
Perspectives on detector networks, and noise

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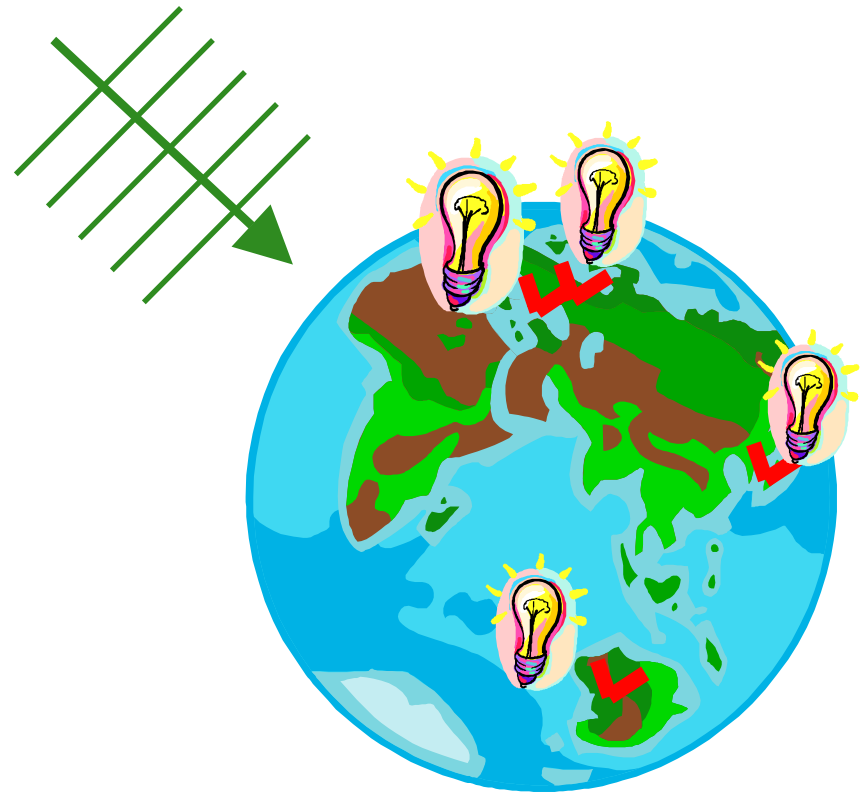
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Overview

- Coincidence, Correlation and the Effective Detector
 - Is the whole less or greater than the sum of the parts?
- In the presence of non-Gaussian noise
 - The effectiveness of the Normal distribution, and going beyond
- Cf. Finn, “Aperture synthesis for grav. wave data analysis...”, PRD **63**, 102001 (2001)

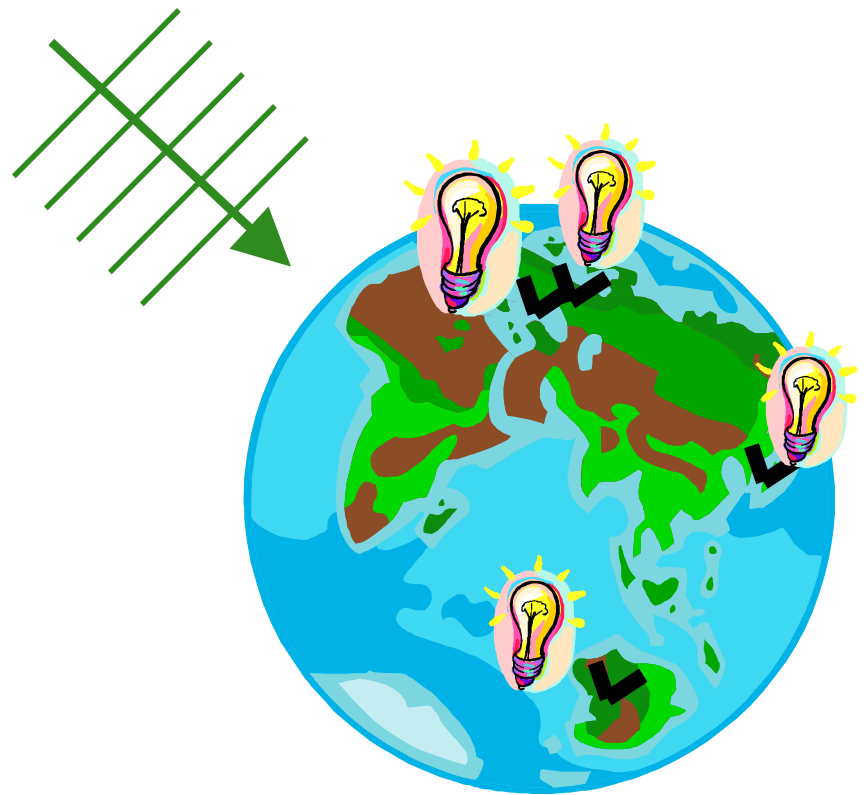
Coincidence ...

- Incident plane wave interacts with an array of detectors
- Coincidence
 - Detection if each detector registers wave's passage
- Sensitivity limited by weakest link ...
 - No detection, no coincidence
- ... or gain nothing by including weaker detectors
 - Too high false rate leaves statistics unimproved



Coincidence ...

- Incident plane wave interacts with an array of detectors
- Coincidence
- Sensitivity limited by weakest link ...
- ... or gain nothing by including weaker detectors
- But what if we shouldn't have seen anything in a given detector?
 - Polarization, amplitude, signal bandwidth, *etc.*



... and Correlation

- Incident plane wave interacts with array of detectors
- Detector array responds *coherently*
 - I.e., relative timing, polarizations, amplitudes
- Treat network response via likelihood techniques
 - I.e., optimal filter when signal known
- Sensitivity?



Correlation

- Recall optimal filter for single detector:

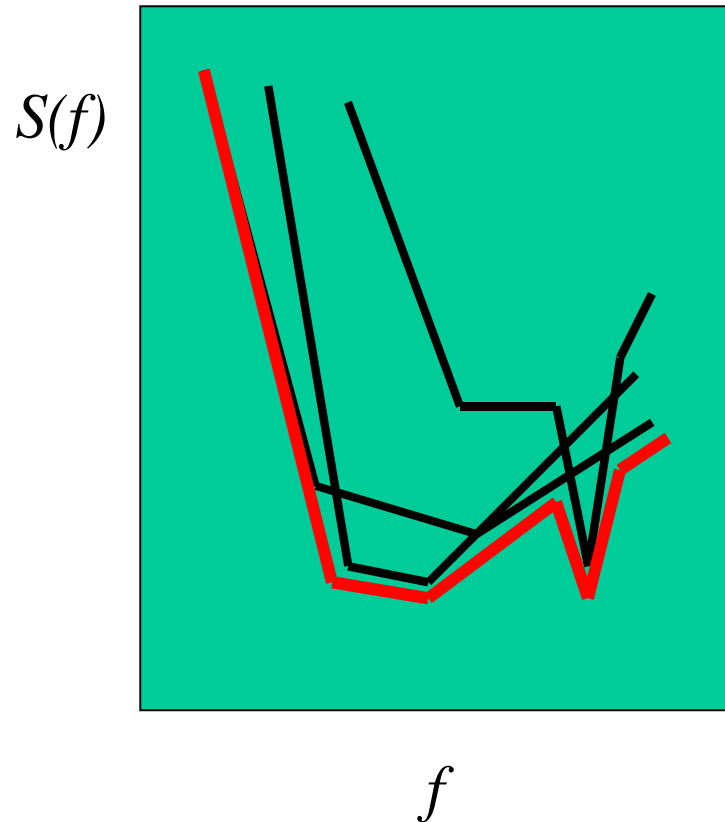
$$\langle h, m \rangle_k = \int df \frac{\tilde{h}(f) \tilde{m}^*(f)}{S_k(f)}$$

- For multiple detectors and no inter-detector noise correlations:

$$\langle \vec{h}, \vec{m} \rangle = \sum_k \langle h_k, m_k \rangle_k$$

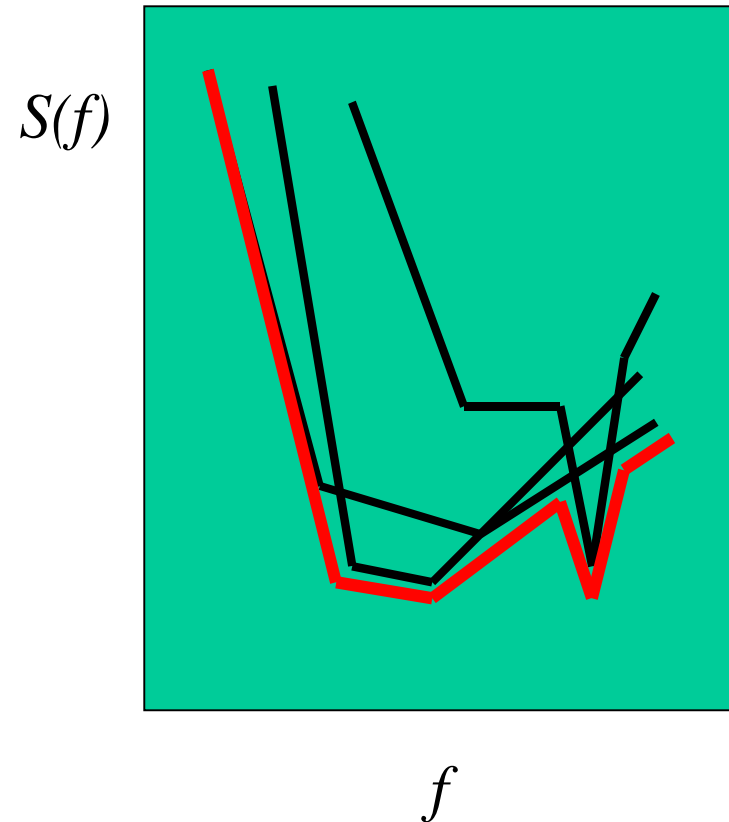
- The *effective detector* noise spectrum:

$$S(f) \approx \left[\sum_k S_k^{-1}(f) \right]^{-1}$$



Correlation

- Additional detectors *always* reduce effective detector noise
 - No more “weakest link” *vis a vis* sensitivity
- Effective detector best envelope of component detectors
 - Get Virgo’s low frequency response, LIGO’s mid-frequency, GEO’s RSE high-frequency, etc.
- Polarization, amplitude, signal bandwidth all accommodated
 - Effective detector noise direction, orientation dependent



What if noise non-Gaussian?

- What *if* noise non-Gaussian!
 - As long as stationary...
- Suppose know only mean, variance (i.e., psd $S(f)$)
 - What probability distribution $P(x)$ makes least assumptions about noise character? The *maximum entropy* distribution:

$$0 = \delta \int dx \left[P(x) \log P(x) + \lambda_0 P(x) + \lambda_1 x P(x) + \lambda_2 x^2 P(x) \right]$$

$$P(x) \propto \exp \left[-\frac{(x - \mu)^2}{2\sigma^2} \right]$$

- Extensible to higher-order moments
 - As long as highest order even

Summary

- Coincidence
 - Effective only when working with detectors of comparable sensitivity in comparable bands with comparable polarization ...
- Correlation
 - The whole is greater than the sum of the parts
 - Failure to “see” in a given detector is positive information!
- In the presence of non-Gaussian noise
 - The Normal distribution, and “optimal” filtering, is the most robust choice if only noise mean, PSD are known
 - Can do better (with cost!) if additional information available. Critical information is higher-order moments of distribution