



Advanced LIGO and further

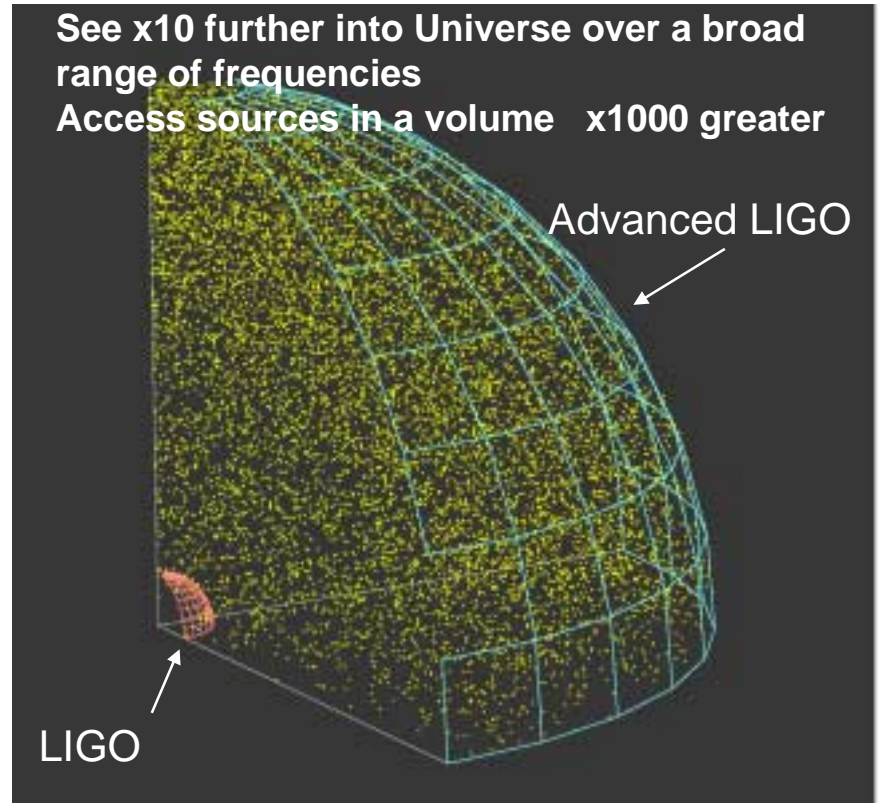
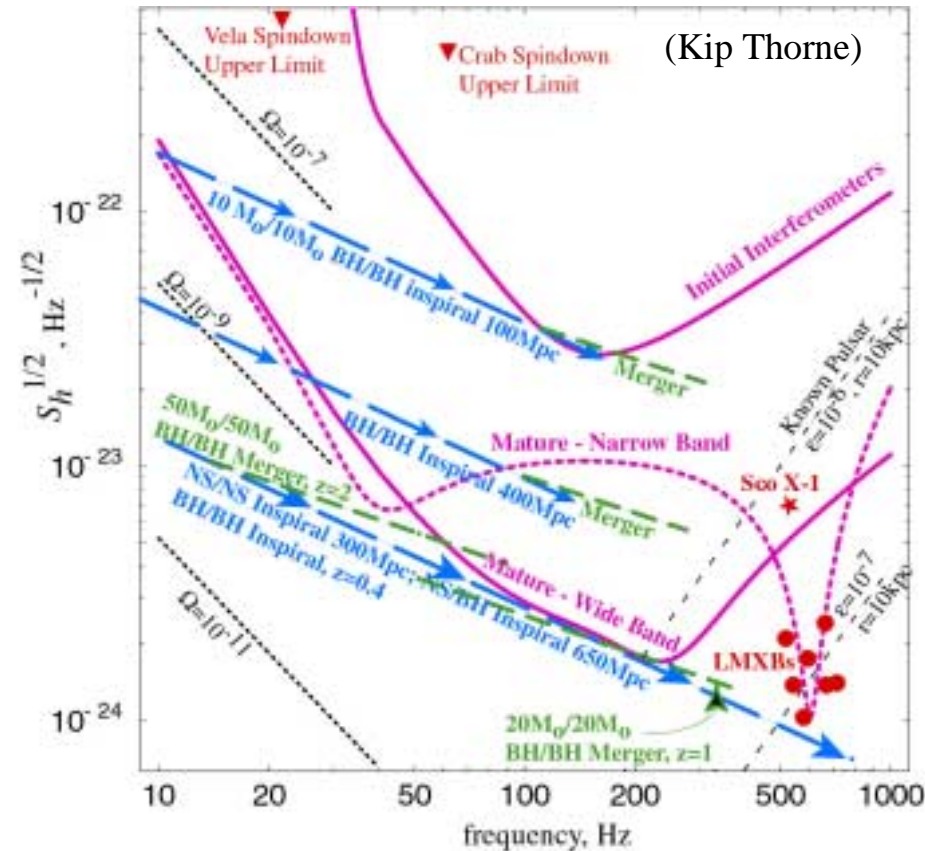
Sheila Rowan, Stanford University/Univ. of Glasgow
on behalf of the LIGO Scientific Collaboration

APS Annual Meeting

6th April 2003

LIGO-G030183-00-Z

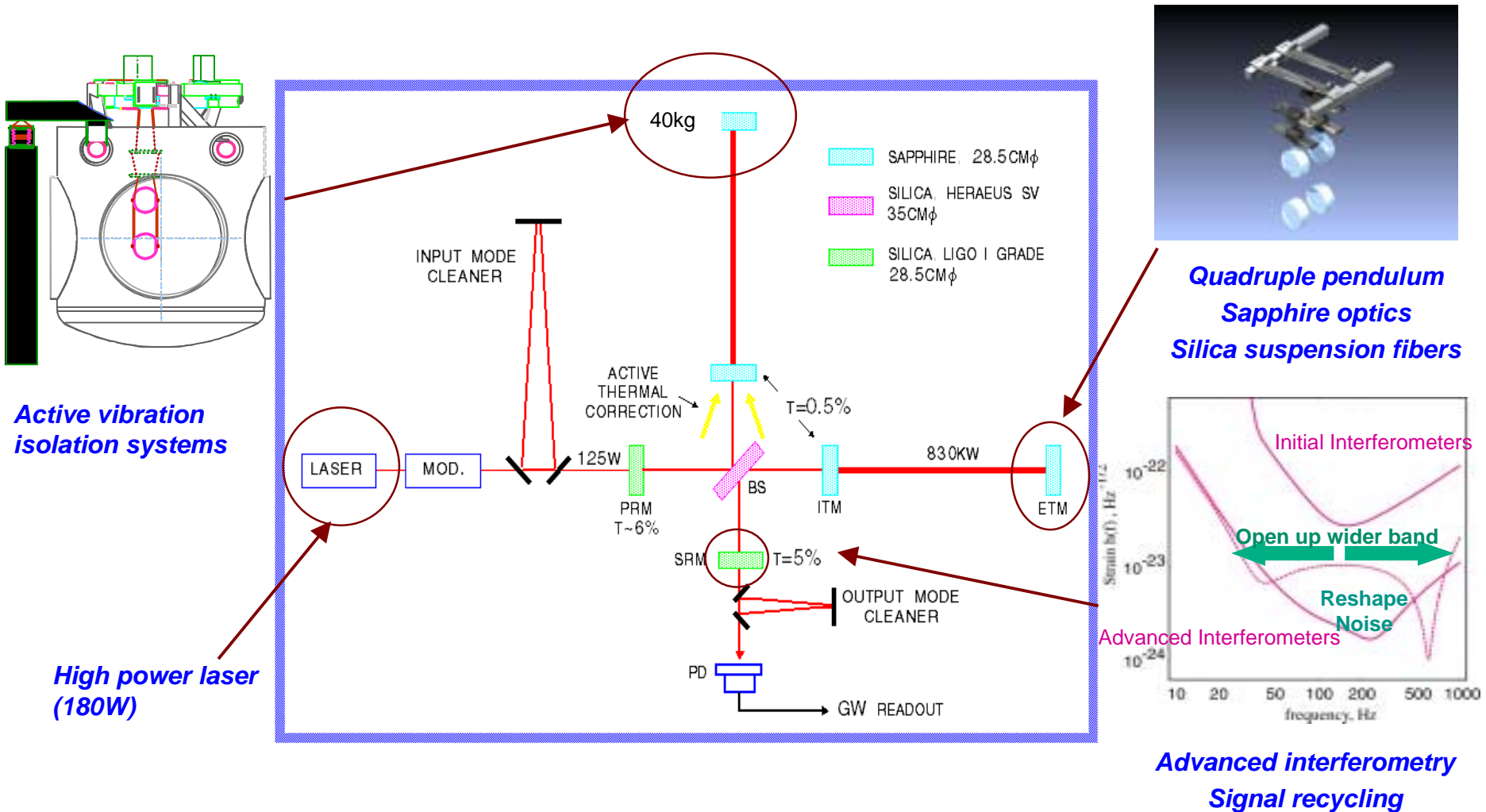
From LIGO to Advanced LIGO



- Initial LIGO observation 2002 – 2006
 - » One complete year of data by 2007
 - » A few hours of observing with Advanced LIGO equivalent to 1 year observation with initial LIGO
- Advanced LIGO: Move from gravitational wave detection to GW astronomy
- Start coincident observations with Advanced LIGO detectors in 2009

What is 'Advanced LIGO'?

Major technological differences between LIGO and Advanced LIGO



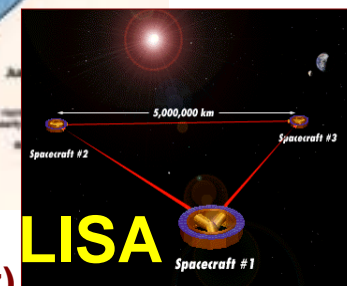
LIGO Advanced LIGO in the worldwide network



AIGO

- Detection confidence
- Source polarization
- Sky location

(Sources $\ll 10\text{Hz}$)

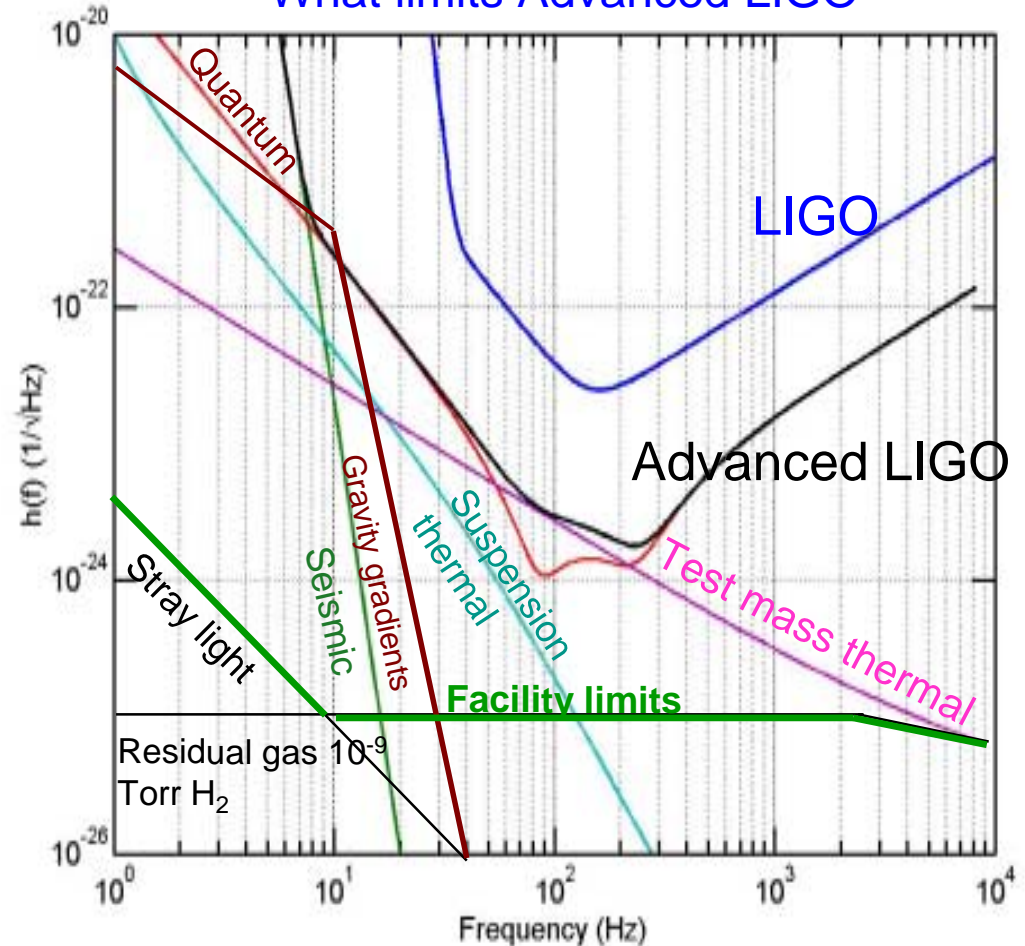


- Techniques to bypass Heisenberg
 - » squeezed light
 - » 'speed-meters'
- Subtracting effects of gravity gradients
- Cryogenics
 - » cooling of test masses and suspensions
- Diffractive optics
 - » alternative topologies
- New materials
 - » silicon



6" silicon mirror

What limits Advanced LIGO



- Research on these technologies starting now so that they are mature on the timescales needed (2010+)