

# The Thermal Noise Interferometer

MEASURING DISPLACEMENT NOISE IN ADVANCED SUSPENDED INTERFEROMETERS

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LIGO/CALTECH

## OBJECTIVES:

### ↓ CHARACTERIZE ADVANCED DETECTORS

VERIFY DESIGN SPECIFICATIONS

MEASURE NOISE SOURCES

MEASURE NON-THERMAL NOISE PROPERTIES

### ↓ PHYSICS OF FUNDAMENTAL NOISE SOURCES

NOISE PHYSICS, STATISTICS

REACH (AND EXCEED) THE SQL



# TNI Design Elements

The many advantages of using a short cavity length ( $L \lambda \ll 1\text{cm}$ )

↓ **SHORT CAVITY STORAGE TIME**

- ∪ **USE HIGH FINESSE CAVITIES**
- ∪ **NO POWER RECYCLING**

↓ **INDEPENDENT CAVITIES**

- ∪ **NO RECOMBINATION**
- ∪ **INDEPENDENT CONTROLS**

↓ **REDUCED OPTICAL POINTING REQUIREMENTS - NO WFS**

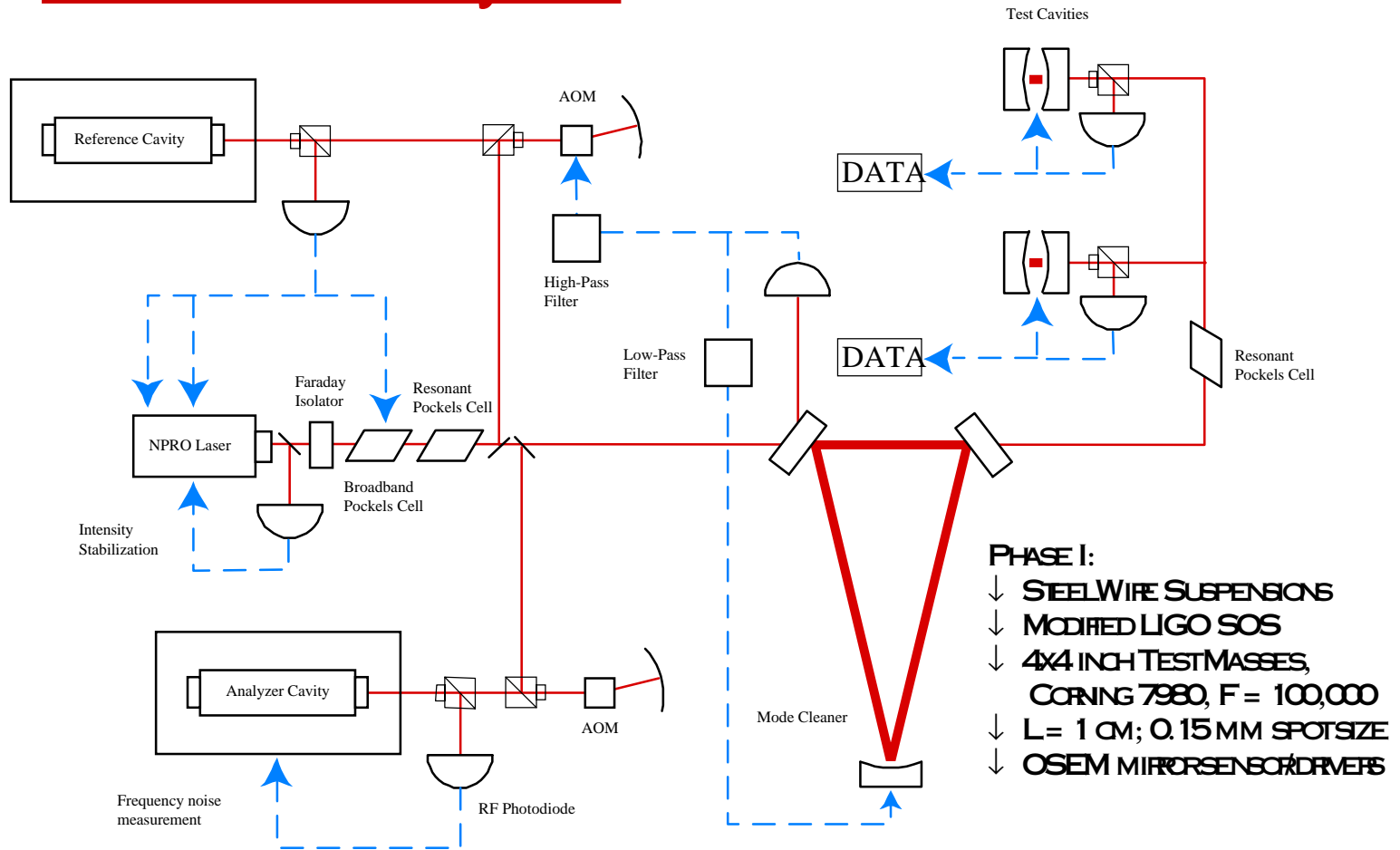
↓ **REDUCED LASER STABILITY REQUIREMENTS**

↓ **COMMON SUPPORT FOR TEST MASSES**

- ∪ **REDUCED SEISMIC NOISE**
- ∪ **LOWER SUSPENSION RECOIL THERMAL NOISE**

**BUT.. SMALLER LASER SPOT SIZE ∪ HIGHER INTERNAL THERMAL NOISE**

# TNI Initial Layout



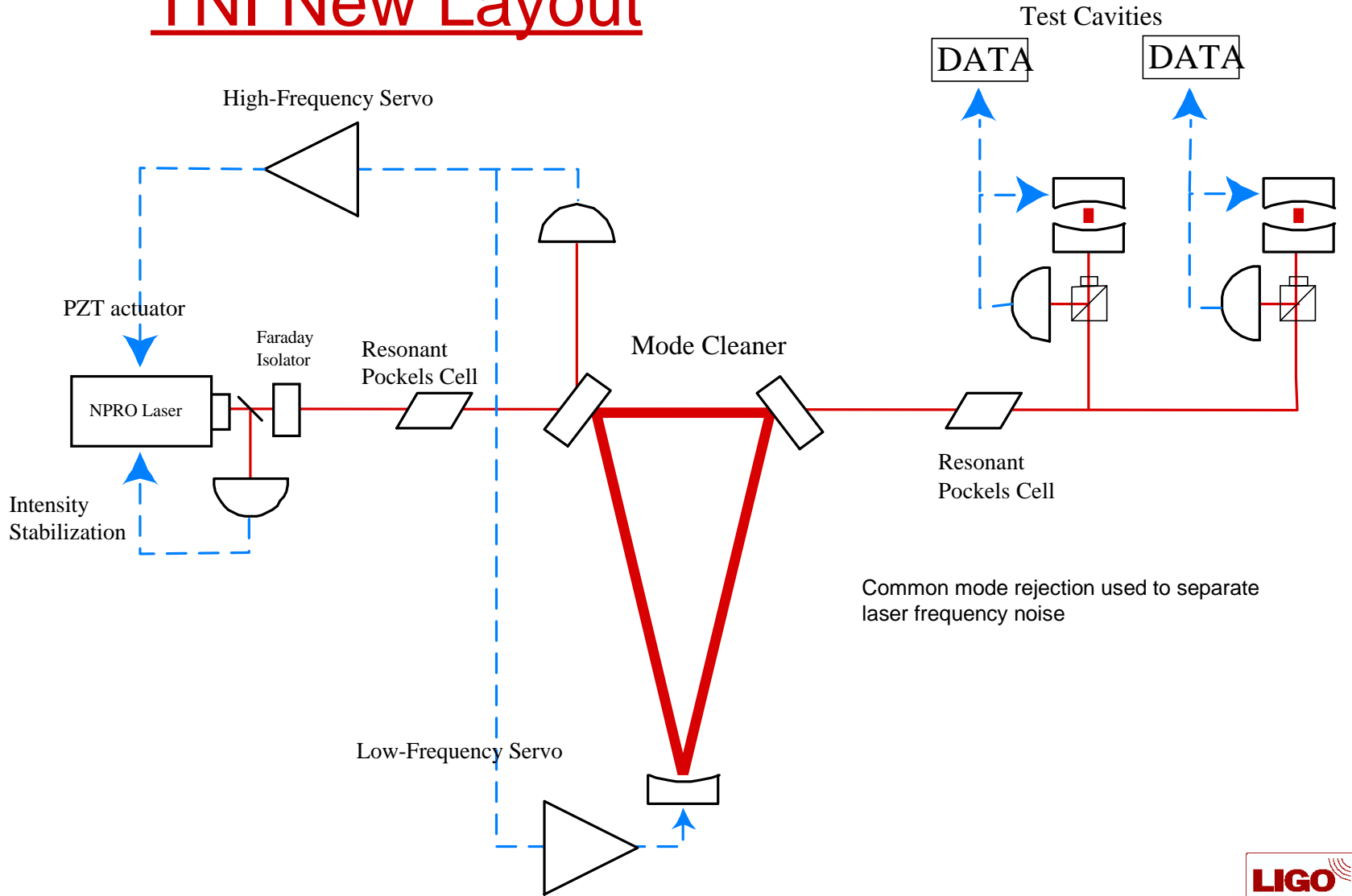
## PHASE I:

- ↓ STEELWIRE SUSPENSIONS
- ↓ MODIFIED LIGO SOS
- ↓ 4x4 INCH TEST MASSES,  
CORNING 7980, F = 100,000
- ↓ L = 1 CM; 0.15 MM SPOTSIZE
- ↓ OSEM MIRROR SENSOR DRIVERS

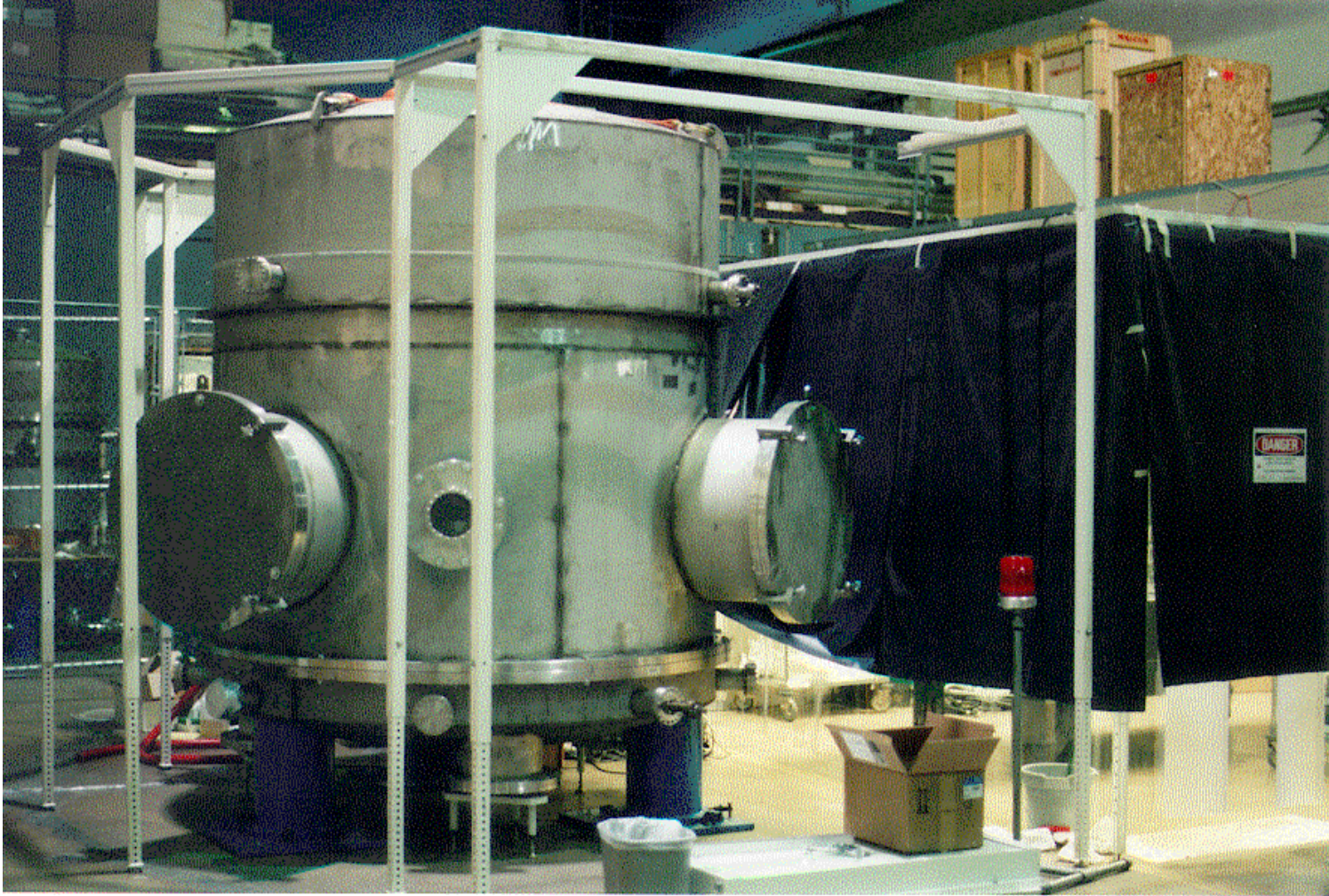
← LIGO PSL copy →



# TNI New Layout

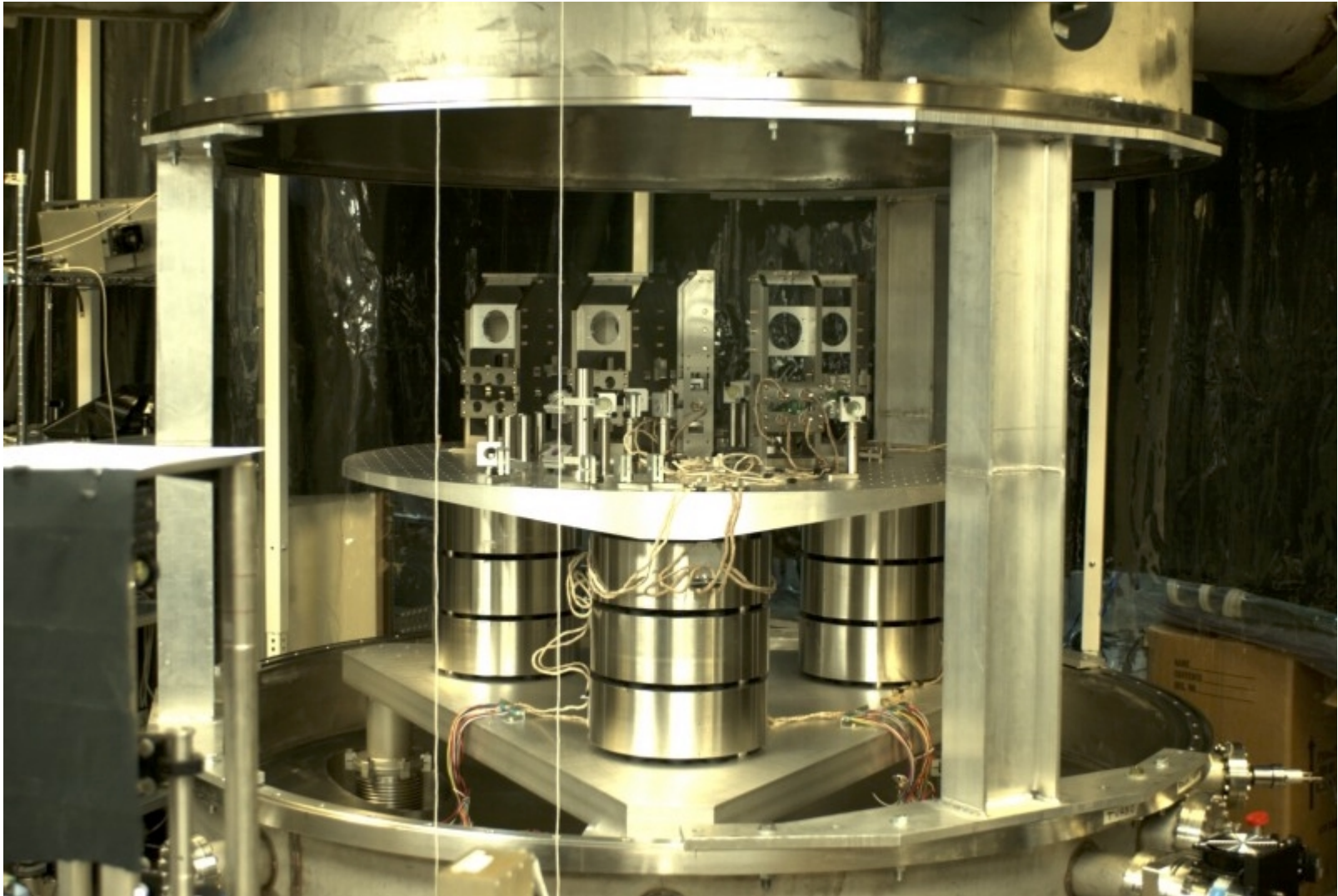


## TNI Outside View

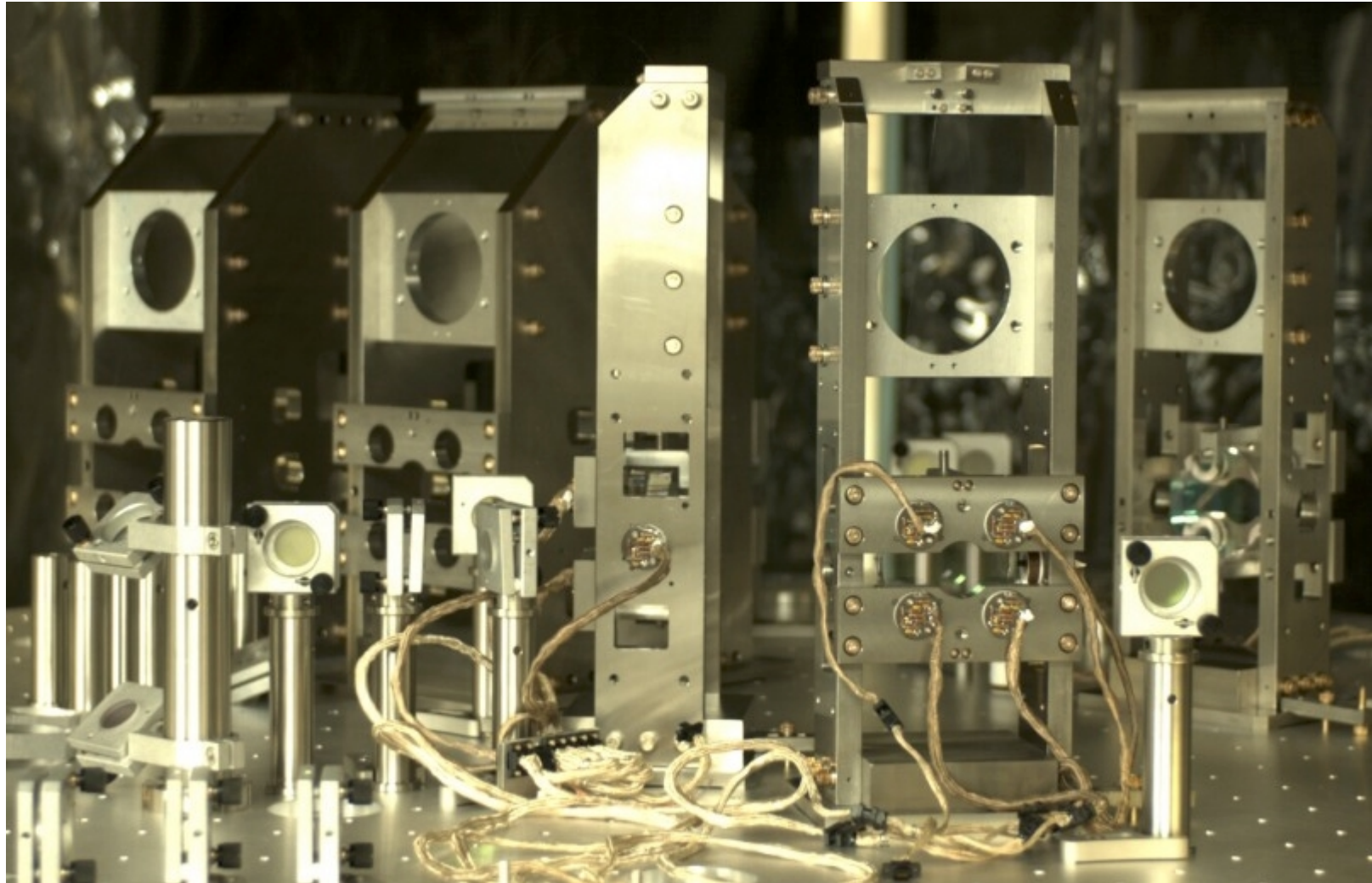


...before clean-room cover around vacuum chamber

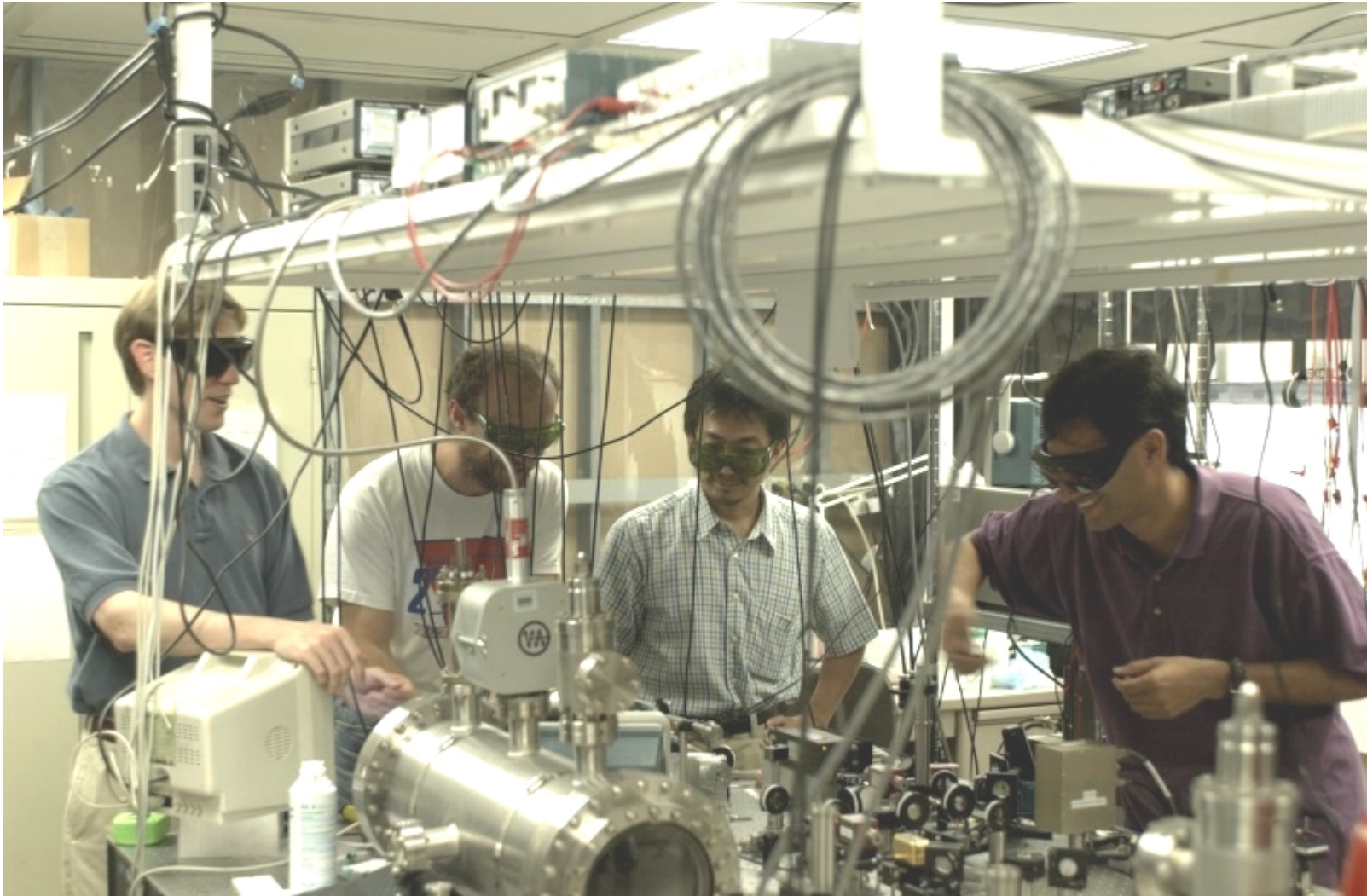
TNI View inside vacuum chamber



TNI View inside vacuum chamber



The TNI crew



Eric Black

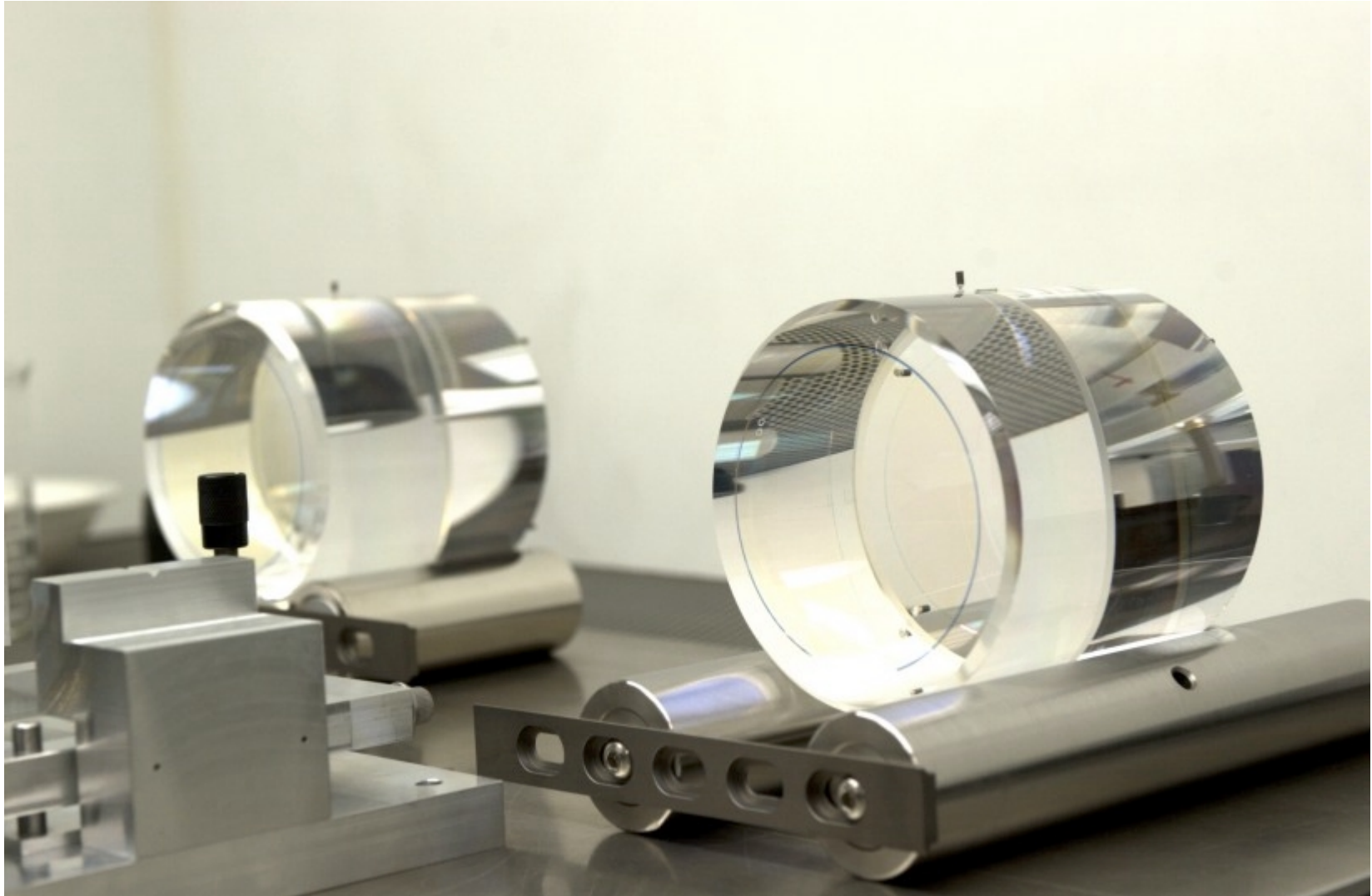
Luca Matone

Seiji Kawamura

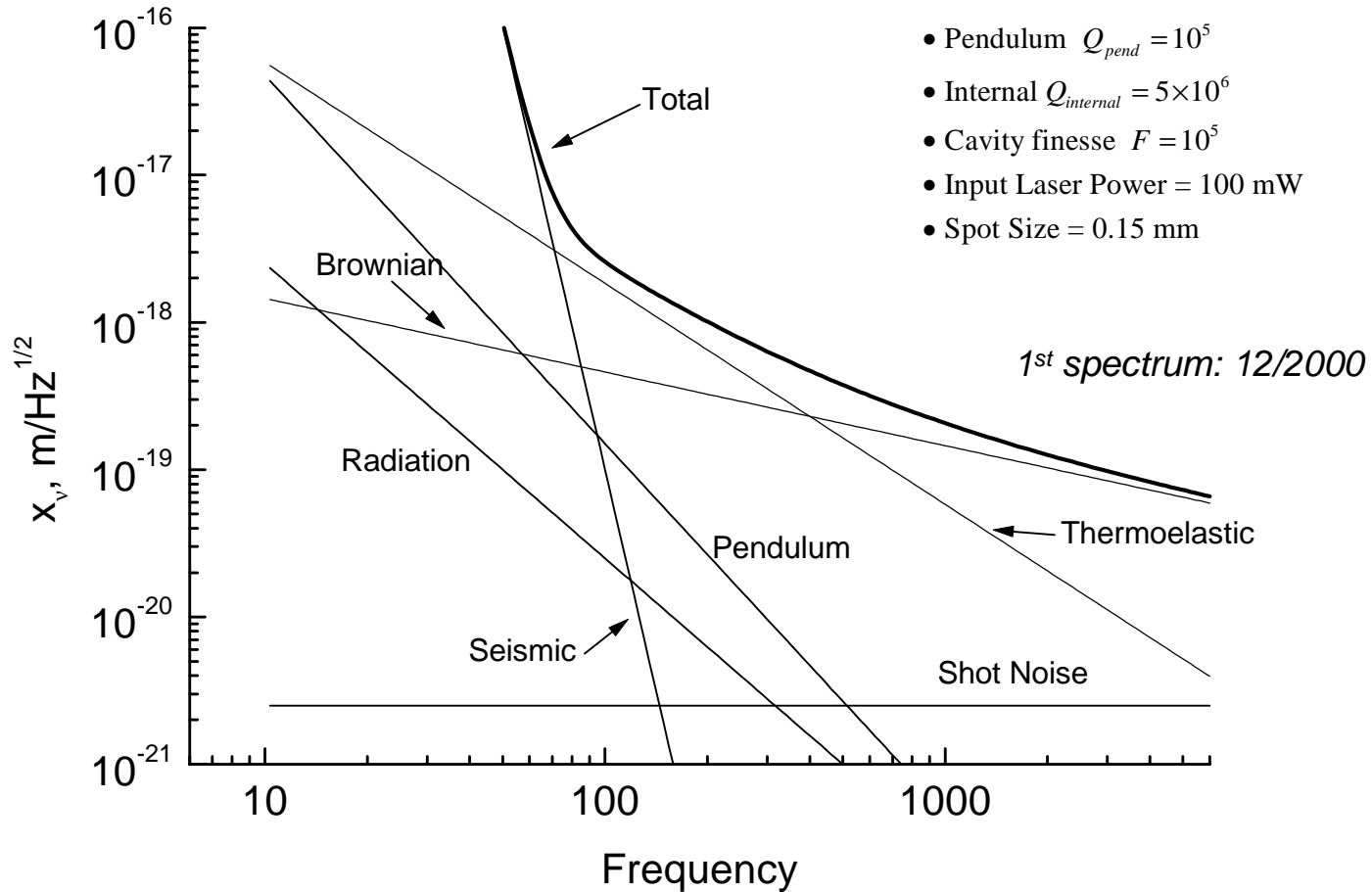
Shanti Rao



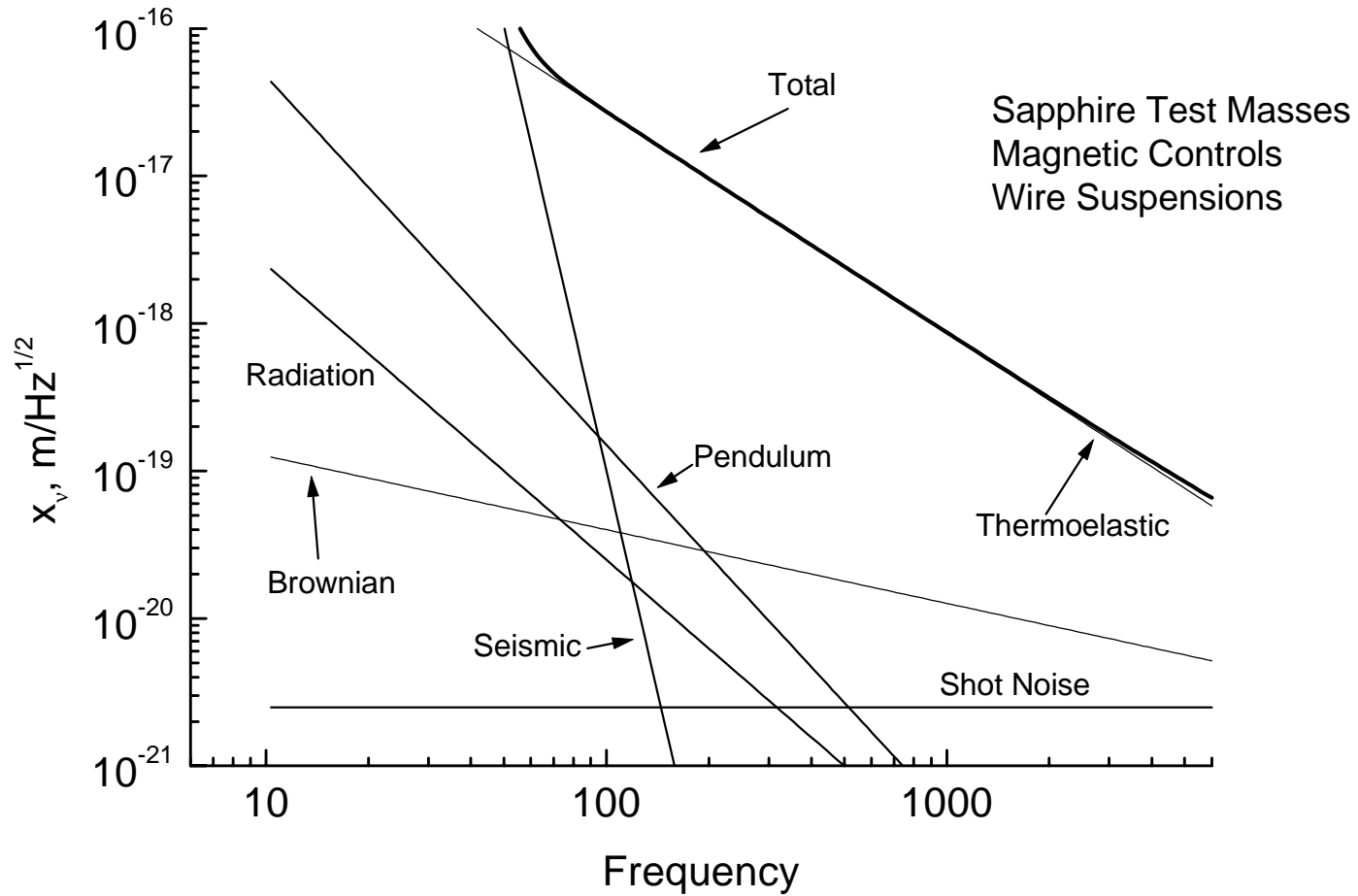
## TNI Phase I Test Masses



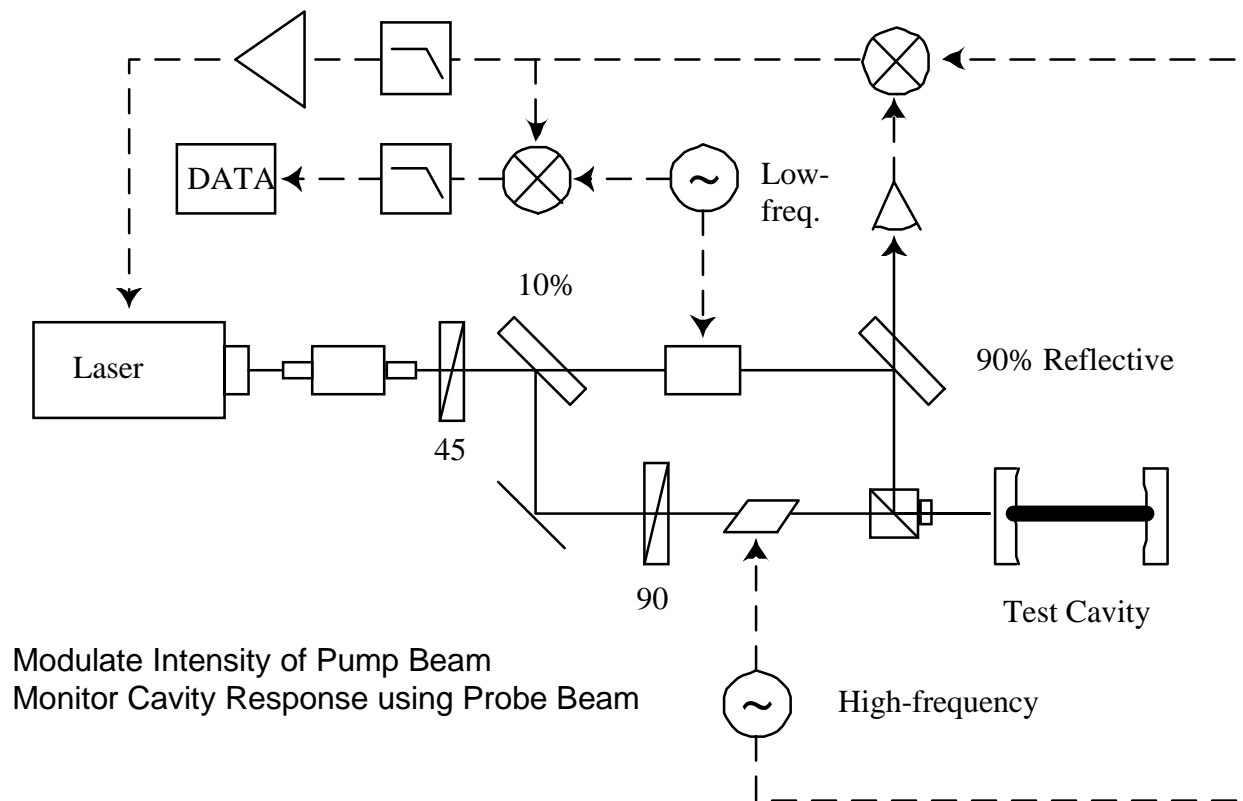
# TNI Phase I Expected Spectrum



# TNI Phase II Expected Spectrum



# PhotoThermal Noise Measurement



# The International Thermal Noise Interferometer (ITNI)

## PRINCIPAL OBJECTIVE:

EXAMINE FUSED SILICA SUSPENSION NOISE

PENDULUM THERMAL NOISE NOT YET MEASURED

NON-GAUSSIAN EVENTS PROBABLY DOMINATE DATA ANALYSIS

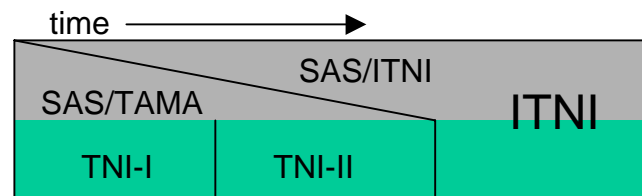
PHYSICS NOT WELL UNDERSTOOD

DETERMINE OPTIMAL STRESS IN FIBERS

## EXPERIMENT UPGRADE:

PUSH TO LOWER FREQUENCIES

COMBINE SAS AND TNI, PLUS COLLABORATORS



### LIGO/Caltech

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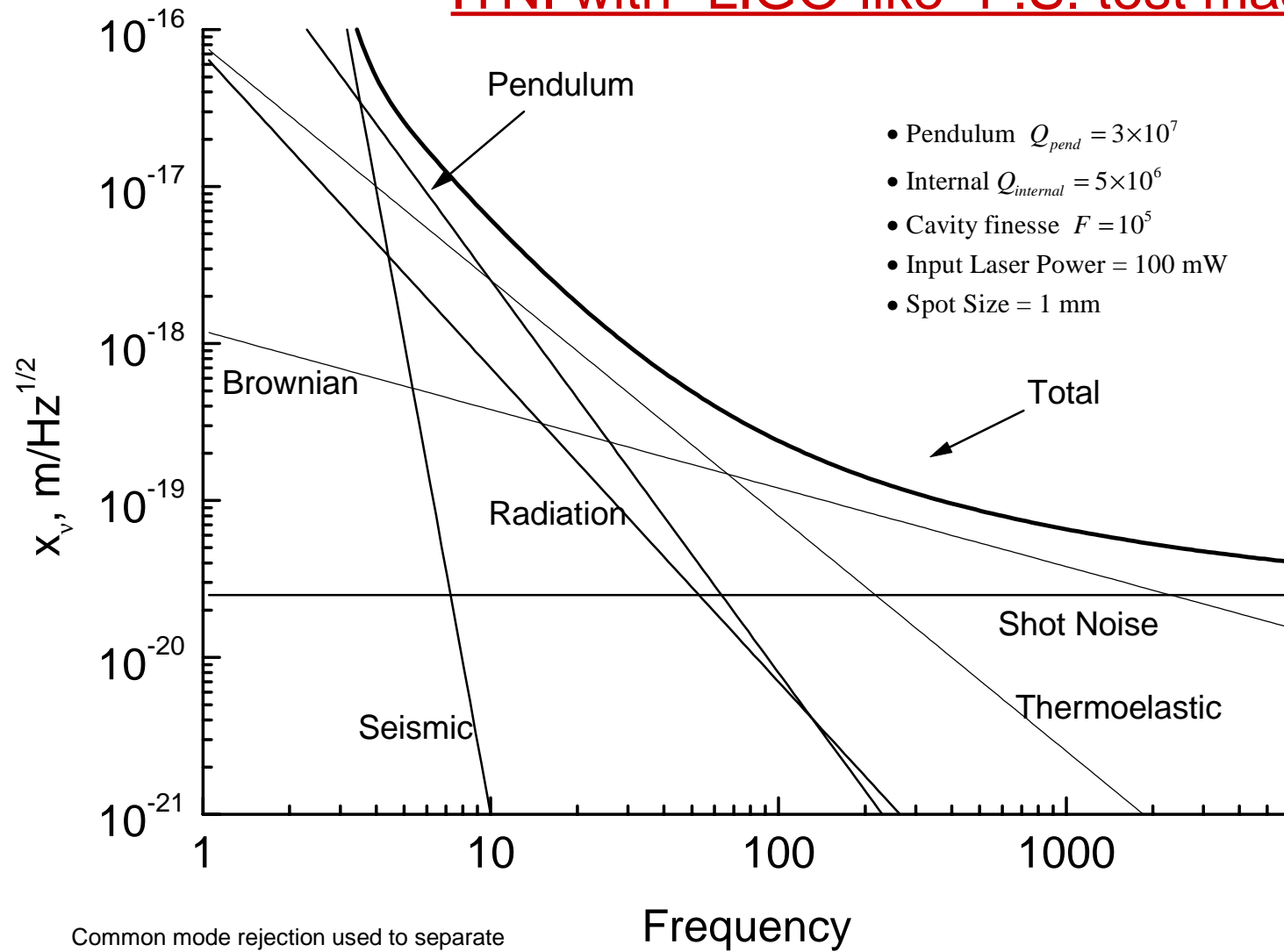
Flavio Vetrano  
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Seiji Kawamura



## ITNI with "LIGO-like" F.S. test masses



Common mode rejection used to separate seismic suspension noise from F.S. suspension noise



## ITNI with flat/flat cavities

