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Sources workshop

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4th roundtable:

inspiral of stellar mass

objects into supermassive  
black holes

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# TEMPLATE COUNTING & DATA ANALYSIS WITH KLUDGED WAVEFORMS

Plan: Use incorrect (but qualitatively plausible) waveforms to determine the costs and performances for various analysis methods

- E.g. Scott Hughes' fake R.R. waves  
Curt Cutler's fake P.N. waves
- Method and performance should not change significantly when Correct™ waveforms become available

Key issue: Waveforms are complicated!

- Ideally, match waveform to within 1 cycle over entire length

→ Each parameter  $\lambda$  requires a number of samples

$$\sim \max_{\lambda, t} \{ \Delta N_{\text{cycles}} \}$$

- Over a dozen search parameters!  
( $\geq 9$  significantly affect wave phase)

Ignorant estimate of # flops / datum:

$$N_{\text{flops}} \sim (N_{\text{cycles}})^9 \log(N_{\text{cycles}}) \times \left(\frac{V_{\oplus}}{c}\right)^2$$

E.g.  $N_{\text{cycles}} \sim 10^5 \rightarrow N_{\text{flops}} \sim 10^{38}$

## Data analysis tricks:

### Semi-coherent analysis:

- Divide data into  $M$  shorter stretches; combine these incoherently
- Effectively reduce  $N_{cycles}$  by factor  $M$ , raise amplitude threshold by factor  $\sqrt[4]{M}$ .
- To get  $N_{flops}$  down to around  $10^{12}$ , need  $N_{cycles} \sim 100 \quad \Rightarrow \sqrt[4]{M} \lesssim 10$

### Correlated parameters:

- When two parameters affect the waveform similarly, we can reduce (or even eliminate) search over one of them.
- E.g. early work with geodesics shows that radius and eccentricity have 2 to 10-fold correlations.

Position &  
orientation  
System constants

Initial conditions

## PARAMETER

## # SEARCH POINTS

Right ascension	$\sim \left(\frac{v_0}{c}\right) N_{\text{cycles}}$
Declination	$\sim \left(\frac{v_0}{c}\right) N_{\text{cycles}}$
Polarization angle	$\sim 1$
Inclination angle	$\sim 1$
Axial rotation angle	$\sim 1$
Central body mass	$\sim N_{\text{cycles}}$
Central body spin	$\sim N_{\text{cycles}}$
Orbiting body mass	$\sim N_{\text{cycles}}$
Initial orbit eccentricity	$\sim N_{\text{cycles}}$
Initial orbit inclination	$\sim N_{\text{cycles}}$
Initial apsidal anomaly	$\sim N_{\text{cycles}}$
Initial nodal angle	$\sim N_{\text{cycles}}$
Start time	$\sim \log(N_{\text{cycles}})$