

E2 Calibration of the Hanford 2km IFO

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calibration procedure

- idea: apply known force to a mass and measure displacement
- methods: swept sine, fixed frequency calibration lines, photon calibrator
- how good? Assume 1-10 events with SNR 10-100, the physical results should not be limited by the systematic uncertainties of the calibration
- +/- 1% amplitude, +/-10 microseconds

Sigg, D. LIGO-T970101-A-D Allen, B. LIGO-T960189-00-E



hubble constant

- Determine the Hubble constant H_o
- require 3-4 interferometers
- amplitude accuracy required: 3%, timing accuracy required:1% of max timing range (-22 to +22 milliseconds)
- three detectors: 3 amplitudes and 2 time delays yields arrival directions (2), polarization amplitudes (2), and phase lag (1).
- four detectors: error box of 6°x6° degrees: search box for galaxy to get source redshift
- ten events out to 100 Mpc determines H_o to a few percent

Shutz, B.F., Nature 323 (1986) 310



itm calibration

Drive the ITM with a slow sinusoid (0.1 Hz) and count the number of fringes that are read out at the antisymmetric port.

Calibrations for the ITM's are then ITMx = 3.6 ± 0.2 nm/count ITMy = 3.5 ± 0.2 nm/count





etm calibration

To extrapolate the calibration of the input test masses (ITM's) to the end test masses:

Assuming identical coil drivers (E1=E2) and pendulum transfer functions (P1=P2), and using equal excitations on both masses,

$$\alpha_{\rm etm} = \frac{E_{\rm etm}}{E_{\rm itm}} \alpha_{\rm itm}$$

Calibrations for the ETM's are then

ETMx = 2.3 nm/count

ETMy = 2.0 nm/count



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measurement stability



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itm vs etm



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LIGO fit to etm_y transfer function



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complicated



sensitivity



Scale set by absolute calibration
Shape set by parametrization
Visible calibration lines ("c")
30% calibration accuracy