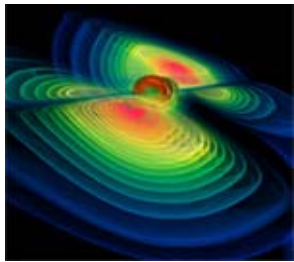




## How I spent this time?

Essentially learning a lot of stuff about gravitational waves interferometers and the *end-to-end (e2e)* simulation package developed to model LIGO detectors.





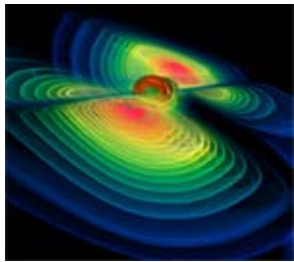
## What is e2e ?

### A time domain simulation program

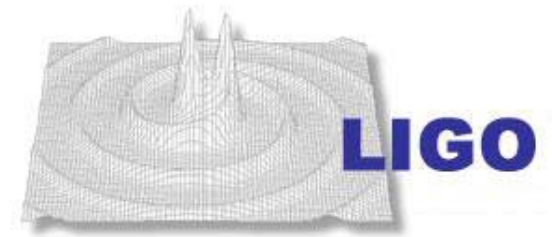
(Time-domain simulation is necessary to understand non-linear behavior)



- **Fields**
- **Optics**
- **Mechanics**
- **Electronics**
- **Control syst.**



# How it works :



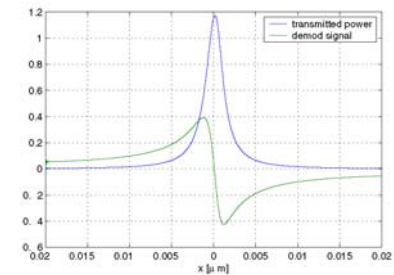
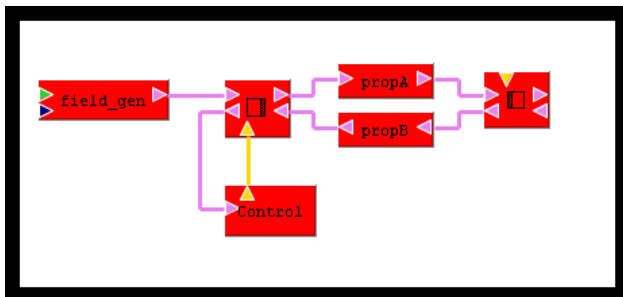
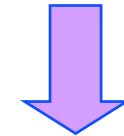
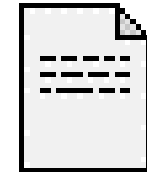
Setup a configuration  
using Graphical Interface  
or Text Editor

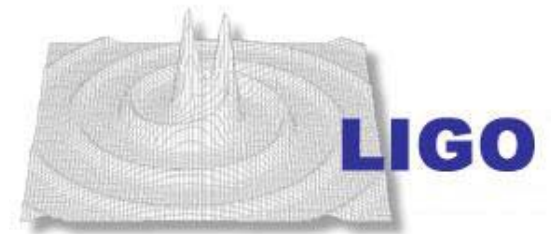
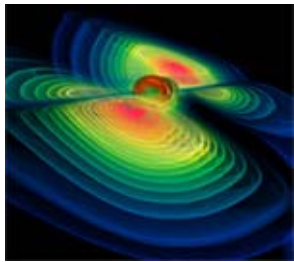


Time domain  
simulation



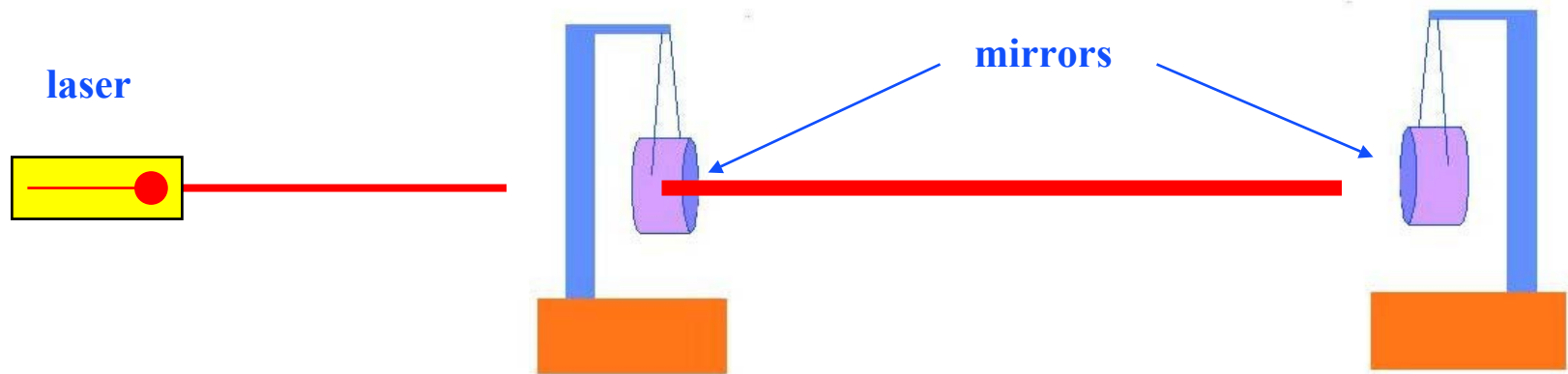
Simulated  
time series of data



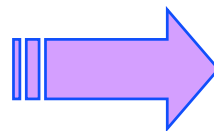


# My personal training with e2e .....

## Fabry-Perot cavity with 3D suspended masses

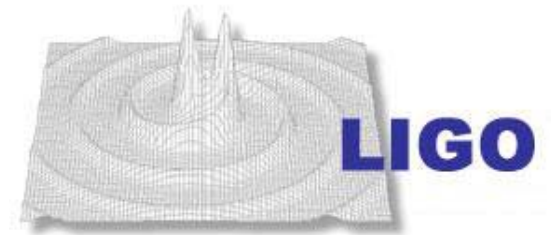
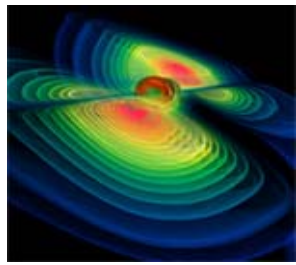


Unexpected results !



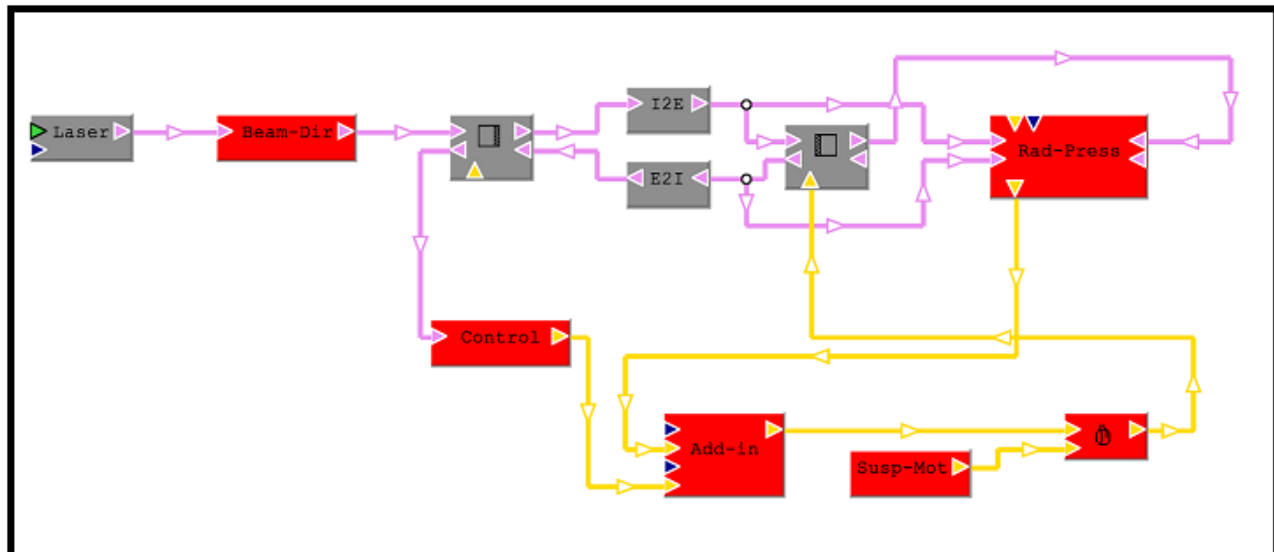
Bugs fixed ...

Training is useful !

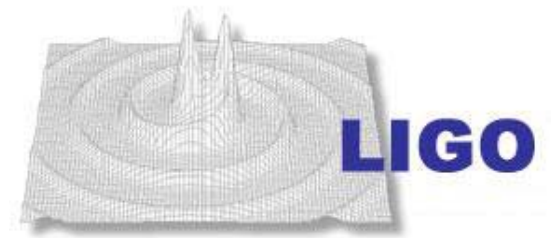
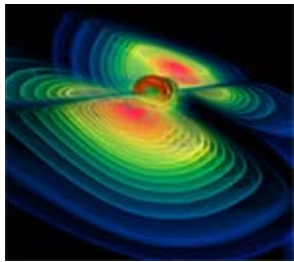


After these preliminary studies I have started to build a *FP cavity simulation with radiation pressure effect*.

This is to validate and extend the analysis by Danniell Sigg about **angular instabilities** induced in a FP cavity by radiation pressure.



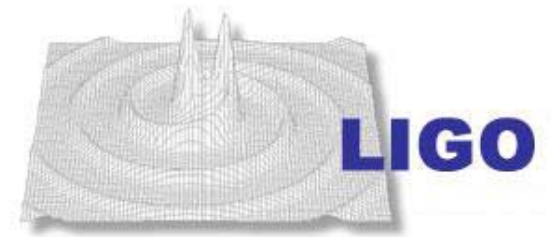
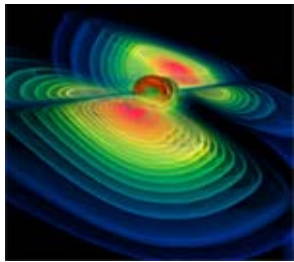




## The analysis is just begun .....

- Freeze all the degrees of freedom of the mirror but one (yaw, pitch, z-motion,.....)
- Add a control system
- Add noise source (seismic motion....)

**Cavity locking.....Under which conditions ?**

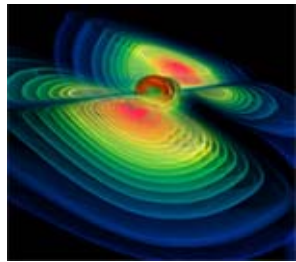


**Power-Induced angular shift can be significant for  
LIGO I** ( laser power in the arms cavities  $\sim 10$  kW)

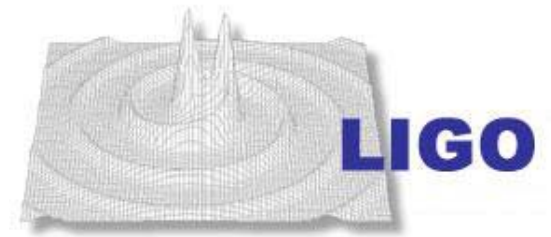
**...but could be very important for LIGO II**  
( power  $\sim$  MW )



**A lot of work to do !**

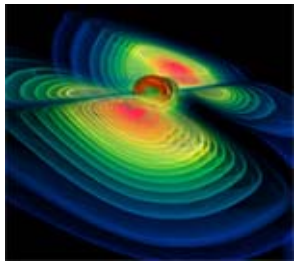


**Thanks to :**

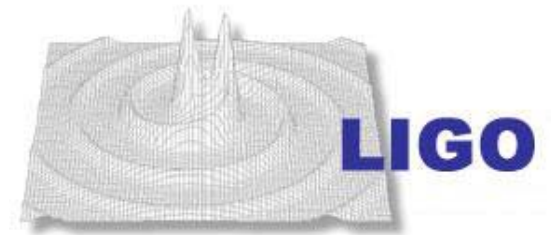


- **Hiro Yamamoto** ( he is very kind and patient )
- **Riccardo DeSalvo and family** (they accommodated us for a month)





# Thanks to (2) :



- **Chiara** ( for being as you are)
- **All new friends met in Caltech**
- **Albert** .... ( he gave me something to have fun..)

