

#### **External Trigger (S)News**

#### LSC 2001 August Meeting LIGO Hanford Observatory

#### Szabolcs Márka

and the the External Triggers Group

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LIGO/CalTech



# (Practical) SNEWS



#### **NEWS**!

•(I believe) Successful high rate test during the summer

•Soon fully automatic

•SNO is onboard!

### International collaboration of SN sensitive neutrino detectors

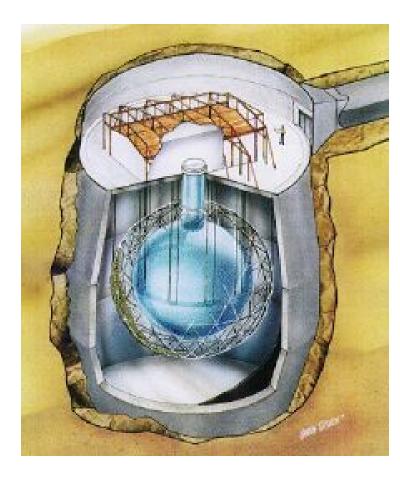
- » Super-K
- » LVD
- » Amanda
- » SNO
- » LIGO (observer)
- Provides near-real time coincidence based SN alarm
  - » Timing and pointing information
  - » Very high confidence
    - Less than 1 false alarm/100y !
- Coordinates detector downtime
- Centralized timing verification
- Privacy is ensured
  - » Input data is strictly secured

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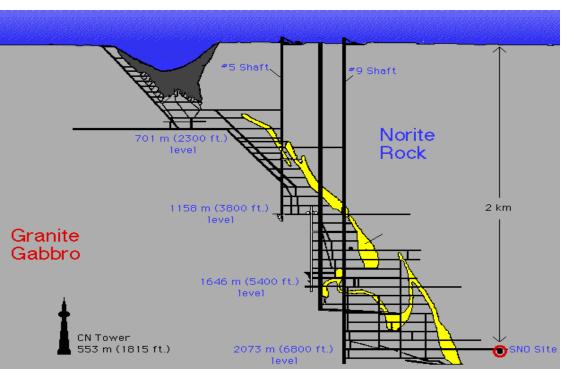


#### **SNO (Sudbury Neutrino Observatory)**



- Creighton mine; Sudbury, ON, Canada
- 1000 Tons of heavy water
- Surrounded by several kTons of water
  - ~9700 PMTs

~2Km below surface



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### **SNO** supernova sensitivity

Number of Particles From 10 kpc Supernova:

Neutrino Reaction	Туре	SNO Cou [ε = 100%		•
$\overline{\nu}_{e} + p_{H_{2}0} \rightarrow n + e^{+}$	СС	356	331	
$\overline{v}_{e} + p_{p_{2}0} \rightarrow \mathbf{n} + \mathbf{e}^{+}$	CC	0.2		•
$\overline{\nu}_{e} + p_{AV} \rightarrow \mathbf{n} + \mathbf{e}^{+}$	CC	5		
$\nu_e + d \rightarrow p + p + e^-$	CC	83	72	
$\overline{\nu}_{e} + d \rightarrow n + n + e^{+}$	CC	53 (x 3) 8	32 [D <sub>2</sub> 0] 138 [sait] 90 [NCD]*	
$\nu_e$ + <sup>16</sup> O $\rightarrow$ <sup>16</sup> F + $e^-$	CC	1		•
$\overline{\nu}_{e}$ + <sup>16</sup> O $\rightarrow$ <sup>16</sup> N + e <sup>+</sup>	CC	3		
$\nu_e + d \rightarrow \nu_e + p + n$	NC	36	12 [D <sub>2</sub> O] 30 (salt) 20 [NCD]-	•
$\overline{\nu}_{e} + d \rightarrow \overline{\nu}_{e} + p + n$	NC	36	12 [D <sub>2</sub> 0] 32 (sait) 21 [NCD]*	
" $\nu_{\mu}$ " + d $\rightarrow$ " $\nu_{\mu}$ " + p + n	NC	192	60 [D <sub>2</sub> O] 164 [sait] 110 [NCD]*	
" $\nu_{\mu}$ " + <sup>16</sup> O $\rightarrow$ ( <b>n</b> , $\gamma$ , <b>n</b> +	γ) NC	7		
$v_e + e^- \rightarrow v_e + e^-$	ËS	26	20	
$\overline{\nu}_{e} + e^{-} \rightarrow \overline{\nu}_{e} + e^{-}$	ES	9	8	
" $v_{\mu}$ " + $e^{-} \rightarrow "v_{\mu}$ " + $e^{-}$	ES	12	9	
TOTAL SNO SN COU	NTS:	917	606 [D20] 804 [sait] 681 [NCD]	

Still "only galactic" sensitivity...

but we can count on exceptional results from SNO

- » ~1000 neutrinos are expected (10Kpc)!
- » Excellent sensitivity
- "Rock solid" SNEWS coincidence trigger
  - » Super-K provides comparable number of events
  - » LVD and Amanda will have considerably less
- Detailed information on SN evolution is expected!
- It is probably conceivable to build an accurate numerical model for this particular SN based on this info...



From the talk by R. Trafirout at MarinaDel Rey, 2001



# **GRB** (notices from GCN)

#### We receive a fair number of triggers

- » Have a chance for half a dozen or more during E6
- » Mostly HETE related
  - Must coordinate E6 with HETE runs
- We receive
  - » Arrival time
  - » Direction

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- » and more...
- Utilize Finn/Mohanty/Romano method
  - » However, we must ...
    - Tune it to low number of events
    - Find best use of info provided
    - Develop a full practical implementation
    - before E6...
- Survey predicted GW waveforms associated with GRBs

TITLE: GCN/HETE BURST POSITION NOTICE				
NOTICE_DATE: Mon 02 Jul 01 03:39:30 UT				
NOTICE_TYPE: HETE S/C_Last				
TRIGGER_NUM: 1576, Seq_Num: 2				
GRB_DATE: 12092 TJD; 183 DOY; 01/07/02				
GRB TIME: 12846.52 SOD {03:34:06.52} UT				
TRIGGER SOURCE: FREGATE Trigger,				
GAMMA RATE: 850 [cnts/s] on a 0.020 [sec] timescale				
XG_TRIG_TIME: 3.600 [sec]				
WXM SIG/NOISE: 28 trigger sig/noise				
SCZ_RA: 280 [deg]				
SC - Z DEC: -22 [deg]				
WXM_CNTR_RA: 281.982d {+18h 47m 56s} (J2000),				
282.003d {+18h 48m 01s} (current),				
281.280d {+18h 45m 07s} (1950)				
WXM_CNTR_DEC: -12.779d {-12d 46' 42"} (J2000),				
-12.777d {-12d 46' 36"} (current),				
-12.835d {-12d 50' 04"} (1950)				
WXM_MAX_SIZE: 38.17 [arcmin] diameter				
WXM_LOC_SN: 2 image sig/noise				
SUN_POSTN: 101.24d {+06h 44m 58s} +23.04d {+23d 02' 10"}				
SUN_DIST: 169.72 [deg]				
MOON_POSTN: 238.64d {+15h 54m 33s} -17.12d {-17d 07' 27"}				
MOON_DIST: 42.04 [deg]				
MOON_ILLUM: 88 [%]				
GAL_COORDS: 21.13,-5.02 [deg] lon,lat of the burst				
COMMENTS: Possible GRB.				
COMMENTS: Possible XRB.				
COMMENTS: WXM error box is circular; not rectangular.				
COMMENTS: There is no WXM or SXC position in this notice.				



#### **Summary (of interesting questions)**

- All right we can get excellent triggers...
  - » Neutrino, Gamma Ray Burst, Optical,...
- How can we use the information the best?
  - » Can we construct an accurate enough simulation of SN gravity wave signature based on the result of neutrino searches?
  - » Is it advantageous to use more than the direction and arrival time?
  - » Can we use a single external event or we need plenty?
    - Can we use half a dozen event?
      - + Is the Finn/Mohanty/Romano method applicable?
  - » Should we construct filters for GWs associated with specific GRB emitters?